Technical Report 32-1180

## Determination of Footpad Penetration Depth from Surveyor Spacecraft Shadows



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# Determination of Footpad Penetration Depth from Surveyor Spacecraft Shadows 

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#### Abstract

A simple form of optical triangulation, using the Surveyor TV camera view angles and sun shadows, is used to determine the elevation of the lunar surface relative to the Surveyor spacecraft. A group of these elevation points, near the spacecraft footpads 2 and 3 , were calculated to indicate the footpad penetration depths. The limitations and potential applications of the method of analysis are discussed. A computer program designed to handle the shadow analysis calculations on the IBM 1620 computer is presented in Appendixes A, B, and C.


## Determination of Footpad Penetration Depth from Surveyor Spacecraft Shadows

## I. Introduction

More than 11,000 TV pictures were obtained during Surveyor Mission A. These TV pictures were taken by one fixed-position survey camera. Despite this voluminous collection of images, answers to many questions about the spacecraft penetration depth and the lunar surface elevation and range remain doubtful. This report presents an attempt to extract more definitive answers from these pictures.

The method of solution is a form of optical triangulation based on a knowledge of (1) the geometric relationship between the TV camera and other fixed elements of the Surveyor spacecraft, (2) elevation and direction of the sun as a function of time at the landing site, and (3) TV photographic images of spacecraft components and their shadows on the lunar surface.

Many of the Surveyor TV frames display useful shadow images. If knowledge of the other factors is correct, the shadowed surface could be located in three-dimensional space. A series of pictures of several components taken over an extended time period would give a family of data points that could be combined to form a simple elevation map of the lunar surface around the spacecraft. The
penetration depth of the Surveyor footpads could be estimated from this profile map.

The methodology for getting three-dimensional data from individual, two-dimensional TV pictures, known as the Surveyor Shadow Analysis, has progressed through three main stages. First, the general mathematical relationships needed to solve the problem were selected and test problems were tried. The solution uses analytical geometry and trigonometry. The general approach is simple; however, since directional vectors are involved, signs and directions must be checked carefully and kept consistent throughout the calculations.

Next, the solution method for the analysis of several typical TV pictures was applied. The Surveyor Mission A picture data file was scanned and 22 frames were selected. An electronic desk calculator was used for all numerical calculations. Results of this analysis are presented in later sections of this report.

This error analysis proved more complex than the original problem solution and much too time consuming for manual calculation. Therefore, the third step involved preparation of a computer program to handle the basic solution and also run the error analysis.

## II. Shadow Analysis Procedure

The basic technique used for the shadow analysis can be more clearly understood by referring to the first two illustrations. Figure 1 is a wide-angle photograph taken on Day 157 of Mission A. This frame shows the auxiliary battery box and its shadow on the surface of the moon. The rows and columns of small, reference reseau marks distributed over the frame should also be noted. The GMT time of transmission, as recorded in the TV data file, is also of importance.

The photographic data reduction provides the view angle, $V$, as seen from the TV position, between the image of the spacecraft object and its shadow on the lunar surface. This is determined by measuring the linear spacing between the object and its shadow in the photograph and converting this length into angular units by reference to the reseau marks. The angle between adjacent reseau marks is a calibrated function of the focal length of the TV camera lens

$$
\begin{equation*}
\text { Angle } V=\frac{(R F)(S W)}{R} \tag{1}
\end{equation*}
$$



Fig. 1. Surveyor I TV photograph showing the auxiliary battery and its shadow on the lunar surface
(Day: 157, Time: 13:59:34)
where

$$
\begin{aligned}
R F & =\text { reseau reference angle } \\
S W & =\text { length of shadow view line } \\
R & =\text { reseau mark spacing }
\end{aligned}
$$

The sun elevation angle and azimuth direction, in selenographic coordinates, must be determined for the GMT time of the TV frame. When the orientation of the spacecraft on the lunar surface has been defined, the sun angles can be related to the picture data by converting all angles into a common reference coordinate system.

The space geometry for the shadow analysis could have been related to any of several coordinate systems, including selenographic, TV camera centered or spacecraft centered. A modified spacecraft coordinate system was chosen to simplify the calculations. In this system, the vertical Z axis is equivalent to the normal spacecraft system. The $X Y$ plane, or the O level for the Z axis, has been positioned at the level of the lowest rigid spaceframe members. This plane also bisects the landing leg hinge centerlines. Positions above this plane have a positive value for $Z$. The $X$ axis, the $X Z$ plane, and the $Y Z$ plane are the same as in the normal spacecraft system. However, the $Y$ axis was reversed from the usual spacecraft orientation to simplify a problem in the computer program. Any position on the spacecraft or on the surface around it can be located by its unique XYZ coordinate values. Directional vector lines can be identified by their respective $\alpha, \beta$, and $\gamma$ direction cosine angles made with the $X, Y$, and $Z$ axes.

The geometric relationships between the TV camera and the shadow object are illustrated in Fig. 2. In the figure, the Surveyor spacecraft is shown in the postlanded position. The auxiliary battery box is the spacecraft component used in this example, and its shadow can be seen on the ground below. If straight lines were drawn between the TV camera $(V)$, the outboard corner of the battery box $(N)$, and its shadow on the surface ( $D$ ), they would form a plane triangle that could be solved to locate the surface position at $D$.

The location of the TV camera elevation mirror reference point ( $X_{1} Y_{1} Z_{1}$ ) and the position of the battery box corner ( $X_{2} Y_{2} Z_{2}$ ) can be established in the modified spacecraft coordinate system. Assume that the view angle $V$, and the sun's direction angles have been determined from TV picture data. The straight line between the TV camera mirror $\left(X_{1} Y_{1} Z_{1}\right)$ and the spacecraft component considered $\left(X_{2} Y_{2} Z_{2}\right)$ is identified as $L$ and its length and


Fig. 2. Surveyor I spacecraft with shadow analysis triangle
direction cosine angles are obtained from the following equations:

$$
\begin{align*}
& \text { length } L=\left[\left(X_{2}-X_{1}\right)^{2}+\left(Y_{2}-Y_{1}\right)^{2}+\left(Z_{2}-Z_{1}\right)^{2}\right]^{1 / 2}  \tag{2}\\
& \cos \alpha=\frac{X_{2}-X_{1}}{L} \quad \text { (angle with } X \text { axis) }  \tag{3}\\
& \cos \beta=\frac{Y_{2}-Y_{1}}{L} \quad \text { (angle with } Y \text { axis) }  \tag{4}\\
& \cos \gamma=\frac{Z_{2}-Z_{1}}{L} \quad \text { (angle with } Z \text { axis) } \tag{5}
\end{align*}
$$

For a better understanding of the next step in the shadow solution, the geometry of the TV camera optical system must be considered. The front nodal point of the camera lens, represented by the apex of the viewing angle $V$, is located in the camera housing, several inches below the elevation mirror. Its exact location depends on the focal length and focus settings of the camera lens. The distance between the elevation mirror and the front nodal point can range from 11.4 in . for a narrow-angle $(100 \mathrm{~mm}), 4-\mathrm{ft}$ focus setting, to approximately 6.4 in . for a wide-angle ( 25 mm ), infinity focus combination (Fig. 3).

The image of the front nodal point $\left(X_{p} Y_{p} Z_{p}\right)$ is located an equivalent distance $\left(L_{p}\right)$ behind the elevation mirror and changes position each time the mirror is moved in


Fig. 3. Optical geometry of the survey TV camera
azimuth or elevation. The total distance from the nodal point ( $X_{p} Y_{p} Z_{p}$ ) to the object point $\left(X_{2} Y_{2} Z_{2}\right)$ is the combined sum of $L$ and $L_{p}$.

