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# Technical Memorandum 33-397 

## A Unified Quasilinear Theory of Weakly Turbulent Plasmas

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## N69-21092



JET PROPULSION LABORATORY CALIFORNIA INSTITUTE OF TECHNOLOGY

August 1, 1968

Technical Memorandum 33-397

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## Acknowledgments

The present research was initiated in the spring of 1087, when the author was visiting the University of Maryland, and completed after his return to the Jet Propulsion Laboratory in the fall. He enjoyed many fruitful discussions with the members of the plasma and space physies group at the Institute for Fluid Dynamics and Applied Mathematics. He is also grateful to his colleagues of the theoretical physics group at the Jet Propulsion Laboratory for their constant encouragement and stimulating conversations. The work was financially supported in part by the National Science Foundation Grant, NSF-GP-4921 at the University of Maryland, and by National Aeronautics and Space Administration Contract No. NAS 7-100 at the Jet Propulsion Laboratory, California Institute of Technology.

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#### Abstract

A quasilinear theory is formulated which includes the microscopic fluctuation fields as well as the coherent waves. The study emphasizes those cases in which the propagating mode of the fluctuation field gives rise to a dominant contribution to the particle correlation. In other words, in these cases the spontaneous Cerenkov emission of plasmons can play a more important role than the twoparticle collision process. The instability which is responsible for the turbulence can be of either electrostatic or electromagnetic, or of mixed nature. The effect of an external magnetic field is also considered in the present theory. The final sesult is valid for an arbitrary ratio of the energy density of the coherent wave to that of the fluctuation field, although the condition of weak turbulence is imposed throughout the analysis.


## A Unified Quasilinear Theory of Weakly Turbulent Plasmas

## i. Introduction

The existing theories concerning weakly turbulent plasmas may be grouped into two categories; in one, the interaction between the coherent wave and the resonant particles is emphasized, and the other is concerned with the effect of correlation. In the former, which is called the quasilinear theory in the literature (Refs. 1-7), ${ }^{1}$ correlation is completely neglected, and in the latter (Refs. 9-16), the macroscopic waves are totally ignored under the assumption of homogeneous plasma. Evidently the two theories are concerned with two extreme situations. For instance, if we define a ratio $\Gamma$ such that


[^0]the quasilinear theory deals with the case $\mathrm{r} \rightarrow 0$ and, on the other hand, the theory emphasizing correlation is concerned with the limit $\Gamma \rightarrow \infty$.

Intuitively one may think that the condition $\mathrm{T} \ll 1$ describes a more realistic plasma, and thus the quasilinear theory seems to be more interesting. However, we want to point out that the two theories are complementary. For example, in the quasilinear theory, we assume implicitly that the quasilinear stabilization process is characterized by a time scale much shorter than the characteristic time for the establishment of particle correlation. This implies that the validity of the quasilinear theory cannot bo fully appreciated unless we can first understand how the plasma turbulence affects the particle correlation. In general, three fundamental processes determine the particle correlation; their relation may be expressed as follows:

where
$(\Lambda)=$ propagnating mode of the microscopic fluctuntion field originated by both stimulated emission (or absorption) and spontancous emission
$(B)=$ nonpropagating (heavily damped) mode of the fluctuntion field, contributing to the dynamic shiolding phenomenon
$(C)=$ direct particle encounters
For turbulent plasmas in many cases, process (A) may provail over processes (B) and (C). In this memorandum we shall pay special attention to these eases. Since the effective correlation time may be grently shortened in these cases (this phenomenon is attributed to the enhanced fluctuation of the miceroscopic density and field due to the presence of instability; see Ref, 17 for publications on the effect of collective processes on various relaxation times in a stable plasma), it is desirable to include the contribution from the fluctuation field associated with the instability.

In the subsequent discussion, we shall restrict our analysis to the weakly turbulent plasmas, in which both mode-coupling and particle-trapping processes (Refs. 8 and 18) are negligible. Moreover, since most instabilities exist in magnetized plasmas, we shall include the effect of an external magnetic field. For the sake of generality, both electrostatic and electromagnetic modes will be considered so that the theory is useful for plasmas with an arbitrary value of $\beta$ (i.e., the ratio of fluid pressure to magnetic pressure). The present theory is intended to bridge the usual quasilinear theory and the kinetic theory for a homogeneous plasma. Consequently, the result is valid over the complete range of $T$.

From the result of this work, we shall see that the smallness of the ratio defined in Eq. (1) alone is not sufficient to justify the neglect of correlation. An additional criterion must be considered; that is,

$$
\left(\begin{array}{c}
\text { stimulated emission }  \tag{2}\\
\text { of plasmons with } \\
\text { momentum k }
\end{array}\right) \gg\left(\begin{array}{c}
\text { spontaneous emission } \\
\text { of plasmons with the } \\
\text { same momentum }
\end{array}\right)
$$

Evidently, in many cases, this condition is violated at a later stage of the relaxation process in the usual quasilinear theory. For instance, (1) in the one-dimensional "bump-on-the-tail" problem (Ref. 18), when the plateau is asymptotically formed and (2) in the case of ion-wave

Instability, when the spontaneous emission may become signifiennt shortly pfter an initial stage of evolution because of the high population of resonant electrons. It is true that in the first instance the contribution from the propagating mode of the fluctuntion fied to the correlation may be expected to be of the same order of magnitude as tho collisional process and, in order to be consistent, one should also include the latter effeet.

To improve the present work and to extend it to strong turbulence is indeed desirable. Perhaps the methods suggested by Dupree (Ref, 19) and Nishikawn (Ref. 20) can be unified and generalized.

## II. Mathematical Formulation

The Klimontovich equations (Refs. 21 and 22) of a multi-species plasma including both clectrostatic and electromagnetic interactions can be expressed as follows:

$$
\begin{align*}
& \frac{\partial N_{*}}{\partial l}+\mathrm{v} \cdot \nabla N_{*}+\frac{c_{n}}{m_{*}} \frac{v}{c} \times \mathrm{B}_{0} \frac{\partial N_{*}}{\partial \mathrm{v}} \\
& +\frac{\boldsymbol{e}_{*}}{m_{*}}\left(\mathbf{E}^{m}+\frac{\mathrm{v}}{c} \times \mathbf{B}^{m}\right) \cdot \frac{\partial N_{\mu}}{\partial \mathrm{v}}=0 \\
& \frac{1}{c} \frac{\partial \mathbf{E}^{m}}{\partial t}+\frac{4 \pi}{c} \sum_{n} e_{A} \int d^{3} v N_{A}(\mathbf{r}, \mathrm{v}, t) \mathbf{v}-\nabla \times \mathrm{B}^{m}=0  \tag{4}\\
& \frac{\mathrm{l}}{c} \frac{\partial \mathbf{B}^{m}}{\partial l}+\nabla \times \mathbf{E}^{m}=0  \tag{5}\\
& \nabla \cdot B^{m}=0  \tag{6}\\
& \nabla \cdot \mathbf{E}^{m}=4 \pi \sum_{A} e_{B} \int d^{3} v N_{A}(\mathbf{r}, \mathbf{v}, t) \tag{7}
\end{align*}
$$

where the subscript $s$ designates the particle species, $\mathbf{B}_{0}$ denotes a uniform external magnetic field, $\mathbf{E}^{m}$ and $\mathbf{B}^{m}$ are the microscopic electromagnetic fields, and $N_{s}(\mathbf{r}, \mathrm{v}, t)$ is the random density, which may be defined as

$$
\begin{equation*}
N_{s}(\mathrm{r}, \mathrm{v}, t)=\sum_{s} \delta\left[\mathrm{r}-\mathrm{r}_{s i}(t)\right] \delta\left[\mathrm{v}-\mathrm{v}_{s i}(t)\right] \tag{8}
\end{equation*}
$$

In expression (8), $\mathbf{r}_{k i}(t)$ and $\mathbf{v}_{8 i}(t)$ are the position and velocity vectors at a time $t$ of the $i$ th particles which belong to the s-species. The derivation of these equations
may be found in Ref. 23. Now we Introduce the Fourier spectral resofution of the mieroscopic quantilies

$$
\begin{align*}
& \mathbf{E}^{m}(\mathrm{r}, t)=\sum_{k} \mathrm{E}_{k}^{m}(t) e^{\mathrm{l} k} \cdot \mathrm{r}  \tag{9}\\
& \mathrm{~B}^{m}(\mathrm{r}, t)=\sum_{k} \mathrm{~B}_{k}^{m}(t) e^{1 \mathrm{k} \cdot \mathrm{r}}  \tag{10}\\
& N_{n}(\mathrm{r}, \mathrm{v}, t)=\sum_{k} N_{n}^{k}(v, t) e^{\mathrm{ik} \cdot \mathrm{r}} \tag{11}
\end{align*}
$$

Correspondingly, we obtain

$$
\begin{align*}
& \frac{\partial N_{A}^{0}}{\partial t}+\frac{\boldsymbol{c}_{n}}{m_{n}}\left(\frac{\mathrm{v}}{c} \times \mathrm{B}_{3}\right) \cdot \frac{\partial N_{n}^{0}}{\partial \mathrm{v}} \\
& =-\sum_{k} \frac{e_{n}}{m_{n}}\left(\mathbf{E}_{k}^{m}+\frac{\mathrm{v}}{c} \times \mathbf{B}_{k}^{m}\right) \cdot \frac{\partial N_{k}^{k}}{\partial \mathbf{v}}  \tag{12}\\
& \frac{\partial N_{A}^{k}}{\partial t}+i k \cdot v N_{n}^{k}+\frac{e_{*}}{m_{*}} \frac{v}{c} \times \mathrm{B}_{0} \frac{\partial N_{n}^{k}}{\partial \mathrm{v}} \\
& +\frac{e_{a}}{m_{A}}\left(\mathrm{E}_{k}^{m}+\frac{\mathrm{v}}{c} \times \mathrm{B}_{k}^{m}\right) \cdot \frac{\partial N_{a}^{n}}{\partial \mathrm{v}}+\frac{e_{a}}{m_{a}} \sum_{\mathbf{k}^{\prime}}\left(\mathrm{E}_{k i k}^{m},\right. \\
& \left.+\frac{v}{c} \times \mathrm{B}_{k_{-k k^{\prime}}^{m}}\right) \cdot \frac{\partial N_{n^{\prime}}^{k^{\prime}}}{\partial \mathrm{v}}=0  \tag{13}\\
& \frac{1}{c} \frac{\partial \mathbf{E}_{k}^{m}}{\partial t}-i \mathrm{k} \times \mathrm{B}_{k}^{m}+\frac{4 \pi}{c} \sum_{s} e_{s} \int d^{n} v \vee N_{n}^{k}=0  \tag{14}\\
& \frac{1}{c} \frac{\partial \mathbf{B}^{m}}{\partial t}+i \mathrm{k} \times \mathrm{E}_{k}^{m}=0  \tag{15}\\
& \mathrm{k} \cdot \mathrm{~B}_{k}^{m}=0  \tag{16}\\
& \mathrm{k} \cdot \mathrm{E}_{\vec{k}}^{m}=4 \pi \sum_{n} c_{s} \int d^{3} v N_{!}^{k}(\mathrm{v}, t) \tag{17}
\end{align*}
$$

Here we consider

$$
\mathbf{B}_{0}^{m}=\mathbf{E}_{0}^{m}=0
$$

In the following, we shall assume that the last term in Eq. (13) may be dropped, in other words, we consider that mode-coupling interaction is negligible. Moreover, we shall take the ensemble-averaged value of Eq. (12) and obtain
$\frac{\partial F_{A}}{\partial t}+\frac{e_{A}}{m_{A}}\left(\frac{\mathrm{v}}{c} \times \mathbf{B}_{0}\right) \cdot \frac{\partial F_{s}}{\partial \mathrm{v}}$
$=-\sum_{\mathbf{k}} \frac{e_{s}}{m_{s} n_{s}} \frac{\partial}{\partial \mathbf{v}} \cdot\left[\left\langle\mathbf{E}_{-k}^{m} N_{\sharp}^{k}\right\rangle+\frac{\mathrm{v}}{c} \times\left\langle\mathbf{B}_{-k}^{m} N_{s}^{k}\right\rangle\right]$
where the symbol () denotes the ensemble-avernged value and the relation

$$
\begin{equation*}
\left\langle N_{i}^{0}(v, t)\right\rangle n_{n} F_{n}(v, t) \tag{10}
\end{equation*}
$$

has been used, In Equ, (19), $F_{n}$ is the one-particle distribution function of the s-spucies particles, which is independent of spatinal coordinates; $n_{n}$ is the average number density of the same species.

