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Spaceborne GPS receiver antenna phase center offset and variation estimation for the Shiyan 3 satellite

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Abstract In determining the orbits of low Earth orbit (LEO) satellites using spaceborne GPS, the errors caused by receiver antenna phase center offset (PCO) and phase center variations (PCVs) are gradually becoming a major limiting factor for continued improvements to accuracy. Shiyan 3, a small satellite mission for space technology experimentation and climate exploration, was developed by China and launched on November 5, 2008. The dual-frequency GPS receiver payload delivers 1 Hz data and provides the basis for precise orbit determination within the range of a few centimeters. The antenna PCO and PCV error characteristics and the principles influencing orbit determination are analyzed. The feasibility of PCO and PCV estimation and compensation in different directions is demonstrated through simulation and in-flight tests. The values of receiver antenna PCO and PCVs for Gravity Recovery and Climate Experiment (GRACE) and Shiyan 3 satellites are estimated from one month of data. A large and stable antenna PCO error, reaching up to 10.34 cm in the z-direction, is found with the Shiyan 3 satellite. The PCVs on the Shiyan 3 satellite are estimated and reach up to 3.0 cm, which is slightly larger than that of GRACE satellites. Orbit validation clearly improved with independent k-band ranging (KBR) and satellite laser ranging (SLR) measurements. For GRACE satellites, the average root mean square (RMS) of KBR residuals improved from 1.01 cm to 0.88 cm. For the Shiyan 3 satellite, the average RMS of SLR residuals improved from 4.95 cm to 4.06 cm.

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1. Introduction

Given its all-weather availability, high precision and good continuity, spaceborne dual-frequency GPS has been successfully applied in the field of satellite navigation. With the development of GPS precise orbit determination (POD) technology, an increasing number of low Earth orbit (LEO) satellites¹⁻ are being equipped with dual-frequency GPS receivers. Such

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1000-9361 © 2016 Chinese Society of Aeronautics and Astronautics. Production and hosting by Elsevier Ltd. This is an open access article under the CC BY-NC-ND license (http://creativecommons.org/licenses/by-nc-nd/4.0/). receivers have become among the most important instruments in LEO satellite POD. Orbits established with spaceborne dual-frequency GPS technology are precise to within several centimeters.⁴⁻⁶ Neglected or mismodeled antenna phase center offset (PCO) and phase center variations (PCVs), which can be attributed to ground calibrations, deformation, fuel consumption and satellite in-flight environment, have gradually become among the most important systematic error sources in LEO GPS data processing.^{7,8} This has lead to a need to perform in-flight determination of antenna PCO and PCVs. The first in-flight calibrations for spaceborne receiver antennas were generated for JASON-1.9 In-flight calibrations have been adopted for more and more satellites to meet high precision orbit requirements, including Gravity Recovery and Climate Experiment (GRACE),^{10–12} TerraSAR-X,^{10,13,14} COSMIC,¹ GOCE¹⁶ and Haiyang 2A.¹⁷

Antenna PCO is generally defined as an offset from antenna reference point (ARP) to mean center of the wave front in the respective antenna-fixed system (AFS).¹⁸ ARP is a mechanical point that has been conventionally defined by the International GNSS Service (IGS) as the intersection of the vertical antenna axis of symmetry with the antenna bottom.¹⁹ A priori PCO vectors in AFS and the coordinates of ARP with respect to the satellite center of mass (COM) in the satellite body system (SBS) are usually provided by ground calibrations. In this study, PCO is redefined as an offset from COM to the mean phase center in spaceborne application. Thus, an a priori PCO vector with respect to COM in SBS can be obtained before launching. Three main error sources for the a priori PCO vector should be considered. The first is ground calibration error, especially for COM. To remain consistent with inflight conditions, solar panels will be kept unfolded during ground calibrations. The second source is fuel consumption caused by orbital maneuvers. The third is deformation caused by gravity and in-flight temperature variations. COM coordinates of the TanDEM-X satellite were off by 7 mm after launch. The fuel tank of the satellite will be empty due to regular orbit and formation control maneuvers, thereby causing a shift in the COM of up to 10 cm over the life of the satellite. A regular update of PCO coordinates is therefore required to ensure high accuracy in orbit determination for the TanDEM-X satellite.¹⁴

In practice, the wave front will fluctuate because of the antenna manufacturing characteristics, and the signal reception locations do not coincide. The receiver antenna phase center is the instantaneous location where the GPS signal is actually received, and is dependent on signal frequency and reception direction. Antenna PCVs describe the directiondependent distortions of the wave front, which can be modeled as the distance between the instantaneous location and the mean phase center. An a priori absolute PCV pattern for each frequency can be obtained from ground robot-field calibrations or anechoic chamber calibrations in AFS before launching.¹⁷ However, ground calibrations are performed without or with only limited information related to the in-flight environment and are insufficiently accurate.¹⁶ Systematic effects (e.g., near-field multipath) encountered in the actual antenna environment must be considered. Therefore, in-flight estimation and compensation for antenna PCVs are necessary in high-precision orbit determination applications.

