## VIBRATION-ROTATION BANDS OF $\mathrm{AsH}_{3}$ <br> IN THE $2 \mu$ REGION

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## REPORT <br> by

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## ABSTRACT

The vibration-rotation bands of $\mathrm{AsH}_{3}$ in the $2 \mu$ region have tentatively been analyzed. The resonance effect, especially the $\nu_{1}$ and the $\nu_{3}$ bands for the K -type doubling and the Giant $l$-type doubling in the $4 \mu$ region have been described briefly. Some new combination difference relations in obtaining ground state constants from a perpendicular band have also been presented.

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# VIBRATION-ROTATION BANDS OF AsH3 IN THE $2 \mu$ REGION 

## I. INTRODUCTION

It seems from a preliminary investigation that there is a possibility of laser action from an $A s H_{3}-\mathrm{N}_{2}$ system. However, aside from other problems that might arise we must first of all have enough information on the molecular potential so that vibrational levels can be predicted to see if there is indeed a level close to that of $\mathrm{N}_{2}\left(\mathrm{v}_{2}=1\right)$.

Arsine ( $\mathrm{AsH}_{3}$ ) is a pyramidal-type molecule similar to ammonia with the three hydrogen atoms forming the base of the pyramid and the As atom located at the apex. There are four fundamental vibrationrotation bands, two of which (designated as $\nu_{1}$ and $v_{3}$ ) occur in the region of $5 \mu$ and the other two (designated as $\nu_{2}$ and $\nu_{4}$ ) are doubly degenerate and are observed in the region of $10 \mu$. All four fundamental bands plus some overtones have been studied in a series of reports by McConaghie and Nielsen ${ }^{1,2,3,4}$ in 1948 .with very low resolution.

In the past thirty years or so, many interesting features of the symmetric top molecule have been discovered. The inversion doubling, in $\mathrm{NH}_{3}$ has been studied in considerable details both experimentally ${ }^{5,6,7,8}$ and theoretically. ${ }^{9,10}$ The Coriolis interaction which splits the degenerate vibrational levels was first realized by Teller and Tisza ${ }^{11}$ in 1932. More recently, due to the availability of high resolution spectrometers the higher order interaction of two different vibrations due to Coriolis forces have been observed, namely the Giant $\ell$-type doubling and the K -type doubling. The theoretical developments are due to Nielsen ${ }^{12}$ and Hoffman ${ }^{13}$ respectively.

We shall report here the first clear observation of K -type doubling in a symmetric top molecule and the first observation of Giant $\ell$-type doubling in the $\nu_{3}$ band of $\mathrm{AsH}_{3}$. These doublings have been observed through the transitions of $\mathrm{R}_{\mathrm{P}}(\mathrm{J}, \mathrm{K})$ and $\mathrm{P}_{\mathrm{R}}(\mathrm{J}, \mathrm{K})$. We have also been able to derive some new combination difference relations by utilizing these transistions to obtain ground state molecular constants from a perpendicular band.

In the $2 \mu$ region there are three possible bands which lie close together. These are the $2 \nu_{1}$, a parallel band of species $A, \nu_{1}+\nu_{3}$, a doubly degenerate perpendicular band of species $E$, and $2 \nu_{3}$, a triply degenerate band of species $A+A+E$. Owing to the complex appearance of the spectrum, they have been only tentatively analyzed and we have not been able to obtain all the molecular constants.

## II. K-TYPE DOUBLING IN THE $\gamma_{3}$ BAND

The doublets of $P_{P(J, 3)}$ and $P_{R(J, 3)}$ transitions are indicated in the partial spectrum presented in Fig. 1. It can be seen from an energy level diagram that the spacing between the doublets $P_{P(J, 3)}$ and $P_{R(J-2,3)}$ should be equal if these splittings are due to the resonance effect of K-type doubling between $\nu_{1}$ and $\nu_{3}$. The experimental finding which is given below indicates that this is indeed the case.

TABLE I
OBSERVED K-TYPE DOUBLING SPLITTINGS

|  | $\mathrm{P}_{\mathrm{P}(8,3)} \cdot \mathrm{P}_{\mathrm{R}}(6,3)$ | $\mathrm{Pr}_{\mathrm{P}}(9.3), \mathrm{P}_{\mathrm{R}(7,3)}$ | $\mathrm{Pr}_{\mathrm{P}(10,3)}, \mathrm{P}_{\mathrm{R}(8.3)}$ | $\mathrm{P}_{\mathrm{P}(11,3)} \cdot \mathrm{P}_{\mathrm{R}(9,3)}$ |
| :---: | :---: | :---: | :---: | :---: |
| $\left\lvert\, \begin{gathered} \Delta v \\ \left(\mathrm{~cm}^{-1}\right) \end{gathered}\right.$ | 0.115, 0.090 | 0.173, 0.174 | $0.275,0.261$ | 0.402, 0.388 |

This doubling was indicated in the $v_{4} \mathrm{R}_{\mathrm{R}(\mathrm{J}, 3)}$ transitions of $\mathrm{PH}_{3}$ molecule in Hoffman's work. ${ }^{13}$ He found that the matrix elements which will split $K=2$ and $K=4 A+A$ levels are the following.

$$
\begin{align*}
& \left.<v_{s}-1, v_{t}, \ell_{t}, \mathrm{~K}\left|\mathrm{H}^{(2)} / \mathrm{Kc}_{\mathrm{c}}\right| \mathrm{v}_{\mathrm{s}}, \mathrm{vt}^{-1}, \ell_{\mathrm{t}} \mp, \mathrm{~K} \pm 2\right\rangle  \tag{1}\\
& =\left\langle\mathrm{v}_{\mathrm{s}}, \mathrm{v}_{\mathrm{t}}-1, \ell_{\mathrm{t}} \overline{+} 1, \mathrm{~K} \pm 2\right| \mathrm{H}^{(2)} / \mathrm{Kc}_{\mathrm{c}}\left|\mathrm{v}_{\mathrm{s}}-1, \mathrm{v}_{\mathrm{t}}, \ell_{\mathrm{t}}, \mathrm{~K}\right\rangle \\
& =\gamma\left[\mathrm{v}_{\mathrm{s}}\left(\mathrm{v}_{\mathrm{s}} \pm \ell_{\mathrm{t}}\right)\right]^{\frac{1}{2}}[(\overline{\mathrm{~J}} \mathrm{~K})(\mathrm{J}+\mathrm{K}-1)(\mathrm{J} \pm \mathrm{K}+1)(\mathrm{J} \pm \mathrm{K}+2)]^{\frac{1}{2}}
\end{align*}
$$

