## General Disclaimer

One or more of the Following Statements may affect this Document

- This document has been reproduced from the best copy furnished by the organizational source. It is being released in the interest of making available as much information as possible.
- This document may contain data, which exceeds the sheet parameters. It was furnished in this condition by the organizational source and is the best copy available.
- This document may contain tone-on-tone or color graphs, charts and/or pictures, which have been reproduced in black and white.
- This document is paginated as submitted by the original source.
- Portions of this document are not fully legible due to the historical nature of some of the material. However, it is the best reproduction available from the original submission.

Produced by the NASA Center for Aerospace Information (CASI)

# THE CANONIZATION OF NICE VARIABLES 

DAVID P. STERN

AUGUST 1970

GODDARD SPACE FLIGHT CENTER
GREEMBELT, MARYLAND


David P. Stern<br>Laboratory for Space Physics Goddard Space Flight Center Greenbelt, Maryland 20771

## Abstract

In a perturbed periodic classical motion the angle variable may be eliminated either by Kruskal'a transformation to "nice variables" or - if the system is canonical - by the Poincaré-Von Zeipel method. For systems that possess a Hamiltonian, the present work (1) shows that Kruskal's transformation may be made canonical order by order, (2) derives a practical formula for achieving this result and (3) shows that the two methods are equivaleat and may be matched order by order.

## INTRODUCTION

A basic perturbation problem in celestial mechanics and in guiding center motion involves a set of $n$ first-order differential equations which can be represented vectorially as

$$
\begin{equation*}
d x / d t=E(\underline{x}) \tag{1}
\end{equation*}
$$

and which has the following properties:
(1) $\underline{F}(\underline{x})$ depends on a small parameter $\mathcal{<}<1$ and may be expanded in it:

$$
\begin{equation*}
\underline{E}(\underline{x})=\underline{p}^{(0)}(\underline{x})+\varepsilon F^{(1)}(x)+\ldots \tag{2}
\end{equation*}
$$

(2) In the limit $\varepsilon \rightarrow 0$ ("unperturbed case") the system may be solved and its solution is then periodic in time.

The problem is then to find an approximate solution valid for amall but finite $\mathcal{E}$ and useful for time intervals of the order of $\varepsilon^{-1}$ periods.

Kruskal ${ }^{(1)}$ devised a method for achieving this, which in many ways resembles the method of Bogoliubor and Zubarev and of Krylov and Bogolium bot ${ }^{(2)}$. The calculation in this case proceede in two stepe. Firat, a transformation to "intermediate variables" $\underline{y}(\underline{x})$ is performed, such that in the limit $\varepsilon \rightarrow 0$, $\mathrm{Y}_{\mathrm{n}}$ is an angle variable linear in time, while the remaining ( $n-1$ ) coaponents of $\underline{y}$ (which we shall collectively denote by $\underline{\tilde{y}}$ ) are constant. The equations according to which $\underline{y}$ evolves, derived frome eq. (1), then have the form

$$
\begin{equation*}
d \underline{y} / d t=\sum_{k=0} \varepsilon^{k} \underline{g}^{(k)}(\underline{y}) \tag{3}
\end{equation*}
$$

In these equations, $y_{n}$ appears in the $\underline{E}^{(k)}$ only as the angle argument of periodical functions and $\underline{g}^{(0)}$ has only one non-rero components, the last ond ${ }^{(3)}$

$$
\begin{equation*}
\underline{\underline{E}}^{(0)}=(0,0, \ldots 0,1) \tag{4}
\end{equation*}
$$

Thas second step involves a near-identity tranaformation to "nice variablea" g

$$
\begin{equation*}
\underline{z}=\underline{z}+\sum_{k=1} \varepsilon^{k} \underline{S}^{(k)}(\underline{y}) \tag{5}
\end{equation*}
$$

such that in the now equations of evolution the transformed angle variable $z_{n}$ no longer appears on the right-hand side

$$
\begin{equation*}
d \underline{z} / \mathrm{dt}=\sum_{k=0} \varepsilon^{k} \underline{\underline{h}}^{(\mathrm{k})}(\underline{\underline{z}}) \tag{6}
\end{equation*}
$$

where as before

$$
\underline{\underline{z}}=\left(\tilde{\underline{\underline{z}}}, z_{n}\right)
$$

and

$$
\begin{equation*}
\underline{\underline{h}}^{(0)}=\underline{\underline{g}}^{(0)} \tag{7}
\end{equation*}
$$

A general recursion scheme for deriving $z$ order by order has been described in an articie by Stern ${ }^{(4)}$ (henceforth referred to as I) and we shall follow the notation introduced there, which differs slightly from Xruskal's (a similar scheme for the related Krylor-Bogoliubov expansion has been given by Kusen ${ }^{(5)}$ ). After equations (6) have been derived in this fashion, their firat ( $n-1$ ) components constitute an autonomous system for deriving the componente of $\tilde{\underline{\underline{Z}}}$, which can then be independently solved. The problem is thus reduced to one with ( $n-1$ ) variables.