This study focuses on the in-flight calibrations of antenna PCO and PCVs of the Shiyan 3 satellite. In order to avoid

the influence of measurement errors on the location of the signal transmitting reference points from GNSS satellites, transmitting antenna PCO and PCVs of GNSS satellites need to be given first. The IGS provides the antenna information exchange file, i.e. igs08.atx,²⁰ which consists of a set of receivers and GNSS satellite absolute antenna PCO and PCVs. Ground-based absolute calibration results from the anechoic chamber of Shiyan 3 choke ring antenna are available. However, calibrations were performed with limited information related to the satellite environment, and were inconsistent with in-flight calibrations. For validation of general effectiveness and comparison of in-flight calibration results, GRACE data tests were also performed. First, an antenna PCO and PCV estimation model was established. The error characteristics and influence principles of this estimation model on orbit determination were analyzed. The feasibility of PCO and PCV in-flight estimation and compensation in different directions was demonstrated by simulation tests. Subsequently, the values of antenna PCO and PCVs for GRACE and Shiyan 3 satellites were estimated from one month of in-flight data. Finally, orbit improvement is validated through orbit overlap analysis and independent k-band ranging (KBR) and satellite laser ranging (SLR) measurements.

2. Antenna PCO and PCV estimation model

The adopted antenna PCO and PCV model is based on the well-known antenna PCV correction equation¹⁷ where the total antenna phase center correction in the direction of the GPS satellite consists of an absolute mean PCO plus the elevation and azimuth dependent PCVs. Fig. 1 shows Antenna PCO and PCV estimation model. When an a priori PCO vector or a priori PCV pattern in AFS is derived from ground calibrations, the vector or pattern has to be converted to SBS. The relation between AFS and SBS should be determined. For convenience, AFS is defined as follows: the origin is ARP, the positive z-axis coincides with the mechanical symmetry axis and points along the boresight direction, and the x- and y-axes point from ARP into their respective directions depending on the specific mounting of the antennas. For the Shiyan 3 satellite, the three axes of AFS coincide with SBS. For GRACE satellites, the x-axis of AFS coincides with the x-axis of SBS, but the y-axis of AFS is in opposite direction of the y-axis of SBS, and the z-axis yields a right-handed coordinate system. In AFS, the azimuth angle α of a unit



Fig. 1 Antenna PCO and PCV estimation model.

line-of-sight (LOS) vector e is defined as an angle between the projection of e in the *xOy*-plane and positive *x*-axis. This angle is counted in a counter-clockwise direction from the *x*-axis to the *y*-axis. The elevation angle β of vector e is defined as an angle between e and the *xOy*-plane.

The first step involves the estimation of the PCO error components. r_{PCO} is the real PCO vector with respect to COM in SBS; $r_{PCO,0}$ is an a priori PCO vector that can be calculated from ground calibrations; and Δr_{PCO} represents the PCO bias caused by ground calibrations or in-flight variation, i.e. $\Delta r_{PCO} = r_{PCO} - r_{PCO,0}$ (see Fig. 1). Δr_{PCO} makes the a priori PCO vector $r_{PCO,0}$ inaccurate, and will induce an elevation and azimuth dependent distance bias in the antenna PCO correction. The real distance correction caused by PCO is given by

$$\delta\rho_{\rm PCO}(\boldsymbol{e}^{j}) = -(\boldsymbol{e}^{j})^{1}\boldsymbol{M}_{\rm B}^{\rm J2000}(\boldsymbol{r}_{\rm PCO,0} + \Delta\boldsymbol{r}_{\rm PCO}) \tag{1}$$

where e^{j} is a LOS vector from the COM of LEO satellite to the COM of GPS satellite *j* in the inertial reference system J2000; $M_{\rm B}^{12000}$ is the attitude rotation matrix from SBS to J2000; and $\Delta r_{\rm PCO}$ is a systematic bias and can be considered as a constant vector.

In the second step, the elevation and azimuth dependent phase pattern can be modeled and estimated by following the 'residual approach' described by Jäggi et al.^{8,10} The unknown PCVs can be considered as a function $d\varphi(\alpha^i, \beta^i)$. The real distance correction caused by PCVs is

$$\delta \rho_{\text{PCV},i}(\boldsymbol{e}^j) = \lambda_i \mathrm{d}\varphi_i(\boldsymbol{\alpha}^j, \boldsymbol{\beta}^j) \tag{2}$$

where α^{j} and β^{j} represent azimuth angle and elevation angle of e^{j} ; λ_{i} is the carrier wavelength of frequency *i*.