Equation (18), in principle, represents the desired kinetic equation, although the right-hand terms are to bo determined. The correlation functions $\left\langle\mathrm{E}_{\vec{k}}^{m} \mathrm{~N}_{\boldsymbol{\prime}}\right\rangle$ ) and $\left\langle B_{r ;}^{m} N_{n}^{k}\right\rangle$ can be derived from Eqs. (13) through (17). However, before going further, some useful definitions and relations should be introduced. First we shall designate

$$
\begin{equation*}
\left\langle N_{n}^{*}(\mathrm{v}, t)\right\rangle=n_{r} f_{n}(\mathrm{k}, \mathrm{v}, t) \tag{20}
\end{equation*}
$$

Evidently, if $\xi_{A}(r, v, t)$ is the distribution function of the $s$-species, $\mathcal{J}_{a}$ may be expressed as

$$
\begin{equation*}
\delta_{s}(\mathrm{r}, \mathrm{v}, t)=\frac{\left\langle N_{*}(\mathrm{r}, \mathrm{v}, t)\right\rangle}{n_{*}}=F_{*}(\mathrm{v}, t)+\sum_{\mathrm{k}} f_{*}(\mathrm{k}, \mathrm{v}, t) e^{i \mathrm{k} \cdot \mathrm{r}} \tag{21}
\end{equation*}
$$

Moreover, we see that

$$
\begin{align*}
& \left\langle N_{*}(\mathbf{r}, \mathrm{v}, t) N_{\mathrm{r}}(\mathbf{r}, \mathrm{v}, t)\right\rangle \\
& =\sum_{k} \sum_{k^{\prime}}\left\langle N_{s}^{k}(v, t) N_{n}^{k^{\prime}-k}\left(v^{\prime}, t\right)\right\rangle \exp \left[i k \cdot r+i\left(k^{\prime}-k\right) \cdot r^{\prime}\right] \\
& =\left\langle N_{d}^{0}(v, t) N_{r}^{0}\left(v^{\prime}, t\right)\right\rangle \\
& +\sum_{k \neq 0}\left\langle N_{s}^{k}(\mathbf{v}, t) N_{r}^{-k}\left(\mathbf{v}^{\prime}, t\right)\right\rangle \exp \left[i \mathbf{k} \cdot\left(\mathbf{r}-\mathbf{r}^{\prime}\right)\right] \\
& +\sum_{\mathrm{k} \neq 0} \sum_{\mathrm{k}^{\prime}}\left\langle N_{B}^{l k}(\mathrm{v}, t) N_{r}^{k_{r}^{\prime}-k}\left(\mathrm{v}^{\prime}, t\right)\right\rangle \exp \left[i \mathrm{k} \cdot \mathrm{x}+i\left(\mathrm{k}^{\prime}-k\right) \cdot \mathrm{r}^{\prime}\right] \\
& =n_{s} n_{r} F_{H}(\mathrm{v}, t) F_{r}\left(\mathrm{v}^{\prime}, t\right)+n_{H} n_{r} G_{H}\left(0, \mathrm{v}, \mathrm{v}^{\prime}, t\right) \\
& +n_{H} \delta_{\beta r}\left(v-v^{\prime}\right) F_{B}(v, t) \\
& +\sum_{\mathrm{k} \neq 0}\left\{n_{H} n_{r} G_{s r}\left(\mathrm{k}, \mathrm{v}, \mathrm{v}^{\prime}, t\right)+n_{H} \delta\left(\mathrm{v}-\mathrm{v}^{\prime}\right) \delta_{s r} F_{s}(\mathrm{v}, t)\right. \\
& \left.+n_{s} n_{r} f_{s}(\mathbf{k}, \mathrm{v}, t) f_{r}\left(-\mathrm{k}, \mathrm{v}^{\prime}, t\right)\right\} \exp \left[i \mathbf{k} \cdot\left(\mathbf{r}-\mathbf{r}^{\prime}\right)\right] \\
& +\sum_{k \neq 0} \sum_{k^{\prime}}\left\{n_{s} n_{r} g_{g r}\left(k, k^{\prime}-\mathbf{k}, \mathrm{v}, \mathrm{v}^{\prime}, t\right)\right. \\
& +n_{d} \delta\left(\mathrm{v}-\mathrm{v}^{\prime}\right) f_{s}\left(\mathrm{k}^{\prime}-\mathrm{k}, \mathrm{v}, t\right) \\
& \left.+n_{s} n_{r} f_{s}(\mathbf{k}, \mathrm{v}, t) f_{r}\left(\mathrm{k}^{\prime}-\mathbf{k}, \mathrm{v}^{\prime}, t\right)\right\} \\
& \times \exp \left[i k \cdot r+i\left(k^{\prime}-k\right) \cdot r^{\prime}\right] \tag{22}
\end{align*}
$$

where $G_{a r}$ and gar represent the Fourter components of the pair-correlation function $P_{\text {ar }}$ (see Ref. 24 for dellnition), or wo may write

$$
\begin{align*}
P_{a r}\left(r, r^{\prime}, v, v^{\prime}, l\right)= & \sum_{k} G_{a r}\left(k, v, v^{\prime}, l\right) e^{i k \cdot r} \\
& +\sum_{k \neq 0} \sum_{k} \mu_{a r}\left(k, k^{\prime}-k, v, v^{\prime}, l\right) \\
& \times \exp \left[i k \cdot r+i\left(k^{\prime}-k\right) \cdot r^{\prime}\right] \tag{23}
\end{align*}
$$

From Eq. (22), we obtain by comparison that

$$
\begin{align*}
\left\langle N_{s}^{k}(v, t) N_{r}^{\prime}\left(v^{\prime}, t\right)\right\rangle= & n_{n} n_{r} f_{n}(k, v, t) f_{r}\left(-k, v^{\prime}, t\right) \\
& +n_{n} n_{r} G_{n r}\left(k, v, v^{\prime}, t\right) \\
& +n_{A} \delta_{n r} \delta\left(v-v^{\prime}\right) F_{n}(v, t) \tag{24}
\end{align*}
$$

This relation will be useful in discussions presented in later sections.

## III. Initial-Vạlue Problem

Now we intend to solve an initial-value problem necording to the following equations:

$$
\begin{gather*}
\frac{\partial N_{s}^{k}}{\partial t}+i \mathrm{k} \cdot \mathrm{y} N_{n}^{k}+\frac{e_{A}}{m_{n}} \frac{\mathrm{v}}{c} \times \mathbf{B}_{0} \cdot \frac{\partial N_{n}^{k}}{\partial \mathrm{v}} \\
+\frac{e_{n} n_{n}}{m_{n}}\left(\mathrm{E}_{k}+\frac{\mathrm{v}}{c} \times \mathrm{B}_{k}\right) \cdot \frac{\partial F_{n}}{\partial \mathrm{v}}=0  \tag{25}\\
\frac{1}{c} \frac{\partial \mathrm{E}_{k}}{\partial t}+\frac{4 \pi}{c} \sum_{n} e_{A} \int d^{3} \cup v N_{n}^{k}(\mathrm{v}, t)-i \mathrm{k} \times \mathrm{B}_{k}=0  \tag{26}\\
\frac{1}{c} \frac{\partial \mathbf{B}_{k}}{\partial t}-i \mathrm{k} \times \mathrm{E}_{k}=0 \tag{27}
\end{gather*}
$$

Two points should be noted: (1) for simplicity, we have dropped the superscript $m$ in the fields $\mathbf{E}^{m}$ and $\mathbf{B}^{m}$ and should keep in mind that these fields are microscopic ones, and (2) in Eq. (25) we have replaced $N_{j}^{0}$ by $n_{s} F_{a}$ (or, in other words, neglected $\delta N_{R}^{0}$ ). Such an approximation does not affect the first-order theory (in the sense discussed by Dupree in Ref, 25) since the error in the evaluation of the function $\left\langle E_{-k} N_{k}^{k}\right\rangle$ is of second order.

Solving Eq. (25) by charneteristic integration nlong the particlo orbit, wo find a formal solution of $N_{i}^{\prime \prime}$; that is,

$$
\begin{align*}
N_{A}^{k}(v, t) & =\exp \left[-i k \cdot r_{n}(t)\right] N_{A}^{\prime}\left(v_{A}(t), 0\right) \\
& =\frac{n_{a} c_{a}}{m_{A}} \int_{0}^{t} d \tau \exp \left[-i k \cdot r_{A}(\tau)\right]\left\{\mathrm{E}_{k}(l-\tau)\right. \\
& \left.+\frac{v_{n}(\tau)}{c} \times \mathrm{I}_{k}(t-\tau)\right\} \cdot \frac{\partial F_{n}\left(v_{a}(r)\right)}{\partial v_{n}(\tau)} \tag{28}
\end{align*}
$$

where $r_{n}(t)$ and $v_{n}(t)$ satisfy the following equations of motion:

$$
\begin{gather*}
\frac{d r_{n}(t)}{d t}=\mathbf{v}_{n}(t)  \tag{20}\\
\frac{d \mathbf{v}_{n}(l)}{d t}=\frac{e_{n}}{m_{n}}\left[\mathbf{E}_{k}(t)+\frac{\mathbf{v}_{n}(t)}{c} \times \mathbf{B}_{k}(t)\right] \tag{30}
\end{gather*}
$$

together with the initial conditions

$$
\begin{equation*}
\mathbf{r}_{n}(0)=0 \tag{31}
\end{equation*}
$$

and

$$
\begin{equation*}
v_{n}(0)=v \tag{32}
\end{equation*}
$$

We notice that Eqs. (25), (26), and (27) resemble the linenrized Vlasov equations." This explains why the instability associated with the macroscopic wave may also happen to the microscopic wave. In the following we postulate that ench Fourier component $\mathbf{E}_{k}$ and $\mathbf{B}_{k}$ may be described by two distinct time variables, say $t$ and $e t$, where $e$ is a small dimensionless parameter; one describes the fast oscillations and the other records the slow amplification. Although the fundamental notion of the multiple time variables is similar to that of Krylov and Bogoliubov (Ref. 26), the only purpose of doing this here is to simplify the Laplace transform method to be used later.3

To proceed with this discussion, we shall write

$$
\begin{align*}
& \mathbf{E}_{k}=\mathbf{E}_{k}(t, a t)  \tag{33}\\
& \mathbf{B}_{k}=\mathbf{B}_{k}(t, a t) \tag{34}
\end{align*}
$$

[^1]Henco we shall introduce a Laplace transform with respect to the fast time variable $t$; l.e.,

$$
\begin{align*}
& \left.N_{i}^{k}(v, t), t\right)=\lim _{\Delta \rightarrow 0} \int_{0}^{\infty} d t \exp (l, t-\Delta t) N_{k}^{k}(v, t, e t)  \tag{35}\\
& \mathrm{E}_{k}(t, r, t) r \operatorname{llm}_{\Delta \rightarrow 0} \int_{i=1}^{\infty} d t \exp (t, t \infty \Delta t) \mathbf{E}_{k}(t, e t)  \tag{30}\\
& B_{k}(\omega t, a t)=\lim _{\Delta \rightarrow 0} \int_{0}^{\infty} d t \exp \left(f(t, t-\Delta t) B_{k}(t, a t)\right. \tag{37}
\end{align*}
$$

where ois real. From (28) and (35), we obtain

$$
\begin{align*}
& N_{\Delta}(v, \omega, e t)=\lim _{\Delta \rightarrow 0} \int_{0}^{\infty} d t \exp \left[i \omega t-\Delta t-i \mathrm{k} \cdot \mathrm{r}_{1}(t)\right] \\
& \times\left\{N_{s}^{k}\left(v_{s}(t), 0\right)-\frac{n_{s} c_{n} l}{m_{n}(\omega)}\left(\frac{v_{a}(t)}{c} \times B_{k}(l=0)\right)\right. \\
& \text { - } \left.\frac{\partial F_{A}\left(v_{0}(t)\right)}{\partial \mathrm{V}_{A}(t)}\right\} \cdots \lim _{\Delta \rightarrow 0} \frac{m_{1} e_{n}}{m_{\Delta}} \int_{0}^{\infty} d t \exp \left[\left\{\omega \tau-\Delta r-i \mathbf{k} \cdot r_{A}(\tau)\right]\right. \\
& \times \int_{0}^{\infty} d r^{\prime} \exp \left(l \omega r^{\prime}-\Delta r^{\prime}\right)\left[\mathbf{E}_{k}\left(r^{\prime}, \varepsilon r^{\prime}\right)-\frac{v_{0}(l)}{\omega}\right. \\
& \left.\times\left(E_{k}\left(T^{\prime}, E T^{\prime}\right) \times k\right)\right] \cdot \frac{\partial F_{s}\left(v_{s}(t)\right)}{\partial v_{s}(t)} \tag{38}
\end{align*}
$$

where we have made use of the relation

$$
\begin{equation*}
\int_{0}^{\infty} d \tau e^{i \omega \tau} B_{k}=\frac{c}{\omega} \int_{0}^{\infty} d r e^{i \omega \tau}\left(k \times \mathrm{E}_{k}\right)+\frac{i}{\omega} \mathrm{~B}_{k}(l=0) \tag{39}
\end{equation*}
$$

which is obtainable from Eq. (27).
If a denotes a unit polarization vector which is parallel to the field $\mathbb{E}_{k}$, the last term in Eq. (35) can be written as

$$
\begin{aligned}
& {\left[N_{n}^{k}(v, w, e t)\right]_{\text {lant }}} \\
& =-\lim _{\Delta \rightarrow 0} \frac{n_{n} e_{A}}{m_{A}} \int_{n}^{\infty} d \tau \exp \left[i \omega r-\Delta r-i k \cdot \mathrm{r}_{A}(\tau)\right] \\
& \times\left[\mathfrak{a}+\frac{\left(\mathbf{a} \cdot \mathbf{v}_{\mathbf{a}}(\tau)\right) \mathbf{k}-\left(\mathbf{k} \cdot \mathbf{v}_{\mathrm{a}}(\tau)\right) \mathbf{a}}{\omega}\right] \\
& \text { - } \frac{\partial F_{s}\left(v_{s}(t)\right)}{\partial v_{s}(\tau)} \int_{0}^{\infty} d \tau^{\prime} \exp \left(i \omega \tau^{\prime}-\Delta \tau^{\prime}\right) E_{k}\left(\tau^{\prime}, e t\right)
\end{aligned}
$$

$$
\begin{align*}
& \square G(\omega) \int_{a}^{\alpha} d r^{\prime} \exp \left(b, r^{\prime}-\Delta T^{\prime}\right) E_{k}\left(r^{\prime}, e r^{\prime}\right) \\
& \simeq \lim _{\Delta \rightarrow 0} G(0) \int_{0}^{\infty} d r^{\prime} \exp \left(i \omega r^{\prime}-\Delta r^{\prime}\right) \\
& \times\left\{E_{k}\left(r^{\prime}, r t\right)+n\left(r^{\prime}-t\right) \cdot \frac{\partial E_{k}\left(r^{\prime}, t\right)}{\partial t}+O\left(r^{2}\right)\right\} \\
& \simeq G(\omega) E_{k}(\omega, e t)+i \frac{\partial G}{\partial(t)} \frac{\partial E_{k}(\omega, e t)}{\partial t}+O\left(\varepsilon^{2}\right) \\
& =-\lim _{\Delta \rightarrow 0} \frac{n_{4} C_{A}}{\frac{n_{s}}{m_{s}}} \int_{0}^{\infty} d{ }^{2}\left\{\left[E_{k}(\omega, t t)+\frac{i \partial E_{k}(\omega, t t)}{\partial t} \frac{\partial}{\partial \omega}\right]\right. \\
& \times\left[\operatorname{Opp}\left[\omega_{\omega r}-\Delta r-i k \cdot r_{1}(r)\right]\right. \\
& \times\left(n+\frac{\left(u \cdot v_{A}(\tau)\right) k-\left(k \cdot v_{a}(\tau)\right) a}{\omega}\right) \\
& \text { - } \left.\left.\frac{\partial F_{a}\left(v_{n}(\tau)\right)}{\partial v_{A}(\tau)}\right]\right\}+O\left(\varepsilon^{a}\right) \tag{40}
\end{align*}
$$