In the case of $P_{P(J, 3)}$ transitions the following determinant governs the splittings
(2)
$\left|\begin{array}{ccc}-\Delta+E & {[J(J+1)]^{\frac{1}{2}} \alpha} & {[J(J+1)(J+2)(J-1)]^{\frac{1}{2}} \gamma} \\ {[J(J+1)]^{\frac{1}{2}} \alpha} & 1+E+J(J+1) \beta & 0 \\ {[J(J+1)(J+2)} \\ (J-1)]^{\frac{1}{2}} \gamma & 0 & -2+E\end{array}\right|=0 \quad$.


Fig. 1. K-type doubling and Giant $\ell$-type doubling in $\nu_{3}$ of $\mathrm{AsH}_{3}$.

To determine the parameter $\gamma$ associated with K-type doubling, one could first employ the second order petturbation technique to obtain an approximate value of $\gamma$. Then by iteration one would eventually arrive at a best value for $\gamma$. which, in this case, comes out to be $1.91 \times 10^{-2} \mathrm{~cm}^{-1}$. These results are tabulated in Table II.

TABLE II
OBSERVED AND CALCULATED $P_{P(J, 3) ~ S P L I T T I N G S ~}^{\text {S }}$

|  | $P_{P(8,3)}$ | $P_{P(9,3)}$ | $P_{P(10,3)}$ | $P_{P(11,3)}$ | $P_{P(12,3)}$ | $P_{P(13,3)}$ | $P_{P(14,3)}$ | $P_{P(15,3)}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Observed <br> $\left(\mathrm{cm}^{-1}\right)$ <br> Calculated <br> $\left(\mathrm{cm}^{-1)}\right.$ | 0.115 | 0.174 | 0.275 | 0.402 | 0.614 | 0.832 | 0.995 | 1.174 |

## III. COMBINATION DIFFERENCE RELLATIONS <br> AND THE GIANT $\ell$-TYPE DOUBLING

The most desirable way of obtaining the ground state constants is to use certain combination relations in which the perturbation in the upper state energy levels is not involved. Therefore, in the past the ground state constants are obtained from parallel bands since the upper state of a parallel band is nondegenerate. However, if the transitions of $P_{R}(J, K)$ and ${ }^{R_{P}} P(J, K)$ are observed in a perpendicular band one can obtain some combination difference relations.which only involve the ground state levels. The fact that each pair of the transitions $\left[R_{P}(J+2, K), R_{R}(J, K)\right]$ and $\left[P_{R}(J-2, K), P_{P(J, K)}\right]$ do have a common upper state energy level leads to the following relations.

$$
\begin{align*}
& \frac{R_{R(J-1, K)}-R_{P(J+1, K)}}{2 J+1}=2 B^{\prime \prime}-4 D_{J}^{\prime}\left(J^{2}+J+1\right)-2 D_{J K}^{\prime \prime} K^{2}  \tag{3}\\
& \frac{P_{R(J-1, K)}-P_{P(J+1, K)}}{2 J+1}=2 B^{\prime \prime}-4 D_{J}^{\prime \prime}\left(J^{2}+J+1\right)-2 D_{J}^{\prime} K^{2}
\end{align*}
$$

The ground state constants obtained from $v_{3}(\perp)$ band by using the relations given by Eq. (3) and that obtained from $v_{2}$ (II) band are very close as can be seen in Table III.

TABLE III
GROUND STATE CONSTANTS OBTAINED
FROM $\nu_{2}$ AND $v_{3}$

|  | $\nu_{2}$ | $\nu_{3}$ |
| :--- | :--- | :--- |
| $\mathrm{~B}^{\prime \prime}\left(\mathrm{cm}^{-1}\right)$ | 3.7512 | 3.7510 |
| $\mathrm{D}_{\mathrm{J}}^{\prime \prime}\left(\mathrm{cm}^{-1}\right)$ | $9.4 \times 10^{-5}$ | $9.3 \times 10^{-5}$ |
| $\mathrm{D}_{\mathrm{JK}}\left(\mathrm{cm}^{-1}\right)$ | $-1.16 \times 10^{-4}$ | $-1.20 \times 10^{-4}$ |

In addition, these $P_{R(J, K)}, R_{P(J, K)}$ transitions provide the relations

$$
\begin{align*}
R_{R(J-1,0)} & +R_{P(J, 0)}=2\left[v_{o}-A^{\prime}+B^{\prime}\right]+4\left[A^{\prime}(1-\zeta)-B^{\prime}\right]  \tag{4}\\
& +2\left(B^{\prime}-B^{\prime \prime}-D_{J}^{\prime}+D_{J}^{\prime \prime}+\frac{1}{2} q\right) J^{2}-2\left(D_{J}^{\prime}-D_{J}^{\prime \prime}\right) J^{4}
\end{align*}
$$

and

$$
\begin{equation*}
\frac{R_{R(J, 0)}-R_{P(J, 0)}}{2 J+1}=\left(2 B^{\prime}+q\right)-4 D_{J}^{\prime}\left(J^{2}+J+1\right) \tag{5}
\end{equation*}
$$

which can be used in obtaining better values for the parameter $q$ of the Giant $\ell$-type doubling as well as the band center.

