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## EVOLUTION OF WEAK DISTURBANCES

IN INERT BINARY MIXTURES

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## AUGUST 1977

# EVOLUTION OF WEAK DISTURBANCES IN INERT BINARY MIXTURES 

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August 1977

The work here presented was written while the author was a visitor from the University of Oklahoma to the Joint Institute for Aeronautics and Acoustics at Stanford University and NASA Ames Research Center. The author is grateful to Professor K. Karamcheti for organizing that visit and giving him the opportunity for valuable discussions with members of the Institute.

## ABSTRACT

The evolution of weak disturbances in inert binary mixtures is determined for the one-dimensinnal pistar problem. The interaction of the dissi' t tive and nonlinear mechanisms is described by Burgers' equation. The binary-mixture diffusion mechanisms enter as an additive term in an effective diffusivity. Results for the impulsive motion of a piston moving into an ambient medium and the sinusoidally oscillating piston are used to illustrate the results and elucidate the incorrect behavior pertaining to the associated linear theory.

## ACKNOWLEDGEMENT


#### Abstract

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## 1. INTRODUCTION

Combined dissipative effects of viscosity, thermal conduction, and mass diffusion play fundamental roles in the propagation of disturbances within mixtures of gases. In many practical applications the behavior of weak disturbances is of primary concern. Whereas the acoustic, or linear, limit leads' to great mathematical simplications, there are situations in which nonlinear cumulative effects cannot be ignored. In this paper we wish to descrin the interaction of nonlinear and the alovementioned dissipative mechanisms as they pertain to the evol ition of weak plane waves in inert binary mixtures. This is done hy means of the classical piston problem.

Some of the basics of acoustics for binary mixtures, especially with regards to sound absorption, were set forth by Kohler (1949). The propagation of spherical acoustic disturbances in binary mixtures, with thermal diffusion ignored, was studied by Rasmussen and Frair (1976). For binary mixtures, the study of nonlinear behavior appears to be limited to steady-state shock waves. Dyakov (1954) studied weak shocks and Sherman (1960) studied both weak and strong shocks. On the other hand, for pure gases, when mass diffusion does not play a role, combined nonlinear and dissipative effects on the evolution of finice disturbances have been the subject of investigation for some time. The classic early work is that of Lighthill (1956), and a summary of other related work dealing witn Burgers' equation is given by Benton and Platzmann (1972). A more recent work is that of Halabisky and Sirovich (1973). Also, for pure gases, Shidlovsky $(1975,1977)$ has investigated the evolution of shock waves and other regions of nonuniformity by means of boundarylayer singular-perturbation methods.

General model equations have been proposed by Blythe (1969) and Ockendon and Spence (1969) for the evolution of waves in relaxing gases. The thrust of their work was toward inviscid flows, but their models could also deal with the viscous-type relaxation under consideration in the present work.

In this investigation, we describe how weak, one-dimensional disturbances in binary inert mixtures evolve with time. In particular, we proceed from the basic linear theory and then illustrate how the smallest nonlinear terms modify the evolutionary description. For the piston problem, the wave front is shown to be governed by Burgers' aquation. The contributions of the binary-mixture diffusion mechanisms enter as an additive term in an effective diffusivity that is a combination of the Schmidt number and pressure and thermal diffusion coefficients. Besides the mass average velocity perturbation, results are obtained for the species mass-fraction perturbation, which is pertinent to binary mixtures.

## 2. BASIC EQUATIONS

The equations of motion for mass, species, momentum, and energy for an inert binary mixture are (with body forces neglected)

$$
\begin{gather*}
\frac{D \rho}{D t}+\rho \operatorname{div} \overrightarrow{\mathrm{V}}=0  \tag{1}\\
\rho \frac{\mathrm{DC}}{\mathrm{Dt}}=-\operatorname{div} \vec{i}  \tag{2}\\
\rho \frac{D \vec{V}}{D t}=-\operatorname{grad} p+\operatorname{div} \stackrel{\leftrightarrow}{\tau}  \tag{3}\\
\rho \frac{D h}{D t}=\frac{D p}{D t}+\overleftrightarrow{\tau}: \leftrightarrow \stackrel{\leftrightarrow}{\varepsilon}-\operatorname{div} \vec{q} \tag{4}
\end{gather*}
$$

where $\rho, p, h$, and $\vec{v}$ are the density, pressure, specific enthalpy, and velocity of the mixture. The mass fractions of the two inert species of the binary mixture are denoted by $c_{1}=c$ and $c_{2}=1-c$. The rate of strain tensor, $\overleftrightarrow{\varepsilon}$, and viscous stress tensor, $\stackrel{\rightarrow}{\tau}$, are given by

$$
\begin{align*}
& \overleftrightarrow{E}=\frac{1}{2}\left[\operatorname{grad} \vec{v}+(\operatorname{grad} \vec{v})^{t}\right]  \tag{5}\\
& \overleftrightarrow{\tau}=2 \overleftrightarrow{\varepsilon}+\lambda(\operatorname{div} \overrightarrow{\mathrm{V}}) \overleftrightarrow{\mathrm{I}} \tag{6}
\end{align*}
$$

where $\mu$ and $\lambda$ are the first and second coefficients of viscosity. The heat-flux vector, $\vec{q}$, and binary diffusion-flux vector, $\vec{i}$, are determined by means of kinetic theory (Hirschfelder et al., 1954) or principles of continuum mechanics (Landau and Lifshitz, 1959):

$$
\begin{gather*}
\vec{q}=-k \operatorname{grad} T+\left[h_{2}-h_{1}+M_{12} k_{T} \mid \vec{i}\right.  \tag{7}\\
\vec{i}=-\rho D_{12}\left[\operatorname{grad} c+k_{p} \operatorname{grad}(\ln p)+k_{T} \operatorname{grad}(\ln T)\right] \tag{8}
\end{gather*}
$$