Hereafter we shall ituglect the second-order terms. Returning to Eqs. (26) and (27), and making use of Eqs. (38) and (40), we find

$$
\begin{align*}
& {\left[\omega^{2} \epsilon_{i j}(k, \omega)-c^{2} k^{2} \delta_{1 /}+c^{2} k_{i} k_{i}\right] E \|_{k}(\omega, a t)} \\
& +i \omega \frac{\partial}{\partial w}\left[\frac{c^{2} k_{i} k_{i}-c^{2} k^{2} \delta_{i j}}{\omega}+\omega \epsilon_{j}(k, \omega)\right] \frac{\partial \mathbb{E}_{t}^{\prime}(\omega, t t)}{\partial t} \\
& =H_{i}(\mathrm{k}, \omega)+K_{i}(\mathrm{k}, \omega) \tag{41.}
\end{align*}
$$

where

$$
\begin{align*}
& \mathrm{H}^{+}=-\lim _{\Delta \rightarrow 0} d_{\sin } \sum_{n} \sum_{n}^{m} e_{n} \int d^{d} v \mathrm{v} \\
& \times \int_{0}^{\infty} d t \exp \left[i \omega t-\Delta t-i k \cdot r_{n}(t)\right] N_{\theta}^{k}\left(v_{s}(t), 0\right) \tag{42}
\end{align*}
$$

$$
\begin{align*}
\mathrm{K}^{+}= & -i c\left[\mathrm{k} \times \mathbf{B}_{k}(t=0)\right]+i_{\omega} \mathrm{E}_{k}(t=0) \\
& -4 \pi \lim _{\Delta \rightarrow 0} \sum_{a} e_{s} \int d^{\mathrm{a}} v \mathrm{v} \\
& \times \int_{0}^{\infty} d t \exp \left[i_{\Delta} t-\Delta t-i \mathbf{k} \cdot \mathrm{r}_{s}(t)\right] \\
& \times \frac{n_{a} e_{s}}{m_{a}}\left(\frac{\mathrm{v}_{a}(t) \times \mathbf{B}_{k}(t=0)}{c}\right) \cdot \frac{\partial F_{n}\left(\mathrm{v}_{s}(t)\right)}{\partial \mathbf{v}_{A}(t)} \tag{43}
\end{align*}
$$

$$
\begin{align*}
& \otimes_{i j}(k, \omega)=\delta_{i j}+\frac{4 \pi i}{\omega} \lim _{\Delta \rightarrow t} \sum_{A} \frac{n_{A} c_{a}^{a}}{m_{A}} \int d^{3} v v_{1} \\
& \times \int_{0}^{\infty} d t \exp \left\{k_{n y} t-\Delta t-i k \cdot r_{A}(t)\right] \\
& \times\left\{\left[1-\frac{k \cdot v_{A}(t)}{\omega}\right] \frac{\partial F_{s}\left(v_{A}(t)\right)}{\partial v_{n j}(t)}+\frac{v_{A j}(t)}{\omega}\left(k \cdot \frac{\partial F_{A}\left(v_{A}(t)\right)}{\partial v_{B}(t)}\right)\right\} \tag{44}
\end{align*}
$$

In $\epsilon_{1 j}(k,(1)$, which is the usual dielectric tensor, the subscript + indicates that the real frequency w should be constdered as the limit of a complex frequency with a vanishing positive imaginary part. Let us write

$$
\begin{equation*}
\epsilon_{i j}^{+}(\mathrm{k}, \omega)=\operatorname{Re} \epsilon_{j j}^{+}(\mathrm{k}, \omega)+i \operatorname{Im} \epsilon_{\dagger_{j}}(\mathrm{k}, \omega) \tag{45}
\end{equation*}
$$

Hereafter we shall consider that $I m \epsilon_{i j}^{+}$is of $O(c)$ in the transparent region. Neglecting the high-order term in Eq. (41), we find, by collecting the real and imaginary parts, that

$$
\begin{align*}
& a_{1}\left[\omega^{2} R e \epsilon_{i j}^{+}-c^{2} k^{2} \delta_{i j}+c^{2} k_{i} k_{j}\right] a_{j} E_{k}^{n}= \\
& \operatorname{Rc}\left\{\left[H_{j}^{\dagger}(\mathrm{k}, \omega)+K_{i}^{+}(\mathrm{k}, \omega)\right] E_{k}^{*} a_{i}\right\}  \tag{46}\\
& 2 \omega \omega^{2} a_{i} I m \epsilon_{i j}^{+} a_{j} E_{i}^{\top}+\omega \frac{\partial}{\partial \omega}\left[a _ { i } \left(\frac{-c^{2} k^{2} \delta_{i j}+c^{3} k_{i} k_{j}}{\omega}\right.\right. \\
& \left.\left.\quad+\omega R e \epsilon_{i j}^{\dagger}\right) a_{j}\right] \frac{\partial E_{k}^{2}}{\partial t} \\
& = \tag{47}
\end{align*}
$$

where $E_{k}=E_{k}(\omega, t) E_{k}^{*}(\omega, t)$, and $E_{k}^{*}$ denotes the complex conjugate of $E_{k}$ Obviously, from Eq. (46), we obtain the dispersion equation

$$
\begin{equation*}
\left|u^{2} R c \epsilon_{i j}^{\dagger}(\mathrm{k}, \omega)-c^{3} k^{3} \delta_{1 j}+c^{2} k_{i} k_{j}\right|=0 \tag{48}
\end{equation*}
$$

Thus, in the transparent region, Eq. (17) may be rewritten as

$$
\begin{gather*}
\frac{\partial}{\partial \omega}\left(\omega_{1} a_{i} R e \epsilon_{i j} a_{j}\right) \frac{\partial E_{k}^{2}}{\partial t} \\
=-2(\omega)^{2} a_{i} I m \epsilon_{i j}^{+} a_{j} E_{k}^{2}+2 \operatorname{lm}\left[\left(H_{i}^{+}+K_{i}^{+}\right) a_{l} E_{k}^{*}\right] \tag{49}
\end{gather*}
$$

Equation (49) is useful in the subsequent discussion.

## IV. The Equation of Wave Amplitude

If we consistently keep all the terms in Eq. (49) to the lowest order in $\varepsilon$, then we heed to insert only the zerothorder expression for $E_{k}^{*}$ into the right-hand side. According to Eq. (41), we lave

$$
\begin{equation*}
E_{k}^{*}(\omega, t)=\frac{\left[H_{i}(k, \omega)+K_{i}(k, \omega)\right] a_{l}}{a_{\mu} \alpha_{\mu \nu}^{-} a_{\nu}} \tag{50}
\end{equation*}
$$

where $\mathrm{H}^{-}$and $\mathrm{K}^{-}$represent the complex conjugate of $\mathrm{H}^{+}$ and $K^{+}$, respectively. Moreover, for simplicity, we have defined

$$
\begin{equation*}
\alpha_{\mu \nu}^{ \pm} \equiv \omega^{2} \epsilon_{\mu \nu}^{ \pm}-c^{2} k^{2} \delta_{\mu \nu}+c^{3} k_{\mu} k_{\nu} \tag{51}
\end{equation*}
$$

Hence, after taking the ensomble-averaged value, we can write Eq. (47) as

$$
\begin{equation*}
\frac{\partial\left\langle E_{k}^{n}(\omega, e t)\right\rangle}{\partial t}=2 \gamma\left\langle E_{k}^{\prime}(\omega, e t)\right\rangle+2 I m\left\{\frac{\left\langle\left[H_{l}^{+}+K_{l}^{+}\right] a_{l}\left[H_{m}^{-}+K_{m}^{-}\right] a_{m}\right\rangle}{\left(a_{\mu} \alpha_{\mu \nu}^{-} a_{\nu}\right)\left[\omega \frac{\partial}{\partial \omega} a_{i}\left(\frac{c^{2} k_{i} k_{j}-c^{2} k^{2} \delta_{i j}}{\omega}+\omega R e \epsilon_{i j}^{+}(k, \omega)\right) a_{j}\right]}\right\} \tag{52}
\end{equation*}
$$

where

$$
\begin{equation*}
\gamma=-\frac{\omega^{2} a_{i} I m \epsilon_{i j}^{+}(\mathbf{k}, \omega) a_{j}}{\omega \frac{\partial}{\partial \omega} a_{\mu}\left(\frac{c^{2} k_{\mu} k_{\nu}-c^{2} k^{2} \delta_{\mu \nu}}{\omega}+\omega \operatorname{Re} \epsilon_{\mu \nu}^{+}\right) a_{\nu}} \tag{53}
\end{equation*}
$$

Let us restrict our discussion to the field associated with the unstable mode. For that case, we may write

$$
E_{k}(t, e t)=E_{0 k}(e t) \exp \left(-i_{\omega_{q}} t\right)
$$

Thus,

$$
\begin{align*}
\left\langle E_{k}^{\prime}(\omega, \varepsilon t)\right\rangle= & \lim _{\Delta \rightarrow 0} \int_{0}^{\infty} d \tau \exp \left[i\left(\omega-\omega_{q}\right) \tau-\Delta \tau\right] \int_{0}^{\infty} d \tau^{\prime} \\
& \times \exp \left[-i\left(\omega-\omega_{q}\right) \tau^{\prime}-\Delta \tau^{\prime}\right] \\
& \times\left\langle E_{0 k}^{*}(\varepsilon \tau) E_{0 k}\left(\varepsilon \tau^{\prime}\right)\right\rangle \\
= & \lim _{\Delta \rightarrow 0} \int_{0}^{\infty} \int_{0}^{\infty} d \tau d \tau^{\prime} \exp \left[i\left(\omega-\omega_{q}\right)\left\langle\tau-\tau^{\prime}\right)-\Delta\left(r+\tau^{\prime}\right)\right] \\
& \times\left\{\left\langle E_{0 k}^{*}(\varepsilon t) E_{0 k}\left(\varepsilon t^{\prime}\right)\right\rangle+O(\varepsilon)\right\} \\
= & \lim _{\Delta \rightarrow 0} \frac{\pi}{\Delta} \delta\left(\omega-\omega_{q}\right)\left\langle E_{k}^{R}(\varepsilon t)\right\rangle+O\left(\varepsilon^{\prime}\right) \tag{54}
\end{align*}
$$

When we substitute expression (54) into Eq. (52), we may ignore the first-order terms in (54) since both $\partial\left\langle E_{k}^{\prime}\right\rangle / \partial t$ and $2 \gamma\left\langle E_{k}^{2}\right\rangle$ are already spontancously first order in $e$. Furthermore, we know that when $\omega=\omega_{q}(k)$, we may write

$$
\begin{align*}
\frac{1}{a_{\mu} \alpha_{\mu \nu}^{-}\left(k, \omega_{q}\right) a_{\nu}} & =\frac{i \pi \delta\left(\omega-\omega_{q}\right)}{\frac{\partial}{\partial \omega_{q}}\left[a_{\mu} R e \alpha_{\mu \nu}^{-}\left(k, \omega_{q}\right) a_{\nu}\right]} \\
& =\frac{i \pi \delta\left(\omega-\omega_{q}\right)}{\frac{\partial}{\partial \omega_{q}}\left[\omega_{\eta}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{v}\right]} \tag{55}
\end{align*}
$$

and

$$
\begin{gather*}
\frac{1}{\omega_{q} \frac{\partial}{\partial \omega_{q}}\left[a_{i}\left(\frac{c^{2} k_{i} k_{j}-c^{2} k^{2} \delta_{i j}}{\omega_{q}}+\omega_{q} R e \epsilon_{i j}^{+}\right) a_{j}\right]} \\
=\frac{1}{\frac{\partial}{\partial \omega_{q}}\left[\omega_{\eta}^{2} a_{i} R e \epsilon_{i j}^{+}\left(k, \omega_{q}\right) a_{j}\right]} \tag{56}
\end{gather*}
$$

From Eq. (5i) and relations (54), (55), and (56), we find, after integrating over $\omega$ throughout Eq. (52), that

$$
\begin{align*}
\frac{\partial\left\langle E_{k}^{\prime}(c t)\right\rangle}{\partial t}= & 2 \gamma\left(\mathbf{k}, \omega_{q}\right\rangle\left\langle E_{k}^{u}(c t)\right\rangle \\
& +\lim _{\Delta \rightarrow 0} \frac{2 \Delta\left\langle H_{l}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{l} H_{m}^{-}\left(\mathbf{k}, \omega_{q}\right) a_{m}\right\rangle}{\left|\frac{\partial}{\partial \omega_{q}}\left[\omega_{\eta}^{2} a_{\mu} R c \epsilon_{\mu \nu}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{\nu}\right]\right|^{\prime 2}} \tag{57}
\end{align*}
$$

Notice that in the source term of Eq. (57) all terms proportional to $K^{ \pm}$vanish. The reason $\Delta\left\langle H_{\downarrow}^{+} a_{l} H_{m}^{-} a_{m}\right\rangle$ survives is that it contains a part which behaves like $\Delta^{-1}$ (see Appendix A). The final form of Eq. (57) is

$$
\begin{align*}
& \frac{\partial}{\partial t}\left\langle E_{k}^{\prime}\langle e t)\right\rangle=2 \gamma\left(\mathrm{k}, \omega_{q}\right)\left\langle E_{k}^{\prime}\left({ }_{k} t\right\rangle\right\rangle \\
& +64 \pi^{4} \omega_{n}^{2} \sum_{s} \sum_{n=-\infty}^{+\infty} n_{s} e_{n}^{2} \int_{0}^{\infty} d v_{\lrcorner} v_{⿺} \int_{-\infty}^{+\infty} d v_{z} \\
& \times \frac{\left.\delta\left(\omega_{q}-n \Omega_{s}-k_{2} v_{2}\right) \mid Q_{n}^{+}\right]^{2} F_{s}\left(v_{z}, v_{\perp}^{2}\right)}{\left|\frac{\partial}{\partial \omega_{q}}\left[\omega_{\tilde{I}}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{\nu}\right]\right|^{2}} \tag{58}
\end{align*}
$$

where

$$
\begin{align*}
\left|Q_{n}^{\dagger}\right|^{2}= & {\left[v_{s} a_{z}+\frac{n \Omega_{n} a_{\perp}}{k_{\perp}} \cos \psi\right]^{n} J_{n}^{*}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{n}}\right) } \\
& +a_{L}^{n} v_{L}^{\prime} \sin ^{2} \psi J_{n}^{\prime 2}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{n}}\right) \tag{59}
\end{align*}
$$

In Eq. (58) and expression (59), the subscripts $z$ and $\perp$ denote the components of vectors parallel and perpendicular to the magnetic field, $\psi$ is the angle between the vectors $\Omega_{\perp}$ and $\mathrm{k}_{\perp}, \Omega_{s}=e_{s} B_{0} / m_{s} c$, and $J_{n}$ and $J_{n}^{\prime}$ are the Bessel function of $n$th order and its derivative, respectively. The first term on the right-hand side may be conceived to be proportional to the stimulated emission of plasmons associated with momentum k and the second term proportional to the spontaneous emission.

