# Absorption Coefficients of $0_2$ at the Lyman-alpha Line

and its Vicinity

M. Ogawa Department of Physics, University of Southern California Los Angeles, California

## ABSTRACT

Using a 3-m vacuum spectrometer, absorption coefficients of  $0_2$  at the Lyman-alpha line and its vicinity have been remeasured between 1214.8 and 1217.8 A. Coefficients at the wavelengths of the apparent doublet peaks of the Lyman-alpha line, caused by selfabsorption and separated by 0.09 A, are pressure independent at relatively low pressure (p<6.3 torr). The absorption coefficient at the shorter wavelength peak is 0.304 ± 0.002 cm<sup>-1</sup> and that at the longer wavelength peak 0.278 ± 0.002 cm<sup>-1</sup>. The minimum of the transmission window is located at about 1216.0 A and its value is 0.254 cm<sup>-1</sup>.

(THRU) 8 (CODE) FORM FACILITY (CATEGORY)



## 1. INTRODUCTION

Absorption coefficients of  $0_2$  at the Lyman-alpha line, Ly- $\alpha$ , and its vicinity are of considerable importance in the study of the upper atmosphere, because the Ly- $\alpha$  line is located in one of the deepest transmission windows of  $0_2$  and because the observed width of the solar Ly-a line at half-maximum is very large, of the order of 1.0 A (Purcell et al., 1960). The coefficients in this region have been studied by many investigators (Preston, 1940; Watanabe et al., 1953; Ditchburn et al., 1954; Lee, 1955; Watanabe et al., 1958; Metzger et al., 1964; Shardanand, 1967). Although the coefficients at the Ly- $\alpha$  line reported in the literatures agree to within  $\pm$  14%, the coefficients measured by a photoelectric method showed a linear pressure dependence. On the other hand, the coefficients measured by photographic methods did not show such a pressure effect. All previous photoelectric measurements were carried out at relatively high pressures of  $0_2$  (p>10 torr) by using relatively low resolution equipment (less than 0.1 A) with the Ly- $\alpha$  line or the hydrogen many line spectrum as a background.

In the present work, the absorption coefficients of  $0_2$  at the Ly- $\alpha$  line and its vicinity have been reinvestigated by using a high resolution spectrometer, relatively low pressures of  $0_2$  (p<6.3 torr), with the Ly- $\alpha$  line and separately an argon continuum as background.

#### 2. EXPERIMENTAL PROCEDURES

A 3-m vacuum monochromator equipped with a grating of 1200 lines/mm blazed at 2950 A was used. Reciprocal linear dispersions were 1.42 A/mm in the second order and 0.95 A/mm in the third order spectrum.

The discharge tube was an ordinary water cooled  $\pi$ -shaped tube with a LiF window. For an argon continuum source, argon gas was continuously flowed through the tube and was excited by a neon sign transformer, 900 W, 15 KV, in combination with a 500 µµF capacitor and 3-mm air-blown spark gap. The argon pressure was maintained at 18 torr. For the Ly- $\alpha$  line source, a trace of H<sub>2</sub> gas was introduced into argon at about 4 torr. In this way, we were able to obtain a very clean Ly- $\alpha$  line both in the second and third order spectra without any spectral interference. An enlargement of the Ly- $\alpha$  line in the third order spectrum is shown in Fig. 1. As seen in this figure, the line shows the apparent doublet due to self-absorption, and these components were separated by 0.09 A. This separation could be changed by the mixing ratio of H<sub>2</sub> Ar and the discharge condition, but a separation of 0.09 A was adopted throughout the present experiments.

For the measurement of the absorption coefficients at the Ly- $\alpha$  line, an additional absorption tube, 15 cm long, was inserted between the entrance slit and the light source. By flowing  $0_2$  at 1.2 torr through this tube, it was used as a filter in order to minimize the dissociation of  $0_2$  along the absorption path within the spectrometer between the entrance and exit slits and thus suppress the formation of  $0_3$  as a secondary process. This filter action is, of course, due to the fact that  $0_2$  has very high absorption coefficients in the vacuum UV region, except for a few transmission windows, and thus filters out among other wavelengths any residual light from the source in the Schuman-Runge region.

