## ON THE TESSERALHARMONICS RESONANCE PROBLEM IN ARTIFICIALSATELLITE THEORY

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## TABLE OF CONTENTS

Page
ABSTRACT ..... v
1 INTRODUCTION ..... 1
2 CHOICE OF THE DISTURBING FUNCTION AND EQUATIONS OF MOTION ..... 3
2. 1 Resonant Part of a Tesseral Harmonic ..... 5
2.2 The Equations of Motion ..... 8
3 HORI'S PERTURBATION METHOD BY LIE SERIES ..... 11
4 DETAILED EXPRESSIONS OF THE PERTURBATIONS ..... 25
4.1 Expression of $\mathrm{S}_{1 / 2}$. ..... 25
4.2 Expression of $Z_{1}$ and Its Derivatives. ..... 29
4.3 Expression of $\mathrm{Z}_{2}$ and Its Derivatives. ..... 33
5 REMOVAL OF LONG-PERIOD TERMS. ..... 41
6 REFERENCES AND BIBLJOGRAPHY ..... 47
APPENDIX A: Derivatives of $D\left(k_{1}\right)$ and $\psi\left(\mathrm{k}_{1}\right)$ ..... A-1


#### Abstract

The longitude-dependent part of the geopotential usually gives rise only to shortperiod effects in the motion of an artificial satellite. However, when the motion of the satellite is commensurable with that of the earth, the path of the satellite repeats itself relative to the earth and perturbations build up at each passage of the satellite in the same spot, so that there can be important long-period effects.

In order to take these effects into account in deriving a theoretical solution to the equations of motion of an artificial satellite, it is necessary to select terms in the longitude-dependent part of the geopotential that will contribute significantly to the perturbations. We have tried to make a selection that is valid in a general case, regardless of the initial eccentricity of the orbit and of the order of the resonance.

The solution to the equations of motion of an artificial satellite, in a geopotential thus determined, is then derived by using Hori's method by Lie series, which, by its properties regarding canonical invariance, has proved advantageous in the classical theory.


RESUME

La partie du géopotentiel, dépendant de la longitude, ne provoque généralement que des effets à courte période dans le mouvement d'un satellite artificiel. Cependant, quand le mouvement du satellite est commensurable avec celui de la terre, le parcours du satellite se répète par rapport à la terre et les perturbations s'ajoutent à chaque passage du satellite au même endroit, et peuvent ainsi créer d'importants effets à longue période.

Pour tenir compte de ces effets en dérivant une solution théorique des équations du mouvement d'un satellite artificiel, il est nécessaire de choisir ceux des termes de la partie du géopotentiel dépendant de la longitude, qui contribueront d'une façon significative aux perturbations. On a essayé de faire un choix qui soit valable pour le cas général, quelques soient la première excentricité de l'orbite et l'ordre et l'ordre de la résonance.

On a ensuite dérivé la solution des équations du mouvement d'un satellite artificiel dans un géopotentiel ainsi déterminé, en employant la méthode de Hori par la série de Lie, qui a été prouvée avantageuse dans la théorie classique, grâce à ses propriétés concernant l'invariance canonique.

## КОНСПЕКТ

Зависящая от долготы составляюцая геопотенциала обычно вызывает лишь короткопериодные изменения в движении искусственного спутника земли. Однако, если движение искусственного спутника соизмеримо с движением земли, то путв искусственного спутника относительно земли и пертурєаций, построенных на каждом отрезке пути искусственного спутника в той же точке, повторяется; так что воздействия уже могут Єыть долгопериодНБIM .

Для учета этих воздействий при получении теоротического рєшения уравнений движения искусственного спутника неоєходимо выбрать те члены зависящей от долготы составляющей геопотенциала, которые являются важной частью пертурєаций. Міы пытались осуцествить выЄор таким оєразом, чтоЄы он годился для обцего случая, в независимости от начального эксцентриситета и порядка резонанса.

Затем в. геопотенциале, определенном таким спосоєом, выводится репение уравнений движения искусственного спутника по методу Хори с помоцью серии Ли. Этот метод оказался полезным в классической теории єлагодаря своей установившейся неизменности.

# ON THE TESSERAL-HARMONICS RESONANCE PROBLEM IN ARTIFICIAL-SATELLITE THEORY 

Barbara A. Romanowicz

## 1. INTRODUCTION

If the gravitational potential of the earth is expanded in terms of Legendre polynomials and functions, then in order to obtain a good approximation in the determination of the orbit of an artificial satellite, it is usually sufficient to consider the zonal, lon-gitude-free terms of the expansion. However, the influence of the tesseral terms becomes important in the case when the mean motion of the satellite and the rate of rotation of the earth around its axis are in a simple ratio.

This problem has been examined from different points of view. Morando (1963) considered the particular, important case of a 24 -hour satellite; Allan (1967, 1973) worked out perturbations due to resonance in a more general case and to the first order, by using Lagrange's equations. On the other hand, Garfinkel (1974) considered the abstract, mathematical problem of "ideal resonance." We have attempted to derive expressions for the perturbations in the motion of an artificial satellite due to the most general tesseral-harmonics resonance. The calculations are carried out to order $3 / 2$ in the small parameter of the expansion of the perturbing function, and the perturbation method used is that of Hori, which is canonically invariant and avoids mixing of old and new variables, as opposed to the generally used method of Von Zeipel.

[^0]
## 2. CHOICE OF THE DISTURBING FUNCTION AND EQUATIONS OF MOTION

The gravitational potential of the earth at exterior points can be expressed as follows:
$U(r, \theta, \lambda)=\frac{\mu}{r}\left[1-\sum_{n=2}^{\infty} J_{n}\left(\frac{R}{r}\right)^{n} P_{n}(\cos \theta)+\sum_{n=2}^{\infty} \sum_{m=1}^{n} J_{n, m}\left(\frac{R}{r}\right)^{n} P_{n, m}(\cos \theta) \cos m(\lambda-\lambda n, m)\right]$,
where ( $\mathrm{r}, \theta, \lambda$ ) are spherical polar coordinates relative to the center of mass of the earth, the axis of rotation being the pole of coordinates; $\mu$ is the gravitational constant $G$ times the mass of the earth; $R$ is the mean equatorial radius of the earth; $P_{n}(z)$ is the $n^{\text {th }}$ Legendre polynomial; $P_{n, m}(z)$ are associated Legendre functions:

$$
P_{n, m}(z)=\left(1-z^{2}\right)^{m / 2} \frac{d^{n} P_{n}(z)}{d z^{n}} ;
$$

and $J_{n}$ and $J_{n, m}$ are dimensionless coefficients related to the normalized coefficients $\overline{C_{n, m}}, \overline{S_{n, m}}$ by:

$$
\begin{align*}
& J_{n, m}=\overline{J_{n, m}} \sqrt{2(2 n+1)(n-m)!/(n+m)!}, \quad m \neq 0 \\
& -J_{n}=C_{n, 0}=\sqrt{2 n+1} \overline{C_{n, 0}},  \tag{2}\\
& J_{n, m} \cos m\left(\lambda-\lambda_{n, m}\right)=\overline{C_{n, m}} \cos m \lambda+\overline{S_{n, m}} \sin m \lambda
\end{align*}
$$

It can be assumed that the largest perturbing forces are due to the leading zonal harmonic, which contains $\mathrm{J}_{2}$ and corresponds to the oblateness of the earth. The remaining zonal harmonics have much smaller effects and are not considered here. The tesseral harmonics, containing $J_{n, m}$, are the longitude-dependent terms and usually give rise only to short-period effects because they all contain the mean anomaly
and the sidereal angle in their arguments. In the case of resonance, however, a certain number of these terms can produce effects of large amplitude and very long period. It is therefore necessary to select appropriately those tesseral harmonics to be included in our disturbing function.

The general tesseral harmonic can be developed in terms of the osculating elliptic elements ( $a, e, I, \Omega, \omega, M$ ), referred to the equator of the earth, in the form (Kaula, 1966):

$$
\begin{align*}
V_{n, m}=\left(\frac{\mu}{R}\right)\left(\frac{R}{a}\right)^{n+1} J_{n, m}\{ & \sum_{p=0}^{p=n} F_{n, m, p^{(I)}} \sum_{q} G_{n, p, q^{(e)} \cos [(n-2 p) \omega+(n-2 p+q) M} \\
& \left.\left.+m\left(\Omega-v t-\lambda_{n, m}\right)\right]\right\} \tag{3}
\end{align*}
$$

where $v$ is the angular velocity of the rotation of the earth, $t$ is the time, and $F_{n, m, p}{ }^{(I)}$ and $G_{n, p, q^{(e)}}$ are, respectively, the inclination and eccentricity functions as defined by Kaula (1966). The functions $G_{n, p, q^{(e)}}$ are of order $|q|$ in eccentricity.

Resonance occurs when a pair ( $\alpha, \beta$ ) of mutually prime integers exists such that the satellite performs $\beta$ nodal periods while the earth rotates a times relative to the precessing satellite's orbit plane. This can be expressed by

$$
\begin{equation*}
\mathrm{a}(\dot{\omega}+\dot{\mathrm{M}})=\beta(\nu-\dot{\Omega}), \tag{4}
\end{equation*}
$$

where $\dot{\omega}, \dot{\mathrm{M}}$, and $\dot{\Omega}$ are the rates of change with time of $\omega$, M , and $\Omega$, respectively. The corresponding slowly varying arguments are of the form

$$
\mathrm{k}_{1} \Phi_{a, \beta}+\text { const }
$$

where

$$
\Phi_{\alpha, \beta}=a(\omega+M)+\beta(\Omega-v t)
$$

is called the resonant variable and $\mathrm{k}_{1}=1,2,3, \ldots$.

Considering that $\dot{\omega}$ is a small quantity and that the general argument in a tesseral harmonic is

$$
\phi=(n-2 p) \omega+(n-2 p+q) M+m\left(\Omega-\nu t-\lambda_{n, m}\right)
$$

we shall select the tesseral harmonics by keeping those containing arguments such that

$$
\begin{array}{ll}
n-2 p+q=k_{1} a, & k_{1}=1,2,3, \ldots \\
m=k_{1} \beta,
\end{array}
$$

Since $\mathrm{n} \geq \mathrm{m}$ and since lower order tesseral harmonics are bound to have larger effects because of the factor $(R / r)^{n}$, where $R / r<1$, it can be assumed that it is enough to consider only the cases $\mathrm{k}_{1}=1,2$, and 3 .

If $\mathrm{k}_{1}=1$, then $\mathrm{m}=\beta$, and the tesseral harmonics to be considered are

$$
\mathrm{v}_{\beta+\mathrm{k}_{0}, \beta}, \quad \mathrm{k}_{0}=0,1,2, \ldots
$$

In a general manner, the tesseral harmonics to be kept are

$$
\mathrm{V}_{\mathrm{k}_{1} \beta+\mathrm{k}_{0}, \mathrm{k}_{1} \beta}
$$

where $\mathrm{k}_{1}=1,2,3, \ldots$ and for each $\mathrm{k}_{1}$, the index $\mathrm{k}_{0}$ takes values $0,1,2, \ldots$.

### 2.1 Resonant Part of a Tesseral Harmonic

Once we have decided which tesseral harmonics to keep for our disturbing function, we then need to extract for each its most important "resonant" part.

If $V_{n, m}$ is a selected tesseral harmonic, with $n=k_{1} \beta+k_{0}$ and $m=k_{1} \beta$, we look for terms containing "resonant" arguments by solving equation (5); that is, since we already know $m=k_{1} \beta$, we have

$$
n-2 p+q=k_{1} a,
$$

or, in this case,

$$
k_{1}(\beta-a)+k_{0}=2 p-q
$$

As the condition $0 \leq p \leq k_{1} \beta+k_{0}$ must be satisfied, the following pairs ( $p, q$ ) are solutions:

$$
\begin{array}{ll}
p=0, & q=k_{1}(a-\beta)-k_{0}, \\
p=1, & q=k_{1}(a-\beta)-k_{0}+2, \\
\cdot & \cdot \\
\cdot & \cdot  \tag{6}\\
p=x, & q=k_{1}(a-\beta)-k_{0}+2 x, \\
p=k_{1} \beta+k_{0}, & q=k_{1}(a+\beta)+k_{0},
\end{array}
$$

and the resonant part of the tesseral harmonic $V_{n, m}$ is
$R_{n, m}=\mu \frac{R^{n}}{a^{n+1}} J_{n, m} \sum_{x=0}^{\beta k_{1}+k_{0}} F_{n, m, p_{x}}^{(I)} G_{n, p_{x}, q_{x}}(e) \cos \left(k_{1} \Phi_{a, \beta}-q_{x} \omega-m \lambda n, m\right)$,
where

$$
q_{x}=k_{1}(a-\beta)-k_{0}+2 x
$$

and

$$
\mathrm{p}_{\mathrm{x}}=\mathrm{x}
$$

Thus,

$$
V_{n, m}=R_{n, m}+V_{n, m}^{\prime}
$$

where $V_{n, m}^{\prime}$ is the residual, the effect of which is much smaller than that of $R_{n, m}$.
It will be convenient to express the sum of the resonant parts of all the selected tesseral harmonics in the following manner:

$$
\sum R_{n, m}=\sum_{k_{1}=1}^{3} D\left(k_{1}\right) \cos 2 \theta_{1}\left(k_{1}\right)
$$

where

$$
D\left(k_{1}\right) \cos 2 \theta_{1}\left(k_{1}\right)=\sum_{k_{0}\left(k_{1}\right)} R_{k_{1} \beta+k_{0}}, k_{1} \beta
$$

the sum over $\mathrm{k}_{0}\left(\mathrm{k}_{1}\right)$ meaning that we have taken into account all the values of $\mathrm{k}_{0}$ when $\mathrm{k}_{1}$ has a given value. In order to obtain $\mathrm{D}\left(\mathrm{k}_{1}\right)$ and $\theta_{1}\left(\mathrm{k}_{1}\right)$, we can write

$$
\begin{aligned}
& S\left(k_{1}, k_{0}\right)=\frac{\mu R^{n_{J}} n_{2} m}{a^{n+1}}, \quad n=k_{1} \beta+k_{0}, \\
& A\left(k_{1}, k_{0}, x\right)=F_{n, m, x}(I) G_{n, x, q_{x}}^{(e)}, \\
& D\left(k_{1}\right) \exp 2 i \theta_{1}\left(k_{1}\right)=\sum_{k_{0}\left(k_{1}\right)} S\left(k_{1}, k_{0}\right) \sum_{x=0}^{x=n} A\left(k_{1}, k_{0}, x\right) \exp i\left(k_{1} \Phi_{a, \beta}-q \omega-k_{1} \beta \lambda_{n, m}\right),
\end{aligned}
$$

where

$$
q_{x}=k_{1}(a-\beta)-k_{0}+2 x
$$

If we write

$$
\Lambda_{k_{1}, k_{0}}=k_{1} \beta \lambda_{k_{1} \beta+k_{0}, k_{1} \beta}
$$

and

$$
2 \theta_{1}\left(k_{1}\right)=k_{1} \Phi_{a, \beta}-k_{1}(a-\beta)^{\omega}+\psi\left(k_{1}\right),
$$

then

$$
D\left(k_{1}\right) \exp i \psi\left(k_{1}\right)=\sum_{k_{0}\left(k_{1}\right)} S\left(k_{1}, k_{0}\right) \sum_{x=0}^{n} A\left(k_{1}, k_{0}, x\right) \exp i\left[\left(k_{0}-2 x\right) \omega-\Lambda_{k_{1}}, k_{0}\right]
$$

Thus,
$B\left(k_{1}\right)=D\left(k_{1}\right) \cos \psi\left(k_{1}\right)=\sum_{k_{0}} \sum_{\substack{x=0}}^{x=k_{1} \beta+k_{0}} S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right) \cos \left[\left(k_{0}-2 x\right) \omega-\Lambda_{k_{1}}, k_{0}\right]$,
$C\left(k_{1}\right)=D\left(k_{1}\right) \sin \psi\left(k_{1}\right)=\sum_{k_{0}} \sum_{x=0}^{\beta+k_{0}} S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right) \sin \left[\left(k_{0}-2 x\right) \omega-\Lambda_{\left.k_{1}, k_{0}\right]}\right.$,
$\mathrm{D}\left(\mathrm{k}_{1}\right)^{2}=\mathrm{B}\left(\mathrm{k}_{1}\right)^{2}+\mathrm{C}\left(\mathrm{k}_{1}\right)^{2}$,
in which $\mathrm{D}\left(\mathrm{k}_{1}\right)$ and $\psi\left(\mathrm{k}_{1}\right)$ are functions of $\mathrm{a}, \mathrm{e}, \mathrm{I}$, and $\Omega$. Appendix A gives expressions for the first and second derivatives of $D\left(k_{1}\right)$ and $\psi\left(\mathrm{k}_{1}\right)$ with respect to the canonical set of Delaunay variables.