The following matrix element which give rise to Giant $\ell$-type doubling has been derived by Nielsen. ${ }^{12}$
(6)

$$
\begin{aligned}
&\left\langle\mathrm{v}_{\mathrm{s}}-1, \mathrm{v}_{\mathrm{t}}, \ell_{\mathrm{t}}, \mathrm{~K}\right| \mathrm{H}^{(1)} / / h_{\mathrm{c}}\left|\mathrm{v}_{\mathrm{s}}, \mathrm{vt}_{\mathrm{t}}-1, \ell_{\mathrm{t}} \pm 1, \mathrm{~K} \pm 1\right\rangle \\
&=\left\langle\mathrm{v}_{\mathrm{s}}, \mathrm{v}_{\mathrm{t}}-1, \ell_{\mathrm{t}} \pm 1, \mathrm{~K} \pm 1\right| \mathrm{H}^{(1)^{\prime} / h_{\mathrm{c}}}\left|\mathrm{v}_{\mathrm{s}}-1, \mathrm{v}_{\mathrm{t}}, \ell_{\mathrm{t}}, \mathrm{~K}\right\rangle \\
&= \frac{1}{2} \frac{\omega_{\mathrm{s}}+\omega_{\mathrm{t}}}{\left(\omega_{\mathrm{s}} \omega_{\mathrm{t}}\right)^{\frac{1}{2}}}\left[\mathrm{v}_{\mathrm{s}}\left(\mathrm{v}_{\mathrm{t}} \overline{+} \ell_{\mathrm{t}}\right)\right]^{\frac{1}{2}}[(\mathrm{~J} \overline{\mathrm{~K}})(\mathrm{J} \pm \mathrm{K}+1)]^{\frac{1}{2}} \\
&\left(\zeta_{\mathrm{st}}^{\mathrm{x}} \mathrm{~B}_{\mathrm{e}}\right) .
\end{aligned}
$$

The effect of these matrix elements is to split the $K=1$ upper state levels. However, owing to its selection rule, only one transition of the split levels is allowed. Therefore the appearance of this effect on the spectrum is the shift of $R_{R(J, 0)}$ and ${ }^{R_{P}(J, 0) \text { transitions. These shifts can be seen }}$ in Fig . 1 .

## IV. VIBRATION-ROTATION BAND IN THE $2 \mu$ REGION

The spectrum in this region is shown in Fig. 2. One can see that it is complex and puzzling as far as a complete analysis is concerned. As already mentioned in the introduction, there are three possible bands which lie very close to each other. For the convenience of the following discussion we shall list them again.
$2 \nu_{1}$ : a parallel band of species A
$\nu_{1}+v_{3}$ : a doubly degenerate perpendicular band of species $E$
$2 \nu_{3}$ : a triply degenerate band of species $A+A+E$ which has two components, one of which is a parallel band of species $A$, while the other is a doubly degenerate perpendicular band of species $E$. We shall designate them as $2 \nu_{3}(11)$ and $2 \nu_{3}(\perp)$ respectively.

From the theoretical point of view, one sees first of all that Fermi type resonance can occur between the parallel component of the $2 \nu_{3}$ band and the parallel band of $2 v_{1}$ through the cubic anharmonic force constants in the molecular potential. However, the perpendicular component of the $2 \nu_{3}$ band and the perpendicular band $\nu_{1}+\nu_{3}$ can give rise to Fermi type resonance only through very high order anharmonic terms. Therefore the observation of this effect is very unlikely. Secondly, according to 'John's rule all three bands will be coupled through Coriolis force.






Fig. 2. Absorption spectrum of $\mathrm{AsH}_{3}$ in the $2 \mu$ region.


Hence the $K=1,2,4$ upper state levels of $\nu_{1}+v_{3}$ band will split through the interactions with both $2 v_{1}$ and the parallel component of the $2 v_{3}$. However, there is no direct interaction between $2 \nu_{1}$ and $2 \nu_{3}(\perp)$ as one can easily see from the matrix elements given by Eqs. (1) and (6). Nonetheless, splitting of some of the levels of $2 \nu_{3}(\perp)$ will be possible through $\nu_{1}+\nu_{3}$, namely the $K=2,4,5$ levels, and $K=3$, and will be shifted. From group theory consideration, the $2 \nu_{3}\left({ }^{+}\right)$upper state rotational levels have the species given in Table IV in which p is any integer.

TABLE IV
SYMMETRY OF THE ROTATIONAL LEVELS
OF $2 v_{3}(1)$

| $\mathrm{K}=0$ | $\mathrm{~K}=3 \mathrm{p}$ | $\mathrm{K}=3 \mathrm{p}+2$ | $\mathrm{~K}=3 \mathrm{p}-2$ |
| :--- | :--- | :--- | :--- |
| E | $\mathrm{E}+\mathrm{E}$ | $\ell=2 \quad \mathrm{~A}_{1}+\mathrm{A}_{1}$ | $\ell=-2$ |
|  |  | $\mathrm{~A}_{1}+\mathrm{A}_{1}$ |  |
|  |  | $\ell=-2$ | E |

Accordingly, the splitting of $K=2,5$ levels will be seen at ${ }^{R_{R}(J, 1)}$ and $R_{R(J, 4)}$ transitions respectively and that of $K=4$ level will be seen at $P_{P(J, 5)}$ transitions. These predicted splittings are not clearly snown in the spectrum except perhaps in the one instance of $P_{P}(8,5)$ and $P_{P}(9,5)$ which correspond to line numbers $(541,542)$ and $(567,568)$ in Fig. 2.

In assigning the rotational quantum numbers we have found Eq. (3) to be extremely useful since the ground state constants are known from the analysis of the fundamental bands. Table $V$ gives all the lines that are tentatively identified.

It appears from the spectrum there is another band whose band center is around $4247 \mathrm{~cm}^{-1}$.

## v. CONCLUSIONS

The detailed analysis of the four fundamental bands of $\mathrm{AsH}_{3}$, their spectra and frequencies can be obtained in Reference 14. Table VI gives the molecular constants from this analysis. The vibration-rotation band in the $2 \mu$ region whose frequencies are given in Table VII is far from being completely analyzed. The author believes that further theoretical study must be done before thorough analysis can be attempted.

