#### Physics Letters B 732 (2014) 167-168

Contents lists available at ScienceDirect

Physics Letters B

www.elsevier.com/locate/physletb

# Vanishing Higgs one-loop quadratic divergence in the scotogenic model and beyond



## Ernest Ma

Department of Physics and Astronomy, University of California, Riverside, CA 92521, USA

#### ARTICLE INFO

Article history: Received 16 January 2014 Received in revised form 21 March 2014 Accepted 23 March 2014 Available online 27 March 2014 Editor: J. Hisano

#### ABSTRACT

It is shown that the inherent one-loop quadratic divergence of the Higgs mass renormalization of the standard model may be avoided in the well-studied scotogenic model of radiative neutrino mass as well as other analogous extensions.

© 2014 The Author. Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/3.0/). Funded by SCOAP<sup>3</sup>.

In quantum field theory, the additive renormalization of  $m^2$  for a scalar field of mass *m* is a quadratic function of the cutoff scale  $\Lambda$ . The elegant removal of this quadratic divergence is a powerful theoretical argument for the existence of supersymmetric particles. However, given the recent discovery of the 126 GeV particle [1,2] at the Large Hadron Collider (LHC), presumably the long sought Higgs boson of the standard model, and the nonobservation of any hint of supersymmetry, it may be a good time to reexamine an alternative solution of the quadratic divergence problem.

It was suggested a long time ago [3] that in the standard model of quarks and leptons, the condition

$$\frac{3}{2}M_W^2 + \frac{3}{4}M_Z^2 + \frac{3}{4}m_H^2 = \sum_f N_f m_f^2,$$
(1)

where  $N_f = 3$  for quarks and  $N_f = 1$  for leptons, would make the coefficient of the  $\Lambda^2$  contribution to  $m_H^2$  vanish. This would predict  $m_H = 316$  GeV, which we now know to be incorrect.

The same idea may be extended to the case of two Higgs doublets [4–7] where  $\langle \phi_{1,2}^0 \rangle = v_{1,2}$ , with  $v = \sqrt{v_1^2 + v_2^2} = 174$  GeV. In that case, the vanishing of quadratic divergences would also depend on how  $\Phi_{1,2}$  couple to the quarks and leptons. In the scotogenic model of radiative neutrino mass [8], there are two scalar doublets  $(\phi^+, \phi^0)$  and  $(\eta^+, \eta^0)$ , distinguished from each other by a discrete  $Z_2$  symmetry, under which  $\Phi$  is even and  $\eta$  odd. Thus only  $\phi^0$  acquires a nonzero vacuum expectation value v. This same discrete symmetry also prevents  $\eta$  from coupling to the usual quarks and leptons, except for the Yukawa terms

$$\mathcal{L}_{Y} = h_{ij} (\nu_{i} \eta^{0} - l_{i} \eta^{+}) N_{j} + H.c., \qquad (2)$$

where  $N_j$  are three neutral singlet Majorana fermions odd under  $Z_2$ . As a result, neutrinos obtain one-loop finite radiative Majorana masses as shown in Fig. 1. This is a well-studied model



Fig. 1. One-loop generation of neutrino mass with Z<sub>2</sub> symmetry.

which also offers  $\sqrt{2}Re(\eta^0)$  as a good dark-matter candidate [9]. The lightest *N* may also be a dark-matter candidate [10] but is more suitable if the dark-matter discrete symmetry  $Z_2$  is extended to  $U(1)_D$  as proposed recently [11].

The scalar potential of the scotogenic  $Z_2$  model is given by [8]

$$V = m_1^2 \Phi^{\dagger} \Phi + m_2^2 \eta^{\dagger} \eta + \frac{1}{2} \lambda_1 (\Phi^{\dagger} \Phi)^2 + \frac{1}{2} \lambda_2 (\eta^{\dagger} \eta)^2 + \lambda_3 (\Phi^{\dagger} \Phi) (\eta^{\dagger} \eta) + \lambda_4 (\Phi^{\dagger} \eta) (\eta^{\dagger} \Phi) + \frac{1}{2} \lambda_5 [(\Phi^{\dagger} \eta)^2 + (\eta^{\dagger} \Phi)^2].$$
(3)