If the system furthermore possesses a Hamiltaian, an additional variable may be eliminated by leriving a constant $J$ of the motion which Kruskal terms the adiabatic invariant. It is defined as

$$
\begin{equation*}
J=\oint \underline{p} \cdot d \underline{q}=\int_{0}^{1} \underline{p} \cdot\left(\partial \underline{q} / \partial z_{n}\right) d z_{n} \tag{8}
\end{equation*}
$$

with the integration performed using an arbitrary canonical set ( $\underline{p}, \underline{q}$ ), over a group of pointe ("ring") all of which evolve according to (6) and possess the lame $\underline{z}$ but with values of $z_{n}$ that cover the full range for that variable (this property, if once established, is maintained throughout the ring's evolution in time).

If $\underline{z}$ itself forns a canonical set, with $z_{1}$ the momentum conjugate to $z_{n}$, then one may use this set in ( 8 ), leading to

$$
\begin{equation*}
J=\int_{0}^{1} z_{1} d z_{n}=z_{1} \tag{9}
\end{equation*}
$$

wilch is a great ajmplification (if this is not the case, one is forced to retrace the transformations $\underline{\underline{x}} \rightarrow \underline{\underline{V}} \rightarrow \underline{z}$ until some canonical set, with wich (8) can be evaluated, is reached). Kruskal did not derive nice canonical variables, but he showed them to be possible. Speciflcally, he proved that in any "nice" set, the following Poisson bracket relations are always satisfied

$$
\begin{align*}
\left(\partial / \partial z_{n}\right)\left[z_{i}, s_{j}\right] & =0  \tag{10}\\
{\left[s_{i}, j\right] } & =\delta_{\text {in }} \tag{11}
\end{align*}
$$

Following Mordhoin and Pues ${ }^{(6)}$, Kruskal then shows that $J$ and $z_{n}$ may be augnented by ( $n-2$ ) functions of the $z_{i}$ to form a complete canonical set.

In what follows we shall assume that the intermediate variables $y_{i}$ form a canonical set, in which case a near-identity transformation like (5) can, in principle, lead to nice canonical variables. This assumption is not unreasonable since, if a Hamiltonian for the system is known, such $y_{i}$ can in general be derived by solving the unperturbed motion via the Hanilton-Jacobi method.

We then
(1) Show that the freedom allowed by Kruskal's method in the derivation of each order of eq. (5) is sufficient to assure the canonical character of the "nice variables" $z_{i}$.
(2) Derive a mod for obtaining such $z_{i}$.
(3) Show that the result is equivalent to what is obtained by conventional perturbation methods based on the Hanilten-Jacobi equation.

## THE POSSIBILITY OF STEEP-BY-STHP DERIVATION

If $\underline{y}$ is canonical

$$
\underline{\mathbf{y}}=(\underline{p}, \underline{q})
$$

then from (10)

$$
\begin{equation*}
\left(\partial / \partial x_{n}\right)\left\{\left[z_{i}, z_{j}\right]-\left[y_{i}, y_{j}\right]\right\}=0 \tag{12}
\end{equation*}
$$

If (5) is substituted here, the zoro-order part cancels identically and the expression remaining inside the curly brackets separates into orders of $\varepsilon$ and
gives

$$
\begin{equation*}
\left(\partial / \partial z_{n}\right) \sum_{m=1} \varepsilon^{m}\left\{\left[\zeta_{i}^{(m)}, y_{j}\right]-\left[\zeta_{j}^{(m)}, y_{i}\right]+\sum_{s=1}^{m-1}\left[\zeta_{i}^{(s)}, \zeta_{j}^{(m-s)}\right]\right\}=0 \tag{13}
\end{equation*}
$$

This condition is satisfied for any "nice" set $\underline{\underline{Z}}$. If $\underline{z}$ is not merely nice but also canonical, then the expressions in the curly brackets of both (12) and (13) are not only independent of $z_{n}$ but actually vanish.