Finally, the potential in which the satellite moves is taken to be

$$
\begin{equation*}
\mathrm{V}=\frac{\mu}{\mathrm{r}}-J_{2} \frac{\mathrm{R}^{2}}{\mathrm{r}^{3}} \mathrm{P}_{2}(\cos \theta)+\sum \mathrm{R}_{\mathrm{n}, \mathrm{~m}}+\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}^{\prime}, \tag{10}
\end{equation*}
$$

the sums being taken over all the tesseral harmonics selected.

## 2. 2 The Equations of Motion

If we consider the canonical set of Delaunay variables,

$$
\begin{array}{lll}
L_{D}=\sqrt{\mu a}, & G_{D}=\sqrt{\mu a\left(1-e^{2}\right)}, & H_{D}=G_{D} \cos i \\
\ell_{D}=M, & g_{D}=\omega, & h_{D}=\Omega, \tag{11}
\end{array}
$$

then the equations of motion of the satellite are

$$
\begin{array}{lll}
\frac{d L_{D}}{d t}=\frac{\partial F_{D}}{\partial \ell_{D}}, & \frac{d G_{D}}{d t}=\frac{\partial F_{D}}{\partial g_{D}}, & \frac{d H_{D}}{d t}=\frac{\partial F_{D}}{\partial \mathrm{~h}_{\mathrm{D}}}, \\
\frac{d \ell_{D}}{d t}=-\frac{\partial F_{D}}{\partial L_{D}}, & \frac{d g_{D}}{d t}=-\frac{\partial F_{D}}{\partial G_{D}}, & \frac{d h_{D}}{d t}=-\frac{\partial F_{D}}{\partial H_{D}}, \tag{12}
\end{array}
$$

where $F_{D}$ is the Hamiltonian of the problem:

$$
\begin{equation*}
\mathrm{F}_{\mathrm{D}}=\frac{\mu^{2}}{2 \mathrm{~L}_{\mathrm{D}}^{2}}+\mathrm{F}_{1}+\mathrm{F}_{2} \tag{13a}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{F}_{1}=-\mathrm{J}_{2} \mu \frac{\mathrm{R}^{2}}{\mathrm{r}^{3}} \mathbf{P}_{2}(\cos \theta) \tag{13b}
\end{equation*}
$$

which can be expressed in Delaunay variables, and

$$
\begin{equation*}
\mathrm{F}_{2}=\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}} \tag{13c}
\end{equation*}
$$

Here, the Hamiltonian $F_{D}$ depends explicitly on time. To avoid this, we can perform a canonical transformation so that the new angular variables will be

$$
\ell=\ell_{\mathrm{D}}, \quad \mathrm{~g}=\mathrm{g}_{\mathrm{D}}, \quad \mathrm{~h}=\mathrm{h}_{\mathrm{D}}-\nu \mathrm{t} .
$$

To find the new momenta and the new Hamiltonian, we have to solve

$$
L d \ell+G d g+H d h-L_{D} d \ell-G_{D} d g-H_{D} d h-\left(F_{D}-F\right) d t=d V,
$$

where dV is the differential of a function. We keep the solution corresponding to $d V \equiv 0$ as follows:

$$
\begin{aligned}
& \mathrm{L}=\mathrm{L}_{\mathrm{D}} \\
& \mathrm{G}=\mathrm{G}_{\mathrm{D}} \\
& \mathrm{H}=\mathrm{H}_{\mathrm{D}} \\
& \mathrm{~F}=\mathrm{F}_{\mathrm{D}}+\nu \mathrm{H}
\end{aligned}
$$

The equations of motion then become

$$
\begin{array}{lll}
\frac{\mathrm{dL}}{\mathrm{dt}}=\frac{\partial \mathrm{F}}{\partial \ell} \quad, & \frac{\mathrm{dG}}{\mathrm{dt}}=\frac{\partial \mathrm{F}}{\partial \mathrm{~g}}, & \frac{\mathrm{dH}}{\mathrm{dt}}=\frac{\partial \mathrm{F}}{\partial \mathrm{~h}}, \\
\frac{\mathrm{~d} \ell}{\mathrm{dt}}=-\frac{\partial \mathrm{F}}{\partial \mathrm{~L}}, & \frac{\mathrm{dg}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}}{\partial \mathrm{G}}, & \frac{\mathrm{dh}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}}{\partial \mathrm{H}}, \tag{14}
\end{array}
$$

with $\mathrm{F}=\mathrm{F}_{0}+\mathrm{F}_{1}+\mathrm{F}_{2}$ :

$$
\begin{align*}
& \mathrm{F}_{0}=\frac{\mu^{2}}{2 \mathrm{~L}^{2}}+\nu \mathrm{H} \\
& \mathrm{~F}_{1}=-\mathrm{J}_{2} \mu \frac{\mathrm{R}^{2}}{\mathrm{r}^{3}}\left[\frac{1}{4} \frac{\mathrm{H}^{2}}{\mathrm{G}^{2}}-\frac{3}{4}\left(1-\frac{\mathrm{H}^{2}}{\mathrm{G}^{2}}\right) \cos 2 \mathrm{u}\right]  \tag{15}\\
& \mathrm{F}_{2}=\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}
\end{align*}
$$

in the above, $u=g+f$, and $f$ is the true anomaly.

The Hamiltonian is now expanded in terms of powers of the small parameter $\mathrm{J}_{2}$ :

$$
\begin{aligned}
& \mathrm{F}_{1}=\mathrm{O}\left(\mathrm{~J}_{2}\right) \\
& \mathrm{F}_{2}=\mathrm{O}\left(\mathrm{~J}_{2}^{2}\right)
\end{aligned}
$$

## 3. HORI'S PERTURBATION METHOD BY LIE SERIES

In order to remove the variables $\ell$ and h from the Hamiltonian, we shall apply Hori's method by the use of Lie series (Hori, 1966; Aksnes, 1970).

The greatest advantage of this method is the fact that the Poisson brackets that appear are canonically invariant, which means that each bracket can be calculated in terms of the most convenient set of canonical variables. Also, this method avoids mixing old and new variables.

Let us consider a system of canonical equations

$$
\begin{equation*}
\frac{d x_{j}}{d t}=\frac{\partial F}{\partial y_{j}} \quad, \quad \frac{d y_{j}}{d t}=-\frac{\partial F}{\partial x_{j}} \quad, \quad j=1, \ldots, n \tag{16}
\end{equation*}
$$

where the Hamiltonian $F$ is developed in powers of a small parameter $\epsilon$,

$$
\mathrm{F}=\sum_{\mathrm{k}=0}^{\infty} \mathrm{F}_{\mathrm{k}}, \quad \mathrm{~F}_{\mathrm{k}}=\mathrm{O}\left(\epsilon^{\mathrm{k}}\right) ;
$$

and let us also consider a Lie transformation (Hori, 1966)

$$
\begin{aligned}
& \left(x_{j}, y_{j}\right) \longrightarrow\left(x_{j}^{\prime}, y_{j}^{\prime}\right), \\
& x_{j}=x_{j}^{\prime}+\frac{\partial S}{\partial y_{j}^{\prime}}+\frac{1}{2}\left\{\frac{\partial S}{\partial y_{j}^{\prime}}, s\right\}+\frac{1}{6}\left\{\left\{\frac{\partial S}{\partial y_{j}^{\prime}}, s\right\}, s\right\}+\cdots, \\
& y=y_{j}^{\prime}-\frac{\partial S}{\partial x_{j}^{\prime}}-\frac{1}{2}\left\{\frac{\partial S}{\partial x_{j}^{\prime}}, s\right\}-\frac{1}{6}\left\{\left\{\frac{\partial S}{\partial x_{j}^{\prime}}, s\right\}, s\right\}+\cdots
\end{aligned}
$$

where $\{$,$\} are Poisson brackets and S$ is a function of $\left(x^{\prime}, y^{\prime}\right)$ of order 1 in $\epsilon$ :

$$
S=\sum_{K=1}^{\infty} S_{K}
$$

Then, if $f(x, y)$ is a function of $\left(x_{j}, y_{j}\right), j=1, n$, it can be expressed in terms of the new variables ( $\mathrm{x}_{\mathrm{j}}^{\prime}, \mathrm{y}_{\mathrm{j}}^{\prime}$ ) as follows:

$$
f(x, y)=f\left(x^{\prime}, y^{\prime}\right)+\{f, s\}+\frac{1}{2}\{\{f, s\}, s\}+\frac{1}{6}\{\{\{f, s\}, s\}, s\}+\ldots
$$

The equations of motion then become

$$
\begin{equation*}
\frac{\mathrm{dx}_{\mathrm{j}}^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{\prime}}{\partial \mathrm{y}_{\mathrm{j}}^{\prime}} \quad, \quad \frac{\mathrm{dy}_{\mathrm{j}}^{\prime}}{\mathrm{dt}}=-\frac{\partial F^{\prime}}{\partial x_{j}^{\prime}} \tag{17}
\end{equation*}
$$

where $F^{\prime}$ is the new Hamiltonian:

$$
\mathrm{F}^{\prime}=\sum_{\mathrm{k}=0}^{\infty} \mathrm{F}_{\mathrm{k}}^{\prime}
$$

If we assume that $F$ does not depend explicitly on time $t$, we can write the energy integral

$$
F(x, y)=F^{\prime}\left(x^{\prime}, y^{\prime}\right)
$$

Expanding in terms of $\epsilon$ and collecting terms of the same order of magnitude, we obtain

$$
\begin{align*}
& \mathrm{F}_{0}=\mathrm{F}_{0}^{\prime}, \\
& \mathrm{F}_{1}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{1}\right\}=\mathrm{F}_{1}^{\prime},  \tag{18}\\
& \left\{\mathrm{F}_{0}, \mathrm{~S}_{2}\right\}+\frac{1}{2}\left\{\mathrm{~F}_{1}+\mathrm{F}_{1}^{\prime}, \mathrm{S}_{1}\right\}+\mathrm{F}_{2}=\mathrm{F}_{2}^{\prime}
\end{align*}
$$

In order to average these equations, we introduce a pseudo time $\mathrm{t}^{\prime}$ :

$$
\frac{d x_{j}^{\prime}}{d t^{\prime}}=\frac{\partial F_{0}}{\partial y_{j}^{\prime}}, \quad \frac{d y_{j}^{\prime}}{d t^{\prime}}=-\frac{\partial F_{0}}{\partial x_{j}^{\prime}}
$$

If this system has a solution, then

$$
\left\{\mathrm{F}_{0}, \mathrm{~S}_{\mathrm{k}}\right\}=-\frac{\mathrm{dS}}{\mathrm{k}} \mathrm{dt}^{\prime} \quad, \quad \mathrm{k} \geq 1
$$

and if $\mathrm{A}\left(\mathrm{t}^{\prime}\right)$ is a periodic function of $\mathrm{t}^{\prime}$, with period T , then

$$
\begin{equation*}
A_{S}=\frac{1}{T} \int_{0}^{T} A\left(t^{\prime}\right) d t^{\prime} \tag{19a}
\end{equation*}
$$

is the secular part of $A$ and

$$
\begin{equation*}
A_{p}=A-A_{S} \tag{19b}
\end{equation*}
$$

is the periodic part of A .

We can remove $t^{\prime}$ from $F^{\prime}$ by applying the averaging technique [eqs. (19)] to equations (18) (Hori, 1966; Aksnes, 1970). Then $S_{k}$ and $F_{k}^{\prime}$ are uniquely determined by the following set of equations:

$$
\begin{align*}
& \mathrm{F}_{0}^{\prime}=\mathrm{F}_{0}, \\
& \mathrm{~F}_{1}^{\prime}=\mathrm{F}_{1 \mathrm{~s}}, \\
& \mathrm{~S}_{1}=\int \mathrm{F}_{1 \mathrm{p}} \mathrm{~d} \mathrm{t}^{\prime},  \tag{20}\\
& \mathrm{F}_{2}^{\prime}=\mathrm{F}_{2 \mathrm{~s}}+\frac{1}{2}\left\{\mathrm{~F}_{1}+\mathrm{F}_{1}^{\prime}, \mathrm{S}_{1}\right\}_{\mathrm{s}}, \\
& \mathrm{~S}_{2}=\int \mathrm{F}_{2 \mathrm{p}}+\frac{1}{2}\left\{\mathrm{~F}_{1}+\mathrm{F}_{1}^{\prime}, \mathrm{S}_{1}\right\}_{\mathrm{p}} \mathrm{~d} t^{\prime},
\end{align*}
$$

In the present case,

$$
\mathrm{F}_{2}=\sum_{\mathrm{k}_{1}} \mathrm{D}\left(\mathrm{k}_{1}\right) \cos 2 \theta_{1}\left(\mathrm{k}_{1}\right)+\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}^{\prime}
$$

and the pseudo time $t^{\prime}$ is defined by

$$
\begin{aligned}
& \frac{\mathrm{L}^{\prime}}{\mathrm{dt}^{\prime}}=-\frac{\partial \mathrm{F}_{0}}{\partial \mathrm{~L}^{\prime}}=\frac{\mu^{2}}{\mathrm{~L}^{\prime 3}}, \\
& \frac{\mathrm{dh}^{\prime}}{\mathrm{dt}^{\prime}}=-\frac{\partial \mathrm{F}_{0}}{\partial \mathrm{H}^{\prime}}=-v,
\end{aligned}
$$

so that

$$
\begin{equation*}
\mathrm{dt}^{\prime}=\frac{\mathrm{L}^{\prime 3}}{\mu^{2}} \mathrm{~d} \ell^{\prime}-\frac{1}{\nu} \mathrm{dh}^{\prime} \tag{21}
\end{equation*}
$$

In the process of calculating $\mathrm{S}_{2}$, we have to evaluate the following integral:

$$
\int D\left(k_{1}\right) \cos 2 \theta_{1}\left(k_{1}\right) d t^{\prime}=\frac{D\left(k_{1}\right) \sin 2 \theta^{\prime}\left(k_{1}\right)}{2 \delta}
$$

where

$$
\delta=\left[\left(\frac{\mu^{2}}{L^{\prime}}\right) a-\nu \beta\right] \mathbf{k}_{1}
$$

In the case of resonance of order $(a, \beta)$,

$$
\mathrm{n}=\frac{\mu^{2}}{\mathrm{~L}^{3^{3}}} \simeq \frac{\nu \beta}{\mathrm{a}}
$$

where $n$ is the mean motion of the satellite; thus, $\delta$ is a small quantity appearing in the denominator, and the preceding method cannot be used.

To avoid this situation, we shall introduce a procedure that is commonly used in different resonance problems (Garfinkel, 1973; Hori, 1960). The method consists of developing the Hamiltonian and the determining function $S$ in powers of the square root of the small parameter:

$$
S=S_{1 / 2}+S_{1}+S_{3 / 2}+S_{2}+\ldots
$$

If $F^{\prime}$ is the new Hamiltonian, then

$$
F^{\prime}=F_{0}^{\prime}+F_{1}^{\prime}+F_{3 / 2}^{\prime}+F_{2}^{\prime}+\ldots
$$

and equations (18) now become

$$
\begin{align*}
\mathrm{F}^{\prime}= & \mathrm{F}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{1 / 2}\right\}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{1}\right\}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{3 / 2}\right\}+\left\{\mathrm{F}_{1}, \mathrm{~S}_{1}\right\}+\left\{\mathrm{F}_{1}, \mathrm{~S}_{1 / 2}\right\}+\ldots \\
& +\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1}\right\}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{6}\left\{\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1}\right\}, \mathrm{S}_{1}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{1}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\ldots \tag{22}
\end{align*}
$$

Equation (22) immediately yields to order 0:

$$
\mathrm{F}_{0}^{\prime}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}\right)=\mathrm{F}_{0}\left(\mathrm{~L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}\right)
$$

We shall group the different terms in equation (22) according to their order, but first we need to develop some of the Poisson brackets in the following manner:

First consider

$$
\left\{\mathrm{F}_{0}, \mathrm{~S}_{1 / 2}\right\}=-\frac{\mu^{2}}{\mathrm{~L}^{\prime 3}} \frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}+v \frac{\partial \mathrm{~S}_{1 / 2}}{\partial \mathrm{~h}^{\prime}}
$$

As $S_{1 / 2}$ arises strictly from the existence of a resonance and the resonant terms are functions of the resonant parameter $\Phi_{a, \beta}^{\prime}=a\left(\ell^{\prime}+g^{\prime}\right)+\beta h^{\prime}$, we look for a solution $S_{1 / 2}=S_{1 / 2}\left(L^{\prime}, G^{\prime}, H^{\prime}, g^{\prime}, \Phi_{a, \beta}^{\prime}\right)$ such that

$$
\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}=\frac{\alpha}{\beta} \frac{\partial S_{1 / 2}}{\partial \mathbf{h}^{\prime}}
$$

Thus,

$$
\left\{\mathrm{F}_{0}, \mathrm{~S}_{1 / 2}\right\}=\left(-\frac{\mu^{2}}{L^{\prime 3}}+\frac{\nu \beta}{a}\right) \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}=\gamma \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}
$$

where $\gamma=(\nu \beta / a)-\left(\mu^{2} / L^{\prime}\right)$ is a small parameter that can be assumed to be of order $1 / 2$, so that $\left\{\mathrm{F}_{0}, \mathrm{~S}_{1 / 2}\right\}$ is of order 1 in equation (22).

