

# Toward large FOV high-resolution X-ray imaging spectrometer: microwave multiplexed readout of 32 TES microcalorimeters

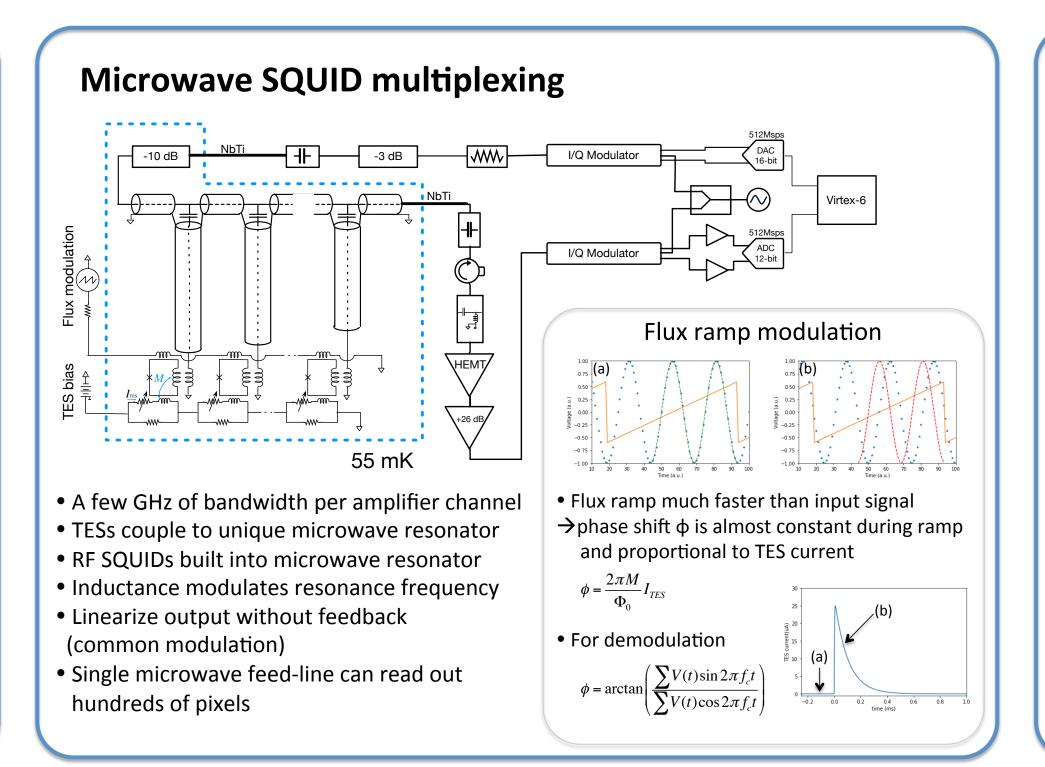
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#### **Abstract**

We performed a small-scale demonstration at GSFC of highresolution x-ray TES microcalorimeters read out using a microwave SQUID multiplexer. This work is part of our effort to develop detector and readout technologies for future space based x-ray instruments such as the microcalorimeter spectrometer envisaged for Lynx, a large mission concept under development for the Astro 2020 Decadal Survey. In this paper we describe our experiment, including details of a recently designed, microwave-optimized lowtemperature setup that is thermally anchored to the 50 mK stage of our laboratory ADR. Using a ROACH2 FPGA at room temperature, we simultaneously read out 32 pixels of a GSFC-built detector array via a NIST-built multiplexer chip with Nb coplanar waveguide resonators coupled to RF SQUIDs. The resonators are spaced 6 MHz apart (at ~5.9 GHz) and have quality factors of ~15,000. Using fluxramp modulation frequencies of 160 kHz we have achieved spectral resolutions of 3-5 eV FWHM on each pixel at 6 keV. We will present the measured system-level noise and maximum slew rates, and briefly describe the implications for future detector and readout design.