Let us now assume that the expansion (5) has already been derived and brought to canonical form, up to and including order ( $k-1$ ). Then the first ( $k-1$ ) orders of (13) do, in fact, vanish, leaving (after division by $\varepsilon^{k}$ )
$\left(\partial / \partial z_{n}\right)\left\{\left[\zeta_{i}^{(k)}, y_{j}\right]-\left[\zeta_{j}^{(k)}, y_{i}\right]+\sum_{s=1}^{k-1}\left[\zeta_{i}^{(s)}, \zeta_{j}^{(k-s)}\right]\right\}+0(\varepsilon)=0$
As the next step, one may derive $\underline{\zeta}^{(k)}$ and thus extend the calculation of $\underline{z}$ one more order. The equation satisfied by $\underline{\zeta}^{(k)}$ is (I , eq. 15)

$$
\begin{equation*}
\partial \underline{y}^{(k)} / \partial y_{n}-\underline{n}^{(k)}(\underline{\tilde{y}})=\underline{\lambda}^{(k)} \tag{15}
\end{equation*}
$$

where $\underline{\lambda}^{(k)}$ depends only on lower orders of $\underline{\underline{\zeta}}^{(m)}$ and $\underline{h}^{(m)}$, assumed to be known at this stage. If $\left\langle\underline{\lambda}^{(k)}\right\rangle$ denotes the result of averaging over the angle variable $y_{n}$

$$
\begin{equation*}
\left\langle\underline{\lambda}^{(k)}\right\rangle=\int_{0}^{1} \underline{\lambda}^{(k)}(\underline{y}) d y_{n} \tag{16}
\end{equation*}
$$

one gets (see I, eq. 19)

$$
\begin{align*}
\underline{\underline{S}}^{(k)} & =\int_{0}^{y_{n}}\left(\underline{\lambda}^{(k)}-\left\langle\underline{\lambda}^{(k)}\right\rangle\right) d y_{n}^{\prime}+\mu^{(k)}(\underline{\tilde{y}}) \\
& =\underline{\zeta}^{(k)}+\underline{\mu}^{(k)}(\underline{\tilde{y}}) \tag{17}
\end{align*}
$$

Here $\mu^{(k)}$ is an arbitrary additive vector independent of $y_{n}$, allowed by the fact that the derivative of $\underline{\zeta}^{(k)}$ which enters (15) is unaffected by such an addition. The question now arises whether $\mu^{(k)}$ may be selected so as to make the expansion (5) canonical to order $k$.

If this occurs, then the $O(1)$ part of (14) must vanish. It helps here to introduce the concept of the conjugate vectors $\overline{\mathbf{y}}$ (I, eq. 4)

$$
\begin{equation*}
\overline{\mathbf{y}}=(\underline{q},-\underline{p}) \tag{18}
\end{equation*}
$$

Then

$$
\begin{equation*}
\left[\zeta_{i}^{(m)}, y_{j}\right]-\left[\zeta_{j}^{(m)}, y_{i}\right]=\partial \zeta_{i}^{(m)} / \partial \bar{y}_{j}-\partial \zeta_{j}^{(m)} / \partial \bar{y}_{i} \tag{19}
\end{equation*}
$$

This has the form of a component of a curl dyadic in $\underline{\overline{\mathbf{y}}}$ space, which one may call the conjugate curl, with component emoted by

$$
\begin{equation*}
\left(\bar{\nabla} \times \underline{\underline{s}}^{(\mathbb{m})}\right)_{i j} \tag{20}
\end{equation*}
$$

Now it has been shown by Stern ${ }^{\text {in an article }}{ }^{(7)}$ (henceforth referred to as II) that the general condition for an expansion (5) to be canonical is

$$
\begin{align*}
\underline{\zeta}^{(m)}=\underline{\underline{f}}^{(m)}(\underline{\zeta})+ & \bar{\nabla} X^{(m)}  \tag{21}\\
& (m=1,2, \ldots)
\end{align*}
$$

where $\bar{\nabla}$ is a gradient operator in $\underline{\bar{y}}$ space, $\chi^{(m)}$ is an arbitrary scalar and $\underline{\underline{L}}^{(m)}(\underline{\zeta})$ are vectors of a certein form, depending on orders of $\underline{\zeta}$ lower than the $m$-th and on their derivatives. Various choices of $f^{(m)}$ are derived in II, all of which satisfy the identity (for canonical $\underline{S}^{(s)}$ )

$$
\begin{equation*}
\left(\bar{\nabla} \times{\underset{m}{ }}_{(m)}^{(m j}=-\sum_{s=1}^{m-1}\left[\zeta_{i}^{(s)}, \zeta_{j}^{(m-s)}\right]\right. \tag{22}
\end{equation*}
$$