Let us write

$$
\mathrm{S}_{1}=\mathrm{SJ}_{1}+\mathrm{SR}_{1}
$$

where $\mathrm{SJ} J_{1}$ arises only from the effect of $J_{2}$, which can be obtained by solving

$$
F_{1}^{\prime}=\left\{\mathrm{F}_{0}, \mathrm{SJ} \mathrm{l}_{1}\right\}+\mathrm{F}_{1} \longrightarrow\left\{\begin{array}{l}
\mathrm{F}_{1}^{\prime}=\mathrm{F}_{1} \mathrm{~s}  \tag{23}\\
\mathrm{SJ} J_{1}=\int_{1 \mathrm{p}} \mathrm{dt}^{\prime}
\end{array}\right.
$$

and SR comes from the existence of the resonant terms that combine their influence with that of $\mathrm{J}_{2}$. We shall look for a solution of the form

$$
\mathrm{SR}_{1}=\mathrm{SR}_{1}\left(\mathrm{~L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \mathrm{g}^{\prime}, \Phi_{\alpha, \beta}^{\prime}\right)
$$

such that

$$
\left\{\mathrm{F}_{0}, \mathrm{SR}_{1}\right\}=\gamma \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}
$$

is of order $3 / 2$.

We can then write

$$
\begin{align*}
& \gamma=\gamma\left(\mathrm{L}^{\prime}\right), \\
& \begin{aligned}
\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\} & =\frac{1}{2}\left\{\gamma \frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\} \\
& =\frac{1}{2} \underbrace{\partial \mathrm{~L}^{\prime}}_{\text {of order } 1}\left(\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{2}
\end{aligned}+\underbrace{\frac{1}{2} \gamma\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\} ;}_{\text {of order } 3 / 2} \\
& \begin{aligned}
\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}= & \frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2}\left\{\gamma \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\} \\
= & \frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}} \\
& +\frac{1}{2} \gamma\left\{\left\{\frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\},\right.
\end{aligned} \tag{24}
\end{align*}
$$

where

$$
\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{S}_{1 / 2}\right\} \quad \text { and } \quad \frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}
$$

are of order $3 / 2$ and

$$
\frac{1}{2} \gamma\left\{\frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}
$$

is of order 2;

$$
\begin{align*}
\frac{1}{6}\left\{\left\{\left\{\mathrm{~F}_{0}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}= & \frac{1}{6} \frac{\partial^{2} \gamma}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{3}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{6} \gamma\left\{\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\} \tag{26}
\end{align*}
$$

where

$$
\frac{1}{6} \frac{\partial^{2} \gamma}{\partial \mathrm{~L}^{\prime 2}}\left(\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{3}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}
$$

is of order $3 / 2$ and

$$
\frac{1}{6} Y\left\{\left\{\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}, S_{1 / 2}\right\}, S_{1 / 2}\right\}
$$

is of order 2. In the expressions above,

$$
\frac{\partial \gamma}{\partial \mathrm{L}}=\frac{3 \mu^{2}}{\mathrm{~L}^{\prime}}, \quad \frac{\partial^{2} \gamma}{\partial \mathrm{~L}^{\prime 2}}=-\frac{12 \mu^{2}}{\mathrm{~L}^{5}}
$$

We write $\mathrm{F}_{2}$ in the following manner:

$$
F_{2}=2 \sum_{k_{1}} D\left(k_{1}\right) \sin ^{2} \theta\left(k_{1}\right)-\sum_{k_{1}} D\left(k_{1}\right)+\sum V_{n, m}^{\prime}
$$

where

$$
\theta\left(k_{1}\right)=\frac{\pi}{2}-\theta_{1}\left(k_{1}\right)
$$

and $D\left(k_{1}\right)$ is a function of ( $\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \mathrm{g}^{\prime}$ ); we assume here that $\mathrm{g}^{\prime}$ is reduced to its constant part $\mathrm{g}_{0}^{\prime}$. The secular part of $\mathrm{g}^{\prime}$ should be taken into account when the long-period perturbations are removed.

We shall assume for the time being that our Hamiltonian contains terms for only one particular value of $k_{1}$, and thus we can drop the $k_{1}$ index for convenience.

We now proceed to splitting equation (22) by grouping terms of the same order.
A. We shall allow the first approximation of the determining function, $\mathrm{s}_{1 / 2}$, to contain the resonant part of the Hamiltonian, which means that it will contain secular terms. We then have, to order 1 ,

$$
\left\{\mathrm{F}_{0}, \mathrm{~S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{2}+2 \mathrm{D} \sin ^{2} \theta=0
$$

or

$$
\begin{equation*}
\gamma\left(\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}\right)+\frac{3}{2} \frac{\mu^{2}}{L^{\prime}}\left(\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}\right)^{2}+2 D \sin ^{2} \theta=0 . \tag{27}
\end{equation*}
$$

This is called the "resonance equation."

If

$$
\begin{align*}
& \mathrm{A}=\frac{2 \mathrm{\gamma L}^{\prime}{ }^{4}}{3 \mu^{2} \mathrm{ak}_{1}},  \tag{28}\\
& \mathrm{k}^{2}=\frac{12 \mathrm{D} \mu^{2}}{\mathrm{~L}^{\prime 4} \mathrm{y}^{2}}
\end{align*}
$$

and we assume that $S_{1 / 2}=S_{1 / 2}\left({ }^{(\theta)}\right.$, then by solving the quadratic equation (27), we get

$$
\begin{equation*}
\frac{\partial S_{1 / 2}}{\partial \theta}=-A \pm A \sqrt{1-k^{2} \sin ^{2} \theta} \tag{29a}
\end{equation*}
$$

or

$$
\begin{equation*}
\frac{\partial S_{1 / 2}}{\partial \theta}=\mathrm{A}(\Delta-1) \tag{29b}
\end{equation*}
$$

if

$$
\begin{equation*}
\Delta= \pm \sqrt{1-\mathrm{k}^{2} \sin ^{2} \theta} . \tag{29c}
\end{equation*}
$$

We discuss this in detail in Section 4.
B. To order $3 / 2$, we have the following:

$$
\begin{aligned}
\mathrm{F}_{3 / 2}^{\prime}= & \gamma \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}+\frac{\gamma}{2}\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}+\frac{\partial \gamma}{\partial \mathrm{L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}} \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}} \\
& +\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}} \frac{\partial \mathrm{S}_{1}}{\partial \ell^{\prime}}+\left\{\mathrm{F}_{1}, \mathrm{~S}_{1 / 2}\right\} \\
& +\frac{1}{6} \frac{\partial^{2} \gamma}{\partial \mathrm{~L}^{\prime 2}}\left(\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{3}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\},
\end{aligned}
$$

or

$$
\begin{align*}
\mathrm{F}_{3 / 2}^{\prime}= & \gamma \Delta \frac{\partial \mathrm{SR}_{1}}{\partial \ell^{\prime}}+\frac{\gamma \Delta}{2}\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{6} \frac{\partial^{2} \gamma}{\partial \mathrm{~L}^{\prime 2}}\left(\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{3} \\
& +\left\{\mathrm{F}_{1}, \mathrm{~S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}}, \frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}} \frac{\partial \mathrm{SJ}_{1}}{\partial \ell^{\prime}} \tag{30}
\end{align*}
$$

We notice that in equation (30), some terms contain only $S_{1 / 2}$ and are therefore due only to the resonant tesseral harmonics, while others express the interaction between the resonant terms and $\mathrm{J}_{2}$. Let us therefore write

$$
\mathrm{SR}=\mathrm{Z}_{1}+\mathrm{Z}_{2},
$$

where $Z_{1}$ contains only those terms that are functions of $S_{1 / 2}$, and $Z_{2}$ contains all the others. If we write $F_{3 / 2}^{\prime}=X+Y$, we then obtain two equations:

$$
\begin{equation*}
X=\gamma \Delta \frac{\partial Z_{1}}{\partial \ell^{\prime}}+\frac{\gamma \Delta}{2}\left\{\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}, S_{1 / 2}\right\}+\frac{1}{6} \frac{\partial^{2} \gamma}{\partial L^{\prime}}\left(\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}\right)^{3} \tag{31}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{Y}=\left\{\mathrm{F}_{1}, \mathrm{~S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}{ }_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \mathrm{Y}}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}} \frac{\partial \mathrm{SJ}_{1}}{\partial \ell^{\prime}}+\frac{\gamma \Delta \partial \mathrm{Z}_{2}}{\partial \ell^{\prime}} . \tag{32}
\end{equation*}
$$

C. Finally, to order 2, we have

$$
\begin{equation*}
\mathrm{F}_{2}^{\prime}=\mathrm{F}_{2 \mathrm{~s}}=-\mathrm{D}\left(\mathrm{~L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \mathrm{g}^{\prime}\right) \tag{33}
\end{equation*}
$$

and

$$
\begin{align*}
0 & =\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}^{\prime}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{3 / 2}\right\}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{2}\right\}+\left\{\mathrm{F}_{0}, \mathrm{~S}_{2}^{\prime}\right\} \\
& +\frac{1}{2} \gamma\left\{\frac{\partial \mathrm{SR}^{\prime}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{SJ}_{1}\right\} \\
& +\frac{1}{6} \gamma\left\{\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\left\{\mathrm{F}_{1}, \mathrm{~S}_{1}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{1}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{6}\left\{\left\{\frac{\partial \mathrm{Y}}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1}\right\}+\frac{1}{6}\left\{\left\{\frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1}\right\}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{4!}\left\{\left\{\frac{\partial \mathrm{V}}{\partial \mathrm{~L}^{\prime}},\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)^{2}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{4!}\left\{\frac{\partial \mathrm{y}}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right)\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}} ; \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{4!} \frac{\partial \gamma}{\partial \mathrm{L}^{\prime}\left\{\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}\left(\frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}\right) .} . \tag{34}
\end{align*}
$$

Equation (34) defines $S_{2}, S_{2}^{\prime}$, and $S_{3 / 2}$. The term $S_{2}$ arises from the existence of $\mathrm{J}_{2}$ only:

$$
\begin{equation*}
\left\{\mathrm{F}_{0}, \mathrm{~S}_{2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{SJ}_{1}\right\}+\left\{\mathrm{F}_{1}, \mathrm{SJ}_{1}\right\}=0 ; \tag{35}
\end{equation*}
$$

$\mathrm{S}_{2}^{\prime}$, which takes care of the nonresonant tesseral part, is obtained by solving

$$
\begin{equation*}
\left\{\mathrm{F}_{0}, \mathrm{~S}_{2}^{\prime}\right\}+\sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}^{\prime}=0 \tag{36a}
\end{equation*}
$$

or

$$
\begin{equation*}
\mathrm{s}_{2}^{\prime}=\int \sum \mathrm{V}_{\mathrm{n}, \mathrm{~m}}^{\prime} \mathrm{dt} \tag{36b}
\end{equation*}
$$

and $\mathrm{S}_{3 / 2}$ is due to the resonant tesseral harmonics, which we assume to be a function of $\theta$, such that

$$
\left\{\mathrm{F}_{0}, \mathrm{~S}_{3 / 2}\right\}=\mathrm{y} \frac{\partial S_{3 / 2}}{\partial \ell^{\prime}}
$$

and

$$
\begin{align*}
0= & \gamma \frac{\partial S_{3 / 2}}{\partial \ell^{\prime}}+\frac{1}{2} \gamma\left\{\frac{\partial S_{1}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2} \frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}}\left(\frac{\partial \mathrm{SJ}_{1}}{\partial \ell^{\prime}}\right) \frac{\partial \mathrm{S}_{1}}{\partial \ell^{\prime}} \\
& \left.+\frac{1}{6} \gamma\left\{\frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{0}, \mathrm{SJ}_{1}\right\}, \mathrm{SR}_{1}\right\}+\frac{1}{2}\left\{\left\{\mathrm{~F}_{1}, \mathrm{~S}_{1 / 2}\right\}, \mathrm{S}_{1 / 2}\right\} \\
& +\frac{1}{6}\left\{\left\{\frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}, \mathrm{S}_{1}\right\}+\frac{1}{6}\left\{\left\{\frac{\partial \gamma}{\partial \mathrm{~L}^{\prime}} \frac{\partial \mathrm{S}_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1}\right\}, \mathrm{S}_{1 / 2}\right\}+\ldots \cdot \tag{37}
\end{align*}
$$

If the Hamiltonian contains more than one value of $\mathrm{k}_{1}$ (usually $\mathrm{k}_{1}=1$ is enough), we can write

$$
\begin{align*}
& \mathrm{s}_{1 / 2}=\sum_{\mathrm{k}_{1}} \mathrm{~s}_{1 / 2, \mathrm{k}_{1}} \\
& \mathrm{z}_{1}=\sum_{\mathrm{k}_{1}} \mathrm{z}_{1, \mathrm{k}_{1}}  \tag{38}\\
& \mathrm{z}_{2}=\sum_{\mathrm{k}_{1}} \mathrm{z}_{2, \mathrm{k}_{1}}
\end{align*}
$$

We first obtain $S_{1 / 2}$ by solving equation (27), replacing $S_{1 / 2}$ by $S_{1 / 2, k_{1}}$ for each value of $k_{1}$, and adding the different $S_{1 / 2, k_{1}}$ terms. This amounts to neglecting mixed terms of the form

$$
\frac{\partial \mathrm{S}_{1 / 2, \mathrm{k}_{1}}}{\partial \ell^{\prime}} \times \frac{\partial \mathrm{S}_{1 / 2, \mathrm{k}_{2}}}{\partial \ell^{\prime}}
$$

in this theory.

We then solve equations (31) and (32) in the same way, again for each value of $\mathrm{k}_{1}$, and add the perturbations obtained; the neglected mixed terms are reasonably small.