Here $k$ is the thermal conductivity, $D_{12}$ is the binary diffusion coefficient, and $M_{12}$ is given by

$$
\begin{equation*}
M_{12} \equiv \frac{T}{k_{p}}\left(R_{1}-R_{2}\right) \tag{9}
\end{equation*}
$$

where $T$ is the temperature, $R_{1}$ and $R_{2}$ are the specific gas constants for species 1 and 2 , and $k_{p}$ and $k_{T}$ are the pressure-diffusion and thermaldiffusion suefficients, given by

$$
\begin{gather*}
\mathbf{k}_{\mathbf{p}} \equiv \frac{\left(R_{1}-R_{2}\right)\left[\left(R_{1}-R_{2}\right) c+R_{2}\right] c(1-c)}{R_{1} R_{2}}  \tag{10}\\
\mathbf{k}_{T} \equiv \alpha c(1-c) \tag{11}
\end{gather*}
$$

The thermal diffusion factor $\alpha$ is usually positive when $c_{1}=c$ refers to the heavier molecular species, but may be slightly negative for exceptional gas pairs. For Maxwellian interaction potentials, $\alpha$ vanishes. The set of equations becomes complete with the addition of thermal and caloric equations of state. For a mixture of thermally perfect gases, we have

$$
\begin{equation*}
p=\rho T\left[\left(R_{1}-R_{2}\right)^{\prime} c+R_{2}\right] \quad, \quad h=\left(h_{1}-h_{2}\right) c+h_{2} \tag{12}
\end{equation*}
$$

## 3. LINEARIZED EQUATIONS

Consider a uniform ambient state denoted by the subscript naught. We consider perturbations about this state and write

$$
\left.\begin{array}{c}
p=p_{0}\left(1+p^{\prime}\right) \quad, \rho=\rho_{0}\left(1+\rho^{\prime}\right) \quad, \quad T=T_{0}\left(1+T^{\prime}\right)  \tag{13}\\
c=c_{0}+c^{\prime} \quad, \quad \vec{v}^{\prime}=\vec{v} / a_{f_{0}}
\end{array}\right\}
$$

where $a_{f_{0}}$ denotes the frozen speed of sound in the ambient state,

$$
\begin{equation*}
a_{f_{0}}^{2} \equiv \gamma p_{o} / \rho_{0} \quad, \frac{\gamma-1}{\gamma} \equiv \frac{\left(R_{1}-R_{2}\right) c_{o}+R_{2}}{\left(C_{p_{1}}-C_{p_{2}}\right) c_{0}+c_{p_{2}}} \tag{14}
\end{equation*}
$$

and where $C_{p_{1}}$ and $C_{p_{2}}$ are the ambient-state constant-pressure specific heats of species 1 and ${ }^{\circ}$. Further, we introduce a nondimensional time, $\tau$, and position vector, $\vec{r}$, as

$$
\begin{equation*}
\tau \equiv a_{f_{0}}^{2} t / \tilde{v}_{0} \quad, \quad \vec{r} \equiv a_{f_{0}} \stackrel{\rightharpoonup}{r} / \tilde{v}_{0} \tag{15}
\end{equation*}
$$

such that $\nabla \equiv \partial / \partial \vec{r}$ (the barred space variables are dimensional). Chariacteristic Prandtl and Schmidt numbers are defined as

$$
\begin{equation*}
P_{r} \equiv \rho_{0} \tilde{v}_{0} c_{p_{0}} / k_{0} \quad, \quad s_{c} \equiv \tilde{v}_{0} / D_{12} \tag{16}
\end{equation*}
$$

where

$$
\begin{gather*}
\tilde{\nu}_{0} \equiv\left(2 \mu_{0}+\lambda_{0}\right) / \rho_{0}=\left(\frac{4}{3} \dot{\rho}_{0}+k_{0}\right) / \rho_{0}, \\
c_{p_{0}} \equiv\left(c_{p_{10}}-c_{p_{20}}\right) c_{0}+c_{p_{20}}, \tag{17}
\end{gather*}
$$

and $k_{0}$ is the bulk mudulus of viscosity in the ambient medium.
W , ow consider flows that are irrotational (such as occur with planar, cylindrical, or spherical symmetry) and introduce a velocity potential such that $\vec{v}^{\prime}=\nabla \phi$. The linearized momentum equation can thus be integrated once, and the governing linearized system of equations becomes

