# Integral Averaging and Bifurcation* 

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## 1. Introduction

One of the simplest topological variations of the phase space of a one-parameter family of differential equations (vector fields, flows) is the creation of periodic orbits from equilibria as the parameter crosses a critical value. The study of such topological variations about an equilibrium was initiated and developed by Poincaré perhaps 90 years ago and belongs today to the classical theory of periodic solutions. It was Hopf [23] who presented the bifurcation theorem in 1942 and it is now commonly known as the Hopf bifurcation theorem. Specifically, consider a one-parameter family of ODE (ordinary differential equations)

$$
\dot{x}(t)=f(\alpha, x(t)), \quad \alpha \in R, x \in R^{n} .
$$

Suppose that $f(\alpha, 0) \equiv 0$ and $f$ admits the linearization

$$
\dot{y}(t)=A(\alpha) y(t) .
$$

Assume that $A(\alpha)$ has a pair of complex conjugate eigenvalues $\lambda(\alpha)$ and $\bar{\lambda}(\alpha)$ such that

$$
\operatorname{Re} \lambda^{\prime}(0)>0, \quad \operatorname{Re} \lambda(0)=0, \quad \text { and } \quad \operatorname{Im} \lambda(0) \neq 0 .^{1}
$$

[^0]

Fig. 1. Amplitude-parameter graph: (a) supercritical; (b) subcritical; (c) vertical.

Then, it was shown by Hopf that there are periodic orbits bifurcating from the zero solution. If one plots the amplitude-parameter graph (see Fig. 1), the three situations (a), (b), and (c) are all possible. Moreover, all three cases are of great physical interest. For example, Fig. la represents the first stage of transition to turbulance in fluids as postulated by Landau [31]; Fig. 1b represents an "inverted Hopf bifurcation" which often occurs for flows that exhibit an immediate transition belavior (see [38]); Fig. 1c represents a "degenerate Hopf bifurcation" which is related to the Liapunov center theorem (see [1]). The existence of Hopf bifurcation is a very elementary application of the implicit functions theorem. However, to determine the specific type of bifurcation depends upon certain analytic conditions which involve nonlinearities in the equation.

The purpose of this paper is to describe how, using the classical method of averaging (see $[16,18]$ ) one can give conditions on the vector fields which ensure a supercritical or subcritical Hopf bifurcation. This is done in such a way that the theorems for FDE's (functional differential equations) and PDE's (partial differential equations) are essentially analogous to those of ODE's. The basic idea is to decompose the equation into three coupled equations which are equations for the amplitude $r$, the phase angle $\theta$, and the stable part $y$. The decomposition is natural in the method of integral averaging. By means of a
series of coordinate changes one then decouples the $r$ equation up to a certain order in $r$. Examples from FDE's and PDE's are also included. Many of the transformations and techniques used in this paper are useful for problems involving a vector parameter $\alpha$, as well as radial and angle coordinates ( $r, \theta$ ) which are vectors. This situation arises, for example, in the study of the bifurcation of an invariant torus from around a periodic orbit. See Lanford [32], Ruelle and Takens [40], and Sacker [41].

Since the appearance of Hopf's paper, there have been many papers related to similar problems, notably, Alexander and Yorke [1], Brunovsky [4], Bruslinskaya [5], Chafee [6, 7], Freedman [12], Friedrichs [13], Iooss [24], Josepg and Sattinger [26], Jost and Zehnder [27], Judovich [28], Kopell and Howard [30], Lanford [32], Marsden [35], Marsden and McCraken [36], McCracken [37], Ruelle and Takens [40], Sacker [41], Sattinger [42], Schmidt [43], Sotomayor [44], and Takens [46].

This paper is organized as follows. In Sections 2, 3 the method of averaging is described, without regard to any specific bifurcation problem. Although the exposition here is for ordinary differential equations, the mechanics of the averaging procedure carry over directly to infinite-dimensional systems (functional and partial differential equations). That is, many infinite-dimensional systems can be studied with averaging simply by rewriting them as an ordinary differential equation in a Banach space and proceeding formally from there. A more precise description of these ideas, as well as examples, is found in Sections 7-11.

In Section 4, the averaging method is applied to study the Hopf bifurcation in $R^{2}$, and in Section 5 these ideas are extended to the situation in $R^{n}$. Although a rather detailed study of the Hopf bifurcation is presented here, it should be noted that the techniques used have a much wider application. For example, the problem of bifurcation of an invariant torus from a periodic orbit can be treated in essentially the same manner, as is described at the end of Section 5. Thus the Hopf bifurcation (although an extremely interesting and important phenomenon) is presented here essentially as one illustration of the averaging technique.

In Section 6 the center manifold theorem is described. With this theorem one can give a rigorous proof of the existence of the periodic solutions formally obtained in the previous sections. This idea is especially important for infinitedimensional systems.

The relation between infinite-dimensional systems and ordinary differential equations in a Banach space is explored in Sections 7, 8. The emphasis here is on writing such systems so that averaging can be applied as for finite-dimensional systems. Finally, in Sections 9-11 some specific examples of Hopf bifurcation are studied.

## 2. The Method cf Averaging

Let us begin with a description of the averaging procedure for a class of ordinary differential (ODE's); in later sections this will be generalized and applied to solve bifurcation problems in both ODE's as well as functional differential equations (FDE's) and certain partial differential equations.

Consider first a two-dimensional system in polar coordinates $(r, \theta)$ of period $2 \pi$ in $\theta$, given by

$$
\begin{align*}
& \dot{r}=\epsilon R_{1}(r, \theta, \alpha)+\epsilon^{2} R_{2}(r, \theta, \alpha)+\cdots, \\
& \dot{\theta}=\omega+\epsilon W_{1}(r, \theta, \alpha)+\epsilon^{2} W_{2}(r, \theta, \alpha) \div \cdots, \tag{2.1}
\end{align*}
$$

where $\epsilon$ and $\alpha$ are parameters, $\epsilon \in\left(-\epsilon_{0}, \epsilon_{0}\right) \subseteq R$, and $\omega \neq 0$ is constant. We assume this differential equation is sufficiently smooth for the following calculations to be performed; moreover, since only a finite number of coefficients of $\epsilon^{k}$ will be considered, it is sufficient that (2.1) represent a finite Taylor series, with a remainder term. We seek periodic solutions of (2.1), or integral manifolds in more general systems. In bifurcation problems, $\alpha$ is the bifurcation parameter while $\epsilon$ represents a scaling factor so that we need only consider $r$ near a constant $r_{0}>0$; later, $\alpha$ will be chosen as a particular function of $\epsilon$.

Now if each $R_{j}$ is independent of $\theta$, so $R_{j}(r, \theta, \alpha)=R_{j}(r, \alpha)$, in principle we are done, for the periodic solutions are preciscly those circles $r=r_{0}$ satisfying

$$
\epsilon R_{1}\left(r_{0}, \alpha\right)+\varepsilon^{2} R_{2}\left(r_{0}, \alpha\right)+\cdots=0 .
$$

Thus, we strive to find new coordinates $(\bar{r}, \bar{\theta})$ for (2.1) in which enough of the $R_{j}$ 's are independent of $\bar{\theta}$. (In general, we can only hope a finite number of them will be, but this is sufficient.) In the Hopf bifurcation problem treated in this paper, it is necessary only to transform $r \rightarrow \bar{\eta}$, as the dependence of the $W_{j}$ on $\theta$ is not important. For completeness however, in this section we transform $\theta \rightarrow \bar{\theta}$ as well, as this is important in considering bifurcation from a periodic orbit to a torus. These coordinate changes are constructed via integral averaging; herewe present a description of them.
Suppose the coefficients of $\epsilon^{j}$ for $1 \leqslant j \leqslant k-1$ are independent of $\theta$, so that. we have

$$
\begin{align*}
& \dot{r}=\epsilon R_{1}(r, \alpha)-\cdots+\epsilon^{k-1} R_{k-1}(r, \alpha)+\epsilon^{k} R_{k}(r, \theta, \alpha)+O\left(\epsilon^{k+1}\right), \\
& \dot{\theta}=\omega+\epsilon W_{1}(r, \alpha)+\cdots+\epsilon^{k-1} W_{k-1}(r, \alpha) \div \epsilon^{k} W_{k}(r, \theta, \alpha)+O\left(\epsilon^{k+1}\right) . \tag{2.2}
\end{align*}
$$

Consider a transformation of the form

$$
\begin{equation*}
\ddot{r}=r+\epsilon^{k} u(r, \theta, \alpha), \quad \bar{\theta}=\theta+\epsilon^{k} v(r, \theta, \alpha) . \tag{2.3}
\end{equation*}
$$

Clearly (2.3) brings (2.2) to the form

$$
\begin{aligned}
& \dot{\bar{r}}=\epsilon R_{\mathbf{1}}(\bar{r}, \alpha)+\cdots+\epsilon^{k-1} R_{k-1}(\bar{r}, \alpha)+\epsilon^{k} \bar{R}_{k}(\bar{r}, \bar{\theta}, \alpha)+O\left(\epsilon^{k \mid 1}\right), \\
& \dot{\bar{\theta}}=\omega+\epsilon W_{1}(\bar{r}, \alpha)+\cdots+\epsilon^{k-1} W_{k-1}(\bar{r}, \alpha)+\epsilon^{k} \bar{W}_{k}(\bar{r}, \bar{\theta}, \alpha)+O\left(\epsilon^{k+1}\right),
\end{aligned}
$$

where

$$
\begin{aligned}
& \bar{R}_{k}(\bar{r}, \bar{\theta}, \alpha)=R_{k}(\bar{r}, \bar{\theta}, \alpha)+\omega(\partial u / \partial \theta)(\bar{r}, \bar{\theta}, \alpha) \\
& \bar{W}_{k}(\bar{r}, \bar{\theta}, \alpha)=W_{k}(\bar{r}, \bar{\theta}, \alpha)+\omega(\partial v / \partial \theta)(\bar{r}, \bar{\theta}, \alpha) .
\end{aligned}
$$

Thus $u$ and $v$ must be chosen to make $\bar{R}_{k}$ and $\bar{W}_{k}$ independent of $\bar{\theta}$; this choice is given in the following lemma.

## Lemma 2.1. Consider the relation

$$
\bar{A}(r, \theta, \alpha)=A(r, \theta, \alpha)+\omega(\partial b / \partial \theta)(r, \theta, \alpha)
$$

where $A$ is given and all functions are $2 \pi$-periodic in $\theta$. If $b$ is chosen as

$$
b\left(r, \theta, \alpha_{\alpha}\right)=\frac{-1}{\omega} \int_{0}^{\theta} A(r, s, \alpha) d s+\frac{\theta}{2 \pi \omega} \int_{0}^{2 \pi} A(r, s, \alpha) d s
$$

then $\bar{A}(r, \theta, \alpha)=\bar{A}(r, \alpha)$ is independent of $\theta$. In fact, $\bar{A}$ is the mean value of $A$, so

$$
\bar{A}(r, \alpha)=(1 / 2 \pi) \int_{0}^{2 \pi} A(r, \theta, \alpha) d \theta
$$

Proof. All that has to be checked is the easy fact that $b$ is $2 \pi$-periodic in $\theta$. This lemma then guarantees the existence of a transformation (2.3) so that the coefficients of $\epsilon^{j}, 1 \leqslant j \leqslant k$, in (2.1) are independent of $\theta$. Observe that the sequence of transformations averaging the $\epsilon, \epsilon^{2}, \ldots, \epsilon^{k}$ terms may be written as a single one,

$$
\begin{aligned}
& \bar{r}=r+\epsilon u_{1}(r, \theta, \alpha)+\cdots+\epsilon^{k} u_{7_{k}}(r, \theta, \alpha) \\
& \bar{\theta}=\theta+\epsilon v_{1}(r, \theta, \alpha)+\cdots+\epsilon^{k} v_{k}(r, \theta, \alpha) .
\end{aligned}
$$

The situation in dimension greater than two is somewhat more complicated. The equations here take the form

$$
\begin{align*}
\dot{r} & =\epsilon R_{1}(r, \theta, y, \alpha)+\epsilon^{2} R_{2}(r, \theta, y, \alpha)+\cdots \\
\dot{\theta} & =\omega+\epsilon W_{1}(r, \theta, y ; \alpha)+\epsilon^{2} W_{2}(r, \theta, y, \alpha)+\cdots  \tag{2.4}\\
\dot{y} & =A_{Q} y+\epsilon Y_{1}(r, \theta, y, \alpha)+\epsilon^{2} Y_{2}(r, \theta, y, \alpha)+\cdots
\end{align*}
$$

where $(r, \theta)$ is the rotational part of the differential equation, and $y \in R^{q}$ is the stable (or saddle point) part. It is thus assumed that the constant matrix $A_{Q}$
has no eigenvalues on the imaginary axis. This implies that for any solution $(r(t), \theta(t), y(t))$ of (2.4) bounded for all real $t$, we must have $y=O(\epsilon)$, since (if $A_{O}$ is stable)

$$
\begin{equation*}
y(t)=\epsilon \int_{-\infty}^{t} e^{A_{Q}(t-s)}\left[Y_{1}(r(s), \theta(s), y(s), \alpha)+O(\epsilon)\right] d s \tag{2.5}
\end{equation*}
$$

with the analogous formula in case of a saddle point. It is not clear how to entirely eliminate the presence of $\theta$ and $y$ in $R_{j}$ and $W_{j}$; however, under very general conditions (but not always) we shall see it is possible to average (2.4) so that in the new coordinates,

$$
\begin{align*}
& \bar{R}_{j}(\bar{r}, \bar{\theta}, y, \alpha)=R_{j 0}(\bar{r}, \alpha)+O\left(|y|^{m}\right) \\
& \bar{W}_{j}(\bar{r}, \bar{\theta}, y, \alpha)=W_{j 0}(\bar{r}, \alpha)+O\left(|y|^{m}\right) \tag{2.6}
\end{align*}
$$

where $m=m(j)$ is given and $1 \leqslant j \leqslant k$. In view of (2.5), this is sufficient. In particular, this can always be done if $A_{Q}$ is a stable matrix.

Assuming the coefficients of $\epsilon^{j}, 1 \leqslant j \leqslant k-1$, are already averaged and so satisfy (2.6), let us describe the averaging of $R_{k}(\gamma, \theta, y, \alpha)$; the situation for $W_{T}$ is the same, so it is omitted. First expand $R_{k}$ in powers of $y$,

$$
R_{k}(r, \theta, y, \alpha)=\sum_{l=0}^{m-1} R_{k l}(r, \theta, \alpha) y^{l}+O\left(|y|^{n k}\right)
$$

so $R_{k l}$ takes values in the vector space $\Lambda_{q}{ }^{l}$ of symmetric $l$-linear maps from $R^{q} \times \cdots \times R^{q}=R^{q l}$ into $R$. Letting

$$
\bar{r}=r+\epsilon^{k} u_{k}(r, \theta, y, \alpha)=r+\epsilon^{k} \sum_{l=0}^{m-1} u_{k i l}(r, \theta, \alpha) y^{l}
$$

it is secn that in the transformed variables $(\bar{r}, \theta, y)$, the coefficient of $\epsilon^{k}$ in $\dot{r}$ is

$$
\begin{align*}
\bar{R}_{k}(\bar{r}, \theta, y, \alpha)= & R_{k}(\bar{r}, \theta, y, \alpha)+\omega \frac{\partial u_{k}}{\partial 0}(\bar{r}, \theta, y, \alpha)+\frac{\partial u_{k}}{\partial y}(\bar{r}, \theta, y, \alpha) A_{o} y \\
= & \sum_{i=0}^{m-1}\left[R_{k l}(\bar{r}, \theta, \alpha)+\omega \frac{\partial u_{k l}}{\partial \theta}(\bar{r}, \theta, \alpha)+l u_{k l}(\bar{r}, \theta, \alpha) A_{o}\right] y^{l}  \tag{2.7}\\
& +O\left(|y|^{m}\right)
\end{align*}
$$

The notation $l u_{k l}(\bar{r}, \theta, \alpha) A_{Q}$ of (2.7) needs some explanation. Recall that $u_{k l}$ (here we suppress the arguments $(\bar{r}, \theta, \alpha)$ ) is a symmetric $l$-linear map

$$
u_{k l}: R^{q} \times \cdots \times R^{q}=R^{q l} \rightarrow R
$$

By $u_{k i} A_{Q}$ we mean that element of $A_{q}{ }^{l}$ given by

$$
\left(u_{k l} A_{\varrho}\right)\left(a_{1}, \ldots, a_{l}\right)=(1 / l) \sum_{i=1}^{l} u_{k l}\left(a_{1}, \ldots, A_{\varrho} a_{i}, \ldots, a_{l}\right)
$$

If $l=1, u_{k l}$ is a linear functional on $R^{a}$ and thus may be denoted by a row vector; in this case, $u_{k l} A_{Q}$ is the usual matrix multiplication. Observe also that the map

$$
A_{Q^{l}}: \Lambda_{Q}^{l} \rightarrow A_{a}^{b} \quad \text { by } \quad U \rightarrow U A_{Q}
$$

is linear. Clearly $A_{Q}{ }^{1}$ is just the adjoint of $A_{Q}: R^{q} \rightarrow R^{q}$.
Now, in order to obtain (2.6), it is necessary that

$$
\begin{gather*}
R_{k 0}(\bar{r}, \theta, \alpha)+\omega \frac{\partial u_{k 0}}{\partial \theta}(\bar{r}, \theta, \alpha)=\text { independent of } \theta,  \tag{2.8}\\
R_{k l}(\bar{r}, \theta, \alpha)+\omega \frac{\partial u_{k l}}{\partial \theta}(\bar{r}, \theta, \alpha)+l u_{k l}(\bar{r}, \theta, \alpha) A_{o}=0 \quad \text { for } \quad 1 \leqslant l \leqslant m-1 . \tag{2.9}
\end{gather*}
$$

To get (2.8), simply choose $u_{k i 0}$ as described in Lemma 2.1. The right-hand side of (2.8) is thus the average

$$
(1 / 2 \pi) \int_{0}^{2 \pi} R_{k 0}(\bar{r}, \theta, \alpha) d \theta=(1 / 2 \pi) \int_{0}^{2 \pi} R_{k}(\bar{r}, \theta, 0, \alpha) d \theta
$$

To obtain (2.9), observe that for each ( $\bar{r}, \alpha$ ), $u_{k l}$ satisfies a linear inhomogeneous equation in $\theta$, with periodic forcing term $R_{k l}$. A well-known result in differential equations asserts that (2.9) has a unique $2 \pi$-periodic solution if and only if the homogeneous equation

$$
\omega(d u / d \theta)+l u A_{Q}=0, \quad u \in A_{Q}^{l}
$$

has no nontrivial $2 \pi$-periodic solutions; and this is true if and only if the linear $\operatorname{map} A_{O}^{l}: \Lambda_{q}^{l} \rightarrow \Lambda_{q}^{l}$ has no eigenvalues of the form $i \omega n / l$, for all integers $n$. This is certainly true for $l=1$ since $A_{Q}{ }^{1}: \Lambda_{q}{ }^{1} \rightarrow \Lambda_{q}{ }^{1}$ is just right-matrix multiplication of row vectors by $A_{Q}$, and $A_{0}$ was assumed to have no pure imaginary eigenvalues. However, for $l>1, A_{Q}$ may have $i \omega n / l$ as an eigenvalue, and this motivates the following definition.

