Physics Letters B 753 (2016) 161-165

Contents lists available at ScienceDirect

Physics Letters B

www.elsevier.com/locate/physletb

Probing models of neutrino masses via the flavor structure of the mass matrix



Shinya Kanemura, Hiroaki Sugiyama*

Department of Physics, University of Toyama, 3190 Gofuku, Toyama 930-8555, Japan

ARTICLE INFO

Article history: Received 29 October 2015 Received in revised form 30 November 2015 Accepted 6 December 2015 Available online 10 December 2015 Editor: J. Hisano

ABSTRACT

We discuss what kinds of combinations of Yukawa interactions can generate the Majorana neutrino mass matrix. We concentrate on the flavor structure of the neutrino mass matrix because it does not depend on details of the models except for Yukawa interactions while determination of the overall scale of the mass matrix requires to specify also the scalar potential and masses of new particles. Thus, models to generate Majorana neutrino mass matrix can be efficiently classified according to the combination of Yukawa interactions. We first investigate the case where Yukawa interactions with only leptons are utilized. Next, we consider the case with Yukawa interactions between leptons and gauge singlet fermions, which have the odd parity under the unbroken Z_2 symmetry. We show that combinations of Yukawa interactions for these cases can be classified into only three groups. Our classification would be useful for the efficient discrimination of models via experimental tests for not each model but just three groups of models. © 2015 The Authors. Published by Elsevier B.V. This is an open access article under the CC BY license

(http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.

1. Introduction

Thanks to the discovery of a Higgs boson h at the CERN Large Hadron Collider (LHC) [1], we have entered the era to explore the origin of particle masses. Coupling constants of W^{\pm} , Z, t, b, and τ with *h* are measured at the LHC [2], and they are consistent with predicted values in the Standard Model (SM). These results strongly suggest that masses of gauge bosons and charged fermions are generated by the vacuum expectation value of the Higgs field, which provides h, as predicted in the SM. Thus, the mechanism to generate their masses in the SM was confirmed. On the other hand, neutrino masses are not included in the SM although neutrino oscillation data uncovered that neutrinos have their masses [3,4]. It is easy to add neutrino mass terms $m_v \overline{\nu_I} \nu_R$ to the SM similarly to the other fermion mass terms by introducing right-handed neutrinos v_R . However, since the neutrino is a neutral fermion in contrast to the other fermions in the SM, another possibility of its mass term exists. That is the Majorana mass term, $(1/2)m_{\nu}\overline{\nu_{L}}(\nu_{L})^{c}$. This unique possibility could be the reason why neutrinos are much lighter than the other fermions. New physics models for the Majorana neutrino mass can be found in e.g. Refs. [5-57].

The overall scale of the neutrino mass matrix m_{ν} generated in new physics models is determined by the structure (tree level, one-loop level, and so on) of the diagram to generate m_{ν} , masses of new particles in the diagram and coupling constants in the diagram. This means that the determination of the overall scale of m_{ν} requires to specify many parts of the Lagrangian of each model. On the other hand, the flavor structure (ratios of elements) of m_{ν} is simply determined by the product of Yukawa coupling matrices and fermion masses. Thus, models to generate m_{ν} can efficiently be classified according to the combination of Yukawa coupling matrices and fermion masses without the detail of these models. When we construct a new model to generate neutrino masses, it will be noticed indeed that the flavor structure is the key to find an appropriate set of model parameters although the overall scale of m_{ν} can be easily tuned by using some parameters in the scalar potential.

In this letter, we first classify models for Majorana neutrino masses according to combination of Yukawa interaction between leptons without introducing new fermions. Next, we do the classification for the case where gauge singlet fermions are introduced such that they have the odd parity under the unbroken Z_2 symmetry which can be utilized to stabilize the dark matter. For Yukawa interactions of these new fermions with leptons, Z_2 -odd scalars are also introduced. We find that models can be classified into only three groups. The classification could be useful to approach efficiently the origin of Majorana neutrino masses with experimental tests of not each model but each group of models.

E-mail addresses: kanemu@sci.u-toyama.ac.jp (S. Kanemura), sugiyama@sci.u-toyama.ac.jp (H. Sugiyama).

http://dx.doi.org/10.1016/j.physletb.2015.12.012

0370-2693/© 2015 The Authors. Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.



^{*} Corresponding author.

Table 1

Scalar bosons which can have Yukawa interactions with leptons without introducing new fermions. The Yukawa matrix Y_A is antisymmetric, while Y_S^s and Y_S^{Δ} are symmetric. The lepton number (L#) is assigned to each of scalar fields such that the Yukawa interactions conserve the L# as a convention. Then, the L# is broken in the scalar potential.