The spectrometer itself was used as an absorption cell whose path length was 618 cm. Extra-dry oxygen gas, obtained from the Matheson Co., was introduced into the spectrometer, and its pressure was measured by a diaphragm type pressure gauge (CVC, GHD-100) and an oil manometer. Maximum pressures were 6.3 torr and 4.75

torr for the measurements with the  $Ly-\alpha$  line and the argon continuum, respectively.

The transmitted light was measured behind the exit slit by a photomultiplier tube, type EMI 9514S, attached behind a LiF window with sodium salicylate on its back surface. Current from the tube was amplified by an electrometer and monitored by a 10 mV stripchart recorder.

For the Ly- $\alpha$  line, the widths of both entrance and exit slits were adjusted to  $2l\mu$  which corresponded to a band width of 0.03 A in the second order and 0.02 A in the third order. For the argon continuum, both slits were adjusted to  $50\mu$  which corresponded to a 0.07 A band width in the second order.

Absorption coefficients, k, are defined by the relation  $\ln(I_o/I)$ = kx, where  $I_o$  and I are light intensities before and after absorption in a path length x reduced to the standard conditions of 0°C and 760 torr. The reduced path is given by  $x = lpT_o/Tp_o$ , where 1 is the actual cell length and the subscript o refers to standard conditions. In the pressure range of the present experiment, the absorption coefficient followed Beer's low within experimental error. Therefore, the coefficients were obtained from the slope of a straight line in a plot of  $\ln(I_o/I)$  vs p.

### 3. RESULTS AND DISCUSSIONS

As seen in Fig 1 (b), the intensity of the shorter wavelength peak of the apparent doublet of the Ly- $\alpha$  line is stronger than the other component. However, in Fig. 1 (a), the intensity of the shorter wavelength component decreases more rapidly than the other as  $0_2$ pressure increases. The measured coefficients at the peaks in the second and third order agreed within the experimental errors. The

values obtained are  $0.304 \pm 0.002$  cm<sup>-1</sup> and  $0.278 \pm 0.002$  cm<sup>-1</sup> at the shorter and longer wavelength peaks, respectively.

The coefficients of the transmission window from 1214.8 to 1217.8 A region are listed in Table 1 at 0.2 A intervals and are shown in Fig. 2 together with those at the Ly- $\alpha$  line. In the last column of the table, the absorption cross sections calculated from  $\sigma = k/n_{o}$  are also listed, where n is Loschmidt's number (2.687x10<sup>19</sup> molecules/ cm<sup>3</sup>). Wavelengths of the apparent doublet of the Ly- $\alpha$  line listed in the table were calculated by  $\lambda = \lambda \pm \frac{1}{2}\Delta\lambda$ , where  $\lambda_{o} = 1215.671$  A and  $\Delta\lambda = 0.09$  A, the separation of the apparent doublet components.

As shown in Fig. 2, the shape of the transmission window is very smooth, the shorter wavelength side is steeper than the other, and its minimum value is  $0.254 \text{ cm}^{-1}$  at around 1216.0 A. For comparison with the present results, the coefficients obtained by Lee (1955) and Watanabe et al. (1958) are also plotted in the same figure. Present values at the shorter wavelength side are larger and those at the longer wavelength side are smaller than the previous ones.

The center of the Ly- $\alpha$  line is located about 0.33 A below the minimum of the transmission window, and the coefficient of the shorter wavelength peak is about 10% higher than that of the other peak, even though their separation is only 0.09 A. Therefore, the apparent coefficient at the Ly- $\alpha$  line may change with the resolution of the apparatus and with the width of the Ly- $\alpha$  source line, if the resolution is not high enough to separate the apparent doublet components.

