## THEORETICAL CHEMISTRY INSTITUTE

## the UNIVERSTTY OF WISCONSIN

NOTE ON THE SYMMETRY OF PERTURBED HARTREE-FOCK AND
X-a WAVEFUNCTIONS

John O. Eaves and Saul T. Epstein


29 March 1974

```
(MASA-CR-138259) NOTE ON THE SYAMETRU
N74-23292 OF GERTURBED HAETREE-FOCK ARD X-ALPEA GAYEFDACTIONS (Gisconsin Univo) \(10-\mathrm{p}\) HC

\section*{MADISON, WISCONSIN}

\title{
NOTE ON THE SYMMETRY OF PERTURBED HARTREE-FOCK
} X- \(\alpha\) WAVEFUNCTIONS*

John O. Eaves and Sau1 T. Epstein
Theoretical Chemistry Institute
University of Wisconsin
Madison, Wisconsin 53706

ABSTRACT

It is shown that the first order orbitals for \(X-\alpha\) or Hartree-Fock atoms perturbed by multipole electric fields have the expected symmetry properties.
Research supported by the National Aeronautics and Space
Administration Grant NGL 50-002-001 and the National
Science Foundation Grant GP-28213.
* - - -

Recently it has been suggested \({ }^{1}\) that the first order orbitals describing the perturbation of Hartree-Fock or \(\mathrm{X}-\alpha\) closed-shell atoms (more generally atoms with all occupied spatialshells closed) by a multipole electric field do not have the expected symmetry properties. In this note we will show that these fears are groundless. More precisely we will show that the assumption of the expected symmetry is a selfconsistent one. This of course does not immediately show that it is the only solution of the perturbation equations, but since the latter are inhomogeneous linear equations, we expect that this is probably the case. [Ahlberg and Goscinski (private communication) have now also reached the same conclusiop, as regards \(X-\alpha]\) We consider the perturbation of an orbital \(U_{n_{i}} \ell_{i} m_{i}\) of the form
\[
\begin{equation*}
\mathrm{U}_{\mathrm{n}_{\mathrm{i}} \ell_{i} \mathrm{~m}_{\mathrm{i}}}^{(0)}(\overrightarrow{\mathrm{I}})=\mathrm{Y}_{\ell_{i} \mathrm{~m}_{\mathrm{i}}}(\hat{1}) \mathrm{R}_{\mathrm{n}_{\mathbf{i} \ell}{ }_{i}}\left(\mathrm{r}_{1}\right) \tag{1}
\end{equation*}
\]
by a perturbation proportional to \(P_{L}(\hat{I})\) where \(\vec{I}\) represents all the cartesian coordinates of particle 1 and where \(\hat{1}\) is a unit vector in the \(\overrightarrow{1}\) direction and hence represents the angular coordinates. Then writing
\[
P_{L}(\hat{I}) Y_{\ell_{i} m_{i}}(\hat{l})=\sum_{j m_{j}} \sum_{j=j}^{j} Y_{j m_{j}}(\hat{I})\left(Y_{j m_{j}}, P_{L} Y_{\ell_{i} m_{1}}\right)
\]
one expects \({ }^{2}\) that the first order perturbed orbital will have the form
\[
\begin{equation*}
U_{n_{i} \ell_{i} m_{i}}^{(1)}=\sum_{j} R_{j n_{i} \ell_{i} L}^{(1)}\left(r_{I}\right) \sum_{m_{j}}^{j} Y_{j-j}(\hat{1})\left(Y_{j m_{j}}, P_{L} Y_{\ell_{i} m_{i}}\right) \tag{3}
\end{equation*}
\]

To see whether or not this assumption is self-consistent we first examine the first-order correction to the charge density. The contribution from a closed-shell of orbital angular momentum \(k\) is then
\[
\begin{equation*}
\sum_{m_{k}}^{k}=-k U_{n_{k} k m_{k}}^{(1)}(\overrightarrow{1}) U_{n_{k} k m_{k}}^{(0) *}(\overrightarrow{1})+c 。 c 。 \tag{4}
\end{equation*}
\]
and in turn the contribution to this from a given \(j\) in (3) is then, in its angular dependence, proportional to
\[
Q_{j k}^{L 0}+Q_{j k}^{L 0 *}
\]
where
\[
\begin{equation*}
Q_{j k}^{L M}(\hat{1}) \equiv \sum_{m_{k}=-k m_{j}}^{k} \sum_{j-j}^{j} Y_{j m_{j}}(\hat{1})\left(Y_{j m_{j}}, Y_{L M^{\prime} m_{k}}\right) Y_{k m_{k}}^{*}(\hat{1}) \tag{5}
\end{equation*}
\]

We will now prove that \(Q_{j k}^{L M}(\hat{l})\) is an \(M\) independent multiple of \(\mathrm{Y}_{\mathrm{LM}}\) (1) and this, as one can then readily verify, is sufficient to completely guarantee the consistency of our assumption as far as \(X-\alpha\) is concerned, and, is sufficient for Hartree-Fock exclusive of the exchange terms. (Note for example that the potential produced by \(Q_{j k}^{L 0}\) will also be proportional to \(\left.P_{L}(\hat{1})\right)\).