Because the $\underline{乌}^{(s)}$ in this case are known to be cenonical up to and iraluding order ( $k-1$ ), eq. (22) will hold for the lowest ( $k-1$ ) values of $m$; however, it will also hold for $m=k$, for even then the orders of $\underline{S}$ appearing in the equation are all lower than the $k$-th. Substitution in (14) then yields

$$
\begin{equation*}
\left(\partial / \partial z_{n}\right)\left\{\bar{\nabla} \times\left[\underline{\zeta}^{(k)}-\underline{\underline{f}}^{(k)}(\underline{\zeta})\right]\right\}+0(\varepsilon)=0 \tag{23}
\end{equation*}
$$

Now let the transformatinn inverse to (5) be

$$
\begin{equation*}
y=\underline{z}+\sum_{s=1}^{1} \varepsilon^{s} \underline{\eta}^{(s)}(\underline{z}) \tag{24}
\end{equation*}
$$

Given the expansion (5), the $\underline{\eta}^{(k)}$ may easily be derived (see appendix). Alternatively, they may be directly obtained from expanding the relation between (3) and (6), in a manner similar to what has been done for the expansion (5) in I . This is essentially the method of Krylov and Bogoliubov (2), as expanded by Musen ${ }^{(5)}$. Indeed, the elimination of the angle variable by Kruskal's method so resembles the Krylov-Bogoliubov approach that the two ought perhaps to be regarded as one method (which could be called the Krylov-Bogoliubov-Kruskal method; however, Kruskal's work proceeds past the elimination to the derivation of adiabatic invarianta).

Since the inner part of (23) is a function of $\underline{y}$, one must transform

$$
\begin{align*}
\partial / \partial z_{n} & =\sum 1\left(\partial y_{i} / \partial z_{n}\right) \partial / \partial y_{i} \\
& =\sum\left\{\delta_{i n}+\sum \varepsilon^{\theta}\left(\partial \eta_{i}^{(a)} / \partial z_{n}\right)\right\} \partial / \partial y_{i} \tag{25}
\end{align*}
$$

or

$$
\begin{equation*}
\partial / \partial z_{n}=\partial / \partial y_{n}+\partial(\varepsilon) \tag{26}
\end{equation*}
$$

Eq:2tion (23) thus becomes

$$
\begin{equation*}
\partial / \partial y_{n}\left\{\bar{\nabla} \times\left[\underline{\zeta}^{(k)}-{\underline{f}^{(k)}}_{(\underline{\zeta})]\} \quad+o(\varepsilon)=0.000}\right.\right. \tag{27}
\end{equation*}
$$

One may now assume that all variables in (27) are expressed in terms of $\underline{y}$; in that case, each under in $\mathcal{E}$ vanishes separately, including the zeroth. Because differential operators commute, this means

$$
\begin{equation*}
\bar{\nabla} \times\left\{\left(\partial / \partial y_{n}\right)\left[\underline{\zeta}^{(k)}-\underline{\underline{e}}^{(k)}(\underline{\zeta})\right]\right\}=0 \tag{28}
\end{equation*}
$$

This integrates to

$$
\begin{align*}
\left(\partial / \partial y_{n}\right)\left[\underline{\zeta}^{(k)}-\underline{f}^{(k)}(\underline{S})\right]= & \bar{\nabla} \psi  \tag{29}\\
& \underline{\zeta}^{(k)}-\underline{f}^{(k)}(\underline{\zeta})=\bar{\nabla} \tau+\underline{u}(\tilde{z}) \tag{30}
\end{align*}
$$

where $\tau$ is the indefinite integral of $\Psi$ and $\underline{u}$ is an additive function independent of $z_{n}$, allowed by the integration.