From the analysis of the four fundamentals of $\mathrm{AsH}_{3}$, it seems that the K-type doubling as well as the Giarlt l-type doubling can be observed more readily in $\nu_{3}$ than $\nu_{4}$ because of the fact that the K structure of $\nu_{1}$ which lies very close to $\nu_{3}$ does not spread out as does that of $\nu_{2}$ so that the rotational lines of $\nu_{3}$ are left clean. It is therefore interesting to further confirm the findings of this work by studying the $\nu_{1}$ and $\nu_{3}$ vibrationrotation bands of similar molecules, namely $\mathrm{PH}_{3}$.

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# TABLE .V <br> ASSIGNED TRANSITIONS OF THE VIBRATION-ROTATION BANDS OF $\mathrm{AsH}_{3}$ IN THE $2 \mu$ REGION 

| Serial <br> No. | Transition | Serial No. | Transition | Serial No. | Transition |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 213 | $\mathrm{R}(9, \mathrm{~K})$ | 328 | $\mathrm{P}_{\mathrm{R}}(2,1)$ | $\{494$ | $\mathrm{P}_{\mathrm{P}(6,3)}$ |
| 235 | $\mathrm{R}(8, \mathrm{~K})$ | 330 | $\mathrm{R}_{\mathrm{R}(2,0)}$ | 495 |  |
| 255 | $\mathrm{R}(7, \mathrm{~K})$ | 331 | $\mathrm{R}_{\mathrm{R}(2,1)}$ | 501 | $\mathrm{R}_{\mathrm{P}(6,0)}$ |
| 274 | $\mathrm{R}(6, \mathrm{~K})$ | 332 | $\mathrm{R}_{\mathrm{R}(2,2)}$ | 502 | $\mathrm{R}_{\mathrm{P}(6,1)}$ |
| 294 | $\mathrm{R}(5, \mathrm{~K})$ | 337 | $\mathrm{P}_{\mathrm{R}(1,1)}$ | 503 | $\mathrm{R}_{\mathrm{P}(6,2)}$ |
| 308 | $\mathrm{R}(4, \mathrm{~K})$ | 339 | $\mathrm{R}_{\mathrm{R}(1,0)}$ | 504 | $\mathrm{R}_{\mathrm{P}(6,3)}$ |
| 325 | $\mathrm{R}(3, \mathrm{~K})$ | 340 | $\mathrm{R}_{\mathrm{R}}(1,1)$ | 511 | ${ }^{P_{P(7,7)}}$ |
| 333 | $\mathrm{R}(2, \mathrm{~K})$ | 343 | $\mathrm{R}_{\mathrm{R}(0.0)}$ | 512 | $\mathrm{P}_{P(7,6)}$ |
| 342 | $\mathrm{R}(1, \mathrm{~K})$ | 406 | ${ }_{P}{ }_{P}(1,1)$ | 513 | $\mathrm{P}_{\mathrm{P}(7,5)}$ |
| 345 | $\mathrm{R}(0,0)$ | 437 | ${ }_{P}{ }_{P}(2,2)$ | 516 | PP(7, 4) |
| 410 | $\mathrm{P}(1,0)$ | 438 | ${ }^{P}{ }_{P}(2,1)$ | 520 | $\mathrm{P}_{P(7,3)}$ |
| 442 | $\mathrm{P}(2, \mathrm{~K})$ | 440 | $\mathrm{RPP}_{\mathrm{P}}(2,0)$ | 521 | $\mathrm{Pr}_{P}(7,2)$ |
| 451 | $P(3, K)$ | 443 | ${ }_{P} \mathrm{P}(3,3)$ | 526 | ${ }_{P} \mathrm{P}_{\mathrm{P}}(7,1)$ |
| 465 | $P(4, K)$ | 445 | ${ }^{P} P(3,2)$ | 536 | ${ }_{P} \mathrm{P}(8,8)$ |
| 485 | $\mathrm{P}(5, \mathrm{~K})$ | 446 | $\mathrm{P}_{\mathrm{P}}(3,1)$ | 537 | ${ }^{P} \mathrm{P}(8,7)$ |
| 506 | $\mathrm{P}(6, \mathrm{~K})$ | 447 | $\mathrm{R}_{\mathrm{R}} \mathrm{P}(3,0)$ | 538 | ${ }^{P} \mathrm{P}(8,6)$ |
| 530 | $P(7, K)$ | 448 | $\mathrm{R}_{\mathrm{P}} \mathrm{P}(3,1)$ | 541 | ${ }_{P} \mathrm{P}(8,5)$ |
| 557 | $\mathrm{P}(8, \mathrm{~K})$ | 456 | ${ }_{P} \mathrm{P}(4,4)$ | 544 | ${ }^{P} P(8,4)$ |
| 583 | $\mathrm{P}(9, \mathrm{~K})$ | 458 | ${ }^{P} \mathrm{P}(4,3)$ | 546 | ${ }^{P} \mathrm{P}(8,3)$ |
| 613 | $\mathrm{P}(10, \mathrm{~K})$ | 459 | $\mathrm{P}_{\mathrm{P}}(4,2)$ | 550 | ${ }^{P_{P}(8,2)}$ |
| 267 | $\mathrm{R}_{\mathrm{R}(6,0)}$ | 460 | ${ }_{P}{ }_{P}(4,1)$ | 562 | ${ }^{P_{P}(9,9)}$ |
| 278 | $\mathrm{P}_{\mathrm{R}(5,5)}$ | 461 | $\mathrm{P}_{\mathrm{R}(4,0)}$ | 564 | ${ }^{P_{P}(9,8)}$ |
| 281 | $\mathrm{P}_{\mathrm{R}(5,4)}$ | 462 | ${ }^{P_{R}(4,1)}$ | 565 | ${ }_{P}{ }_{P}(9,7)$ |
| 283 | $\mathrm{P}_{\mathrm{R}(5,3)}$ | 464 | $\mathrm{P}_{\mathrm{R}(4,2)}$ | 564 | ${ }_{P}{ }_{P}(9,6)$ |
| 286 | ${ }_{R_{R}(5,0)}$ | 469 | ${ }_{P}{ }_{P}(5,5)$ | 585 | ${ }^{P_{P}(10,10)}$ |
| 297 | ${ }_{P}{ }_{R}(4,4)$ | 471 | ${ }_{P}{ }_{P}(5,4)$ | 588 | ${ }^{P_{P}(10,9)}$ |
| 298 | $\mathrm{P}_{\mathrm{R}(4,3)}$ | 473 | $\mathrm{P}_{\mathrm{P}(5,3)}$ | 590 | ${ }^{P_{P}(10,8)}$ |
| 300 | $\mathrm{P}_{\mathrm{R}(4,2)}$ | 474 | ${ }_{P}^{P}(5,2)$ | 592 | ${ }^{P_{P}(10,7)}$ |
| 303 | $\mathrm{R}_{\mathrm{R}(4,0)}$ | 476 | $\mathrm{P}_{\mathrm{P}}(5,1)$ | 594 | ${ }_{P}{ }_{P}(10,6)$ |
| 307 | $\mathrm{R}_{\mathrm{R}(4,4)}$ | 478 | $\mathrm{R}_{\mathrm{P}(5,0)}$ | 608 | ${ }_{P}{ }_{P}(11,11)$ |
| 317 | $\mathrm{P}_{\mathrm{R}}(3,2)$ | 479 | $\mathrm{R}_{\mathrm{P}}(5,1)$ | 612 | $\mathrm{P}_{\mathrm{P}(11,10)}$ |
| 320 | $\mathrm{R}_{\mathrm{R}(3,0)}$ | 480 | $\mathrm{R}_{\mathrm{P}(5,2)}$ | 617 | $\mathrm{P}_{\mathrm{P}(11,9)}$ |
| 321 | $\mathrm{R}_{\mathrm{R}(3,1)}$ | 481 | $\mathrm{R}_{P(5,3)}$ | 620 | $\mathrm{P}_{\mathrm{P}(11,8)}$ |
| 322 | ${ }_{\mathrm{R}}^{\mathrm{R}(3,2)}$ | 489 | ${ }^{\text {P }} \mathrm{P}(6,6)$ | 621 | $\mathrm{P}_{\mathrm{P}}(11,7)$ |
| 323 | $\mathrm{R}_{\mathrm{R}(3,3)}$ | 490 | ${ }^{P} P(6,5)$ |  |  |
| 327 | $\mathrm{P}_{\mathrm{R}(2,2)}$ | 492 | $\mathrm{P}_{\mathrm{P}(6,4)}$ |  |  |