We first use the spherical harmonic addition theorem to rewrite (5), to within a constant factor as
\[
\begin{equation*}
Q_{j k}^{L M}(\hat{1})=\int \mathrm{d} \hat{2} P_{j}(\hat{1} \cdot \hat{2}) Y_{L M}(\hat{2}) P_{k}(\hat{1} \cdot \hat{2}) \tag{6}
\end{equation*}
\]

It is then easy to show \({ }^{3}\) that \(Q_{j k}^{L M}(\hat{I}\); transforms under rotation precisely like \(\mathrm{Y}_{\mathrm{LM}}(\hat{\mathrm{I}})\) and hence as claimed must be an M-inOGplswbeir numerical multiple of \(Y_{L M}(\hat{1})\) 。 The numerical coefficient is evaluated in the appendix of this article.

Turning now to the exchange terms of Hartree-Fock, the contribution from a given \(j\) and \(k\) in the equation for \(U_{n_{i} \ell_{i} m_{i}}^{(1)}\) is readily found to be the sum of two pieces one of which, insofar as is its angular dependence is concerned, being proportional to
\[
\begin{equation*}
\mathrm{T}_{\mathrm{jk}}^{\mathrm{LO}}(\overrightarrow{\mathrm{I}}) \tag{7}
\end{equation*}
\]
and the other to
\[
\mathrm{T}_{\mathrm{kj}}^{\mathrm{LO}}(\overrightarrow{\mathrm{I}})
\]
where
\[
\begin{equation*}
T_{j k}^{I M M}(\vec{I})=\sum_{m_{k}=-k}^{k} \sum_{m_{j}=-j}^{j} \int \frac{d \hat{2}}{r_{12}} Y_{j m_{j}}(\hat{1})\left(Y_{j m_{j}}, Y_{L M} Y_{k m_{k}}\right) Y_{l_{i} m_{i}}(\hat{2}) Y_{k m_{k}}^{*}(\hat{2}) \tag{8}
\end{equation*}
\]

We will now show that \({ }_{j k}^{I M}(\overrightarrow{1})\) is of the form
\[
\begin{equation*}
T_{j k}^{L M}(\overrightarrow{\mathrm{I}})=\sum_{\lambda} \rho_{\lambda l_{i}}^{j k L_{1}}\left(r_{1}\right) \sum_{\mu=-\lambda}^{\lambda} Y_{\lambda \mu}(\hat{1})\left(Y_{\lambda \mu}, Y_{L M} Y_{l_{i} m_{i}}\right) \tag{9}
\end{equation*}
\]
where, as indicated, \(\rho\) is independent of \(\mu, M\) and \(m_{i}\). It is then easy to show that this is sufficient to ensure the consistency of the exchange term since with this result the \(m_{i}, \mu\) dependent coefficients \(\left(Y_{\lambda \mu}, P_{L} Y_{\ell_{i} m_{i}}\right)\) will correctly cancel out of the equations one gets by equating the coefficient of each spherical harmonic separately equal to zero, the result in each case then being the same set of coupled equations for the radial functions \(R_{j n_{i} \ell_{i}}^{(1)}\).

To derive (9) we note that from the spherical harmonic addition theorem we have, to within an \(m_{i}, M\) independent factor that
\[
T_{j k}^{L M}(\overrightarrow{1})=\int \frac{d \hat{2}}{r_{12}} d \hat{3} P_{j}(\hat{1} \cdot \hat{3}) P_{k}(\hat{2} \cdot \hat{3}) Y_{L M}(\hat{3}) Y_{\ell_{i} m_{i}}(\hat{2})
\]
in which form it is clear \({ }^{4}\) that under a rotation \(T_{j k}^{L M}(\overrightarrow{1})\) transforms like \(Y_{L M}(\hat{1}) Y_{\ell_{i} m_{i}}(\hat{1})\). Hence writing \(T_{j k}^{L M}(\overrightarrow{1})\) as
\[
T_{j k}^{L M}(\vec{I})=\sum_{\lambda} \sum_{\mu} Y_{\lambda \mu}(\hat{1})\left(Y_{\lambda \mu}, T_{j k}^{L M}\right)
\]
this means that (the argument is essentially the same as that in footnote 3)
\[
\sum_{\mu} Y_{\lambda \mu}(\hat{I})\left(Y_{\lambda \mu}, T_{j k}^{L M}\right)=\hat{4 \pi \mu} \int \hat{\mathrm{~d}}^{2} \mathrm{P}_{\lambda}(\hat{I} \cdot \hat{2}) \mathrm{T}_{\mathrm{jk}}^{\mathrm{LM}}(\hat{2)}
\]
transforms like \(\sum_{\mu} Y_{\lambda \mu}(\hat{1})\left(Y_{\lambda \mu}, Y_{L M} Y_{\ell_{i}} m_{i}\right)\) which in turn means, since the \(Y_{\lambda \mu}\) for a given \(\lambda\) yield an irreducible representation of the rotation group, that to within a \(\mu, M\) and \(m_{i}\) independent factor
\[
\begin{equation*}
\left(Y_{\lambda \mu}, T_{j k}^{L M}\right)=\left(Y_{\lambda \mu}, Y_{L M^{\prime}} Y_{\ell_{i} m_{i}}\right) \tag{1.0}
\end{equation*}
\]
which proves the point. The coefficients \(\rho_{\lambda \ell_{\perp}}^{j k L}\) are given explicitly in the appendix.