The measured absorption coefficients of  $0_2$  at the Ly- $\alpha$  line are summarized in Table 2, which includes the method of measurement and the pressure range of  $0_2$  used. In this table, (P) indicates that the coefficient was pressure dependent. However, as also seen in the table, the pressure dependent measurements were carried out

at relatively high pressures compared with those which were pressure independent. The coefficients measured by Watanabe et al. (1958), or more recently by Shardanand (1967), increased at a rate of  $5.0 \times 10^{-4}$  cm<sup>-1</sup>/torr or  $4.5 \times 10^{-4}$  cm<sup>-1</sup>/torr, respectively. Therefore, the pressure dependence is too small to be detected at low pressures and the results shown in the table are consistent in regard to the observation of pressure effect.

Preston (1940) suggested that the pressure broadening of the rotational lines of the 0, absorption bands may account for the pressure dependence of the absorption coefficient. Recently, the pressure dependence has been ascribed by Shardanand (1967) to the formation of the  $0_4$  molecule. We are not in a position to discuss the origin of the pressure dependence. However, we can give the position of the rotational lines of the 0, absorption band in the transmission window in question. Very recently, Alberti et al. (1968) and independently we have analyzed the rotational structure of the Tanaka progression (II), the forbidden transition  $\alpha^{1}\Sigma_{u}^{+} - X^{3}\Sigma_{g}^{-}$ , whose 3-0 band is located in this transmission window. Alberti et al. observed the rotational lines of the <sup>S</sup>R-branch up to N=15 and we observed them up to N=29. The observed and the calculated wavelengths of the rotational lines of the <sup>s</sup>R-branch which are located in the region of the present experiment, are the following: (N=33) [1215.61 A, (N=31)] [1216.16A, (N=29)<sub>obs</sub> 1216.69 A, (N=27)<sub>obs</sub> 1217.20A, (N=25)<sub>obs</sub> 1217.70 A, (N=23)<sub>obs</sub> 1218.20A. The position of these lines is indicated by their N-numbers in Fig. 2. These lines are quite sharp on the photographic plate (Kodak SWR), but they are not in evidence on the strip-chart trace in the present experiment with a band width of 0.07 A.

Acknowledgments. The author wishes to thank Professor G. L. Weissler and Professor K. Watanabe for their suggestions and discussions.

This research was supported in part by Contract No. NG2-05-018-044 of the National Aeronautics and Space Administration and by Contract No. 228(27) of the Office of Naval Research.

## REFERENCES

- 1. Alberti, F., R. A. Ashby and A. E. Douglas, Absorption Spectra of  $0_2$  in the  $a^{1}\Delta g$ ,  $b^{1}\Sigma g^{+}$ , and  $X^{3}\Sigma g$  States, Can. J. Phys. <u>46</u> 337, 1968.
- Ditchburn, R. W., J. E. S. Bradley, C. G. Cannon and G. Munday, Rocket Exploration of the Upper Atmosphere, Eds. R. L. E. Boyd, and M. J. Seaton, Pergamon Press, London. pp. 327-334, 1954.
- Lee Po, Photodissociation and Photoionization of Oxygen as Inferred from Measured Absorption Coefficients, J. Opt. Soc. Am. <u>45</u>, 703, 1955.
- Metzger, P. H., and G. R. Cook, A Reinvestigation of the Absorption Cross Sections of Molecular Oxygen in the 1050-1800 Å Region,
  J. Quant. Spectr. Radiative Transfer, 4, 107, 1964.
- 5. Ogawa, M., and K. Yamawaki, to be published.
- Preston, W. M., The Origin of Radio Fade-outs and the Absorption Coefficient of Gases for Light Wavelength 1215.7Å, Phys. Rev., <u>57</u>, 887, 1940.
- Purcell, J. D., and R. Tousey, The Profile of Solar Hydrogen-Lyman-α, J. Geophys. Research, <u>65</u>, 370, 1960.
- Shardanand, Absorption Coefficients of 0<sub>4</sub> at Oxygen Windows, Private Communication; see also NASA Technical Note, NASA TM D-4225, 1967.
- 9. Watanabe, K., C. Y. Inn, and M. Zelikoff, Absorption Coefficients of Oxygen in the Vacuum Ultraviolet, J. Chem. Phys., 21, 1026, 1953.
- Watanabe, K., J. Mottl, and H. Sakai, (unpublished work), see
  K. Watanabe, Advance in Geophysics, Eds. H. E. Landsberg, and
  J. Van Mieghem, Academic Press Inc., New York, pp. 183-184,
  1958.