Scalar	SU(2) _L	U(1) _Y	L#	Yukawa	Note
s_{1}^{+}	<u>1</u>	1	-2	$(Y_A^s)_{\ell\ell'} \left[\overline{L_\ell} \in L_{\ell'}^c s_1^- \right]$	Antisymmetric
s ⁺⁺	<u>1</u>	2	-2	$(Y_S^s)_{\ell\ell'} \left[\overline{(\ell_R)^c} \ell'_R s^{++} \right]$	Symmetric
$\Phi_2 = \begin{pmatrix} \phi_2^+ \\ \phi_2^0 \end{pmatrix}$	<u>2</u>	$\frac{1}{2}$	0	$\mathbf{y}_{\ell} \Big[\overline{L_{\ell}} \Phi_2 \ell_R \Big]$	Diagonal
$\Delta = \begin{pmatrix} \frac{\Delta^+}{\sqrt{2}} & \Delta^{++} \\ \Delta^0 & -\frac{\Delta^+}{\sqrt{2}} \end{pmatrix}$	<u>3</u>	1	-2	$(\mathbf{Y}^{\Delta}_{S})_{\ell\ell'} \left[\overline{L_{\ell}} \Delta^{\dagger} \boldsymbol{\epsilon} L^{c}_{\ell'} \right]$	Symmetric

Models of neutrino masses can also be classified according to topologies of diagrams [58] or decompositions of higher massdimensional operators [59]. They seem useful to find new models and increase the number of models in order to exhaust all possibilities. In contrast with these classifications, ours would be useful to simplify the situation where many models exist.

2. Classification of flavor structure

First, we introduce only scalar fields listed in Table 1, which have Yukawa interactions with leptons. We do not always introduce all of them, and we utilize only scalar bosons for required Yukawa interactions. For the Yukawa interaction with the second $SU(2)_L$ -doublet scalar field Φ_2 , the flavor changing neutral current is forbidden by utilizing a softly-broken Z₂ symmetry as usually done in the two Higgs doublet models. In order to obtain m_{ν} , we try to connect ν_{I} to $(\nu_{I})^{c}$ by using these Yukawa interactions and the weak interaction. We do not care how scalar lines are closed because we concentrate on the flavor structure of m_{ν} . Each charged lepton $(\ell_L, \ell_R, (\ell_L)^c, (\ell_R)^c)$ should appear only once on the fermion line in order to obtain the simplest combinations, which would give the largest contribution to $m_{\rm p}$. In addition, ℓ_L and ℓ_R should appear only in the next to each other on the fermion line. If they do not, the replacement of the structure between them with the mass term of ℓ can give the simpler combination.¹ It is assumed that m_{ν} is generated via a solo mechanism (a solo kind of fermion lines). Then, we find that only the following five combinations² connect v_L and $(v_L)^c$:

$$m_{\nu} \propto Y^{s}_{A} y_{\ell} Y^{s}_{S} y_{\ell} (Y^{s}_{A})^{T}, \qquad (1)$$

$$m_{\nu} \propto y_{\ell} \left(Y_{S}^{s} \right)^{*} y_{\ell}, \tag{2}$$

$$m_{\nu} \propto g_2 \, y_{\ell} \, (Y_S^s)^* \, y_{\ell} \, g_2, \tag{3}$$

$$m_{\nu} \propto Y_{S}^{\Delta},$$
 (4)

$$m_{\nu} \propto Y_A^s y_{\ell}^2 + (Y_A^s y_{\ell}^2)^T,$$
 (5)

where Yukawa matrices Y_A , Y_S^s , y_ℓ , and Y_S^{Δ} are defined in Table 1. Diagrams of fermion lines for combinations in eqs. (1)–(5) are shown in Figs. 1–5, respectively. The SU(2)_L gauge coupling constant g_2 is shown for clarity although the weak interaction is flavor blind. The combination in eq. (3) gives at least a dimension-9 operator for the Majorana neutrino mass while the others can be a dimension-5 one.

The combination in eq. (5) is the one in the Zee–Wolfenstein model [5,6] of the Majorana neutrino mass at the one-loop level,

Fig. 1. The diagram of the fermion line for the combination in eq. (1).

$$\nu_L \stackrel{\phi_2^{+} = s^{--} = \phi_2^{+}}{\underbrace{\begin{array}{c} \mathbf{I} & \mathbf{I} \\ -\mathbf{I} & \ell_R & \mathbf{I} & (\ell_R)^c & \mathbf{I} \\ y_{\ell} & (Y_S^{s)*} & y_{\ell} \end{array}}_{(Y_S^{s)*} (\nu_L)^c} (\nu_L)^c$$

Fig. 2. The diagram of the fermion line for the combination in eq. (2).

$$\nu_L - \frac{ \begin{array}{ccc} W^+ & \langle \phi^0 \rangle & s^{--} & \langle \phi^0 \rangle & W^+ \\ \end{array}}{g_2 & \ell_L & \frac{1}{2} & \ell_R & \frac{1}{2} (\ell_R)^c & \frac{1}{2} (\ell_L)^c \overset{}{\underset{\scriptstyle \leq}{\underset{\scriptstyle \leq}{\atop\scriptstyle S}}} & (\nu_L)^c \end{array}}{g_2 & (\nu_L)^c & \frac{1}{2} (\ell_R)^c & \frac{1}{2} (\ell_R)$$

Fig. 3. The diagram of the fermion line for the combination in eq. (3).

$$Y_L = \frac{ \begin{pmatrix} \Delta^0 \\ \mathbf{I} \end{pmatrix}}{Y_S^{\Delta}} (\nu_L)^c$$

Fig. 4. The diagram of the fermion line for the combination in eq. (4).