Our interest in these questions was aroused by conversations
with R. Ahlberg, Also it is a pleasure to acknowledge further correspondence with him and Dr. Goscinski。

\section*{FOOTNOTES AND REFERENCES}
1. R. Ahlberg and O. Goscinski, Jo Phys. B. Aton. Molec. Phys. 6, 2254 (1973)。
2. See for example A. Dalgarno, Adv. in Phys. 21, \(_{1} 281\) (1962) Eq. (98).
3. Let \(\hat{R} \hat{1}\) denote the effect of rotating \(\hat{1}\). Then evidently
\[
\begin{aligned}
Q_{j k}^{L M}(\hat{R I}) & =\int \hat{d 2} P_{j}(\hat{R I} \cdot \hat{2}) Y_{L M}(\hat{2}) P_{k}(\hat{R 1} \cdot \hat{2}) \\
& =\int d \hat{2} P_{j}\left(\hat{1} \cdot R^{-1} \hat{2}^{2}\right) Y_{L M}(\hat{2}) P_{k}\left(\hat{1} \cdot R^{-1} \hat{2}\right)
\end{aligned}
\]
which upon changing variables according te \(:{ }^{\hat{1} \hat{2}-2}\)
becomes
\(Q_{j k}^{L M}(\hat{R 1})=\int \mathrm{d} \hat{2} P_{j}(\hat{1} \cdot \hat{2}) Y_{L M}(K \hat{2}) P_{k}(\hat{1} \cdot \hat{2})\)
which proves the point since the transfors \(1.10 n=0\) efficients are independent of coordinates.
4. Proceeding similarly as in 3 one finds thiai.
\(T_{j k}^{L M}(R \overrightarrow{1})=\int \frac{\mathrm{d} \hat{2}}{r_{12}} \mathrm{~d} \hat{3} P_{j}(\hat{1} \cdot \hat{3}) P_{k}(\hat{2} \cdot \hat{3}) Y_{L M}(R \hat{3}) Y_{\ell_{i} m_{i}}(R \hat{2})\)
which proves the point since the transforraition coefficients are independent of coordinates.
5. A. Messiah, Quantum Mechanics, Vol. II (Joha Viley \& Sons, Inc., New York, 1966), Appendix C.

\section*{APPENDIX}

We use the conventions of Messiah \({ }^{5}\) in evaluating \(Q_{j k}^{L M}(\hat{1})\) and \(\rho_{\lambda l_{i}}^{j k L}\left(r_{1}\right)\) and refer to his equations. We evaluate \(Q_{j k}^{L M}(\hat{1})\) of (6) by the spherical harmonic composition relation. From Messiah C. 1.6 and C.17b one obtains

The identity C. 15 a gives the desired result
\[
\begin{equation*}
Q_{j k}^{\mathrm{LM}}(\hat{1})=4 \pi\left(\left(_{000}^{\mathrm{jkL}}\right)^{2} \mathrm{Y}_{\mathrm{LM}}(\hat{1})\right. \tag{A-1}
\end{equation*}
\]

To evaluate \(\rho_{\lambda l_{i}}^{j k L}\left(r_{1}\right)\) we note by comparing (9) and (10) it follows that
\[
\begin{equation*}
\rho_{\lambda l_{i}}^{j k L}\left(r_{1}\right)\left(Y_{\lambda \mu}, Y_{L M} Y_{\ell_{i}} m_{i}\right)=\left(Y_{\lambda \mu}, T_{j k}^{L M}\right) \tag{A-2}
\end{equation*}
\]

We now evaluate the right hand side of this equation. From the definition (8) and Messiah C. 16 it follows that
\[
\begin{aligned}
& \left(Y_{\lambda \mu}, T_{j k}^{L M}\right)=(-)^{\mu}(4 \pi)^{-1 / 2}\left(2_{j}+1\right)(2 k+1)\left[(2 \lambda+1)\left(2 \ell_{i}+1\right)(2 L+1)\right]^{1 / 2}\binom{j k L}{000}
\end{aligned}
\]
mince
\[
\mathscr{E}_{g_{3}}\left(x_{2}\right)=\int \frac{\Sigma_{<}^{e^{3}}}{x_{y}^{l}+1} r_{2}^{2} d r_{2}
\]

Use of Messiah C. 33 then yields

มกผอver
and therefore, from ( \(\mathrm{A}-2\) ), we have as our final result```