The above condition is satisfied by any $\underline{\$}^{(k)}$ belonging to a nice set of variables which is canonical to order $(k-1)$. For instance, $\hat{\zeta}^{(k)}$ defined in (17) will correspond to a certain $\tau$ and to a certain additive function, which may be denoted by $\hat{\underline{u}}$ :

If one now chooses in (17)

$$
\mu^{(k)}=-\underline{\underline{u}}
$$

then

$$
\begin{align*}
\underline{\zeta}^{(k)} & =\underline{\hat{\zeta}}^{(k)}-\hat{\underline{u}} \\
& =\underline{\underline{f}}^{(k)}(\underline{\zeta})+\bar{\nabla} \tau \tag{32}
\end{align*}
$$

and by (21), $\underline{\underline{y}}^{(k)}$ satisfies the condition for canonical variables, making $\underline{z}$ canonical to order $k$. Thus the requirement can be met.

## PRACTICAL CANONIZATION

In order to actually derive the "canonizing" $\mu^{(k)}$, one must first investigate the amount of arbitrariness inherent in that vector. Let

$$
\begin{equation*}
\underline{\zeta}^{(k)}=\hat{\zeta}^{(k)}+{\mu_{1}^{(k)}(\tilde{J})}^{(k)} \tag{33}
\end{equation*}
$$

extend the canonical properties to $0\left(\varepsilon^{k}\right)$, i.e. let it satisfy

$$
\begin{equation*}
\hat{\zeta}^{(k)}+\underline{\mu}^{(k)}(\underline{\tilde{y}})=\underline{\underline{e}}^{(\dot{k})}(\underline{\Sigma})+\bar{\nabla} \chi^{(k)} \tag{34}
\end{equation*}
$$

Then if one replaces $\mu^{(k)}$ with

$$
\begin{equation*}
\mu^{(k)}=\mu^{(k)}(\underline{\tilde{y}})+\bar{\nabla} \Phi(\tilde{y}) \tag{35}
\end{equation*}
$$

with $\Phi$ an arbitrary function independent of $y_{n}$, one Ganda

$$
\begin{equation*}
\hat{\zeta}^{(k)}+\underline{\mu}^{(k)}(\underline{\tilde{y}})=\underline{\underline{f}}^{(k)}(\underline{\jmath})+\bar{\nabla}\left(\chi^{(k)}+\bar{\Phi}\right) \tag{36}
\end{equation*}
$$

which still has the required form (21). The canonizing choice of $\mu^{(k)}$ is thus arbitrary within the addition of a conjugate gradient of some scalar $\Phi$ which does not contain $y_{n}$ 。

In order to isolate $\mu^{(k)}$ one operates on (34) with the avernging operator of (16); since $\mu^{(k)}$ is independent of the ancle variable, it equals its own average, giving

$$
\begin{align*}
\underline{\mu}^{(k)}(\underline{\tilde{y}}) & =\left\langle\underline{f}^{(k)}(\underline{\jmath})-\underline{\hat{\zeta}}^{(k)}\right\rangle+\left\langle\bar{\nabla} \chi^{(k)}\right\rangle \\
& =\left\langle\underline{f}^{(k)}(\underline{\grave{\zeta}})-\underline{\hat{\zeta}}^{(k)}\right\rangle+\bar{\nabla}\left\langle\chi^{(k)}\right\rangle \tag{37}
\end{align*}
$$

The last term on the right is a conjugate gradient of some function of $\underline{\tilde{v}}$, and it has already been established that such rectors, when added to $\underline{\mu}^{(k)}$, do not affect canonization. One may thus drop this part end obtain

$$
\begin{equation*}
\underline{\mu}^{(k)}=\left\langle\underline{\underline{f}}^{(k)}(\underline{\zeta})-\underline{\zeta}^{(k)}\right\rangle \tag{38}
\end{equation*}
$$

 obtained by methods given in II. The most general canonizing additive function is then

$$
\begin{equation*}
\underline{\underline{\mu}}^{(k)}=\left\langle\underline{\underline{f}}^{(k)}(\underline{\zeta})-\hat{\zeta}^{(k)}\right\rangle+\bar{\nabla} \Phi(\tilde{y}) \tag{39}
\end{equation*}
$$

with $\Phi$ arbitrary,

## EQUIVALENCE TO CANONICAL PERTURBATION THEORY

A widely used method for solving perturbed periodic canonical systems is due to Poincare and Yon Zeipel $(5),(8)-(11)$ and operates in the following manner. First, one expresses the Hamiltonian in terms of the solution

$$
\underline{\mathbf{y}}=(\underline{p}, \underline{q})
$$

of the unperturbed Hamilton-Jacobi equation, with $y_{n}$ an angle variable and $y_{1}$ the conjugate action variable. In the absence of "slowly varying" quantities the Hamiltonian then assumes the form

$$
\begin{equation*}
\mathrm{H}=y_{i}+\sum_{k=1} \varepsilon^{k} H^{(k)}(\underline{y}) \tag{40}
\end{equation*}
$$