The phase residuals can be obtained by observation minus computation from orbit post-fit model (O-C). Dual-frequency ionosphere-free (IF) combination observations are adopted. For carrier phase observations the ionosphere-free combination yields

$$\begin{aligned} L_{\rm IF}^{j}(t_{i}) &= \frac{f_{1}^{2}}{f_{1}^{2} - f_{2}^{2}} L_{1}^{j}(t_{i}) - \frac{f_{2}^{2}}{f_{1}^{2} - f_{2}^{2}} L_{2}^{j}(t_{i}) \\ &= \rho^{j}(t_{i}) + c(\delta t_{\rm r}(t_{i}) - \delta t^{j}(t_{i})) + \lambda_{\rm IF} A_{\rm IF}^{j} + \delta \rho_{\rm corr}(t_{i}) + \varepsilon^{j}(t_{i}) \end{aligned}$$
(3)

where the subscripts 1 and 2 denote the different frequencies; $L_{\rm IF}$ is the phase ionosphere-free combination; t_i is the signal reception time; f is the carrier frequency; L is the phase observation of frequency f; ρ^j is the geometric distance from the LEO satellite to the GPS satellite; δt_r is the receiver clock error; δt^j is the GPS satellite clock error; c is the speed of light; $\lambda_{\rm IF}A_{\rm IF}$ is the ambiguity of phase ionosphere-free combination; and ε^j contains thermal measurement noise, multipath and all other unmodeled errors. The geometric range

$$\rho^{j}(t_{i}) = \|\mathbf{r}(t_{i}) - \mathbf{r}_{G}^{j}(t_{i} - \tau^{j})\|$$
(4)

is simply given as the distance between the antenna phase center location of the GPS receiver $\mathbf{r}(t_i)$ at signal reception time, and the GPS satellite $\mathbf{r}_G^j(t_i - \tau^j)$ at signal emission time. τ^j is the signal path delay, which can be obtained by iterative calculation. $\delta \rho_{corr}$ represents all kinds of correction,

$$\delta\rho_{\rm corr}(t_i) = -c\delta\rho_{\rm clk}^j(t_i - \tau^j) + \delta\rho_{\rm PCV,G}^j(\boldsymbol{e}^j) + \delta\rho_{\rm rel}(t_i) + \delta\rho_{\rm PCO}(\boldsymbol{e}^j) + \delta\rho_{\rm PCV,IF}(\boldsymbol{e}^j)$$
(5)

where $\delta \rho_{\text{clk}}^{j}(t_{i} - \tau^{j})$ is the GPS clock error correction; $\delta \rho_{\text{PCV,G}}^{j}(e^{j})$ is the transmitter antenna PCV correction of the GPS satellite; $\delta \rho_{\text{rel}}(t_{i})$ is the relativity correction; $\delta \rho_{\text{PCO}}(e^{j})$ is the receiver antenna PCO correction; and $\delta \rho_{\text{PCV,IF}}(e^{j})$ is the receiver antenna PCV correction for ionosphere-free combination, which can be expressed as

$$\delta\rho_{\rm PCV,IF}(\boldsymbol{e}^{j}) = \frac{f_{1}^{2}}{f_{1}^{2} - f_{2}^{2}} \delta\rho_{\rm PCV,1}(\boldsymbol{e}^{j}) - \frac{f_{2}^{2}}{f_{1}^{2} - f_{2}^{2}} \delta\rho_{\rm PCV,2}(\boldsymbol{e}^{j})$$
(6)

where $\delta \rho_{\text{PCV},i}(e^{i})$ is the receiver antenna PCV correction for the phase observation L_{i} .

Assuming that the distance computation by orbit post-fit model without PCV correction in epoch t_i is $z_{IF}^j(t_i)$ and the ionosphere-free phase observation is $L_{IF}^j(t_i)$, then ionosphere-free phase O-C residuals can be expressed as

$$\delta \rho_{\text{PCV,IF}}(\boldsymbol{e}^{j}) + \varepsilon^{j}(t_{i}) = L_{\text{IF}}^{j}(t_{i}) - z_{\text{IF}}^{j}(t_{i})$$
(7)

The existence of $\delta \rho_{\text{PCV,IF}}(e^j)$ will add inconsistency between observation $L_{IF}^{j}(t_{i})$ and computation $z_{IF}^{j}(t_{i})$. Some might be absorbed by other parameters of orbit post-fit model, whereas the remaining will make O-C residuals larger. We can reconstruct the remaining phase pattern iteratively through the O-C residuals. As PCVs depend on the direction of GPS reception signal, i.e., the azimuth and elevation angles of the LOS vector in AFS, the phase pattern can be estimated by azimuth/elevation bins. In this study, the phase O-C residuals are sorted in the azimuth/elevation bins of $\Delta A \times \Delta E$. In the region of $[(n-1)\Delta A, n\Delta A] \times [(m-1)\Delta E, m\Delta E]$ $(n = 1, n\Delta E)$ 2, ..., $360^{\circ}/\Delta A$; $m = 1, 2, ..., 90^{\circ}/\Delta E$), the phase pattern is obtained on the basis of the mean value of all the phase O-C residuals that fall into this region. If observations falling into some regions are insufficient, the phase pattern in that bin can be set to zero directly or constrained to zero with a priori variance. Most of the random noise $\varepsilon^{j}(t_{i})$ will be smoothed out by long-term statistics (one month or longer). The values of ΔA and ΔE are both set as 5° in this study.