The value of $L_{p}$ is an independent input for each individual solution, and its value can be determined from a table of focus-focal length factors. However, for purposes of discussion, $L_{p}$ is assumed to be 6.5 in. for all wide-angle frames and 11 in . for narrow-angle frames.

If it is assumed that the camera optical axis passes through the elevation mirror pivot axis, then the line segments $L_{p}$ and $L$ will be parallel and will meet at $X_{1} Y_{1} Z_{1}$. The location of $X_{p} Y_{p} Z_{p}$ can then be determined by the equations

$$
\begin{align*}
& X_{p}=(\cos \alpha)\left(L+L_{p}\right)-X_{2}  \tag{6}\\
& Y_{p}=(\cos \beta)\left(L+L_{p}\right)-Y_{2}  \tag{7}\\
& Z_{p}=(\cos \gamma)\left(L+L_{p}\right)-Z_{2} \tag{8}
\end{align*}
$$

The sun's direction angles are often given in the spherical coordinates of the elevation above the horizon ( $\theta$ ) and the azimuth from the spacecraft $X$ axis $(\phi)$. These angles can be converted into polar form by the relations

$$
\begin{array}{ll}
\cos \alpha^{\prime}=\cos \theta \cos \phi & \text { (sun angle with } X \text { axis) } \\
\cos \beta^{\prime}=\cos \theta \sin \phi & \text { (sun angle with } Y \text { axis) } \\
\cos \gamma^{\prime}=\sin \theta & \text { (sun angle with } Z \text { axis) } \tag{11}
\end{array}
$$

It should be noted that the prime symbol is used for the sun angles. The length of line $N D$ is still unknown, but its direction in space must be the same as the sun vector line with angles $\alpha^{\prime}, \beta^{\prime}$, and $\gamma^{\prime}$. With the direction cosine angles for the two intersecting lines, VN and $N D$, their intersect angle ( $N$ ) can be calculated as follows:

$$
\begin{equation*}
\cos N=\cos \alpha \cos \alpha^{\prime}+\cos \beta \cos \beta^{\prime}+\cos \gamma \cos \gamma^{\prime} \tag{12}
\end{equation*}
$$

With angle $V$, this gives two angles and one side of a plane triangle

$$
\begin{equation*}
\text { angle } D=\pi-(N+V) \tag{13}
\end{equation*}
$$

Other parts of the triangle can be solved by the law of sines

$$
\left.\begin{array}{rl}
\frac{S}{\sin V} & =\frac{\left(L+L_{p}\right)}{\sin D}  \tag{14}\\
S & =\frac{\left(L+L_{p}\right) \sin V}{\sin D}
\end{array}\right\}
$$

Side $S$ is the length of the shadow line in threedimensional space. The coordinates for the shadow point ( $X_{3} Y_{3} Z_{3}$ ) can now be calculated by using a variation of the direction cosine formulas

$$
\begin{array}{ll}
\cos \alpha^{\prime}=\frac{X_{2}-X_{3}}{S} & X_{3}=X_{2}-S \cos \alpha^{\prime} \\
\cos \beta^{\prime}=\frac{Y_{2}-Y_{3}}{S} & Y_{3}=Y_{2}-S \cos \beta^{\prime} \\
\cos \gamma^{\prime}=\frac{Z_{2}-Z_{3}}{S} & Z_{3}=Z_{2}-S \cos \gamma^{\prime} \tag{17}
\end{array}
$$

All of the calculated solutions for $Z_{3}$ have a negative value, since the ground level is several inches below the reference $X Y$ plane. Conversion of $Z_{3}$ values into relative footpad penetration depths can be done by subtracting the normal elevation value for the bottom of the footpads.

This step completes the general procedure for the shadow analysis. It can be used with selected Surveyor TV pictures from Mission A or from subsequent Surveyor missions. The images of a known spacecraft component and its shadow on the lunar surface should be contained within the same TV frame to minimize the errors in the angle $V$. When the images are in adjacent frames, careful mosaic work can combine the two photographs for analysis. This limits the analysis of the surface to the area under or near the spacecraft on which shadows fall.

However, this is the area about which information was needed to determine the footpad penetration depth.

The TV azimuth and elevation angles were not used as data inputs. These angles were avoided for two reasons. The tilted camera axis would greatly complicate the calculations and require conversion into and out of the spacecraft coordinate system. The TV identification data, as recorded during Mission A, contained occasional dropouts and gross errors and some unexplained angular shifts that would cause large errors in the calculated results. Should these problems be eliminated by improved mission operations or post-mission data reduction, then the shadow analysis could be extended to more distant areas where shadows will be cast when the sun is at lower angles. In such cases, the TV camera pointing coordinates would locate the shadow and establish an angle $V$ even when the object, such as the solar panel, is not directly visible to the camera. This would be the next logical stage of development for the shadow analysis. A discussion of this phase of analysis will not be undertaken herein.

## III. Data Assumptions

The general approach for the shadow analysis is simple. The accuracy of the results depends on the care exercised in the selection and interpretation of the data sources. When numerical data from several sources must be evaluated and selected, it is inevitable that many assumptions that will influence the results will be made. Before discussing the results of calculations, based on data from Mission A, the assumptions that preceded them should be examined. The general solution equations were kept free of limiting factors that only apply to Mission A.

The first assumption concerns the orientation of the spacecraft on the lunar surface. For Mission A, the orientation angles adopted were those reported in the Hughes Aircraft Company memorandum on sun/earth positions (Ref. 1). Conclusions reached by Hughes Aircraft Company personnel were based on data inputs from the spacecraft gyros, the antenna and solar panel positions, and the TV camera star sightings. The slightly different angles recorded in the JPL Mission Report (Ref. 2) were apparently based only on the star sightings. The differences are small, but the deciding factor was that the Hughes Aircraft Company memorandum contained a complete list of sun and earth direction angles, calculated in spacecraft coordinates, which could be used with a minimum of additional conversion.

The variable focal-length TV lens is normally used at the wide-angle ( 25 mm ) and the narrow-angle ( 100 mm ) settings. It was assumed that the operation of the lens was normal and that all pictures were taken at one of these two positions. At these focal lengths, the angular spacing between the rows and columns of reseau marks (parameter RF in Eq. 1) is assumed to be 5 deg at WA and 1.25 deg at $N A$. These values would be limited to the footpad focus distances. The reseau angle varies with focus changes and would be approximately $10 \%$ larger at infinity focus. The effects of optical or geometric distortions in the camera lens were neglected. On later Surveyor missions, a test record of geometric distortion will be included as part of the camera calibration; this information, however, was not available for Mission A. The reference point for the TV camera was assumed to be the center of the elevation mirror hinge line. The coordinate values for this point were used as $X_{1}, Y_{1}$, and $Z_{1}$ in the equations.