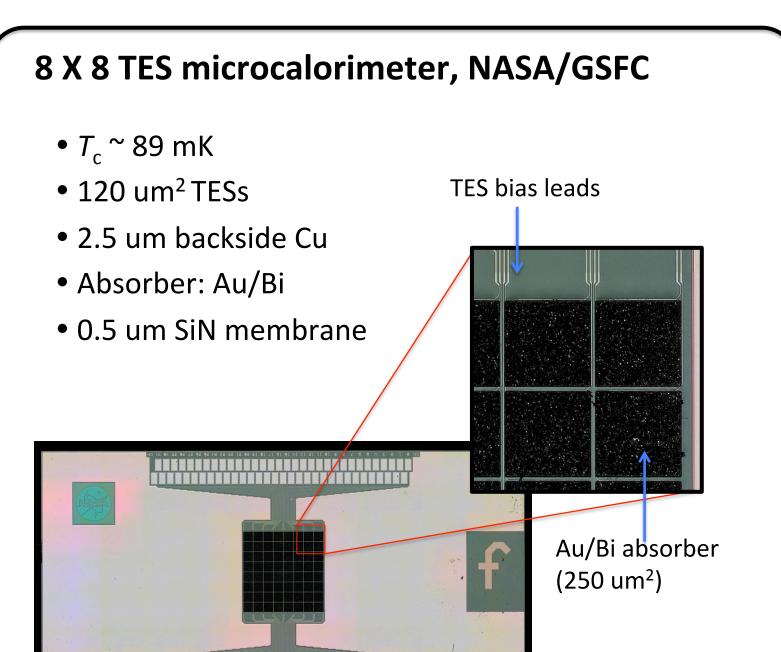


#### Lynx

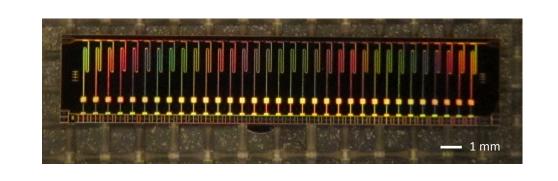
A large mission concept under development by NASA for the Astro 2020 Decadal Survey



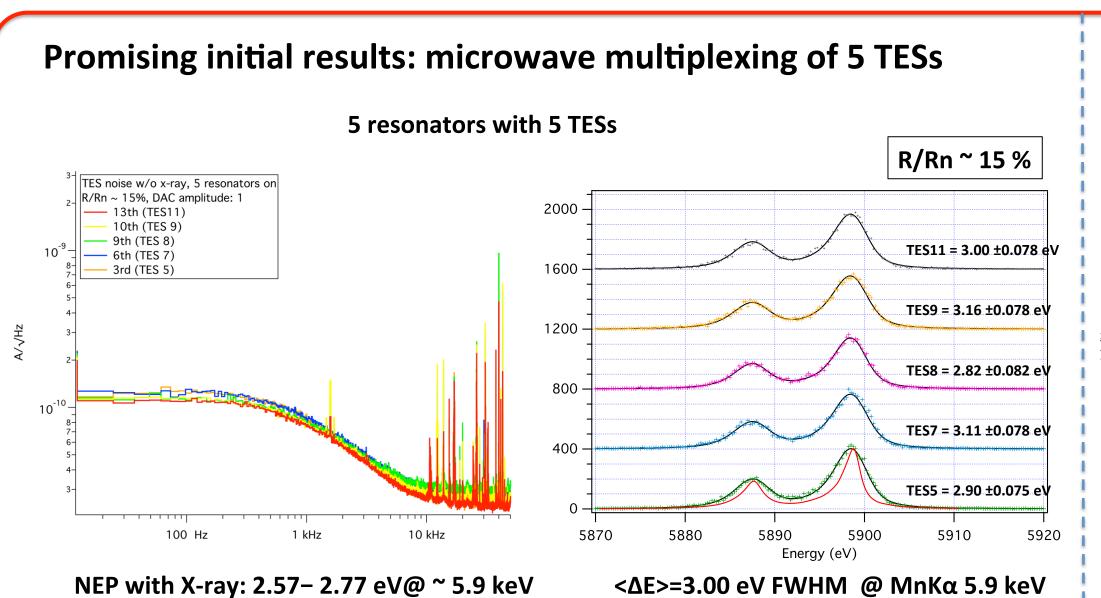
- Energy resolution: better than 3 eV FWHM at 0.2 10 keV
- Number of readout channel > 56,000 (number of pixel > 150,000 with hydra)
- Various angular resolution, energy resolution and count rates
- Sensor: Transition-edge sensor(TES) or magnetically coupled calorimeter (MCC)
- Initial approach: Use position-sensitive TES microcalorimeter, "Hydras". These have multiple absorbers attached to each sensor
- See also
- PE-46, "The Design of the Lynx X-ray Microcalorimeter," Simon Bandler et al.
- PE-59, "Design and optimization of multi-pixel transition-edge sensors for X-ray astronomy application," Stephen J. Smith et al.