According to this model, antenna PCO and PCV estimation has been added to NUDTTK, which is a POD toolkit developed by the National University of Defense Technology. A summary of the dynamic and measurement models used for GPS based LEO satellite orbit determination is given in Table 1.

3. Antenna PCO estimation test

3.1. Simulation I: effect of PCO error on orbit determination

The effect of PCO error in different directions on orbit determination is analyzed in Simulation I. One week dual-frequency GPS observations are first simulated. Three different simulations are performed in Simulation I, in which a 10 cm antenna PCO error is added to the simulated observations in all x, y and z directions, $\Delta \mathbf{r}_{PCO}^{(1)} = [10, 0, 0] \text{ cm}$, $\Delta \mathbf{r}_{PCO}^{(2)} = [0, 10, 0] \text{ cm}$, $\Delta \mathbf{r}_{PCO}^{(3)} = [0, 0, 10] \text{ cm}$. The reference LEO satellite orbit, SIMU-REF, is obtained by the post-fit orbit of the GRACE A satellite science orbit product over a period of 7 days (January 1, 2006 to January 7, 2006). Without receiver antenna PCO correction, GRACE A orbit determination results, SIMU-noPCO, are then obtained through the reduced

Item	Description
GPS measurement model	Undifferenced ionosphere-free code and phase observations; 10 s sampling; igs08.atx PCO and PCVs correction of GPS transmitter antennas; relativity correction; phase wind-up correction; receiver antenna PCO and PCVs estimation; CODE final GPS orbits and 30 s clocks interpolated linearly to 10 s steps
Gravitational forces	Earth gravity, GGM02C 100×100 Solid-earth (IERS Conventions 2010, 4 × 4), pole (IERS Conventions 2010) and ocean tides (FES2004) Luni-solar-planetary gravity, DE405 Relativity, only the Schwarzschild item
Non-gravitational forces	Solar radiation pressure, ball, Cr is estimated Atmospheric Drag, Jacchia 71 density model, Cd is estimated per 3 h Empirical force, piecewise linear spline in T and N directions per 30 min
Reference frames	ITRF2008 reference frame Precession_Nutation, IAU 2000A EOP Parameters, IERS Standard Rapid EOP

 Table 1
 Summary of dynamic and measurement models used for GPS-based LEO satellite orbit determination in NUDTTK.

dynamic approach on the basis of simulated data. Comparing SIMU-noPCO and SIMU-REF, the simulation results for the effect of PCO error in different directions on orbit determination are shown in Table 2. The RMSs of orbital position errors are given in radial (R), transverse (T) and normal (N) directions. Here, a three-axis stabilized attitude is used by the LEO satellite, and the *x*-direction coincides with T-direction, the *y*-direction corresponding to N-direction, the *z*-direction corresponding to R-direction.

The effects of antenna PCO error in different directions on orbit determination are significantly different. The x-direction coincides with satellite flight direction. A 10 cm PCO error in the x-direction, $\Delta r_{PCO}^{(1)}$, can induce a constant orbit offset of the same size in the transverse (T) direction. The average RMSs of radial (R), transverse (T) and normal (N) orbital position errors are 0.010, 10.000 and 0.007 cm, respectively. The y-direction corresponds to the normal orbit direction. A 10 cm PCO error in the y-direction, $\Delta r_{PCO}^{(2)}$, can induce a constant orbit offset of the same size in the N-direction. The average RMSs of R, T and N orbital position errors are 0.005, 0.015 and 10.004 cm, respectively. The z-direction corresponds to the orbit R-direction. A 10 cm PCO error in the z-direction, $\Delta \mathbf{r}_{PCO}^{(3)}$, does not induce a constant orbit offset of the same size in one direction but results in irregular fluctuations in three directions. The average RMSs of R, T and N orbital position errors are 1.234, 1.479 and 1.794 cm, respectively. The average RMS in three dimensions is 2.634 cm, only a quarter of the 10 cm simulation, and some are smoothed out by dynamic orbit fitting.

3.2. Simulation II: feasibility of PCO error estimation

The feasibility of PCO error estimation in different directions is analyzed in Simulation II. Aside from the 10 cm antenna PCO error, random Gaussian white noise is also added to the simulated observations with the standard deviation values of $\sigma_{P1} = \sigma_{P2} = 50$ cm, $\sigma_{L1} = \sigma_{L2} = 0.2$ cm. One week dualfrequency GPS observations are re-simulated, and three simulations are conducted with different PCO error vectors: $\Delta r_{PCO}^{(1)}$, $\Delta r_{PCO}^{(2)}$ and $\Delta r_{PCO}^{(3)}$. According to Eq. (2), antenna PCO error is considered as a constant vector. Parameter estimation in the *x*, *y* and *z*-directions is performed. When one of the directions is estimated, the other two directions are fixed at zero. The daily antenna PCO error estimation results are shown in Table 3. The column '*x*-direction estimation' refers to the simulation with $\Delta r_{PCO}^{(2)}$, the column '*y*-direction estimation' refers to the simulation with $\Delta r_{PCO}^{(2)}$, and the column '*z*-direction estimation' refers to the simulation with $\Delta r_{PCO}^{(2)}$.