In deriving Eq. (58), we have assumed that

$$
\begin{equation*}
F_{s}(v)=F_{n}\left(v_{z}, v_{\perp}^{a}\right) \tag{60}
\end{equation*}
$$

The validity of this postulated condition may be questioned if the unstable mode is propagated in a direstion neither parallel nor perpendicular to the external magnetic field $\mathbf{B}_{0}$. One important point is that if the plasma is initially cylindrically symmetrical, there should be no preferred direction of propagation in the plane of symmetry. This is to say that if the aforementioned unstable mode exists, then other modes must exist which have the same propagation characteristics but are cylindrically symmetrical with respect to the magnetic field. Consequently the symmetry condition is not violated.

Finally, since $E_{k}$ is the microscopic field, the quantity $\left\langle E_{n}^{*}\right\rangle$ may be expressed as the sum of two parts, i.e.,

$$
\left\langle E_{k}^{\prime}\right\rangle=\tilde{E}_{k}^{\prime}+\left\langle\delta E_{\hat{k}}\right\rangle
$$

where $\widetilde{E}_{k}$ is the ensemble-averaged field or the macroscopic field and $\delta E_{k}$ the fluctuation field. In the usual quasilinear theory, it is assumed that

$$
\tilde{E}_{k}^{n} \gg\left\langle\delta E_{k}^{n}\right\rangle
$$

and, on the other hand, the theory with correlation considers

$$
\left\langle\delta E_{k}\right\rangle \gg \tilde{E}_{k}^{\prime} \rightarrow 0
$$

## V. The Kinetic Equation

Formally the kinetic equation may be expressed as

$$
\begin{align*}
\frac{\partial F_{s}}{\partial t}= & -\sum_{\mathbf{k}} \frac{e_{s}}{m_{s} n_{s}} \frac{\partial}{\partial v} \int_{-\infty}^{+\infty} d \omega \lim _{\Delta \rightarrow 0} \frac{\Delta}{\pi} \\
& \times\left[\mathfrak{a}+\frac{(v \cdot a) k-(v \cdot k) \mathfrak{a}}{\omega}\right]\left\langle E_{k}^{*}(\omega, e t) N_{s}^{k}(v, \omega, e t)\right\rangle \tag{61}
\end{align*}
$$

Since $F_{\Delta}$ is considered to be symmetrical in the $v_{1}$-plane and $\left\langle E_{-k}^{*} N_{d}^{-k}\right\rangle=\left\langle E_{k}^{*} N_{s}^{*}\right\rangle^{*}$, we can average the right-hand term of Eq. (61) over the azimuthal angle and retain only the real part of it. Thus,

$$
\begin{align*}
& \frac{\partial F_{s}}{\partial t}=-\sum_{k .} \frac{e_{s}}{m_{s} n_{s}} \lim _{\Delta \rightarrow 0} \frac{\Delta}{2 \pi^{2}} \int_{0}^{2 \pi} d \phi \int_{-\infty}^{+\infty} d \omega \frac{\partial}{\partial \mathrm{v}} \\
& \quad \cdot \operatorname{Re}\left\{\left[\mathrm{a}+\frac{(\mathrm{v} \cdot \mathrm{a}) \mathrm{k}-(\mathrm{v} \cdot \mathrm{k}) \mathrm{a}}{\omega}\right]\left\langle\mathrm{E}_{k}^{*}(\omega, e t) N_{\Delta}^{k}(\mathrm{v}, \omega, t)\right\rangle\right\} \tag{62}
\end{align*}
$$

The evaluation of the right-hand term is shown in Appendix $B$. Here we shall simply state the result:

$$
\begin{align*}
\frac{\partial F_{s}}{\partial t}= & \sum_{k} \frac{\pi e_{s}^{z}}{m_{s}} \sum_{n=-\infty}^{+\infty}\left(k_{z} \frac{\partial}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right) \\
& \times\left\{\left|Q_{n}^{+}\right|^{2} \delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right)\right. \\
& \times\left[\frac{4 \pi F_{s}\left(v_{z}, v_{\perp}^{2}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{v}\right)}\right. \\
& \left.\left.+\frac{\langle E k}{\left.\omega_{q}^{2}(z t)\right\rangle}\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right]\right\} \tag{63}
\end{align*}
$$

The quantity $\left|Q_{n}^{+}\right|^{=}$is defined by Eq. (59).
Equation (63) is obtained in the fuasilinear approximation in which the resonant diffusion process is considered to be most important and the nonresonant contribution and the ordinary collision integral are neglected. The expression for the nonresonant contribution is somewhat lengthy and may be found in Appendix B. The term proportional to $\left\langle E_{k}^{?}\right\rangle$ is obtainable from the conventional quasilinear theory. The extra term, or what we call the friction term, is attributed to correlation.

## VI. Conservation of Energy

It is important to examine the law of conservation of energy from the equation just derived. However, since the resonant diffusion process is important only in a restricted region in the velocity space, it is conceivable that in the discussion of conservation of energy and momentum one must consider both resonant and nonresonant contributions to the kinetic equation.

Let us study the resonant contribution by making use of Eq. (63). We obtain

$$
\begin{align*}
\left(\frac{\partial U}{\partial t}\right)_{\text {resonant }}= & \sum_{s} \frac{\partial}{\partial t}\left[\frac{n_{s} m_{s}}{2} \int l^{3} v v^{2} F_{s}\right]_{\text {resonnat }} \\
= & -\sum_{k} \sum_{s} \frac{2 \pi^{2} n_{s} e_{s}^{\frac{2}{s}}}{m_{s}} \int_{0}^{\infty} d p_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \\
& \times \sum_{n^{2}-\infty}^{+\infty}\left(k_{z} v_{z}+n \Omega_{s}\right)\left|Q_{n}^{+}\right|^{2} \delta\left(\omega_{q}-n \Omega_{s}\right. \\
& \left.-k_{z} v_{z}\right)\left[\frac{4 \pi m_{s} F_{s}}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{\nu}\right)}\right. \\
& \left.+\frac{\left(E_{\tilde{s}}^{n}(\varepsilon t)\right\rangle}{\omega_{q}^{2}}\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right] \tag{64}
\end{align*}
$$

However, from Eq. (58) and definition (53), we find

$$
\begin{align*}
& \frac{\partial}{\partial \omega_{q}}\left\langle\omega_{\eta}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{\nu}\right) \frac{\partial}{\partial t} \frac{\left\langle E_{k}^{2}\right\rangle}{8 \pi}=\sum_{s} \frac{2 \pi^{2} n_{s} e_{s}^{2}}{m_{s}} \\
& \times \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \sum_{n=-\infty}^{+\infty}\left|Q_{n}^{+}\right|^{2} \delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right) \\
& \times\left[\frac{4 \pi m_{s} F_{8} \omega_{q}^{2}}{\frac{\partial}{\partial \omega_{q}}\left\langle\omega_{\square}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{\nu}\right)}+\left\langle E_{k}^{2}(\varepsilon t)\right\rangle\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right] \tag{65}
\end{align*}
$$

In Eq. (65), we have made use of expression for

$$
a_{i} \operatorname{Im} \epsilon_{i j}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{j}
$$

which is shown in Appendix C.

Comparing Eqs. (64) and (65), we find

$$
\begin{equation*}
\left(\frac{\partial U}{\partial t}\right)_{\text {reßənnnt }}=-\sum_{k} \frac{1}{\omega_{q}} \frac{\partial}{\partial \omega_{q}}\left[\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu v}^{+}\left(k, \omega_{q}\right) a_{v}\right] \frac{\partial}{\partial t} \frac{\left\langle E_{k}^{2}\right\rangle}{8 \pi} \tag{68}
\end{equation*}
$$

To calculate the nonresonant contribution, we shall employ the result given by Eq. (B-19) in Appendix B. We see that

$$
\begin{align*}
& \left(\frac{\partial U}{\partial t}\right)_{n o n r e s o n n n t}=\sum_{s} \frac{n_{A} m_{s}}{2} \int d^{3} v v^{2}\left(I_{s}^{I I}\right)_{\text {nonresonnat }} \\
& =-\sum_{k} \sum_{s} \frac{4 \pi n_{s} e_{A}^{2}}{4 m_{s} \omega_{\square}^{2}} \sum_{n=-\infty}^{+\infty} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \\
& \quad \times\left\{a_{z}^{2} v_{z} J_{n}^{2}\left(v_{z}\right) \frac{\partial F_{s}}{\partial v_{z}}+a_{\perp}^{n} \cos ^{2} \psi\left(\frac{n \Omega_{s}}{k_{\perp}}\right)^{2} J_{n}^{2} \frac{1}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right. \\
& \quad+a_{1}^{n} v_{\perp}^{2} \sin ^{2} \psi J_{n}^{\prime 2} \frac{1}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}} \\
& \quad+\left[-\omega_{q} \frac{\partial}{\partial \omega_{q}} P \frac{1}{\omega_{q}-n \Omega_{s}--v_{z} k_{z}}+P \frac{1}{\omega_{q}-k_{z} v_{z}-n \Omega_{s}}\right] \\
& \left.\quad \times\left|Q_{n}^{+1}\right|^{2}\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right\} \tag{67}
\end{align*}
$$

However, from Appendix C, we know that

$$
\begin{align*}
& a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathrm{k}, \omega_{q}\right) a_{\nu} \\
&= 1+2 \pi \sum_{s} \frac{\omega_{B}^{2}}{\omega_{\square}^{2}} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \\
& \times \sum_{n=-\infty}^{+\infty}\left\{P \frac{\left|Q_{n}^{+}\right|^{2}}{\omega_{q}-n \Omega_{s}-k_{z} v_{z}}\left(k \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right. \\
&+a_{\tilde{z}}^{2} v_{z} J_{n}^{2} \frac{\partial F_{s}}{\partial v_{z}}+\frac{a_{\perp}^{2} n^{2} \Omega_{\AA}^{2}}{k_{\perp}^{2}} \cos ^{2} \psi J_{n}^{2} \frac{1}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}} \\
&\left.+a_{\perp}^{2} v_{\perp}^{2} \sin ^{2} \psi J_{n}^{\prime 2} \frac{1}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right\} \tag{68}
\end{align*}
$$

From (67) and (68), we obtain

$$
\begin{align*}
\left(\frac{\partial U}{\partial t}\right)_{\text {nourononant }}= & -\sum_{k}\left\{\left[a_{\mu} R e \epsilon_{p \nu}^{+}\left(\mathrm{k}, \omega_{q}\right) a_{\nu}\right] \frac{\partial}{\partial t} \frac{\left\langle E_{k}^{R}\right\rangle}{8 \pi}\right. \\
& +\frac{1}{\omega_{q}} \frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{\nu}\right) \\
& \left.\times \frac{\partial}{\partial t} \frac{\left\langle E_{k}^{2}\right\rangle}{8 \pi}-\frac{\partial}{\partial t} \frac{\left\langle E_{k}^{\imath}\right\rangle}{8 \pi}\right\} \tag{69}
\end{align*}
$$

However, we know that

$$
\omega_{\square}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathrm{k}, \omega_{q}\right)=c^{2} k^{2}-c^{2} a_{\mu} k_{\mu} k_{\nu} a_{\nu}
$$

Thus,

$$
\begin{align*}
& \left(\frac{\partial U}{\partial t}\right)_{\text {nonrobonnnt }} \\
& =-\sum_{\mathrm{k}}\left\{\left[\frac{c^{2} k^{2}-c^{2} a_{\mu} k_{\mu} k_{\nu} a_{\nu}}{\omega_{q}^{2}}+1\right] \frac{\partial}{\partial t} \frac{\left\langle E_{k}^{2}\right\rangle}{8 \pi}\right\} \tag{70}
\end{align*}
$$

However, it is easy to verify that

$$
\begin{equation*}
\left[\frac{c^{2} k^{2}-c^{2} a_{\mu} k_{\mu} k_{\nu} a_{\nu}}{\omega_{\eta}^{2}}\right]\left\langle E_{k}^{2}\right\rangle=\frac{c^{2}}{\omega_{\|}^{2}}(\mathbf{k} \times a)^{2}\left\langle E_{k}^{2}\right\rangle=\left\langle B_{k}^{2}\right\rangle \tag{71}
\end{equation*}
$$

From Eqs. (66), (69), and (70), we conclude that

$$
\begin{equation*}
\left(\frac{\partial U}{\partial t}\right)_{\mathrm{totnl}}+\frac{\partial}{\partial t} \sum_{\mathbf{k}}\left(\frac{\left\langle E_{k}^{2}\right\rangle}{8 \pi}+\frac{\left\langle B_{k}^{2}\right\rangle}{8 \pi}\right)=0 \tag{72}
\end{equation*}
$$

This proves that the sum of the particle energy and energy associated with the wave fields is constant in time.