The location of the TV camera ( $X_{1} Y_{1} Z_{1}$ ) and of other spacecraft components ( $X_{2} Y_{2} Z_{2}$ ), referenced to the modified spacecraft coordinate system, should ideally be based on measurements of the actual flight hardware or on the best documentation record of this hardware. However, such direct contact with assembled flight spacecraft is discouraged for obvious reasons. The other space frames and test vehicles available at the time of the search contained structural differences that would have been misleading. The assembly drawing recommended as the best source for the needed information was the Surveyor Spacecraft A-21 Configuration Drawing (Ref. 3). Scaling measurements from a print are usually discouraged for many reasons including that of paper shrinkage. However, this drawing contains a built-in linear scale that can be used to correct for some of the paper shrinkage. Therefore, for the purpose of the calculations on the Mission A data, the spacecraft location measurements were taken from a J-size print of the referenced drawing. The reduced microfilm copies contained noticeable distortion and would not be usable for this purpose.

The results could also be influenced by the final deflection angle of the landing legs after the spacecraft comes to rest. The legs normally are at an angle of 18 deg, as measured from a level surface. Leg angle pots provide a telemetry voltage that is transmitted after landing and that is calibrated during spacecraft checkout. However, the information from this source is not without its limitations. The pot voltage is not set for zero, but just measured when the legs are hanging free with no contact with the ground. The telemetry data have an
error band equivalent to $\pm 1$ deg. The telemetry data from Mission A could be interpreted as being somewhere between 16 and 18 deg , depending on the choice of tolerances.

Therefore, for purpose of these calculations, it was assumed that all three legs are resting at the normal angle of 18 deg . Different angles will affect the $Z_{2}$ values for the footpad and other components mounted on the lower portions of the leg. The errors in $Z_{3}$ would be mainly from the direct contribution of $Z_{2}$ in Eq. (17).

The top of footpad 2 is tilted, toe down and heel up, at an angle of approximately 6 deg referenced to the $X Y$ plane. This was determined experimentally by comparing shadow angles of pictures taken late in the lunar day. This assumed tilt angle affects the $Z_{2}$ value for portions of the footpad top.

## IV. Data Calculations and Results

Twenty-two TV pictures were selected from the Mission A data. These pictures were chosen to help define the area around footpads 2 and 3 . Some of these pictures contained two or more possible shadow data points so that 30 solutions were calculated. The shadows were cast by eight different spacecraft components. The numerical values assigned to each $X_{2} Y_{2} Z_{2}$ position, and the parameters calculated at several steps of the solutions, are tabulated in Tables 1, 2, and 3 . The modified spacecraft coordinate system and the definition for angular measurements are illustrated in Fig. 4. This figure also indicates the surface area included in the several solutions.

The interpretation of the view angle $V$ has a potential source of error that should be pointed out. This error concerns the reseau spacing width, parameter $R$, in

Table 1. Spacecraft I.D. parameters used in the shadow solutions ${ }^{\text {a }}$

| Spacecraft item | S/C positions |  |  | Distance and direction from TV camera |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\begin{aligned} & X_{2,}, \\ & \text { in. } \end{aligned}$ | $\begin{aligned} & \mathbf{Y}_{2}, \\ & \text { in. } \end{aligned}$ | $\begin{aligned} & Z_{2 r} \\ & \text { in. } \end{aligned}$ | $\begin{gathered} L_{1} \\ \text { in. } \end{gathered}$ | $\operatorname{Cos} \alpha$ | Cos $\beta$ | $\operatorname{Cos} \gamma$ |
| TV camera reference ( $X_{1} Y_{1} Z_{1}$ ) | 13.0 | 27.4 | 44.0 | - | - | - | - |
| Footpod 2 gas jet | 60.5 | 39.6 | $-9.0$ | 72.20 | 0.6578 | 0.1689 | -0.7340 |
| Footpad 3 gas jet | $-60.5$ | 39.6 | -9.0 | 91.43 | -0.8041 | 0.1334 | -0.5798 |
| Leg 2 lock fixture | 58.5 | 34.0 | -4.1 | 66.53 | 0.6842 | 0.0992 | -0.7233 |
| Leg 3 lock fixture | $-58.5$ | 34.0 | -4.1 | 86.42 | -0.8272 | 0.0764 | -0.5565 |
| Leg 2 TV target (top) | 57.0 | 36.3 | $-5.3$ | 66.67 | 0.6606 | 0.1336 | -0.7402 |
| Footpad 2, top-center | 66.3 | 38.4 | -13.6 | 79.2 | 0.6729 | 0.1388 | $-0.7272$ |
| Footpad 2, top (position 10) | 67.8 | 44.3 | $-13.9$ | 81.49 | 0.6724 | 0.2074 | $-0.7105$ |
| Footpad 2, top (position 3) | 69.2 | 33.0 | $-13.6$ | 80.66 | 0.6972 | 0.0694 | -0.7146 |
| Footpad 3, lop-center | -66.3 | 38.4 | $-13.6$ | 98.6 | -0.8042 | 0.1115 | -0.5841 |
| Footpad 3, top (position 3) | -65.0 | 43.8 | -13.6 | 98.33 | -0.7932 | 0.1667 | -0.5857 |
| a Positions on top of footpad are referenced as per clock positions. |  |  |  |  |  |  |  |

Eq. (1). For purposes of this report, the separations between the etched reseau marks on the face of the TV camera tube can be considered equal. The separations observed in a typical Surveyor $I$ photograph are noticeably unequal, with the difference approaching $10 \%$ in some frames. This distortion is a combination of all the nonlinear deflection and scan problems in the camera and the ground recording equipment. As a result, no one value for $R$ can be used for all sections of a Surveyor photograph. To reduce errors in the angle $V$, the reseau
marks measured are those that are over or adjacent to the portion of the picture area containing the shadow image. For the 22 pictures considered in this study, enlarged $8 \times 10-\mathrm{in}$. glossy prints and a pair of drafting dividers were used to make the data measurements. When a diagonal shadow ran between reseau marks with widely different row and column spacing, the calculated reseau diagonal spacing distance was used for parameter $R$. Careful attention at this stage of data reduction will keep the errors in $R$ below $1 \%$.