#### uMUX chip, NIST



- Fabricated by NIST
- 33 Nb microwave resonators
- Resonance freq.: 5.7-6.0 GHz
- 300 kHz bandwidth
- Frequency spacing: ~6 MHz
- Quality factor ~ 15,000
- Coupling constant: 8.73 (SQUID input – flux ramp circuit)



16 resonator multiplexing (5 TESs, 11 without TES) TES noise w/o x-ray, Resonator 1- 16 on R/Rn ~ 15%, DAC amplitue 0.5 9th (TES 8) 100 Hz

NEP with X-ray: 2.60 - 2.81 eV @ ~ 5.9 keV No NEP degradation compared to five resonator multiplexing!

## Low temperature uMUX setup uMUX chip Nb shield output Coil & aperture 10 dB attenuator T = 55 mK

#### **Room temperature electronics**

- ROACH 2 with MKID ADC/DAC board
- ADC/DAC sampling rate: 512 Msps
- Number of channels: 32
- Bin select: 8 MHz sampling per channel
- Signal bandwidth: 1 MHz



• Flux ramp frequency: 160 kHz, mixed frequency ~ 330 kHz

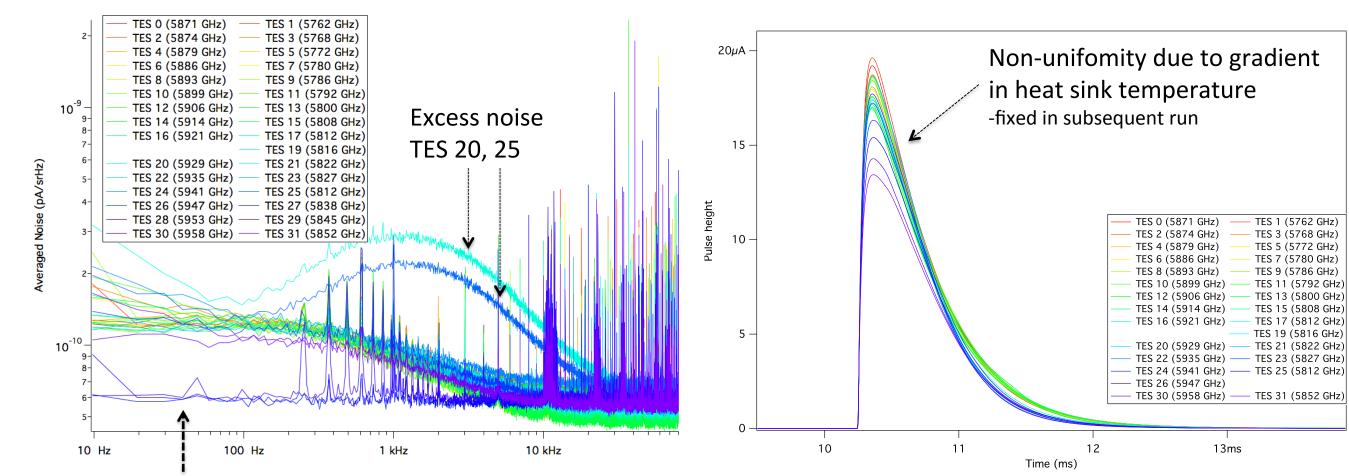
• Throw away transient part, used only 1  $\Phi_0$ 

### Ongoing work: microwave multiplexing of 32 TES microcalorimeters

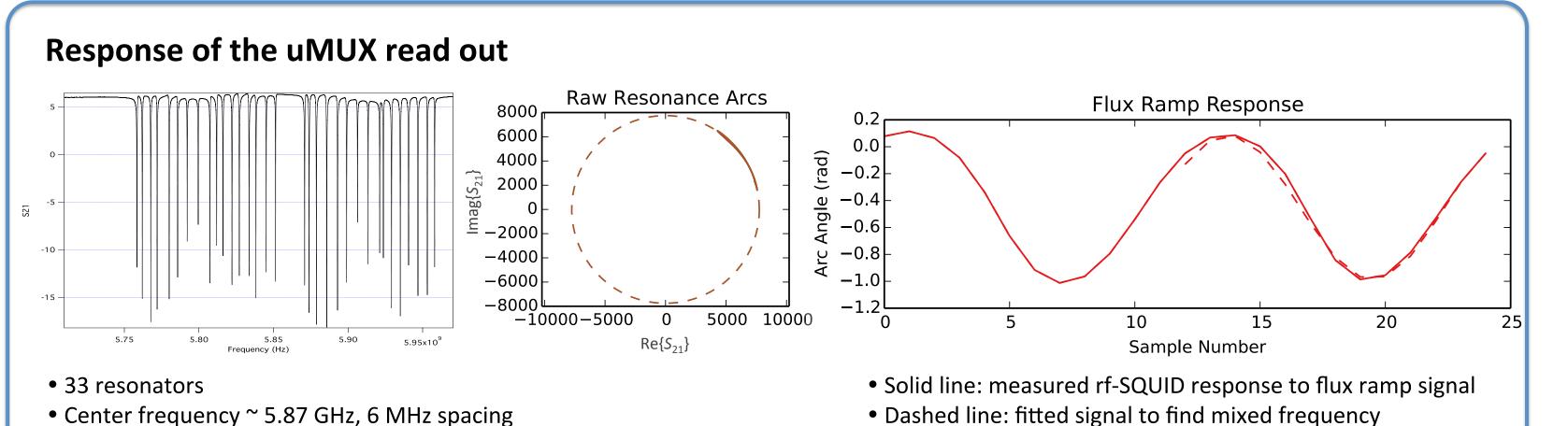
**Changed TES detector chip with the goals of:** 