$$
\begin{align*}
& \frac{\partial \rho^{\prime}}{\partial \tau}+\nabla^{2} \phi=0 \quad, \quad D_{s^{\prime}} c^{\prime}=k_{p_{0}} \nabla^{2} p^{\prime}+k_{T_{0}} \nabla^{2} T^{\prime}  \tag{18}\\
& \mathrm{p}^{\prime}=\gamma\left[\nabla^{2} \phi-\frac{\partial \phi}{\partial \tau}\right] \quad, \mathrm{p}^{\prime}=\rho^{\prime}+\mathrm{T}^{\prime}+\mathrm{N}_{\mathrm{o}} \mathrm{c}^{\prime} \tag{19}
\end{align*}
$$

where

$$
\begin{gather*}
M_{o} \equiv M_{12} / C_{P_{0}} T_{o}=\frac{\gamma-1}{\gamma} \frac{N_{o}}{k_{P_{0}}}  \tag{21}\\
N_{0} \equiv \frac{R_{1}-R_{2}}{\left(R_{1}-R_{2}\right) C_{o}+R_{2}}  \tag{22}\\
D_{s}=S_{C} \frac{\partial}{\partial t}-V^{2}, D_{p} \not P_{r} \frac{\partial}{\partial t}-V^{2} \tag{23}
\end{gather*}
$$

The above equations can be manipulated so that the following single seventh-order equation for $\phi$ is obtained:

$$
\begin{array}{r}
\gamma\left(1+N_{0} k_{p_{0}}\right) V^{\sigma_{Q_{1}}+V^{4}\left|\nabla^{2} \phi-\tilde{\alpha} \phi_{\tau \tau}\right|} \\
+V^{2}\left|\beta_{1} \phi_{\tau \tau}-B_{2} V^{2} \phi_{\phi}\right|_{\tau}+p_{r} S_{c}\left|\nabla^{2} \phi-\phi_{\tau \tau}\right|_{\tau \tau}=0 \tag{24}
\end{array}
$$

## where

$$
\begin{gather*}
\tilde{\alpha} \equiv P_{r}\left(1-N_{0} k_{T_{0}}\right)+\gamma\left(1+S_{c}\right)+\gamma N_{0} k_{p_{0}}+\gamma P_{r}\left(k_{p_{0}}+k_{T_{0}}\right)\left(N_{0}+M_{0} k_{T_{0}}\right) \\
\beta_{1} \equiv \tilde{\alpha}+S_{c} P_{r}-\gamma\left(1+N_{0} k_{D_{0}}\right)  \tag{25}\\
B_{2} \equiv S_{c}+P_{r}\left(1+M_{0} k_{T}^{2}\right)
\end{gather*}
$$

When $k_{T_{0}}=0$, the above linear equations reduce to those of Pasmussen and Frair (1976) who studied spherical waves by means of these equations.

The solution of the linearized equations for a one-dimensional piston moving into an ambient medium requires three boundary conditions at the piston surface: (1) the mass mean velocity, $\phi_{x}$, is equal to the piston velociiy, (2) the diffusion-flux normal to the piston vonishes, $\vec{i} \cdot \hat{n}=0$, and (3) either the temperature or the normal heat flux, $\vec{q} \cdot \hat{A}$, is specified. The solution is obtained analogously to the spherical explosion problem of Rasmussen and Frair (1976). The solution can b. written as a part that describes a diffused wave front and two coupled parts that aescribe heat and mass-diffusion boundary layers adjacent to the piston face. The heat and mass-diffusion boundary layers spread from the piston face as the square root of time whereas the position of the bulk wave front travels like the time itself. When we fix our attention on the wave front for large times, therefore, we can ignore the boundary layers since they are qreatly outdistanced by the wave front. The asymptotic lonq-time wave behavior can be obtained by laplace-transform techniques (Rasmassen 1975). For the velocity and nass-fraction perturbition wr have for the impulsive motion of a piston, wher. $\$_{x}(x=0,1>0)$ $\mathrm{U}_{\mathrm{p}} / \mathrm{fo}^{\prime}$

$$
\begin{align*}
& v^{\prime}(x, \tau) \sim \frac{1}{2} \frac{U_{p}}{a_{f o}}\left[\operatorname{erfc}\left\{\frac{b(x-\tau)}{\sqrt{\tau}}\right\}+e^{4 b^{2} x^{2}} \operatorname{erfc}\left\{\frac{b(x+\tau)}{\sqrt{\tau}}\right\}\right]  \tag{26}\\
& c^{\prime}(x, \tau) \sim \frac{U_{p}\left[\gamma k_{p o}+(\gamma-1) k_{T o}\right]}{a_{f o} S_{c}} \frac{b x}{\sqrt{\pi \tau^{3}}} \exp \left\{-\frac{b^{2}(x-\tau)^{2}}{\tau}\right\} \tag{27}
\end{align*}
$$

where

$$
\begin{equation*}
b \equiv\left[2\left(\frac{\beta_{1}-\beta_{2}}{P_{r} S_{c}}\right)\right]^{-\frac{1}{2}} \tag{28}
\end{equation*}
$$