Table 2. Surveyor shadow study data on photographs near footpads 2 and 3

| I.D. | Day | GMT | Sun angles |  | Sun direction cosines |  |  | Solution |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $\begin{aligned} & \text { EI } \theta, \\ & \text { deg } \end{aligned}$ | $\begin{gathered} \text { Az } \phi, \\ \operatorname{deg} \end{gathered}$ | $\operatorname{Cos} \alpha^{\prime}$ | $\operatorname{Cos} \beta^{\prime}$ | $\operatorname{Cos} \boldsymbol{\gamma}^{\prime}$ |  |
| Area east of footpad 3 |  |  |  |  |  |  |  |  |
| 22124 | 161 | 09:43:05 | 51.7 | 264.7 | $-0.0573$ | -0.6171 | 0.7848 | 1 |
| 22125 | 161 | 09:43:11 | 51.7 | 264.7 | -0.0573 | -0.6171 | 0.7848 | 2 |
| 27627 | 161 | 14:42:53 | 49.13 | 265.05 | -0.0565 | -0.6519 | 0.7562 | 3 |
| 17264 | 161 | 15:01:18 | 49.0 | 265.1 | -0.0560 | -0.6537 | 0.7547 | 4 |
| 22614 | 162 | 10:21:32 | 39.18 | 266.15 | -0.0520 | -0.7733 | 0.6318 | 5 |
| 22615 | 162 | 10:21:40 | 39.18 | 266.15 | -0.0520 | -0.7733 | 0.6318 | 6 |
| 25247 | 162 | 17:11:12 | 35.7 | 266.43 | -0.0505 | -0.8105 | 0.5835 | 7 |
| 30233 | 163 | 10:46:17 | 26.78 | 267.1 | -0.0452 | -0.8915 | 0.4506 | 8 |
| 33166 | 164 | 12:45:26 | 13.58 | 267.83 | -0.0367 | -0.9713 | 0.2349 | 9 |
| Area east of footpad 2 |  |  |  |  |  |  |  |  |
| 22044 | 161 | 09:29:06 | 51.8 | 264.7 | -0.0571 | -0.6157 | 0.7859 | 10 |
| 22532 | 162 | 10:00:17 | 39.37 | 266.2 | $-0.0512$ | -0.7714 | 0.6343 | 11 |
| 30271 | 163 | 10:50:17 | 26.75 | 267.1 | $-0.0452$ | -0.8918 | 0.4501 | 12 |
| 34005 | 164 | 18:51:20 | 10.43 | 267.95 | -0.0352 | -0.9829 | 0.1811 | 13 |
| 34103 | 164 | 19:11:42 | 10.28 | 267.97 | -0.0349 | -0.9833 | 0.1785 | 14 |
| Area west of footpad 2 |  |  |  |  |  |  |  |  |
| 00722 | 153 | 07:43:09 | 29.4 | 91.2 | -0.0182 | 0.8710 | 0.4909 | 15 |
| 01324 | 153 | 09:45:06 | 30.42 | 91.5 | -0.0226 | 0.8621 | 0.5063 | 16 |
| 02151 | 154 | 04:09:59 | 39.78 | 92.2 | -0.0295 | 0.7679 | 0.6399 | 17 |
| 07115 | 155 | 10:55:18 | 55.3 | 94.1 | -0.0407 | 0.5678 | 0.8221 | 18 |
| 07201 | 156 | 06:22:34 | 65.2 | 96.5 | -0.0475 | 0.4167 | 0.9078 | 19 |
| 07225 | 156 | 06:41:30 | 65.38 | 96.6 | -0.0477 | 0.4137 | 0.9091 | 20 |
| 10347 | 156 | 11:22:35 | 67.77 | 97.7 | -0.0501 | 0.3751 | 0.9256 | 21 |
| 10414 | 156 | 11:29:49 | 67.82 | 97.7 | -0.0504 | 0.3742 | 0.8260 | 22 |

Table 3. Surveyor shadow data

| Solution | Item | $s w,$ in. | $\begin{aligned} & R F, \\ & \text { deg } \end{aligned}$ | $R_{r}$ <br> in. | $\begin{gathered} V_{1} \\ \text { deg } \end{gathered}$ | 5, in. | $x_{3}$ <br> in. | $\mathbf{Y}_{3}$ <br> in. | $\mathbf{z}_{3}$ in. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Footpad 3 shadow solutions |  |  |  |  |  |  |  |  |  |
| 1 | Footpad 3 gas jet | 1.375 | 5 | 1.344 | 5.11 | 10.599 | -59.89 | 46.14 | -17.31 |
| 2 | Footpad 3 gas jel | 1.406 | 5 | 1.375 | 5.11 | 10.594 | $-59.89$ | 46.13 | $-17.31$ |
| 3 | Footpad 3 gas jet | 6.125 | 3.75 | 4.125 | 5.56 | 12.024 | $-50.82$ | 47.43 | -18.09 |
| 4 | Footpad 3 gas jet | 6.219 | 5 | 5.50 | 5.65 | 11.681 | -59.84 | 47.23 | -17.81 |
| 5A | Footpad 3 gas jet | 1.813 | 5 | 1.313 | 6.90 | 13.914 | $-59.77$ | 50.36 | $-17.79$ |
| 5B | Leg 3 lock fixture | 2.906 | 5 | 1.313 | 11.06 | 21.180 | $-57.39$ | 50.38 | $-17.48$ |
| 6 | Footpad 3 gas jet | 1.844 | 5 | 1.344 | 6.86 | 13.819 | $-59.78$ | 50.28 | $-17.73$ |
| 7 A | Footpad 3 gas jet | 1.906 | 5 | 1.313 | 7.25 | 14.464 | -59.76 | 51.32 | -17.44 |
| 7B | Leg 3 lock fixture | 3.156 | 5 | 1.313 | 12.01 | 22.863 | -57.34 | 52.53 | $-17.44$ |
| 8 | Leg 3 lock fixture | 4.125 | 5 | 1.313 | 15.70 | 29.685 | -57.15 | 60.46 | $-17.47$ |
| 9 | Footpad 3, top (position 3) | 2.250 . | 5 | 1.313 | 8.56 | 17.131 | -64.36 | 60.44 | -17.62 |
| Footpad 2 shadow solutions |  |  |  |  |  |  |  |  |  |
| 10A | Footpad 2 gas jet | 1.156 | 5 | 1.344 | 4.30 | 9.225 | 61.02 | 45.28 | -16.24 |
| 10B | Leg 2 TV target | 1.75 | 5 | 1.344 | 6.51 | 13.187 | 57.75 | 44.42 | $-15.66$ |
| 11 A | Footpad 2 gas jet | 1.563 | 5 | 1.313 | 5.95 | 11.537 | 61.09 | 48.49 | -16.31 |
| 11 B | Leg 2 lock fixture | 2.75 | 5 | 1.313 | 10.47 | 18.849 | 59.46 | 48.54 | -16.05 |
| 11 C | leg 2 TV target | 2.406 | 5 | 1.313 | 9.16 | 16.910 | 57.86 | 49.34 | $-16.02$ |
| 12A | Leg 2 lock fixture | 4.094 | 5 | 1.344 | 15.23 | 25.673 | 59.65 | 56.89 | $-15.65$ |
| 12 B | Leg 2 TV target | 3.613 | 5 | 1.344 | 13.44 | 22.973 | 58.03 | 56.78 | -15.64 |
| 13A | Footpad 2, top (position 10$)$ | 1.656 | 5 | 1.281 | 6.46 | 11.149 | 68.19 | 55.25 | $-15.91$ |
| 13B | Footpad 2 (same) to ridge | 0.438 | 5 | 1.281 | 1.70 | 2.842 | 67.90 | 47.09 | $-14.41$ |
| 14A | Footpad 2, top (position 10) | 1.813 | 5 | 1.406 | 6.44 | 11.108 | 68.18 | 55.22 | -15.88 |
| 14B | Footpad 2 (same) to ridge | 0.469 | 5 | 1.406 | 1.66 | 2.770 | 67.89 | 47.02 | -14.39 |
| 15 | Footpad 2, top (position 3) | 0.844 | 5 | 0.969 | 4.35 | 7.138 | 69.33 | 26.78 | -17.10 |
| 16 | Footpad 2, top (position 3) | 0.781 | 5 | 0.969 | 4.02 | 6.632 | 69.34 | 27.28 | -16.95 |
| 17 | Footpad 2, top (position 3) | 0.844 | 5 | 1.313 | 3.21 | 5.551 | 69.36 | 28.73 | $-17.15$ |
| 18 | Footpad 2, top (position 3) | 0.613 | 5 | 1.281 | 2.39 | 4.590 | 69.38 | 30.39 | $-17.37$ |
| 19 | Footpad 2, top (position 3) | 0.531 | 5 | 1.313 | 2.02 | 4.187 | 69.39 | 31.25 | $-17.40$ |
| 20 | Footpad 2, top (position 3) | 2.063 | 1.25 | 1.25 | 2.06 | 4.503 | 69.41 | 31.13 | -17.69 |
| 21 | Footpad 2, top (position 3) | 0.469 | 5 | 1.219 | 1.92 | 4.068 | 69.40 | 31.47 | -17.36 |
| 22 | Foatpad 2, top (position 3) | 0.531 | 5 | 1.25 | 2.12 | 4.508 | 69.42 | 31.31 | -17.77 |



The final results of all the calculations are tabulated in the $X_{3}, Y_{3}$, and $Z_{3}$ columns of Table 3. The data points are plotted to show their relative positions, in both the $X Y$ plane and the vertical profile view, in Figs. 5 through 10. In these figures, outlines of the footpads indicate
their position relative to the calculated surface points. As an additional aid in visualizing the area represented by these figures, several of the Surveyor I pictures can be studied in Ref. 4. The GMT times and picture numbers for examples of each area are noted in Table 4.