Fig. 5. The diagram of the fermion line for the combination in eq. (5).

which has been excluded already by the neutrino oscillation data [60]. Thus, this combination is ignored below. An example for m_{ν} in eq. (1) is the Zee–Babu (ZB) model [7,8], which generates m_{ν} at the two-loop level. The structure in eq. (2) is given in a model in Ref. [9] by Cheng and Li (the CL model), which also generates m_{ν} at the two-loop level.³ The Gustafsson–No–Rivera (GNR) model [10] is an example for the combination in eq. (3), in which m_{ν} is generated at the tree-loop level. Scalar lines of W^+ and s^{--} are connected at the one-loop level by introducing the unbroken Z_2 symmetry and Z_2 -odd scalar fields, which provide a dark matter candidate. The structure in eq. (4) is given at the tree level, and an example is the Higgs triplet model (HTM) [9,11]. Since eqs. (2) and (3) have the same flavor structure, that of m_{ν} is given

¹ Although the electron Yukawa coupling is small, the diagonal matrix y_{ℓ} would not be negligible because of the tau Yukawa coupling.

² Notice that another possible combination $Y_A^s g_2 + (Y_A^s g_2)^T$ becomes zero.

³ In Ref. [9], scalar lines of ϕ_2^+ and s^{--} are closed in a little bit complicated way. Instead of that, it seems the simplest to introduce an SU(2)_L-doublet scalar field with the hypercharge Y = 3/2.

Table 2

Scalar bosons for Yukawa interactions of gauge singlet fermion ψ_{iR}^0 with leptons. These scalar bosons and ψ_{iR}^0 are Z_2 -odd fields. Structures of Yukawa matrices Y^s and Y^η are arbitrary. When ψ_R^0 has $L^{\#} = x$, lepton numbers -x - 1 and x - 1 are assigned to s_2^+ and η , respectively, such that their Yukawa interactions conserve the L# as a convention. The L# is broken in the scalar potential and/or M_{ψ} .



Fig. 6. The diagram of the fermion line for the combination in eq. (6). Bold red lines are for the Z_2 -odd particles.

$$\nu_L \stackrel{\phi_2^+}{\underbrace{ \begin{array}{c} & s_2^- & s_2^- & \phi_2^+ \\ & \mathbf{l} & \ell_R & (\psi_R^0)^c & \psi_R^0 & (\ell_R)^c & \mathbf{l} \\ & y_\ell & (Y^s)^* & \mathcal{M}_{\ell_*} & (Y^s)^\dagger & y_\ell \end{array}}_{} (\nu_L)^c$$

Fig. 7. The diagram of the fermion line for the combination in eq. (7). Bold red lines are for the Z_2 -odd particles.

$$\nu_L \stackrel{W^+}{\longrightarrow} \begin{array}{c} \langle \phi^0 \rangle & s_2^- & s_2^- & \langle \phi^0 \rangle & W^+ \\ \hline & & & \\ g_2 & \ell_L & \frac{1}{\ell_R} & \ell_R & (\psi_R^0)^c & \psi_R^0 & (\ell_R)^c & \frac{1}{\ell} & (\ell_L)^c \overset{}{\underset{\scriptstyle \leftarrow}{\underset{\scriptstyle \leftarrow}{\atop\scriptstyle \leftarrow}}} \\ \hline & & & \\ M_{\pm} & & (Y^s)^\dagger & y_\ell & g_2 \end{array} (\nu_L)^c$$

Fig. 8. The diagram of the fermion line for the combination in eq. (8). Bold red lines are for the Z_2 -odd particles.

by only three combinations of Yukawa matrices: $Y_A^s y_\ell Y_S^s y_\ell (Y_A^s)^T$, $y_\ell (Y_S^s)^* y_\ell$, and Y_S^{Δ} .