The above results show that the wave front is centered at $x=\tau$ and spreads out with time. The peak of the mass-fraction perturbation at $x=\tau$ dies out like $\tau^{-\frac{1}{2}}$ and thus ultimately vanishes. If we define the thickness of the wave front as $\delta \equiv\left(\frac{u_{p}}{a_{f o}}\right) /\left|\partial v^{\prime}, \partial \bar{x}\right|_{\text {max }}$, we obtain for large times

$$
\begin{equation*}
\frac{a_{\mathrm{fo}_{o}} \delta}{\tilde{v}_{\mathrm{o}}}=\frac{\sqrt{\pi \tau}}{\mathrm{b}} \tag{29}
\end{equation*}
$$

Thus the wave front in the linear theory spreads out like the squareroot of time. It is known, however, that a constant piston speed should generate a steady-state shock front. The continual spreading of the wave front and the dying out of the mass-fraction perturbation are incorrect behaviors that arise out of the omission of the convective nonlinearities.

## 4. NONLINEAR INTERACTION

In order to delineate the balance between the dissipative transport terms and the nonlinear terms, let us first renormalize the variables appearing in (24). We introduce a characteristic length $L$ and define nondimensional time and distance as

$$
\begin{equation*}
\tilde{\tau} \equiv a_{f} t / L=\varepsilon \tau \quad, \quad \overrightarrow{\tilde{r}} \equiv \overrightarrow{\bar{r}} / L=\varepsilon \vec{r} \tag{30}
\end{equation*}
$$

where the parameter $\varepsilon \equiv \tilde{v}_{0} / a_{f_{0}} L$ is to be regarded as small. We note that $\nabla=s \tilde{\nabla}$. In the new variables, nth-order texms are proportional to $\varepsilon^{n}$, and we can rewrite (24) and display the lowest two orders as

$$
\begin{equation*}
\left|\tilde{\nabla}^{2} \phi-\phi_{\tilde{\tau} \tilde{\tau}}\right|_{\tilde{\tau} \tilde{\tau}}=\frac{\varepsilon}{P_{r} S_{c}} \tilde{\nabla}^{2}\left[\beta_{2} \tilde{\nabla}^{2} \phi-\beta_{1} \phi_{\tilde{\tau} \tilde{\tau}}\right]_{\tilde{\tau}}+o\left(\varepsilon^{2}\right) \tag{31}
\end{equation*}
$$

Thus as the viscous (transport) effects go to zero, the classical wave operator prevails, and the first-order correction for small $\varepsilon$ arises because of the next higher-order derivatives. The terms of order $\varepsilon^{2}$ in (31) correspond to the seventh- and six-order derivatives in (24). Equation (31) does not account for the nonlinearities in the problem. We account for the lower-order nonlinearities by replacing the perturbation quantities in (13) by $p^{\prime}=\varepsilon \tilde{p}$, and so on for the other primed variables, and utilizing the normalizations (30). The full nonlinear equations then lead to the following equation for the perturbation potential:

$$
\begin{align*}
& \left|\tilde{\nabla}^{2} \phi-\phi_{\tilde{\tau} \tilde{\tau}}\right|_{\tilde{\tau} \tilde{\tau}}=\varepsilon\left\{\frac { \tilde { v } ^ { 2 } } { P _ { r } S _ { c } } \left[\beta_{2} \tilde{\nabla}^{2} \phi-\left.\beta_{1} \phi_{\tilde{\tau} \tilde{\tau}}\right|_{\tilde{\tau}}\right.\right. \\
& \left.+\left|(\gamma-1) \phi_{\tilde{\tau}} \tilde{\nabla}^{2} \phi+\frac{\partial}{\partial \tilde{\tau}}(\tilde{\nabla} \phi)^{2}\right|_{\tilde{\tau} \tilde{\tau}}\right\}+0\left(\varepsilon^{2}\right) \tag{32}
\end{align*}
$$

In Equation (32) the second-degree nonlinear terms appear to order $\varepsilon$, as shown, and the third and higher degree nonlinearities appear to order $\varepsilon^{2}$. The nonlinear terms of order $\varepsilon$ in (32) are the same as occur in inviscid potential theory. Thus, the first-order corrections to inviscid acoustics are the sum of the first-order linear viscous correction and the firstorder nonlinear inviscid correction, which of course might have been anticipated a priori. Our goal now is to examine Equation (32) with the terms of order $\varepsilon^{2}$ ignored.

## 5. PISTON PROBLEM

The boundary condition for the one-dimensional impulsive motion of a piston is $v^{\prime}(\tilde{x}=0, \tilde{\tau})=\left(U_{p} / a_{f o}\right) f(\tilde{\tau})$, where $U_{p}$ is the sharacteristic speed of the piston and $f(\tilde{\tau})$ is of order unity. The characteristic length is taken to be $L=\tilde{v}_{0} / U_{p}$, and it follows from (30) that $\tilde{\tau}=U_{p} a_{f} t / \tilde{\nu}_{o}, \tilde{x}=U_{p} \bar{x} / \tilde{\nu}_{0}$, and $\varepsilon=U_{p} / a_{f}$. The small parameter $\varepsilon$ is thus the frozen Mach number of the piston, and we note that $v^{\prime}=\varepsilon_{\tilde{x}}$.