If ( $\mathrm{L}, \mathrm{G}, \mathrm{H}, \mathrm{\ell}, \mathrm{~g}, \mathrm{~h}$ ) is the set of modified Delaunay variables as defined in Section 2, and ( $\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}$ ) are the new variables, then after applying Hori's transformation, the solution of the equations of motion (14) becomes

$$
\begin{align*}
& \mathrm{L}=\mathrm{L}^{\prime}+\frac{\partial \mathrm{S}}{\partial \ell^{\prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \ell^{\prime}}, \mathrm{S}\right\}+\frac{1}{3!}\left\{\left\{\frac{\partial \mathrm{S}}{\partial \ell^{\prime}}, \mathrm{S}\right\}, \mathrm{S}\right\}+\ldots, \\
& \mathrm{G}=\mathrm{G}^{\prime}+\frac{\partial \mathrm{S}}{\partial \mathrm{~g}^{\prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~g}^{\prime}}, \mathrm{S}\right\}+\ldots, \\
& \mathrm{H}=\mathrm{H}^{\prime}+\frac{\partial \mathrm{S}}{\partial \mathrm{~h}^{\prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~h}^{\prime}}, \mathrm{S}\right\}+\ldots,  \tag{39}\\
& \ell=\ell^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{~L}^{\prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~L}^{\prime}}, \mathrm{S}\right\}-\ldots, \\
& \mathrm{g}=\mathrm{g}^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{G}^{\prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{G}^{\prime}}, \mathrm{S}\right\}-\ldots, \\
& \mathrm{h}=\mathrm{h}^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{H}^{\prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{H}^{\prime}}, \mathrm{S}\right\}-\ldots,
\end{align*}
$$

where $\mathrm{S}=\mathrm{S}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)$ is the generating function

$$
\mathrm{S}=\mathrm{S}_{1 / 2}+\mathrm{SJ}+\mathrm{SR}+\ldots
$$

Then, if

$$
\mathrm{F}^{\prime}=\frac{\mu^{2}}{2 \mathrm{~L}^{\prime 2}}+\nu \mathrm{H}^{\prime}-\sum_{\mathrm{k}_{1}} \mathrm{D}\left(\mathrm{k}_{1}\right)+\mathrm{F}_{1}^{\prime}
$$

is the new Hamiltonian, we get

$$
\begin{aligned}
& \frac{\mathrm{dL}^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{\prime}}{\partial \ell^{\prime}}=0, \\
& \frac{\mathrm{dG}^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{\prime}}{\partial \mathrm{g}^{\prime}}=0, \\
& \frac{\mathrm{dH}^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{\prime}}{\mathrm{gn}^{\prime}}=0, \\
& \frac{\mathrm{~d} \ell^{\prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{\prime}}{\partial \mathrm{L}^{\prime}}=\frac{\mu^{2}}{\mathrm{~L}^{\prime}}{ }^{3}-\frac{\partial \mathrm{F}_{1}^{\prime}}{\partial \mathrm{L}^{\prime}}+\sum_{\mathrm{k}_{1}} \frac{\partial \mathrm{D}\left(\mathrm{k}_{1}\right)}{\partial \mathrm{L}^{\prime}}=\ell_{1}^{\prime}, \\
& \frac{\mathrm{dg}^{\prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{\prime}}{\partial \mathrm{G}^{\prime}}=-\frac{\partial \mathrm{F}_{1}^{\prime}}{\partial \mathrm{G}^{\prime}}+\sum_{\mathrm{k}_{1}} \frac{\partial \mathrm{D}\left(\mathrm{k}_{1}\right)}{\partial \mathrm{G}^{\prime}}=\mathrm{g}_{1}^{\prime}, \\
& \frac{d h^{\prime}}{d t}=-\frac{\partial \mathrm{F}^{\prime}}{\partial \mathrm{H}^{\prime}}=-\frac{\partial \mathrm{F}_{1}^{\prime}}{\partial \mathrm{H}^{\prime}}-v+\sum_{\mathrm{k}_{1}} \frac{\partial \mathrm{D}\left(\mathrm{k}_{1}\right)}{\partial \mathrm{H}^{\prime}}=\mathrm{h}_{1}^{\prime},
\end{aligned}
$$

where $\ell_{1}^{\prime}, g_{1}^{\prime}$, and $h_{1}^{\prime}$ are constants such that

$$
\begin{aligned}
& \ell^{\prime}=\ell_{1}^{\prime} \mathrm{t}+\ell_{0}^{\prime}, \\
& \mathrm{g}^{\prime}=\mathrm{g}_{1}^{\prime} \mathrm{t}+\mathrm{g}_{0}^{\prime}, \\
& \mathrm{h}^{\prime}=\mathrm{h}_{1}^{\prime} \mathrm{t}+\mathrm{h}_{0}^{\prime},
\end{aligned}
$$

with $\ell_{0}^{\prime}, g_{0}^{\prime}$, and $\mathrm{h}_{0}^{\prime}$ also constants.

In Section 4, we will give precise expressions of the perturbations up to order $3 / 2$. The perturbations due to $\mathrm{J}_{2}$ only are not considered in this paper, but can be found in a classical theory of artificial satellites, for instance Brouwer's. However, the interactions between $J_{2}$ and the resonant tesseral harmonics are expressed with the generating function $\mathrm{Z}_{2}$.

## 4. DETAILED EXPRESSIONS OF THE PERTURBATIONS

In this section, we drop the primes in the variables ( $L^{\prime}, G^{\prime}, H^{\prime}, \ell^{\prime}, g^{\prime}, h^{\prime}$ ), since no confusion exists here with the original Delaunay variables, which do not appear.

### 4.1 Expression of $S_{1 / 2}$

As we already have seen,

$$
\frac{\partial \mathrm{S}_{1 / 2}}{\partial \theta}=\mathrm{A}(\Delta-1),
$$

where

$$
\Delta=\epsilon \sqrt{1-\mathrm{k}^{2} \sin ^{2} \theta}, \quad \epsilon= \pm 1
$$

so that

$$
\begin{equation*}
S_{1 / 2}=-A+\epsilon A \int_{0}^{\theta} \sqrt{1-k^{2} \sin ^{2} x} d x \tag{40}
\end{equation*}
$$

Let us discuss the value of $\epsilon$ in equation (40). First, we consider the case when $\mathrm{k}<1$. For any value of $\theta, 1-\mathrm{k}^{2} \sin ^{2} \theta>0$, and we are in the circulation case. In order to have continuity with the classical solution, valid away from resonance, it is necessary that $S_{1 / 2} \rightarrow 0$ when $k \rightarrow 0$. Since $A \rightarrow \infty$ when $k \rightarrow 0$, then, necessarily, $\epsilon=+1$. Otherwise, $\mathrm{S}_{1 / 2} \rightarrow-\infty$ when $\mathrm{A} \rightarrow \infty$.

For the case when $\mathrm{k}=1$,

$$
\begin{aligned}
& \Delta=\epsilon \cos \theta \\
& \int_{0}^{\theta} \Delta(x) d x=\epsilon \sin \theta
\end{aligned}
$$

In order to have continuity when $k \rightarrow 1$ and $k<1$, we must have $\epsilon=+1$.

When $\mathrm{k}>1$,

$$
1-k^{2} \sin ^{2} x \geq 0 \Longleftrightarrow-\arcsin \frac{1}{k} \leq x \leq \arcsin \frac{1}{k}
$$

and we are in the libration case.

We need to know the expressions of the elliptic integrals of the first and second kind $-F(\theta, k)$ and $E(\theta, k)$, respectively - for the cases $k>1$ and $\pi / 2<\theta$. For these, we have the following formulas (Gradshteyn and Ryzhik, 1966):

$$
\begin{align*}
& \mathrm{E}(\mathrm{~m} \pi \pm \theta, \mathrm{k})=2 \mathrm{mE} \pm \mathrm{E}(\theta, \mathrm{k}), \\
& \mathrm{F}(\mathrm{~m} \pi \pm \theta, \mathrm{k})=2 \mathrm{mk} \pm \mathrm{F}(\theta, \mathrm{k}) \tag{41}
\end{align*}
$$

where $E$ and $K$ are the complete elliptic integrals of

$$
\begin{aligned}
\mathrm{E} & =\mathbf{E}\left(\frac{\pi}{2}, \mathrm{k}\right), \\
\mathrm{F} & =\mathrm{F}\left(\frac{\pi}{2}, \mathrm{k}\right) .
\end{aligned}
$$

Then, if $k_{1}=1 / k,{k^{\prime}}^{2}=1-k^{2}$, and $\theta_{1}$ is defined by $\sin \theta_{1}=k \sin \theta$, we have

$$
\begin{align*}
& E(\theta, k)=k_{1}\left[k^{2} E\left(\theta^{\prime}, k_{1}\right)+{k^{\prime}}^{2} F\left(\theta^{\prime}, k_{1}\right)\right]  \tag{42}\\
& F(\theta, k)=k_{1} F\left(\theta^{\prime}, k_{1}\right)
\end{align*}
$$

Returning to our problem, we see that in order to have continuity when $\mathrm{k} \rightarrow 1$ and $k>1$, we can take $\epsilon=+1$ in this case also, because, from equations (42),

$$
\lim _{k \rightarrow 1} E(\theta, k)=\lim _{k \rightarrow 1} E\left(\theta^{\prime}, k\right)=E(\theta, 1)=\sin \theta
$$

In conclusion, whatever the value of $k$, we can write $S_{1 / 2}$ in the form

$$
\mathrm{S}_{1 / 2}=-\mathrm{A} \theta+\mathrm{AE}(\theta, \mathrm{k})
$$

To simplify this, we can set

$$
\begin{aligned}
& \mathrm{I}_{2}=\mathrm{E}(\theta, \mathrm{k}), \\
& \mathrm{I}_{0}=\mathrm{F}(\theta, \mathrm{k})
\end{aligned}
$$

Then the derivatives of $\mathrm{S}_{1 / 2}$ are readily obtained:

$$
\begin{align*}
& \frac{\partial S_{1 / 2}}{\partial l}=a c_{2} \gamma(\Delta-1), \\
& \frac{\partial S_{1 / 2}}{\partial g}=d c_{2} \gamma(\Delta-1)+c_{2} \gamma a_{4}\left(I_{2}-I_{0}\right), \\
& \frac{\partial S_{1 / 2}}{\partial \mathrm{~h}}=\beta c_{2} \gamma(\Delta-1), \\
& \frac{\partial S_{1 / 2}}{\partial \mathrm{~L}}=\frac{2}{a}\left(I_{0}-\theta\right)+8 c_{1} \gamma\left(I_{2}-\theta\right)+c_{2} p_{1} \gamma(\Delta-1)+\left(c_{2} \gamma a_{1}-4 c_{1} \gamma\right)\left(I_{2}-I_{0}\right),  \tag{43}\\
& \frac{\partial S_{1 / 2}}{\partial G}=c_{2} \gamma p_{2}(\Delta-1)+c_{2} \gamma a_{2}\left(I_{2}-I_{0}\right), \\
& \frac{\partial S_{1 / 2}}{\partial H}=\gamma c_{2} p_{3}(\Delta-1)+c_{2} \gamma a_{3}\left(I_{2}-I_{0}\right),
\end{align*}
$$

where

$$
c_{2}=\frac{L^{4}}{3 \mu^{2}} \quad, \quad c_{1}=\frac{c_{2}}{L}
$$

and

$$
\begin{array}{ll}
a_{1}=\frac{1}{\bar{D}} \frac{\partial \mathrm{D}}{\partial \mathrm{~L}}, & \mathrm{p}_{1}=\frac{\partial \psi}{\partial \mathrm{L}}, \\
\mathrm{a}_{2}=\frac{1}{\mathrm{D}} \frac{\partial \mathrm{D}}{\partial \mathrm{G}}, & \mathrm{p}_{2}=\frac{\partial \psi}{\partial \mathrm{G}}, \tag{44}
\end{array}
$$

$$
\begin{array}{ll}
a_{3}=\frac{1}{D} \frac{\partial D}{\partial H}, & p_{3}=\frac{\partial \psi}{\partial H}, \\
a_{4}=\frac{1}{D} \frac{\partial D}{\partial g}, & p_{4}=\frac{\partial \psi}{\partial g},  \tag{44}\\
d=\beta+p_{4}, &
\end{array}
$$

To calculate the elliptic integrals, we use the following expansions (from Gradshteyn and Ryzhik):

$$
\begin{align*}
& 0<\theta<\frac{\pi}{2} \quad, \quad \mathrm{k}^{2}<1 \\
& F(\theta, k)=\sum_{m=0}^{\infty}\binom{-1 / 2}{m}\left(-\mathrm{k}^{2}\right)^{m} \mathrm{t}_{2 \mathrm{~m}^{(\theta)}}  \tag{45}\\
& E(\theta, k)=\sum_{m=0}^{\infty}\binom{1 / 2}{m}\left(-\mathrm{k}^{2}\right)^{m} t_{2 m}(\theta)
\end{align*}
$$

where

$$
\begin{align*}
& \mathrm{t}_{0}(\theta)=\theta \\
& \mathrm{t}_{2}(\theta)=\frac{1}{2}(\theta-\sin \theta \cos \theta), \\
& \mathrm{t}_{2 \mathrm{~m}}(\theta)=\frac{2 \mathrm{~m}-1}{2 \mathrm{~m}} \mathrm{t}_{2(\mathrm{~m}-1)^{(\theta)}-\frac{1}{2 \mathrm{~m}} \cos \theta \sin ^{2 \mathrm{~m}-1} \theta},  \tag{46}\\
& \binom{a}{\mathrm{n}}=\frac{(-1)^{n}(a)_{n}}{\mathrm{n}!}
\end{align*}
$$

in which

$$
\begin{aligned}
& \text { (a) }{ }_{n}=a(a+1) \cdots(a+n-1), \quad n=1,2, \cdots, \\
& \text { (a) })_{0}=1,
\end{aligned}
$$

and

$$
\begin{align*}
& \mathrm{K}(\mathrm{k})=\frac{\pi}{2}+\sum_{\mathrm{m}=1}^{\infty} \frac{(1 / 2)_{\mathrm{m}}(1 / 2)_{\mathrm{m}} \mathrm{k}^{2 \mathrm{~m}}}{\mathrm{~m}!\mathrm{m}!}, \\
& \mathrm{E}(\mathrm{k})=\frac{\pi}{2}+\sum_{\mathrm{m}=1}^{\infty} \frac{1}{(1-2 \mathrm{~m})}\binom{-1 / 2}{\mathrm{~m}}^{2} \mathrm{k}^{2 \mathrm{~m}} . \tag{47}
\end{align*}
$$

### 4.2 Expression of $\mathrm{Z}_{1}$ and Its Derivatives

We first calculate the Poisson brackets that appear in equation (31). After some algebra,

$$
\begin{aligned}
\left\{\frac{\partial S_{1 / 2}}{\partial \ell^{\prime}}, \mathrm{S}_{1 / 2}\right\}= & a \mathrm{ac}_{2} \gamma\left[(\Delta-1)\left(\frac{1}{\Delta}-1\right)+\mathrm{k}^{2}\left(\mathrm{I}_{0}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}\right. \\
& +4 \mathrm{ac}_{1} \gamma(\Delta-1)^{2}+\mathrm{c}_{2} \mathrm{x}_{\mathrm{\gamma}}(\Delta-1)\left(\Delta-\frac{1}{\Delta}\right) \\
& \left.+\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right) \mathrm{c}_{2}{\gamma x k^{2}}^{2} \frac{\sin \theta \cos \theta}{\Delta}+4 a \mathrm{c}_{1} \gamma^{2}{ }^{2}\left(\mathrm{I}_{2}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}\right],
\end{aligned}
$$

where $c_{2}, \gamma, k$, and $\Delta$ are as defined earlier, and

$$
\begin{align*}
& c_{1}=\frac{L^{\prime 3}}{3 \mu^{2} a} \\
& x=\frac{1}{2}\left(d a_{2}+a a_{1}+\beta a_{3}-\frac{4}{L^{\prime}} a\right),  \tag{49}\\
& d=\beta+p_{4}
\end{align*}
$$

The equation that gives X and $\mathrm{Z}_{1}$ is then

$$
\begin{align*}
X= & \gamma \Delta \frac{\partial Z_{1}}{\partial \ell^{\prime}}+\frac{a c_{2} \gamma^{2} \Delta}{2}\left[(\Delta-1)\left(\frac{1}{\Delta}-1\right)+k^{2}\left(I_{0}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}\right. \\
& +4 a c_{1} \gamma(\Delta-1)^{2}+c_{2} x \gamma(\Delta-1)\left(\Delta-\frac{1}{\Delta}\right)+c_{2} x \gamma k^{2}\left(I_{2}-I_{0}\right) \frac{\sin \theta \cos \theta}{\Delta} \\
& \left.+\frac{1}{3} \frac{\partial^{2} \gamma}{\partial L^{2}} a^{3} c_{2}^{2} \gamma\left(\frac{\Delta-1}{\Delta}\right)^{3}+4 a c_{1} \gamma k^{2}\left(I_{2}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}\right] . \tag{50}
\end{align*}
$$