Definition. A $q \times q$ matrix $M$ is called $l$-simple if the induced linear transformation $M^{l}: \Lambda_{q}{ }^{l} \rightarrow \Lambda_{q}{ }^{l}$ by $U \rightarrow U M$ has no eigenvalues of the form in/l for all integers $n$.

Thus (2.9) has a unique solution if and only if $(1 / \omega) A_{Q}$ is $l$-simple. In Section 3, we shall give a necessary and sufficient criterion for a matrix to be $l$-simple; in particular, any stable matrix (all eigenvalues in the left half-plane) will be shown to be $l$-simple for all $l \geqslant 1$.

Here let us summarize the above as a theorem.

## Theorem 2.2. Consider the differential equation

$$
\begin{aligned}
& \dot{r}=\epsilon R_{1}(r, \theta, y, \alpha)+\epsilon^{2} R_{2}(r, \theta, y, \alpha)+\cdots \\
& \dot{\theta}=\omega+\epsilon W_{1}(r, \theta, y, \alpha)+\epsilon^{2} W_{2}(r, \theta, y, \alpha)+\cdots \\
& \dot{y}=A_{Q} y+\epsilon Y_{1}(r, \theta, y, \alpha)+\epsilon^{2} Y_{2}(r, \theta, y, \alpha)+\cdots
\end{aligned}
$$

for which the matrix $(1 / \omega) A_{Q}$ is $l$-simple for each $1 \leqslant l \leqslant m-1$. Then there exists a transformation

$$
\bar{r}=r+\epsilon^{2} u(r, \theta, y, x),
$$

where $u$ is a polynomial in $y$ of degree at most $m-1$, such that in the new $(\bar{r}, \theta, y)$ coordinates, the term $R_{l c}$ becomes $R_{k 0}(\bar{r}, \alpha)+\left(|y|^{m}\right)$, where

$$
R_{k: 0}(\bar{r}, \alpha)=(1 / 2 \pi) \int_{0}^{2 \pi} R_{k}(\bar{r}, \theta, 0, \alpha) d \theta
$$

By means of a transformation

$$
\bar{\theta}=\theta+\epsilon^{T} v(r, \theta, y, \alpha)
$$

where v satisfies the same conditions as $u$, the term $W_{k}$ may be similarly averaged.
We shall see that even for the simplest bifurcation problems, it is necessary to average not only $\theta$ terms, but also $y$ terms as above. The problems associated with $(1 / \omega) A_{Q}$ not being $l$-simple, however, do not arise in the generic case.

In studying the bifurcation from a periodic orbit to an invariant torus, one arrives at a system of the form (2.4), except that $R_{j}, W_{j}$, and $Y_{j}$ are also $2 \pi$ periodic in $t$. Consider then

$$
\begin{align*}
\dot{r} & =\epsilon R_{1}(r, \theta, y, t, \alpha)+\epsilon^{2} R_{2}(r, \theta, y, t, \alpha)+\cdots \\
\dot{\theta} & =\omega+\epsilon W_{1}(r, \theta, y, t, \alpha)+\epsilon^{2} W_{2}(r, \theta, y, t, \alpha)+\cdots  \tag{2.10}\\
\dot{y} & =A_{Q} y+\epsilon Y_{1}(r, 0, y, t, \alpha)+\epsilon^{2} Y_{2}(r, 0, y, t, \alpha)+\cdots
\end{align*}
$$

where $\omega, A_{Q}$ are constant as hefore, and all other terms are $2 \pi$-periodic in $t$. To average the term $\epsilon^{k} R_{k}(r, \theta, t, \alpha) y^{l}$, consider a transformation

$$
\bar{r}=r+\epsilon^{k} u_{k i l}(r, \theta, t, \alpha) y^{l}
$$

of period $2 \pi$ in both $\theta$ and $t$. The analogs of (2.8), (2.9) in this case are then

$$
\begin{align*}
& R_{k o}+\omega \frac{\partial u_{k o}}{\partial \theta}+\frac{\partial u_{k o}}{\partial t}=\text { independent of }(\theta, t)  \tag{2.11}\\
& R_{k t}+\omega \frac{\partial u_{k l}}{\partial \theta}+\frac{\partial u_{k i t}}{\partial t}+l u_{k t} \Lambda_{O}=0, \quad l \div 0 \tag{2.12}
\end{align*}
$$

By expanding $R_{k o}$ and $u_{k o}$ in Fourier series

$$
\begin{align*}
R_{k o}(r, \theta, t, \alpha) & =\sum R_{k o m n}(r, \alpha) e^{i(m \theta+n t)}  \tag{2.13}\\
u_{k o}(r, \theta, t, \alpha) & =\sum u_{k o m n}(r, \alpha) e^{i(m \theta+n t)}
\end{align*}
$$

one sees that (2.11) is equivalent to

$$
u_{\text {komin }}(r, \alpha)=[i /(m \omega+n)] R_{\text {komn }}(r, \alpha), \quad(m, n) \neq(0,0)
$$

which in general can be solved only if $\omega m+n \neq 0$ for all integers $(m, n) \neq$ $(0,0)$ occurring in the expansion (2.13) for $R_{k o}$. Very often this expansion is simply a trigonometric polynomial in 0 , so this imposes only finitely many resonance conditions on $\omega$. How such conditions arise in the bifurcation to a torus are described in Section 5.

To solve (2.12), again the Fourier expansion shows that one must assume a condition analogous to the $l$-simplicity of Theorem 2.2, namely, that the induced linear transformation $A_{o}{ }^{l}: \Lambda_{q}^{l} \rightarrow \Lambda_{q}{ }^{l}$ have no eigenvalues of the form $i(m \omega+n) / l$ for integers ( $m, n$ ) appearing in the Fourier expansion of $R_{k l}$. In practice, this restriction is generally of little consequence, as it appears only if terms of sufficiently high order must be averaged. Results of Section 3 show there is never any restriction when $A_{Q}$ is a stable matrix.

## 3. Characterization of $l$-Simplicity

In this section, we prove the following result.
Theorem 3.1. Let $M$ be a $q \times q$ matrix with eigenvalues $\lambda_{1}, \ldots, \lambda_{q}$, and $M^{l}: \Lambda_{q}^{l} \rightarrow \Lambda_{q}^{l}$ the induced linear transformation described in Section 2. Then the eigenvalues of $M^{l}$ are precisely

$$
\left\{(1 / l) \sum_{i=1}^{l} \lambda_{\alpha_{i}} \mid 1 \leqslant \alpha_{1} \leqslant \cdots \leqslant \alpha_{l} \leqslant q\right\} .
$$

Two immediate consequences of this theorem are stated without proof.
Corollary 3.2. The matrix $M$ is $l$-simple if and only if

$$
\sum_{j=1}^{q} n_{j} \lambda_{j} \neq \text { in }
$$

for all integers $n_{j} \geqslant 0$ with $\sum n_{j}=l$ and all integers $n$.

Corollary 3.3. If all eigenvalues of $M$ lie in the left half-plane, then $M$ is $l$-simple for each $l \geqslant 1$.

Proof of Theorem 3.1. It suffices to consider the case where $M$ is diagonalizable, since any $M$ can be approximated by a diagonalizable matrix, and the eigenvalues of $M^{l}$ vary continuously with $M$.

Assume then there is a basis $\left\{e_{1}, \ldots, e_{q}\right\}$ such that $M e_{i}=\lambda_{i} e_{i}$ for each $i$. Letting $\mathscr{A}$ be the set of multiindices

$$
\mathscr{A}=\left\{\alpha=\left(\alpha_{1}, \ldots, \alpha_{\eta}\right) \mid 1 \leqslant \alpha_{1} \leqslant \cdots \leqslant \alpha_{l} \leqslant q\right\}
$$

it is clear that any $U \in \Lambda_{q}{ }^{l}$ is uniquely determined by the values

$$
U\left(e_{\alpha_{1}}, \ldots, e_{\alpha_{l}}\right), \quad \alpha \in \mathscr{A}
$$

For any $\beta \in \mathscr{A}$, let $U_{\beta} \in \Lambda_{q}^{l}$ be defined by

$$
U_{\beta}\left(e_{\alpha_{1}}, \ldots, e_{\alpha_{l}}\right)= \begin{cases}1 & \text { if } \beta=\alpha  \tag{3.1}\\ 0 & \text { if } \beta \neq \alpha\end{cases}
$$

Clearly, the $U_{\beta}$ form a basis for $\Lambda_{q}{ }^{l}$. Now we compute

$$
\begin{aligned}
\left(U_{\beta} M\right)\left(e_{\alpha_{1}}, \ldots, e_{\alpha_{l}}\right) & =\frac{1}{l} \sum_{i=1}^{l} U_{\beta}\left(e_{\alpha_{1}}, \ldots, M e_{\alpha_{i}}, \ldots, e_{\alpha_{l}}\right) \\
& =\left(\frac{1}{l} \sum_{i=1}^{l} \lambda_{\alpha_{i}}\right) U_{\beta}\left(e_{\alpha_{1}}, \ldots, e_{\alpha_{l}}\right) \\
& =\left(\frac{1}{l} \sum_{i=1}^{l} \lambda_{\beta_{i}}\right) U_{\beta}\left(e_{\alpha_{1}}, \ldots, e_{\alpha_{l}}\right)
\end{aligned}
$$

by (3.1). Hence $U_{\beta}$ is an eigenvector of $M^{l}$, with eigenvalue $(1 / l) \sum_{i=0}^{l} \lambda_{\beta_{i}}$. There are no other eigenvectors of $M^{l}$ since the $U_{\beta}$ form a basis.

Remark. In order to solve Eq. (2.11), we must assume

$$
\begin{equation*}
m \omega+n \neq 0 \tag{3.2}
\end{equation*}
$$

for all ( $m, n$ ) $\neq(0,0)$ in the Fourier expansion of $R_{k o}$. In order to solve (2.12), we must assume

$$
\sum_{j=1}^{q} n_{j} \mu_{j} \neq i(m \omega+n)
$$

for all $n_{j} \geqslant 0, \sum n_{j}=l$, where $\left\{\mu_{j}\right\}$ are the eigenvalues of $A_{Q}$.
Remark. Professor Y. Bibikov has pointed out to us the possibility of
dispensing with some of the above restrictions if one is willing to transform the $y$ variable as well as ( $r, \theta$ ), namely, to consider transformations

$$
\bar{y}=y+\epsilon^{k} u(r, \theta, y, \alpha) .
$$

Such ideas are briefly described by Pyartli [39]. In order to average (2.4), the $l$-simplicity restrictions are not necessary. Instead, one further decomposes $y=\left(y_{1}, y_{2}\right)$ into a stable and unstable part, corresponding to eigenvalues with real parts negative and positive. In the equation for $\dot{r}$, the terms involving $y$ are not averaged; rather, in the equations for $\dot{y}_{j}$ one averages those terms which are independent of $y_{j}$. After sufficiently many averagings one sees as in (2.5) that in fact $y(t)=O\left(\epsilon^{N}\right)$. Some of the restrictions that arise in the time dependent case (2.10), specifically (3.2), are, however, essential and cannot be eliminated,

## 4. Hopf Bifurcation for an ODE in $R^{2}$

Our study of bifurcation begins with a discussion of the Hopf bifurcation for an autonomous ODE in the plane $R^{2}$. Consider such a system depending on a scalar parameter $\alpha$ near zero, such that the origin $x=0$ is a fixed point for all $\alpha$. To be specific, consider

$$
\begin{gathered}
\dot{x}=f(x, \alpha), \quad x \in R^{2}, \quad \alpha \in\left(-\alpha_{0}, \alpha_{0}\right) ; \\
f(0, \alpha) \equiv 0 .
\end{gathered}
$$

Assume the linearized equation about $x=0$ is an exponentially stable spiral for $\alpha<0$, a center when $\alpha=0$ with eigenvalues $\pm i \omega_{0} \neq 0$, and an unstable spiral when $\alpha>0$. The eigenvalues of this equation are thus $\gamma(\alpha) \pm i \omega(\alpha)$ where $\gamma(0)=0, \alpha \gamma(\alpha)>0$ for $\alpha \neq 0$, and $\omega(0)=\omega_{0}$. Wc stipulate that $\gamma^{\prime}(0) \neq 0$, and in fact, by using $\gamma(\alpha)$ instead of $\alpha$ as a bifurcation parameter we assume $\gamma(\alpha) \equiv \alpha$. The differential equation then takes the form

$$
\begin{aligned}
\dot{x} & =A(\alpha) x+F(x, \alpha) \\
|F(x, \alpha)| & =O\left(|x|^{2}\right)
\end{aligned}
$$

where $A(\alpha)$ has eigenvalues $\alpha \pm i \omega(\alpha)$ with $\omega(0)=\omega_{0}$. By means of a linear coordinate change $x \rightarrow P(\alpha) x$, where $P(\alpha)$ is an appropriate $2 \times 2$ matrix, we may assume $A(\alpha)$ is in Jordan form

$$
A(\alpha)=\left(\begin{array}{cc}
\alpha & -\omega(\alpha)  \tag{4.1}\\
\omega(\alpha) & \alpha
\end{array}\right) .
$$

Let us write

$$
x=\binom{x^{1}}{x^{2}}, \quad F(x, \alpha)-\binom{F^{1}\left(x^{1}, x^{2}, \alpha\right)}{F^{2}\left(x^{1}, x^{2}, \alpha\right)} .
$$

Expanding in a Taylor series yields

$$
\begin{aligned}
\dot{x}^{1} & =\alpha x^{1}-\omega(\alpha) x^{2}+\sum_{j=2}^{\infty} B_{j}^{1}\left(x^{1}, x^{2}, \alpha\right) \\
\dot{x}^{2} & =\omega(\alpha) x^{1}+\alpha x^{2}+\sum_{j=2}^{\infty} B_{j}^{3}\left(x^{1}, x^{2}, \alpha\right) \\
B_{j}^{i} & =\text { homogeneous polynomial of order } j \text { in }\left(x^{1}, x^{2}\right) .
\end{aligned}
$$

As usual it is sufficient that this expansion be only a finite series. Passing to polar coordinates $\left(x^{1}, x^{2}\right)=(r \cos \theta, r \sin \theta)$ gives

$$
\begin{align*}
& \dot{r}=\alpha r+r^{2} C_{3}(\theta, \alpha)+r^{3} C_{4}(\theta, \alpha)+\cdots \\
& \dot{\theta}=\omega(\alpha)+r D_{3}(\theta, \alpha)+r^{2} D_{4}(\theta, \alpha)+\cdots \tag{4.2}
\end{align*}
$$

where

$$
\begin{aligned}
& C_{j}(\theta, \alpha)=(\cos \theta) B_{j-1}^{1}(\cos \theta, \sin \theta, \alpha)+(\sin \theta) B_{j-1}^{2}(\cos \theta, \sin \theta, \alpha) \\
& D_{\lambda}(\theta, \alpha)=(\cos \theta) B_{j-1}^{2}(\cos \theta, \sin \theta, \alpha) \cdots(\sin \theta) B_{j-1}^{1}(\cos \theta, \sin \theta, \alpha)
\end{aligned}
$$

Observe that $C_{j}$ and $D_{j}$ are homogeneous polynomials of degree $j$ in $(\cos \theta, \sin \theta)$. We seek, for $\alpha \rightarrow 0$, periodic solutions of (4.2) with $r \rightarrow 0$. Scale $r$ and $\alpha$ by replacing

$$
r \rightarrow \epsilon r, \quad \alpha \rightarrow \epsilon \alpha,
$$

where the new $r$ is to be considered near a constant $r_{0}>0$, to be determined later; we shall also later specify $\alpha$ as a function of $\epsilon$, but it is not clear how to do this yet. After scaling, (4.2) becomes

$$
\begin{align*}
& \dot{r}=\epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)\right]+\epsilon^{2} r^{3} C_{1}(\theta, \epsilon \alpha)+\cdots \\
& \dot{\theta}=\omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon \alpha)\right]+\cdots \tag{4.3}
\end{align*}
$$

Although this is not quite in the form (2.1) (since $C_{j}$ and $D_{j}$ depend on $\epsilon$ ) it is clear that the averaging procedure still works.