Next a near-identity transformation to nev variables

$$
\underline{\underline{\varepsilon}}=(\underline{p}, \underline{\underline{g}})
$$

is sought, produced by the generating function

$$
\begin{equation*}
\sigma(\underline{p}, \underline{q})=\sum p_{i} q_{1}+\sum_{k=1} \varepsilon^{k} \sigma^{(k)}(\underline{p}, \underline{q}) \tag{41}
\end{equation*}
$$

and having the property that the new Hamiltonian $H^{*}(z)$ is independent of the transformed angle variable $z_{n}$, making its conjugate $z_{1}$ a constant of the motion. Since the transformation is a near-identity one, the lowest order of $H^{*}$ has the same form as that of H, giving

$$
\begin{equation*}
H^{*}(\tilde{\underline{z}})=z_{1}+\sum_{k=1} \varepsilon^{k} H^{(k)}(\tilde{y}) \tag{42}
\end{equation*}
$$

Methods then exist for deriving $\sigma^{(k)}$ and $H^{*(k)}$ order by order

Suppose now that two near-identity canonical transformations are given

$$
\begin{align*}
& \underline{\mathbf{z}}=\underline{y}+\sum_{s=1} \varepsilon^{s} \underline{Y}^{(s)}(\underline{y})  \tag{43}\\
& \underline{\mathbf{v}}=\underline{y}+\sum_{s=1} \varepsilon^{s} \underline{\Psi}^{(s)}(\underline{y})
\end{align*}
$$

either of which eliminates the angle variable from the Hamiltonian. Let furthermore

$$
\begin{equation*}
\underline{S}^{(s)}(\underline{y})=\underline{\Psi}^{(s)}(\underline{y}) \tag{44}
\end{equation*}
$$

for

$$
\mathrm{s}=1,2, \ldots(\mathrm{k}-1)
$$

Then it will be shownthat $\underline{\zeta}^{(k)}$ differs from $\underline{\Psi}^{(k)}$ at most by a conjugate gradient of a function $\Phi(\underline{\tilde{y}})$ independent of $y_{n}$.

Clearly such a property would allow the two methods discussed in this work to be matched order by order. If $\underline{\underline{~}}$, for instance, represents a solution of a given problem by the Poincaré-Von Zeipel method and $\underline{z}$ represents a solution by the Kruskal method, then $\underline{z}$ can be made equal to $\underline{w}$ to any desired order. To achieve this one must first choose $\mu^{(k)}$ so as to make $z$ canonical and then add to each order the appropriate $\Phi(\underline{\tilde{y}})$ which makes the corresponding $\underline{S}^{(k)}$ and $\underline{\Psi}^{(k)}$ equal to each other. Actually, one could also work in the opposite direction, since the Poincere-Von Zeipel method also contains a certain amount of arbitrariness in each order, but we shall not consider this possibility here.

Let

$$
H^{*}(\underline{\tilde{z}})=\sum_{s=0} \varepsilon^{s} H^{*(s)}(\underline{\tilde{z}})
$$

and

$$
H^{* *}(\tilde{\underline{w}})=\sum_{s=0} \varepsilon^{s} H^{* *(s)}(\tilde{\tilde{W}})
$$

be the two alternative forms of the Hamiltonian. Since the transformation is time-independent

$$
\begin{equation*}
\mathbf{H}^{*}(\underline{\tilde{z}})=\mathbf{H}^{* *}(\underline{\underline{u}})=\mathbf{H}(\underline{y}) \tag{45}
\end{equation*}
$$

Also, since it is a near-identity transformation, the $O(1)$ parts of the above equations must be equal, which leads to

$$
\begin{align*}
& \mathbf{H}^{*(0)}\left(\underline{\underline{\underline{\alpha}})}=\varepsilon_{1}\right. \\
& \mathbf{H}^{*(0)}\left(\tilde{\underline{\underline{v}})}=w_{1}\right. \tag{46}
\end{align*}
$$

We shall furthermore assume (and later justify)

$$
\begin{aligned}
& H^{*(s)}=H^{* *(s)} \\
& \quad(s=1,2, \ldots k-1)
\end{aligned}
$$