Fig. 5. Surveyor I shadow data - XY position of lunar surface points on east side of footpad 3


Fig. 6. Surveyor I shadow data - profile plot of lunar surface points on east side of footpad 3


Fig. 7. Surveyor I shadow data - XY position of lunar surface points on east side of footpad 2


Fig. 8. Surveyor I shadow data - profile plot of lunar surface points on east side of footpad 2

Before discussing the error probabilities, some of the general aspects and implications of these profiles should be considered. The vertical spread of the data points in each group is approximately 1 in . This spread could be caused by random errors in the solutions, actual variations in the vertical positions of the surface points, or a combination of both. Is it reasonable to conclude that the surface granular material in an undisturbed area could have such a large vertical irregularity? Reference 4
presents pictures of the nearby lunar surface taken with the low-angle sunlight of the lunar afternoon. The GMT times and numbers for some of these examples are noted in Table 4. These pictures suggest that the surface is very irregular, and that even larger variations could be considered reasonable.

The general position that is assumed for spacecraft orientation will naturally influence the calculated results.


Fig. 9. Surveyor I shadow data - XY position of lunar surface points on west side of footpad 2


Fig. 10. Surveyor I shadow data - profile plot of lunar surface points on west side of footpad 2

However, the shifts in the assumed position would have to amount to several degrees before they become a problem.

If the landing legs are not resting at an angle of 18 deg , as had been originally assumed, they are probably at a slightly smaller angle. Variations in the leg angle about the nominal 18 deg will change the footpad $Z_{2}$ value by the ratio of $0.7 \mathrm{in} . / \mathrm{deg}$. For example, reducing leg 3 from 18 to 17 deg would reduce the $Z_{2}$ for the footpad top from -13.6 to -12.9 in ., or a change of
more than $5 \%$. This will cause an equivalent change in $Z_{3}$ since it is the sum of $Z_{2}$ and the product of $S\left(\cos \gamma^{\prime}\right)$. However, the relative position of footpad 3 and the surface points will not change by more than $1.5 \%$ because $S\left(\cos \gamma^{\prime}\right)$ is less sensitive to changes in $Z_{2}$. In other words, the indicated elevation of the surface will be strongly dependent upon the assumed leg angle, but the indicated footpad penetration depth is less dependent on this factor.

## V. Error Analysis

The computer program and the IBM 1620 computer were used in the analysis of errors in the shadow solutions. This computer program is described in the appendixes.

The analysis consisted of calculating the magnitude and percentage changes in $X_{3}, Y_{3}$ and $Z_{3}$ when the input parameters are varied. The computer program can handle eight different percentage variations, or plus and minus four different levels, for the ten main input parameters. This range of percentage levels was chosen to include the plus and minus values of (A) the smallest measurable increment of each parameter, (B) and (C) two separate choices of the probable input errors (maximum and minimum) for each parameter, and (D) a larger, gross error level which exceeds all probable parameter errors. The computer considers each parameter variation as a separate problem and calculates all outputs. This amounts to 80 problem calculations for each solution. However, the computer completes the total solution in seconds.

The computer prints out the results of all these error calculations as columns of percentages and final values for $X_{3}, Y_{3}$, and $Z_{3}$. The final step of selecting a combination of ten output errors (one for each parameter at the chosen input error level) and calculating the RMS or algebraic sum of the individual errors, is left as a manual calculation for the investigator.

The computer program could be refined by including the calculation of the combined RMS errors. However, this would require the selection of a particular set of individual error percentages that could be applied to all solutions. At this time, the range of possible values has not been narrowed sufficiently to mechanize this step.

A shadow solution was selected from each of the areas under consideration to demonstrate the application of the computer error analysis. The resulting printouts are included in Appendix C. Solution 8 was east of footpad 3, solution 10A was east of footpad 2, and solution 15 was west of footpad 2 .

The input parameter variations picked for these computer error solutions are listed in Table 5. The calculated output errors are unique to these solutions and could not be generally applied to all shadow problems. However, they indicate the relative effects of each input parameter

Table 5. Input parameter variations for error analysis

| Parameter | Symbol | $\pm$ Percentage increments $^{\mathrm{a}}$ |  |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: |
|  |  | (A) | (B) | (C) | (D) |
| Sun elevation angle |  | 0.2 | 0.4 | 1 | 2 |
| Sun azimuth angle | $\phi$ | 0.1 | 0.2 | 0.5 | 2 |
| Reseau scale | $R$ | 0.5 | 1.0 | 2 | 5 |
| Shadow length | $S W$ | 0.5 | 1.0 | 2 | 5 |
| S/C measurements | $X_{1} X_{2} Y_{1} Y_{2} Z_{1} Z_{2}$ | 0.2 | 0.5 | 1 | 5 |

a The percentage values chosen for each parameter are representative of:
(A) The smallest increment that can be easily measured.
(B) The probable error level (minimum range).
(C) The probable error level (maximum range).
(D) A large error level to indicate gross effects on output.

Table 6. Results of computer error analysis

| Output parameter | RMS total errors in each range |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | (A) |  | (B) |  | (C) |  | (D) |  |
|  | \% | in. | \% | in. | \% | in. | \% | in. |
| Solution 8 |  |  |  |  |  |  |  |  |
| $\chi_{3}$ | 0.29 | 0.17 | 0.66 | 0.38 | 1.5 | 0.84 | 6.7 | 3.8 |
| $\gamma_{3}$ | 0.40 | 0.24 | 0.83 | 0.50 | 1.7 | 1.03 | 5.6 | 3.4 |
| $\mathrm{Z}_{3}$ | 0.67 | 0.12 | 1.36 | 0.24 | 2.8 | 0.50 | 8.2 | 1.4 |
| Solution 10A |  |  |  |  |  |  |  |  |
| $x_{3}$ | 0.20 | 0.12 | 0.50 | 0.31 | 1.02 | 0.62 | 5.05 | 3.1 |
| $r_{3}$ | 0.22 | 0.10 | 0.52 | 0.24 | 1.04 | 0.47 | 4.91 | 2.2 |
| $Z_{3}$ | 0.41 | 0.07 | 0.86 | 0.14 | 1.81 | 0.29 | 6.13 | 0.99 |
| Solution 15 |  |  |  |  |  |  |  |  |
| $x_{3}$ | 0.20 | 0.14 | 0.50 | 0.35 | 1.0 | 0.70 | 5.0 | 3.5 |
| $Y_{3}$ | 0.31 | 0.08 | 0.72 | 0.19 | 1.4 | 0.39 | 6.6 | 1.8 |
| $Z_{3}$ | 0.23 | 0.04 | 0.52 | 0.09 | 1.1 | 0.18 | 4.5 | 0.76 |

and the error sensitivity slope as the percentage is increased.