The PCO error estimation results in different directions significantly differ. The daily estimation results in the z-direction are stable and almost equal to 10 cm. The average PCO estimation result in the z-direction is 10.01 cm, which is the value nearest to the original 10 cm antenna PCO error among all three directions. The standard deviation of daily estimation

Item	$\Delta \mathbf{r}_{ m PCO}^{(1)} = [10, 0, 0] m cm$			$\Delta \mathbf{r}_{PCO}^{(2)} = [0, 10, 0] \text{ cm}$				$\Delta \mathbf{r}_{PCO}^{(3)} = [0, 0, 10] \text{ cm}$				
	R (cm)	T (cm)	N (cm)	All (cm)	R (cm)	T (cm)	N (cm)	All (cm)	R (cm)	T (cm)	N (cm)	All (cm)
Day 1	0.012	9.991	0.007	9.991	0.005	0.018	10.003	10.003	1.080	1.429	1.879	2.596
Day 2	0.011	10.010	0.007	10.010	0.005	0.011	10.004	10.004	1.069	1.358	1.574	2.338
Day 3	0.011	10.007	0.007	10.007	0.005	0.015	10.004	10.004	1.197	1.478	1.828	2.638
Day 4	0.010	9.995	0.006	9.995	0.005	0.013	10.002	10.002	1.218	1.438	1.794	2.602
Day 5	0.010	9.995	0.007	9.995	0.005	0.028	10.004	10.004	1.276	1.493	1.786	2.654
Day 6	0.010	10.009	0.006	10.009	0.005	0.011	10.005	10.005	1.377	1.566	1.852	2.789
Day 7	0.009	9.990	0.006	9.990	0.005	0.012	10.003	10.003	1.423	1.591	1.845	2.822
Mean	0.010	10.000	0.007	10.000	0.005	0.015	10.004	10.004	1.234	1.479	1.794	2.634

 Table 2
 Simulated effect of 10 cm-size antenna PCO error in different directions on orbit determination.

Table 3 Simula	e 3 Simulation of antenna PCO error estimation in different directions.							
Item	$\Delta \mathbf{r}_{\text{PCO}}^{(1)} = [10, 0, 0] \text{ cm}$ <i>x</i> -direction (cm)	$\Delta \mathbf{r}_{PCO}^{(2)} = [0, 10, 0] \text{ cm}$ y-direction (cm)	$\Delta \mathbf{r}_{\text{PCO}}^{(3)} = [0, 0, 10] \text{ cm}$ z-direction (cm)					
Day 1	1500.6	4.53	10.03					
Day 2	1141.1	15.00	9.99					
Day 3	473.4	-2.71	10.02					
Day 4	312.5	17.29	10.01					
Day 5	347.2	14.35	10.01					
Day 6	172.4	6.32	9.99					
Day 7	-17.7	16.76	9.99					
Mean	561.4	10.22	10.01					
Standard	550.8	7.61	0.016					

results in this direction is only 0.016 cm, which indicates that the PCO estimation accuracy in the z-direction is very high. In NUDTTK software, empirical accelerations in the R-direction are not used. Consequently, the PCO in the z-direction can be estimated without any conditions. If R-direction empirical accelerations are switched on, we must set a condition for the PCO estimate in the z-direction, e.g., the mean value of empirical accelerations in the R-direction is constrained to zero.

The daily estimation results in the y-direction are unstable and not very close to 10 cm. Although the average PCO estimation result in the y-direction is near 10 cm, the standard deviation of daily estimation results is 7.61 cm, which is significantly larger than that in the z-direction. The main reason for this is that the y-direction PCO estimation is coupled with the N-direction empirical force parameters (see Table 1). Estimating these two kinds of parameters together without any conditions will make the solution unstable.