## VII. Stationary Solution

In this section, we shall discuss the form of the timeindependent solution of $F_{\mathrm{a}}$. From the amplitude equation, Eq. (59), we find that

$$
\begin{gather*}
\int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \sum_{n=-\infty}^{+\infty} \delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right)\left|Q_{n}^{+}\right|^{2} \\
\times\left\{4 \pi\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\left\langle E_{k}^{2}\right\rangle+m_{s}\left(16 \pi^{2} \omega_{q}^{2}\right)\right. \\
\times \frac{F_{s}\left(v_{z}, v_{\perp}^{2}\right)}{\left.\frac{\partial}{\partial \omega_{q}}\left[\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{\nu}\right]\right\}=0} \tag{73}
\end{gather*}
$$

However, from the kinetic equation we also conclude that any moment $T_{s}$, i.e.,

$$
T_{A}=2_{\pi} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int d v_{z} \Phi_{A}\left(v_{z}, v_{\perp}\right) F_{s}\left(v_{z}, v_{\perp}^{2}\right)
$$

must be independent of time. Thus,

$$
\begin{align*}
& \int_{0}^{\infty} d v_{\perp} v_{!!} \int_{-\infty}^{+\infty} d v_{z} \sum_{n=-\infty}^{+\infty} \delta\left(\omega_{q}-n \Omega_{A}-k_{z} v_{z}\right)\left|Q_{n}^{+}\right|^{2} \\
& \times\left(\frac{\partial \Phi}{\partial v_{z}}+\frac{n \Omega_{*}}{v_{1}} \frac{\partial \Phi}{\partial v_{\perp}}\right)\left\{4 \pi\left(k_{z} \frac{\partial F_{n}}{\partial v_{z}}+\frac{n \Omega_{n}}{v_{\perp}} \frac{\partial F_{\beta}}{\partial v_{\perp}}\right)\left\langle E_{k}^{*}\right\rangle\right. \\
& \left.+m_{s}\left(16 \pi^{2} \omega_{q}^{2}\right) \frac{F_{B}\left(v_{z}, v_{i}^{n}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{\eta}^{2} a_{\beta} R e \epsilon_{\mu \nu}^{+}, a_{v}\right)}\right\}=0 \tag{74}
\end{align*}
$$

From Eqs. (73) and (74), we see that because $\Phi$ is arbitrary, it must be true that

$$
\begin{align*}
& \left\{\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{B}}{v_{\perp}} \frac{\partial F_{\beta}}{\partial v_{\perp}}\right)\left\langle E_{k}^{2}\right\rangle\right. \\
+ & \left.\frac{4 \pi m_{s} \omega_{q}^{2} F_{s}\left(v_{z}, v_{\perp}^{q}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{l}^{2} a_{\mu} R e \epsilon_{\mu 1}^{+} a_{y}\right)}\right\}_{R_{z} v_{z}+\eta \Omega_{s}=\omega_{q}}=0 \tag{75}
\end{align*}
$$

Let us assume that $F_{s}\left(v_{z}, v_{\perp}^{2}\right)=g_{1}\left(v_{z}\right) g_{2}\left(v_{\perp}^{2}\right)$. Hence,

$$
\left\{\left[k_{z} g_{1}^{\prime} g_{2}+2 n \Omega_{s} g_{1} g_{2}^{\prime}\right] \alpha+g_{1} g_{s}\right\}_{k s} v_{s}+n \Omega=\omega_{q}=0
$$

or

$$
\begin{equation*}
\left[k_{z} \frac{g_{1}^{\prime}}{g_{1}}+2 n \Omega_{8} \frac{g_{2}^{\prime}}{g_{2}}\right]_{k_{z} v_{z}+n \Omega_{z}=\omega \omega_{n}} \alpha=-1 \tag{76}
\end{equation*}
$$

where

$$
\begin{equation*}
\alpha=\frac{\left\langle E_{\hbar}^{\prime}\right\rangle}{4 \pi} \frac{1}{m_{\mu} \omega_{\square}^{2}} \frac{\partial}{\partial \omega_{q}}\left\langle\omega_{\AA}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}, a_{\nu}\right\rangle \tag{77}
\end{equation*}
$$

We now define

$$
\begin{equation*}
\frac{g_{1}^{\prime}}{g_{1}}=f_{1}\left(v_{z}\right) \quad \text { and } \quad \frac{g_{2}^{\prime}}{g_{2}}=f_{2}\left(v_{1}^{2}\right) \tag{78}
\end{equation*}
$$

Since $\alpha$ is independent of $v_{z}, v_{1}^{2}$ and $n$, the only choice of $f_{1}$ and $f_{2}$ is

$$
\begin{align*}
& f_{1}=-2 v_{z} A_{A}  \tag{79}\\
& f_{z}=-A_{*} \tag{80}
\end{align*}
$$

where $A_{s}$ is a constant. Thus, from Eq. (76), we find the condition

$$
\begin{equation*}
2 \omega_{q} \alpha A_{n}=1 \tag{81}
\end{equation*}
$$

From Eqs. (78), (79), and (80) we can easily show that

$$
\begin{equation*}
F_{B}=g_{1}\left(v_{z}\right) g_{2}\left(v_{\perp}^{2}\right)=B_{s} \exp \left[-A_{s}\left(v_{z}^{2}+v_{\perp}^{2}\right)\right] \tag{82}
\end{equation*}
$$

In Eq. (82), $B_{\mathrm{a}}$ is a normalization coefficient. Since we require

$$
\begin{equation*}
2 \pi \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} F_{y}\left(v_{z}, v_{\perp}^{4}\right)=1 \tag{83}
\end{equation*}
$$

we find that

$$
\begin{equation*}
B_{y}=\left(\frac{A_{s}}{\pi}\right)^{3 / 2} \tag{84}
\end{equation*}
$$

and $A_{8}$ may be identified with the thermal energy, such that

$$
A_{s}=\frac{m_{s}}{2 \times{ }^{T}}
$$

Therefore, Eq. (82) shows that the time-independent solution of $F_{\Delta}$ is the Maxwellian distribution. When this state is reached, we see that, from Eq. (77),

$$
\begin{equation*}
\left\langle E_{k}^{\mu}\right\rangle=\left\langle\delta E_{k}^{\prime}\right\rangle=\frac{4 \pi \chi T}{\frac{\mathrm{I}}{\omega_{q}} \frac{\partial}{\partial \omega_{q}}\left\langle\omega_{q}^{2} a_{\mu} R e \epsilon_{p}^{+}, a_{v}\right)} \tag{85}
\end{equation*}
$$

for the mode with wave vector k and frequency $\omega_{q}$. However, since in this case $-\omega_{q}(\mathrm{k})$ is also a root of the dispersion equation, the total value of $\left\langle\delta E_{k}^{2}\right\rangle$ should be

$$
\begin{equation*}
\left\langle\delta E_{k}^{2}\right\rangle_{l}=\frac{8 \pi \chi T \omega_{q}}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+} a_{v}\right)} \tag{86}
\end{equation*}
$$

Equation (86) represents the energy spectrum of the fluctuation field. When there is no external field and $\delta \mathbf{E}$ is longitudinal, (86) reduces to

$$
\begin{equation*}
\left\langle\delta E_{k}^{R}\right\rangle_{t} \simeq 4 \pi \chi^{\prime} T \tag{87}
\end{equation*}
$$

which is well known.

## VIII. Special Cases

In this section, we shall consider a number of special cases for which the equations yield simpler forms, Discussion of these cases may facilitate the future applications of the theory.

## A. Electrostatic Instabilities Without an External Magnetic Field

The quasilinear equations for this case reduce to the following form:

$$
\begin{align*}
\frac{\partial F_{s}}{\partial t}= & \sum_{\mathrm{k}} \frac{\pi c_{z}^{2}}{m_{s} k^{2}}\left(\mathrm{k} \cdot \frac{\partial}{\partial \mathrm{v}}\right) \\
& \times\left\{\delta(\omega \mathrm{k} \cdot \mathrm{v})\left[\frac{4 \pi F_{s}(\mathrm{v})}{\frac{\partial}{\partial \omega_{q}}\left(R e \epsilon^{+}\left(\mathrm{k}, \omega_{q}\right)\right)}+\frac{\left\langle E_{k}^{\prime \prime}\right\rangle}{m_{s}}\left(\mathrm{k} \cdot \frac{\partial F_{b}}{\partial \mathrm{v}}\right)\right]\right\} \tag{88}
\end{align*}
$$

$$
\begin{align*}
\frac{\partial\left\langle E_{k}^{n}\right\rangle}{\partial t}= & 2 \gamma\left\langle E_{k}^{a}\right\rangle+8 \pi^{2} \sum_{B} \frac{\omega_{n}^{2} m_{s}}{k^{2}} \\
& \times \int d^{3} v \frac{\delta\left(\omega_{q}-k \cdot v\right) F_{B}(v)}{\left\lvert\, \frac{\partial}{\partial \omega_{q}}\left(\left.R e \epsilon^{+}\left(k, \omega_{q}\right)\right|^{2}\right.\right.} \tag{89}
\end{align*}
$$

where

$$
\begin{equation*}
R e \epsilon^{+}\left(k, \omega_{q}\right)=1+\sum_{\beta} \frac{\omega_{A}^{2}}{k^{2}} p \int d^{3} v\left(k \cdot \frac{\partial F_{s}}{\partial v}\right) \frac{1}{(\omega-k \cdot v)} \tag{90}
\end{equation*}
$$

$$
\begin{equation*}
\gamma=\frac{\pi}{\frac{\partial}{\partial \omega_{q}}\left(R e \epsilon^{+}\left(k, \omega_{q}\right)\right)} \sum_{s} \frac{\omega_{\Delta}^{2}}{k^{2}} \int d^{3} v \delta(\omega-\mathrm{k} \cdot \mathrm{v})\left(\mathrm{k} \cdot \frac{\partial F_{s}}{\partial \mathrm{v}}\right) \tag{91}
\end{equation*}
$$

If we consider electron oscillations and approximate

$$
\begin{equation*}
\frac{\partial}{\partial \omega_{q}}\left(\operatorname{Re} \epsilon^{+}\left(k, \omega_{q}\right)\right) \simeq \frac{2}{\omega_{q}} \tag{92}
\end{equation*}
$$

then it is found that the form of Eqs. (88) and (89) reduces to that obtained previously by Frarris, ${ }^{4}$ who formulated the problem from a quantum mechanical approach, and other authors (Refs. 12 and 13): However, two points should be noted. First, in our theory the ffeld $E_{k}$ is the total wave field (macroscopic field plus fluctuation field). In other words, we have

$$
\left\langle E_{k}^{2}\right\rangle=\widetilde{E}_{k}^{n}+\left\langle\delta E_{k}^{v}\right\rangle
$$

where $\widetilde{E}_{k}$ is the usual macroscopic field. If the plasma satisfies the conditions

$$
\begin{equation*}
\tilde{E}^{2} \gg\left\langle\delta E_{k}^{q}\right\rangle \tag{94}
\end{equation*}
$$

and

$$
\begin{equation*}
\tilde{E}^{a} \gg\left[\frac{4 \pi m_{A} F_{s}(v)}{\left(\mathrm{k} \cdot \frac{\partial F_{s}}{\partial \mathrm{v}}\right) \frac{\partial}{\partial \omega_{q}}\left(R e \epsilon^{+}\left(\mathrm{k}, \omega_{q}\right)\right)}\right]_{v-\omega_{q} / k} \tag{95}
\end{equation*}
$$

then Eqs. (88) and (89) reduce to the conventional quasilinear equations. From this point of view, we can also see that condition (94) alone is not sufficient to validate the usual quasilinear equations. Condition (95) indicates that when $\left(k \cdot \partial F_{s} / \partial v\right)_{v=w / k}$ becomes very small-for instance, when the "plateau" is asymptotically formed in the onedimensional "bump-on-the-tail" problem (Refs. 1 and 2), the efficiency of stimulated emission of plasmons becomes very low, and the spontaneous emission may become significant.

Second, in deriving the equations, we have assumed that the unstable mode has a frequency $\omega_{q}(k)$. In some cases, when $F_{s}$ is symmetrical in velocity space, $-\omega_{q}(\mathrm{k})$ can also be a root of the dispersion equation. Then, if we define each propagating mode as consisting of both plus and minus frequencies, the friction term in Eq. (88) and the source term in Eq. (89) should be modified by a factor of 2 and the energy density is defined as twice the present value.

Finally, it is important to note that in the study of a turbulent plasma in which ion-wave instability is playing a major role, the spontaneous emission can be more significant than that in the case of a growing Langmuir wave, since the phase velocity of the ion wave is low compared with the electron thermal velocity.

[^2]
## B. Electrostatic Instabilities With an External Magnetic Field

The unified quasilinear equations in this case may be obtained by setting $\psi=0$ (longitudinal waves):

$$
\begin{align*}
\frac{\partial F_{s}}{\partial t}= & \sum_{k} \sum_{n=-\infty}^{+\infty} \frac{\pi c_{n}^{n}}{m_{n} k^{n}}\left(k_{z} \frac{\partial}{\partial v_{z}}+\frac{n \Omega_{A}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right) \\
& \times\left\{\delta\left(\omega_{q}-n \Omega_{\pi}-k_{z} v_{z}\right) J_{n}^{n}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{A}}\right)\right. \\
& \left.\times\left[\frac{4 \pi F_{B}\left(v_{z}, v_{\perp}^{2}\right)}{\frac{\partial}{\partial \omega_{\eta}}\left[R e \epsilon^{+}\left(k, \omega_{q}\right)\right]}+\frac{\left\langle E_{k}^{*}\right\rangle}{m_{s}}\left(k_{z} \frac{\partial F_{\beta}}{\partial v_{z}}+\frac{n \Omega_{R}}{v_{\perp}} \frac{\partial F_{\beta}}{\partial v_{\perp}}\right)\right]\right\} \tag{96}
\end{align*}
$$

$$
\begin{align*}
\frac{\partial\left\langle E_{k}^{2}\right\rangle}{\partial t}= & 2 \gamma\left\langle E_{k}^{2}\right\rangle+16 \pi^{3} \sum_{n} \frac{m_{s} \omega_{s}^{2}}{k^{2}} \int_{0}^{\infty} d v_{1} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \\
& \times \sum_{n=-\infty}^{+\infty} \frac{\delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right) J_{n}^{2} F_{s}\left(v_{z}, v_{1}^{2}\right)}{\left|\frac{\partial}{\partial \omega_{q}}\left[R e \epsilon^{+}\left(k, \omega_{q}\right)\right]\right|^{2}} \tag{97}
\end{align*}
$$

where the growth rate is

$$
\begin{align*}
\gamma= & 2 \pi \sum_{s} \frac{\omega_{\bar{z}}^{\frac{2}{2}}}{k^{2}} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \\
& \because \sum_{n=\infty}^{+\infty} \delta\left(\omega_{q}-n \Omega_{\sharp}-k_{z} v_{z}\right) J_{n}^{u}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{s}}\right) \\
& \times\left(k_{z} \frac{\partial F_{\sharp}}{\partial v_{z}}+\frac{n \Omega_{A}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right) \frac{1}{\frac{\partial}{\partial \omega_{q}}\left[R e \epsilon^{+}\left(\mathrm{k}, \omega_{q}\right)\right]} \tag{98}
\end{align*}
$$

We can see easily that when $\left(\omega_{q} / k_{z}\right)<v_{s} \equiv$ thermal velocity of the $s$-species particles, the friction term in Eq. (96) and the source term in Eq. (97) can be very important. Yet, these terms are not included in the usual quasilinear theory.

## C. Cyclotron or Alfvén Waves in a Magnetized Plasma

In this case, $a_{z}=k_{\perp}=0$ and $\psi=\pi / 2$. Thus, we find that $\left|Q_{n}^{+}\right|^{2}=0$ if $n \neq \pm 1$, and $\left|Q_{n}^{+}\right|^{2}=v_{\perp}^{n} / 4$ if $n= \pm 1$.