Table 6 presents a summary of the total RMS errors for these three shadow solutions. Each of the four percentage columns were calculated as a separate RMS total to indicate how each level of input errors can affect the final output results. In a typical error analysis, only one RMS total would be needed.

The effect of errors in the angle of the landing legs could be included as part of the $X_{2} Y_{2} Z_{2}$ increments; however, it would be better to keep the angle as a separate error input. As previously mentioned, the effect on $Z_{2}$ is about $0.7 \mathrm{in} . /$ deg change; however, this is a fixed bias on all solutions that will cause a common shift in all the $Z_{3}$ levels. Therefore, this bias error must be remembered although it is not specifically included in every total.

## VI. Conclusions and Potential Applications

The application of the shadow analysis to some typical Surveyor Mission A data was demonstrated by the 30 shadow solutions and the 3 error calculations. The calculated lunar surface elevations are plotted in Figs. 5 through 10. The penetration depth of the footpads is indicated by the relative position of the pads in the profile views.

Some important limitations in measuring actual footpad penetration should be pointed out. At best, the analysis is an attempt to determine the final resting position of the footpads and not the maximum penetration that occurred during the first landing contact. The closest shadow data points are still inches away from the edge of the footpads, so that penetration numbers are relative to the nearby surface and not to the actual soil under the pads.

With these limitations in mind, some probable conclusions about footpad penetration can be considered. The data profile in Fig. 6 would indicate an average penetration of 1 in . for footpad 3 . Figure 8 shows a partial profile through part of the ridge thrown up by footpad 2. The data profiles in Figs. 8 and 10 suggest penetration values of from 0.8 to 3 in . for footpad 2 . Would it have been even possible for footpads 2 and 3 to have penetrations as different as 3 to 1? Studies of the Surveyor I strain gauge data and landing dynamics by F. B. Sperling of JPL have ruled out this possibility. A large difference in penetration depth could be expected to produce noticeably different leg force levels. The strain gauge data indicate similar loads in all three legs (Ref. 5).

Therefore, it can be concluded that the data profile from the west side of footpad 2 (Fig. 10) is more indicative of the actual penetration of this pad. This is in the 0.8 - to $1.6-\mathrm{in}$. range. In this case, the difference between the two profiles, east and west of the pad, would indicate an actual difference in elevation, because of a shallow crater or rise. Again, a visual study of some of the Surveyor I pictures, as noted for footpad 2 in Table 4, can give a better understanding of the situation. Figures I-179 and I-180 (pp. 218 and 219, Ref. 4) would particularly be helpful in determining, by visual interpretation, the presence of a shallow depression in the area to the southwest of footpad 2. Mosaic 29 (p. 354, Ref. 4), combines these photographs and increases the perception of a depression. The pictures showing the east side are not as clear or as well lighted; however, the presence of a higher area seems a reasonable interpretation.

Some additional work would be necessary to explain the situation around footpad 2 ; however, the data support the conclusion that there is a significant variation in the surface elevation around the pad. However, there is still considerable doubt about the original level of the material directly under the footpad. It may have had a sloping surface. At present, it could be speculated that footpad 2 impacted in an area of slightly lower elevation and came to rest in contact with the eastern boundary of this area.

The computer program presented in Appendix B reduces the time required for calculations and data conversion to a minimum, and makes it possible to run many solutions rapidly. The material presented in the
main part of this report plus the appendix should be sufficient to assist interested investigators in conducting analyses of similar data. The shadow analysis method is suggested as a possible tool for photographic data reduction.

There are still hundreds of photographs from Surveyor Mission A, not considered in this report, that could be used for additional shadow analysis work. Similar pictures from Surveyor Missions C through $\mathbf{G}$ could also be analyzed by this method.

For a more complete analysis of the footpad interaction with the lunar surface, E. Christensen of JPL has proposed attaching several thin radial rods or whiskers to the tops of footpads 2 and 3 . These rods would cast shadows on the surface area adjacent to the footpads. The shadow analysis method would then make it possible (1) to draw cross-section profiles through the disturbed material ridge pushed up by the footpad, (2) to construct such disturbed material profiles on two, and maybe three, sides of the footpad, and (3) to determine footpad penetration relative to the material immediately adjacent to the footpad. This relatively small investment could produce a bonanza of useful scientific data. The shadow analysis program would be directly applicable for the reduction of photographic data of such whisker-type devices.

The computer program, the printout of the complete program, and the tabulated printout for the several computer solutions are discussed in Appendixes A, B, and C.

## References

1. Attitude Matrix and Sun/Earth Positions, Interdepartmental Correspondence, 2292/20, SC-10, Hughes Aircraft Company, El Segundo, Calif., Aug. 19, 1966.
2. Scientific Data and Results, Surveyor I Mission Report, Part II. Technical Report 32-1023, Jet Propulsion Laboratory, Pasadena, Calif., 1966.
3. Surveyor Spacecraft A-21 Configuration, Drawing 276806A (J-size edition), Hughes Aircraft Company, El Segundo, Calif.
4. Television Data, Surveyor I Mission Report, Part III. Technical Report 32-1023, Jet Propulsion Laboratory, Pasadena, Calif., 1966.
5. Sperling, F. B., and Garba, J. A., A Description of the Surveyor Lunar Landing Dynamics and an Evaluation of Pertinent Telemetry Data Returned by Surveyor I. Technical Report 32-1035, Jet Propulsion Laboratory, Pasadena, Calif., 1967.

## Appendix A

## Computer Program for Surveyor Shadow Analysis

## I. Introduction

To reduce the time and effort required for the calculation of shadow solutions and probable errors, a computer program was prepared by L. I. Busch, JPL's Section 314 (Computation and Analysis Section). The program, in FORTRAN II language, was written for the IBM 1620 (Monitor II) computer. The program can be also converted for the IBM 7094 computer.

The symbols, nomenclature, and calculations used in the computer analysis are the same as discussed previously and are summarized on the following pages. The methods for determining the various data input items are discussed in the main sections of this report.

A complete printout of the computer program is presented in Appendix B. This consists of the error analysis subroutine, which is read into the computer memory disc for semi-permanent storage, and the main shadow analysis program, which must be used with each batch of shadow calculations. Once the error analysis has been stored, it can be recalled by the computer whenever needed. The inclusion of the error analysis is an optional choice that is made at the time of each data run. The format that must be used in preparing data input cards is also illustrated in Appendix B. A total of 245 cards, plus 2 cards for each solution input, are used.

Appendix C presents the computer printout for Surveyor $I$ shadow solutions $8,10 \mathrm{~A}$, and 15. All input parameters and the calculation results are listed in these printouts. The error analysis then gives the output percentage changes and the new output values resulting from each input parameter change. The total error effects can then be obtained by combining a set of these individual parameter errors. The investigator is free to make his own choice of the most probable error magnitude for each parameter and to use addition or RMS summation of the combination. A list of RMS error calculations for these three solutions is given in Table 6.

The computer program can be used for the analysis of Surveyor shadow data or of any similar photographic data, assuming that the characteristics and limitations of this approach are fully considered. Although the dimensional unit of inches was used for the calculations, other units, such as metric units, could be used if they are kept consistent throughout a solution.