The daily estimation results in the *x*-direction are worse. The standard deviation of daily estimation results reaches 550.8 cm. The estimation results of days 1 and 2 are larger than 1.0 m, which indicates that the *x*-direction PCO estimation without any condition will make the algorithm difficult to converge. This is because the *x*-direction PCO estimation is coupled with T-direction empirical force parameters. Simulation II indicates that the use of GPS data to estimate PCO error precisely is feasible in the *z*-direction, but is unstable in the *y*- and *x*-directions. Hence, we sometimes only estimate *z*-direction PCO error in certain LEO satellite orbit determination applications by GPS.⁸

3.3. Antenna PCO estimation results for GRACE satellites

GRACE²¹ is a satellite gravity field measurement system developed by Germany's Deutsches Zentrum für Luft- und Raumfahrt and America's National Aeronautics and Space Administration. It consists of two identical formation-flying satellites, GRACE A and GRACE B. On March 17, 2002, the twin satellites were launched into an almost circular, near polar orbit with an inclination of 89° and an altitude of 500 km. The distance between the two satellites is approximately 220 km. The GRACE satellite payload includes a KBR system, dual-frequency GPS receiver, laser reflector, accelerometer and star sensor. The KBR, which is mainly for detecting changes in Earth's gravity field, can be used to verify inter-satellite baselines with precision.

The original antenna PCO vectors for GRACE satellites are given in the related SBS published by Kroes.²² Over a period of 31 days (January 1, 2006 to January 31, 2006), GRACE A and GRACE B dual-frequency GPS observations were processed to estimate daily antenna PCO errors in the z-direction with respect to the a priori PCO vector (see Fig. 2). The mean value of daily antenna PCO error estimation results for GRACE A is 0.59 cm, which equals the mean PCO estimation (-40.81 cm) minus the original PCO of GRACE A (-41.40 cm), and the standard deviation is 0.12 cm; the mean value of daily PCO error estimation results for GRACE B is 0.34 cm, which equals the mean PCO estimation (-41.09 cm) minus the original PCO value of GRACE B (-41.43 cm), and the standard deviation is 0.12 cm. The mean values are very small, which means the a priori antenna PCO value for GRACE satellites from ground calibrations is accurate. The standard deviations are also small, indicating that the consistency of daily estimation results is good. Based on GRACE in-flight data tests, it is feasible to use GPS data to estimate PCO error precisely in the z-direction.

3.4. Antenna PCO estimation results for Shiyan 3 satellite

Shiyan 3 is a small satellite mission developed by China and used to test new technologies for exploring space, the atmosphere and its environment.²³ On November 5, 2008, the Shiyan 3 satellite was launched into a sun-synchronous orbit with an inclination of 90° and an altitude of 791 km. The Shiyan 3 satellite payload includes a dual-frequency GPS receiver, laser reflector and star sensor.



Fig. 2 Daily GPS receiver antenna PCO estimation results in *z*-direction for GRACE satellites.



Fig. 3 Daily GPS receiver antenna PCO estimation results in *z*-direction for Shiyan 3 satellite.

The original antenna PCO vector for the Shiyan 3 satellite is given by the satellite manufacturer. Over a period of 31 days (December 1, 2008 to December 31, 2008), Shiyan 3 dualfrequency GPS observations were processed to estimate daily antenna PCO errors in the z-direction with respect to the a priori PCO vector (see Fig. 3). The mean value of daily antenna PCO error estimation results is -10.34 cm, and the standard deviation is 0.21 cm. A large mean value of 10 cm level is found, which reflects the inaccuracy of the a priori antenna PCO value from ground calibrations. However, the standard deviation is small, which means that the consistency of daily estimation results is good.

4. Antenna PCV estimation test

4.1. Simulation III: feasibility of PCV estimation

The feasibility of PCV estimation in different directions is analyzed in Simulation III. First, an antenna phase pattern is selected as a reference. The effects of this phase pattern on distance are added into the simulated observations according to Eq. (2). The phase pattern of SEN67-1575-14 + CRG antenna from ground calibrations by automated absolute field approach^{8,24} is selected as a reference, while SEN67-1575-14 + CRG is a typical choke ring antenna, and has been successfully used in TanDEM-X mission for dual-frequency GPS signal reception, see Fig. 4. Fig. 4(a) shows that the phase pattern of the ionosphere-free combination is almost perfectly symmetric with respect to the boresight axis. The phase pattern has the obvious character of systematic deviation. The maximum value for the mean PCVs on ionosphere-free combination can reach 1.5 cm.

Subsequently, without receiver antenna PCV correction, GRACE A orbit determination results, SIMU-noPCV, are obtained by a reduced dynamic approach using simulated observations. Comparing SIMU-noPCV and SIMU-REF, the effect of the reference phase pattern on orbital position in three directions is shown in Fig. 5(a). The average RMSs of R, T and N orbital position errors are 2.6, 4.6 and 3.0 mm, respectively. The average RMS in three dimensions is 6.1 mm.

After three iterations, the phase pattern estimation results are obtained with $5^{\circ} \times 5^{\circ}$ resolution by residual approach^{6,10} (see Fig. 4(b)). The shape of the pattern corresponds well to



Fig. 4 Phase pattern of SEN67-1575-14+CRG antenna in TanDEM-X mission.

the ground calibrated PCVs shown in Fig. 4(a). Distortions of both phase patterns are similar. Most of the antenna PCVs can be effectively estimated except for *y*-direction, for which parts are absorbed by carrier-phase ambiguities and receiver clock parameters.