Let us now work through the averaging of (4.3). The generic situation will be completely determined by averaging the $\epsilon$ and $\epsilon^{2}$ terms in $\dot{r}$, and, in fact, there is no need to average the terms in $\dot{\theta}$. Thus the coordinate change

$$
\begin{equation*}
\bar{r}=r+\epsilon u_{1}(r, \theta, \alpha, \epsilon)+\epsilon^{2} u_{2}(r, \theta, \alpha, \epsilon) \tag{4.4}
\end{equation*}
$$

is considered. The argument $\epsilon$ appears in $u_{1}$ and $u_{2}$ since it appears in the coefficients of the expansion in (4.3). Note that the inverse of (4.4) is

$$
r=\bar{r}-\epsilon u_{1}(\bar{r}, \theta, \alpha, \epsilon)+O\left(\epsilon^{2}\right)
$$

Substitution into (4.3) yields

$$
\begin{aligned}
\dot{r}= & \epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)\right]+\epsilon^{2} r^{3} C_{4}(\theta, \epsilon \alpha) \\
& +\epsilon^{2} \frac{\partial u_{1}}{\partial r}(r, \theta, \alpha, \epsilon)\left[\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)\right] \\
& +\epsilon \frac{\partial u_{1}}{\partial \theta}(r, \theta, \alpha, \epsilon)\left[\omega_{0}+\epsilon \alpha \omega^{\prime}(0)+\epsilon r D_{3}(\theta, \epsilon \alpha)\right] \\
& +\epsilon^{2} \frac{\partial u_{2}}{\partial \theta}(r, \theta, \alpha, \epsilon) \omega_{0}+O\left(\epsilon^{3}\right) \\
= & \epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)+\frac{\partial u_{1}}{\partial \theta}(r, \theta, \alpha, \epsilon) \omega_{0}\right] \\
& +\epsilon^{2}\left[r^{3} C_{4}(\theta, \epsilon \alpha)+\frac{\partial u_{1}}{\partial r}(r, \theta, \alpha, \epsilon)\left(\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)\right)\right. \\
& \left.+\frac{\partial u_{1}}{\partial r}(r, \theta, \alpha, \epsilon)\left(\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon \alpha)\right)+\frac{\partial u_{2}}{\partial \theta}(r, \theta, \alpha, \epsilon) \omega_{0}\right] \\
== & {\left[\alpha \bar{r}+\bar{r}^{2} C_{3}(\theta, \epsilon \alpha)+\frac{\partial u_{1}}{\partial \theta}(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right] } \\
& +\epsilon^{2}\left[\bar{r}^{3} C_{4}(\theta, \epsilon \alpha)+\frac{\partial u_{1}}{\partial r}(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha \bar{r}+\bar{r}^{2} C_{3}(\theta, \epsilon \alpha)\right)\right. \\
& +\frac{\partial u_{1}}{\partial \theta}(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha \omega^{\prime}(0)+\bar{r} D_{3}(\theta, \epsilon \alpha)\right) \\
& -u_{1}(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha+2 \bar{r} C_{3}(\theta, \epsilon \alpha)+\frac{\partial^{2} u_{1}}{\partial r}(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right) \\
& \left.+\frac{\partial u_{2}}{\partial \theta}(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right]+O\left(\epsilon^{3}\right) .
\end{aligned}
$$

Following Lemma 2.1, $u_{1}$ is given by

$$
u_{1}(r, \theta, \alpha, \epsilon)=-\left(r^{2} / \omega_{0}\right) \int_{0}^{\theta} C_{3}(s, \epsilon \alpha) d s
$$

since $C_{3}$, being a homogeneous trigonometric polynomial of degree 3 , has mean value zero. The coefficient of $\epsilon$ is thus $\alpha \vec{r}$. Next, $u_{2}$ is chosen as in the lemma; it is not necessary to determine $u_{2}$ explicitly since the coefficient of $\epsilon^{2}$ is the mean value

$$
\text { mean } \begin{aligned}
& {\left[\vec{r}^{3} C_{4}+\left(\frac{\partial u_{1}}{\partial r}\right)\left(\alpha \bar{r}+\bar{r}^{2} C_{3}\right)+\left(\frac{\partial u_{1}}{\partial \theta}\right)\left(\alpha \omega^{\prime}(0)+\bar{r} D_{3}\right)\right.} \\
& \left.-\left(u_{1}\right)\left(\alpha+2 \tilde{r} C_{3}+\frac{\partial^{2} u_{1}}{\partial r \partial \theta} \omega_{0}\right)\right] \\
= & \text { mean }\left[\bar{r}^{3} C_{4}-\frac{\bar{r}^{3}}{\omega_{0}} C_{3} D_{4}\right] \stackrel{\text { def }}{=} \vec{r}^{3} K .
\end{aligned}
$$

This is summarized as a theorem.

Theorem 4.1. Consider the differential equation

$$
\begin{align*}
& \dot{r}=\epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon \alpha)\right]+\epsilon^{2} r^{3} C_{4}(\theta, \epsilon \alpha)+O\left(\epsilon^{3}\right), \\
& \dot{\theta}=\omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon \alpha)\right]+O\left(\epsilon^{2}\right) \tag{4.5}
\end{align*}
$$

arising from the Hopf bifurcation problem in $R^{2}$ described above, and the scaling $r \rightarrow \epsilon r, \alpha \rightarrow \epsilon \alpha$. Then there exists a coordinate change

$$
\bar{r}==r+\epsilon u_{1}(r, \theta, \alpha, \epsilon)+\epsilon^{2} u_{2}(r, \theta, \alpha, \epsilon)
$$

transforming (4.5) into the averaged system of the form

$$
\begin{align*}
& \dot{\vec{r}}=\epsilon \alpha \bar{r}+\epsilon^{2} \vec{r}^{3} K+O\left(\epsilon^{3}\right) \\
& \dot{\theta}=\omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon \alpha)\right]+O\left(\epsilon^{2}\right) \tag{4.6}
\end{align*}
$$

where $K$ is the constant

$$
\begin{equation*}
K=(1 / 2 \pi) \int_{0}^{2 \pi} C_{4}(\theta, 0)-\left(1 / \omega_{0}\right) C_{3}(\theta, 0) D_{3}(\theta, 0) d \theta \tag{4.7}
\end{equation*}
$$

The generic case occurs when $K \neq 0$; for definiteness suppose $K<0$. This suggests the choice $\alpha=\epsilon$, for then (4.6) becomes (dropping the bars)

$$
\begin{aligned}
& \dot{r}=\epsilon^{2}\left(r+r^{3} K\right)+O\left(\epsilon^{3}\right), \\
& \dot{\theta}=\omega_{0}+O(\epsilon)
\end{aligned}
$$

so a periodic solution for $r$ near

$$
r_{0}=(-K)^{-1 / 2}
$$

seems likely. This is, indeed, the case although we must verify that all periodic solutions are obtained in this manner (i.e., none are lost in scaling).

To see this, consider any periodic solution of the unscaled equation (4.2) bifurcating from $r=0, \alpha=0$. At some point $\left(r_{1}, \theta_{1}\right)$ of the solution $\dot{r}$ must vanish, and after scaling by $\epsilon=r_{1} / r_{0}, \dot{r}$ vanishes at $\left(r_{0}, \theta_{1}\right)$. Thus

$$
0=\epsilon \alpha r_{0}+\epsilon^{2} r_{0}^{3} K+O\left(\epsilon^{3}\right)=\epsilon r_{0}\left(\alpha-\epsilon+O\left(\epsilon^{2}\right)\right)
$$

so that $\alpha=\epsilon+O\left(\epsilon^{2}\right)$. Consider an annulus of the form

$$
\mathscr{A}\left\{\begin{array}{l}
(1-\gamma) r_{0} \leqslant r \leqslant(\mathrm{l}+\gamma) r_{0},  \tag{4.8}\\
\gamma \rightarrow 0 \text { as } \epsilon \rightarrow 0 .
\end{array}\right.
$$

We see that for an appropriate $\gamma=\gamma(\epsilon)$, $\alpha$ must be positively invariant since

$$
\begin{aligned}
& r=(1+\gamma) r_{0} \Rightarrow \dot{r}=\epsilon^{2}(1+\gamma) r_{0}\left(\frac{\alpha}{\epsilon}-(1+\gamma)^{2}+O(\epsilon)\right)<0, \\
& r=(1-\gamma) r_{0} \Rightarrow \dot{r}=\epsilon^{2}(1-\gamma) r_{0}\left(\frac{\alpha}{\epsilon}-(1-\gamma)^{2}+O(\epsilon)\right)>0
\end{aligned}
$$

hence the periodic solution lies entirely in $\mathscr{A}$. This then implies the following theorem.

Theorem 4.2. Let the constant $K$ defined in (4.7) satisfy $K<0$, and let $r_{0}=(-K)^{-1 / 2}$. Then all periodic solutions of the original bifurcation problem

$$
\hat{x}=f(x, \alpha), \quad x \in R^{2}
$$

bifurcating from the origin $r=0, \alpha=0$ may be obtained by scaling $r \rightarrow \epsilon \gamma, \alpha \rightarrow \epsilon \alpha$ and averaging as above to obtain

$$
\begin{aligned}
& \dot{\vec{r}}=\epsilon \alpha \bar{r}+\epsilon^{2} \bar{r}^{3} K+O\left(\epsilon^{3}\right), \\
& \dot{\theta}=\omega_{0}+O(\epsilon),
\end{aligned}
$$

then letting $\alpha=\epsilon$ and considering $\bar{r}$ near $r_{0}$. About each such solution, there is a positively invariant annulus $\mathscr{A}$ as in (4.8).

If $K>0$, the same result holds except now $r_{0}=K^{-1 / 2}, \alpha=-\epsilon$, and $\mathscr{A}$ is negatively invariant. Thus in either case, we have in the unscaled variables

$$
\begin{aligned}
& \alpha=-(\operatorname{sgn} K) \epsilon^{2}, \\
& r \sim|K|^{-1 / 2} \epsilon .
\end{aligned}
$$

In the critical case $K=0$ nothing more can be said until more terms of (4.5) are averaged. In order to study this situation, the following lemma is useful.

Lemma 4.3. Consider the system (4.3), but expanded in powers of $\epsilon$

$$
\begin{align*}
& \dot{r}=\epsilon\left[\alpha r+r^{2} C_{3}(\theta)\right]+\epsilon^{2}\left[r^{3} C_{4}(\theta)+\alpha r^{2} C_{3}^{1}(\theta)\right]+\cdots,  \tag{4.9}\\
& \dot{\theta}=\omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta)\right]+\cdots,
\end{align*}
$$

where $C_{j}(\theta)=C_{j}(\theta, 0), C_{j}^{1}(\theta)=\left(\partial / \partial \partial_{x}\right) C_{j}(\theta, 0)$, etc., and let the coefficients of $\epsilon, \ldots, \epsilon^{k}$ in $\dot{r}$ be averaged by a series of coordinate changes

$$
\bar{r}=r+\epsilon^{j} u_{j}(r, \theta, \alpha), \quad 1 \leqslant j \leqslant k
$$

Then the following properties hold for the averaged system:
(1) For all $m$, the coefficient $R_{m}(r, \theta, \alpha)$ of $\epsilon^{m i n}$ in $\dot{\gamma}$ is a polynomial in $(r, \alpha)$ of the form

$$
R_{m}(r, \theta, \alpha)=\sum_{j=0}^{m+1} r^{m+1-j} \alpha^{j} P_{m i j}(\theta)
$$

(2) For all $m$, the coefficient $W_{m}(r, \theta, \alpha)$ of $\epsilon^{m i}$ in $\dot{\theta}$ is a polynomial in $(r, \alpha)$ of the form

$$
W_{m}(r, \theta, \alpha)=\sum_{j=0}^{m} r^{m-j} \alpha^{j} V_{m j}^{\prime}(\theta) .
$$

(3) The terms $P_{m j}(\theta)$ and $V_{m j}(\theta)$ are polynomials in $(\cos \theta, \sin \theta)$ and satisfy

$$
\begin{aligned}
& P_{m j}(\theta+\pi)=(-1)^{m+j} P_{m i}(\theta) \\
& V_{m j}(\theta+\pi)=(-1)^{n+j} V_{m j}(\theta)
\end{aligned}
$$

That is, only the terms $\cos ^{p} \theta \sin ^{q} \theta$ appear, where $p+q=m+j(\bmod 2)$.
Proof. Observe that 1 and 2 simply say that under the reverse scaling $r \rightarrow(1 / \epsilon) r, \alpha \rightarrow(1 / \epsilon) \alpha$, that $\epsilon$ is absent from the differential equation. Condition 3 says that under the substitution $\alpha \rightarrow-\alpha, \epsilon \rightarrow-\epsilon, \theta \rightarrow \theta+\pi$, the equation remains unchanged. It is thus clear that the unaveraged equation (4.9) (see also (4.2), (4.3)) satisfies these three properties. We induct on $k$, so assume the $\epsilon, \ldots, \epsilon^{2 \cdot I}$ terms have been averaged, and 1,2 , and 3 hold for all terms. Clearly, then, upon averaging the $\epsilon^{k}$ term, $u_{k}$ must have the same form as the coefficient $R_{k}$ in the sense that

$$
\begin{align*}
u_{k j}(r, \theta, \alpha) & =\sum_{j=0}^{k+1} r^{k+1-j} \alpha^{j} v_{k j}(\theta) \\
v_{k j}(\theta+\pi) & =(-1)^{k+j} v_{k j}(\theta)  \tag{4.10}\\
v_{k j}(\theta) & =\text { polynomial in }(\cos \theta, \sin \theta)
\end{align*}
$$

This is because Lemma 2.1 shows that $u_{k}$ is obtained basically by integrating $R_{k}$ with respect to $\theta$; in particular, each $v_{k j i}(\theta)$ (like $P_{k j j}(\theta)$ ) involves only terms $\cos ^{p} \theta \sin ^{q} \theta$ with $p+q=k+j(\bmod 2)$. The coordinate change

$$
\begin{equation*}
\bar{r}=\boldsymbol{r}+\epsilon^{k} u_{k}(\boldsymbol{r}, \theta, \alpha)-\boldsymbol{r}+\epsilon^{k} \sum_{j=0}^{k+1} r^{k+1-\beta_{\alpha}} v_{k j}(\theta) \tag{4.11}
\end{equation*}
$$

has the property that under the reverse scaling, $\epsilon$ is absent; this property thus carries over when the $\epsilon^{k}$ term is averaged, that is, the averaged equation satisfies 1 and 2. We also see that from (4.10), the transformation (4.11) remains unchanged under the substitution $\alpha \rightarrow-\alpha, \epsilon \rightarrow-\epsilon, 0 \rightarrow \theta+\pi$; thus 3 also holds. This then proves the lemma.

Lemma 4.3 implies that when (4.9) has been averaged, it must have the form (for $r$ bounded)

$$
\begin{aligned}
\dot{r}= & \epsilon \alpha r+\epsilon^{2}\left[r^{3} K_{2}+O(\alpha)\right]+O\left(\epsilon^{3} \alpha\right) \\
& +\epsilon^{4}\left[r^{5} K_{4}+O(\alpha)\right]+\cdots+\epsilon^{2 p}\left[r^{2 p+1} K_{2 p}+O(\alpha)\right] \\
& +O\left(\epsilon^{2 p+1}\right), \\
\dot{\theta}= & \omega_{0}+O(\epsilon),
\end{aligned}
$$

where $K_{2}, K_{4}, \ldots, K_{2 p}$ are computable constants and $k=2 p$. The generic case $K_{2} \neq 0$ was analyzed above, so here assume

$$
K_{2}=\cdots=K_{2 p-2}=0, \quad K \stackrel{\text { def }}{=} K_{2 p} \neq 0
$$

Thus we are considering

$$
\begin{align*}
& \dot{r}=\epsilon \alpha r+O\left(\epsilon^{2} \alpha\right)+\epsilon^{2 p} r^{2 p+1} K+O\left(\epsilon^{2 p+1}\right) \\
& \dot{\theta}=\omega_{0}+O(\epsilon) \tag{4.12}
\end{align*}
$$

For the same reasons as in the generic case, the choice

$$
\alpha=\left\{\begin{align*}
\epsilon^{2 p-1}, & K<0  \tag{4.13}\\
-\epsilon^{2 p-1}, & K>0
\end{align*}\right.
$$

is made, and we work near $r_{0}=|K|^{-1 / 2 p}$. This leads immediately to the natural generalization of Theorem 4.2.

Theorem 4.4. Let the averaging procedure be performed on the terms in $\dot{r}$ in (4.3) until (4.12) is obtained for some $p \geqslant 1, K \neq 0$. Then the conclusions of Theorem 4.2 hold with the following changes:
(1) Choose

$$
\begin{aligned}
\alpha & =-(\operatorname{sgn} K) \epsilon^{2 p-1} \quad(\text { scaled }) \\
r_{0} & =|K|^{-1 / 2 p}
\end{aligned}
$$

(2) The annulus $\mathscr{A}$ is positively invariant if $K<0$ and negatively invariant if $K>0$, as before.
(3) In the unscaled variables,

$$
\begin{aligned}
\alpha & =-(\operatorname{sgn} K) \epsilon^{2 p} \\
r & \sim|K|^{-1 / 2 p} \epsilon
\end{aligned}
$$

We omit any further justification of this theorem.
Of course Theorems 4.2 and 4.4 do not establish the existence of periodic solutions, but merely estimate the region where they may be found. In Section 6, existence of such solutions will be proved.