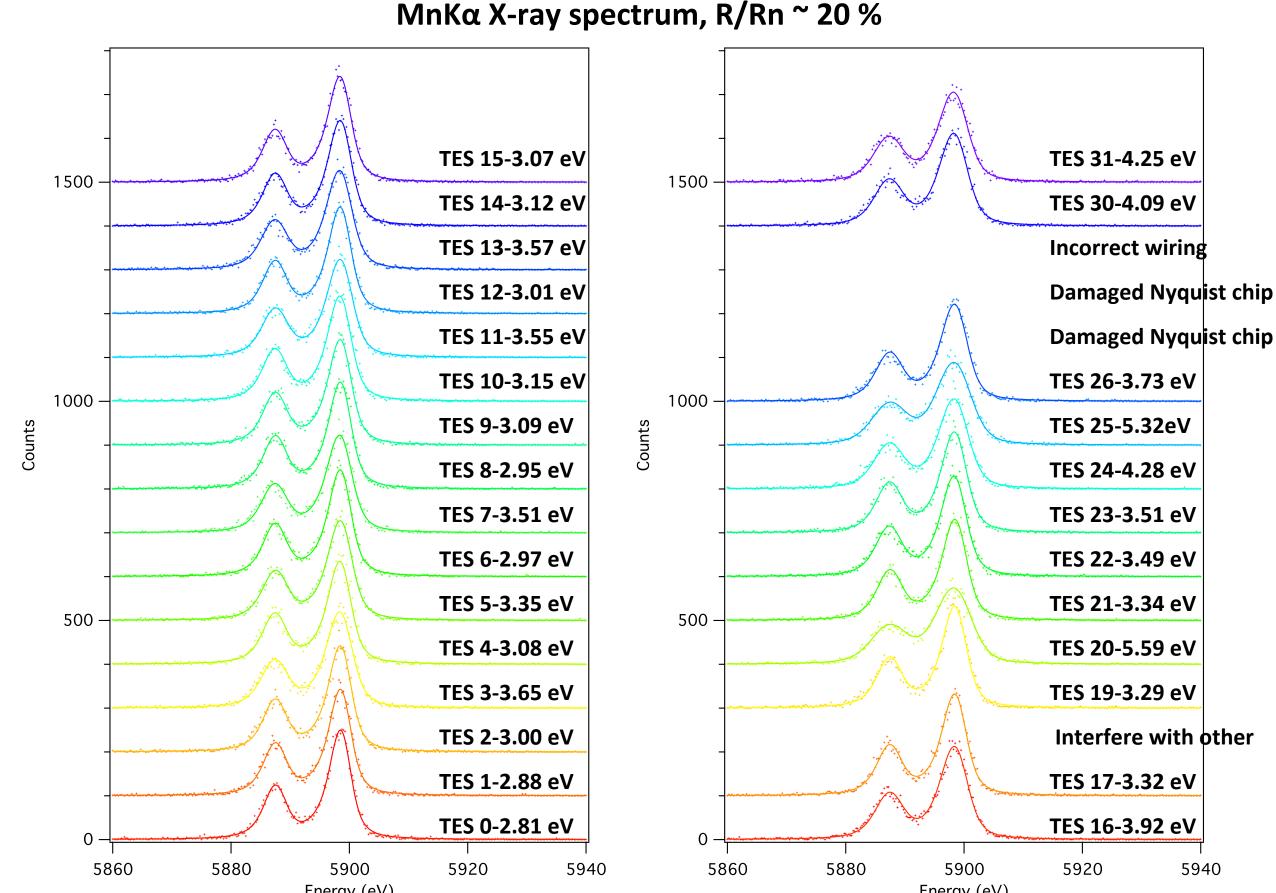
1. Improved energy resolution – expect non-multiplexed  $\Delta E_{FWHM}$ =1.6 eV based on measurements of similar chip

