



Planetary Crater Detection and Registration Using Marked Point Processes, Multiple Birth and Death Algorithms, and Region-based Analysis

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Introduction

Crater Detection

- Marked Point Process Model
- Energy Function
- Multiple Birth and Death Algorithm
- Region-of-Interest Approach
- Experimental Results

Image Registration

- 2-step Approach
- Experimental Results









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Introduction



Need for automated methods for image registration



Objective

- Crater detection in planetary images
- Development of an image registration method based on the extracted features









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Crater detection based on a marked point process (MPP) model

MPP: Stochastic Process Realizations Configurations of objects, each described by a marked point

Mathematical Formulation

A **point process** *X*, defined over a bounded subset *P* of \mathbb{R}^2 maps from a probability space to a **configuration of points** in *P*.

Realizations of the process *X* are random configurations *x* of points, $x = \{x_1, ..., x_n\}$, where x_i is the location of the *i*th point in the image plane $(x_i \in P)$

A configuration of an MPP consists of a point process whose points are enriched with additional parameters, called marks and aimed at parameterizing objects linked to the points.

Bayesian approach: Maximum *a posteriori* (MAP) rule to fit the model to the image is equivalent to minimizing an energy function (computationally challenging)













Energy function of the configuration $X = \{x_i, x_2, ..., x_n\}$ wrt the extracted set *C* of **contour pixels** (Canny):

 $U(X|C) = U_P(X) + U_L(C|X)$

Prior

Repulsion coefficient based on the overlapping of the ellipses (overlapping craters are quite unlikely)

$$U_P(X) = \frac{1}{n} \sum_{x_i \wedge x_j > 0} \frac{x_i \wedge x_j}{x_i \vee x_j}$$

 $x_i \lor x_j$ = area of union of ellipses x_i and x_j $x_i \land x_j$ = area of intersection of x_i and x_j



Likelihood

Two terms, one based on a **correlation** measure, the other based on a **distance** measure (fit between contours and realization of *X*)

$$U_{L}(C|X) = \sum_{i=1}^{n} \left[\frac{d_{\mathcal{H}}(x_{i}^{0}, C)}{na_{i}} - \frac{|x_{i}^{0} \cap C|}{|C|} \right]$$

 x_i^0 = set of pixels corresponding to ellipse x_i in the image plane $d_{\mathcal{H}}(x_i^0, C)$ = Hausdorff distance between ellipse x_i and the contours:

$$d_{\mathcal{H}}(A,B) = \max \left\{ \sup_{\alpha \in A} \inf_{\beta \in B} d(\alpha,\beta) ; \sup_{\beta \in B} \inf_{\alpha \in A} d(\alpha,\beta) \right\}$$

Classical distance between sets d(A, B) = 0





Markov chain Monte Carlo-type method Simulated Annealing scheme



Markov chain sampled by a multiple birth and death (MBD) algorithm







Birth Step

For each pixel *s* in the image, compute the birth probability as $\min\{\delta \cdot B(s), 1\}$, where:

$$B(s) = \frac{b(s)}{\sum_{s} b(s)}$$

b(s) is the **birth map** computed from the contour map using generalized Hough transform and Gaussian filtering



Death Step

For each ellipse x_i in the configuration, compute the death probability as $d(x_i)$:

$$d(\mathbf{x}_i) = \frac{\delta \cdot a(\mathbf{x}_i)}{1 + \delta \cdot a(\mathbf{x}_i)}$$

$$a(\mathbf{x}_i) = \exp\left[-\beta\left(U_L(X \setminus \{\mathbf{x}_i\} | C) - U_L(X | C)\right)\right]$$









Region Based Flowchart and Example









- 6 THEMIS (Thermal Emission Imaging System) images, TIR, 100m resolution, Mars Odissey mission
- 7 HRSC (High Resolution Stereo Color) images, VIS, ~20m resolution, Mars Express mission
- Image sizes from 1581×1827 to 2950×5742 pixels

Quantitative Performance Assessment of the crater detection algorithm: Detection Percentage (D), Branching Factor (B), and Quality Percentage (Q)

Data	$D = \frac{TP}{TP + FN}$	$B = \frac{FP}{TP}$	$Q = \frac{TP}{TP + FP + FN}$
Avg on all THEMIS	0.91	0.10	0.83
Avg on all HRSC	0.89	0.06	0.85
Avg on all images	0.90	0.09	0.84