Let us close this section by briefly examining van der Pol's equation

$$
\begin{equation*}
\ddot{x}-\epsilon\left(1-x^{2}\right) \dot{x}+x=0 \tag{4.14}
\end{equation*}
$$

from the point of view of averaging; we wish to compare this with the equations obtained in the Hopf bifurcation. It is known that (4.14) has a unique periodic solution for all $\epsilon$, stable when $\epsilon>0$, and with amplitude near 2 for small $\epsilon$. Only $\epsilon$ near zero is considered here. In polar coordinates $(x, \dot{x})=(y \cos \theta$, $r \sin \theta$ ), (4.14) becomes

$$
\begin{aligned}
& \dot{r}=r\left(1-r^{2} \cos ^{2} \theta\right) \sin ^{2} \theta \\
& \dot{\theta}=--1+\epsilon\left(1-r^{2} \cos ^{2} \theta\right) \cos \theta \sin \theta
\end{aligned}
$$

Upon averaging via a transformation

$$
\bar{r}=r+\epsilon u(r, \theta)
$$

we obtain

$$
\begin{align*}
& \dot{r}=\epsilon\left((\bar{r} / 2)-\left(\bar{r}^{3} / 8\right)\right)+O\left(\epsilon^{2}\right) \\
& \dot{\theta}=-1+O(\epsilon) \tag{4.15}
\end{align*}
$$

because of the computation of the mean value

$$
(1 / 2 \pi) \int_{0}^{2 \pi} r\left(1-r^{2} \cos ^{2} \theta\right) \sin ^{2} \theta d \theta=(r / 2)-\left(r^{3} / 8\right)
$$

This suggests the existence of a periodic solution, for small $\epsilon$, near $r=2$, the unique positive root of $(r / 2)-\left(r^{3} / 8\right)=0$. Iy is instructive to compare (4.15) with the normal forms listed below, obtained for the Hopf bifurcation by Theorems 4.2 and 4.4.

$$
\begin{array}{ll}
\dot{r}=\epsilon^{2}\left(r+r^{3} K\right)+O\left(\epsilon^{3}\right) & \text { (generic case, } K<0), \\
\dot{r}=\epsilon^{2}\left(-r+r^{3} K\right)+O\left(\epsilon^{3}\right) & \text { (generic case, } K>0), \\
\dot{r}=\epsilon^{2 p}\left(r+r^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right) & (p \geqslant 2, K<0), \\
\dot{r}=\epsilon^{2 p}\left(-r+r^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right) & \\
\dot{\theta}=\omega_{0}+O(\epsilon) & \text { (all cases) } .
\end{array}
$$

## 5. Hopf Bifurcation in Higher Dimensions and More General Bifurcations

Our object in this section is first to carry over the results in Section 4 to ODE's in $R_{P}(n \geqslant 3$ ) and then study more general systems such as bifurcation to an invariant torus. This will pave the way to considering infinite-dimensional evolution systems (FDE's and PDE's) in Section 7.

Consider an ODE in coordinates $(x, y) \in R^{2} \times R^{n-2}=R^{n}$, depending on the parameter $\alpha$, such that the origin $(x, y)=(0,0)$ is a fixed point for all $\alpha$, with linearized equation

$$
\begin{align*}
& \dot{x}=A_{P}(\alpha) x+\alpha E(\alpha) y  \tag{5.1}\\
& \dot{y}=\alpha H(\alpha) x+\left[A_{O}+\alpha M(\alpha)\right] y
\end{align*}
$$

with $A_{p}(\alpha)$ the matrix in (4.1), and where $A_{Q}$ has no pure imaginary eigenvalues. When $\alpha=0$ these linear equations decouple, and it is seen we are considering the appropriate generalization of the situation in Section 4. Expand the nonlinear equation as follows:

$$
\begin{align*}
& \dot{x}=B_{0}(y, \alpha)+B_{1}(y, \alpha) x+B_{2}(y, \alpha) x^{2}+\cdots \\
& \dot{y}=\Gamma_{0}(x, \alpha)+\Gamma_{1}(x, \alpha) y+\Gamma_{2}(x, \alpha) y^{2}+\cdots \tag{5.2}
\end{align*}
$$

Then the origin is a fixed point of (5.2) with variational equation (5.1) if and only if

$$
\begin{array}{rlrl}
B_{0}(0, \alpha) & =0, & \Gamma_{0}(0, \alpha)=0, \\
\left(\partial B_{0} / \partial y\right)(0, \alpha) & =\alpha E(\alpha), & & \left.\partial \Gamma_{0} \mid \partial x\right)(0, \alpha)=\alpha H(\alpha), \\
B_{1}(0, \alpha) & =A_{P}(\alpha), & \Gamma_{1}(0, \alpha)=A_{Q}+\alpha M(\alpha),
\end{array}
$$

and this implies $B_{0}, B_{1}, \Gamma_{0}, \Gamma_{1}$ have the form

$$
\begin{aligned}
& B_{0}(y, \alpha)=\alpha E(\alpha) y+F(y, \alpha) y^{2} \\
& B_{1}(y, \alpha)=A_{P}(\alpha)+G(y, \alpha) y \\
& \Gamma_{0}(x, \alpha)=\alpha H(\alpha) x+J(x, \alpha) x^{2} \\
& \Gamma_{1}(x, \alpha)=A_{O}+\alpha M(\alpha)+N(x, \alpha) x
\end{aligned}
$$

for some functions $F, G, J, N$. The differential equation (5.2) then takes the form

$$
\begin{aligned}
\dot{x}= & \alpha E(\alpha) y+F(y, \alpha) y^{2}+A_{P}(\alpha) x+G(y, \alpha) x y \\
& +B_{2}(y, \alpha) x^{2}+B_{3}(y, \alpha) x^{3}+\cdots, \\
\dot{y}= & \alpha H(\alpha) x+J(x, \alpha) x^{2}+A_{a} y+\alpha M(\alpha) y+N(x, \alpha) x y \\
& +\Gamma_{2}(x, \alpha) y^{2}+\Gamma_{3}(x, \alpha) y^{3}+\cdots,
\end{aligned}
$$

which in polar coordinates $x=(r \cos \theta, r \sin \theta)$ becomes

$$
\begin{align*}
\dot{r}= & {\left[\alpha E_{1}(\theta, \alpha) y+F_{1}(\theta, y, \alpha) y^{2}\right]+r\left[\alpha+G_{2}(\theta, y, \alpha) y\right] } \\
& +r^{2} C_{3}(\theta, y, \alpha)+r^{3} C_{4}(\theta, y, \alpha)+\cdots, \\
\dot{\theta}= & \frac{1}{r}\left[\alpha E_{1}{ }^{*}(\theta, \alpha) y+F_{1}^{*}(\theta, y, \alpha) y^{2}\right]+\left[\omega(\alpha)+G_{2}^{*}(\theta, y, \alpha) y\right]  \tag{5.3}\\
& +r D_{3}(\theta, y, \alpha)+r^{2} D_{4}(\theta, y, \alpha)+\cdots, \\
\dot{y}= & \text { as above but with } x=(r \cos \theta, r \sin \theta) .
\end{align*}
$$

The notation is such that $E_{1}, E_{1}{ }^{*}, F_{1}$, etc., are computed from $E, F, \ldots$, just as $C_{j}$ and $D_{j}$ are computed from $B_{j-1}$ as in Section 4. Moreover, the subscript $j$ in (5.3) (such as on $E_{1}, E_{1}^{*}, F_{1}, C_{3}$, etc.) means the indicated function is homogeneous of degree $j$ in $(\cos \theta, \sin \theta)$. Scale (5.3) by

$$
r \rightarrow \epsilon r, \quad y \rightarrow \epsilon y, \quad \alpha \rightarrow \epsilon \alpha
$$

to get

$$
\begin{align*}
\dot{r}= & \epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon y, \epsilon \alpha)+\alpha E_{1}(\theta, \epsilon \alpha) y\right. \\
& \left.+F_{1}(\theta, \epsilon y, \epsilon \alpha) y^{2}+r G_{2}(\theta, \epsilon y, \epsilon \alpha) y\right]+\epsilon^{2} r^{3} C_{4}(\theta, \epsilon y, \epsilon \alpha) \div O\left(\epsilon^{3}\right), \\
\dot{\theta}= & \omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon y, \epsilon \alpha)+\frac{\alpha}{r} E_{1}^{*}(\theta, \epsilon \alpha) y\right. \\
& \left.+\frac{1}{r} F_{1}^{*}(\theta, \epsilon y, \epsilon \alpha) y^{2}+G_{2}^{*}(\theta, \epsilon y, \epsilon \alpha) y\right]+O\left(\epsilon^{2}\right),  \tag{5.4}\\
\dot{y}= & A_{0} y+\epsilon\left[\alpha H(\epsilon \alpha) x+J(\epsilon x, \epsilon \alpha) x^{2}+\alpha M(\epsilon \alpha) y\right. \\
& \left.+N(\epsilon x, \epsilon \alpha) x y+\Gamma_{2}(c x, \epsilon \alpha) y^{2}\right]+O\left(\epsilon^{2}\right),
\end{align*}
$$

and we are ready to average. The generic case ought to be determined by averaging the $\epsilon, \epsilon y$, and $\epsilon^{2}$ terms in (5.4), as we anticipate $y=O(\epsilon)$. Thus the associated transformation has the form

$$
\bar{r}=r+\epsilon u_{1}(r, \theta, \alpha, \epsilon)+\epsilon \tau v(r, \theta, \alpha, \epsilon) y+\epsilon^{2} u_{2}(r, \theta, \alpha, \epsilon)
$$

with the inverse satisfying

$$
r=\vec{r}-\epsilon u_{1}(\bar{r}, \theta, \alpha, \epsilon)+O(\epsilon|y|)+O\left(\epsilon^{2}\right) .
$$

Substituting into (5.4) yields

$$
\begin{aligned}
\dot{F}= & \epsilon\left[\left[\alpha r+r^{2} C_{3}\right]+\epsilon\left[\alpha E_{1}+r G_{2}\right] y+\epsilon^{2} r^{3} C_{4}\right. \\
& +\epsilon^{2}\left(\partial u_{1} / \partial r\right)(r, \theta, \alpha, \epsilon)\left[\alpha r+r^{2} C_{3}\right] \\
& +\epsilon\left(\hat{\partial} u_{1} / \partial \theta\right)(r, \theta, \alpha, \epsilon)\left[\omega_{0}+\epsilon \alpha \omega^{\prime}(0)+\epsilon r D_{3}\right] \\
& +\epsilon(\partial w / \partial \theta)(r, \theta, \alpha, \epsilon) \omega_{0} y+\epsilon \tau(r, \theta, \alpha, \epsilon) A_{O} y \\
& +\epsilon^{2} w(r, \theta, \alpha, \epsilon)\left[\alpha H x+J x^{2}\right] \\
& +\epsilon^{2}\left(\partial u_{2} / \partial \theta\right)(r, \theta, \alpha, \epsilon) \omega_{0}+O\left(\epsilon|y|^{2}\right)+O\left(\epsilon^{2}|y|\right)+O\left(\epsilon^{3}\right) \\
= & \epsilon\left[\alpha \bar{r}+\bar{r}^{2} C_{3}+\left(\partial u_{1} / \partial \theta\right)(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right] \\
& +\epsilon\left[\alpha E_{1}+\bar{r} G_{2}+(\partial w / \partial \theta)(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}+w(\bar{r}, \theta, \alpha, \epsilon) A_{o}\right] y \\
& +\epsilon^{2}\left[\bar{r}^{3} C_{4}+\left(\partial u_{1} / \partial r\right)(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha \bar{r}+\bar{r}^{2} C_{3}\right)\right. \\
& +\left(\partial u_{1} / \partial \theta\right)(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha \omega^{\prime}(0)+\bar{r} D_{3}\right) \\
& +z v(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha H \bar{x}+J \bar{x}^{2}\right) \\
& -u_{1}(\bar{r}, \theta, \alpha, \epsilon)\left(\alpha+2 \bar{r} C_{3}+\left(\partial^{2} u_{1} / \partial r \partial \theta\right)(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right) \\
& \left.+\left(\partial u_{2} / \partial \theta\right)(\bar{r}, \theta, \alpha, \epsilon) \omega_{0}\right]+O\left(\epsilon|y|^{2}\right)+O\left(\epsilon^{2}|y|\right)+O\left(\epsilon^{3}\right),
\end{aligned}
$$

where $C_{3}, E_{1}, \ldots$ are evaluated at $(\theta, 0, \epsilon \alpha)$ and $\bar{x}=(\bar{r} \cos \theta, \bar{r} \sin \theta)$. As before, the coefficient of $\epsilon$ (when $y=0$ ) is averaged by letting

$$
u_{1}(r, \theta, \alpha, \epsilon)=\left(-r^{2} / \omega_{0}\right) \int_{0}^{\theta} C_{3}(s, 0, \epsilon \alpha) d s
$$

this coefficient then becomes the mean value $\alpha \vec{r}$. To average the coefficient of $\epsilon y$, neglect the lower order term $\alpha E_{1}$ and let $w(r, \theta, \alpha, \epsilon)$ be the unique $2 \pi$-periodic row vector solution of

$$
\begin{equation*}
r G_{2}(\theta, 0, \epsilon \alpha)+(\partial w / \partial \theta)(r, \theta, \alpha, \epsilon) \omega_{0}+w(r, \theta, \alpha, \epsilon) A_{Q}=0 \tag{5.5}
\end{equation*}
$$

Finally, average the coefficient of $\epsilon^{2}$ (with $y=0$ ) by choosing $u_{2}$ so as to obtain the mean value

$$
\begin{gathered}
\text { mean }\left[\bar{r}^{-3} C_{4}+\left(\frac{\partial u_{1}}{\partial r}\right)\left(\alpha \bar{r}+\bar{r}^{3} C_{3}\right)+\left(\frac{\partial u_{1}}{\partial \theta}\right)\left(\alpha \omega^{\prime}(0)+\bar{r} D_{3}\right)+w\left(\alpha H \bar{x}+J \bar{x}^{2}\right)\right] \\
\left.==\bar{r}^{3} \text { mean }\left[C_{4}-\frac{1}{\omega_{0}} C_{3} D_{3}\right]+\operatorname{mean}[w] \bar{x}^{2}\right]+O(\alpha) .
\end{gathered}
$$

Observe this is not the same quantity obtained in Section 4, due to the additional term mean $\left(w / \bar{x}^{2}\right)$. Thus even in the generic case, the terms involving $y$ cannot be ignored. It is clear from (5.5) that

$$
w(\bar{r}, \theta, \alpha, \epsilon)=\bar{r} \bar{w} *(\theta, \alpha, \epsilon)
$$

for some $w^{*}$, and so, since $\bar{x}=(\bar{r} \cos \theta, \vec{r} \sin \theta)$, we have

$$
\operatorname{mean}\left(w J \bar{x}^{2}\right)=\dot{r}^{3} K^{* *}
$$

for some constant $K^{* *}$. This gives the analog of Theorem 4.1.

## Theorem 5.1. Consider the differential equation

$$
\begin{align*}
\dot{r}= & \epsilon\left[\alpha r+r^{2} C_{3}(\theta, \epsilon y, \epsilon \alpha)+\alpha E_{1}(\theta, \epsilon \alpha) y\right. \\
& \left.+F_{1}(\theta, \epsilon y, \epsilon \alpha) y^{2}+r G_{2}(\theta, \epsilon y, \epsilon \alpha) y\right]+\epsilon^{2} r^{3} C_{4}(\theta, \epsilon y, \epsilon \alpha)+O\left(\epsilon^{3}\right), \\
\dot{\theta}= & \omega_{0}+\epsilon\left[\alpha \omega^{\prime}(0)+r D_{3}(\theta, \epsilon y, \epsilon \alpha)+\frac{\alpha}{r} E_{1}^{*}(\theta, \epsilon \alpha) y\right.  \tag{5.6}\\
& \left.+\frac{1}{r} F_{1}^{*}(\theta, \epsilon y, \epsilon \alpha) y^{2}+G_{2}^{*}(\theta, \epsilon y, \epsilon \alpha) y\right]+O\left(\epsilon^{2}\right), \\
\dot{y}= & A_{O} y+\epsilon\left[\alpha H(\epsilon \alpha) x+J(\epsilon x, \epsilon \alpha) x^{2}+\alpha M(\epsilon \alpha) y\right. \\
& \left.+N(\epsilon x, \epsilon \alpha) x y+\Gamma_{2}(\epsilon x, \epsilon \alpha) y^{2}\right]+O\left(\epsilon^{2}\right)
\end{align*}
$$

arising from the Hopf bifurcation problem in $R^{n}$ described above in Section 5, and the scaling $r \rightarrow \epsilon T, y \rightarrow \epsilon y, \alpha \rightarrow \epsilon \alpha$. Then there exists a coordinate change

$$
\bar{r}=r+\epsilon u_{1}(r, \theta, \alpha, \epsilon)+\epsilon w(r, \theta, \alpha, \epsilon) y+\epsilon^{2} u_{2}(r, \theta, \alpha, \epsilon)
$$

transforming (5.6) into the averaged system of the form

$$
\begin{aligned}
& \dot{r}=\epsilon \alpha \bar{r}+\epsilon^{2} \vec{r}^{3} K+O\left(\epsilon|y|^{2}\right)+O\left(\epsilon^{2}|y|\right)+O\left(\epsilon^{3}\right) \\
& \dot{\theta}=\omega_{0}+O(\epsilon) \\
& \dot{y}=A_{Q} y+O(\epsilon)
\end{aligned}
$$

where $K$ is the constant

$$
\begin{aligned}
K & =K^{*}+K^{* *} \\
K^{*} & =\frac{1}{2 \pi} \int_{0}^{2 \pi} C_{4}(\theta, 0,0)-\frac{1}{\omega_{0}} C_{3}(\theta, 0,0) D_{3}(\theta, 0,0) d \theta \\
K^{* *} & =\frac{1}{2 \pi} \int_{0}^{2 \pi} w^{*}(\theta) J(0,0)(\cos \theta, \sin \theta)^{2} d \theta
\end{aligned}
$$

where $w^{*}(\theta)$ is the unique $2 \pi-$-periodic solution of

$$
\begin{equation*}
G_{2}(\theta, 0,0)+w^{*}(\theta) \omega_{0}+w^{*}(\theta) A_{Q}=0 \tag{5.7}
\end{equation*}
$$

We recall that for each $(x, \alpha), J(x, \alpha)$ is a bilinear form in the $x$-space $R^{2}$, taking values in the $y$-space; in the theorem $J(0,0)$ acts on the point $(\cos \theta, \sin \theta) \in R^{2}$. Also note that the easiest way of solving (5.7) and computing $K^{* *}$ may be to expand $G_{2}$ and $w^{*}$ in Fourier series; see for example, Wright's equation in Section 9. Observe the following interesting fact: the property $K \neq 0$ depends only on the differential equation at $\alpha=0$ and not on the particular parameterization passing through this equation, since the formulas for $K^{*}$ and $K^{* *}$ do not involve derivatives of terms with respect to $\alpha$. We are assuming $\operatorname{Re} \lambda(\alpha)=\alpha$ for the eigenvalues $\lambda(\alpha)$ of $A_{P}(\alpha)$, but if more generally, we have $\operatorname{Re} \lambda^{\prime}(0)=$ $\nu \neq 0$, then a generic bifurcation will still occur if $K \neq 0$. In this case the averaged equation for $\dot{r}$ would be

$$
\begin{aligned}
\dot{r} & =\operatorname{Re}(\lambda(\epsilon \alpha)) r+\epsilon^{2} r^{3} K+O\left(\epsilon^{3}\right) \\
& =\epsilon \alpha \nu r+\epsilon^{2} \dot{r}^{3} K+O\left(\epsilon^{3}\right) \\
& =\epsilon^{2}\left( \pm \nu r+r^{3} K\right)+O\left(\epsilon^{3}\right), \quad \pm=-\operatorname{sgn}(\nu K)
\end{aligned}
$$

where $\alpha=-\operatorname{sgn}(\nu K) \epsilon$. Thus, if $K \neq 0$ for the equation at $\alpha=0$, then a generic bifurcation should occur if $\operatorname{Re} \lambda^{\prime}(0) \neq 0$. We recall that the assumption here that $\operatorname{Re} \lambda(\alpha)=\alpha$ is simply for convenience. As long $\operatorname{Re} \lambda^{\prime}(0) \neq 0$, all of the averaging techniques described here apply.