Substituting the expansions in (45)

$$
\begin{equation*}
\sum \varepsilon^{s} E^{(s)}\left(\tilde{\tilde{y}}+\sum \varepsilon^{m} \underline{\zeta}^{(m)}\right)=\sum_{\alpha} \varepsilon^{s} \mathbf{a}^{*}(\mathrm{~s})\left(\underset{\mathrm{y}}{ }+\sum \varepsilon^{\mathrm{a}} \underline{Y}^{(n)}\right) \tag{48}
\end{equation*}
$$

By means of expansion operators $\mathrm{T}^{(\mathrm{m})}$ and $\mathrm{s}^{(\mathrm{m})}$ (eec I) this may be broken up into a series of equations, one for each order in $\varepsilon$. The equation
for $O\left(\varepsilon^{k}\right)$ then is

$$
\begin{equation*}
\sum_{m=0}^{k} T^{(m)} * H^{*(k-m)}(\underline{\tilde{y}})=\sum_{m=0}^{k} S^{(m)} * H^{* *(k-m)}(\underline{\tilde{y}}) \tag{49}
\end{equation*}
$$

where marks "operates on" and where

$$
\begin{aligned}
& T^{(0)}=S^{(0)}=1 \\
& T^{(1)}=\underline{乌}^{(1)} \cdot \nabla \\
& S^{(1)}=\underline{\Psi}^{(1)} \cdot \nabla
\end{aligned}
$$

and in general

$$
\begin{align*}
& T^{(m)}=\underline{\zeta}^{(m)} \nabla+N^{(m)}(\underline{\zeta})  \tag{50}\\
& S^{(m)}=\underline{\Psi}^{(m)} \cdot \nabla+N^{(m)}(\underline{\Psi})
\end{align*}
$$

with $N^{(m)}$ an operator depending only on orders of its argument that are lower than the moth. Substitution and use of (46) give

$$
\begin{align*}
& H^{*(k)}+\zeta^{(k)} \cdot \nabla y_{1}+\left\{\mathbb{N}^{(k)}(\underline{\zeta}) * y_{1}+\sum_{m=1}^{k-1} T^{(m)} H^{*(k-m)}\right\}= \\
& =H^{* *(k)}+\Psi^{(k)} \cdot \nabla y_{1}+\left\{N^{(k)}(\underline{\Psi}) * y_{1}+\sum_{m=1}^{k-1} s^{(m)} * H^{* *(k-m)}\right\} \tag{51}
\end{align*}
$$

The two expressions in curly brackets depend only on lower orders; they are therefore equal to each other and may be dropped, leaving

$$
\begin{equation*}
\mathbf{H}^{*(\mathbf{k})}(\underline{\tilde{y}})=\mathbf{H}^{* *(\mathbf{k})}(\underline{\tilde{y}})=\left(\underline{\Psi}^{(\mathbf{x})}-\underline{\zeta}^{(\mathbf{k})}\right) \cdot \nabla \mathrm{y}_{1} \tag{52}
\end{equation*}
$$

Now by (24), scalars $\chi^{(k)}$ and $\tau^{(k)}$ must exist such that

$$
\begin{aligned}
& \underline{\zeta}^{(k)}=\underline{\underline{f}}^{(k)}(\underline{\zeta})+\bar{\nabla} \chi^{(k)} \\
& \underline{\Psi}^{(k)}=\underline{\underline{f}}^{(k)}(\underline{\Psi})+\bar{\nabla} \tau^{(k)}
\end{aligned}
$$

Since lower orders of the expansions are equal, the two $\underline{f}^{(k)}$ vectors are equal too, leaving

$$
\begin{equation*}
\underline{\zeta}^{(k)}-\underline{\Psi}^{(k)}=\bar{\nabla}\left(X^{(k)}-\tau^{(k)}\right)=\bar{\nabla} \Phi \tag{54}
\end{equation*}
$$

To prove our assertion we must show that $\Phi$ is independent of $y_{n}$. Substituting in (52)

$$
\begin{equation*}
\mathbf{H}^{*(k)}(\underline{\tilde{y}})-\mathbf{H}^{* *(k)}(\underline{\tilde{y}})=\partial \Phi / \partial \overline{\mathrm{y}}_{1}=\partial \Phi / \partial \mathrm{y}_{\mathrm{n}} \tag{55}
\end{equation*}
$$

Now $\Phi$ is allowed to depend on $y_{n}$ only in a periodic manner, from which it follows that $\partial \Phi / \partial y_{n}$ also depends periodically on $y_{n}$. However, the left-hand side is independent of that variable, so that $\Phi$ must be independent of $y_{n}$. This proves the main assertion. Incidentally, one also finds.

$$
\begin{equation*}
H^{*(k)}=H^{* *(k)} \tag{56}
\end{equation*}
$$

which justifies equation (47).