So far we have considered that the wave is linearly polarized. In order to discuss a cyclotron wave, which is
circularly polarized, we may consider it as the superposition of two linearly polarized modes with a phase difference of $\pi / 2$. This can be readily studied, and the result is given as follows:

$$
\begin{align*}
& \frac{\partial F_{n}}{\partial t}=\sum_{k} \frac{\pi e_{k}^{2}}{m_{n}}\left(\dot{\kappa}_{z} \frac{\partial}{\partial v_{z}} \pm \frac{\Omega_{n}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right) \\
& \times\left\{\delta\left(\omega_{q}-\left( \pm \Omega_{n}\right)-k_{z} v_{z}\right) v_{i}^{z}\right. \\
& \times\left[\frac{2 \pi F_{*}\left(v_{z}, v_{ \pm}^{2}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{\eta}^{2} a_{\mu} R e \epsilon_{\mu \mu^{\prime}}^{+} a_{v}\right)}\right. \\
& \left.\left.+\frac{\left\langle E_{k}^{\prime \prime}\right\rangle}{2 \omega_{\eta}^{2} m_{A}}\left(k_{z} \frac{\partial F_{A}}{\partial v_{z}} \pm \frac{\Omega_{A}}{v_{\perp}} \frac{\partial F_{A}}{\partial v_{\perp}}\right)\right]\right\} \tag{99}
\end{align*}
$$

$$
\begin{align*}
& \times \delta\left(\omega_{q}-\left( \pm \Omega_{s}\right)-k_{z} v_{z}\right) \frac{v_{1}^{2} F_{s}\left(v_{z}, v_{1}^{2}\right)}{\left|\frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(k_{z}, \omega_{q}\right) a_{v}\right)\right|^{2}} \tag{100}
\end{align*}
$$

where

$$
\begin{align*}
\gamma= & \frac{\pi^{2}}{2} \sum_{n} \omega_{\beta}^{\frac{2}{\beta}} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \delta\left[\omega_{q}-\left( \pm \Omega_{s}\right)-k_{z} v_{z}\right] \\
& \times v_{\perp}^{2} \frac{\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}} \pm \frac{\Omega_{s}}{v_{\perp}} \frac{\partial F_{q}}{\partial v_{\perp}}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{q}^{2} a_{\mu} R c \epsilon_{\mu v}^{+} a_{v}\right)} \tag{101}
\end{align*}
$$

It is understood that, in the expression

$$
\frac{\partial}{\partial \omega_{q}}\left(\omega_{\bar{q}}^{\frac{2}{2}} a_{\mu} R e \epsilon_{\mu \nu}^{+}, a_{v}\right)
$$

we should set $a_{z}=k_{\perp}=0$. Furthermore, the proper sign to be chosen depends upon whether the wave is right- or left-circularly polarized. It is useful to remember that $\Omega_{8}$ is negative for the electrons and positive for the ions. Thus from the argument of the delta function in Eq. (99) or (100), we see that for the right-circularly polarized wave (electron cyolotron wave) we should choose the minus sign, otherwise the plus sign.

## IX. Summary and Discussion

Ono flaw in the usual quasilinear theory is that the theory breaks down when the growth rate diminishes. As this condition is approached, the process of resonant diffusion becomes more and more inefficient. For the one-dimensional "bump-on-the-tail" problem, the "dying" diltusion process eventually leads to an ambiguous result, namely, the formation of a plateau at a certain part of the distribution function; and then the "quasilinear" interaction comes to a complete stop. It is true that this difficulty may be resolved by considering the threedimensional problem. However, the basic problem is still there, although it appears less severe.

In this memorandum, we present a unified theory which includes both the macroscopic coherent field and the microscopic fluctuation field. We are especially interested in the case in which the contribution to the pair correlation from the propagating mode of the fluctuation field is considerably more important than that due to the nonpropagating modes and direct particle encounters. For the sake of generality, we have considered both electrostatic and electromagnetic interactions, and also a magnetized plasma. The result may be summarized as follows:

$$
\begin{align*}
& \frac{\partial F_{B}}{\partial t}=\sum_{k} \frac{\pi \mathcal{C}_{n}^{n}}{m_{s}} \sum_{n=-\infty}^{+\infty}\left(k_{2} \frac{\partial}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right) \\
& \times\left\{\left|Q_{n}^{+}\right|^{n} \delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right)\right. \\
& \times\left[\frac{4 \pi F_{s}\left(v_{z}, v_{1}^{2}\right)}{\frac{\partial}{\partial \omega_{q}}\left(\omega_{\eta}^{2} a_{\mu} \operatorname{Re} \epsilon_{\mu v}^{+}\left(\mathbf{k}, \omega_{q}\right) a_{v}\right)}\right. \\
& \left.\left.+\frac{\left\langle E_{k}^{2}(\varepsilon t)\right\rangle}{\omega_{\eta}^{2} m_{s}}\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{\mathrm{s}}}{v_{\perp}} \frac{\partial F_{8}}{\partial v_{\perp}}\right)\right]\right\}  \tag{102}\\
& \frac{\partial}{\partial t}\left\langle E_{k}^{2}(\varepsilon t)\right\rangle=2 \gamma\left(k, \omega_{q}\right)\left\langle E_{k}^{2}(\varepsilon t)\right\rangle \\
& +64 \pi^{4} \omega_{\square}^{2} \sum_{s} \sum_{n=-\infty}^{+\infty} n_{s} e_{\phi}^{2} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{\pi} \\
& \times \frac{\delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right)\left|Q_{\eta}^{+}\right|^{2} F_{s}\left(v_{z}, v_{\perp}^{2}\right)}{\left\lvert\, \frac{\partial}{\partial \omega_{q}}\left[\left.\omega_{\eta}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathrm{k}, \omega_{q}\right) a_{\nu}\right|^{2}\right.\right.} \tag{103}
\end{align*}
$$

where

$$
\begin{aligned}
\left|Q_{n}^{+}\right|^{2}= & {\left[v_{z} a_{z}+\frac{n \Omega_{\Delta} a_{\perp}}{k_{\perp}} \cos \psi\right]^{2} J_{n}^{2}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{\mathrm{a}}}\right) } \\
& +a_{\perp}^{2} v_{\perp}^{2} \sin ^{2} \psi J_{n}^{\prime 2}\left(\frac{k_{\perp} v_{\perp}}{\Omega_{\mathrm{s}}}\right)
\end{aligned}
$$

Equations (102) and (103) show that the time-inclependent solution of $F_{n}$ is the Maxwellian distribution. The conservation of particle and wave energy is also proved by taking into account both the resonant and nonresonant contributions.

The "friction" term in Eq. (102) and the source term in Eq. (103) represent the spontaneous Cerenkov emission of the unstable mode. The usual "collision integral" is considered to be negligible in this case.

For the high-frequency electron wave, the spontaneous emission term in Eq. (102) is admittedly small and of the same order of magnitude as the collision term (Ref. 29). However, for the unstable ion wave, the spontaneous emission can be much more important than the collisional contribution since the population of the emitting electrons is high.

In deriving the amplitude equation, we have made use of the adiabatic approximation; that is, we have treated the distribution function as a time-independent quantity. Strictly speaking, such an approximation may be inconsistent with the theory and, in principle, we should include the effect of slow variation of $F_{s}$ in the derivation. However, such correction turns out to be rather insignificant in the quasilinear theory, as pointed out by several authors (Refs. 6 and 30). Intuitively, this consequence is conceivable from the following point of view. It is expected that the correction terms to the Landau growth rate are proportional to the moments of $\partial F_{s} / \partial t$ since $\gamma$ is independent of particle velocity. However, according to the usual quasilinear theory, we know that only a narrow region of the velocity distribution function evolves significantly due to the resonant diffusion process, and the change of the moment of the entire distribution function with respect to time cannot be very large under ordinary

[^3]circumstances. Thus, the correction to the growth rate cannot be important,

The theory established in the present work is fairly general except for the assumption of weak turbulence.

However, it is possible to improve the theory so that it ean describe strong turbulence. It also appenrs desirnble to generalize the present theory to study instabilities originated by plasma inhomogencity such as, for instance, the drift-wave instability.

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## Appendix A

## Evaluation of the Source Term

The source term may be represented ns

$$
\begin{align*}
& =\lim _{\Delta \rightarrow 0} 2 \Delta \int_{-\infty}^{+\infty} d(0) \frac{\delta\left(\omega-\omega_{q}\right) 16 \pi^{2} \omega^{2}}{\frac{\partial}{\partial \omega}\left|\left[\omega^{2} a_{1} R c \epsilon_{1}(k, \omega) a_{j}\right]\right|^{2}} \\
& \times \sum_{a} e_{n} \sum_{r} e_{r} \int d^{3} v(v \cdot a) \int d^{v^{\prime}} v^{\prime}\left(v^{\prime} \cdot a\right) \\
& \times \int_{0}^{\infty} d t \exp \left[i_{\omega} t-\Delta t-i k \cdot r_{s}(t)\right] \\
& \times \int_{0}^{\infty} d t^{\prime} \exp \left[-i \omega t^{\prime}-\Delta t^{\prime}+i \mathrm{k} \cdot \mathrm{r}_{r}^{\prime}\left(t^{\prime}\right)\right] \\
& \times n_{s} \delta_{a r} \delta\left[v_{s}(t)-v_{r}^{\prime}\left(t^{\prime}\right)\right] F_{s}\left(v_{s}(t)\right) \tag{A-1}
\end{align*}
$$

where we have : nade use of Eq. (24). (Notice: other terms do not contribute.)

If we "translate" the velocities such that $\mathrm{v}_{s}(t) \rightarrow \mathbf{v}$ and $v_{s}\left(t^{\prime}\right) \rightarrow v^{\prime}$, then correspondingly, we should translate $v$ to $v_{s}(-t)$ and $v^{\prime}$ to $v_{r}\left(-t^{\prime}\right)$. Hence, we can write

$$
\begin{align*}
S= & \left.\lim _{\Delta \rightarrow 0} \Delta \frac{32 \pi^{3} \omega_{a}^{2}}{\left\lvert\, \frac{\partial}{\partial \omega_{q}}\left[\omega_{\eta}^{a} a_{1} R e \epsilon_{i}^{\prime}\left(\mathrm{k}, \omega_{a}\right) a_{a}\right]\right.}\right|^{2}
\end{align*} \sum_{a} n_{A} e_{i}^{u} \int d^{3} v^{\prime} .
$$

In Eq. (A-2), we have replaced $\mathrm{r}_{s}(t)$ by $-\mathrm{r}_{s}(-t)$, and $r_{a}^{\prime}(t)$ by $-r_{a}^{\prime}\left(-t^{\prime}\right)$ because of the velocity translation.

Integrating over $v^{\prime}$ and utilizing the relations

$$
\begin{gather*}
v_{t}(-t) \cdot a=v_{z} a_{z}+v_{A} a_{1}\left[\cos \left(\phi+\Omega_{A} t\right) \cos \psi\right. \\
\left.-\sin \left(\phi+\Omega_{4} t\right) \sin \psi\right]  \tag{A-3}\\
i k \cdot\left[r_{4}(-t)-r_{4}^{\prime}\left(-t^{\prime}\right)\right]=-i k_{s} v_{z}\left(t-t^{\prime}\right) \\
-i \frac{k_{1} v_{1}}{\Omega_{s}}\left[\sin \left(\phi+\Omega_{A} t\right) \cdots \sin \left(\phi+\Omega_{s} t^{\prime}\right)\right] \tag{A-4}
\end{gather*}
$$

where $\phi$ is the angle between the vectors $k_{L}$ and $v_{\perp}$ and other quantities are defned after Eq. (59), we find that

$$
\begin{align*}
S= & \lim _{\Delta \rightarrow 0} \frac{32_{0}^{2} \omega_{n}^{2} \Delta \sum_{n} n_{s} e_{n}^{2}}{\left.\left\lvert\, \frac{\partial}{\partial \omega_{\eta}}\left[\omega_{q}^{2} a_{1} R c \epsilon_{j}^{\prime}\left(k, \omega_{q}\right) a_{1}\right]\right.\right]^{2}} \sum_{m=-\infty}^{+\infty} \sum_{n=-\infty}^{+\infty} \int_{0}^{2 \pi} d \phi \\
& \times \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{-\infty}^{+\infty} d v_{z} \int_{0}^{\infty} d t \int_{0}^{\infty} d l^{\prime} \\
& \times \exp \left[i_{\omega_{\eta}}\left(t-t^{\prime}\right)-\Delta\left(t+t^{\prime}\right)-i n \Omega_{n} t\right. \\
& \left.+i m \Omega_{A} t^{\prime}-i m \phi+i n \phi\right]\left\{\left[a_{z} v_{z}+a_{\perp} \cos \psi\left(\frac{n \Omega_{s}}{k_{\perp}}\right)\right]\right. \\
& \left.\times J_{n}(\alpha)-i a_{1} v_{\perp} \sin \psi J_{n}^{\prime}(\alpha)\right\} \\
& \times\left\{\left[a_{z} v_{z}+a_{\perp} \cos \psi\left(\frac{m \Omega_{s}}{k_{\perp}}\right)\right]\right. \\
& \left.\times J_{m}(\alpha)+i a_{\perp} v_{\perp} \sin \psi J_{m}^{\prime}(\alpha)\right\} F_{s}(v) \tag{A-5}
\end{align*}
$$

where we have used the identity

$$
e^{i \alpha \ln \varphi}=\sum_{n=-\infty}^{+\infty} J_{n}(\alpha) e^{i n \phi}
$$

and $\phi$ is defined as

$$
\alpha=\frac{k_{1} v_{\perp}}{\Omega_{s}}
$$

Since we consider that

$$
F_{s}(v)=F_{a}\left(v_{z}, v_{1}^{z}\right)
$$

the $\phi$-integration is non-\%ero only when $m=n$. We finally concludo that

$$
\begin{align*}
S= & \frac{0 \pi^{4} \omega_{\eta}^{2} \sum n_{s} c_{n}^{2}}{\left\lvert\, \frac{\partial}{\partial \omega_{\eta}}\left[\left.\omega_{\eta}^{2} a_{1} R c \epsilon_{j}^{+}\left(k, \omega_{q}\right)\left(a_{j}\right]\right|^{2}\right.\right.} \\
& \times \sum_{n=m \infty}^{+\infty} \int_{0}^{\infty} d v_{\perp} v_{\perp} \int_{m \infty}^{+\infty} d v_{z}\left|Q_{n}^{+}\right|^{n} \\
& \times \delta\left(\omega_{q}-n \Omega_{\pi}-k_{x} v_{s}\right) F_{d}\left(v_{z}, v_{l}^{n}\right) \tag{A-B}
\end{align*}
$$

where

$$
\begin{align*}
& Q_{n}=\left[v_{z} a_{z}+a_{1} \cos \psi\left(\frac{n \Omega_{1}}{k_{1}}\right)\right] J_{n}(\alpha) \pm l a_{L} v_{L} \sin \psi J_{n}^{\prime}(\alpha)  \tag{A-7}\\
& \text { or } \\
& \left|Q_{n}^{n}\right|^{2}=\left[v_{s} a_{x}+a_{1} \cos \psi\left(\frac{n \Omega_{1}}{k_{1}}\right)\right]^{2} J_{n}^{n}(x) \\
& +a_{1}^{4} v_{1}^{2} \sin ^{2} \psi J_{n}^{\prime \prime}(\alpha)
\end{align*}
$$