Crater geometric properties extracted by the proposed method						
Crater	$\boldsymbol{C}=(\boldsymbol{x_0},\boldsymbol{y_0})$	Semi-axes (<i>a</i> , <i>b</i>)	Orientation $ heta$			
Crater 1	(139, 393)	(35, 33)	64°			
Crater 2	(258,756)	(51,50)	115°			
Crater 3	(343,23)	(13, 12)	180°			
Crater 4	(591,215)	(19, 18)	31°			
Crater 5	(919, 157)	(15, 14)	106°			

HRSC Sensor



THEMIS Sensor







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Why a 2-step Optimization?

Feature-based registration

- Min Hausdorff distance $(d_{\mathcal{H}})$ between extracted craters through genetic algorithm
- Fast but sensitive to accuracy of crater maps

Area-based registration

- Max Mutual Information (*MI*) through genetic algorithm
- Highly accurate but computationally heavy















Semi-simulated image pairs

20 pairs composed of one real THEMIS or HRSC image and of an image obtained by applying a synthetic transform and AWGN

Quantitative validation with respect to the true transform (RMSE)



Real multi-temporal image pairs

Real multi-temporal pair of LROC (Lunar Reconnaissance Orbiter Camera) images

100m resolution

Only qualitative visual analysis is available, as no ground truth is available





			Left Image	Right Image	
Data set	RMSE [pixel]	p_{GT}	(7.05, 35.91, 0.18°, 1.071)	(76.59, 19.96, 2.17°, 1.031)	
THEMIS (10 data sets)	0.31	p^{*}	(7.04, 35.92, 0.19°, 1.071)	(76.41, 20.06, 2.18°, 1.031)	
HRSC (10 data sets)	0.22	RMSE 1 st Step	0.79	0.51	
Average (20 data sets)	0.26	RMSE 2 nd Step	0.16	0.33	





Registration Results with Real Data





Visually accurate matching between reference and registered images in the real multitemporal data set





Checkerboard representation of the registered images (zoom on details)





Registration Results with Real Data





Visually accurate matching between reference and registered images in the real multitemporal data set









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Conclusions

- Accurate crater maps, useful for both image registration and planetary science, were obtained from data from different sensors.
- **Higher accuracy** as compared to previous work on crater detection (not shown for brevity)
- Reduced time for convergence thanks to a region-based approach
- Sub-pixel accuracy and visual precision in registration: effectiveness of the proposed 2-step registration method

Future Developments

- Test in conjunction with a **parallel** implementation (e.g. computer cluster)
- Validation with multi-sensor real images
- Extension to **other applications** requiring the extraction of ellipsoidal or circular features, e.g. optical Earth observation images or medical images







- G. Troglio, J. A. Benediktsson, G. Moser and S. B. Serpico, "Crater Detection Based on Marked Point Processes," in *Signal and Image Processing for Remote Sensing*, CRC Press, 2012, p. 325–338.
- G. Troglio, J. Le Moigne, J. A. Benediktsson and G. S. S. B. Moser, "Automatic Extraction of Ellipsoidal Features for Planetary Image Registration," *IEEE Geoscience and Remote Sensing Letters*, vol. 9, pp. 95-99, 2012.
- S. Descamps, X. Descombes, A. Bechet and J. Zerubia, "Automatic Flamingo detection using a multiple birth and death process," in *IEEE International Conference on Acoustics, Speech and Signal Processing*, Las Vegas, NV, 2008.
- X. Descombes, R. Minlos and E. J. Zhizhina, "Object Extraction Using a Stochastic Birth-and-Death Dynamics in Continuum," *Journal of Mathematical Imaging and Vision,* vol. 33, p. 347–359, 2009.
- E. Zhizhina and X. Descombes, "Double Annealing Regimes in the Multiple Birth-and-Death Stochastic Algorithms," *Markov Processes and Related Fields, Polymath,* vol. 18, pp. 441-456, 2012.
- J. Le Moigne, N. S. Netanyahu and R. D. Eastman, Image Registration for Remote Sensing, Cambridge University Press, 2011.
- I. Zavorin and J. Le Moigne, "Use of multiresolution wavelet feature pyramids for automatic registration of multisensor imagery," *IEEE Transactions on Image Processing,* vol. 14, no. 6, pp. 770 782, 2005.
- J. M. Murphy, J. Le Moigne and D. J. Harding, "Automatic Image Registration of Multimodal Remotely Sensed Data With Global Shearlet Features," *IEEE Transactions on Geoscience and Remote Sensing*, vol. 54, no. 3, pp. 1685 1704, 2016.
- J. H. Holland, Adaptation in natural and artificial systems: an introductory analysis with applications to biology, control, and artificial intelligence, University of Michigan Press, 1975, p. 183.