TABLE VI
MOLECULAR CONSTANTS FOR As $\mathrm{H}_{3}$ (in $\mathrm{cm}^{-1}$ )

|  | Ground State | $v_{1}$ | $v_{2}$ | $\nu_{3}$ | $v_{4}$ | $2 \nu_{1}$ | $2 v_{3}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| B | 3.7513 | 3.7108 | 3.5848 | 3.7153 | 3.886 | 3.639 | 3.688 |
| $\mathrm{D}_{\mathrm{J}}$ | $9.4 \times 10^{-5}$ | $1.10 \times 10^{-4}$ | $-1.13 \times 10^{-4}$ | $0.86 \times 10^{-4}$ |  | $-3.6 \times 10^{-4}$ |  |
| $\mathrm{D}_{\mathrm{JK}}$ | -1.18 $\times 10^{-4}$ |  | $3.71 \times 10^{-4}$ | $-2.19 \times 10^{-4}$ |  |  |  |
| $\mathrm{D}_{\mathrm{K}}$ | $1.58 \times 10^{-4}$ |  |  | $2.05 \times 10^{-4}$ |  |  |  |
| A | 3.4866 |  |  | 3.4273 | 3.571 |  |  |
| $D_{K}^{\prime}-D_{k}^{\prime}$ |  |  | $-2.66 \times 10^{-4}$ |  |  |  |  |
| $A^{\prime}-A^{\prime \prime}$ |  |  | 0.0243 |  |  |  | - . 0433 |
| (1-5) A |  |  |  | 3.5076 | 5.088 |  | 3.477 |
| q |  |  |  | -7.0 $\times 10^{-4}$ | 0.061 |  |  |
| Y |  |  |  | $1.91 \times 10-2$ |  |  |  |
| $\nu_{1}$ |  | 2115.191 |  |  |  |  |  |
| $v_{2}$ |  |  | 906.736 |  |  | . |  |
| $\nu_{3}$ |  |  |  | 2126.402 |  |  |  |
| $v_{4}$ |  |  |  |  | 999.409 |  |  |
| $2 v_{1}$ |  |  |  | . |  | 4167.01 |  |
| $2 v_{3}$ | - |  | 1 |  |  |  |  |

## TABLE VII <br> MEASURED FREQUENCIES OF $\mathrm{AsH}_{3}$ BANDS IN THE $2 \mu$ REGION

| Serial <br> No. | Wave No. $\left(\text { in vacc. } \mathrm{cm}^{-1}\right)$ | Serial <br> No | Wave No. $\text { (in vacc. } \mathrm{cm}^{-1} \text { ) }$ |
| :---: | :---: | :---: | :---: |
| 1 |  | 36 | 4269.53 |
| 2 |  | 37 |  |
| 3 | 4292.34 | 38 | 4268.99 |
| 4 | 4292.15 | 39 | 4268.84 |
| 5 | 4286.83 | 40 | 4268.54 |
| 6 | 4286.53 | 41 | 4268.38 |
| 7 | 4286.29 | 42 | 4268.25 |
| 8 | 4286.11 | 43 | 4267.98 |
| 9 | 4285.92 | 44 | 4267.75 |
| 10 | 4283.06 | 45 | 4267.38 |
| 11 | 4282.03 | 46 | 4267.25 |
| 12 | 4280.18 | 47 | 4267.05 |
| 13 | 4279.92 | 48 | 4266.68 |
| 14 | 4279.70 | 49 | 4266.42 |
| 15 | 4279.52 | 50 | 4266.30 |
| 16 | 4278.52 | 51 | 4266.22 |
| 17 | 4276.91 | 52 | 4265.90 |
| 18 | 4275:95 | 53 | 4265.55 |
| 19 | 4275.31 | 54 | 4265.35 |
| 20 | 4274.91 | 55 | 4265.25 |
| 21 | 4274.44 | 56 | 4265.09 |
| 22 | 4273.88 | 57 | 4264.90 |
| 23 | 4273.63 | 58 | 4264.46 |
| 24 | 4273.17 | 59 | 4264.21 |
| 25 | 4272.96 | 60 | 4264.03 |
| 26 | 4272.67 | 61 | 4263.79 |
| 27 | 4272.37 | 62 | 4263.53 |
| 28 | 4271.97 | 63 | 4263.36 |
| 29 | 4271.75 | 64 | 4263.22 |
| 30 | 4271.37 | 65 | 4262.88 |
| 31 | 4271.21 | 66 | 4262.60 |
| 32 | 4270.70 | 67 | 4262.46 |
| 33 | 4270.41 | 68 | 4262.17 |
| 34 | 4270.13 | 69 | 4261.98 |
| 35 | 4269.77 | 70 | 4261.78 |