Next, we impose the unbroken Z_2 symmetry to models and introduce gauge singlet fermions ψ_{iR}^0 as the Z_2 -odd fields. The fermions have Majorana mass terms, $(1/2)M_{\psi i}(\psi_{iR}^0)^c\psi_{iR}^0$. We can take the basis where M_{ψ} is diagonalized without loss of generality. For Yukawa interactions of ψ_{iR}^0 with leptons, scalar fields in Table 2 are also introduced as Z_2 -odd fields. Scalar fields in Table 1 and the SM fields are Z_2 -even ones. Then, the lightest Z_2 -odd particle becomes stable. If the lightest Z_2 -odd particle is neutral one, it can be a dark matter candidate. We find that the Majorana neutrino mass matrix can be obtained by the following four kinds of combinations of Yukawa matrices and the weak interaction in addition to the five combinations in eqs. (1)–(5):

$$m_{\nu} \propto Y_{A}^{s} y_{\ell} Y^{s} M_{\nu}^{-1} (Y^{s})^{T} y_{\ell} (Y_{A}^{s})^{T},$$
(6)

$$m_{\rm v} \propto \gamma_{\ell} \, (Y^{\rm s})^* \, M_{\rm v}^{-1} \, (Y^{\rm s})^\dagger \, \gamma_{\ell}, \tag{7}$$

$$m_{\nu} \propto g_2 \, y_{\ell} \, (Y^s)^* \, M_{\nu}^{-1} \, (Y^s)^{\dagger} \, y_{\ell} \, g_2, \tag{8}$$

$$m_{\nu} \propto Y^{\eta} M_{\mu}^{-1} (Y^{\eta})^{T},$$
 (9)

where Yukawa matrices Y^s and Y^η are defined in Table 2. Figs. 6–9 correspond to diagrams of fermion lines for combinations in eqs. (6)–(9), respectively. The part M_{ψ}^{-1} is given by assuming ψ_{iR}^0 are heavier than the other particles. If it is not the case, M_{ψ}^{-1} can be replaced with M_{ψ} .

The Krauss–Nasri–Trodden (KNT) model [12] of m_{ν} at the threeloop level is an example for the combination in eq. (6). The struc-

$$\nu_L \xrightarrow[Y^{\eta}]{} \frac{\psi_R^0}{M_{\psi}} \underbrace{(\psi_R^0)^c}_{(Y^{\eta})^T} (\nu_L)^c$$

Fig. 9. The diagram of the fermion line for the combination in eq. (9). Bold red lines are for the Z_2 -odd particles.

ture in eq. (7) is realized, for example, in the Aoki–Kanemura–Seto (AKS) model [13] at the three-loop level by introducing the Z_2 -odd real singlet scalar boson. Since the three-loop diagram utilizes the scalar interaction with two Higgs doublet fields, the AKS model can explain not only m_{ν} and the dark matter but also the baryon asymmetry of the universe via the electroweak baryogenesis scenario. An example of the combination in eq. (9) is the Ma model [14], where m_{ν} is generated at the one-loop level. No model is known for m_{ν} in eq. (8).⁴ Flavor structures of combinations in eqs. (7) and (8) are the same because the weak interaction does not change the flavor. Therefore, the flavor structure of m_{ν} is determined by three combinations when we use the Yukawa interactions in Table 2: $Y_A^s y_\ell Y^s M_{\psi}^{-1} (Y^s)^T y_\ell (Y_A^s)^T$, $y_\ell (Y^s)^* M_{\psi}^{-1} (Y^s)^{\dagger} y_\ell$, and $Y^{\eta} M_{\psi}^{-1} (Y^{\eta})^T$.

It is clear that combinations in eqs. (1)-(4) and eqs. (6)-(9) can be classified further to only the following three groups:

Group-I:
$$m_{\nu} \propto Y_A^s y_{\ell} X_{SR} y_{\ell} (Y_A^s)^T$$
, (10)

Group-II:
$$m_{\nu} \propto y_{\ell} X_{SR}^* y_{\ell}$$
, (11)

Group-III:
$$m_{\nu} \propto X_{SL}$$
, (12)

where symmetric matrices X_{SR} and X_{SL} are given by

$$X_{SR} = Y_S^s, \quad Y^s M_{\psi}^{-1} (Y^s)^T, \quad Y^s M_{\psi} (Y^s)^T,$$
(13)

$$X_{SL} = Y_S^{\Delta}, \quad Y^{\eta} M_{\psi}^{-1} (Y^{\eta})^T, \quad Y^{\eta} M_{\psi} (Y^{\eta})^T.$$
(14)

The matrix X_{SR} is for the effective interactions of right-handed charged leptons while the matrix X_{SL} is for the ones of left-handed leptons. As long as we concentrate on the flavor structure, it seems difficult to discriminate the origin of X_{SR} (X_{SL}) in eq. (13) (eq. (14)).