A series solution of Equation (32) by a straight-forward expansion in powers of $\varepsilon$ will lead to secular behavior such that the first-order correction becomes as large as the zeroth-order term when $\tilde{x}=0\left(s^{-1}\right)$. We account for this behavior by introducing new variables defined as

$$
\begin{equation*}
\xi=\tilde{x}-\tilde{\tau} \quad, \quad x=\varepsilon \tilde{x} \tag{33}
\end{equation*}
$$

We further expand $\phi$ in a series of the form

$$
\begin{equation*}
\phi(\tilde{x}, \tilde{\tau} ; \varepsilon)=\phi_{0}(\xi, x)+\varepsilon \phi_{1}(\xi, x)+\ldots . \tag{34}
\end{equation*}
$$

Expressions (33) and (34) applied to Equation (32) lead to the following equation for $\phi_{O}(\xi, x)$, which has been integrated twice with respect to $\xi$ and the functions of integration set equal to zero:

$$
\begin{equation*}
\phi_{o_{X \xi}}+\frac{\gamma+1}{2} \phi_{o_{\xi}} \phi_{O_{\xi \xi}}=\frac{\beta_{1}-\beta_{2}}{2 P_{r^{\prime}}{ }_{c}} \phi_{O_{\xi \xi \xi}} \tag{35}
\end{equation*}
$$

We now note that $U_{0} \equiv \phi_{O_{\xi}}$ is the lowest order velocity contribution, that is,

$$
\begin{equation*}
\mathbf{v}^{\prime}=\varepsilon U_{0}+o\left(\varepsilon^{2}\right) \tag{36}
\end{equation*}
$$

If we make the further definitions

$$
\begin{gather*}
\eta \equiv \frac{\gamma+1}{2} x=\frac{\gamma+1}{2} \varepsilon \tilde{x} \\
\nu^{*} \equiv \frac{\beta_{1}-\beta_{2}}{(\gamma+1) P_{r} S_{c}} \tag{37}
\end{gather*}
$$

then Equation (35) can be expressed in the standard form for Burgers' equation:

$$
\begin{equation*}
\mathrm{U}_{\mathrm{o}_{\eta}}+\mathrm{u}_{\mathrm{o} \mathrm{U}_{\mathrm{o}_{\xi}}=\mathrm{v}^{*} \mathrm{U}_{\mathrm{o}_{\xi \xi}} .} \tag{38}
\end{equation*}
$$

The nonlinear terms of order $\varepsilon$ in (32) lead to the convective nonlinear term $\mathrm{U}_{\mathrm{O}} \mathrm{U}_{\mathrm{O}_{\xi}}$ in (38), and the dissipative terms of order $\varepsilon$ in (32) lead to the diffusion term $U_{O_{\xi \xi}}$ in (38). The origin of the diffusivity factor $\nu^{*}$ can be seen by identifying the factors $\beta_{1}$ and $\beta_{2}$ in the original linear Equation (24). To lowest order, the other perturbation variables are found to be

$$
\begin{align*}
& \mathrm{p}^{\prime}=\varepsilon \gamma U_{0}+o\left(\varepsilon^{2}\right) \\
& \rho^{\prime}=\varepsilon U_{0}+0\left(\varepsilon^{2}\right)  \tag{39}\\
& T^{\prime}=\varepsilon(\gamma-1) U_{o}+0\left(\varepsilon^{2}\right) \\
& c^{\prime}=-\frac{\varepsilon^{2}}{S_{c}}\left|\gamma{ }_{p_{0}}+(\gamma-1) k_{T_{o}}\right|_{U_{\xi}}+o\left(\varepsilon^{3}\right)
\end{align*}
$$

### 5.1 Impulsive Compression

For impulsive motion of the piston into the fluid, the function $U_{0}(\xi, \eta)$ must satisfy the boundary condition $U_{0}(\xi, 0)=H(-\xi)$, where $H(\xi)$ is the Heaviside unit step function. The solution to (38) in this case is derived in Chapter 4 of Whitham (1974) and is

$$
\begin{equation*}
U_{0}\left(\xi^{\star}, \eta^{\star}\right)=\frac{\operatorname{erfc}\left\{\frac{\xi^{\star}-\eta^{\star}}{\sqrt{4 \eta^{\star}}}\right\}}{\operatorname{erfc}\left\{\frac{\xi^{\star}-\eta^{\star}}{\sqrt{4 \eta^{\star}}}\right\}+\operatorname{erfc}\left\{\frac{-\xi^{\star}}{\sqrt{4 \eta^{\star}}}\right\} \exp \left\{\frac{1}{2}\left(\xi^{\star}-\frac{1}{2} \eta^{\star}\right)\right\}} \tag{40}
\end{equation*}
$$