## TABLE VII (Cont.)

| Serial No. | Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ ) | Serial No. | $\begin{gathered} \text { Wave No. } \\ \text { (in vacc. } \mathrm{cm}^{-1} \text { ) } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| 71 | 4261.64 | 111 | 4252.56 |
| 72 | 4261.34 | 112 | 4252.48 |
| 73 | 4261.17 | 113 | 4252.29 |
| 74 | 4260.97 | 114 | 4252.15 |
| 75 | 4260.68 | 115 | 4251.98 |
| 76 | 4260.22 | 116 | 4251.78 |
| 77 | 4260.01 | 117 | 4251.45 |
| 78 | 4259.67 | 118 | 4251.38 |
| 79 | 4259.50 | 119 | 4251.17 |
| 80 | 4259.20 | 120 | 4250.96 |
| 81 | 4258.85 | 121 | 4250.77 |
| 82 | 4258.61 | 122 | 4250.64 |
| 83 | 4258.46 | 123 | 4250.25 |
| 84 | 4258.15 | 124 | 4250.03 |
| 85 | 4257.92 | 125 | 4249.88 |
| 86 | 4257.66 | 126 | 4249.67 |
| 87 | 4257.66 | 127 | 4249.57 |
| 88 | 4257.54 | 128 | 4249.30 |
| 89 | 4257.35 | 129 | 4249.13 |
| 90 | 4256.95 | 130 | 4248.99 |
| 91 | 4256.69 | 131 | 4248.82 |
| 92 | 4256.59 | 132 | 4248.67 |
| 93 | 4256.36 | 133 | 4248.57 |
| 94 | 4256.22 | 134 | 4248.40 |
| 95 | 4255.96 | 135 | 4248.00 |
| 96 | 4255.75 | 136 | 4247.67 |
| 97 | 4255.70 | 137 | 4247.43 |
| 98 | 4255.35 | 138 | 4247.23 |
| 99 | . 4255.12 | 139 | 4247.03 |
| 100 | 4254.92 | 140 | 4246.85 |
| 101 | 4254.73 | 141 | 4246.77 |
| 102 | 4254.47 | 142 | 4246.49 |
| 103 | 4254.35 | 143 | 4246.29 |
| 104 | 4254.14 | 144 | 4246.07 |
| 105 | 4253.87 | 145 | 4245.82 |
| 106 | 4253.54 | 146 | 4245.69 |
| 107 | 4253.33 | 147 | 4245.55 |
| 108 | 4253.19 | 148 | 4245.28 |
| 109 | 4252.94 | 149 | 4245.08 |
| 110 | 4252.86 | 150 | 4244.87 |

## TABLE VII (Cont.)

| Serial No. | Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ ) | Serial <br> No. | Wave No. (in vacc. $\mathrm{cm}^{-1}$ ) |
| :---: | :---: | :---: | :---: |
| 151 | 4244.68 | 191 | 4236.12 |
| 152 | 4244.39 | 192 | 4235.89 |
| 153 | 4244.14 | 193 | 4235.64 |
| 154 | 4244.02 | 194 |  |
| 155 | 4243.80 | 195 | 4235.46 |
| 156 | 4243.72 | 196 | 4235.12 |
| 157 | 4243.47 | 197 | 4234.93 |
| 158 | 4243.20 | 198 | 4234.77 |
| 159 | 4243.10 | 199 | 4234.56 |
| 160 | 4242.80 | 200 | 4234.38 |
| 161 | 4242.74 | 201 | 4234.17 |
| 162 | 4242.63 | 202 | 4233.98 |
| 163 | 4242.42 | 203 | 4233.81 |
| 164 | 4242.25 | 204 | 4233.75 |
| 165 | 4242.11 | 205 | 4233.52 |
| 166 | 4241.79 | 206 | 4233.33 |
| 167 | 4241.63 | 207 | 4232.97 |
| 168 | 4241.52 | 208 | 4232.80 |
| 169 | 4241.24 | 209 | 4232.62 |
| 170 | 4241.05 | 210 | 4232.41 |
| 171 | 4240.79 | 211 | 4232.08 |
| 172 | 4240.55 | 212 | 4232.00 |
| 173 | 4240.38 | 213 | $4231.85 \mathrm{R}(9, \mathrm{~K})$ |
| 174 | 4240.14 | 214 | 4231.39 |
| 175 | 4239.98 | 215 | 4231.29 |
| 176 | 4239.85 | 216 | 4231.05 |
| 177 | 4239.63 | 217 | 4230.85 |
| 178 | 4239.33 | 218 | 4230.66 |
| 179 | 4239.11 | 219 | 4230.28 |
| 180 | 4238.85 | 220 | 4230.20 |
| 181 | 4238.62 | 221 | 4229.95 |
| 182 | 4237.68 | 222 | 4229.67 |
| 183 | 4237.43 | 223 | 4229.21 |
| 184 | 4237.31 | 224 | 4228.95 |
| 185 | 4237.24 | 225 | 4228.67 |
| 186 | 4236.97 | 226 | 4228.46 |
| 187 | 4236.75 | 227 | 4228.13 |
| 188 | 4236.57 | 228 | 4227.93 |
| 189 | 4236.44 | 229 | 4227.60 |
| 190 | 4236.35 | 230 | 4237.28 |