Since ( $\Delta-1$ ) is periodic in $\theta$, we do not find any constant terms:

$$
\mathrm{x}=0 .
$$

We assumed earlier that $\mathrm{SR}_{1}$ was a function of $\theta$, and we will now assume that $Z$ is a function of $\theta$ only. Therefore,

$$
\frac{\partial Z_{1}}{\partial \ell^{\prime}}=\frac{\partial Z_{1}}{\partial \theta} \frac{\partial \theta}{\partial \ell^{\prime}}=\frac{\mathbf{a}}{2} \frac{\partial Z_{1}}{\partial \theta},
$$

and the following equation gives $\mathrm{Z}_{1}$ :

$$
\begin{align*}
\frac{\partial \mathrm{Z}_{1}}{\partial \theta}= & -\mathrm{c}_{2} \gamma\left[(\Delta-1)\left(\frac{1}{\Delta}-1\right)+\mathrm{k}^{2}\left(\mathrm{I}_{0}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}+4 \mathrm{a} c_{1} \gamma(\Delta-1)^{2}\right. \\
& +\mathrm{c}_{2} \times \gamma(\Delta-1)\left(\Delta-\frac{1}{\Delta}\right)+\mathrm{c}_{2} \gamma \times \mathrm{k}^{2}\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right) \frac{\sin \theta \cos \theta}{\Delta} \\
& \left.+4 \mathrm{ac}_{1} \gamma \mathrm{k}^{2}\left(\mathrm{I}_{2}-\theta\right) \frac{\sin \theta \cos \theta}{\Delta}-\frac{4}{3} a c_{1} \gamma\left(\Delta^{2}-3 \Delta+3-\frac{1}{\Delta}\right)\right] . \tag{51}
\end{align*}
$$

In order to integrate this equation and obtain $Z_{1}$, the following expressions are useful:

$$
\begin{align*}
& \int \frac{x \sin x \cos x d x}{\sqrt{1-k^{2} \sin ^{2} x}}=-\frac{x \Delta(x)}{k^{2}}+\frac{I_{2}(x)}{k^{2}} \\
& \int \frac{I_{2}(x) \sin x \cos x d x}{\sqrt{1-k^{2} \sin ^{2} x}}=-\frac{\Delta(x) I_{2}}{k^{2}}+\frac{1}{k^{2}}\left(x-\frac{k^{2}}{2} x+\frac{k^{2}}{4} \sin 2 x\right)  \tag{52}\\
& \int \frac{I_{0}(x) \sin x \cos x d x}{\sqrt{1-k^{2} \sin ^{2} x}}=-\frac{I_{0} \Delta+x}{k^{2}}
\end{align*}
$$

Thus, by setting

$$
q_{1}=a c_{1} \gamma \quad \text { and } \quad q_{2}=c_{2} \gamma x
$$

we obtain

$$
\begin{align*}
Z_{1}= & c_{2} \curlyvee\left\{I_{2}\left(2+q_{2}+8 q_{1}\right)-\theta\left[3+\frac{20}{3} q_{1}-\frac{k^{2}}{2}\left(\frac{20}{3} q_{1}+2 q_{2}\right)\right]\right. \\
& -\frac{\mathrm{k}^{2}}{4} \sin 2 \theta\left(\frac{20}{3} q_{1}+2 q_{2}\right)+\Delta \mathrm{I}_{2}\left(q_{2}+4 q_{1}\right)-\Delta \theta\left(1+4 q_{1}\right) \\
& \left.+I_{0}\left(1-q_{2}-\frac{4}{3} q_{1}\right)+\Delta I_{0}\left(1-q_{2}\right)\right\} . \tag{53}
\end{align*}
$$

Then the derivatives of $Z_{1}$ with respect to the variables $L, G, H, \ell, g$, and $h$ are

$$
\begin{align*}
\frac{\partial Z_{1}}{\partial X}=c_{2} \gamma & \left\{\frac{\partial I_{2}}{\partial X}\left[2+q_{2}+8 q_{1}+\Delta\left(q_{2}+4 q_{1}\right)\right]\right. \\
& -\frac{\partial \theta}{\partial X}\left[3+\frac{20}{3} q_{1}-\frac{k^{2}}{2}\left(\frac{20}{3} q_{1}+2 q_{2}\right)+\Delta\left(1+4 q_{1}\right)\right] \\
& +\frac{\partial I_{0}}{\partial X}\left[1-q_{2}-\frac{4}{3} q_{1}+\Delta\left(1-q_{2}\right)\right] \\
& +\frac{\partial \Delta}{\partial X}\left[I_{2}\left(q_{2}+4 q_{1}\right)-\theta\left(1+4 q_{1}\right)+I_{0}\left(1-q_{2}\right)\right] \\
& -\left(\frac{20}{3} q_{1}+2 q_{2}\right)\left(\frac{k^{2}}{2} \frac{\partial \theta}{\partial X} \cos 2 \theta+\frac{1}{4} \frac{\partial k^{2}}{\partial X} \sin 2 \theta\right) \\
& \left.+\frac{\partial q_{2}}{\partial X}\left(I_{2}+k^{2} \theta-\frac{k^{2}}{2} \sin 2 \theta+\Delta I_{2}-I_{0}-\Delta I_{0}\right)\right\} \tag{54}
\end{align*}
$$

[eq. (54) cont. on next page]

$$
\begin{align*}
+\epsilon\left(-c_{2} \gamma \frac{\partial q_{1}}{\partial L^{\prime}}\left[8 I_{2}\right.\right. & \left.-\frac{20}{3}\left(1-\frac{k^{2}}{2}\right) \theta-\frac{20}{3} \frac{\mathrm{k}^{2}}{4} \sin 2 \theta+4 \Delta \mathrm{I}_{2}-4 \Delta \theta-\frac{4}{3} \mathrm{I}_{0}\right] \\
+\left(4 \mathrm{c}_{1} \gamma+\frac{1}{\mathrm{a}}\right) & \left\{\mathrm{I}_{2}\left(2+\mathrm{q}_{2}+8 \mathrm{q}_{1}\right)-\theta\left[3+\frac{20}{3} \mathrm{q}_{1}-\frac{\mathrm{k}^{2}}{2}\left(\frac{20}{3} \mathrm{q}_{1}+2 \mathrm{q}_{2}\right)\right]\right. \\
& -\frac{\mathrm{k}^{2}}{4} \sin 2 \theta\left(\frac{20}{3} \mathrm{q}_{1}+2 \mathrm{q}_{2}\right)+\Delta \mathrm{I}_{2}\left(\mathrm{q}_{2}+4 \mathrm{q}_{1}\right) \\
& \left.\left.-\Delta \theta\left(1+4 \mathrm{q}_{1}\right)+\mathrm{I}_{0}\left(1-\mathrm{q}_{2}-\frac{4}{3} \mathrm{q}_{1}\right)+\Delta \mathrm{r}_{0}\left(1-\mathrm{q}_{2}\right)\right\}\right) \tag{54}
\end{align*}
$$

for $\mathrm{X}=\mathrm{L}, \mathrm{G}, \mathrm{H}, \mathrm{l}, \mathrm{g}$, and h , where

$$
\begin{align*}
& \epsilon=1 \quad \text { if } \quad X=L^{\prime} ; \\
& \epsilon=0 \quad \text { if } \quad X \neq L^{\prime} ; \\
& \frac{\partial \Delta}{\partial X}=\frac{a}{2}\left(\Delta-\frac{1}{\Delta}\right)-\frac{k^{2} \sin \theta \cos \theta}{\Delta} \frac{\partial \theta}{\partial X} ;  \tag{55a}\\
& \frac{\partial I_{2}}{\partial X}=\frac{\partial \theta}{\partial X} \Delta+\frac{a_{X}}{2}\left(I_{2}-I_{0}\right) ;  \tag{55b}\\
& \frac{\partial I_{0}}{\partial X}=\frac{1}{\Delta} \frac{\partial \theta}{\partial X}+\frac{a}{2}\left(I_{-2}-I_{0}\right),
\end{align*}
$$

with

$$
\begin{align*}
& I_{-2}(\theta)=\int_{0}^{\theta} \frac{d x}{\Delta^{3}}=\frac{1}{1-k^{2}}\left(I_{2}-k^{2} \frac{\sin \theta \cos \theta}{\Delta}\right) ;  \tag{55c}\\
& \frac{\partial q_{2}}{\partial \mathrm{X}}=c_{2} \gamma \frac{\partial \mathrm{x}}{\partial \mathrm{X}}+\epsilon\left(4 \mathrm{c}_{1} \gamma+\frac{1}{\mathrm{a}}\right) \mathrm{x} ;  \tag{55d}\\
& \frac{\partial \theta}{\partial \mathrm{X}}= \begin{cases}\frac{\partial \theta}{\partial \mathrm{L}^{\prime}}=\frac{p_{1}}{2}, & \frac{\partial \theta}{\partial \ell^{\prime}}=\frac{\mathrm{a}}{2}, \\
\frac{\partial \theta}{\partial G^{\prime}}=\frac{\mathrm{p}_{2}}{2}, & \frac{\partial \theta}{\partial \mathrm{~g}^{\prime}}=\frac{\mathrm{p}_{4}}{2} \\
\frac{\partial \theta}{\partial \mathrm{H}^{\prime}}=\frac{p_{3}}{2}, & \frac{\partial \Theta}{\partial \mathrm{~h}^{\prime}}=\frac{\beta}{2} ;\end{cases} \tag{55e}
\end{align*}
$$

$$
a_{x}=\left\{\begin{array}{lll}
0 & \text { if } & X=\ell^{\prime} \text { or } h^{\prime}  \tag{55f}\\
a_{1}-\frac{4}{L^{\prime}}-\frac{2}{a} & \text { if } & X=L^{\prime} \\
a_{2} & \text { if } & X=G^{\prime} \\
a_{3} & \text { if } & X=H^{\prime} \\
a_{4} & \text { if } & X=g^{\prime} \quad ;
\end{array}\right.
$$

and

$$
\begin{equation*}
\frac{\partial x}{\partial X}=\frac{1}{2}\left(a \frac{\partial a_{1}}{\partial X}+\beta \frac{\partial a_{3}}{\partial X}+d \frac{\partial a_{2}}{\partial X}+\beta \frac{\partial p_{4}}{\partial X}+4 \epsilon \frac{a}{L^{2}}\right) \tag{55~g}
\end{equation*}
$$

### 4.3 Expression of $Z_{2}$ and Its Derivatives

From equation (32) and from the fact that $\left\{\mathrm{F}_{0}, \mathrm{SJ}_{1}\right\}=-\mathrm{F}_{1 \mathrm{p}}$, we can derive $\mathrm{Z}_{2}$ as follows:

$$
\begin{equation*}
\mathrm{Y}=\left\{\mathrm{F}_{1}, \mathrm{~S}_{1 / 2}\right\}-\frac{1}{2}\left\{\mathrm{~F}_{1 p}, \mathrm{~S}_{1 / 2}\right\}+\gamma \Delta \frac{\partial Z_{2}}{\partial \ell}+\frac{1}{2} \frac{\partial \mathrm{y}}{\partial \mathrm{~L}} \frac{\partial \mathrm{~S}_{1 / 2}}{\partial \ell} \frac{\partial S_{J}}{\partial \ell} . \tag{56}
\end{equation*}
$$

Here we can neglect $F_{1 p}$, the periodic part of $F_{1}$, as it is of order 1 in eccentricity and therefore very small for the satellites we are dealing with. We then have

$$
\begin{equation*}
\mathrm{Y}=\left\{\mathrm{F}_{1 \mathrm{~s}}, \mathrm{~S}_{1 / 2}\right\}+\gamma \Delta \frac{\partial \mathrm{Z}_{2}}{\partial \ell} \tag{57}
\end{equation*}
$$

Since there are no constant terms,

$$
\mathrm{Y}=0
$$

and therefore,

$$
\begin{equation*}
\gamma \Delta \frac{\partial Z_{2}}{\partial \ell}=-\left(b_{1} \frac{\partial S_{1 / 2}}{\partial \ell}+b_{2} \frac{\partial S_{1 / 2}}{\partial g}+b_{3} \frac{\partial S_{1 / 2}}{\partial \mathrm{~h}}\right), \tag{58}
\end{equation*}
$$

where

$$
\begin{aligned}
& \mathrm{b}_{1}=\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{~L}}=-\frac{3 \mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}}{\mathrm{G}^{3} \mathrm{~L}^{4}}\left(-\frac{1}{4}+\frac{3}{4} \frac{\mathrm{H}^{2}}{\mathrm{G}^{2}}\right), \\
& \mathrm{b}_{2}=\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{G}}=-\frac{3 \mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}}{\mathrm{~L}^{3} \mathrm{G}^{4}}\left(-\frac{1}{4}+\frac{5}{4} \frac{\mathrm{H}^{2}}{\mathrm{G}^{2}}\right), \\
& \mathrm{b}_{3}=\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{H}}=+\frac{3}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}}{\mathrm{~L}^{3} \mathrm{G}^{5}}
\end{aligned}
$$

and, since $Z_{2}$ is assumed to be a function of $\theta$,

$$
\begin{equation*}
\frac{\partial Z_{2}}{\partial \theta}=-\frac{2 c_{2}}{a}\left[\left(b_{1} a+b_{2} \beta+b_{3} d\right)\left(\frac{\Delta-1}{\Delta}\right)+b_{2} a_{4}\left(\frac{I_{2}-I_{0}}{\Delta}\right)\right] \tag{59}
\end{equation*}
$$

If we write

$$
\begin{align*}
& B_{1}=-\frac{2 c_{2}}{a}\left(b_{1} a+b_{2} d+b_{3} \beta\right), \\
& B_{2}=-\frac{2 c_{2}}{a} b_{2} a_{4} \tag{60}
\end{align*}
$$

we can obtain $\mathrm{Z}_{2}$ from equation (59):

$$
\begin{equation*}
\mathrm{Z}_{2}=\mathrm{B}_{1}\left(\theta-\mathrm{I}_{0}\right)+\mathrm{B}_{2} \int_{0}^{\theta} \frac{\mathrm{I}_{2}(\mathrm{x})-\mathrm{I}_{0}(\mathrm{x})}{\Delta(\mathrm{x})} \mathrm{dx} \tag{61}
\end{equation*}
$$

Now we can write the integrals (Gradshteyn and Ryzhik, 1966)

$$
\begin{align*}
& \int_{0}^{\theta} \frac{I_{0}(x)}{\Delta(x)} d x=\frac{I_{0}^{2}}{2} \\
& \int_{0}^{\theta} \frac{I_{2}(x)}{\Delta(x)} d x=\frac{I_{0}^{2} E}{2 K}+\log \frac{\Theta\left(I_{0}\right)}{\Theta(0)}, \tag{62}
\end{align*}
$$

where $E$ and $K$ are complete elliptic integrals of the first and second kind, respectively, and $\Theta$ is the following theta function:

$$
\Theta(u)=1+2 \sum_{m=1}^{\infty}(-1)^{m} q^{m^{2}} \cos (2 m v)
$$

in which

$$
\begin{equation*}
\mathrm{q}=\exp \left(\frac{-\pi \mathrm{k}^{\prime}}{\mathrm{k}}\right), \quad \mathrm{k}^{\prime}=\sqrt{\mathrm{l-k}^{2}} \tag{63a}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{v}=\frac{\pi \mathrm{u}}{2 \mathrm{~K}} \tag{63b}
\end{equation*}
$$