## TABLES and FIGURES

Fig. 1. The Ly- $\alpha$  line spectrum in the third order.

- (a) Pressures of 0<sub>2</sub> are 0, 7, and 14 torr from the top. Exposure times are 90 seconds.
- (b) Exposure times are 60, 10, and 2 seconds from the top. Pressure of  $0_2$  is zero torr.
- Fig. 2. Absorption coefficients of 0 in the range of 1214.8-1217.8A Present work, o; Watanabe, Mottl, and Sakai, x; Lee
- TABLE 1. Absorption Coefficients of  $0_2$  at the Lyman-alpha Line and its Vicinity
- TABLE 2. Absorption Coefficients of  $0_2$  at the Lyman-alpha Line

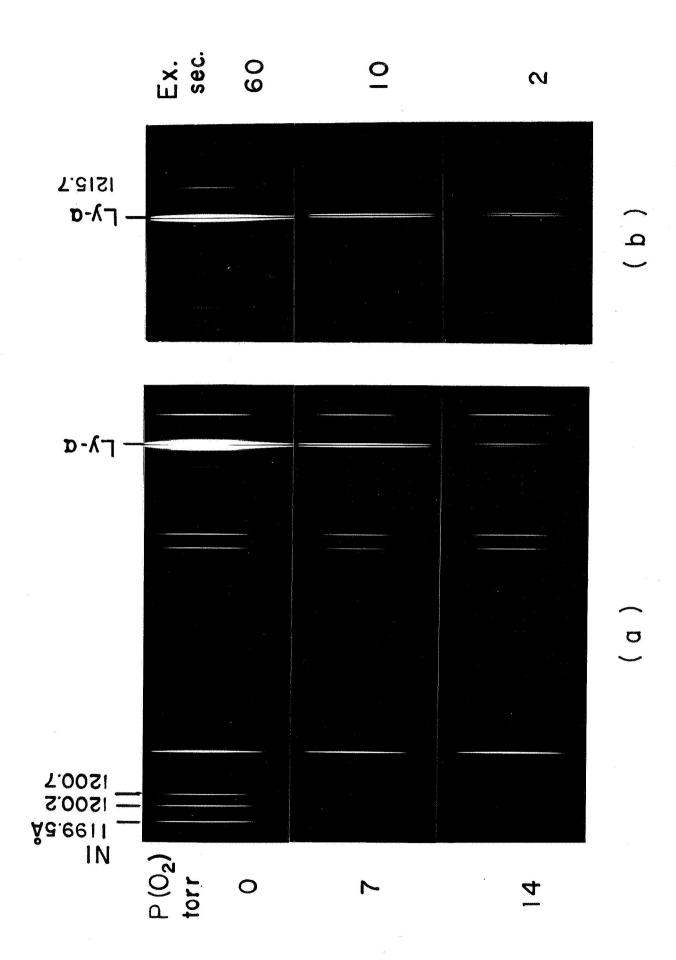
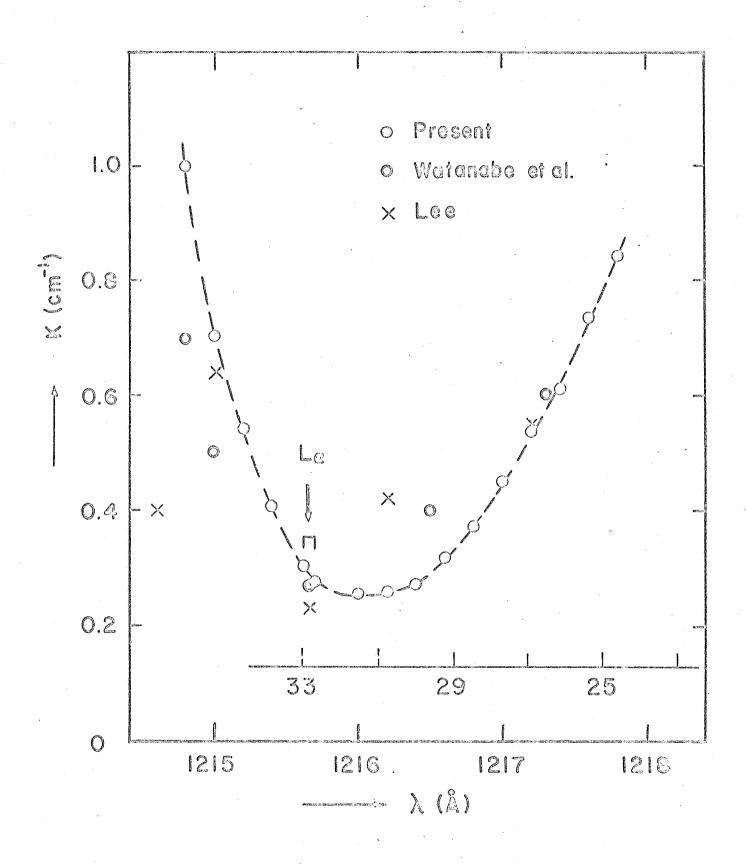