where $\xi^{\star}=\xi / \nu^{\star}$ and $\eta^{*}=\eta / \nu^{*}$. The binary-mixture contribution appears in the parameter $\nu^{*}$ and hence amounts to a stretching of the coordinates. The derivative $U_{o^{\star}}\left[\begin{array}{l}\text { is given by } \\ \left.U_{O_{\xi^{\star}}}\left(\xi^{\star}, \eta^{\star}\right)=--\frac{1}{\sqrt{\pi \eta^{\star}}} \mathrm{e}^{-\left(\xi^{\star}-\eta^{\star}\right)^{2} / 4 \eta^{\star}}+\frac{1}{2} \mathrm{U}_{0} \operatorname{erfc}\left\{\frac{-\xi^{\star}}{\sqrt{4 \eta^{\star}}}\right\} \mathrm{e}^{\frac{1}{2}\left(\xi^{\star}-\frac{1}{2} \eta^{\star}\right)}\right] \\ \operatorname{erfc}\left\{\frac{\xi^{\star}-\eta^{\star}}{\sqrt{4 \eta^{\star}}}\right\}+\operatorname{erfc}\left\{\frac{-\xi^{\star}}{\sqrt{4 \eta^{\star}}}\right\} \exp \left\{\frac{1}{2}\left(\xi^{\star}-\frac{1}{2} \eta^{\star}\right)\right\}\end{array}\right.$

The function $U_{0}\left(\xi,{ }^{\star}, \eta^{\star}\right)=U_{0}\left(x^{\star}-\tau^{\star}, \varepsilon^{\star} x^{*}\right)$ is plotted in Figure 1 as a function of $x^{*} \equiv \tilde{x} / V^{*}$ for various times $\tau^{*} \equiv \tilde{\tau} / V^{*}$ and for $\varepsilon^{\star} \equiv(\gamma+1) \varepsilon / 2=0.1$, where $\eta^{\star} \equiv \varepsilon^{*} x^{*}$. In the early stages, evolution of the wave front is dominated by diffusion and hence flattens until a balance is achieved by the steepening effects of nonlinear convection. The $\xi$-derivative of the velocity, $U_{O_{\xi^{\star}}}$, which is a measure of the slope of the wave front for small $\varepsilon$, is snown in Figure 2. This figure indicates how the maximum slope decreases and the wave front broadens as time increases. The evolution is dramatic for $\tau^{*} \leq 14$ shown in Figures 1 and 2, but for larger times changes occur more slowly. The maximum value of $-{ }^{[J}{ }_{O_{\xi, \star}}$ approaches 0.125 as $\tau^{\star} \rightarrow \infty$, but at $\tau^{\star}=140$ the maximum value of
$-U_{O_{\xi *}}$ has decreased to only 0.141 , for $\varepsilon^{\star}=0.1$. Ultimately, diffusion and nonlinear convection balance and a steady-state shape is attained with $-U_{O_{\xi}}$ having a symmetric shape about a maximum value of $1 / 8$.

The function ${ }^{[ }{ }_{O_{\xi,}}$ also describes the mass-fraction perturbation, $c^{\prime}$, as seen from expressions (39). For the heavier species $\left[\gamma_{k_{0}}+(\gamma-1) k_{T_{0}}\right]$ is negative, and hence $c^{\prime}$ is negative in the wave front. Thus, the mass fraction of the heavy species decreases in the wave front, and the mass fraction of the light species increases. The maximum value of these species perturbations is skewed toward the piston side of the wave front. As time increases, the distribution broadens and ultimately becomes symmetric, achieving a steady-state form.
5.2 Thickness of Steady-State Wave Front

As $\tau^{*} \rightarrow \infty$, the wave front evolves into a steady-state form. The
thickness $\delta$ of the wave front can be defined as

$$
\begin{equation*}
\delta=\frac{1}{\left|\frac{\partial U_{0}}{\partial \bar{x}}\right|_{\max }} \tag{42}
\end{equation*}
$$

For $\tau^{*} \rightarrow \infty$ and $\varepsilon \rightarrow 0$, we obtain

$$
\begin{gather*}
\frac{U_{p} \delta}{\tilde{\nu}_{o}}=8 v^{*}=\frac{8\left(\beta_{1}-\beta_{2}\right)}{(\gamma+1) P_{r} S_{c}} \\
=\frac{8}{\gamma+1}\left[\frac{\gamma-J+P_{r}}{{ }^{\prime} P_{r}}+\frac{\gamma N_{o}}{S_{c} k_{p}}\left(k_{p_{0}}+\frac{\gamma-1}{\gamma} k_{T_{o}}\right)^{2}\right] \tag{43}
\end{gather*}
$$

Note here that $P_{r}$ and $S_{c}$ are defined in terms of the reduced viscosity $\tilde{v}_{0} \equiv\left(2 \mu_{0}+\lambda_{0}\right) / \rho_{0}$. Formula (43) is identical to previous results
obtained for weak steady shocks, such as by Sherman (1960) when rewritten in our notation, which is a partial check on the correctness of our present analysis. Note that the parameter $V^{*}$ is related to the parameter $b$ in the linear theory, Equation (28), by the relation $2(\gamma+1) b^{2} v^{\star}=1$. Thus linear theory yields the correct combination of terms arising from viscous, thermal, and mass-diffusion dissipation that contribute to the breadth of the wave front, but raised to the wrong power.