MBD – Birth Step



For each pixel in the image compute the Birth Probability as $\min\{\delta \cdot B(s), 1\}$, where:

$$B(s) = \frac{b(s)}{\sum_{s} b(s)}$$

Being b(s) the **Birth Map** computed from the **Canny Contour Map**









For each ellipse x_i in the configuration compute the **Death Probability** as $d(x_i)$, where

$$d(x_i) = \frac{\delta \cdot a(x_i)}{1 + \delta \cdot a(x_i)} \quad \text{and} \quad a(x_i) = e^{-\beta \left(U_L(\{x \setminus x_i\} | I_g) - U_L(x | I_g) \right)} = e^{\beta \cdot U_L^i(x_i | I_g)}$$

The complete **Flowchart** of the **Death Step** is as follows:









Hausdorff Distance

Similarity = mean_c
$$\left\{ \sum_{i=1}^{N^{c}} \sum_{t=1}^{P} \left[d_{H}(\underline{x}_{i}^{c}, \underline{x}_{t}) \right] \right\}$$

c = craters in Input Image N^{c} = sum(pixels in crater c in Input Image) P = sum(craters'border pixels in Ref Image) \underline{x}_{i}^{c} = coord of pixel i in crater c in Input Image \underline{x}_{t} = coord of pixel t in Ref Image's craters



Mutual Information

$$MI(X,Y) = \sum_{x \in X} \sum_{y \in Y} p_{X,Y}(x,y) \log\left(\frac{p_{X,Y}(x,y)}{p_X(x) \, p_Y(y)}\right)$$

X: pixel intensity in Reference Image *Y*: pixel intensity in Input Image $p_X(x)$: probability density function (pdf) of *X* $p_Y(y)$: probability density function (pdf) of *Y* $p_{X,Y}(x, y)$: joint pdf of *X* and *Y*







Rotation – Scale – Translation Transformation







Region of Interest Approach









Ground Truth Transformation $p_{GT} = (t_{x1}, t_{y1}, \theta_1, k_1) \rightarrow T_{p_{GT}}(x, y) = Q_{p_{GT}} \cdot [x, y, 1]^T$ ComputedTransformation $p = (t_x, t_y, \theta, k) \rightarrow T_p(x, y) = Q_p \cdot [x, y, 1]^T$ Erorr Transformation $p = (t_{x1}, t_{y1}, \theta_1, k_1) \rightarrow T_{p_{GT}}(x, y) = Q_{p_{GT}} \cdot [x, y, 1]^T$ $mathbf{intermation}$ $p = (t_x, t_y, \theta, k) \rightarrow T_p(x, y) = Q_p \cdot [x, y, 1]^T$ $mathbf{intermation}$ $p = (t_x, t_y, \theta, k) \rightarrow T_p(x, y) = Q_p \cdot [x, y, 1]^T$

RMS Error:
$$E(p_e) = \sqrt{\frac{1}{AB} \int_0^A \int_0^B (x' - x)^2 + (y' - y)^2 dx dy}, \qquad \alpha = A^2 + B^2$$

 $E^{2}(p_{e}) = \frac{1}{AB} \int_{0}^{A} \int_{0}^{B} (k_{e} \cos(\theta_{e}) x + k_{e} \sin(\theta_{e}) y + t_{xe} - x)^{2} + (-k_{e} \sin(\theta_{e}) x + k_{e} \cos(\theta_{e}) y + t_{ye} - y)^{2} dxdy$

$$E^{2}(p_{e}) = \frac{\alpha}{3}(k_{e}^{2} - 2k_{e}\cos(\theta_{e}) + 1) + (t_{xe}^{2} + t_{ye}^{2}) - (At_{xe}^{2} + Bt_{ye}^{2})(1 - k_{e}\cos(\theta_{e})) - k_{e}(At_{ye} - Bt_{xe})\sin(\theta_{e})$$