## TABLE VII (Cónt.)

| Serial No. | Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ ) | Serial No. | $\begin{gathered} \text { Wave No. } \\ \text { (in vacc. } \mathrm{cm}^{-1} \text { ) } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| 231 | 4226.85 | 271 | 4214.76 |
| 232 | 4226.74 | 272 | 4214.62 |
| 233 | 4226.56 | 273 | 4214.31 |
| 234 | 4226.30 | 274 | 4214.15 |
| 235 | 4226.16 | 275 | 4213.64 |
| 236 | 4225.88 | 276 | 4213.29 |
| 237 | 4225.54 | 277 | 4213.11 |
| 238 | 4225.09 | 278 | 4212.63 |
| 239 | 4224.94 | 279 | 4212.28 |
| 240 | 4224.74 | 280 | 4212.02 |
| 241 | 4224.30 | 281 | 4211.89 |
| 242 | 4223.97 | 282 | 4211.39 |
| 243 | 4223.63 | 283 | 4211.19 |
| 244 | 4222.97 | 284 | 4210.68 |
| 245 | 4222.73 | 285 | 4210.44 |
| 246 | 4222.59 | 286 | 4209.49 |
| 247 | 4222.41 | 287 | 4209.28 |
| 248 | 4222.25 | 288 | 4209.02 |
| 249 | 4222.12 | 289 | 4208.74 |
| 250 | 4221.88 | 290 | 4208.25 |
| 251 | 4221.68 | 291 | 4208.14 |
| 252 | 4221.55 | 292 | 4208.04 |
| 253 | 4220.82 | 293 | 4207.91 |
| 254 | 4220.37 | 294 | 4207.00 |
| 255 | 4220.28 | 295 | 4206.84 |
| 256 | 4219.66 | 296 | 4206.21 |
| 257 | 4219.49 | 297 | 4205.43 |
| 258 | 4219.27 | 298 | 4204.59 |
| 259 | 4219.11 | 299 | 4204.33 |
| 260 | 4218.86 | 300 | 4203.81 |
| 261 | 4218.16 | 301 | 4203.57 |
| 262 | 4217.79 | 302 | 4203.13 |
| 263 | 4217.47 | 303 | 4202.70 |
| 264 | 4217.25 | 304 | 4202.49 |
| 265 | 4216.77 | 305 | 4202.01 |
| 266 | 4216.58 | 306 | 4201.71 |
| 267 | 4216.10 | 307 | 4201.61 |
| 268 | 4215.95 | 308 | 4201.53 |
| 269 | 4215.81 | 309 | 4200.43 |
| 270 | 4215.40 | 310 | 4200.21 |


| Serial No. | Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ ) | Serial No. | $\begin{gathered} \text { Wave No. } \\ \text { (in vacc. } \mathrm{cm}^{-1} \text { ) } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| 311 | 4199.62 | 351 | 4172.26 |
| 312 | 4198.96 | 352 | 4172.96 |
| 313 | 4198.76 | 353 | 4171.74 |
| 314 | 4197.83 | 354 | 4171.56 |
| 315 | 4197.68 | 355 | 4171.39 |
| 316 | 4197.41 | 356 | 4171.00 |
| 317 | 4196.92 | 357 | 4170.80 |
| 318 | 4196.65 | 358 | 4170.53 |
| 319 | 4196.20 | 359 | 4170.27 |
| 320 | 4195.92 | 360 | 4170.13 |
| 321 | 4195.73 | 361 | 4169.83 |
| 322 | 4195.19 | 362 | 4169.64 |
| 323 | 4195.11 | 363 | 4169.41 |
| 324 | 4195.02 | 364 | 4169.33 |
| 325 | 4190.92 | 365 | 4169.23 |
| 326 | 4190.74 | 366 | 4169.02 |
| 327 | 4190.56 | 367 | 4168.96 |
| 328 | 4189.92 | 368 | 4168.75 |
| 329 | 4189.65 | 369 | 4168.48 |
| 330 | 4189.27 | 370 | 4168.30 |
| 331 | 4189.02 | 371 | 4168.06 |
| 332 | 4188.76 | 372 | 4167.83 |
| 333 | 4188.35 | 373 | 3167.63 |
| 334 | 4188.22 | 374 | 4167.38 |
| 335 | 4183.02 | 375 | 4167.08 |
| 336 | 4182.79 | 376 | 4166.87 |
| 337 | 4182.52 | 377 | 4166.64 |
| 338 | 4182.35 | 378 | 4166.28 |
| 339 | 4182.24 | 379 | 4166.10 |
| 340 | 4181.93 | 380 | 4165.74 |
| 341 | 4181.60 | 381 | 4165.57 |
| 342 | 4181.32 | 382 | 4165.33 |
| 343 | 4175.09 | 383 | 4165.22 |
| 344 | 4174.50 | 384 | 4165.02 |
| 345 | 4174.15 | 385 | 4164.86 |
| 346 | 4174.00 | 386 | 4164.72 |
| 347 | 4173.84 | 387 | 4164.50 |
| 348 | 4173.68 | 388 | 4164.35 |
| 349 | 4173.38. | 389 | 4164.15 |
| 350 | 4172.80 | 390 | 4163.93 |