Fig. I



É

Fig. 2

		• *
λ(Α)	k(cm <sup>-1</sup> )	$\sigma(10^{-20} \text{ cm}^2)$
1217.8	0.840	3.12
1217.6	0.732	2.72
1217.4	0.610	2.27
1217.2	0.537	2.00
1217.0	0.450	1.67
1216.8	0.370	1.38
1216.6	0.317	1.18
1216.4	0.271	1.01
1216.2	0.259	0.963
1216.0	0.254	0.945
1215.72* 1215.63*	0.278 0.304	1.03 1.13
1215.4	0.409	1.52
1215.2	0.542	2.02
1215.0	0.704	2.62
1214.8	1.000	3.72

TABLE 1. Absorption Coefficients of 0<sub>2</sub> at the Lyman-alpha and its Vicinity

\* The Lyman-alpha line

C

			an a	المراجع المراجع المراجع المراجع
Investigator	k (cm <sup>-1</sup> )	σ (10 <sup>-20</sup> cr	pressure 2 (torr)	method
Preston (1940)	0.28(P) <sup>a</sup>	1.04	30-290	photoelectric, poly. <sup>b</sup>
Watanabe et al. (1953)	0.27(P)	1.00	70-490	${\tt photoelectric,mono.}^{{\tt b}}$
Ditchburn et al.(1954)	0.226	0.84	-20	photographic, poly.
Lee (1955)	0.23	0.85	-23	photographic, poly.
Watanabe et al. (1958)	0.27(P)	1.00	18-500	photoelectric, mono.
Metzger et al. (1964)	0.28(P)	1.04		photoelectric, mono.
Shardanand (1967)	0.29(P)	1.08	25-400	photoelectric, mono.
Present Work	$\{ \begin{smallmatrix} 0.278 \\ 0.304 \end{smallmatrix} \}$	1.03 1.13	1.2-6.3	{ photoelectric, poly. with $0_2$ filter

TABLE 2. Absorption Coefficients of  $0_2$  at the Lyman-alpha Line

a; (P) refers to observed pressure dependence. b; mono. and poly. indicate "monochromatic" or "polychromatic" light passing through the absorption cell.