## APPENDIX : THE INVERSE TRANSFORMATION

Adding (5) and (24) and cancelling zeroth order terms gives

$$
\begin{align*}
\sum^{\prime} \varepsilon^{k} \underline{\eta}^{(k)}(\underline{z}) & =-\sum^{1} \varepsilon^{k} \underline{\zeta}^{(k)}(\underline{y}) \\
& =-\sum^{1} \varepsilon^{k} \underline{\zeta}^{(k)}\left(\underline{z}+\sum \varepsilon^{m} \underline{\eta}^{(\underline{m})}(\underline{z})\right) \tag{A-1}
\end{align*}
$$

If $\nabla$ is in z-space and * detiotes operation

$$
\begin{align*}
\zeta^{(k)}\left(\underline{z}^{(k)} \varepsilon^{m} \underline{\eta}^{(m)}(\underline{z})\right) & =\exp \left(\sum_{m=1} \varepsilon^{m} \underline{\eta}^{(m)}(\underline{z}) \cdot \nabla\right) * \underline{\zeta}^{(k)}(\underline{z}) \\
& =\sum_{m=0} \varepsilon^{m} s^{(m)} \underline{\zeta}^{(k)}(\underline{z}) \quad \tag{A-2}
\end{align*}
$$

where

$$
\begin{align*}
& s^{(0)}=1 \quad s^{(1)}=\eta^{(1)} \cdot \nabla  \tag{A-3}\\
& s^{(2)}=\eta^{(2)} \nabla+\frac{1}{2} \eta^{(1)} \eta^{(1)}: \nabla \nabla
\end{align*}
$$

and so forth. If this is substituted in (A-1), orders of $\varepsilon$ may be individually equated since everything is now expressed in $\underline{z}$. Separating the $=0$ term from the rest then brings the $O\left(\varepsilon^{k}\right)$ relation to the form

$$
\begin{equation*}
\underline{\eta}^{(k)}=-\underline{\zeta}^{(k)}-\sum_{m=1}^{k-1} s^{(m)} \underline{\zeta}^{(k-m)} \tag{A-4}
\end{equation*}
$$

If all lower orders of $\underline{S}^{(m)}$ are known, those of $\overbrace{}^{(m)}$ may be derived and used for constructing the $s^{(m)}$.

## REFGREMNCES

(1) M. Kruskal, Jo Math. Phys. 3, 806 (1962)
(2) N. N. Bogoliubov and D. N. Zubarev, Ukr. Math. J. VII, 5 (1955) ; Translation by Burton Fried published 1960 by Space Technology Laboratories. The method of Krylov and Bogoliubov is described in N. M. Krylov and N.N. Bogoliubov, Introduction to Non-Linear Mechanics, Kiev 2936, and also in N.N. Bogoliubov and Y. A. Mitropolsky, Asymptotic Methods in the Theory of Nonlinear Oscillations, Gordon \& Breach, HiY. 1961, sect. 25.
(3) Kruskal's original method is slightly more general than this in allowing the non-zero component of (4) to be a function independent of the angle variable. The recursion can then still be performed, but in each order the derivation of the $n$-th component must be postponed until all other components have been found. If $\underline{y}$ is canonical, this more general behavior only occurs if some of the variables are "slowly varying."
(4) D. Stern, On Kruskal's Perturbation Method, Jo Math. Phys. (to be published)
(5) P. Musen, J. of the Astronautical Sciences, 12, 129 (1965)
(6) L. Nordheim and E. Fues, Handbuch der Physik, Ed. S. Flügge (Julius Springer, Berlin 1927)
(7) D. Stern, Direct Canonical Transformations, J. Kath. Phys. (to be published)
(8) H. Poincaré, Les Methodes Nouvelles de la Mecanique Celeste, vol. II , Gauthier-Villars, Paris 1893; reprinted by Dover Publications, New York, 1957; NASA translation TT F-451, 1967 •
(9) H. Von Zeipel, Ark. Astr. Mat. Fys. 11, no. 1, (1916)
(10) H. C. Corben and P. Stehle, Classical Mechanics, 2nd. Ed.; John Wiley and Sons, 1960; Section 75.
(11) D. Ter Haar, Elements of Hamiltonian Mechanics, North Holland Publishing Co. 1961 ; Chapter 7 .