Let  $\phi^0 = v + H/\sqrt{2}$  and  $\eta^0 = (\eta_R + i\eta_I)/\sqrt{2}$ , then

$$m^2(H) = 2\lambda_1 \nu^2, \tag{4}$$

$$m^2(\eta^{\pm}) = m_2^2 + \lambda_3 v^2,$$
 (5)

$$m^{2}(\eta_{R}) = m_{2}^{2} + (\lambda_{3} + \lambda_{4} + \lambda_{5})v^{2},$$
(6)

$$m^{2}(\eta_{I}) = m_{2}^{2} + (\lambda_{3} + \lambda_{4} - \lambda_{5})v^{2}.$$
(7)

The corresponding two conditions for the vanishing of quadratic divergences are

$$\frac{3}{2}M_W^2 + \frac{3}{4}M_Z^2 + \frac{3}{4}m_H^2 + \left(\lambda_3 + \frac{1}{2}\lambda_4\right)\nu^2 = 3m_t^2,$$
(8)

http://dx.doi.org/10.1016/j.physletb.2014.03.047

0370-2693/© 2014 The Author. Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/3.0/). Funded by SCOAP<sup>3</sup>.





Fig. 2. One-loop generation of electron mass with soft Z<sub>2</sub> breaking.

$$\frac{3}{2}M_W^2 + \frac{3}{4}M_Z^2 + \left(\frac{3}{2}\lambda_2 + \lambda_3 + \frac{1}{2}\lambda_4\right)v^2 = \sum_{i,j}h_{ij}^2v^2.$$
 (9)

Consequently, the following two sum rules are obtained:

$$\lambda_3 + \frac{1}{2}\lambda_4 = \frac{3}{\nu^2} \left( m_t^2 - \frac{1}{2}M_W^2 - \frac{1}{4}M_Z^2 - \frac{1}{4}m_H^2 \right) = 2.063, \quad (10)$$

$$h^{2} - \frac{1}{2}\lambda_{2} = \frac{1}{\nu^{2}} \left( m_{t}^{2} - \frac{1}{4}m_{H}^{2} \right) = 0.863,$$
(11)

where  $3h^2 = \sum_{i,j} h_{ij}^2$ . Since  $\lambda_2$  must be positive, Eq. (11) cannot be satisfied without the Yukawa couplings of Eq. (2). In other words, the existence of *N*, hence the radiative generation of neutrino mass, is necessary for this scenario. In a model with simply a second "inert" scalar doublet [12,13], vanishing quadratic divergence will not be possible. To test Eq. (10), Eqs. (5) to (7) may be used, i.e.

$$2\lambda_4 v^2 = m_R^2 + m_I^2 - 2m_+^2. \tag{12}$$

As for  $\lambda_3$ , it may be extracted [14,15] from  $H \rightarrow \gamma \gamma$  using also  $m_+$ . However Eq. (11) is very difficult to test, because  $h^2$  and  $\lambda_2$  are not easily measurable.

Analogous extensions of the scotogenic model may also accommodate vanishing quadratic divergences. As an example, consider the addition of a charged scalar  $\chi^+$  odd under  $Z_2$ , then the electron may acquire a radiative mass by assigning  $e_R$  to be odd with the Yukawa couplings  $f\bar{e}_R N_L \chi^-$  as shown in Fig. 2, where  $N_L$  is even under  $Z_2$ , but the soft Dirac mass term  $\bar{N}_L N_R$  breaks  $Z_2$  explicitly. With the addition of  $\chi^+$ , the scalar potential has the extra terms

$$V' = m_3^2 \chi^+ \chi^- + \frac{1}{2} \lambda_6 (\chi^+ \chi^-)^2 + \lambda_7 (\chi^+ \chi^-) (\Phi^{\dagger} \Phi) + \lambda_8 (\chi^+ \chi^-) (\eta^{\dagger} \eta) + [\mu (\eta^+ \phi^0 - \eta^0 \phi^+) \chi^- + H.c.].$$
(1)