## II. Surveyor Shadow Analysis

## A. Inputs

## Symbol

$X_{1} Y_{1} Z_{1} \quad$ location of a point in space within an XYZ axis system
$X_{2} Y_{2} Z_{2}$ location of other points in the $X Y Z$ axis system (described by I.D. below)
$L_{p}$ distance from elevation mirror pivot axis to TV camera front nodal point
I.D. identification of item located at $X_{2} Y_{2} Z_{2}$

GMT GMT time (hr-min-sec)
$\theta$ sun elevation angle above $X Y$ plane
$\phi$ sun azimuth angle, from $+X$ axis
$R$ reseau mark spacing width
$R F$ reseau reference angle
SW length of shadow view line

## B. Outputs and Other Parameters

$X_{p} Y_{p} Z_{p}$ location of TV camera front nodal point for the camera azimuth, elevation and focal length being considered
$X_{3} Y_{3} Z_{3}$ location of a point in $X Y Z$ axis system
$S$ length of shadow
$L$ length of line from $X_{1} Y_{1} Z_{1}$ to each $X_{2} Y_{2} Z_{2}$
$\cos$ of $\alpha \beta \gamma \quad$ direction cosine angles from line $L$ to $X Y Z$ axes
$\cos$ of $\alpha^{\prime} \beta^{\prime} \gamma^{\prime}$ direction cosines of sun line to $X Y Z$ axes
$N$ intersect angle between line $L$ and sun line
$\cos N \quad$ cosine of angle $N$
Angles $V$ and $D \quad V$ is shadow view angle (angles $V, D$, and $N$ form triangle)

## Calculations

$$
\begin{align*}
\operatorname{angle}(V) & =\frac{(R F)(S W)}{(R)}  \tag{1}\\
L & =\left[\left(X_{2}-X_{1}\right)^{2}+\left(Y_{2}-Y_{1}\right)^{2}+\left(\mathrm{Z}_{2}-\mathrm{Z}_{1}\right)^{2}\right]^{1 / 2} \tag{2}
\end{align*}
$$

$\cos \alpha=\frac{X_{2}-X_{1}}{L}$
$\cos \beta=\frac{Y_{2}-Y_{1}}{L}$
$\cos \gamma=\frac{Z_{2}-Z_{1}}{L}$

$$
\begin{equation*}
X_{p}=(\cos \alpha)\left(L+L_{p}\right)-X_{2} \tag{5}
\end{equation*}
$$

$$
\begin{equation*}
Y_{p}=(\cos \beta)\left(L+L_{p}\right)-Y_{\underline{D}} \tag{6}
\end{equation*}
$$

$$
\begin{align*}
Z_{p} & =(\cos \gamma)\left(L+L_{p}\right)-Z_{2}  \tag{8}\\
\cos \alpha^{\prime} & =\cos \theta \cos \phi  \tag{9}\\
\cos \beta^{\prime} & =\cos \theta \sin \phi  \tag{10}\\
\cos \gamma^{\prime} & =\sin \theta
\end{align*}
$$

$\cos (N)=\cos \alpha \cos \alpha^{\prime}+\cos \beta \cos \beta^{\prime}+\cos \gamma \cos \gamma^{\prime}$

$$
\begin{equation*}
\text { Angle }(D)=\pi-(N+V) \tag{12}
\end{equation*}
$$

$$
\begin{equation*}
S=\frac{\left(L+L_{p}\right)(\sin V)}{(\sin D)} \tag{13}
\end{equation*}
$$

$$
\begin{equation*}
X_{3}=X_{2}-S \cos \alpha^{\prime} \tag{14}
\end{equation*}
$$

$Y_{3}=Y_{2}-S \cos \beta^{\prime}$
$\mathrm{Z}_{3}=\mathrm{Z}_{2}-\mathrm{S} \cos \gamma^{\prime}$

## Appendix B <br> Computer Program

0000000000,0000000000100001100001000010000000000000000000000000000000000000000 | 123 |
| :--- | 11111111111111111111111111111111111111111111111111111111111111111111111111111111


 $44444444444444^{\prime \prime} 4444$ ! 444444444444444444444444444444444444444444444444444444444444 5555555555555555555555555555555555555555555555555555555555555555555555555555555







| CARD | PARAMETER |
| :---: | :---: |
| A | $\theta$ (SUN EL) |
| B | $\phi$ (SUN AZ) |
| C | $R$ (RESEAU) |
| D | SW (SHADOW) |
| E | $x_{1}$ |
| F | $x_{2}$ |
| G | $r_{1}$ |
| H | $r_{2}$ |
| I | $Z_{1}$ |
| $J$ | $Z_{2}$ |

DATA CARD I

0000000000000000000000000000000000.200000000000000000000000000000000000000000000

 222222222222222222222222222222222222222222222222222222222222222222222222222
 444444444444444444444444444444444 . 444444444,444444444444444444444 ; 444444 1444444 555555555555555 555555.55..5555555555555-55555555555555555555555555555555555555555
 77771777771777777777717771717777777777777717717777777177777177777771777717 88888888888886188888888 , $18888888881688888888188888388818888888881888388888168888 \quad 1$

 (FIELD LOCATIONS)

DATA CARD 2


000000000000000000000000000000000000000000000000000030000000000003307500030000
 © ill $1111111111111111111111111111111111111111111111111111111111 / 111111111$









* RIGHT ADJUST
** LEFT ADJUST


```
    8 FORMAT(26X1HL,9X1HN,9X1HD,9X1HV,9X1HS,18\times2HX3,8\times2HY3,8\times2HZ3,/19X
        1 4F10.2,F11.3,9\times3F10.2./1
    10 FORMAT(10A1,7F10.0)
    11 FORMAT (413,8\times5F10.0,5\times5A1)
    12 FORMAT(76HIF ERROR ANALYSIS IS DESIRED,PUT SENSE SWITCH 3 ON,AND P
        IUSH START ON CONSOLE/ 23HIF NOT, JUST PUSH START.)
    13 FORMAT (8F5.0)
    14 FORMAT (25X2HLP,58\times2HXP,8\times2HYP, 8X2HZP,/19XF10.2,50\times3F10.2,//)
    C
        DIMENSION ID(10), CODE(5),THETAE(8),PHIE(8),RE(8),SWE(8), X1E(8),
        1 Y1E(8),Z1E(8),X2E(8),Y2E(8),Z2E(8)
        COMMON THETAE,PHIE,RE,SWE,X1E,Y1E,Z1E,X2E,Y2E,Z2E,CONVF,ELP
        CONVF=.01745329
        READ 13,THETAE,PHIE,RE,SWE,XIE,Y1E,Z1E,XZE,Y2E,Z2E
    15 PRINT 4
        TYPE 12
        PAUSE
    1 READ 10,10,ELP,X2,Y2,22,X1,Y1, 21
        READ 11,IDAY,IHR,MIN,ISEC,THETA,PHI,R,RF,SW,CODE
        EL =SQRT ((X2-X1)**2+(Y2-Y1)**2+(Z2-Z1)**2)
        COSA=(X2-X1)/EL
        COSB=(Y2-Y1)/EL
        COSG=(Z2-Z1)/EL
        XP=X2-COSA*(EL+ELP)
        YP=Y2-COSB*(EL+ELP)
        ZP=Z2-COSG*(EL+ELP)
        COSTH=COS(THETA*CONVF)
        SINTH=SIN(THETA*CONVF)
        COSPH=COS(PHI*CONVF)
        SINPH=SIN(PHI*CONVF)
        V=(RF*SW)/R
        SINV=SIN(V*CONVF)
        COSAP=COSTH#COSPH
        COSBP =COSTH*SINPH.
        COSGP=SINTH
        COSN=COSA*COSAP + COSB *COSBP + COSG*COSGP
        EN=ATAN(SQRT(1.0-COSN**2)/COSN)*57.29578 + 180.0
        D=180.0-EN-V
        SIND=SIN(D*CONVF)
        S=(EL+ELP)#SINV/SIND
        X3=X2-S*COSAP
        Y3=Y2-S*COSBP
        Z3=22-S*COSGP
        PRINT 5
        PRINT 6,IDAY,IHR,MIN,ISEC,THETA,PHI,R,RF,SW,CODE
        PRINT 7,1D,X2,Y2,Z2,X1,Y1,Z1
        PRINT 8,EL,EN,D,V,S,X3,Y3,Z3
        PRINT 14,ELP,XP,YP,ZP
        IF(SENSE SWITCH 3)2,3
    2.CALL ERRAN(THETA,PHI,R,SW,RF,X1,Y1,Z1,X2,Y2,Z2,X3,Y3,Z3)
    3 IF(SENSE SWITCH 3)15,1
        END
```