Finally, with receiver antenna PCV correction from estimation results, GRACE A orbit determination results, SIMU-PCV, are obtained using simulated observations. Comparing SIMU-PCV and SIMU-REF, the improvement of orbital position accuracy by PCV estimation is analyzed (see Fig. 5(b)). The average RMSs of R, T and N position errors are 0.7, 1.3 and 1.3 mm, respectively. The average RMS in three dimensions is 1.9 mm. Compared with SIMU-noPCV (Fig. 5 (a)), position accuracy is significantly improved by PCV estimation. The average percentage of R, T and N position improvements are 73%, 71% and 57%, respectively. The average percent in three dimensions is 69%. By Simulation III, the improvement of orbital position accuracy by PCV estimation is obvious. Therefore, in-flight PCV calibration and compensation are necessary in high precision orbit determination applications.



Fig. 5 Influence of PCV estimation of SEN67-1575-14+CRG antenna on orbit accuracy.

4.2. Antenna PCV estimation results for GRACE satellites

First, new PCO vectors after in-flight calibration and compensation for GRACE satellites can be obtained by adding the mean value of daily PCO error estimation results into the original antenna PCO vectors (see Table 4). With new PCO correction and without PCV correction, GRACE satellite orbit determination results, GRACEA-PCO-noPCV and GRACEB-PCO-noPCV, are obtained.

Subsequently, after three iterations, the ionosphere-free phase pattern estimation results are given in Fig. 6. There is no a priori PCV from ground calibrations used as an initial reference for PCV iterative estimation. The phase patterns have the obvious character of systematic deviation, closely matching the phase patterns based on 365 days of data set in 2007 as previously obtained by Jäggi et al.⁸ though with a different data set and software. The phase pattern distortions from NUDTTK software with $5^{\circ} \times 5^{\circ}$ resolution look a little different from those of Bernese software with $1^{\circ} \times 1^{\circ}$ resolution. These differences in PCV estimates might be caused by insuf-



Fig. 6 Ionosphere-free phase pattern estimation results for GRACE satellites.

ficient observation in some regions, such as low elevation angles. Observation data numbers in low-elevation bins are usually small, and variances of phase O-C residuals in lowelevation bins are usually large.

Finally, with the estimated PCO (see Table 4) and the ionosphere-free phase pattern estimation results (see Fig. 6), new PCV corrections can be performed again according Eq. (5), and GRACE satellite orbit determination results, GRACEA-PCO-PCV and GRACEB-PCO-PCV, are obtained. The influence of PCV estimation on the orbital position of GRACE satellites is given in Fig. 7.

Table 4 Antenna PCO for GRACE and Shiyan 3 satellites.								
Satellite	Original PCO ²²	(cm)		Estimated PCO	Estimated PCO (cm)			
	x-direction	y-direction	z-direction	x-direction	y-direction	z-direction		
GRACE A	0.04	-0.04	-41.40	0.04	-0.04	-40.81		
GRACE B	0.06	-0.08	-41.43	0.06	-0.08	-41.09		
Shiyan 3	-1.60	-1.60	-51.46	-1.60	-1.60	-61.80		



Fig. 7 Influence of PCV estimation on orbital position for GRACE satellites.

Comparing GRACEA-PCO-noPCV and GRACEA-PCO-PCV, the average RMSs of R, T and N position differences are 1.6, 3.1 and 3.5 mm, respectively. The RMS in three dimensions is 4.9 mm (see Fig. 7(a)). Comparing GRACEB-PCO-noPCV and GRACEB-PCO-PCV, the average RMSs of R, T and N position differences are 2.1, 4.1 and 9.2 mm, respectively. The RMS in three dimensions is 10.3 mm (see Fig. 7(b)).

4.3. Antenna PCV estimation results for Shiyan 3 satellite

First, new PCO vectors after in-flight calibration and compensation for the Shiyan 3 satellite can be obtained by adding the mean value of daily PCO error estimation results into the original antenna PCO vectors (see Table 4). With new PCO correction and without PCV correction, Shiyan 3 satellite orbit determination results, Shiyan3-PCO-noPCV, are obtained. After three iterations, the ionosphere-free phase pattern estimation result is then given in Fig. 8. The phase pattern has the obvious character of systematic deviation. The maximum value for the mean PCVs can reach 3.0 cm.

Finally, with PCV correction from the phase pattern estimation result, the Shiyan 3 satellite orbit determination result, Shiyan3-PCO-PCV, is obtained. The influence of PCV estimation on the orbital position for the Shiyan 3 satellite is given in Fig. 9. Comparing Shiyan3-PCO-noPCV and Shiyan3-PCO-PCV, the average RMSs of R, T and N position differences are 2.8, 7.9 and 5.4 mm, respectively. The RMS in three dimensions is 10 mm. The average difference of three dimensions is 10 mm.



Fig. 8 Ionosphere-free phase pattern estimation results for Shiyan 3 satellite.