We mention here the type-I [15] and the type-III seesaw [16] models, where gauge singlet fermions (for the type-I) or SU(2)_L-triplet Majorana fermions (for the type-III) are introduced. The structure of m_{ν} in these models can be included in the Group-III because Yukawa matrices Y_A and y_{ℓ} are not used to generate m_{ν} . However, they are exceptions because new scalar fields are not introduced. Discussion in the next section (namely, $\tau \rightarrow \bar{\ell}_1 \ell_2 \ell_3$ ($\ell_1, \ell_2, \ell_3 = e, \mu$) for the Group-III) is not applicable for these models.⁵

3. Discussion

The neutrino mass matrix m_{ν} is expressed as U_{MNS}^* diag $(m_1 e^{i\alpha_{12}}, m_2, m_3 e^{i\alpha_{32}})U_{\text{MNS}}^{\dagger}$, where m_i (i = 1-3) are the neutrino mass eigenvalues, α_{12} and α_{32} are the Majorana phases [61], and U_{MNS} is the Maki–Nakagawa–Sakata (MNS) matrix [62] of the lepton flavor mixing. The Group-I gives $m_1 = 0$ or $m_3 = 0$

⁴ The combination in eq. (8) gives at the least a dimension-9 operator for m_{ν} , and it might be four-loop realization at the least. Then, too small neutrino masses might be generated.

⁵ There is the box diagram with the *W* boson and neutral fermions from $SU(2)_L$ -singlet or triplet, but the interaction of the neutral fermions with *W* is suppressed by $\sqrt{m_v/M_R}$ (the mixing between v_L and the fermions), where M_R denotes the fermion mass.

because of $\text{Det}(m_{\nu}) \propto \text{Det}(Y_A) = 0$. Although this has been known for the Zee–Babu model [8] (an example of models in the Group-I), our statement is more model-independent. The Group-I is excluded if the absolute neutrino mass is directly measured at the KATRIN experiment [63] whose estimated sensitivity is 0.35 eV at 5 σ confidence level. The indirect bound on the sum of neutrino masses, $\sum_i m_i < 0.23$ eV (90% confidence level), was obtained by cosmological observations [64], and sensitivity to $\sum_i m_i = \mathcal{O}(0.01)$ eV is expected in future experiments [65].

The flavor structure of m_{ν} is constrained by the neutrino oscillation data, and the constrained structure can be translated into constraints on the flavor structure (ratios of elements) of X_{SR} of the Group-II and X_{SL} of the Group-III. Hereafter, we denote X_{SR} of the Group-II and X_{SL} of the Group-III as X for simplicity. These interactions can cause the lepton flavor violating (LFV) decays $\tau \rightarrow \overline{\ell}_1 \ell_2 \ell_3$ $(\ell_1, \ell_2, \ell_3 = e, \mu)$. Ratios of the decay branching ratios (BR) of these LFV decays can be determined by the flavor structure of X independently on the overall scale of m_{ν} . In order to evade the strong constraint BR($\mu \rightarrow \overline{e}ee$) < 1.0 × 10⁻¹² [66], LFV decays $\tau \rightarrow \overline{\ell}_1 \ell_2 \ell_3$ can be observed at the Belle II experiment [67] only for $X_{ee} = 0$ or $X_{e\mu} = 0$, which constrains ratios of BR($\tau \rightarrow \overline{\ell}_1 \ell_2 \ell_3$) as discussed in the HTM (included in the Group-III) [68]. For $X_{ee} = 0$ ($X_{e\mu} = 0$), LFV decays $\tau \to \overline{\ell} ee$ ($\tau \to \overline{\ell} e\mu$) do not occur. Since $X_{e\ell}$ elements for the Group-II are enhanced by $1/m_e$ for a given m_{ν} , it is likely that BR $(\tau \rightarrow \overline{e}e\mu)$ for $X_{ee} = 0$ or BR $(\tau \rightarrow \overline{e}ee)$ for $X_{e\mu} = 0$ is larger than the others. For $X_{ee} = X_{e\mu} = 0$, only $au
ightarrow \overline{e} \mu \mu$ can be observed for the Group-II as shown in the GNR model [10], while $\tau \rightarrow \overline{\mu}\mu\mu$ is also possible for the Group-III. Notice that $X_{ee} = 0$ for the Group-II and III results in $(m_v)_{ee} = 0$, which is excluded if the neutrinoless double beta decay (see e.g. Ref. [69]) is observed or $m_3 < m_1$ (the inverted mass ordering of neutrinos) is determined by neutrino oscillation experiments (see e.g. Ref. [70]). Notice also that $(X_{SR})_{ee} = 0$ for the Group-I does not mean $(m_{\nu})_{ee} = 0$. Therefore, if $(m_{\nu})_{ee} = 0$ is excluded by these neutrino experiments, the observation of $\tau \rightarrow \overline{\ell} e e$ indicates the Group-I because the situation is inconsistent for the Group-II and III.