2. 32-channel multiplexing – bond pad layout compatible with μMUX chip layout



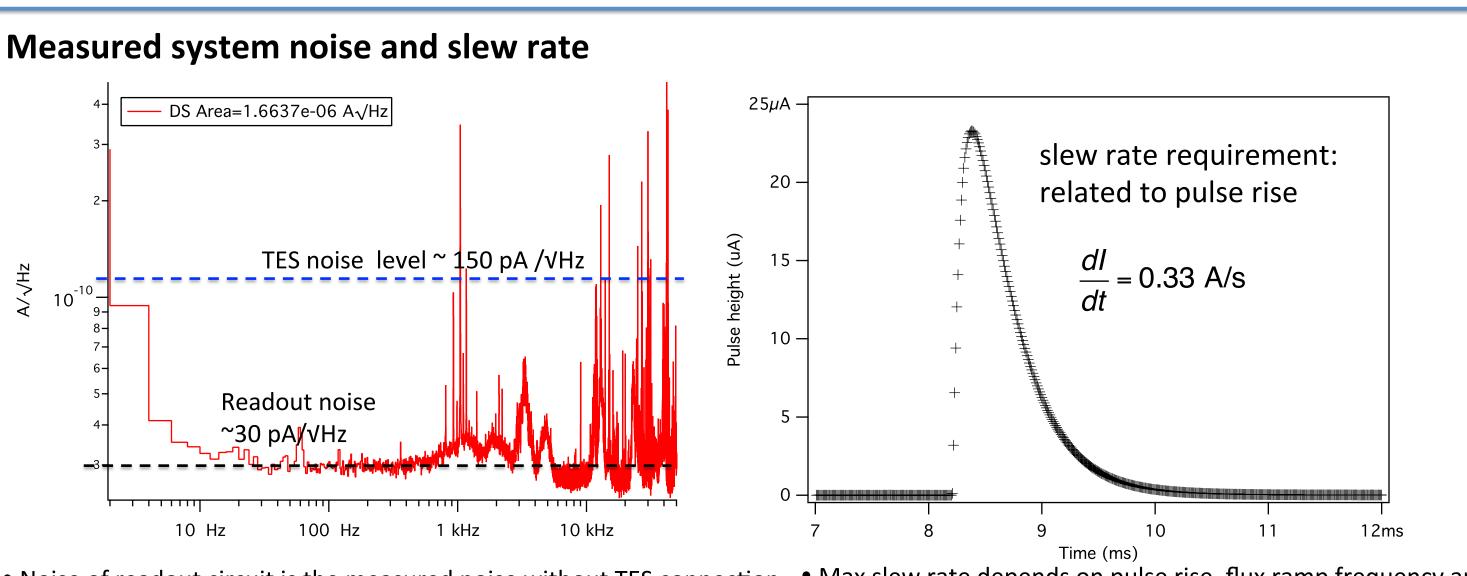
Due to the damaged Nyquist inductor and incorrect wiring





- $<\Delta E> = 3.53$  eV FWHM @ 5.9 keV including all measured pixels - TES 18: Resonator intentionally turned off to avoid interference with other resonator
- TES 20 / 25: Relatively poor energy resolution due to excess detector noise

In progress: fixed wiring, replaced damaged Nyquist chip (4/2017), required new μMUX chip due to subsequent damage during handling, μMUX screening underway (4-7/2017) **Future work**: microwave readout of 'Hydra' pixels for *Lynx* (see Bandler, Smith)



- Noise of readout circuit is the measured noise without TES connection - microwave readout circuit: 30 pA/VHz
- TES noise level: ~ 150 pA/vHz

• Q ~ 13500

 $-Q_{c} \sim 14500 \rightarrow Q_{i} \sim 1.5 - 2 \times 10^{5}$ 

- → uMUX readout noise is a factor of ~5 below the TES noise level
- Max slew rate depends on pulse rise, flux ramp frequency and
  - resonator bandwidth
  - Trade between slew rate capability and resonator packing density Approaches for Lynx:
- slow the rise of the hydra pixels, increase the resonator bandwidth