The combination $N_{0} k_{p_{0}}$ can be written as

$$
\begin{equation*}
N_{0} k_{p_{0}} \equiv \frac{m_{1}}{m_{2}}\left(1-\frac{m_{2}}{m_{1}}\right)^{2} c_{0}\left(1-c_{0}\right) \tag{44}
\end{equation*}
$$

where $m_{1}$ and $m_{2}$ are the molecular masses of the species. Since this combination is always positive, it can be seen from (43) that pressure and thermal diffusion always tend to broaden the wave front, the more so the greater the disparity in masses of the species. As pointed out $\mathrm{b}_{\mathrm{y}}$ Sherman (1960), $\mathrm{k}_{\mathrm{p}_{\mathrm{o}}}$ and $\mathrm{k}_{\mathrm{T}_{\mathrm{o}}}$ are usually of opposite signs and tend to counteract one another.

### 5.3 Oscillating Piston

For the oscillating piston problem, the boundary condition is $\mathrm{v}^{\prime}(0, \tilde{\tau})=\varepsilon \sin \tilde{\omega} \tilde{\tau}$, where $\varepsilon \equiv{U_{\mathrm{p}}} / \mathrm{a}_{\mathrm{f}}$ as before, and $\tilde{\omega}$ is the nondimensional frequency defined in terms of the physical frequency, $\omega$, by $\tilde{\omega} \equiv \tilde{\nu}_{o} \omega / a_{f}{ }_{0} \mathrm{U}_{\mathrm{p}}$. The initial condition on the Burgers' function is thus $\mathrm{U}_{\mathrm{o}}(\xi, \eta=0)=$ - sin $\tilde{\omega} \xi$. By means of Fourier series expansions, we find the solution of (38) to be (Benton and Platzmann, 1972)

$$
\begin{equation*}
U_{0}(\xi, \eta)=\frac{4 \tilde{\omega} \nu^{\star} \sum_{n=1}^{\infty} n A_{n} \exp \left\{-n^{2} \tilde{\omega}^{2} \nu^{\star} \eta\right\}_{\sin (n \tilde{\omega} \xi,)}}{1+2 \sum_{n=1}^{\infty} A_{n} \exp \left\{-n^{2} \tilde{\tilde{\omega}}^{2} \nu^{\star} \eta\right\} \cos (n \tilde{\omega} \xi,)} \tag{45}
\end{equation*}
$$

where

$$
\begin{equation*}
A_{n}\left(\tilde{\omega} V^{*}\right) \equiv(-1)^{n_{n}} I_{n}\left(\frac{1}{2 \tilde{\omega} v^{*}}\right) / I_{0}\left(\frac{1}{2 \tilde{\omega} V^{*}}\right) \tag{46}
\end{equation*}
$$

and $\tau_{n}$ is the modified Bessel function of the first kind of order $n$. The appearance of the factor $v^{*}$ delineates the hinary-mixture, or mass diffusion, $\in f$ fects.

When $\tilde{\omega}^{2} v \star \eta \gg 1$, the term $n=1$ dominates the series, and thus in the limit of ultimate decay expression (45) becomes

$$
\begin{equation*}
U_{0}(\xi, \eta) \sim 4 \tilde{\omega} \nu^{\star} A_{1}(\tilde{\omega} \nu \star) \exp \left(-\tilde{\omega}^{2} v^{\star} \eta\right) \sin \tilde{\omega} \xi \tag{47}
\end{equation*}
$$

Aside from the exponential damping, the amplitude of the disturbance in the far field is a function of the frequency of oscillation. This is a nonlinear effect and is akin somewhat to the dependency of the amplituie on frequency in the nonlinear oscillations of spring-mass systems. The amplitude function $4 \tilde{\omega} V^{\star} A_{1}\left(\tilde{\omega} v^{\star}\right)$ is plotted as a function of $\tilde{\omega} V^{*}$ in Figure 3. The function is a maximum of unity when $\tilde{u} \nu^{*} \rightarrow \infty$, and decreases as $\tilde{\omega} u^{*}$ becomes smaller. When $\tilde{\omega} v^{\star} \gg 1$, expression (47) approaches the form

$$
\begin{equation*}
u_{0}(\xi, \eta) \sim-\exp \left(-\tilde{\omega}^{2} v^{\star} \eta\right) \sin \tilde{\omega} \xi \tag{48}
\end{equation*}
$$

This is the solution that is obtained from the linear theory, described by Equation (27), in the limit $\varepsilon \tilde{\omega} \nu^{*} \ll 1$. Thus, the linear theory according to Equation (27) is valid in the final stage of decay when $\tilde{\omega} \nu^{*} \gg 1$ and $\varepsilon \ll\left(\tilde{\omega} \nu^{*}\right)^{-1}$. Except for this, special limit, nonlinear effects always make a contribution, even in the final stages of decay. Even though small, the nonlinear effects are cumulative. At high
frequencies the steepening of the compression part of the wave tends to cancel the flattening of the expansion part., whereas for low frequencies the compressions and expansions are less closely spaced and hence lead to a nonvanishing cumulative effect in the far field.