Higher order terms of (5.6) can be averaged provided $A_{Q}$ is $l$-simple for appropriate $l$. In the following lemma, the terms $\epsilon, \epsilon^{2}, \ldots, \epsilon^{2 p}$ are averaged. The proof involves essentially the same arguments as in Section 4, especially Lemma 4.3, so it is omitted.

Lemma 5.2. Let $A_{Q}$ be l-simple for all $1 \leqslant l \leqslant 2 p-1$. Then the coefficients of $\epsilon^{i} y^{j}, 1 \leqslant i \leqslant 2 p, 0 \leqslant j \leqslant 2 p-i$ in the expression for $\dot{r}$ in (5.6) may be averaged with a transformation of the form

$$
\begin{aligned}
& \bar{r}=r+\sum_{i=1}^{2 p} \epsilon^{i} u_{i}(r, \theta, y, \alpha, \epsilon) \\
& u_{i}=\text { polynomial in } y \text { of degree }(2 p-i)
\end{aligned}
$$

The resulting averaged equations have the form

$$
\begin{aligned}
\dot{r}= & \epsilon \alpha r+O\left(\epsilon^{2} \alpha\right)+\epsilon^{2} r^{3} K_{2}+\epsilon^{4} r^{5} K_{4}+\cdots+\epsilon^{2 p} r^{2 p+1} K_{2 p} \\
& +\sum_{i=1}^{2 p} O\left(\epsilon^{i}|y|^{2 p-i+1}\right)+O\left(\epsilon^{2 p+1}\right) \\
\dot{\theta}= & \omega_{0}+O(\epsilon) \\
\dot{y}= & A_{Q} y+O(\epsilon)
\end{aligned}
$$

where $K_{2}, K_{4}, \ldots, K_{2 p}$ are computable constants.
In the nongeneric case $K_{2}=0$; if $K_{2}=\cdots=K_{2 p-2}=0, K \stackrel{\text { def }}{=} K_{2 p} \neq 0$ then we are considering (dropping the bars)

$$
\begin{align*}
& \dot{r}=\epsilon \alpha r+O\left(\epsilon^{2} \alpha\right)+\epsilon^{2 p} r^{2 p+1} K+\sum_{i=1}^{2 p} O\left(\epsilon^{i}|y|^{2 p-i+1}\right)+O\left(\epsilon^{2 p+1}\right) \\
& \dot{\theta}=\omega_{0}+O(\epsilon)  \tag{5.8}\\
& \dot{y}=A_{Q} y+\epsilon Y(r, \theta, y, \alpha, \epsilon)
\end{align*}
$$

for some function $Y$.
If $(r(t), \theta(t), y(t))$ is any solution bounded on $(-\infty, \infty)$ (e.g., a periodic solution) then the equation for $y$ can be written in integrated form

$$
\begin{equation*}
y(t)=\epsilon \int_{-\infty}^{t} e^{A_{Q}(t-s)} Y(r(s), \theta(s), y(s), \alpha, \epsilon) d s \tag{5.9}
\end{equation*}
$$

if $A_{Q}$ is a stable matrix. (If $A_{Q}$ is hyperbolic then two integrals are needed, one from $-\infty$ to $t$, the other from $t$ to $+\infty$.) Temporarily choose $\epsilon$ as the supremum
of $|r(t)|+|y(t)|$ (unscaled), so that upon scaling, $|r(t)|+|y(t)|$ has 1 as its supremum. But from (5.9), we have for appropriate constants $q>0, \Omega>0$

$$
|y(t)| \leqslant|\epsilon| \int_{-\infty}^{t} e^{-\alpha(t-s)} d s \sup _{|r|+|| | \leqslant 1}|Y(r, \theta, y, \alpha, \epsilon)| \leqslant \Omega|\in|,
$$

so that $|r(t)| \geqslant 1-2 \Omega|\epsilon|$ for some $t$. Thus as long as we scale so that $r$ remains in a bounded region, as $\epsilon \rightarrow 0$, we have (uniformly) $y=O(\epsilon)$. In particular, in (5.8), $\sum_{i=1}^{2 p} O\left(\epsilon^{i}|y|^{2 p-i+1}\right)=O\left(\epsilon^{2 p+1}\right)$, and this justifies the choice of $\alpha=-(\operatorname{sgn} K) \epsilon^{2 p-1}$. Finally, the annulus $\mathscr{A}^{*}$ surrounding a periodic solution is given by

$$
\mathscr{A}^{*} \begin{cases}(1-\gamma) r_{0} \leqslant r \leqslant(1+\gamma) r_{0}, & \gamma \rightarrow 0 \text { as } \epsilon \rightarrow 0,  \tag{5.10}\\ |y| \leqslant \Omega|\epsilon|, & \Omega=\text { constant } .\end{cases}
$$

Observe that $\mathscr{A}^{*}$ may not be invariant as before; if $K<0$ but $A_{Q}$ is not stable, for example, then solutions will enter $\mathscr{A}^{*}$ along the boundary $r=(1 \pm \gamma) r_{0}$, but may leave along $|y|=\Omega|\in|$. We obtain then the following theorem.

Theorem 5.3. Let (5.6) be averaged as in Lemma 5.2, and assume the form (5.8) for some $K \neq 0$. Then all periodic solutions of the unscaled equation (5.3) bifurcating from $r=0, y=0, \alpha=0$ may be obtained by letting $\alpha=-(\operatorname{sgn} K) \epsilon^{2 p-1}$ in (5.8) to give

$$
\begin{align*}
& \dot{r}=\epsilon^{2 p}\left( \pm r+r^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right), \quad \pm=- \text { sgn } K, \\
& \dot{\theta}=\omega_{0}+O(\epsilon,  \tag{5.11}\\
& \dot{y}=A_{0} y+O(\epsilon),
\end{align*}
$$

and by considering $r$ near $r_{0}=|K|^{-1 / 2 p}$ and y near zero. The annulus $\mathscr{A} *$ in (5.10) is positively invariant if $K<0$ and $A_{Q}$ is stable.

Let us now briefly describe the phenomenon of bifurcation of an invariant torus from a periodic orbit. Suppose in $R^{n}$ an autonomous differential equation $\dot{x}=f(x, \alpha)$ has for $\alpha=0$ a nonconstant periodic solution $p(t)$, which is nondegenerate, that is, the characteristic multiplier $\mu=1$ is simple. It is well known that for $|\alpha|$ small there is a unique periodic solution $p(t, \alpha)$, smooth in $(t, \alpha)$, with $p(t, 0)=p(t)$. We may assume also by rescaling the time that $p(t, \alpha)$ has period $2 \pi$. In an appropriate coordinate system around the periodic orbit, the autonomous equation may be rewritten as a nonautonomous equation

$$
\begin{equation*}
\dot{x}_{1}=f_{1}\left(t, x_{1}, \alpha\right), \tag{5.12}
\end{equation*}
$$

where $x_{1} \in R^{n+1}, f_{1}$ is $2 \pi$-periodic in $t$, and $f_{1}(t, 0, \alpha) \equiv 0$. The solution $x_{1}=0$
corresponds in this coordinate system to the periodic orbit $p(t, \alpha)$. We may write (5.12) as

$$
\begin{align*}
\dot{x}_{1} & =A(\alpha) x_{1}+F\left(x_{1}, t, \alpha\right), \\
\left|F\left(x_{1}, t, \alpha\right)\right| & =O\left(\left|x_{1}\right|^{2}\right) . \tag{5.13}
\end{align*}
$$

The linearized equation $\dot{x}_{1}=A(\alpha) x_{1}$ here has been made autonomous by a linear transformation, using Floquet theory. Assume the same conditions as before on the eigenvalues of $A(\alpha)$. In particular, the eigenvalues $\pm i \omega_{0}$ at $\alpha=0$ correspond to characteristic multipliers $\mu=e^{ \pm 2 \pi i \omega_{0}}$ of the original periodic orbit. The averaging of (5.13) then proceeds exactly as for the Hopf bifurcation, but with the appropriate modifications described in Section 2. In particular, in order to average the terms $\epsilon r^{2} C_{3}(\theta, t)$ and $\epsilon^{2} r^{3} C_{4}(0, t)$ (as in (5.6)) it is necessary to assume $m \omega_{0}+n \neq 0$ for ( $\left.m, n\right) \neq(0,0)$, and $|m| \leqslant 4$. That is, one must assume the critical characteristic multipliers $\mu=e^{ \pm 2 \pi i \omega_{0}}$ satisfy

$$
\mu^{N} \neq 1, \quad N=1,2,3,4 .
$$

Under these conditions, averaging and scaling gives rise to a normal form as in (5.11), where $p=1, K$ is constant, and the higher order terms are periodic in $t$. If $K \neq 0$, one may expect a two-dimensional invariant manifold near the torus

$$
r=|K|^{-1 / 2}, \quad y=0, \theta, t=\text { arbitrary },
$$

where $(\theta, t)$ are the coorinates on the torus. This is indeed the case and follows from standard results on invariant manifolds.

## 6. Existence of the Bifircating Solutions

We have yet to actually prove that the system (5.11) obtained by scaling and averaging (5.2), (5.3) has a periodic solution bifurcating from the fixed point $r=0, y=0$; this will now be shown. Briefly, the system (5.11) possesses an invariant manifold $\Sigma$, the center manifold, given by $y=y(r, \theta, \epsilon)$. It is defined near $r=0, \epsilon=0$, and passes through the origin for each $\epsilon$, so that $y(0,0, \epsilon)=0$. All orbits which stay near the origin for all $t \in(-\infty, \infty)$ lie on the center manifold; in particular, all periodic solutions lie on $\Sigma$. By substituting $y=$ $y(r, \theta, \epsilon)$ into the differential equation (5.11) the search for periodic solutions has been reduced to a two-dimensional problem, as the equations now involve only $(r, \theta)$.

Strictly speaking, it is not necessary to use the center manifold, as it is not difficult to prove the existence of periodic solutions of (5.11) directly. Its advantage lies in the fact that many infinite-dimensional systems (such as functional and partial differential equations) have center manifolds. Obtaining
periodic solutions of such systems directly may be very difficult; by looking only on the center manifold, however, the problem is reduced to a two-dimensional ODE.

To obtain the center manifold, augment the original system (5.2) by considering the parameter $\alpha$ as a state variable satisfying $\dot{\alpha}=0$. The linearized equation about the origin $(x, y, \alpha)=(0,0,0)$ is then

$$
\begin{aligned}
& \dot{x}=A_{p}(0) x, \quad A_{p}(0)=\left(\begin{array}{cc}
0 & -\omega_{0} \\
\omega_{0} & 0
\end{array}\right), \\
& \dot{y}=A_{O} y \\
& \dot{\alpha}=0 .
\end{aligned}
$$

Since no eigenvalues of $A_{Q}$ are purely imaginary, the center manifold theorem guarantees the existence of a smooth invariant three-dimensional manifold $\Sigma$, passing through $(x, y, \alpha)=(0,0,0)$ and tangent to the ( $x, \alpha$ ) space. In polar coordinates $x=(r \cos \theta, r \sin \theta)$ then $\Sigma$ has the form

$$
\begin{aligned}
\Sigma: y & =y^{*}(r, \theta, \alpha) \\
0 & =y^{*}(0, \theta, 0)=\frac{\partial y^{*}}{\partial r}(0, \theta, 0)=\frac{\partial y^{*}}{\partial \alpha}(0, \theta, 0),
\end{aligned}
$$

for some function $y^{*}$. Since the fixed point $r=0, y=0$ must lie on $\Sigma$ for all $\alpha$ we also have $y^{*}(0, \theta, \infty)=0$, so that

$$
\begin{aligned}
& y^{*}(r, \theta, \alpha)=r z^{*}(r, \theta, \alpha) \\
& z^{*}(0, \theta, 0)=0
\end{aligned}
$$

for some smooth $z^{*}$. After scaling $r \rightarrow \epsilon r, y \rightarrow \epsilon y, \alpha \rightarrow \epsilon \alpha$ and averaging by $r \rightarrow r+O(\epsilon)$, and setting $\alpha= \pm \epsilon^{2 p-1}$, the equation for $\Sigma$ takes the form

$$
y=r z^{*}\left(\epsilon r+O\left(\epsilon^{2}\right), \theta, \pm \epsilon^{2 p}\right)+O(\epsilon) \stackrel{\text { dof }}{=} y(r, \theta, \epsilon)
$$

Note that $y(r, \theta, \epsilon)=O(\epsilon)$ uniformly, from (6.1). Thus substitution of $y:=$ $y(r, \theta, \epsilon)$ into (5.11) yields the differential equation on $\Sigma$

$$
\begin{align*}
& \dot{r}=\epsilon^{2 p}\left( \pm r+r^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right), \quad \pm=\operatorname{sgn} K, \\
& \dot{\theta}=\omega_{0}+O(\epsilon) . \tag{6.2}
\end{align*}
$$

The above reduction of the bifurcation problem to $\Sigma$ carries over for more general (infinite-dimensional) systems whenever the following hold:
(1) there exists a smooth invariant manifold $\Sigma$ given by $y=y^{*}(r, \theta, \alpha)$ through the origin $(r, y, \alpha)=(0,0,0)$ and tangent to the $(r, \theta, \alpha)$-space;
(2) each point of $\Sigma$ lies on a trajectory of the differential equation, that is, the differential equation induces a smooth flow on $\Sigma$; and
(3) all orbits lying near the origin for all time $t \in(-\infty, \infty)$ lie on $\Sigma$.

Quite generally, if the equations generate a semigroup $T(t, x, \alpha)$ such that the $\operatorname{map}(x, \alpha) \rightarrow T(t, x, \alpha)$ is smooth for each fixed $t$, then a center manifold with these properties exists. This is the case for retarded functional differential equations and for many classes of PDE's such as certain nonlinear parabolic equations, and the Navier-Stokes equation.

It is now very easy to obtain periodic solutions for the two-dimensional system (6.2).

Lemma 6.1. For restricted to a sufficiently large bounded region, the system (6.2) has a unique periodic solution $r(t, \epsilon), \theta(t, \epsilon)$ for small $\mid \epsilon!\neq 0 . A s \epsilon \rightarrow 0$,

$$
\begin{aligned}
r(t, \epsilon) \rightarrow r_{0} & =|K|^{-1 / 2 p} \text { uniformly }, \\
\tau(\epsilon) & \stackrel{\text { der }}{=} \text { period of the solution } \rightarrow 2 \pi / \omega_{0} .
\end{aligned}
$$

The solution, restricted to $\Sigma$, is stable when $K<0$ and unstable when $K>0$.
Theorem 6.2. Let the Hopf bifurcation problem (5.6) be averaged, and the substitution $\alpha=-(\operatorname{sgn} K) \epsilon^{2, p-1}$ made, as described in Theorem 5.3, so that

$$
\begin{aligned}
& \dot{r}=\epsilon^{2 p}\left( \pm r+r^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right), \quad \pm=-\operatorname{sgn} K, \\
& \dot{\theta}=\omega_{0}+O(\epsilon), \\
& \dot{y}=A_{Q} y+O(\epsilon),
\end{aligned}
$$

is obtained. Then in the original (unaveraged and unscaled) equation, there is a unique periodic solution bifurcating from the origin, either for $\alpha>0$ (when $K<0$ ) or $\alpha<0$ (when $K>0$ ). More precisely, in the original coordinates $(x, y, \alpha)$ with $x=(r \cos \theta, r \sin \theta)$ the solution has the form

$$
\begin{aligned}
r(t, \epsilon) & =\epsilon r_{0}+O\left(\epsilon^{2}\right), \quad r_{0}=|K|^{-1 / 2 p} \\
\theta(t, \epsilon) & =\omega_{0} t+O(\epsilon) \\
y(t, \epsilon) & =O\left(\epsilon^{2}\right) \\
\tau(\epsilon) & =\text { period of the solution }=\left(2 \pi / \omega_{0}\right)+O(\epsilon),
\end{aligned}
$$

where $\epsilon$ is related to the bifurcation parameter $\alpha$ by

$$
\alpha=-(\operatorname{sgn} K) \epsilon^{2 p}
$$

The solutions obtained for $\epsilon$ and $-\epsilon$ are identical and only differ by a time translation. The solution is stable if and only if all eigenvalues of $A_{Q}$ lie in the left half-plane and $K<0$.