The derivatives of $Z_{2}$, then, are

$$
\begin{align*}
\frac{\partial \mathrm{Z}_{2}}{\partial \mathrm{X}}= & \mathrm{B}_{1}\left(\frac{\partial \theta}{\partial \mathrm{X}}-\frac{\partial \mathrm{I}_{0}}{\partial \mathrm{X}}\right)+\frac{\partial \mathrm{B}_{1}}{\partial \mathrm{X}}\left(\theta-\mathrm{I}_{0}\right) \\
& +\frac{\partial \mathrm{B}_{2}}{\partial \mathrm{X}} \int_{0}^{\theta} \frac{\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right)(\mathrm{x})}{\Delta(\mathrm{x})} \mathrm{dx}+\mathrm{B}_{2} \frac{\partial \theta}{\partial \mathrm{X}}\left(\frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta}\right)+\mathrm{B}_{2} \operatorname{\sigma r}_{(\mathrm{X})} \tag{64}
\end{align*}
$$

where $X=L^{\prime}, G^{\prime}, H^{\prime}, \ell^{\prime}, g^{\prime}, h^{\prime}$ and, if ${k^{\prime}}^{2}=1-k^{2}$,

$$
\begin{align*}
& \mathscr{f}(\mathrm{X})=\frac{\mathrm{a}}{2 \mathrm{x}^{2}}\left[\frac{\mathrm{I}_{0}^{2} \mathrm{k}^{\prime 2}}{2}+\frac{\mathrm{I}_{2}^{2}}{2}-\mathrm{I}_{0} \mathrm{I}_{2}-\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right) \frac{\mathrm{k}^{2} \sin \theta \cos \theta}{\Delta}+\frac{\mathrm{k}^{2} \sin ^{2} \theta}{2}\right] ; \\
& \frac{\partial B_{1}}{\partial \mathrm{X}}=-\frac{2 \mathrm{c}_{2}}{\mathrm{a}}\left(\frac{\partial \mathrm{~b}_{1}}{\partial \mathrm{X}} \mathfrak{a}+\frac{\partial \mathrm{b}_{2}}{\partial \mathrm{X}} d+\frac{\partial \mathrm{p}_{4}}{\partial \mathrm{X}} \mathrm{~b}_{2}+\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{X}} \beta\right)-\frac{\epsilon}{\mathrm{L}^{\prime}} \mathrm{B}_{1} \quad ;  \tag{65a}\\
& \frac{\partial \mathrm{B}_{2}}{\partial \mathrm{X}}=-\frac{2 \mathrm{c}_{2}}{\mathrm{a}}\left(\frac{\partial \mathrm{~b}_{2}}{\partial \mathrm{X}} \mathrm{a}_{4}+\mathrm{b}_{2} \cdot \frac{\partial \mathrm{a}_{4}}{\partial \mathrm{X}}\right)-\frac{\epsilon}{\mathrm{L}^{\prime}} \mathrm{B}_{2},
\end{align*}
$$

where

$$
\begin{array}{lll}
\epsilon=0 & \text { if } & X \neq L^{\prime}, \\
\epsilon=1 & \text { if } & X=L^{\prime},
\end{array}
$$

$\frac{\partial \mathrm{p}_{4}}{\partial \mathrm{X}}=0 \quad$ if $\quad \mathrm{X}=\ell^{\prime}$ or $\mathrm{h}^{\prime} ;$
$\frac{\partial p_{4}}{\partial \mathrm{~L}^{\prime}}, \frac{\partial \mathrm{p}_{4}}{\partial \mathrm{G}^{\prime}}, \frac{\partial \mathrm{p}_{4}}{\partial \mathrm{H}^{\prime}}$, and $\frac{\partial \mathrm{p}_{4}}{\partial \mathrm{~g}^{\prime}}$ are given in Appendix A ;
$\frac{\partial b_{1}}{\partial \mathrm{X}}=\frac{\partial \mathrm{b}_{2}}{\partial \mathrm{X}}=\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{X}}=0 \quad$ if $\quad \mathrm{X}=\ell^{\prime}$ or $\mathrm{h}^{\prime}$ or $\mathrm{g}^{\prime} ;$
$\frac{\partial \mathrm{b}_{1}}{\partial \mathrm{~L}^{\prime}}=-\frac{4 \mathrm{~b}_{1}}{\mathrm{~L}^{\prime}} ;$
$\frac{\partial \mathrm{b}_{2}}{\partial \mathrm{~L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}} \mathrm{b}_{2} \quad ;$
$\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{~L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}}, \mathrm{b}_{3} \quad ;$
$\frac{\partial b_{1}}{\partial G^{\prime}}=-\frac{3 b_{1}}{G^{\prime}}+\frac{9}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime 2}}{\mathrm{~L}^{\prime 4} \mathrm{G}^{\prime 6}} ;$
$\frac{\partial \mathrm{b}_{2}}{\partial \mathrm{G}^{\prime}}=-\frac{4 \mathrm{~b}_{1}}{\mathrm{G}^{\prime}}+\frac{15}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}{H^{\prime}}^{2}}{{\mathrm{~L}^{\prime}}^{3} \mathrm{G}^{\prime 7}} ;$
$\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{G}^{\prime}}=-\frac{5 \mathrm{~b}_{3}}{\mathrm{G}^{\prime}} ;$
$\frac{\partial \mathrm{b}_{1}}{\partial \mathrm{H}^{\prime}}=-\frac{9}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime}}{\mathrm{L}^{\prime^{4} \mathrm{G}^{\prime 5}}} ;$
$\frac{\partial b_{2}}{\partial \mathrm{H}^{\prime}}=-\frac{15}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime}}{{L^{\prime}}^{3} \mathrm{G}^{\prime 6}} ;$
$\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{H}^{\prime}}=\frac{3}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}}{\mathrm{~L}^{\prime 3} \mathrm{G}^{\prime 5}}$.

We also need the second derivatives of $\mathrm{Z}_{2}$ that appear in expressions such as $\left\{\partial Z_{2} / \partial \mathrm{X}, \mathrm{S}_{1 / 2}\right\}$ for $\mathrm{X}=\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}$, as they are used in the final expression of the perturbations. Their equations follow:

$$
\begin{align*}
\frac{\partial^{2} \mathrm{Z}_{2}}{\partial \mathrm{X}} \frac{\partial \mathrm{Y}}{}= & \frac{\partial \mathrm{B}_{1}}{\partial \mathrm{Y}}\left(\frac{\partial \mathrm{~S}}{\partial \mathrm{X}}-\frac{\partial \mathrm{I}_{0}}{\partial \mathrm{X}}\right)+\frac{\partial^{2} \mathrm{~B}_{1}}{\partial \mathrm{X} \partial \mathrm{Y}}\left(\theta-\mathrm{I}_{0}\right)+\mathrm{B}_{1}\left(\frac{\partial^{2} \theta}{\partial \mathrm{X} \partial \mathrm{Y}}-\frac{\partial^{2} \mathrm{I}_{0}}{\partial \mathrm{X} \partial \mathrm{Y}}\right) \\
& +\frac{\partial^{2} \mathrm{~B}_{2}}{\partial \mathrm{X} \partial \mathrm{Y}} \int_{0}^{\theta} \frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta} \mathrm{dx}+\frac{\partial \mathrm{B}_{2}}{\partial \mathrm{X}} \mathscr{H}_{(\mathrm{Y})}+\frac{\partial \mathrm{B}_{2}}{\partial \mathrm{Y}} \frac{\partial \theta}{\partial \mathrm{X}}\left(\frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta}\right) \\
& +\mathrm{B}_{2} \frac{\partial^{2} \theta}{\partial \mathrm{X} \partial \mathrm{Y}}\left(\frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta}\right)+\mathrm{B}_{2} \frac{\partial \theta}{\partial \mathrm{X}}\left[\frac{\left(\partial \mathrm{I}_{2} / \partial \mathrm{Y}\right)-\left(\partial \mathrm{I}_{0} / \partial \mathrm{Y}\right)}{\Delta}-\frac{1}{\Delta^{2}} \frac{\partial \Delta}{\partial \mathrm{Y}}\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right)\right] \\
& +\frac{\partial \mathrm{B}_{2}}{\partial \mathrm{Y}} \mathcal{A}_{(\mathrm{X})}+\mathrm{B}_{2} \int_{0}^{\theta} \frac{\partial^{2}}{\partial \mathrm{Y} \partial \mathrm{X}}\left(\frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta}\right) \mathrm{dx} \tag{66}
\end{align*}
$$

where

$$
\begin{aligned}
& \frac{\partial^{2} \theta}{\partial \mathrm{X} \partial \bar{Y}}=\frac{1}{2} \frac{\partial^{2} \psi}{\partial \mathrm{X} \partial \mathrm{Y}} \quad \text { if } \quad X \text { and } Y \text { are } L^{\prime}, G^{\prime}, H^{\prime} \text {, or } g^{\prime}, \\
& \text { [see eqs. (A-9) in Appendix A] } \\
& \frac{\partial^{2} \theta}{\partial \mathrm{X} \partial \mathrm{Y}}=0 \quad \text { if } \quad(\mathrm{X} \text { or } \mathrm{Y}) \text { is }\left(\ell^{\prime} \text { or } \mathrm{h}^{\prime}\right) ; \\
& \frac{\partial^{2} B_{1}}{\partial X \partial Y}=-\frac{2 c_{2}}{a}\left(\frac{\partial^{2} b_{1}}{\partial X \partial Y} a+\frac{\partial^{2} b_{2}}{\partial X \partial Y} d+\frac{\partial b_{2}}{\partial X} \frac{\partial p_{4}}{\partial Y}+\frac{\partial p_{4}}{\partial X} \frac{\partial b_{2}}{\partial Y}\right. \\
& \left.+\frac{\partial^{2} \mathrm{p}_{4}}{\partial \mathrm{X} \partial \mathrm{Y}} \mathrm{~b}_{2}+\frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{X} \partial \mathrm{Y}} \beta\right)-\frac{\epsilon}{\mathrm{L}^{\prime}} \frac{\partial \mathrm{B}_{1}}{\partial \mathrm{Y}} \\
& +\eta\left[-\frac{4 c_{1}}{a}\left(\frac{\partial b_{1}}{\partial \mathrm{X}} a+\frac{\partial \mathrm{b}_{2}}{\partial \mathrm{X}} \mathrm{~d}+\frac{\partial \mathrm{p}_{4}}{\partial \mathrm{X}} \mathrm{~b}_{2}+\frac{\partial \mathrm{b}_{3}}{\partial \mathrm{X}} \beta\right)+\frac{\epsilon}{{L^{\prime}}^{2}} \mathrm{~B}_{1}\right] \text {, }
\end{aligned}
$$

in which

$$
\begin{array}{llll}
\epsilon=0 & \text { if } & \mathrm{X} \neq \mathrm{L}^{\prime}, \\
\epsilon=1 & \text { if } & \mathrm{X}=\mathrm{L}^{\prime}, \\
\eta=0 & \text { if } & \mathrm{Y} \neq \mathrm{L}^{\prime}, \\
\eta=1 & \text { if } & \mathrm{Y}=\mathrm{L}^{\prime},
\end{array}
$$

$$
\begin{align*}
& \frac{\partial^{2} b_{1}}{\partial \mathrm{X} \partial \bar{Y}}=\frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{X} \partial \mathrm{Y}}=\frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{X} \partial \mathrm{Y}}=0 \quad \text { if } \quad(\mathrm{X} \text { or } \mathrm{Y})=\left(\ell^{\prime} \text { or } \mathrm{g}^{\prime} \text { or } \mathrm{h}^{\prime}\right) ; \\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{~L}^{\prime 2}}=\frac{20}{\mathrm{~L}^{\prime 2}} \mathrm{~b}_{1} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{~L}^{\prime 2}}=\frac{12}{\mathrm{~L}^{\prime 2}} \mathrm{~b}_{2} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{~L}^{\prime 2}}=\frac{12}{\mathrm{~L}^{\prime 2}} \mathrm{~b}_{3} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{G}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{G}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{4}{\overline{\mathrm{~L}}^{\prime}} \frac{\partial \mathrm{b}_{1}}{\partial \mathrm{G}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{H}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{4}{\mathrm{~L}^{\prime}} \frac{\partial \mathrm{b}_{1}}{\partial \mathrm{H}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{G}^{\prime 2}}=\frac{3}{\mathrm{G}^{\prime 2}} \mathrm{~b}_{1}-\frac{3}{\mathrm{G}^{\prime}} \frac{\partial \mathrm{b}_{1}}{\partial \mathrm{G}^{\prime}}-\frac{27 \mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}{\mathrm{H}^{\prime}}^{2}}{\mathrm{~L}^{\prime 4} \mathrm{G}^{\prime 7}} ;  \tag{67}\\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{G}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{H}^{\prime} \partial \mathrm{G}^{\prime}}=\frac{45}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime}}{\mathrm{L}^{\prime 4} \mathrm{G}^{\prime}}{ }^{5} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{1}}{\partial \mathrm{H}^{2}}=-\frac{9}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}}{\mathrm{~L}^{\prime 4} \mathrm{G}^{5}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{G}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{G}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}} \frac{\partial \mathrm{b}_{2}}{\partial \mathrm{G}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{H}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}} \frac{\partial \mathrm{b}_{2}}{\partial \mathrm{H}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{G}^{\prime 2}}=-\frac{105}{2} \frac{J_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime 2}}{{\mathrm{~L}^{\prime}}^{3} \mathrm{G}^{\prime}{ }^{8}}-\frac{4}{\mathrm{G}^{\prime}} \frac{\partial \mathrm{b}_{2}}{\partial \mathrm{G}^{\prime}}+\frac{4}{\mathrm{G}^{\prime}}{ }^{2} \mathrm{~b}_{2} \quad ;
\end{align*}
$$

[eq. (67) cont. on next page]

$$
\begin{align*}
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{G}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{H}^{\prime} \partial \mathrm{G}^{\prime}}=-\frac{4}{\mathrm{G}^{\prime}} \frac{\partial \mathrm{b}_{2}}{\partial \mathrm{H}^{\prime}}+\frac{15 \mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2} \mathrm{H}^{\prime}}{{\mathrm{L}^{\prime}}^{3} \mathrm{G}^{\prime 7}} ; \\
& \frac{\partial^{2} \mathrm{~b}_{2}}{\partial \mathrm{H}^{\prime 2}}=-\frac{15}{2} \frac{\mathrm{~J}_{2} \mu^{4} \mathrm{R}^{2}}{{\mathrm{~L}^{\prime} \mathrm{G}^{\prime}}^{5}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{G}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{G}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}} \frac{\partial \mathrm{b}_{3}}{\partial \mathrm{G}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{~L}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{H}^{\prime} \partial \mathrm{L}^{\prime}}=-\frac{3}{\mathrm{~L}^{\prime}} \frac{\partial \mathrm{b}_{3}}{\partial \mathrm{H}^{\prime}} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{G}^{\prime 2}}=\frac{30}{\mathrm{G}^{\prime 2}} \mathrm{~b}_{3} \quad ; \\
& \frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{G}^{\prime} \partial \mathrm{H}^{\prime}}=\frac{\partial^{2} \mathrm{~b}_{3}}{\partial \mathrm{H}^{\prime} \partial \mathrm{G}^{\prime}}=-\frac{5}{\mathrm{G}^{\prime}} \frac{\partial \mathrm{b}_{3}}{\partial \mathrm{H}^{\prime}} \quad ;  \tag{67}\\
& \frac{\partial^{2} b_{3}}{\partial{H^{\prime}}^{2}}=0 \quad ; \\
& \frac{\partial^{2} I_{0}}{\partial X \partial Y}=\frac{\partial^{2} \psi}{\partial X \partial Y}\left(\frac{1}{2 \Delta}\right)-\frac{1}{2 \Delta^{2}} \frac{\partial \Delta}{\partial Y} \frac{\partial \psi}{\partial X}+\frac{1}{2} \frac{\partial a_{X}}{\partial Y}\left(I_{-2}-I_{0}\right) \\
& +\frac{a_{x}}{2}\left[-\frac{\partial I_{0}}{\partial Y}+\frac{1}{{k^{\prime}}^{2}}\left(\frac{\partial I_{2}}{\partial Y}-\frac{\partial k^{2}}{\partial \bar{Y}} \frac{\sin \theta \cos \theta}{\Delta}+\frac{k^{2} \sin \theta \cos \theta}{\Delta^{2}} \frac{\partial \Delta}{\partial \bar{Y}}\right.\right. \\
& \left.\left.-\mathrm{k}^{2} \frac{\cos 2 \theta}{4 \Delta} \frac{\partial \psi}{\partial \mathrm{Y}}\right)+\frac{1}{\mathrm{k}^{\prime 4}} \frac{\partial \mathrm{k}^{2}}{\partial \mathrm{Y}}\left(\mathrm{I}_{2}-\frac{\mathrm{k}^{2} \sin \theta \cos \theta}{\Delta}\right)\right] ; \\
& \frac{\partial^{2} B_{2}}{\partial X \partial Y}=-\frac{2 c_{2}}{a}\left(\frac{\partial^{2} b_{2}}{\partial X \partial Y} a_{4}+\frac{\partial b_{2}}{\partial X} \frac{\partial a_{4}}{\partial Y}+\frac{\partial b_{2}}{\partial Y} \frac{\partial a_{4}}{\partial X}\right)-\frac{\epsilon}{L^{\prime}} \frac{\partial B_{2}}{\partial Y} \\
& +\eta\left[-\frac{8 c_{1}}{\mathrm{a}}\left(\frac{\partial \mathrm{~b}_{2}}{\partial \mathrm{X}} \mathrm{a}_{4}+\mathrm{b}_{2} \frac{\partial \mathrm{a}_{4}}{\partial \mathrm{X}}\right)+\frac{\epsilon}{\mathrm{L}^{\prime 2}} \mathrm{~B}_{2}\right] .
\end{align*}
$$

Note: In the expression $\partial^{2} Z_{2} / \partial X \partial Y$, we have neglected the term

$$
\mathrm{B}_{2} \int_{0}^{\theta} \frac{\partial^{2}}{\partial \mathrm{Y} \partial \mathrm{X}}\left(\frac{\mathrm{I}_{2}-\mathrm{I}_{0}}{\Delta}\right) \mathrm{dx}
$$

which is a very small quantity.