Fig. 9 Influence of PCV estimation on orbital position for Shiyan 3 satellite.

5. Orbit validation

5.1. Orbit validation with KBR data

The KBR system is one of the key scientific instruments onboard GRACE satellites. It measures the one-way range change between the two GRACE satellites with the precision of approximately 10 μ m at 5 s data intervals. Given the high precision of KBR data, the relative position accuracy of GRACE satellites can be validated.^{13,22}

The relative positions computed by GRACE-noPCOnoPCV, GRACE-PCO-PCV and JPL precise science orbits are validated by KBR data (see Fig. 10). The selected product type of JPL precise science orbits is GNV1B-01, which is produced by JPL's GIPSY-OASIS software.²⁵ For GIPSY-OASIS software, GPS 5-min clock solutions are used, and GRACE data are processed with 5-min samples rather 10 s. Empirical accelerations for unknown perturbations compensated in R, T and N directions are modeled as supplementary processs noise parameters with 30 min time correlation updated every 5-min in the extended Kalman filter.

The average standard deviations of KBR comparison residuals are 10.1, 8.8 and 14.2 mm. Compared to the relative positions obtained without PCO and PCV estimation, the average



Fig. 10 Daily standard deviations of KBR validation for GRACE satellites.

standard deviation of the relative positions obtained with PCO and PCV estimation is improved by 1.3 mm and is better than JPL precise science orbits. Considering that antenna PCO and PCV estimation could remove the phase errors of both GRACE satellites, the relative orbital position accuracy may be improved.

5.2. Orbit validation with SLR data

Another independent way for orbit validation is offered through SLR data. Statistics over the entire SLR data arcs for GRACE and Shiyan 3 satellites are displayed in Table 5. Typical results of an SLR comparison for orbit products with PCO and PCV estimation are shown in Fig. 11. The mean value and the RMS of the SLR comparison residuals for each arc are calculated and used to compute statistics. At the International Laser Ranging Service (ILRS), no information is available for the Shiyan 3 satellite at present. Only the Chinese SLR stations (including Shanghai, Changchun, Beijing and Kunming) track this satellite.

Data sets used here are the same as those used in the PCV estimation. The results reveal that for the two GRACE satellites, the average RMS of orbit products obtained with PCO and PCV estimation is close to the JPL orbit products and better than that of orbit products obtained without PCO and PCV estimation. For the Shiyan 3 satellite, the average RMS of the orbit products obtained with PCO and PCV estimation is improved by 0.9 cm compared with the orbit products obtained without PCO and PCV estimation. Most of these

Table 5SLR validation for different types of GRACE andShiyan 3 orbit products.

Type of orbit	Counts of SLR arc	Mean (cm)	RMS (cm)
GRACEA_noPCO_noPCV	169	-0.37	1.71
GRACEA-PCO-PCV	169	-0.35	1.64
GRACEA-JPL	169	-0.02	1.79
GRACEB_noPCO_noPCV	154	-0.17	1.73
GRACEB-PCO-PCV	154	-0.16	1.67
GRACEB-JPL	154	-0.27	1.62
Shiyan3_noPCO_noPCV	41	-2.89	4.95
Shiyan3-PCO-PCV	41	-2.67	4.06



Fig. 11 SLR comparison statistics for orbit products with PCO and PCV estimation of GRACE and Shiyan 3 satellites.

contributions might come from the finding of a large PCO error value. In addition, the mean value of the Shiyan 3 satellite is significantly larger than that of GRACE satellites (see Table 5), which might be caused by the inaccurate a priori SLR antenna offset vector from COM, e.g., COM error, because GPS antenna PCO vector from COM of the Shiyan 3 satellite is also inaccurate.

6. Conclusions

Antenna PCO and PCV error characteristics and influence principles on orbit determination are analyzed. The feasibility of PCO and PCV estimation and compensation in different directions is demonstrated by simulations and in-flight tests. Antenna PCO precise error estimation is feasible in the z-direction but is unstable in y- and x-directions. Most of the antenna PCVs can be effectively estimated, except for the y-direction.

Ground-based absolute calibration results from the anechoic chamber for Shiyan 3 choke ring antenna are not consistent with in-flight calibrations. A large and stable antenna PCO error in the z-direction is found on the Shiyan 3 satellite, of which the magnitude is up to 10.34 cm. The PCVs on the Shiyan 3 satellite are estimated, of which the magnitude is up to 3.0 cm, larger than those for GRACE satellites.

The significant improvement of orbits through PCO and PCV estimation and compensation can be confirmed by orbit validation with independent KBR and SLR measurements. For GRACE satellites, the average RMS of KBR residuals is improved from 1.01 cm to 0.88 cm and is better than JPL precise science orbits. For the Shiyan 3 satellite, the average RMS of SLR residuals is improved from 4.95 cm to 4.06 cm. The mean value is large, which might be caused by the inaccurate a priori COM.

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