The evolution of the wave disturbances just before final decay can be examined by keeping the term $n=1$ in the denominator as well as the numerator of (45) and ig:oring all terms for $n \geq 2$. Then we have, for $\tilde{\omega}^{2} \nu \star \eta \geq 1$,

$$
\begin{equation*}
\mathrm{U}_{0}(\xi, \eta) \sim \frac{-\left(\frac{4 \tilde{\omega} v^{\star} I_{1}}{I_{0}}\right) \exp \left(-\tilde{\omega}^{2} v^{\star} \eta\right) \sin \tilde{\omega} \xi}{1-\left(\frac{2 I_{1}}{I_{0}}\right) \exp \left(-\tilde{\omega}^{2} v^{\star} \eta\right) \cos \tilde{\omega} \xi} \tag{49}
\end{equation*}
$$

The wave forms produced by expression (49) are shown in Figure 4. The curves are shown for $\tilde{\omega}^{2} v^{\star} \eta=1$ and $\tilde{\omega} v^{\star}=\infty, 0.5,0.25$. The linear sine wave is represented by $\tilde{\omega} \nu^{*}=\infty$. The values $\tilde{\omega} \nu^{*}=0.5$ and 0.25 represent the effects of the nonlinearities, which yield wave forms that are flatter in the middle and steeper on the ends. As $\eta \rightarrow \infty$, for $\tilde{\omega} v^{*}$ fixed, all the curves evolve toward the sine-wave shape.

The other perturbation variables are also given by expressions (39), the pressure, dcnoity, and temperature being proportional to the velocity. The mass-fraction perturbation, $\mathrm{c}^{\prime}$, is proportional to the $\xi$-derivative of $U_{0}$. For the behavior approaching ultimate decay, $\tilde{\omega}^{2} v^{\star} \eta \geqslant 1$, the derivative of (49) yields a valid description:

$$
\begin{equation*}
\mathrm{U}_{\mathrm{o}_{\varepsilon, \star}}(\xi, \eta) \sim-\frac{\left.\left(\frac{2 \tilde{\omega} \nu^{\star} I_{1}}{\mathrm{I}_{0}}\right) \exp \left(-\tilde{\omega}^{2} \nu^{\star} \eta\right) \right\rvert\, 2 \tilde{\omega} \nu^{\star} \cos \tilde{\omega} \xi+\mathrm{U}_{0} \sin \tilde{\omega} \xi,}{1-\frac{2 \mathrm{I}_{1}}{\mathrm{I}_{0}} \exp \left(-\tilde{\omega}^{2} v^{\star} \eta\right) \cos \tilde{\omega} \xi,} \tag{50}
\end{equation*}
$$

The behavior of expression (50), and thus of $c^{\prime}$, is shown in Figure 5 for $\tilde{\omega}^{2} v^{*} \eta=1$ and $\tilde{\omega} \nu^{*}=\infty, 0.5$, and 0.25 . The limiting value $\tilde{\omega} \nu^{*}=\infty$ yields a damped cosine wave that corresponds to the associated linear theory. The perturbation $c^{\prime}$ is thus 90 degrees out of phase with $U_{0}$. When $\tilde{\omega} \nu$ is finite, the magnitude of $U_{O_{\xi^{*}}}$, and thus $c^{\prime}$, is larger in the compression part of the wave than it is in the expansion part. This is a nonlinear effect. If $c^{\prime}$ is identified with the heavier species, such that $\left\{\gamma k_{p_{0}}+(\gamma-1) k_{T_{0}}\right\}$ is negative, then we find that $c^{\prime}$ is negative in the compression part of the wave and positive with a smaller magnitude in the expansion part of the wave. It also can be seen that more time is spent in the expansion part of the wave than in the compression part. As $\eta \rightarrow \infty$, for $\tilde{\omega} V^{*} f i x e d$, all the curves tend toward the cosine-wave shape.

## 6. CONCLUDING REMARKS

The interaction of nonlinear and dissipative mechanisms for wak disturbances in inert binary mixtures has been investigated. is for pure gases, the basic description of one-amensional problens is by means of Burgers' equation. The effects of the binary mixture are delineated $b=$ the effective ditfusion coefficient $\widetilde{\nu}^{*}$ which contains the effects of Schmidt number and pressure and thermal diffusion coefficients associated with binary mixtures as well as the prandtl number and ratio of specific heats associated also with pure gases. The binary-mixture mechanisms tend to broaden the wave fronts and to modify the amplitide-frequency relation and damping rate of oscillatory disturbances. The mass: fraction perturbation is skewed toward the piston side of revolving sho $k$ front:s, but becomes symmetrical as the steady-state is approached. for osciilatory motions, the magnitude of the mas:sfaction pertarbation is larger in the compression part of the wave and of naller macritude and opposite sign in the expansion part of the wave.

The nonzero mass-fraction perturbations suggest that pressure and ther. 1 diffusion may have significant effe:ts in chemically reacting mixtures where nonequilibrium effects play essential roles. Research towards these ends may yield significant results for the behavior of detonation waves and combustion noise.

Perturbation methods applied to hi er-order forms of t... :anliear Equatior (32) may yiold more general mox \& equatir is han 3uraers' equation. Also spherical and cylindrical disturtance $n$ the be stuated fruitfully by this approach. These are aroas for future research.


Figure 2. Evolution of Mass-Fraction Perturbation




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