We also give the second derivatives of $S_{1 / 2}$, so that we can calculate Poisson brackets:

$$
\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \mathrm{X}}, \mathrm{~S}_{1 / 2}\right\}, \quad\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \mathrm{X}}, \mathrm{z}_{1}\right\}, \quad\left\{\frac{\partial \mathrm{S}_{1 / 2}}{\partial \mathrm{X}}, \mathrm{z}_{2}\right\}, .
$$

If the first derivatives of $S_{1 / 2}$ are in compact form,

$$
\frac{\partial S_{1 / 2}}{\partial X}=2 c_{2} \gamma\left(\frac{\partial I_{2}}{\partial X}-\frac{\partial \theta}{\partial X}\right)+\varepsilon\left(I_{2}-\theta\right)\left(8 c_{1} \gamma+\frac{2}{a}\right),
$$

where

$$
\epsilon=0 \text { if } X \neq L^{\prime} \quad \text { and } \quad \epsilon=1 \text { if } X=L^{\prime}
$$

then

$$
\begin{aligned}
\frac{\partial^{2} S_{1 / 2}}{\partial X \partial Y}= & 2 c_{2} \gamma\left(\frac{\partial^{2} I_{2}}{\partial X \partial \bar{Y}}-\frac{\partial^{2} \theta}{\partial X \partial Y}\right)+\epsilon\left(\frac{\partial I_{2}}{\partial Y}-\frac{\partial \theta}{\partial \mathbf{Y}}\right)\left(8 c_{1} \gamma+\frac{2}{a}\right) \\
& +\eta\left[\left(8 c_{1} \gamma+\frac{2}{a}\right)\left(\frac{\partial I_{2}}{\partial X}-\frac{\partial \theta}{\partial X}\right)+8 \epsilon\left(I_{2}-\theta\right)\left(\frac{{L^{\prime}}^{2}}{\mu^{2}} \gamma+\frac{1}{a L^{\prime}}\right)\right]
\end{aligned}
$$

where

$$
\eta=0 \text { if } \mathrm{Y} \neq \mathrm{L}^{\prime} \quad \text { and } \quad \eta=1 \text { if } \mathrm{Y}=\mathrm{L}^{\prime}
$$

and

$$
\frac{\partial^{2} \mathrm{I}_{2}}{\partial \mathrm{X} \partial \mathrm{Y}}=\frac{1}{2} \frac{\partial \mathrm{a}_{\mathrm{x}}}{\partial \mathrm{Y}}\left(\mathrm{I}_{2}-\mathrm{I}_{0}\right)+\frac{\Delta}{2} \frac{\partial \mathrm{p}_{\mathrm{x}}}{\partial \mathrm{Y}}+\frac{\mathrm{p}_{\mathrm{x}}}{2} \frac{\partial \Delta}{\partial \mathrm{Y}}+\frac{\mathrm{a}_{\mathrm{x}}}{2}\left(\frac{\partial \mathrm{I}_{2}}{\partial \mathrm{Y}}-\frac{\partial \mathrm{I}_{0}}{\partial \mathrm{Y}}\right)
$$

## 5. REMOVAL OF LONG-PERIOD TERMS

If we do not assume that $g^{\prime}$ is constant, as we did in Section 4, then, after eliminating $\ell^{\prime}$ and $h^{\prime}$, we are left with the following system of equations:

$$
\begin{array}{ll}
\frac{d L^{\prime}}{d t}=\frac{\partial \mathrm{F}^{*}}{\partial \ell^{\prime}}, & \frac{\mathrm{d} \ell^{\prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{*}}{\partial \mathrm{~L}^{\prime}}, \\
\frac{d G^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{*}}{\partial \mathrm{~g}^{\prime}}, & \frac{\mathrm{dg}^{\prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{*}}{\partial \mathrm{G}^{\prime}},  \tag{68}\\
\frac{d \mathrm{H}^{\prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{*}}{\partial \mathrm{~h}^{\prime}}, & \frac{d \mathrm{~h}^{\prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{*}}{\partial \mathrm{H}^{\prime}},
\end{array}
$$

where

$$
\begin{align*}
& \mathrm{F}^{*}=\mathrm{F}_{0}^{*}+\mathrm{F}_{1}^{*}+\mathrm{F}_{2}^{*} \\
& \mathrm{~F}_{0}^{*}=\frac{\mu^{2}}{2 \mathrm{~L}^{\prime 2}}+\nu \mathrm{H}^{\prime}  \tag{69}\\
& \mathrm{F}_{1}^{*}=\mathrm{F}_{1}^{*}\left(\mathrm{~L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}\right) \\
& \mathrm{F}_{2}^{*}=-\sum_{\mathrm{k}_{1}} \mathrm{D}\left(\mathrm{k}_{1}\right)=\mathrm{F}_{2}^{*}\left(\mathrm{~L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \mathrm{g}^{\prime}\right)
\end{align*}
$$

We now proceed to remove $g^{\prime}$ from the Hamiltonian, by performing a new canonical transformation using Hori's method.

We introduce a determining function $\mathrm{S}^{\prime}$ and a new Hamiltonian $\mathrm{F}^{* *}$ such that, if $\mathrm{L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}, \ell^{\prime \prime}, \mathrm{g}^{\prime \prime}$, and $\mathrm{h}^{\prime \prime}$ are the new variables,

$$
\begin{aligned}
& \mathrm{F}_{0}^{* *}=\mathrm{F}_{0}^{*}\left(\mathrm{~L}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right) \\
& \mathrm{F}_{1}^{* *}=\mathrm{F}_{1}^{*}\left(\mathrm{~L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right)
\end{aligned}
$$

and

$$
\left\{\mathrm{F}_{0}^{*}, \mathrm{~S}^{\prime}\right\}+\mathrm{F}_{2}^{*}=\mathrm{F}_{2}^{* *}
$$

Introducing a pseudotime $t^{* *}$, defined by

$$
\frac{\mathrm{dg}^{\prime \prime}}{\mathrm{dt}}{ }^{* *}=-\frac{\partial \mathrm{F}_{1}^{*}}{\partial \mathrm{G}^{\prime \prime}}\left(\mathrm{L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right)
$$

or

$$
d t^{* *}=\rho d g^{\prime \prime} \quad \text { with } \quad \rho=\left(-\frac{\partial \mathrm{F}_{1}^{*}}{\partial G^{\prime \prime}}\right)^{-1}
$$

we can solve for $\mathrm{F}_{2}^{* *}$ and $\mathrm{S}^{\prime}$ :

$$
\begin{aligned}
& \mathrm{F}_{2}^{* *}=\mathrm{F}_{2 \mathrm{~s}}^{*} \\
& \left\{\mathrm{~F}_{0}^{*}, \mathrm{~S}^{\prime}\right\}=-\mathrm{F}_{2 \mathrm{p}}^{*} \Longrightarrow \mathrm{~S}^{\prime}=\int \mathrm{F}_{2 \mathrm{p}}^{*} \mathrm{dt}
\end{aligned}
$$

Let us find the secular and periodic parts of $\mathrm{F}_{2}^{* *}$. By using expressions for only one fixed value of $\mathrm{k}_{1}$, the complete expressions of $\mathrm{F}_{2}{ }^{* *}$ and $\mathrm{S}^{\prime}$ are then

$$
\mathrm{F}_{2}^{* *}=\sum_{\mathrm{k}_{1}} \mathrm{~F}_{2}^{* *}\left(\mathrm{k}_{1}\right) \quad, \quad \mathrm{S}^{\prime}=\sum_{\mathrm{k}_{1}} \mathrm{~S}^{\prime}\left(\mathrm{k}_{1}\right)
$$

We can assume that $k_{1}$ is fixed and can drop the index $k_{1}$ in the following derivation. We then have

$$
\mathrm{F}_{2}^{*}=-\mathrm{D}
$$

where

$$
D^{2}=\left(\sum_{i} A_{i} \cos \phi_{i}\right)^{2}+\left(\sum_{i} A_{i} \sin \phi_{i}\right)^{2}
$$

in which

$$
\begin{aligned}
& A_{i}=S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right) \\
& \phi_{i}=\left(k_{0}-2 x\right) g-\Lambda_{k_{1}}, k_{0}
\end{aligned}
$$

and the summations are taken over values of x and $\mathrm{k}_{0}$ [see eqs. (8) and (9)]. Hence,

$$
\begin{aligned}
D^{2} & =\sum_{i} A_{i}^{2}+2 \sum_{i<j} A_{i} A_{j}-4 \sum_{i<j} A_{i} A_{j} \sin ^{2}\left(\frac{\phi_{i}-\phi_{j}}{2}\right) \\
& =\left(\sum A_{i}\right)^{2}\left\{1-\frac{4 \sum A_{i} A_{j} \sin ^{2}\left[\left(\phi_{i}-\phi_{j}\right) / 2\right]}{\left(\sum A_{i}\right)^{2}}\right\}
\end{aligned}
$$

To the first order,

$$
\begin{aligned}
& D=\left(\sum_{i} A_{i}\right)\left\{1-\frac{2 \sum A_{i} A_{j} \sin ^{2}\left[\left(\phi_{i}-\phi_{j}\right) / 2\right]}{\left(\sum A_{i}\right)^{2}}\right\} \\
& D=\sum_{i} A_{i}-\frac{\sum_{i} A_{i} A_{j}}{\sum_{i} A_{i}}+\frac{\sum A_{i} A_{j} \cos \left(\phi_{i}-\phi_{j}\right)}{\sum_{i} A_{i}}
\end{aligned}
$$

Therefore,

$$
\begin{align*}
& F_{2 s}=\frac{\sum A_{i} A_{j}}{\sum A_{i}}-\sum A_{i}  \tag{70a}\\
& F_{2 p}=-\frac{\sum A_{i} A_{j} \cos \left(\phi_{i}-\phi_{j}\right)}{\sum A_{i}} \tag{70b}
\end{align*}
$$

and

$$
s_{2}=-\rho \int \frac{\sum A_{i} A_{j} \cos \left(\phi_{i}-\phi_{j}\right)}{\sum A_{i}} d g^{\prime \prime}=-\frac{\rho}{\sum A_{i}} \sum_{i<j}\left[\frac{A_{i} A_{j} \sin \left(\phi_{i}-\phi_{j}\right)}{\left(k_{0}-2 x\right)_{i}-\left(k_{0}-2 x\right)_{j}}\right]
$$

If we set $\left(k_{0}-2 x\right)_{i}-\left(k_{0}-2 x\right)_{j}=a_{i j}$, then

$$
\begin{aligned}
& \frac{\partial S_{2}}{\partial \ell^{\prime \prime}}=\frac{\partial S_{2}}{\partial h^{\prime \prime}}=0 ; \\
& \frac{\partial S_{2}}{\partial g^{\prime \prime}}=-\frac{\rho}{\sum_{i} A_{i}} \sum_{i<j} A_{i} A_{j} \cos \left(\phi_{i}-\phi_{j}\right) ;
\end{aligned}
$$

and, if $\mathrm{X}=\mathrm{L}^{\prime \prime}, \mathrm{G}^{\prime \prime}$, or $\mathrm{H}^{\prime \prime}$,

$$
\begin{align*}
\frac{\partial S_{2}}{\partial X}= & -\frac{\rho}{\sum A_{i}}\left[\sum_{i<j} \frac{\left(A_{i} A_{j, x}+A_{i, x}^{\prime} A_{j}\right) \sin \left(\phi_{i}-\phi_{j}\right)}{a_{i j}}\right] \\
& +\left[\frac{1}{\sum A_{i}} \frac{\partial \rho}{\partial X}-\frac{\rho}{\left(\sum A_{i}\right)^{2}} \sum A_{i, x}\right]\left[\sum \frac{A_{i} A_{j} \sin \left(\phi_{i}-\phi_{j}\right)}{a_{i j}}\right] \tag{71}
\end{align*}
$$

where

$$
\begin{aligned}
& A_{i, x}=\frac{\partial A_{i}}{\partial X}=\frac{\partial S\left(k_{1}, k_{0}\right)}{\partial X} A\left(k_{1}, k_{0}, x\right)+S\left(k_{1}, k_{0}\right) \frac{\partial A}{\partial X}\left(k_{1}, k_{0}, x\right) \\
& \quad \text { [see Appendix A, eqs. (A-2)] ; } \\
& \frac{\partial F_{2}}{\partial X}=\frac{\sum\left(A_{i} A_{j, x}+A_{i, x} A_{j}\right)}{\sum A_{i}}-\sum A_{i, x}-\frac{1}{\left(\sum_{i} A_{i}\right)^{2}}\left(\sum A_{i} A_{j}\right)\left(\sum A_{i, x}\right) .
\end{aligned}
$$

The new Hamiltonian becomes

$$
\mathrm{F}^{* *}=\mathrm{F}_{0}^{* *}+\mathrm{F}_{1}^{* *}+\mathrm{F}_{2}^{* *},
$$

where

$$
\begin{align*}
& \mathrm{F}_{0}^{* *}=\frac{\mu^{2}}{2 \mathrm{~L}^{\prime \prime}}+\nu \mathrm{H}^{\prime \prime}, \\
& \mathrm{F}_{1}^{* *}=\mathrm{F}_{1 \mathrm{~s}}\left(\mathrm{~L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right),  \tag{72}\\
& \mathrm{F}_{2}^{* *}=\mathrm{F}_{2 \mathrm{~s}}^{*}\left(\mathrm{~L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right) ;
\end{align*}
$$

and the passage from the old variables ( $\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}$ ) to the new ones is given by

$$
\begin{align*}
& \mathrm{L}^{\prime}=\mathrm{L}^{\prime \prime}+\frac{\partial \mathrm{S}^{\prime}}{\partial \ell^{\prime \prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \ell^{\prime \prime}}, \mathrm{S}^{\prime}\right\}+\ldots, \\
& \mathrm{G}^{\prime}=\mathrm{G}^{\prime \prime}+\frac{\partial \mathrm{S}^{\prime}}{\partial \mathrm{g}^{\prime \prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \mathrm{g}^{\prime \prime \prime}}, \mathrm{S}^{\prime}\right\}+\ldots, \\
& \mathrm{H}^{\prime}=\mathrm{H}^{\prime \prime}+\frac{\partial \mathrm{S}^{\prime}}{\partial \mathrm{h}^{\prime \prime}}+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \mathrm{h}^{\prime \prime}}, \mathrm{S}^{\prime}\right\}+\ldots,  \tag{73}\\
& \ell^{\prime}=\ell^{\prime \prime}-\frac{\partial \mathbf{S}^{\prime}}{\partial \mathrm{L}^{\prime \prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \mathrm{L}^{\prime \prime}}, \mathrm{S}^{\prime}\right\}-\ldots, \\
& \mathrm{g}^{\prime}=\mathrm{g}^{\prime \prime}-\frac{\partial \mathrm{S}^{\prime}}{\partial \mathrm{G}^{\prime \prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \mathrm{G}^{\prime \prime}}, \mathrm{S}^{\prime}\right\}-\ldots, \\
& \mathbf{h}^{\prime}=\mathrm{h}^{\prime \prime}-\frac{\partial \mathrm{S}^{\prime}}{\partial \mathrm{H}^{\prime \prime}}-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}^{\prime}}{\partial \mathrm{H}^{\prime \prime}}, \mathrm{S}^{\prime}\right\}-\ldots,
\end{align*}
$$