## Appendix B

## Evaluation of the Interaction Term

The interaction term is denoted by $I_{a}$, i,e.

$$
\begin{align*}
I_{s} ص- & \frac{e_{t}}{m_{A} n_{A}} \int_{0}^{2 \pi} d p \frac{\partial}{\partial \mathrm{v}} \cdot \int_{-\infty}^{+\infty} d \omega \lim _{\Delta \rightarrow 0}\left(\frac{\Delta}{2 \pi^{a}}\right) \\
& \times \sum_{\mathrm{k}}\left[\mathrm{a}+\frac{(\mathrm{v} \cdot \mathrm{a}) \mathrm{k}-(\mathrm{v} \cdot \mathrm{k}) \mathrm{a}}{\omega}\right] \\
& \times\left\langle E_{k}^{*}(\omega, e t) N_{\Delta}^{k}(\mathrm{v}, \omega, e t)\right\rangle \tag{B-1}
\end{align*}
$$

where

$$
\begin{align*}
N_{s}^{k}(v, \omega, \varepsilon t)= & \int_{0}^{\infty} d t \exp \left[i \omega t-\Delta t-i \mathrm{k} \cdot \mathrm{r}_{s}(t)\right] \\
& \times N_{s}^{k}\left(\mathrm{v}_{s}(t), 0\right)-\frac{n_{s} e_{B}}{m_{s}} \\
& \times \int_{0}^{\infty} d t\left[E_{k}(\omega, \varepsilon t)+i \frac{\partial \bar{L}_{s}\left(\omega, \varepsilon t^{\prime},\right.}{\partial t} \frac{3}{\partial t}\right] \\
& \times\left\{\exp \left[i \omega \tau-\Delta \tau-i \mathrm{k} \cdot \mathrm{r}_{s}(\tau)\right]\right. \\
& \times\left[\mathbf{a}+\frac{\left(\mathrm{v}_{a}(\tau) \cdot \mathbf{a}\right) \mathbf{k}-\left(\mathrm{v}_{s}(\tau) \cdot \mathrm{k}\right)}{\omega}\right] \\
& \left.\cdot \frac{\partial F_{s}\left(\mathrm{v}_{s}(\tau)\right)}{\partial \mathrm{v}_{s}(\tau)}\right\}+R_{s}^{k}, \quad(\mathrm{~B}-2) \tag{B-2}
\end{align*}
$$

$$
\begin{align*}
E_{:}^{*}(\omega, \mathrm{et})= & \frac{4 \omega i \pi}{a_{\mu} \alpha_{\mu \nu}^{-}(\mathrm{k}, \omega) a_{\nu}} \sum_{r} e_{r} \int d l^{3} v^{\prime}\left(\mathrm{a} \cdot \mathrm{v}^{\prime}\right) \\
& \times \int_{0}^{\infty} d_{\tau} \exp \left[\ell_{\omega \tau}-\Delta \tau+i \mathbf{k} \cdot \mathrm{r}_{r}(\tau)\right] \\
& \times N_{r}^{-k}\left(\mathbf{v}_{r}(\tau), 0\right)+E_{k} \tag{B-3}
\end{align*}
$$

In Eqs. (B-2) and (B-3), $R_{s}^{k}$ and $\mathcal{E}_{k}$ are two remainder: terms. Since, in the limit $\Delta \rightarrow 0, R_{g}^{k}$ and $\delta_{k}$ contribute only to terms of order $O(\Delta)$, we may ignore both of them.

Hence,

$$
\begin{aligned}
& \lim _{\Delta \rightarrow 0} \frac{\Delta}{2 \pi^{2}}\left\langle E_{k}^{*}(\omega, \varepsilon \hat{i}) N_{s}^{k}(\mathbf{v}, \omega, \varepsilon t)\right\rangle \\
&= \lim _{\Delta \rightarrow 0} \frac{\Delta}{2 \pi^{2}}\left\{-\frac{4 \pi^{2} \omega \delta\left(\omega-\omega_{q}\right)}{\frac{\partial}{\partial \omega}\left[\omega^{2} a_{\mu} R e^{d} \varepsilon_{\mu \nu}(\mathrm{k}, \omega) a_{\nu}\right]}\right. \\
& \times \sum_{r} e_{r} \int d^{3} v^{\prime}\left(\mathbf{v}^{\prime} \cdot \mathfrak{a}\right) \int_{0}^{\infty} d \tau \int_{0}^{\infty} d \tau^{\prime} \\
& \quad \times \exp \left[i \omega\left(\tau-\tau^{\prime}\right)-\Delta\left(\tau+\tau^{\prime}\right)-i \mathbf{k} \cdot \mathbf{r}_{s}(\tau)\right.
\end{aligned}
$$

$$
\begin{align*}
& \left.+i \mathrm{k} \cdot \mathrm{r}_{r}\left(\tau^{\prime}\right)\right]\left\langle N_{s}^{k}\left(\mathrm{v}_{s}(\tau), 0\right) N_{r}^{-k}\left(\mathrm{v}_{r}^{\prime}\left(\tau^{\prime}\right), 0\right\rangle\right\rangle \\
& -\frac{n_{A} e_{A}}{m_{n}} \int_{0}^{\infty} d \tau \exp \left[i \omega \tau-\Delta \tau-i k \cdot r_{s}(\tau)\right] \\
& \times\left\langle E_{k}^{*}(\omega, e t) E_{k}(\omega, e t)\right\rangle \\
& \times\left[\frac{\left(v_{a}(\tau) \cdot a\right) k-\left(v_{s}(\tau) \cdot k\right) a}{\omega}+\mathbf{a}\right] \cdot \frac{\partial F_{a}\left(v_{a}(\tau)\right)}{\partial \mathbf{v}_{A}(r)} \\
& -i \frac{n_{s} c_{s}}{m_{n}} \int_{0}^{\infty} d r \frac{1}{2} \frac{\partial\left\langle E_{k}^{*}(\omega, \varepsilon t) E_{k}(\omega, e t)\right\rangle}{\partial t} \\
& \times \frac{\partial}{\partial \omega}\left[\exp \left(i_{\omega \tau}-\Delta_{T}-i k \cdot r_{s}(r)\right)\right. \\
& \left.\times\left(\frac{\left(v_{\mathbf{8}}(\tau) \cdot \mathfrak{a}\right) \mathbf{k}-\left(v_{\mathbf{a}}(\tau) \cdot \mathbf{k}\right) \mathbf{a}}{\omega}+\mathfrak{a}\right)\right] \\
& \text { - } \frac{\partial F_{s}\left(v_{s}(r)\right)}{\partial \mathbf{V}_{s}} \frac{(r)}{\}} \tag{B-4}
\end{align*}
$$

As before, we see that in $\left\langle N_{s}^{k} N_{r}^{-k}\right\rangle$ only the term which consists of the delta function $\delta\left[\mathbf{v}_{s}(\tau)-\mathbf{v}_{r}^{\prime}\left(r^{\prime}\right)\right]$ would survive in the limit $\Delta \rightarrow 0$. Thus, we may simply write
$\left\langle N_{B}^{k}\left(v_{s}(\tau), 0\right) N_{r}^{-k}\left(v_{r}^{\prime}\left(\tau^{\prime}\right), 0\right)\right\rangle$

$$
\begin{equation*}
=n_{s} \delta_{a r} \delta\left[v_{t}(\tau)-v_{r}^{\prime}\left(\tau^{\prime}\right)\right] F_{s}\left(v_{s}(\tau)\right) \tag{B-5}
\end{equation*}
$$

Again, since we are particularly interested in the unstable mode $\omega=\omega_{q}(\mathrm{k})$, we may use expression (54) and retain the lowest-order contribution,

Hereafter, we write $I_{s}=I_{s}^{1}+I_{8}^{\mathrm{II}}$. From Eqs. (B-5) and (B-6), we have

$$
\begin{align*}
I_{B}^{\mathrm{I}}= & \sum_{\mathbf{k}} \frac{e_{s}^{\mathrm{s}}}{m_{s}} \int_{0}^{a \pi} d \phi \frac{\partial}{\partial \mathrm{v}} \cdot \int_{-\infty}^{+\infty} d \omega \lim _{\Delta \rightarrow 0} \\
& \times \frac{2 \Delta \omega \delta\left(\omega-\omega_{q}\right)}{\frac{\partial}{\partial \omega}\left[\omega^{2} a_{\mu} R e \epsilon_{\mu \nu}(\mathbf{k}, \omega) a_{\nu}\right]}\left[\mathfrak{a}+\frac{(\mathrm{v} \cdot \mathbf{a}) \mathbf{k}-(\mathrm{v} \cdot \mathbf{k}) \mathrm{a}}{\omega}\right] \\
& \times \int d^{3} v^{\prime}\left(\mathbf{v}^{\prime} \cdot \mathbf{a}\right) \int_{0}^{\infty} d t \int_{0}^{\infty} d t^{\prime} \\
& \times \exp \left[i_{\omega}\left(t-t^{\prime}\right)-\Delta\left(t+t^{\prime}\right)-i \mathbf{k} \cdot \mathbf{r}_{s}(t)+i \mathbf{k} \cdot \mathbf{r}_{s}^{\prime}\left(t^{\prime}\right)\right] \\
& \times \delta\left[\mathbf{v}_{s}(t)-\mathbf{v}_{s}^{\prime}\left(t^{\prime}\right)\right] F_{s}\left(\mathbf{v}_{s}(t)\right) \tag{B-6}
\end{align*}
$$

$$
\begin{align*}
I_{s}^{\mathrm{I}}= & \sum_{\mathrm{k}} \frac{e_{A}^{2}}{m_{a}} \int_{0}^{a_{0} \pi} d \phi \frac{\partial}{\partial v} \cdot \int_{-\infty}^{+\infty} d \omega \frac{\delta\left(\omega-\omega_{q}\right)}{2 \pi} \\
& \times\left[\mathfrak{n}+\frac{(v \cdot \Omega) k-(v \cdot k) a}{\omega}\right] \\
& \times\left[\left\langle E_{k}^{z}(t)\right\rangle+\frac{i}{2} \frac{\partial\left\langle E_{k}^{2}(t)\right\rangle}{\partial t} \frac{\partial}{\partial \omega}\right] \\
& \times\left\{\int_{0}^{\infty} d l \exp \left[i \omega \tau-0_{+} \tau-i k \cdot r_{s}(\tau)\right]\right. \\
& \left.\times\left[\mathfrak{n}+\frac{\left(v_{s}(\tau) \cdot \mathfrak{a}\right) k-\left(v_{s}(\tau) \cdot k\right) a}{\omega}\right] \frac{\partial F_{s}\left(v_{s}(\tau)\right)}{\partial v_{s}(\tau)}\right\} \tag{B-7}
\end{align*}
$$

In the following, $I_{s}^{I}$ and $I_{s}^{I I}$ will be discussed separately. First, in Eq. (B-6) we notice $F_{s}\left(v_{a}(t)\right)=i_{a}(v)$ because we consider $F_{s}(v)=F_{s}\left(v_{z}, v_{1}^{2}\right)$. Integrating over $v^{\prime}$ and $\omega$, we find that

$$
\begin{align*}
I_{s}^{\mathrm{I}}= & \sum_{\mathrm{k}} \frac{2 e_{\beta}^{2} \omega_{a}}{m_{a}} \int_{0}^{2 \pi} d \phi \frac{\partial}{\partial \mathrm{v}} \cdot\left[\mathrm{a}+\frac{(\mathrm{v} \cdot \mathrm{a}) \mathbf{k}-(\mathrm{v} \cdot \mathrm{k}) \mathfrak{a}}{\omega}\right] \\
& \times \lim _{\Delta \rightarrow 0} \frac{\Delta}{\frac{\partial}{\partial \omega_{q}}\left[\omega_{q}^{2} a_{\mu} R e \epsilon_{\mu \nu}^{+}\left(\mathrm{k}, \omega_{q}\right) a_{\nu}\right]} \int_{0}^{\infty} d t \\
& \times \int_{0}^{\infty} i t^{\prime} \exp \left[i \omega_{q}\left(t-t^{\prime}\right)-\Delta\left(t+t^{\prime}\right)\right. \\
& \left.-i \mathrm{k} \cdot \mathrm{r}_{s}\left(t-t^{\prime}\right)\right]\left[\mathrm{a} \cdot \mathrm{v}_{s}\left(t-t^{\prime}\right)\right] F_{s}(\mathrm{v}) \tag{B-8}
\end{align*}
$$

In obtaining expression (B-8), we have introduced a "velocity translation" by equating $\mathbf{v}_{s}^{\prime}\left(t^{\prime}\right)$ to $\mathbf{v}_{B}(t)$. Correspondingly, we find that

$$
\begin{gather*}
\mathbf{r}_{s}^{\prime}\left(t^{\prime}\right)=\mathrm{r}_{s}^{\prime}\left(t^{\prime}\right)-\mathrm{r}_{s}^{\prime}(0) \rightarrow \mathrm{r}_{s}(t)-\mathrm{r}_{s}\left(t-t^{\prime}\right)  \tag{B-9}\\
\mathrm{v}^{\prime} \rightarrow \mathrm{v}_{s}\left(t-t^{\prime}\right) \tag{B-10}
\end{gather*}
$$

We shall designate $\tau=t-t^{\prime}$ and employ the relations

$$
\begin{align*}
\mathbf{a} \cdot \mathbf{v}_{s}(\tau)= & a_{z} v_{z}+v_{\perp} a_{\perp} \\
& \times\left[\cos \psi \cos \left(\phi-\Omega_{s} \tau\right)-\sin \psi \sin \left(\phi-\Omega_{s} \tau\right)\right] \tag{B-11}
\end{align*}
$$

$$
\begin{equation*}
\mathrm{k} \cdot \mathrm{r}_{\theta}(\tau)=k_{z} v_{z} \tau-\frac{k_{1} v_{\perp}}{\Omega_{n}} \sin \left(\phi-\Omega_{a} \tau\right) \tag{B-12}
\end{equation*}
$$

After some algebra, we obtain

$$
\begin{align*}
I_{a}^{\mathrm{L}}= & \sum_{k} \frac{4 \pi^{2} c_{a}^{3}}{m_{A}} \sum_{n=-\infty}^{+\infty}\left(k_{z} \frac{\partial}{\partial v_{z}}+\frac{n \Omega_{\mathrm{a}}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right) \\
& \times\left\{\left|Q_{n}^{!}\right|^{2} \delta\left(\omega_{q}-n \Omega_{d}-k_{z} v_{k}\right) \frac{F_{s}(v)}{\frac{\partial}{\partial \omega_{q}}\left[\omega_{\eta}^{2} a_{\mu} R e \epsilon_{\mu \nu} a_{v}\right]}\right\} \tag{B-13}
\end{align*}
$$

where all notations have been defined previously (see Appendix A).