The discussion above did not require the discovery of new particles. If a charged scalar boson is discovered and dominantly decays into leptons, the branching ratios are expected to be given by Y_A (y_ℓ) when the Group-I (II) is assumed. The flavor structure of y_ℓ is known, and decays via the y_ℓ are dominated by the decay into τ . The flavor structure of Y_A is determined by the neutrino oscillation data as $(Y_A)_{e\mu}/(Y_A)_{e\tau} = -(U_{MNS})^*_{\tau 1}/(U_{MNS})^*_{\mu 1}$ and $(Y_A)_{\mu\tau}/(Y_A)_{e\tau} = -(U_{MNS})_{e1}/(U_{MNS})^*_{\mu 1}$ for $m_1 < m_3$. For $m_1 > m_3$, they are given by $(Y_A)_{e\mu}/(Y_A)_{e\tau} = -(U_{MNS})_{\tau 3}/(U_{MNS})_{\mu 3}$ and $(Y_A)_{\mu\tau}/(Y_A)_{e\tau} = -(U_{MNS})^*_{e3}/(U_{MNS})^*_{\mu 3}$. Ratios of decay branching ratios BR($s_1^- \rightarrow e\nu$) : BR($s_1^- \rightarrow \mu\nu$) : BR($s_1^- \rightarrow \tau\nu$) are roughly given by 2 : 5 : 5 for $m_1 < m_3$ and 2 : 1 : 1 for $m_1 > m_3$ [71]. Therefore, Group-I and II can be tested by measuring leptonic decays of the charged scalar boson at the collider experiments.

When a group of models is favored by the experiments discussed above, we will try to discriminate models in the group by using details of each model. For example, the doubly-charged scalar boson is introduced in the ZB model in the Group-I while it does not exist in the KNT model of the Group-I. Thus, if the doubly-charged scalar boson is discovered at the collider experiments, the ZB model would be favored among models in the Group-I. This is the same for the CL model and the GNR model in the Group-II and the HTM in the Group-III. Even if groups of models have not been discriminated, collider experiments can test each models by measuring properties (e.g. decay patterns) of new particles as usually studied for model by model.

4. Conclusion

In this letter, we have studied the systematic classification of models for generating Majorana neutrino masses m_{y} according to combinations of Yukawa interactions. If we use Yukawa interactions for leptons by introducing new scalar fields relevant for these Yukawa interactions, the flavor structure of m_{ν} is given by three combinations: $Y_A^s y_\ell Y_S^s y_\ell (Y_A^s)^T$, $y_\ell (Y_S^s)^* y_\ell$, and Y_S^{Δ} . The Yukawa matrix Y_A is antisymmetric while Y_S^s and Y_S^{Δ} are symmetric. The Yukawa couplings y_ℓ are proportional to charged lepton masses. For the case where gauge singlet Z_2 -odd fermions ψ_{iR}^0 and Z_2 -odd scalar fields are addi-tionally introduced, the flavor structure of m_v is determined also by $Y_A^s y_\ell Y^s M_{\psi}^{-1} (Y^s)^T y_\ell (Y_A^s)^T$, $y_\ell (Y^s)^* M_{\psi}^{-1} (Y^s)^{\dagger} y_\ell$, and $Y^{\eta} M_{\psi}^{-1} (Y^{\eta})^{T}$. The Yukawa matrices Y_{S}^{s} and Y_{S}^{η} are symmetric, and M_{ψ} is the Majorana mass matrix for ψ_{iR}^{0} . Combining these results, we have found that models can be classified into only three groups: $m_v \propto Y_A^s y_\ell X_{SR} y_\ell (Y_A^s)^T$, $y_\ell X_{SR}^* y_\ell$, and X_{SL} . Here, X_{SR} and X_{SL} are some symmetric matrices. Although the structure of m_{ν} in the type-I seesaw and the type-III seesaw models can be classified in the Group-III, these models are exceptions to the discussion in this letter. Our classification enable us to approach efficiently to the origin of Majorana neutrino masses by testing not each model but each groups of models.

We concentrated on Majorana neutrino masses in this letter. The similar classification of models for Dirac neutrino masses is also desired because the nature may respect the lepton number conservation. This will be presented elsewhere [72].

Acknowledgements

This work was supported, in part, by Grant-in-Aid for Scientific Research No. 23104006 (SK) and Grant H2020-MSCA-RISE-2014 No. 645722 (Non-Minimal Higgs) (SK).