The new system of equations is given by

$$
\begin{array}{ll}
\frac{\mathrm{dL}^{\prime \prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{* *}}{\partial \ell^{\prime \prime}}, & \frac{\mathrm{d} \ell^{\prime \prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{~L}^{\prime \prime}}, \\
\frac{\mathrm{dG}^{\prime \prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{~g}^{\prime \prime}}, & \frac{\mathrm{dg}^{\prime \prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{G}^{\prime \prime}},  \tag{74}\\
\frac{\mathrm{dH}^{\prime \prime}}{\mathrm{dt}}=\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{~h}^{\prime \prime}}, & \frac{\mathrm{dh}^{\prime \prime}}{\mathrm{dt}}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{H}^{\prime \prime}},
\end{array}
$$

and, since $\mathrm{F}^{* *}=\mathrm{F}^{* *}\left(\mathrm{~L}^{\prime \prime}, \mathrm{G}^{\prime \prime}, \mathrm{H}^{\prime \prime}\right)$, the solution is

$$
\begin{align*}
& \mathrm{L}^{\prime \prime}=\mathrm{L}_{0}^{\prime \prime}=\text { constant }, \\
& \mathrm{G}^{\prime \prime}=\mathrm{G}_{0}^{\prime \prime}=\mathrm{constan} \mathrm{t}, \\
& \mathrm{H}^{\prime \prime}=\mathrm{H}_{0}^{\prime \prime}=\mathrm{constant},  \tag{75}\\
& \ell^{\prime \prime}=\ell_{0}^{\prime \prime}+\ell^{\prime \prime}, \mathrm{t}, \\
& \mathrm{~g}^{\prime \prime}=\mathrm{g}_{0}^{\prime \prime}+\mathrm{g}^{\prime \prime}, \mathrm{t}, \\
& \mathrm{~h}^{\prime \prime}=\mathrm{h}_{0}^{\prime \prime}+\mathrm{h}^{\prime \prime}, \mathrm{t},
\end{align*}
$$

where $\ell_{0}^{\prime \prime}, \mathrm{g}_{0}^{\prime \prime}, \mathrm{h}_{0}^{\prime \prime}, \ell_{1}^{\prime \prime}, \mathrm{g}_{1}^{\prime \prime}, \mathrm{h}_{1}^{\prime \prime}$ are constants and

$$
\begin{align*}
& \ell_{1}^{\prime \prime}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{~L}^{\prime \prime}}=\frac{\mu^{2}}{\mathrm{~L}^{\prime \prime}}-\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{~L}^{\prime \prime}}-\frac{\partial \mathrm{F}_{2 \mathrm{~s}}^{*}}{\partial \mathrm{~L}^{\prime \prime}}, \\
& \mathrm{g}_{1}^{\prime \prime}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{G}^{\prime \prime}}=-\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{G}^{\prime \prime}}-\frac{\partial \mathrm{F}_{2 \mathrm{~s}}^{*}}{\partial \mathrm{G}^{\prime \prime}},  \tag{76}\\
& \mathrm{h}_{1}^{\prime \prime}=-\frac{\partial \mathrm{F}^{* *}}{\partial \mathrm{H}^{\prime \prime}}=-v-\frac{\partial \mathrm{F}_{1 \mathrm{~s}}}{\partial \mathrm{H}^{\prime \prime}}-\frac{\partial \mathrm{F}_{2 \mathrm{~s}}^{*}}{\partial \mathrm{H}^{\prime \prime}} .
\end{align*}
$$

Returning to the initial set of modified Delaunay variables ( $L, G, H, \ell, g, h$ ), we now have

$$
\begin{align*}
& \mathrm{L}=\mathrm{L}^{\prime}+\frac{\partial \mathrm{S}}{\partial \ell^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \ell^{\prime}}, \mathrm{S}\right\}+\ldots, \\
& \mathrm{G}=\mathrm{G}^{\prime}+\frac{\partial \mathrm{S}}{\partial \mathrm{~g}^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~g}^{\prime}}, \mathrm{S}\right\}+\ldots, \\
& \mathrm{H}=\mathrm{H}^{\prime}+\frac{\partial \mathrm{S}}{\partial \mathrm{~h}^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)+\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~h}^{\prime}}, \mathrm{S}\right\}+\ldots,  \tag{77}\\
& \ell=\ell^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{~L}^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{~L}^{\prime}}, \mathrm{S}\right\}-\ldots, \\
& \mathrm{g}=\mathrm{g}^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{G}^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{G}^{\prime}}, \mathrm{S}\right\}-\ldots, \\
& \mathrm{h}=\mathrm{h}^{\prime}-\frac{\partial \mathrm{S}}{\partial \mathrm{H}^{\prime}}\left(\mathrm{L}^{\prime}, \mathrm{G}^{\prime}, \mathrm{H}^{\prime}, \ell^{\prime}, \mathrm{g}^{\prime}, \mathrm{h}^{\prime}\right)-\frac{1}{2}\left\{\frac{\partial \mathrm{~S}}{\partial \mathrm{H}^{\prime}}, \mathrm{S}\right\}-\cdots,
\end{align*}
$$

where $\mathrm{S}=\mathrm{S}_{1 / 2}+\mathrm{SJ}+\mathrm{Z}_{1}+\mathrm{Z}_{2}$, which, with the help of equations (28), gives the complete solution.

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## APPENDIX A

## DERIVATIVES OF D( $\left.\mathrm{k}_{1}\right)$ AND $\psi\left(\mathrm{k}_{1}\right)$

## A. 1 First Derivatives

We recall the following equations from Section 3:

$$
\begin{align*}
& \mathrm{D}^{2}\left(\mathrm{k}_{1}\right)=\mathrm{B}^{2}\left(\mathrm{k}_{1}\right)+\mathrm{C}^{2}\left(\mathrm{k}_{1}\right) \\
& \mathrm{D}\left(\mathrm{k}_{1}\right) \cos \psi\left(\mathrm{k}_{1}\right)=\mathrm{B}\left(\mathrm{k}_{1}\right)  \tag{A-1}\\
& \mathrm{D}\left(\mathrm{k}_{1}\right) \sin \psi\left(\mathrm{k}_{1}\right)=\mathrm{C}\left(\mathrm{k}_{1}\right)
\end{align*}
$$

Then, if X is any of the Delaunay variables $\mathrm{L}, \mathrm{G}, \mathrm{H}$, or g ,

$$
D \frac{\partial D}{\partial X}=B \frac{\partial B}{\partial X}+C \frac{\partial C}{\partial X}
$$

If $X \neq \mathrm{g}$,

$$
\begin{align*}
& \frac{\partial B}{\partial X}=\sum_{k_{0}} \sum_{x=0}^{k_{1} \beta+k_{0}} \frac{\partial}{\partial X}\left[S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right)\right] \cos \left[\left(k_{0}-2 x\right) g-\Lambda_{k_{1}, k_{0}}\right] \\
& \frac{\partial C}{\partial X}=\sum_{k_{0}} \sum_{x=0}^{k_{1} \beta+k_{0}} \frac{\partial}{\partial \bar{X}}\left[S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right)\right] \sin \left[\left(k_{0}-2 x\right) g-\Lambda_{\left.k_{1}, k_{0}\right]}\right. \tag{A-2}
\end{align*}
$$

$\frac{\partial A\left(k_{1}, k_{0}, x\right)}{\partial X}=\frac{\partial F_{n, m, p^{(I)}}^{\partial \bar{I}}}{\partial I} \frac{\partial I}{\partial X} G(e)+F_{n, m, p^{(I)}}^{\partial G_{n, p, q^{(e)}}^{\partial e}} \frac{\partial e}{\partial X} \quad$,
$\frac{\partial S\left(k_{1} k_{0}\right)}{\partial X}=\left\{\begin{array}{lll}0 & \text { if } & X \neq L \quad, \\ \frac{-2\left(\beta k_{1}+k_{0}+1\right)}{L} S\left(k_{1}, k_{0}\right) & \text { if } & X=L .\end{array}\right.$.

If $\mathrm{X}=\mathrm{g}$,

$$
\begin{aligned}
& \frac{\partial B}{\partial g}=-\sum_{k_{0}} \sum_{x=0}^{\beta k_{1}+k_{0}} S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right)\left(k_{0}-2 x\right) \sin \left[\left(k_{0}-2 x\right) g-\Lambda_{\left.k_{1}, k_{0}\right]}\right. \\
& \frac{\partial C}{\partial g}=\sum_{k_{0}} \sum_{x=0}^{\beta k_{1}+k_{0}} S\left(k_{1}, k_{0}\right) A\left(k_{1}, k_{0}, x\right)\left(k_{0}-2 x\right) \cos \left[\left(k_{0}-2 x\right) g-\Lambda_{\left.k_{1}, k_{0}\right]}\right.
\end{aligned}
$$

For the derivatives of $\psi\left(\mathrm{k}_{1}\right)$, we have

$$
\mathrm{C}\left(\mathrm{k}_{1}\right)=\mathrm{D}\left(\mathrm{k}_{1}\right) \sin \psi\left(\mathrm{k}_{1}\right)
$$

so that

$$
\begin{equation*}
\frac{\partial \mathrm{C}\left(\mathrm{k}_{1}\right)}{\partial \mathrm{X}}=\mathrm{D}\left(\mathrm{k}_{1}\right) \frac{\partial \psi\left(\mathrm{k}_{1}\right)}{\partial \bar{X}} \cos \psi\left(\mathrm{k}_{1}\right)+\frac{\partial \mathrm{D}\left(\mathrm{k}_{1}\right)}{\partial \mathrm{X}} \sin \psi\left(\mathrm{k}_{1}\right) \tag{A-4a}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial \psi\left(\mathrm{k}_{1}\right)}{\partial \mathbf{X}}=\frac{1}{\mathbf{B}\left(\mathrm{k}_{1}\right)}\left[\frac{\partial \mathbf{C}\left(\mathrm{k}_{1}\right)}{\partial \mathbf{X}}-\frac{1}{\mathrm{D}\left(\mathrm{k}_{1}\right)} \frac{\partial \mathrm{D}\left(\mathrm{k}_{1}\right)}{\partial \mathbf{X}} \mathrm{C}\left(\mathrm{k}_{1}\right)\right] . \tag{A-4b}
\end{equation*}
$$

As we did in Section 2, we can now write

$$
\begin{array}{ll}
a_{1}=\frac{1}{D_{1}\left(k_{1}\right)} \frac{\partial D}{\partial L}\left(k_{1}\right), & a_{2}=\frac{1}{D\left(k_{1}\right)} \frac{\partial D}{\partial G}\left(k_{1}\right), \\
a_{3}=\frac{1}{D\left(k_{1}\right)} \frac{\partial D}{\partial H}\left(k_{1}\right), & a_{4}=\frac{1}{D\left(k_{1}\right)} \frac{\partial D}{\partial g}\left(k_{1}\right), \\
p_{1}=\frac{\partial \psi}{\partial L}\left(k_{1}\right), & p_{3}=\frac{\partial \psi}{\partial H}\left(k_{1}\right), \\
p_{2}=\frac{\partial \psi}{\partial G}\left(k_{1}\right), & p_{4}=\frac{\partial \psi}{\partial g}\left(k_{1}\right) .
\end{array}
$$

## A. 2 Second Derivatives

Let us drop the index $\mathrm{k}_{1}$ for simplicity:

$$
\begin{align*}
& \frac{1}{D} \frac{\partial D}{\partial \bar{X}}=\frac{1}{D^{2}}\left(B \frac{\partial B}{\partial \bar{X}}+C \frac{\partial C}{\partial X}\right) \\
& \frac{\partial}{\partial Y}\left(\frac{1}{D} \frac{\partial D}{\partial \bar{X}}\right)=-\frac{2}{D^{2}}\left(\frac{1}{D} \frac{\partial D}{\partial \bar{Y}}\right)\left(B \frac{\partial B}{\partial \bar{X}}+C \frac{\partial C}{\partial X}\right) \\
& +\frac{1}{D^{2}}\left(\frac{\partial B}{\partial Y} \frac{\partial B}{\partial \bar{X}}+\frac{\partial C}{\partial Y} \frac{\partial C}{\partial X}+C \frac{\partial^{2} C}{\partial X \partial Y}+B \frac{\partial^{2} B}{\partial X \partial Y}\right) . \tag{A-5}
\end{align*}
$$

If $Y \neq g$ and $X \neq g$, then

$$
\begin{equation*}
\frac{\partial^{2} \mathrm{~B}}{\partial \mathrm{X} \partial \mathrm{Y}}=\sum_{\mathrm{k}_{0}} \sum_{\mathrm{X}} \frac{\partial^{2}}{\partial \mathrm{X} \partial \overline{\mathrm{Y}}}(\mathrm{SA}) \cos \left[\left(\mathrm{k}_{0}-2 \mathrm{x}\right) \mathrm{g}-\Lambda_{\mathrm{k}_{1}}, \mathrm{k}_{0}\right] \tag{A-6a}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial^{2} \mathrm{C}}{\partial \mathrm{X} \partial \mathrm{Y}}=\sum_{\mathrm{k}_{0}} \sum_{\mathrm{x}} \frac{\partial^{2}}{\partial \mathrm{X} \partial \mathrm{Y}}(\mathrm{SA}) \sin \left[\left(\mathrm{k}_{0}-2 \mathrm{x}\right) \mathrm{g}-\Lambda_{\left.\mathrm{k}_{1}, \mathrm{k}_{0}\right]}\right] \tag{A-6b}
\end{equation*}
$$

If $X \neq g$ and $Y=g$, we have

$$
\begin{equation*}
\frac{\partial^{2} B}{\partial \bar{X} \partial g}=-\sum_{k_{0}} \sum_{x} \frac{\partial S A}{\partial \bar{X}}\left(k_{0}-2 x\right) \sin \left[\left(k_{0}-2 x\right) g-\Lambda_{n, m}\right] \tag{A-7a}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial^{2} C}{\partial X \partial g}=\sum_{k_{0}} \sum_{x} \frac{\partial S A}{\partial X}\left(k_{0}-2 x\right) \cos \left[\left(k_{0}-2 x\right) g-\Lambda_{n, m}\right] \tag{A-7b}
\end{equation*}
$$

If $X=g$ and $Y=g$, then

$$
\begin{equation*}
\frac{\partial^{2} \mathrm{~B}}{\partial \mathrm{~g}^{2}}=-\sum_{\mathrm{k}_{0}} \sum_{\mathrm{x}} \mathrm{SA}\left(\mathrm{k}_{0}-2 \mathrm{x}\right)^{2} \cos \left[\left(\mathrm{k}_{0}-2 \mathrm{x}\right) \mathrm{g}-\Lambda_{\mathrm{k}_{1}, \mathrm{k}_{0}}\right] \tag{A-8a}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\partial^{2} C}{\partial g^{2}}=-\sum_{\mathrm{k}_{0}} \sum_{\mathrm{x}} \mathrm{SA}\left(\mathrm{k}_{0}-2 \mathrm{x}\right)^{2} \sin \left[\left(\mathrm{k}_{0}-2 \mathrm{x}\right) \mathrm{g}-\Lambda_{\left.\mathrm{k}_{1}, \mathrm{k}_{0}\right]}\right] \tag{A-8b}
\end{equation*}
$$

We find the following second derivative of $\psi\left(\mathrm{k}_{1}\right)$ from equations (A-4):

$$
\begin{equation*}
\frac{\partial \mathrm{B}}{\partial \mathbf{Y}} \frac{\partial \psi}{\partial \mathbf{X}}+\mathrm{B} \frac{\partial^{2} \psi}{\partial \mathbf{X} \partial \mathbf{Y}}+\frac{\partial \mathrm{C}}{\partial \bar{Y}} \frac{1}{\mathrm{D}} \frac{\partial \mathrm{D}}{\partial \mathbf{X}}+\mathbf{C} \frac{\partial}{\partial \bar{Y}}\left(\frac{1}{\mathrm{D}} \frac{\partial \mathrm{D}}{\partial \mathbf{X}}\right)=\frac{\partial^{2} \mathrm{C}}{\partial \mathbf{X} \partial \mathbf{Y}} \tag{A-9a}
\end{equation*}
$$

so that

$$
\begin{equation*}
\frac{\partial^{2} \psi}{\partial X \partial Y}=\frac{1}{\bar{B}}\left[\frac{\partial^{2} C}{\partial X \partial Y}-\frac{\partial B}{\partial Y} \frac{\partial \psi}{\partial X}-\frac{1}{D} \frac{\partial D}{\partial X} \frac{\partial C}{\partial Y}-C \frac{\partial}{\partial \bar{Y}}\left(\frac{1}{D} \frac{\partial D}{\partial \bar{X}}\right)\right] \tag{A-9b}
\end{equation*}
$$

## BIOGRAPHICAL NOTE

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Ms. Romanowicz was a member of the Analytical Geophysics Department at Smithsonian Astrophysical Observatory in 1973 and 1974, specializing in celestial mechanics. She is currently in a graduate program in applied mathematics at Harvard University.

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