I. Sample Format of Input, Cards A-J, Percentage Increments for Error Analysis

```
    *LDISK
    SUBROUTINE ERRANITHETA,PHI,R,SW,RF,X1,Y1,Z1,X2,Y2,22,X3,Y3,23)
        DIMENSION THETAE(8),PHIE(8),RE(8),SWE(8),X1E(8),Y1E(8),Z1E(8),
        1 X2E(8),Y2E(8),Z2E(8)
        COMMON THETAE,PHIE,RE,SWE,XIE,Y1E,Z1E,X2E,Y2E,Z2E,CONVF,ELP
    C
        20 FORMAT (1H0,8X14HERROR ANALYSIS,6XI 7HVARYING PARAMETER,8X2OHPERCENT
        IAGE CHANGE IN, 16X13HNEW VALUES OF/53\times2HX3, 8X2HY3,8\times2HZ3,10\times2HX3,8X
        2 2HY3,8X2HZ3)
90 FORMAT ( 31X5HTHETA,F6.2,4H PC, 3F10.3,3\times3F10.3)
    91 FORMAT ( }31\times5\textrm{HPHI},F6.2,4\textrm{H PC , 3F10.3,3\times3F10.3)
    92 FORMAT( }31\times5\textrm{HR},F6.2,4H\mathrm{ PC , 3F10.3,3X3F10.3)
    93 FORMAT(31X5HSW ,F6.2,4H PC ,3F10.3,3\times3F10.3)
    94 FORMAT ( 31 X5HX1 ,F6.2,4H PC, 3F10.3,3\times3F10.3)
    95 FORMAT ( }31\times5\mathrm{ HX2 ,F6.2,4H PC , 3F10.3,3X3F10.3)
    96 FORMAT (31X5HY1 , ,F6.2,4H PC, 3F10.3,3\times3F10.3)
    97 FORMAT(31X5HY2 ,F6.2,4H PC ,3F10.3,3\times3F10.3)
    98 FORMAT (31\times5HZ1 ,F6.2,4HPC,3F10.3,3\times3F10.3)
    99 FORMAT(31X5HZ2 ,F6.2,4H PC ,3F10.3,3\times3F10.3)
C
    PRINT }2
    THETAO=THETA
        PHIO=PHI
        RO=R
        SWO=SW
        x10=x1
        Y 10=Y1
        Z10=21
        x20=x2
        Y20=Y2
        Z20=22
        x30=x3
        Y30=Y3
        Z30=23
        K=0
    100 K=K+1
        I =0
    105 I = I +1
        IF(I-8)107,107,100
    107 GO TO(110,120,130,140,150,160,170,180,190,200),K
    110 THETA=THETA+.01*THETAE(I)*THETA
        GO TO 205
    120 PHI=PHI+.01*PHIE(I)*PHI
    GO TO 205
    130 R=R+.01*RE(I)*R
        GO TO 205
    140SW=SW+.01*SWE(I)*SW
        GO TO 205
    150 Xl= X1+.01* X1E(I)* X1
        GO TO 205
    160 X2=X2+.01*X2E(I)*X2
        GO TO 205
    170 Y1=Y1+.01*Y1E(I)*Y1
        GO TO 205
    180 Y 2=Y2+.01*Y2E(I)*Y2
        GO TO 205
    190 21=Z1+.01*Z1E(I)*21
        GO TO }20
    200 Z2=Z2+.01*Z2E(I)*Z2
C
    205 EL=SQRT((X2-X1)**2+(Y2-Y1)**2+(Z2-Z1)**2)
```


## II. Sample Format of Input - Two Cards Required for each Set of Data

```
    COSA=(X2-X1)/EL
    COSB=(Y2-Y1)/EL
    COSG=(Z2-Z1)/EL
    XP=COSA*(EL+ELP)-X2
    YP =COSR* (EL+ELP )-Y2
    ZP=COSG*(EL+ELP)-Z2
    COSTH=COS(THETA*CONVF)
    SINTH=SIN(THETA*CONVE)
    COSPH=COS(PHI*CONVF)
    SINPH=SIN(PHI *CONVE)
    V=(RF*SW)/R
    SINV=SIN(V*CONVF)
    COSAP =COSTH*COSPH
    COSBP=COSTH*SINPH
    COSGP=SINTH
    COSN=COSA*COSAP + COSB*COSBP + COSG*COSGP
    EN=ATAN(SQRT(1.0-COSN**2)/COSN)*57.29578+180.0
    D =180.0-EN-V
    SIND=SIN(D*CONVF)
    S=(EL+ELP)*SINV/SIND
    X3=X2-S*COSAP
    Y3=Y2-S*COSBP
    Z3=22-S*COSGP
    X3C }=(\times3/\times30)*100.0-100.0
    Y3C = (Y3/Y30)*100.0-100.0
    Z3C=(73/730)*100.0-100.0
C
G0 TO(210,220,230,240,250,260,270,280,290,300).K
    210 THETA=THETAO
    PRINT 90,THETAE(I),X3C,Y3C, Z3C, X3,Y3, 23
    GO TO 105
220. PHI=PHIO
    PRINT 91,PHIE(I),X3C,Y3C,Z3C,X3,Y3,Z3
    GO TO 105
    230 R=RO
    PRINT 92,RE(I),X3C,Y3C,Z3C,X3,Y3,Z3
    GO TO 105
    240 SW=SWO
    PRINT 93,SWE(I),X3C,Y3C,Z3C,X3,Y3,Z3
    GO TO }10
    250 X1=X10
    PRINT 94,X1E(1),X3C,Y3C,Z3C,X3,Y3,23
    GO TO 105
    260 <2 = 人20
    PRINT 95,X2E(I),X3C,Y3C,Z3C,X3,Y3,Z3
    GO TO 105
    270 Y1=Y10
        PRINT 96,Y1E(I),X3C,Y3C,Z3C,X3,Y3,Z3
    GO TO 105
    280 Y 2=Y20
        PRINT 97,Y2E(I),X3C,Y3C,Z3C,X3,Y3,Z3
        GO TO 105
    290 21=210
        PRINT 98,21E(I),X3C,Y3C,Z3C,X3,Y3,23
        GO TO 105
    300 22=220
        PRINT 99,Z2E(I),X3C,Y3C,Z.3C,X3,Y3,Z3
        IF(I-8)105,325,325
    325 RETURN
    END
```


## Appendix C

Surveyor I Shadow Analysis Data



III. Solution 15