## TABLE VII (Ćont.)

| Serial <br> No. | Wave No. $\text { (in vacc. } \mathrm{cm}^{-1} \text { ) }$ | Serial No. | Wave No. $\text { (in vacc. } \mathrm{cm}^{-1} \text { ) }$ |
| :---: | :---: | :---: | :---: |
| 391 | 4163.71 | 431 | 4155.46 |
| 392 | 4163.53 | 432 | 4155.22 |
| 393 | 4163.37 | 433 | 4154.97 |
| 394 | 4163.15 | 434 | 4154.74 |
| 395 | 4162.81 | 435 | 4154.41 |
| 396 | 4162.67 | 436 | 4153.91 |
| 397 | 4162.54 | 437 | 4153.54 |
| 398 | 4162.36 | 438 | 4153.04 |
| 399 | 4162.16 | 439 | 4152.73 |
| 400 | 4162.05 | 440 | 4152.56 |
| 401 | 4161.76 | 441 | 4151.96 |
| 402 | 4161.53 | 442 | 4151.62 |
| 403 | 4161.29 | 443 | 4146.33 |
| 404 | 4160.95 | 444 | 4146.04 |
| 405 | 4160.66 | 445 | 4145.77 |
| 406 | 4160.31 | 446 | 4145.27 |
| 407 | 4160.20 | 447 | 4144.72 |
| 408 | 4159.98 | 448 | 4144.38 |
| 409 | 4159.87 | 449 | 4144.07 |
| 410 | 4159.66 | 450 | 4143.88 |
| 411 | 4159.43 | 451 | 4143.79 |
| 412 | 4159.24 | 452 | 4140.24 |
| 413 | 4159.16 | 453 | 4140.02 |
| 414 | 4158.97 | 454 | 4139.62 |
| 415 | 4158.81 | 455 | 4139.35 |
| 416 | 4158.63 | 456 | 4138.96 |
| 417 | 4158.44 | 457 | 4138.52 |
| 418 | 4158.27 | 458 | 4138.38 |
| 419 | 4158.16 | 459 | 4137.97 |
| 420 | 4157.94 | 460 | 4137.39 |
| 421 | 4157.68 | 461 | 4136.76 |
| 422 | 4157.29 | 462 | 4136.49 |
| 423 | 4157.16 | 463 | 4136.22 |
| 424 | 4156.95 | 464 | 4136.07 |
| 425 | 4156.57 | 465 | 4135.84 |
| 426 | 4156.44 | 466 | 4135.71 |
| 427 | 4156.26 | 467 | 4133.53 |
| 428 | 4155.95 | 468 | 4132.23 |
| 429 | 4155.77 | 469 | 4131.53 |
| 430 | 4155.62 | 470 | 4130.98 |


| Serial No. | $\begin{gathered} \text { Wave No. } \\ \text { (in vacc. } \mathrm{cm}^{-1} \text { ) } \end{gathered}$ | Serial No. | Wave No. (in vacc. $\mathrm{cm}^{-1}$ ) |
| :---: | :---: | :---: | :---: |
| 471 | 4130.81 | 511 | 4116.40 |
| 472 | 4130.52 | 512 | 4115.62 |
| 473 | 4130.32 | 513 | 4114.87 |
| 474 | 4129.91 | 514 | 4114.64 |
| 475 | 4129.72 | 515 | 4113.98 |
| 476 | 4129.43 | 516 | 4113.88 |
| 477 | 4129.19 | 517 | 4113.45 |
| 478 | 4128.73 | 518 | 4113.34 |
| 479 | 4128.46 | 519 | 4113.10 |
| 480 | 4128.22 | 520 | 4112.70 |
| 481 | 4128.14 | 521 | 4112.13 |
| 482 | 4127.77 | 522 | 4111.92 |
| 483 | 4127.65 | 523 | 4111.86 |
| 484 | 4127.49 | 524 | 4111.78 |
| 485 | 4127.04 | 525 | 4111.71 |
| 486 | 4126.87 | 526 | 4111.40 |
| 487 | 4125.60 | 527 | 4110.81 |
| 488 | 4125.33 | 528 | 4110.69 |
| 489 | 4124.01 | 530 | 4110.55 |
| 490 | 4123.27 | 531 | 4110.26 |
| 491 | 4122.95 | 532 | 4110.07 |
| 492 | 4122.48 | 533 | 4109.95 |
| 493 | 4122.30 | 534 | 4109.60 |
| 494 | 4122.15 | 535 | 4108.81 |
| 495 | 4121.88 | 536 | 4108.72 |
| 496 | 4121.41 | 537 | 4107.90 |
| 497 | 4121.33 | 538 | 4107.10 |
| 498 | 4121.16 | 539 | 4106.85 |
| 499 | 4121.01 | 540 | 4106.66 |
| 500 | 4120.72 | 541 | 4106.27 |
| 501 | 4120.28 | 542 | 4106.06 |
| 502 | 4120.05 | 543 | 4105.75 |
| 503 | 4119.58 | 544 | 4105.52 |
| 504 | 4119.30 | 545 | 4105.01 |
| 505 | 4119.17 | 546 | 4104.65 |
| 506 | 4119.05 | 547 | 4104.48 |
| 507 | 4118.91 | 548 | 4103.87 |
| 508 | 4118.73 | 549 | 4103.71 |
| 509 | 4118.55 | 550 | 4103.59 |
| 510 | 4118.01 |  |  |

## TABLE VII (Cont.)

\(\left.$$
\begin{array}{lccc}\begin{array}{l}\text { Serial } \\
\text { No. }\end{array} & \begin{array}{c}\text { Wave No. } \\
\text { (in vacc. cm }\end{array} \\
& & \begin{array}{l}\text { Serial }\end{array} & \begin{array}{c}\text { Wave No. } \\
\text { No. }\end{array}
$$ <br>

(in vacc. \mathrm{cm}^{-1} )\end{array}\right]\)|  |
| :--- |
| 551 |

TABLE VII (Cont.)

| Serial  <br> No. Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ )Serial <br> No.. | Wave No. <br> (in vacc. $\mathrm{cm}^{-1}$ ). |  |  |
| :--- | :---: | :---: | :---: |
|  |  |  |  |
| 631 | 4078.16 | 638 | 4075.64 |
| 632 | 4077.91 | 639 | 4075.51 |
| 633 | 4077.80 | 640 | 4075.37 |
| 634 | 4077.64 | 641 | 4075.10 |
| 635 | 4077.05 | 642 | 4074.71 |
| 636 | 4076.12 | 643 | 4074.48 |
| 637 | 4076.07 |  |  |

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