The conditions for vanishing quadratic divergence in this model are then:

$$\frac{3}{2}M_W^2 + \frac{3}{4}M_Z^2 + \frac{3}{4}m_H^2 + \left(\lambda_3 + \frac{1}{2}\lambda_4 + \frac{1}{2}\lambda_7\right)v^2 = 3m_t^2, \quad (14)$$

$$\frac{3}{2}M_W^2 + \frac{3}{4}M_Z^2 + \left(\frac{3}{2}\lambda_2 + \lambda_3 + \frac{1}{2}\lambda_4 + \frac{1}{2}\lambda_8\right)v^2 = \sum_{i,j}h_{ij}^2v^2,$$
(15)

$$3(M_Z^2 - M_W^2) + (\lambda_6 + \lambda_7 + \lambda_8)v^2 = f^2 v^2.$$
(16)

Again, verification is possible, at least in principle. Other more involved scenarios such as the scotogenic  $U(1)_D$  model [11] or that of a recent proposal [16], where all quark and lepton masses are radiative with either  $Z_2$  or  $U(1)_D$  dark matter, may also have similar viable solutions.

It is of course well-known that the one-loop vanishing of the Higgs quadratic divergence is not invariant under the renormalization-group running of the gauge, Yukawa, and quartic scalar couplings. Thus the two-loop SM contribution has also been studied [17,18]. Whereas it is impossible to have both set equal to zero, if the latter is viewed as a perturbation to the first for a physical cutoff [19], then the approximate validity of the Veltman condition remains a plausible solution. Other ideas regarding the inherent quadratic divergence of any scalar mass have also been discussed in the recent literature [20–22].

### Acknowledgements

I thank Maria Krawczyk for discussions at Scalars 2013. This work is supported in part by the U.S. Department of Energy under Grant No. DE-SC0008541.

#### References

3)

- [1] ATLAS Collaboration, G. Aad, et al., Phys. Lett. B 716 (2012) 1.
- [2] CMS Collaboration, S. Chatrchyan, et al., Phys. Lett. B 716 (2012) 30.
- [3] M. Veltman, Acta Phys. Pol. B 12 (1981) 437.
- [4] C. Newton, T.T. Wu, Z. Phys. C 62 (1994) 253.
- [5] E. Ma, Int. J. Mod. Phys. A 16 (2001) 3099.
- [6] B. Grzadkowski, P. Osland, Phys. Rev. D 82 (2010) 125026.
- [7] N. Darvishi, M. Krawczyk, in preparation.
- [8] E. Ma, Phys. Rev. D 73 (2006) 077301.
- [9] L. Lopez Honorez, E. Nezri, J.F. Oliver, M.H.G. Tytgat, J. Cosmol. Astropart. Phys. 02 (2007) 028.
- [10] J. Kubo, E. Ma, D. Suematsu, Phys. Lett. B 642 (2006) 18.
- [11] E. Ma, I. Picek, B. Radovcic, Phys. Lett. B 726 (2013) 744.
- [12] N.G. Deshpande, E. Ma, Phys. Rev. D 18 (1978) 2574.
- [13] R. Barbieri, L.J. Hall, V.S. Rychkov, Phys. Rev. D 74 (2006) 015007.
- [14] A. Arhrib, R. Benbrik, N. Gaur, Phys. Rev. D 85 (2012) 095021.
- [15] B. Swiezewska, M. Krawczyk, Phys. Rev. D 88 (2013) 035019.
- [16] E. Ma, Phys. Rev. Lett. 112 (2014) 091801.
- [17] M.S. Al-sarhi, I. Jack, D.R.T. Jones, Z. Phys. C 55 (1992) 283.
- [18] Y. Hamada, H. Kawai, K.-y. Oda, Phys. Rev. D 87 (2013) 053009.
- [19] D.R.T. Jones, Phys. Rev. D 88 (2013) 098301.
- [20] K. Fujikawa, Phys. Rev. D 83 (2013) 058501.
- [21] H. Aoki, S. Iso, Phys. Rev. D 86 (2012) 013001.
- [22] I. Masina, M. Quiros, Phys. Rev. D 88 (2013) 093003.