In case the eigenvalues $\lambda(\alpha), \overline{\lambda(\alpha)}$ of the $(r, \theta)$-subspace of the linearized equation satisfy merely $\operatorname{Re} \lambda^{\prime}(0) \neq 0$ rather than $\operatorname{Re} \lambda(\alpha) \equiv \alpha$, then all of the above carries over with the obvious modifications.

Only Lemma 6.1 will be proved, as the theorem is a simple application of this lemma to the scaled and averaged system considered in Theorem 5.3.

Proof of Lemma 6.1. Consider the solution $r(t), \theta(t)$ of (6.2) for $\epsilon$ fixed, with initial conditions

$$
r(0)=\rho, \quad \theta(0)=0,
$$

and define $R(\rho, \epsilon)=r(\tau)$ where $\theta(\tau)=2 \tau$. Periodic solutions then are given by solving

$$
\begin{equation*}
R(\rho, \epsilon)=\rho, \tag{6.3}
\end{equation*}
$$

and $\tau=\tau(\rho, \epsilon)$ is the period. From the form of (6.2) it is clear that

$$
\begin{aligned}
\tau(\rho, \epsilon) & =\left(2 \pi / \omega_{0}\right)+O(\epsilon) \\
R(\rho, \epsilon) & =\rho+\epsilon^{2 \rho}\left(2 \pi / \omega_{0}\right)\left( \pm \rho+\rho^{2 p+1} K\right)+O\left(\epsilon^{2 p+1}\right)
\end{aligned}
$$

Thus (6.3) reduces to solving

$$
\begin{aligned}
S(\rho, \epsilon) & \stackrel{\text { def }}{=} \epsilon^{-2 p}(R(\rho, \epsilon)-\rho) \\
& =\left(2 \pi / \omega_{0}\right)\left( \pm \rho+\rho^{2 p+1} K\right)+O(\epsilon) \\
& =0 .
\end{aligned}
$$

It is immediate that $S\left(r_{0}, 0\right)=0,\left(\partial S / \partial \rho_{0}\right)\left(r_{0}, 0\right) \neq 0$. Hence by the implicit function theorem there is a unique zero $\rho=\rho(\epsilon)$ of $S$, with $\rho(0)=r_{0}$. This establishes the existence of the periodic solution. The assertions about stability are a consequence of the existence of the annulus $\mathscr{A}$ in Theorem 4.3; alternatively one may observe that

$$
\begin{aligned}
& K<0 \Rightarrow\left(\partial S / \partial r_{0}\right)(\rho(\epsilon), \epsilon)<0 \Rightarrow \text { stability }(\text { in } \Sigma) \\
& K>0 \Rightarrow\left(\partial S / \partial r_{0}\right)(\rho(\epsilon), \epsilon)>0 \Rightarrow \text { instability }
\end{aligned}
$$

## 7. Infinite-Dimensional Systems

Here we show how the averaging procedure and its application to bifurcation carries over to certain classes of infinite-dimensional evolution equations, such as functional and certain partial differential equations. Assume the equation can be written abstractly as

$$
\begin{gather*}
\dot{z}=f(z, \alpha)=A(\alpha) z+F(z, \alpha),  \tag{7.1}\\
F(z, \alpha)=O\left(|z|^{2}\right),
\end{gather*}
$$

where $z$ and $\dot{z}$ lie in (generally different) Banach spaces. Specifically, assume

$$
f: X_{1} \times\left(-\alpha_{0}, \alpha_{0}\right) \rightarrow X_{2}
$$

is sufficiently smooth, where $X_{1}$ and $X_{2}$ are Banach spaces with $X_{1}$ continuously and densely contained in $X_{2}$. For example, if (7.1) represents a parabolic equation of the form

$$
\begin{aligned}
\partial z / \partial t & =g\left(x, z, \partial z / \partial x, \partial^{2} z / \partial x^{2}, \alpha\right) \\
x \in \Omega & =\text { smooth, open, bounded set in } R^{n}, \\
z & -0 \quad \text { for } \quad x \in \partial \Omega,
\end{aligned}
$$

then choices for $\left(X_{1}, X_{2}\right)$ would possibly be $\left(H_{0}{ }^{2}(\Omega), L^{2}(\Omega)\right)$ or $\left(C_{0}^{2+\alpha}(\Omega), C^{\alpha}(\Omega)\right)$.
We also assume, as before,

$$
\begin{aligned}
f(0, \alpha) & =0 \\
(\partial f / \partial z)(0, \alpha) & =A(\alpha) \\
& =\text { bounded linear operator from } X_{1} \text { to } X_{2}
\end{aligned}
$$

so that the linearized equation about the origin $z=0$, when $\alpha=0$, is

$$
\dot{z}=A(0) z .
$$

By considering $X_{1}$ as a subset of $X_{2}$, then $A(0)$ (or more generally $A(\alpha)$ ) may be regarded as an unbounded closed operator from $X_{2}$ into itself, with domain $X_{1} \subseteq X_{2}$. Typically, the space $X_{1}$ (when not considered a subset of $X_{2}$ ) will be eldowed with the graph norm of $A(0)$. We have the spectral decomposition

$$
X_{2}=P \oplus Q
$$

where $P$ is the two-dimensional eigenspace of $A(0)$ corresponding to simple eigenvalues $\pm i \omega_{0} \neq 0$ and the spectrum of $A_{Q}(0)(=A(0)$ restricted to $Q)$ is assumed to lie a positive distance $\delta$ from the imaginary axis. In fact, we assume the decomposition

$$
\begin{aligned}
X_{2} & =P(\alpha) \oplus Q(\alpha), \\
\operatorname{dim} P(\alpha) & =2
\end{aligned}
$$

holds for all $\alpha$ near 0 , the positive distance $\delta$ holds uniformly, and the eigenvalues of $A(\alpha)$ restricted to $P(\alpha)$ are $\alpha \pm i \omega(\alpha)$. Since the eigenspace $P$ lies in $X_{1} \subseteq X_{2}$ the decomposition restricts to $X_{1}$

$$
X_{1}=P \oplus\left(Q \cap X_{1}\right) \stackrel{\text { def }}{=} P \oplus Q_{1}
$$

It is now clear that (7.1) can be decomposed by writing

$$
\begin{aligned}
& z=x+y \in P \oplus Q_{1}=X_{1}, \\
& \dot{z}=x+y \in P \oplus Q=X_{2}, \\
& x=(r \sin \theta, r \cos \theta) .
\end{aligned}
$$

Assuming all of the above conditions on the spectrum of $A(\alpha)$, and the resulting decomposition, the only other assumption needed is the existence of a center manifold

$$
\begin{aligned}
\Sigma: y & =y^{*}(r, \theta, \alpha) \in X_{1} \\
0 & =y^{*}(0, \theta, \alpha)=\left(\partial y^{*} / \partial r\right)(0, \theta, 0)
\end{aligned}
$$

through the origin $r=0$, and tangent to the $(r, \theta, \alpha)$ space. It is important that $y^{*}$ be a smooth map taking values in $X_{1}$, since after averaging, we must substitute $y=y^{*}(r, \theta, \alpha)$ into the right-hand side of (7.1). As described in Section 6, a smooth flow is induced on $\Sigma$, and all periodic orbits bifurcating from the origin lie on $\Sigma$. The basic assumption on (7.1) necessary for $\Sigma$ to exist is that the nonlinear semigroup $T(t, z, \alpha)$ in $X_{1}$ generated by (7.1) be smooth in $(z, \infty)$ for each fixed $t$.

With the above setup, all there remains to do is to rigorously justify the formal averaging procedure applied to (7.1). Recall from Theorem 5.1 the coordinate change

$$
\begin{equation*}
\bar{r}=r+\epsilon u_{1}(r, \theta, \alpha, \epsilon)+\epsilon w(r, \theta, \alpha, \epsilon) y+\epsilon^{2} u_{2}(r, \theta, \alpha, \epsilon) \tag{7.2}
\end{equation*}
$$

used to compute the first order constant $K$. (For simplicity we consider only this case, as averaging of higher order terms is similar.) Here $u_{1}$ and $u_{2}$ are scalar valued, while $w$ takes values in the dual $Q_{1}{ }^{*}$ of the $y$-space $Q_{1}$. In fact $w(r, \theta, \alpha, \epsilon)=$ $r w^{*}(\theta)$ whereas in (5.7), $w^{*}$ is the unique $2 \pi$-periodic solution of

$$
\begin{equation*}
G_{2}(\theta, 0,0)+w^{*^{\prime}}(\theta) \omega_{0}+w^{*}(\theta) A_{Q}=0 \tag{7.3}
\end{equation*}
$$

Let us consider more carefully the meaning of Eq. (7.3) and trasformation (7.2) in the infinite-dimensional space. Now $G_{2}$ arises as a coefficient of $y$ in the differential equation involving $\dot{r}$ (after decomposing); hence $G_{2}(\theta, 0,0)$ for each $\theta$ is a linear functional acting on $y \in Q_{1}$. In particular, writing $G_{2}$ as a Fourier series yields

$$
\begin{gathered}
G_{2}(\theta, 0,0)=\sum_{n=-\infty}^{\infty} g_{n^{2}} e^{i n \theta} \\
g_{n} \in Q_{1}{ }^{*}, \quad \sum_{n=-\infty}^{\infty}\left|g_{n}\right|_{Q_{1}^{*}}^{2}<\infty
\end{gathered}
$$

By expanding $w^{*}(\theta)$ as a Fourier series

$$
\begin{equation*}
w^{*}(\theta)=\sum_{n \because:-\infty}^{\infty} w_{n} e^{i n \theta} \tag{7.4}
\end{equation*}
$$

inserting this into Eq. (7.3) and equating coefficients, we arrive at

$$
\begin{equation*}
w_{n}=-g_{n}\left(A_{Q}+i n \omega_{0}\right)^{-1} \tag{7.5}
\end{equation*}
$$

Since $A_{Q}$ is a bounded operator from $Q_{1}$ into $Q$, then $\left(A_{Q}+i n \omega_{0}\right)^{-1}$ is bounded from $Q$ into $Q_{1}$. This implies in particular

$$
w_{n} \in Q^{*} \quad\left|w_{n}\right|_{Q^{*}} \leqslant \text { (const.) } n^{-1}\left|g_{n}\right|_{Q_{1}^{*}}
$$

Thus by defining $w^{*}$ by (7.4) and (7.5), both $w^{*}$ and its derivative $w^{* \prime}$ are square integrable functions taking values in $Q^{*}$. This is stronger than saying they take values in $Q_{1}{ }^{*}$, and is due to the presence of the smoothing operator $\left(A_{Q}+i \omega_{0}\right)^{-1}$ in (7.5). This last observation is important since when Eq. (7.1) is rewritten in terms of the new averaged coordinates $(\vec{r}, \theta, y)$, it is seen that the functional $w(r, \theta, \alpha, \epsilon)$ acts on $\dot{y} \in Q$. Indeed, this is what happens when (7.2) is differentiated with respect to time.

One thus concludes that the form of the equation, that is, $z \in X_{1}$, and $\dot{z} \in X_{2}$, is preserved under any sequence of averaging transformations (7.2), and (of course) scaling; the equation $y=y(r, \theta, \epsilon)$ describing the center manifold in scaled, averaged coordinates may be substituted into the infinite-dimensional system, to reduce the problem to the two-dimensional case, as before.

## 8. Functional Differential Equations

In order to average retarded functional differential equations, care must be taken as to how the equation is interpreted as an (abstract) ordinary differential equation in a Banach space, as in (7.1). Consider the RFDE

$$
\begin{equation*}
\dot{z}(t)=f\left(z_{t}, \alpha\right) \tag{8.1}
\end{equation*}
$$

where the notation of Hale is followed. In particular, assume $z \in R^{n}$, and $z_{t}$ is the function defined by

$$
z_{t}(\theta)=z(t+\theta), \quad-r \leqslant \theta \leqslant 0
$$

thus

$$
z_{t} \in C \stackrel{\text { def }}{=} C\left([-r, 0], R^{n}\right)
$$

where $r$ is fixed. The phase space of (8.1) is thus the Banach space $C$, and $f=f(\phi, \alpha)$ satisfies

$$
\begin{gathered}
f: C \times\left(-\alpha_{0}, \alpha_{0}\right) \rightarrow R^{n} \\
f(0, \alpha)=0 .
\end{gathered}
$$

Equation (8.1) can be solved forward in time by specifying an initial condition $z_{0}=\phi \in C$ at time $t=0$.

Some remarks on notation are in order here. As there is some overlap in the notation for FDE's and that of the previous seven sections, it will be necessary to adopt several new conventions for the integral averaging. For example, we use $\theta$ from now on to denote the argument of $z_{t}$, so that $\theta \in[-r, 0]$, and not the angle variable as before. The symbol $\zeta$ henceforth denotes the angle in polar coordinates $(r, \zeta)$.

Write (8.1) as

$$
\begin{align*}
\dot{z}(t) & =L(\alpha) z_{t}+F\left(z_{t}, \alpha\right), \\
L(\alpha) & =(\partial f \mid \alpha \phi)(0, \alpha),  \tag{8.2}\\
|F(\phi, \alpha)| & =O\left(|\phi|^{2}\right),
\end{align*}
$$

so that the linearized equation at the origin is, for each $\alpha$,

$$
\dot{z}(t)=L(\alpha) z_{i} .
$$

Here $L(\alpha)$ is a linear functional on $C$ but takes values in $R^{n}$; it thus has the Stieltjes integral representation

$$
L(\alpha) \phi=\int_{-r}^{0} d_{\theta} \eta(\alpha, \theta) \phi(\theta)
$$

where the $n \times n$ matrix $\eta(\alpha, \theta)$ is of bounded variation in $\theta \in[-r, 0]$, and smooth in $\alpha$ when considered as a $B V[-r, 0]$ valued function.

Equations (8.1), (8.2) certainly do not fit the framework of Section 7, in particular (7.1). The left-hand side $\dot{z}(t)$ lies in $R^{n}$ and this cannot be considered a phase space of (8.2). The clue to writing (8.2) as an abstract ODE comes from the variation of constants formula for retarded equations.

First scale $z \rightarrow \epsilon z, \alpha \rightarrow \epsilon \alpha$ so that (8.2) takes the form (with a different $F$ )

$$
\begin{align*}
\dot{z}(t) & =L z_{t}+\epsilon F\left(z_{t}, \alpha, \epsilon\right), \\
L & =L(0) . \tag{8.3}
\end{align*}
$$

The linear equation $\dot{z}(t)=L z_{t}$ at $\epsilon=0$ generates a strongly continuous semigroup $T(t)$ of bounded linear operators on $C$, with infinitesimal generator $A$ given by

$$
\begin{aligned}
A \phi & =(d / d \theta) \phi, \\
\phi \in C^{1} \cap\{\phi \mid \dot{\phi}(0)=L \phi\} & =\text { domain of } A .
\end{aligned}
$$

In integrated form, (8.3) becomes

$$
\begin{equation*}
z_{t}=T(t) z_{0}+\epsilon \int_{0}^{t} T(t-s) X_{0} F\left(z_{s}, \alpha, \epsilon\right) d s \tag{8.4}
\end{equation*}
$$

where $X_{0}=X_{0}(\theta)$ is given by

$$
X_{0}(\theta)=\left\{\begin{array}{l}
I=n \times n \text { identity matrix, } \quad \theta=0, \\
0, \quad-r \leqslant \theta<0 .
\end{array}\right.
$$

Strictly speaking $X_{0}$ does not belong to $C$ because of the discontinuity at $\theta=0$; nevertheless $X_{0}$ can serve as the initial condition for the linear equation, so the semigroup can act on it to produce $T(t) X_{0}$. Equation (8.4) must be interpreted as an equality for each $\theta \in[-r, 0]$, but one may informally think of it as an equation in $C$. If (8.4) is differentiated with respect to $t$. we obtain the formal expression

$$
\begin{equation*}
(d / d t) z_{t}-A z_{t}+\epsilon X_{0} F\left(z_{t}, \alpha, \epsilon\right) \tag{8.5}
\end{equation*}
$$

As it stands, (8.5) does not make sense, for two reasons:
(1) In general $z_{t}$ does not belong to the domain of $A$; it is certainly $C^{1}$ (at least for $t \geqslant r$ ) but may not satisfy $\dot{z}(t)=L_{z_{t}}$.
(2) The nonlinear term is a multiple of $X_{0}$, hence does not belong to $C$.