References

- G. Aad, et al., ATLAS Collaboration, Phys. Lett. B 716 (2013) 1;
 S. Chatrchyan, et al., CMS Collaboration, Phys. Lett. B 716 (2012) 30.
- [2] The ATLAS and CMS Collaborations, ATLAS-CONF-2015-044.
- [3] Y. Fukuda, et al., Super-Kamiokande Collaboration, Phys. Rev. Lett. 81 (1998) 1562;
- R. Wendell, et al., Super-Kamiokande Collaboration, Phys. Rev. D 81 (2010) 092004.
- [4] Q.R. Ahmad, et al., SNO Collaboration, Phys. Rev. Lett. 89 (2002) 011301;
- B. Aharmim, et al., SNO Collaboration, Phys. Rev. C 88 (2) (2013) 025501. [5] A. Zee, Phys. Lett. B 93 (1980) 389;
- A. Zee, Phys. Lett. B 95 (1980) 461.
- [6] L. Wolfenstein, Nucl. Phys. B 175 (1980) 93.
- [7] A. Zee, Nucl. Phys. B 264 (1986) 99.
- [8] K.S. Babu, Phys. Lett. B 203 (1988) 132.
- [9] T.P. Cheng, L.F. Li, Phys. Rev. D 22 (1980) 2860.
- M. Gustafsson, J.M. No, M.A. Rivera, Phys. Rev. Lett. 110 (21) (2013) 211802;
 M. Gustafsson, J.M. No, M.A. Rivera, Phys. Rev. Lett. 112 (25) (2014) 259902;
 M. Gustafsson, J.M. No, M.A. Rivera, Phys. Rev. D 90 (1) (2014) 013012.
- [11] W. Konetschny, W. Kummer, Phys. Lett. B 70 (1977) 433;
 R.N. Mohapatra, G. Senjanovic, Phys. Rev. Lett. 44 (1980) 912;
 M. Magg, C. Wetterich, Phys. Lett. B 94 (1980) 61;
 G. Lazarides, Q. Shafi, C. Wetterich, Nucl. Phys. B 181 (1981) 287;
- J. Schechter, J.W.F. Valle, Phys. Rev. D 22 (1980) 2227.
- [12] L.M. Krauss, S. Nasri, M. Trodden, Phys. Rev. D 67 (2003) 085002;
 A. Ahriche, S. Nasri, J. Cosmol. Astropart. Phys. 1307 (2013) 035;
 A. Ahriche, S. Nasri, R. Soualah, Phys. Rev. D 89 (9) (2014) 095010.
- M. Aoki, S. Kanemura, O. Seto, Phys. Rev. Lett. 102 (2009) 051805;
 M. Aoki, S. Kanemura, O. Seto, Phys. Rev. D 80 (2009) 033007;
 M. Aoki, S. Kanemura, K. Yagyu, Phys. Rev. D 83 (2011) 075016.

- [14] E. Ma, Phys. Rev. D 73 (2006) 077301;
- J. Kubo, E. Ma, D. Suematsu, Phys. Lett. B 642 (2006) 18.
- [15] P. Minkowski, Phys. Lett. B 67 (1977) 421;
 - T. Yanagida, Conf. Proc. C 7902131 (1979) 95;
 - T. Yanagida, Prog. Theor. Phys. 64 (1980) 1103;
 - M. Gell-Mann, P. Ramond, R. Slansky, Conf. Proc. C 790927 (1979) 315;
- R.N. Mohapatra, G. Senjanovic, Phys. Rev. Lett. 44 (1980) 912. [16] R. Foot, H. Lew, X.G. He, G.C. Joshi, Z. Phys. C 44 (1989) 441.
- [17] R.N. Mohapatra, J.W.F. Valle, Phys. Rev. D 34 (1986) 1642.
- [18] S. Khalil, Phys. Rev. D 82 (2010) 077702.
- [19] S. Fraser, E. Ma, O. Popov, Phys. Lett. B 737 (2014) 280.
- [20] A. Pilaftsis, Z. Phys. C 55 (1992) 275.
- [21] P.S.B. Dev, A. Pilaftsis, Phys. Rev. D 86 (2012) 113001.
- [22] S.M. Barr, Phys. Rev. Lett. 92 (2004) 101601.
- [23] W. Wang, Z.L. Han, Phys. Rev. D 92 (9) (2015) 095001.
- [24] S. Kanemura, T. Nabeshima, H. Sugiyama, Phys. Rev. D 85 (2012) 033004.
- [25] S. Kanemura, T. Matsui, H. Sugiyama, Phys. Rev. D 90 (2014) 013001.
- [26] S. Kanemura, H. Sugiyama, Phys. Rev. D 86 (2012) 073006.
- [27] J. Kubo, D. Suematsu, Phys. Lett. B 643 (2006) 336.
- [28] S. Kanemura, O. Seto, T. Shimomura, Phys. Rev. D 84 (2011) 016004.
- [29] M. Aoki, J. Kubo, H. Takano, Phys. Rev. D 87 (11) (2013) 116001.
- [30] B. Dasgupta, E. Ma, K. Tsumura, Phys. Rev. D 89 (4) (2014) 041702.
- [31] E. Ma, D. Suematsu, Mod. Phys. Lett. A 24 (2009) 583.
- [32] S. Kanemura, T. Ota, Phys. Lett. B 694 (2011) 233.
- [33] M. Aoki, S. Kanemura, K. Yagyu, Phys. Lett. B 702 (2011) 355;
 M. Aoki, S. Kanemura, K. Yagyu, Phys. Lett. B 706 (2012) 495.
- [34] K. Kumericki, I. Picek, B. Radovcic, J. High Energy Phys. 1207 (2012) 039.
- [35] H. Okada, T. Toma, Phys. Rev. D 86 (2012) 033011.
- [36] V. Brdar, I. Picek, B. Radovcic, Phys. Lett. B 728 (2014) 198.
- [37] A. Aranda, E. Peinado, arXiv:1508.01200 [hep-ph].
- [38] M. Lindner, D. Schmidt, T. Schwetz, Phys. Lett. B 705 (2011) 324.
- [39] H. Okada, T. Toma, K. Yagyu, Phys. Rev. D 90 (2014) 095005.
- [40] H. Hatanaka, K. Nishiwaki, H. Okada, Y. Orikasa, Nucl. Phys. B 894 (2015) 268.
- [41] A. Ahriche, C.S. Chen, K.L. McDonald, S. Nasri, Phys. Rev. D 90 (2014) 015024.
- [42] A. Ahriche, K.L. McDonald, S. Nasri, J. High Energy Phys. 1410 (2014) 167.
- [43] C.S. Chen, K.L. McDonald, S. Nasri, Phys. Lett. B 734 (2014) 388.
- [44] A. Ahriche, K.L. McDonald, S. Nasri, T. Toma, Phys. Lett. B 746 (2015) 430.
- [45] P. Culjak, K. Kumericki, I. Picek, Phys. Lett. B 744 (2015) 237.