We next study $I_{n}^{1 I}$. If we define

$$
\begin{equation*}
\Lambda=\mathbf{a}+\frac{\left(\mathbf{v}_{8}(t) \cdot \mathbf{a}\right) \mathbf{k}-\left(\mathrm{v}_{8}(t) \cdot \mathrm{k}\right) \mathbf{a}}{\omega_{q}} \tag{B-14}
\end{equation*}
$$

then we see that

$$
\begin{align*}
\Lambda \cdot \frac{\partial F_{s}\left(\mathrm{~V}_{s}(t)\right)}{\partial \mathrm{v}_{s}(t)} & =\left[(\hat{b} \cdot \Lambda) \hat{b}+(\hat{b} \times \Lambda) \times \hat{b} \cos \Omega_{s} t\right. \\
- & \left.(\Lambda \times \hat{b}) \sin \Omega_{s} t\right] \cdot \frac{\partial F_{s}(\mathrm{v})}{\partial \mathbf{v}} \tag{B-15}
\end{align*}
$$

where $F_{s}(v)=F_{s}\left(v_{s}, v_{\mathrm{L}}^{2}\right)$ and $\hat{b}$ is a unit vector parallel to $\mathbf{B}_{0}$. From Eqs. (B-7) 3-14), and (B-15), we can show that

$$
\begin{align*}
I_{\theta}^{1 \mathrm{I}}= & \sum_{\mathrm{k}} \frac{\frac{e_{8}^{2}}{2 m_{\mathrm{s}}^{2} \pi}}{\int_{0}^{z \pi}} d \phi\left\{\frac{\partial}{\partial v_{z}}\right. \\
& \times\left[\left(a_{z}+\frac{v_{\perp} a_{\perp} k_{z} \cos (\phi+\psi)-k_{\perp} v_{\perp} a_{z} \cos \phi}{u_{q}}\right)\right. \\
& \left.\times \mathbb{P}\left(k_{z}, k_{\perp}, v_{z}, v_{\perp}, \omega_{q}\right)\right]+\frac{1}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\left[v _ { \perp } \left(a_{\perp} \cos (\phi+\psi)\right.\right. \\
& \left.+\frac{v_{z} a_{z} k_{\perp} \cos \phi-v_{z} a_{\perp} k_{z} \cos (\phi+\psi)}{\omega_{q}}\right) \\
& \left.\left.\times \mathbb{P}\left(k_{z}, k_{\perp}, v_{z}, v_{\perp}, \omega_{q}\right)\right]\right\} \tag{B-16}
\end{align*}
$$

where

$$
\begin{aligned}
& \mathbb{P}=\left[\left\langle E_{k}^{*}(t)\right\rangle+\frac{i}{2} \frac{\partial\left\langle E_{k}\right\rangle}{\partial t} \frac{\partial}{\partial \omega_{q}}\right]\left\{\int_{0}^{\infty} d t\right. \\
& \times \exp \left[i \omega_{\eta} t-O_{+} t-i k_{z} v_{z} t\right]
\end{aligned}
$$

$\times \exp \left[i \frac{k_{\perp} v_{\perp}}{\Omega_{g}}\left[\sin \left(\phi-\Omega_{s} t\right)-\sin \phi\right]\right]\left[\left(a_{z}\right.\right.$
$\left.+\frac{v_{\perp} a_{\perp} k_{z} \cos \left(\phi+\psi-\Omega_{s} t\right)-k_{\perp} v_{\perp} a_{z} \cos \left(\phi-\Omega_{B} t\right)}{\omega_{q}}\right)$
$\times \frac{\partial F_{s}}{\partial v_{z}}+\left(a_{\perp} \cos \left(\phi+\psi-\Omega_{s} t\right)\right.$
$\left.\left.\left.+\frac{v_{z} a_{z} k_{\perp} \cos \left(\phi-\Omega_{s} t\right)-v_{z} k_{z} a_{\perp} \cos \left(\phi+\psi-\Omega_{s} t\right)}{\omega_{q}}\right) \frac{\partial F_{s}}{\partial v_{\perp}}\right]\right\}$
(B-17)

After some algebra, we find that $I_{8}^{1 \mathrm{I}}$ may be expressed as the sum of two parts, say $\left(I_{d}^{\mathrm{II}}\right)_{\text {robonant }}$ and $\left(I_{B}^{\mathrm{II}}\right)_{\text {uonrobonant }}$, and they take the following form:

$$
\begin{align*}
\left\langle I_{s}^{\mathrm{II}}\right)_{\text {resonnut }}= & \sum_{\mathrm{k}} \frac{\pi e_{8}^{2}}{m_{s}^{2} \omega_{\square}^{2}} \sum_{n=-\infty}^{+\infty}\left(k_{z} \frac{\partial}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\right), \\
& \times\left\{\delta\left(\omega_{q}-n \Omega_{s}-k_{z} v_{z}\right)\left|Q_{n}^{+}\right|^{2}\right. \\
& \left.\times\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right)\right\}\left\langle E_{k}^{2}\right\rangle \quad \text { (B-18 } \tag{B-18}
\end{align*}
$$

$$
\begin{align*}
\left(I_{n}^{I I}\right)_{\text {nonrononnit }}= & \sum_{k} \frac{e_{\sharp}^{2}}{2 m_{n}^{2} \omega_{\eta}^{2}} \sum_{n=m=\infty}^{+\infty}\left\{\frac { \partial } { \partial v _ { z } } \left[\left(a_{z}\left(\omega_{q}-n \Omega_{n}\right) J_{n}(\alpha)\right.\right.\right. \\
& +a_{\perp} \frac{k_{z}}{k_{\perp}} n \Omega_{A} \cos \psi J_{n}(\alpha) \\
& \left.\left.-i a_{\perp} v_{\perp} k_{z} \sin \psi J_{n}^{\prime}(\alpha)\right) \mathbb{R}\left(v_{z}, v_{\perp}, k_{z}, k_{\perp}, \omega_{q}\right)\right] \\
& +\frac{1}{v_{\perp}} \frac{\partial}{\partial v_{\perp}}\left[\left(a_{\perp}\left(\omega_{q}-v_{z} k_{z}\right) \cos \psi J_{n}(\alpha) \frac{n \Omega_{\varepsilon}}{k_{\perp}}\right.\right. \\
& +a_{z} v_{z} n \Omega_{s} J_{n}(\alpha) \\
& \left.-i a_{\perp} v_{\perp}\left(\omega_{q}-k_{z} v_{z}\right) \sin \psi J_{n}^{\prime}(\alpha)\right) \\
& \left.\left.\times \mathbb{a}\left(v_{z}, v_{\perp}, k_{z}, k_{\perp}, \omega_{q}\right)\right]\right\} \quad \text { (B-19) } \tag{B-19}
\end{align*}
$$

where

$$
\begin{align*}
\mathbb{R}= & \left\{a_{z} J_{n}(\alpha) \frac{\partial F_{R}}{\partial v_{z}}+a_{\perp} \frac{\cos \psi}{k_{\perp}} J_{n}(\alpha) \frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right. \\
& +i a_{\perp} \sin \psi J_{n}^{\prime}(\alpha) \frac{\partial F_{A}}{\partial v_{\perp}}+\left[P \frac{1}{\omega_{q}-} \frac{1}{n \Omega_{B}-k_{z} v_{z}}\right. \\
& \left.-\omega_{q} \frac{\partial}{\partial \omega_{q}} P \frac{1}{\omega_{q}-n \Omega_{s}-k_{z} v_{z}}\right] \\
& \left.\times Q_{n}^{+}\left(k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{\sharp}}{v_{\perp}} \frac{\partial F_{z}}{\partial v_{\perp}}\right)\right\} \frac{\partial\left\langle E_{k}^{2}\right\rangle}{\partial t} \tag{B-20}
\end{align*}
$$

In the quasilinear theory, $\left(I_{s}^{\mathrm{I}}\right)_{\text {retomant }}$ is supposed to be much more important than $\left(I_{g}^{\mathrm{II}}\right)_{\text {nonreonant }}$, which represents the contribution from the nonresonant part, and thus the latter is often neglected.

## Appendix C

Evaluation of $a_{\mu} \epsilon_{f w}^{\dagger}(k, \omega) a_{v}$

By definition, we have

$$
\begin{aligned}
& a_{\mu} \epsilon_{\mu \nu}(k, \omega) a_{v}=a_{\mu} a_{\mu}-\frac{4 \pi l}{\omega} \lim _{\Delta \rightarrow 0} \sum_{s} \frac{n_{s} e_{n}^{u}}{m_{s}} \int d d^{s} v(v \cdot a) \\
& \times \int_{\Delta}^{\infty} d t \exp \left[i \omega t-\Delta t-i \mathrm{k} \cdot \mathrm{r}_{g}(t)\right] \\
& \times\left[\mathrm{a}+\frac{\mathrm{k}\left(\mathrm{a} \cdot \mathrm{v}_{A}(t)\right)-\left(\mathrm{v}_{\mathrm{B}}(t) \cdot \mathrm{k}\right) \mathrm{a}}{\omega}\right] \\
& \text { - } \frac{\partial F_{s}\left(v_{s}(t)\right)}{\partial v_{s}(t)} \\
& =1-\frac{i}{\omega} \sum_{n} \omega_{n} \lim _{\Delta \rightarrow 0} \int d^{s} v \int_{v}^{\infty} d t\left(v_{s}(-t) \cdot a\right) \\
& \times \exp \left[i \omega t-\Delta t+i k \cdot r_{n}(-t)\right] \\
& \times\left[a+\frac{k(v \cdot a)-(v \cdot k) a}{\omega}\right] \frac{\partial F_{n}(v)}{\partial v} \\
& =1-\frac{i}{\omega} \sum_{*} \omega_{\omega_{\bar{x}}^{2}} \lim _{\Delta \rightarrow 0} \int d^{3} v \int_{0}^{\infty} d t \\
& \times\left[v_{z} a_{z}+v_{\perp} a_{\perp}\left(\cos \psi \cos \left(\psi+\Omega_{s} t\right)\right.\right. \\
& \left.-\sin \psi \sin \left(\phi+\Omega_{s} t\right)\right] \exp \left[i \omega t-\Delta t-i k_{z} v_{z} t\right. \\
& \left.-i \frac{k_{\perp} v_{\perp}}{\Omega_{g}}\left(\sin \left(\phi+\Omega_{\theta} t\right)-\sin \phi\right)\right]\left\{\left[a_{*}\right.\right.
\end{aligned}
$$

$$
\begin{align*}
& \left.+\frac{v_{\perp} a_{\perp} k_{z}(\cos \psi \cos \phi-\sin \psi \sin \phi)-k_{\perp} v_{\perp} a_{z} \cos \psi}{\omega}\right] \\
& \times \frac{\partial F_{a}}{\partial v_{z}}+\left[a_{\perp}(\cos \psi \cos \phi-\sin \psi \sin \phi)\right. \\
& +\frac{v_{z} a_{z} k_{\perp} \cos \phi-v_{z} k_{z} a_{\perp} \cos \psi \cos \phi}{\omega} \\
& \left.\left.+\frac{v_{z} k_{z} a_{\perp} \sin \psi \sin \phi}{\omega}\right] \frac{\partial F_{z}}{\partial v_{\perp}}\right\} \tag{C-1}
\end{align*}
$$

where $\omega_{i=}^{2}=4 \pi n_{s} e_{s}^{2} / m_{s}$ and the other notations are similar to those used before. After some manipulations, we obtain

$$
\begin{align*}
& +\sum_{n=-\infty}^{+\infty}\left[\frac{a_{1}^{2} n^{2} \Omega_{n}^{2}}{k_{1}^{2}} \cos ^{2} \psi J_{n}^{2}\left(\frac{k_{\perp} \mathcal{N}_{\perp}}{\Omega_{s}}\right)\right. \\
& \left.+a_{1}^{\#} v_{\perp}^{\prime \prime} \sin ^{2} \psi J_{n}^{\prime \prime}\left(\frac{k_{1} v_{\perp}}{\Omega_{s}}\right)\right] \frac{1}{i_{\perp}} \\
& \left.\times \frac{\partial F_{s}}{\partial v_{\perp}}\right\}+2 \pi \sum_{s} \frac{\omega_{n}^{1}}{\omega^{2}} \int_{0}^{\infty} d v_{v_{\perp}} v_{\perp} \int_{0}^{\infty} d v_{*} \\
& \times \sum_{n=-\infty}^{+\infty} \frac{\left|Q_{n}^{+}\right|^{z}}{\omega-k_{z} v_{z}-n \Omega_{s}+i O_{+}} \\
& \times\left\{k_{z} \frac{\partial F_{s}}{\partial v_{z}}+\frac{n \Omega_{s}}{v_{\perp}} \frac{\partial F_{s}}{\partial v_{\perp}}\right\} \tag{C-2}
\end{align*}
$$


[^0]:    For other publications concerning the applications of the quasilinear theory to specific problems, see Ref. 8.

[^1]:    "1'he Initial conclition of the microscopic density $N$ is not "smooth," as wo can see from expression (8). The ensemble-nveraged value of the binary product of these initial conditions will contain a "selfcorrelation" part. It is this part that eventually gives rise to the new contribution which the usual quasilinar theory does not provide. ${ }^{3}$ This technique was used in Section 10 of Ref. 21. Although the diseussion there is concerned with the stable case, extension to a weakly unstable plasma can be made without diffeulty.

[^2]:    "Notice that the contribution from the nonresonant particles vanishes automatically since it is proportional to $\partial\left\langle E_{k}^{2}\right\rangle / \partial t$.

[^3]:    The general collision integral has not been derived in this memorandum. However, when electromagnetic interaction is not important, the collision integral has been discussed by Rostoker (Ref. 28).