We shall show that (8.5) does make sense if interpreted correctly. Both of the above problems can be remedied at once if we think of $z_{i}$ as belonging to

$$
z_{t} \in C^{1}=\left\{\phi \in C \mid \phi \text { is of class } C^{1}\right\}
$$

and extend the domain of $A$ to all of $C^{1}$.
To extend the domain of $A$, consider the formula for $A^{\mathbf{1}}$; to solve $A \phi=\psi$ for $\phi$ we have

$$
\begin{equation*}
\phi(\theta)=\phi(0)+\int_{0}^{\theta} \psi(s) d s \tag{8.6}
\end{equation*}
$$

with $\phi(0)$ determined by

$$
\begin{align*}
\psi(0) & =L\left[\phi(0)+\int_{0}^{\theta} \psi(s) d s\right] \\
& =\left[\int_{-r}^{0} d \eta(\theta)\right] \phi(0)+\int_{-r}^{0} d \eta(\theta) \int_{0}^{\theta} \psi(s) d s \tag{8.7}
\end{align*}
$$

As long as

$$
\operatorname{det}\left[\int_{-r}^{0} d \eta(\theta)\right] \neq 0
$$

then $\phi(0)$ is uniquely determined, hence $A$ has a bounded inverse. What is crucial is that formulas (8.6), (8.7) are defined even for $\psi=X_{0}$. In this case, $\phi$ is the constant (matrix valued) function

$$
\begin{equation*}
\phi=A^{-1} X_{0}=\left[\int_{-r}^{0} d \eta(\theta)\right]^{-1} \tag{8.8}
\end{equation*}
$$

This means that the domain of $A$ has been extended to include all constant functions, provided we let $A$ take values in.

$$
B C=C \oplus\left\langle X_{0}\right\rangle
$$

the space of all functions continuous on $-r \leqslant \theta<0$, with a jump discontinuity at $\theta=0$. It is easy to see that $A$ is well defined on $C^{1}$, since any $C^{1}$ function $\phi$ can be uniqucly written as

$$
\begin{aligned}
\phi & =\phi^{1}+\phi^{2} \\
\dot{\phi}^{1}(0) & =L \phi^{1}, \\
\phi^{2} & =\text { constant function. }
\end{aligned}
$$

Indeed, to attain this decomposition it is enough to let $\phi^{2}$ be defined by

$$
\dot{\phi}(0)=I\left(\phi-\phi^{2}\right)=L \phi-\left[\int_{-r}^{0} d \eta(\theta)\right] \phi^{2} .
$$

Once this is done, we have $A$ defined on all of $C^{1}$ by

$$
A \phi=A \phi^{1}+A \phi^{2}=\dot{\phi}+X_{0}\left[\int_{-r}^{0} d \eta(\theta)\right] \phi^{2}
$$

from (8.8), which implies that

$$
\begin{equation*}
A \phi=\dot{\phi}+X_{0}[L \phi-\dot{\phi}(0)] \tag{8.9}
\end{equation*}
$$

for all $\phi \in C^{1}$. Observe here that we may now drop the restriction that $A^{-1}$ exist, and let (8.9) define $A$ in its extended domain $C^{1}$. With $A$ interpreted in this extended sense, we claim in the following theorem that (8.5) holds for solutions of the RFDE (8.2), as long as $t \geqslant r$.

Theorem 8.1. Consider the retarded functional differential equation

$$
\begin{equation*}
\dot{z}(t)=L z_{t}+\epsilon F\left(z_{t}, \alpha, \epsilon\right) \tag{8.10}
\end{equation*}
$$

where

$$
L \phi==\int_{-r}^{0} d \eta(\theta) \phi(\theta)
$$

Let the operator $A$ map $C^{1}$ into $B C=C \oplus\left\langle X_{0}\right\rangle$ by

$$
A \phi=\dot{\phi}+X_{0}[L \phi-\dot{\phi}(0)]
$$

Then any solution of (8.10) for $t \geqslant t_{0}$, satisfies

$$
\begin{equation*}
(d / d t) z_{t}=A z_{t}+\epsilon X_{0} F\left(z_{t}, \alpha, \epsilon\right) \tag{8.11}
\end{equation*}
$$

as long as $t \geqslant t_{0}+r$ (or, in fact, as long as $z_{t} \in C^{1}$ ).
The proof of this is immediate, so it is omitted.
With this interpretation of the retarded equation (8.10) as the abstract ODE (8.11), all that needs to be done before averaging is to decompose $B C$ as $P \oplus Q$ with the two-dimensional eigenspace $P \subset C^{1}$ and complement $Q$, according to the spectrum of $A$. This is done by Hale and we review the main points in preparation for the examples of later sections.

The spectrum of $A$ is determined, as with ODE's, by exponential solutions of $\dot{z}(t)=L z_{t}$. In particular, all spectral values $\lambda$ are isolated and of finite multiplicity, and are determined by solving the characteristic equation.

$$
\operatorname{det}\left[\lambda I-\int_{-r}^{0} d \eta(\theta) e^{\lambda \theta}\right]=0
$$

The elements of the eigenspaces are all exponential polynomials. A basis $\Phi=$ ( $\phi^{1}, \phi^{2}$ ) for $P$ is chosen, as well as a dual basis $\Psi=\operatorname{col}\left(\psi^{1}, \psi^{2}\right)$ for the adjoint equation, where

$$
\psi^{i} \in C^{*} \stackrel{\text { def }}{=} C\left([0, r], R^{n}\right) \quad \text { (row vectors) }
$$

With the bilinear form on $C^{*} \times C$ defined by

$$
(\psi, \phi)=\psi(0) \phi(0)-\int_{-r}^{0} \int_{0}^{\theta} \psi(s-\theta) d \eta(\theta) \phi(s) d s
$$

we assume $(\Psi, \Phi)=I$. The projections $\phi^{P}$ and $\phi^{\circ}$ of any $\phi \in B C$ onto $P$ and $Q$ are given by

$$
\begin{aligned}
\phi^{P} & =\Phi(\Psi, \phi), \\
\phi^{O} & =\phi-\phi^{P} .
\end{aligned}
$$

Relative to the basis $\Phi$, the operator $A$ restricted to $P$ may be represented by a matrix $A_{P}$ defined by

$$
\begin{equation*}
A \Phi=(d / d \theta) \Phi=\Phi A_{P} \tag{8.12}
\end{equation*}
$$

we let $A_{Q}$ denote $A$ restricted to $Q$. Thus, for each $t$ we may decompose $z_{t}$ as

$$
\begin{aligned}
z_{t} & =z_{t}^{P}+z_{t}^{0} \\
& =\Phi\left(\Psi, z_{t}\right)+z_{t}^{o} \\
& \stackrel{\text { def }}{=} \Phi x(t)+y_{t}
\end{aligned}
$$

where $x(t)$ lies in $R^{2}$ and $y_{t}$ takes values in $Q$, but does not necessarily satisfy $y_{t}(\theta)=y(t+\theta)$. With this decomposition (8.11) becomes

$$
\begin{aligned}
\dot{x}(t) & =A_{p} x(t)+\epsilon \Psi(0) F\left(\Phi_{x(t)}+y_{t}, \alpha, \epsilon\right), \\
(d d d t) y_{t} & =A_{o} y_{t}+X_{0} O F\left(\Phi x(t)+y_{t}, \alpha, \epsilon\right) .
\end{aligned}
$$

By writing $x$ in polar coordinates $(r, \zeta)$, we may now average over the angle $\zeta$.

## 9. Wright's Equation

We study the Hopf bifurcation for Wright's equation

$$
\begin{equation*}
\dot{z}(t)=-a z(t-1)[1+z(t)] \tag{9.1}
\end{equation*}
$$

with the real parameter $a$. This equation has been studied by many authors; in particular, it is known that the origin $z=0$ is stable when $0<a<\pi / 2$ and unstable when $a>\pi / 2$. Moreover, for all $a>\pi / 2$, (9.1) always has a periodic solution. Such solutions have been shown to exist by means of topological fixed point theorem; however, this method sheds no light on qualitative behavior, such as amplitude, period, or stability. Using integral averaging we shall study the local behavior of periodic solutions near the bifurcation points

$$
\begin{equation*}
a_{N}=(-1)^{N}[(\pi / 2)+N \pi], \quad z=0 \tag{9.2}
\end{equation*}
$$

for each nonnegative integer $N$. In all cases the constant $K$ turns out to be nonzero, so a generic bifurcation occurs. In fact, our computations reveal the following.

Theorem 9.1. Consider Wright's equation (9.1). At each bifurcation point $a=a_{N}$ in (9.2), a generic Hopf bifurcation occurs from $z=0$. The bifurcating solution has the form

$$
z(t)=\left[\frac{20\left(a-a_{N}\right)}{3 a_{N}-1}\right]^{1 / 2} \cos a_{N} t+O\left(a-a_{N}\right)
$$

for a near $a_{N}$. Thus bifurcation occurs to the right of $a_{N}$ for $N$ even, and the left


Figure 9.1
for $N$ odd, in accordance with Fig. 9.1. Only the solution at $N=0$ is stable, and it then has the form

$$
z(t)=\left[\frac{40(a-(\pi / 2))}{3 \pi-2}\right]^{1 / 2} \cos \left(\frac{\pi t}{2}\right)+O\left(a-\frac{\pi}{2}\right)
$$

with the constant

$$
[40 /(3 \pi-2)]^{1 / 2} \cong 2 \cdot 3210701
$$

The proof of this theorem consists in writing (9.1) as an abstract ODE, and then averaging. The characteristic equation near $z=0$ is

$$
\begin{equation*}
\lambda+a e^{-\lambda}=0 \tag{9.3}
\end{equation*}
$$

and at $a=a_{N}$ has imaginary roots

$$
\begin{aligned}
\lambda & = \pm i b_{N} \\
b_{N} & =\left|a_{N}\right|=(\pi / 2)+N \pi .
\end{aligned}
$$

For $a$ near $a_{N}$, there is a unique pair of conjugate roots $\lambda(a), \overline{\lambda(a)}$ near $\pm i b_{N}$ with $\lambda\left(a_{N}\right)=i b_{N}$. This follows from the implicit function theorem; moreover, differentiating (9.3) shows

$$
\begin{equation*}
\nu=\operatorname{Re} \lambda^{\prime}\left(a_{N}\right)=a_{N} /\left(1+a_{N}^{2}\right) \neq 0 \tag{9.4}
\end{equation*}
$$

Let us scale $z \rightarrow \epsilon z$ and set $a=a_{N}$. We compute the constant $K$ of Theorem 5.1 , and as noted before, $K$ depends only on the differential equation at $a=a_{N}$.

Consider then

$$
\dot{z}(t)=-a_{N} z(t-1)-\epsilon a_{N} z(t-1) z(t)
$$

Hale gives a basis

$$
\Phi(\theta)=\left(\cos b_{N} \theta, \sin b_{N} \theta\right)
$$

for the eigenspace corresponding to $\lambda= \pm i b_{N}$, and a dual basis

$$
\Psi(\theta)=\frac{2}{1+b_{N}{ }^{2}}\binom{\cos b_{N} \theta-b_{N} \sin b_{N} \theta}{\sin b_{N} \theta+b_{N} \cos b_{N} \theta}
$$

relative to the bilinear form

$$
(\psi, \phi)=\psi(0) \phi(0)-a_{N} \int_{-1}^{0} \psi(\theta+1) \phi(\theta) d \theta
$$

Thus $(\Psi, \Phi)=I$, the identity matrix. Upon decomposing as in Section 8 we obtain the equations

$$
\begin{align*}
\dot{x}(t) & =A_{p} x(t)+\epsilon \Psi(0) f\left(\Phi x(t)+y_{t}\right) \\
(d / d t) y_{t} & =A_{Q} y_{t}+\epsilon X_{0}{ }^{\circ} f\left(\Phi_{x}(t)+y_{t}\right)  \tag{9.5}\\
f(\phi) & =-a_{N} \phi(-1) \phi(0)
\end{align*}
$$

The defining condition (8.12) for $A_{P}$ implies

$$
A_{P}=\left(\begin{array}{cc}
0 & b_{N} \\
-b_{N} & 0
\end{array}\right)
$$

and we observe

$$
\begin{aligned}
A \phi & =\dot{\phi}-X_{0}\left[a_{N} \phi(-1)+\dot{\phi}(0)\right], \quad \phi \in C^{1} \\
A_{Q} & =A \text { restricted to } O=\left\{\phi \in C^{1} \mid(\Psi, \phi)=0\right\} \\
X_{0}{ }^{\circ} & =X_{0}-X_{0}^{P}=X_{0}-\Phi \Psi(0)
\end{aligned}
$$

In polar coordinates (9.5) becomes

$$
\begin{aligned}
\dot{r} & =\frac{2 \epsilon}{1+b_{N}^{2}}\left(\cos \zeta+b_{N} \sin \zeta\right) f\left(\Phi x+y_{t}\right), \\
\dot{\zeta} & =-b_{N}+\frac{2 \epsilon}{r\left(1+b_{N}^{2}\right)}\left(b_{N} \cos \zeta-\sin \zeta\right) f\left(\Phi x+y_{t}\right), \\
\frac{d}{d t} y_{t} & =\text { as before } \\
x & =\operatorname{col}(r \cos \zeta, r \sin \zeta), \\
f\left(\Phi x+y_{t}\right) & =-a_{N}\left((-1)^{N+1} r \sin \zeta+y_{t}(-1)\right)\left(r \cos \zeta+y_{t}(0)\right)
\end{aligned}
$$

Therefore, with the notation of Theorem 5.1, and in particular (5.6), we have

$$
\begin{aligned}
C_{3}(\zeta, 0,0)= & \frac{2 b_{N}}{1+b_{N}{ }^{2}}\left(\cos \zeta+b_{N} \sin \zeta\right) \sin \zeta \cos \zeta \\
C_{4}(\zeta, 0,0)= & 0 \\
D_{3}(\zeta, 0,0)= & \frac{2 b_{N}}{1+b_{N}^{2}}\left(b_{N} \cos \zeta-\sin \zeta\right) \sin \zeta \cos \zeta \\
\omega_{0}= & -b_{N} \\
G_{2}(\zeta, 0,0) \phi= & \frac{2 b_{N}}{1+b_{N}^{2}}\left(\cos \zeta+b_{N} \sin \zeta\right) \\
& \times\left(\phi(0) \sin \zeta+(-1)^{N+1} \phi(-1) \cos \zeta\right) \\
J_{2}(0,0)(\cos \zeta, \sin \zeta)^{2}= & b_{N}(\cos \zeta \sin \zeta) X_{0} O
\end{aligned}
$$

Direct calculation yields

$$
K^{*}=\left(1 / 2 \pi b_{N}\right) \int_{0}^{2 \pi} C_{3} D_{3} d \zeta=0
$$

Writing $G_{2}$ as

$$
G_{2}(\zeta, 0,0)=\sum g_{n} e^{i n \zeta}
$$

we solve Eq. (5.7)

$$
G_{2}(\zeta)-w^{*^{\prime}}(\zeta) b_{N}+w^{*}(\zeta) A_{O}=0
$$

to yield

$$
w^{*}(\zeta)=\sum g_{n}\left(i n b_{N}-A_{O}\right)^{-1} e^{i n \zeta}
$$

From Theorem 5.1 we then have

$$
\begin{aligned}
K^{* *} & =\frac{b_{N}}{2 \pi} \sum g_{n}\left(i n b_{N}-A_{Q}\right)^{-1} X_{0} o \int_{0}^{2 \pi} e^{i n \zeta} \cos \zeta \sin \zeta d \zeta \\
& =\frac{i b_{N}}{4}\left[g_{2}\left(2 i b_{N}-A_{\varrho}\right)^{-1}+g_{-2}\left(2 i b_{N}+A_{O}\right)^{-1}\right] X_{0} o
\end{aligned}
$$

Since $X_{0}{ }^{\circ}=X_{0}-\Phi \Psi(0)$ and $G_{2}(\zeta, 0,0)$ are real, we have $g_{-2}=\bar{g}_{2}$, which implies

$$
K^{* *}=\left(-b_{N} / 2\right) \operatorname{Im} g_{2}\left(2 i b_{N}-A_{Q}\right)^{-1} X_{0}^{o} .
$$

First let us determine the linear functional $g_{2}$. By writing $\cos \zeta \sin \zeta$ in terms of $e^{ \pm i \zeta}$ it is easy to see

$$
g_{2} \phi=\overline{2(1} \frac{b_{N}}{\left.+b_{N}^{2}\right)}-\left(1 \quad i b_{N}\right)\left(-\phi(0) i \mid(1)^{N+1} \phi(\quad 1)\right)
$$

for all $\phi \in C$. Next we determine

$$
\phi=\left(2 i b_{N} \quad A_{O}\right)^{-1} X_{0}^{o}
$$

To calculate, more generally,

$$
\phi=\left(2 i b_{N}-A\right)^{-1} \psi
$$

we must solve

$$
\dot{\phi}(\theta)=2 i b_{N} \phi(\theta)-\psi(\theta)
$$

subject to the boundary condition

$$
\dot{\phi}(0)=-a_{N} \phi(-1) .
$$

The solution is easily obtained as

$$
\begin{aligned}
\phi(\theta) & =e^{2 i b_{N} \theta} \phi(0)-\int_{0}^{\theta} e^{2 i b_{N}(\theta-s)} \psi(s) d s \\
\phi(0) & =\left[2 i b_{N}+a_{N} e^{-2 i b_{N}}\right]^{-1}\left[\psi(0)+a_{N} \int_{0}^{-1} e^{-2 i b_{N}(1+s)} \psi(s) d s\right] \\
& =[(\pi / 2)+N \pi]^{-1}\left(2 i+(-1)^{N+1}\right)^{-1}\left[\psi(0)+a_{N} \int_{0}^{-1} e^{-2 i b_{N}(1+s)} \psi(s) d s\right]
\end{aligned}
$$

For $\psi=X_{0}$ we then have

$$
\begin{aligned}
\phi(0)= & \left(\frac{\pi}{2}+N \pi\right)^{-1}\left(2 i+(-1)^{N+1}\right)^{-1}, \\
\phi(-1)= & -\phi(0), \\
g_{2} \phi= & g_{2}\left(2 i b_{N}-A\right)^{-1} X_{0} \\
= & \frac{b_{N}}{2\left(1+b_{N}^{2}\right)}\left(1-i b_{N}\right)\left(-1+(-1)^{N}\right) \\
& \times\left(\frac{\pi}{2}+N\right)^{-1}\left(2 i+(-1)^{N+1}\right)^{-1}, \\
\frac{-b_{N}}{2} \operatorname{Im}\left[g_{0}\left(2 i b_{N}-A\right)^{-1} X_{0}\right]= & -\frac{b_{N}^{2}\left[3 b_{N}+(-1)^{N-1}\right]}{20\left(1+b_{N}^{2}\right)\left[(\pi / 2)+N_{\pi}\right]} .
\end{aligned}
$$