- [46] Y. Kajiyama, H. Okada, K. Yagyu, Nucl. Phys. B 874 (2013) 198.
- [47] D. Restrepo, A. Rivera, M. Sánchez-Peláez, O. Zapata, W. Tangarife, Phys. Rev. D 92 (1) (2015) 013005.
- [48] K.L. McDonald, J. High Energy Phys. 1307 (2013) 020.
- [49] S.S.C. Law, K.L. McDonald, Phys. Rev. D 87 (11) (2013) 113003.
- [50] S.S.C. Law, K.L. McDonald, J. High Energy Phys. 1309 (2013) 092.
- [51] Y. Kajiyama, H. Okada, T. Toma, Phys. Rev. D 88 (1) (2013) 015029.
- [52] H. Okada, K. Yagyu, Phys. Rev. D 90 (3) (2014) 035019.
- [53] H. Okada, Y. Orikasa, Phys. Rev. D 90 (7) (2014) 075023.
- [54] H. Okada, N. Okada, Y. Orikasa, arXiv:1504.01204 [hep-ph].
- [55] S. Kashiwase, H. Okada, Y. Orikasa, T. Toma, arXiv:1505.04665 [hep-ph].
- [56] K. Nishiwaki, H. Okada, Y. Orikasa, arXiv:1507.02412 [hep-ph].
- [57] H. Okada, K. Yagyu, arXiv:1508.01046 [hep-ph].
- [58] E. Ma, Phys. Rev. Lett. 81 (1998) 1171;
 F. Bonnet, M. Hirsch, T. Ota, W. Winter, J. High Energy Phys. 1207 (2012) 153;
- D. Aristizabal Sierra, A. Degee, L. Dorame, M. Hirsch, J. High Energy Phys. 1503 (2015) 040.
- [59] K.S. Babu, C.N. Leung, Nucl. Phys. B 619 (2001) 667;
- F. Bonnet, D. Hernandez, T. Ota, W. Winter, J. High Energy Phys. 0910 (2009) 076.
- [60] X.G. He, Eur. Phys. J. C 34 (2004) 371.
- [61] S.M. Bilenky, J. Hosek, S.T. Petcov, Phys. Lett. B 94 (1980) 495;
 - M. Doi, T. Kotani, H. Nishiura, K. Okuda, E. Takasugi, Phys. Lett. B 102 (1981) 323.
- [62] Z. Maki, M. Nakagawa, S. Sakata, Prog. Theor. Phys. 28 (1962) 870.
- [63] A. Osipowicz, et al., KATRIN Collaboration, arXiv:hep-ex/0109033.
- [64] P.A.R. Ade, et al., Planck Collaboration, arXiv:1502.01589 [astro-ph.CO].
- [65] K.N. Abazajian, et al., Topical Conveners: K.N. Abazajian, J.E. Carlstrom, A.T. Lee Collaboration, Astropart. Phys. 63 (2015) 66.
- [66] U. Bellgardt, et al., SINDRUM Collaboration, Nucl. Phys. B 299 (1988) 1.
- [67] T. Abe, et al., Belle-II Collaboration, arXiv:1011.0352 [physics.ins-det].
- [68] E.J. Chun, K.Y. Lee, S.C. Park, Phys. Lett. B 566 (2003) 142;
- A.G. Akeroyd, M. Aoki, H. Sugiyama, Phys. Rev. D 79 (2009) 113010.
- [69] J.D. Vergados, H. Ejiri, F. Simkovic, Rep. Prog. Phys. 75 (2012) 106301.
- [70] M. Blennow, P. Coloma, P. Huber, T. Schwetz, J. High Energy Phys. 1403 (2014) 028.
- [71] K.S. Babu, C. Macesanu, Phys. Rev. D 67 (2003) 073010;
- D. Aristizabal Sierra, M. Hirsch, J. High Energy Phys. 0612 (2006) 052.
- [72] S. Kanemura, H. Sugiyama, in preparation.