The calculation of $\left(2 i b_{N}-A\right)^{-1} X_{0}{ }^{P}$ uses the fact that $X_{0}{ }^{P}=\Phi \Psi(0)$, and that $A$ is represented by the matrix $A_{P}$ relative to the basis $\Phi$ of $I$. Thus

$$
\begin{aligned}
\left(2 i b_{N}-A\right)^{-1} X_{0}^{P} & =\Phi\left(2 i b_{N}-A_{P}\right)^{-1} \Psi(0) \\
& =\Phi\left(\begin{array}{cc}
2 i b_{N} & b_{N} \\
b_{N} & 2 i b_{N}
\end{array}\right)^{-1}\binom{1}{b_{N}} \frac{2}{1+b_{N}^{2}} \\
& =\frac{-2}{3\left(1+b_{N}^{2}\right)} \Phi\binom{\left(2 i j b_{N}\right)+1}{\left(-1 / b_{N}\right)+2 i}
\end{aligned}
$$

and this yields

$$
\begin{aligned}
g_{2}\left(2 i b_{N}-A\right)^{-1} X_{0}^{P}= & \frac{b_{N}}{2\left(1+b_{N}^{2}\right)}\left(1+i b_{N}\right)\left(\frac{-2}{3\left(1+b_{N}^{2}\right)}\right)\left(i+\frac{1}{b_{N}}\right) \\
& -\frac{b_{N}}{2} \operatorname{Im}\left[g_{2}\left(2 i b_{N}-A\right)^{-1} X_{0}^{P}\right]=0 .
\end{aligned}
$$

We therefore finally obtain

$$
K=K^{* *}=-\frac{b_{N}{ }^{2}\left[3 b_{N}+(-1)^{N+1}\right]}{20\left(1+b_{N}^{2}\right)[(\pi / 2)+N \pi]}=-\frac{b_{N}\left[3 b_{N}+(-1)^{N+1}\right]}{20\left(1+b_{N}^{2}\right)}
$$

Thus the averaged equation has the normal form

$$
\begin{aligned}
\dot{r} & =\epsilon \alpha \nu r+\epsilon^{2} r^{3} K+O\left(\epsilon^{3}\right) \\
\dot{\zeta} & =-b_{N}+O(\epsilon) \\
a-a_{N} & =\epsilon \alpha
\end{aligned}
$$

where $\nu$ is as in (9.4). The form of the bifurcating solution (in unscaled coordinates now)

$$
r \sim|\nu / K|^{1 / 2} \epsilon, \quad a-a_{N}=-\operatorname{sgn}(\nu K) \epsilon^{2}
$$

is clear, and this implies the theorem.

## 10. Equations with First Integrals

Consider the equation

$$
\begin{equation*}
z(t)=a \int_{t-L_{1}-L_{2}}^{t-L_{1}} g(z(s)) d s \tag{10.1}
\end{equation*}
$$

where $L_{1}>0$ and $L_{2}>0$ are fixed constants, $a$ is a parameter, and $g$ is a known smooth function mapping the reals into the reals. Observe that (10.1) can be written as a functional differential equation:

$$
\begin{equation*}
\dot{z}(t)=a g\left(z\left(t-L_{1}\right)\right)-a g\left(z\left(t-L_{1}-L_{2}\right)\right) \tag{10.2}
\end{equation*}
$$

with a first integral

$$
I(\varphi)=\varphi(0)-\int_{-L_{1}-L_{2}}^{-L_{1}} g(\varphi(\theta)) d \theta, \quad \varphi \in C\left[-L_{1}-L_{2}, 0\right]
$$

For appropriate choices of $g, L_{1}$, and $L_{2}$, there is an equilibrium $c=c(a)$ obtained by solving

$$
c-a L_{2} g(c)=0 .
$$

The linearized equation about the equilibrium $c=c(a)$ has the following characteristic equation

$$
\lambda-a g^{\prime}(c(a))\left[e^{-L_{1} \lambda}-e^{-\left(L_{1}+L_{2}\right) \lambda}\right]=0
$$

Note that zero is always an eigenvalue because of the first integral $I(\varphi)$. In addition, we may have a pair of conjugate eigenvalues $\lambda(a)$ and $\lambda(a)$ crossing the imaginary axis with nonzero speed at $a=a_{0}$.

In general, one may expect a Hopf bifurcation occurring on each integral surface

$$
I(\varphi)=\text { constant }
$$

where $I$ is near zero. Since the integral surface

$$
I(\varphi)=0
$$

is of interest in applications, we consider only the bifurcation in that surface.
To do the integral averaging on the integral surface, we first decompose

$$
z_{l}-c=\Phi x+\Phi_{0} \sigma+y_{t}
$$

where

$$
\begin{aligned}
\Phi(\theta) & =(\cos \mu \theta, \sin \mu \theta), \quad \mu=\operatorname{Im} \lambda\left(a_{0}\right) \\
\Phi_{0}(\theta) & \equiv 1
\end{aligned}
$$

are the bases for the appropriate eigenspaces, $x \in R^{2}, \sigma \in R^{1}$, and $y_{t}$ is in the complementary subspace. As in the case of Wright's equation, we obtain the following equations after scaling,

$$
\begin{align*}
\dot{x}(t) & =A_{P} x(t)+\epsilon \Psi(0) N\left(x(t), \sigma(t), y_{t}, \epsilon\right), \\
\sigma(t) & =\epsilon \Psi_{0}(0) N\left(x(t), \sigma(t), y_{t}, \epsilon\right)  \tag{10.3}\\
(d / d t) y_{t} & =A_{Q}\left(y_{t}\right)+\epsilon X_{0}^{Q} N\left(x(t), \sigma(t), y_{t}, \epsilon\right)
\end{align*}
$$

where $\Psi(\theta),-\left(L_{1}+L_{2}\right) \leqslant \theta \leqslant 0$, is the basis for the adjoint eigenspace (see Section 9 ), $\Psi_{0}(0)$ is some fixed constant, $N$ is the nonlinearity, $A_{Q}$ is the infinitesimal generator on the complementary subspace, $X_{0}{ }^{\circ}$ is as in Section 9 and

$$
A_{P}=\left(\begin{array}{cc}
0 & \mu \\
-\mu & 0
\end{array}\right)
$$

At the bifurcation,

$$
\begin{aligned}
& I\left(c+\Phi x+\Phi_{0} \sigma+y_{t}\right) \\
& \quad=\left[1-a L_{2} g^{\prime}(c(a))\right] \sigma+O\left(|x|^{2}+\sigma^{2}+\left|y_{t}\right|^{2}\right)
\end{aligned}
$$

So, after scaling we may solve $I=0$ to get

$$
\rho=\epsilon S\left(x, y_{t}, \epsilon\right)
$$

Now, by eliminating $\sigma$ from Eqs. (10.3) we obtain two equations involving $x$ and $y_{t}$ only. Thus proceeding as in Section 9 we obtain the bifurcations. Details are omitted.

## 11. Diffusion Equations

Consider the scalar parabolic equation

$$
\begin{equation*}
u_{t}=u_{x x}, \quad t \geqslant 0, \quad 0<x<1 \tag{11.1}
\end{equation*}
$$

with nonlinear boundary conditions

$$
\begin{align*}
& u_{x}(0, t)=0  \tag{11.2}\\
& u_{x}(1, t)=a g(u(0, t), u(1, t)) \tag{11.3}
\end{align*}
$$

where $a$ is a parameter and

$$
g(u, v)=\alpha u+\beta v+O\left(u^{2}+v^{2}\right)
$$

For simplicity, we assume $\alpha=1$ and $\beta=0$. A necessary and sufficient condition. for $\lambda=i \mu, \mu>0$, to be an eigenvalue is $\left(\sigma==(\mu / 2)^{1 / 2}\right)$

$$
\begin{aligned}
& 2 \sigma \sinh \sigma \cos \sigma=a_{0} \\
& 2 \sigma \cosh \sigma \sin \sigma=-a_{0}
\end{aligned}
$$

It is also not difficult to see that for appropriate parameters the eigenvalues cross the imaginary axis with nonzero speed as $a$ passes through $a_{0}$.

Scaling (11.1), (11.2), and (11.3), we obtain

$$
\begin{align*}
u_{t} & -u_{x x} \\
u_{x}(0, t) & =0  \tag{11.4}\\
u_{x}(1, t) & =a_{0} u(0, t)+\epsilon N(u(0, t), u(1, t), \epsilon)
\end{align*}
$$

where $N$ is the nonlinearity.
In order to write (11.4) as an ordinary differential equation, we consider the operator

$$
\begin{aligned}
\tilde{A}: H^{2}(0,1) \cap \tilde{H} & \rightarrow L^{2}(0,1), \\
u & \rightarrow\left(d^{2} / d x^{2}\right) u
\end{aligned}
$$

where $L^{2}(0,1)$ is the Lebesgue space, $H^{2}(0,1)$ is the standard Sobolev space, and $\tilde{H}$ consists of functions such that

$$
\left.(d \mid d x) u\right|_{x=0}=0,\left.\quad(d / d x) u\right|_{x=1}=\left.a_{0} u\right|_{x=0}
$$

Note that $H^{2}(0,1) \cap \widetilde{H}$ is well defined by the trace theorem and is in fact a subspace of codimension two. The operator $\tilde{A}$ is in fact a linear isomorphism with $\tilde{A}^{-1}$ given by

$$
\begin{align*}
\tilde{A}^{-1}: L^{2}(0,1) & \rightarrow H^{2}(0,1) \cap \tilde{H} \\
u & \rightarrow\left(1 / a_{0}\right) \int_{0}^{1} u(x) d x+\int_{0}^{x} \int_{0}^{t} u(s) d s d t \tag{11.5}
\end{align*}
$$

Because of the nonlinear boundary condition (11.3), the solution $u(\cdot, t)$ is in general not in $H^{2}(0,1) \cap \tilde{H}$. However, the solution $u(\cdot, t) \in H^{2}(0,1)$. Therefore, we extend the definition of $\tilde{A}$ to $H^{2}(0,1)$ by permitting the extended map $A$ to take values in a larger space. Specifically, observe that (11.5) is defined for the Dirac measures $\delta(x)$ and $\delta(x-1)$, and in fact

$$
\begin{align*}
\tilde{A}^{-1}(\delta(x)) & =\left(1 / a_{0}\right)+x \\
\tilde{A}^{-1}(\delta(x-1)) & =\left(1 / a_{0}\right) \tag{11.6}
\end{align*}
$$

Now, for $u \in H^{2}(0,1)$ we have

$$
u(x)=\check{u}(x)+\left(\frac{1}{a_{0}}+x\right) u_{x}(0)+\frac{1}{a_{0}}\left(a_{0} u(0)-u_{x}(1)\right),
$$

where $\tilde{u}(x) \in H^{2}(0,1) \cap \tilde{H}$ and

$$
\check{u}(x)=u(x)-x u_{x}(0)+\frac{1}{a_{0}}\left[u_{x}(1)-u_{x}(0)-a_{0} u(0)\right]
$$

Motivated by (11.6), we define

$$
\begin{aligned}
A: H^{2}(0,1) & \rightarrow L^{2}(0,1) \oplus\langle\delta(x), \delta(x-1)\rangle \\
u(x) & \rightarrow u_{x x}(x)+u_{x}(0) \delta(x)+\left[a_{0} u(0)-u_{x}(0)\right] \delta(x-1)
\end{aligned}
$$

where $\langle\delta(x), \delta(x-1)\rangle$ denotes the vector space generated by $\delta(x)$ and $\delta(x-1)$.
Using the definition of $A$, we may rewrite (11.4) as an ordinary differential equation:

$$
\begin{aligned}
d u / d t & =A u-u_{x}(0) \delta(x)-\left[a_{0} u(0)-u_{x}(1)\right] \delta(x-1) \\
& =A u+\epsilon N(u(0, t), u(1, t), \epsilon) \delta(x-1) \\
& =f(u, \epsilon)
\end{aligned}
$$

where $f: H^{2}(0,1) \times\left(-\epsilon_{0}, \epsilon_{0}\right) \rightarrow L^{2}(0,1) \oplus\langle\delta(x), \delta(x-1)\rangle$. Now, we may proceed to discuss the bifurcations as before.

## 12. Notes and Remarks

Section 2. Changes of coordinates as (2.3) are given in the textbook of Hale [16] and a survey paper of Volosov [47]. For more references on integral averaging, see the references in Hale [15, 16] and Volosov [47].

In the literature, Eq. (2.4) is usually averaged by eliminating the variables $r$ and $\theta$ in the equation for $\dot{y}$ rather than eliminating the variable $y$ in the equations for $\dot{r}$ and $\dot{\theta}$, as was done here. For example, in the book of Lefschetz [33] such an approach is to be used to study stability properties in which no bifurcation parameter appears (see also the papers of Hale [19] and Hausrath [[22]).

Section 3. The resonance condition of Corollary 3.2 for the eigenvalues has been used by a number of authors, notably, Sternberg [45], Hartman [21], and Lefschetz [33]. If the resonance condition is not satisfied, then it is impossible to average the equations as required since certain terms cannot be eliminated. In principle, one can eliminate these terms by a homeomorphic (but not diffeomorphic) change of coordinates. In practice, these changes of coordinates are not necessary as one could analyze the problem even with the presence of these terms.

Section 5. The existence of the invariant forms described at the end of this section follows from Chapter 18 of Hale [15].

Section 6. Apparently, the center manifold technique was first used by Chafee [6, 7] and also by Ruelle and Takens [39] (see also [32]). For PDE's, it was used by Marsden [35] and McCracken and Marsden [36]. Dorroh and Marsden [11] showed that for a large class of PDE, Properties (1), (2), and (3) hold.

Section 8. For the theory of functional differential equations, see Hale [17]. The space $B C=C \oplus\left\langle X_{0}\right\rangle$ was used by Hale [19] in studying stability properties of equilibria of neutral FDE in critical cases.

Section 9. Equation (9.1) has been studied by many authors, notably, Wright [48] and Kakutani and Markus [29]. Jones [25] first showed that there are periodic solutions of (9.1) for $a>\pi / 2$.

Section 10. Equation (10.1) is model of population growth introduced by Cooke and Yorke [10]. Hale [20] showed that for generic $g$ 's there is a Hopf bifurcation. Greenberg [14] studied the existence of periodic solutions for large values of $a$.

Section 11. The problem (11.1), (11.2), and (11.3) is a simplified model of a diffusion equation with nonlinear boundary conditions, occurring in the study of the interaction and production of two enzymes (see [3]). The actual system is:

$$
\begin{aligned}
u_{t} & =u_{x x}-Q^{2} u \\
v_{t} & =v_{n x}-Q^{2} v \\
u_{x}(0, t) & =-P Q f(v(0, t)), \\
u_{x}(1, t) & =0 \\
v_{x}(0, t) & =0 \\
v_{x}(1, t) & =P Q g(u(1, t)),
\end{aligned}
$$

where $P$ and $Q$ are constants and $f$ and $g$ are nonlinearities.
For parabolic equations in $R^{n}$ with nonlinear boundary conditions a similar procedure may be followed. Consider, for example,

$$
\begin{aligned}
u_{t} & =\Delta u \quad \text { in } \Omega \subseteq R^{n}, \\
\partial u / \partial n & =\xi(u) \quad \text { on } \partial \Omega,
\end{aligned}
$$

where $\xi$ is a nonlinear function which may depend on values of $u$ even in $\Omega$. Define

$$
\begin{aligned}
A: H(\Omega) & \rightarrow H^{1}(\Omega)^{*}, \\
u & \rightarrow\left[v \rightarrow-\int_{\Omega} \nabla u \cdot \nabla v\right]
\end{aligned}
$$

where $H^{1}(\Omega)^{*}$ is the dual space of $H^{1}(\Omega)$. Formally, if $\langle$,$\rangle represents the duality$ between $H^{1}(\Omega)^{*}$ and $H^{1}(\Omega)$ we have

$$
\begin{aligned}
\langle A u, v\rangle & =-\int_{\partial \Omega}(\partial u / \partial n) v+\int_{\Omega}(\Delta u) v \\
& =-\int_{\partial \Omega} \xi(u) v+\int u_{t} v^{2}
\end{aligned}
$$

Assume that the nonlinear functional

$$
\xi: H^{1}(\Omega) \rightarrow H^{-1 / 2}(\partial \Omega)
$$

and define

$$
\begin{aligned}
& N: H^{1}(\Omega) \rightarrow H^{1}(\Omega)^{*} \\
& \langle N(u), v\rangle=\int_{\delta \Omega} \xi(u) v .
\end{aligned}
$$

Thus we obtain for the differential equation

$$
\int_{\Omega} u_{t} v=\langle A u+N(u), v\rangle, \quad v \in H^{2}(\Omega)
$$

If we identify the function $u_{t}$ with the element of $H^{1}(\Omega)^{*}$ defined by $v \rightarrow \int_{\Omega} u_{t} v$, $v \in H^{1}(\Omega)$, the differential equation may be written

$$
d u / d t=A u+N(u)
$$

where $u \in H^{1}(\Omega)$, $d u / d t \in H^{1}(\Omega)^{*}$. For more details see, for example, Lions and Magenes [34].

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    ${ }^{1}$ If $\operatorname{Im} \lambda(0)=0$, then it is not periodic but equilibrium states which bifurcate from the zero solution. Such bifurcation theory is not within the scope of the present paper. In $[8,9]$, we consider bifurcation problems under precisely the condition $\operatorname{Im} \lambda(0)=0$.

