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## A LINEAR PHOTODIODE ARRAY EMPLOYED IN A SHORT RANGE LASER TRIANGULATION OBSTACLE AVOIDANCE SENSOR

by

## James Paul Odenthal

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James Paul Odenthal<br>A Project Submitted to the Graduate<br>Faculty of Rensselaer Polytechnic Institute<br>in Partial Fulfillment of the<br>Requirements for the Degree of<br>MASTER OF ENGINEERING



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## ABSTRACT

An opto-electronic receiver incorporating a multi-element linear photodiode array as a component of a laser-triangulation rangefinder is described. Developed as an obstacle avoidance sensor for the Rensselaer Polytechnic Institute's Martian Roving Vehicle, the detector can resolve the angle of laser return in $1.5^{\circ}$ increments within a field of view of $30^{\circ}$ and a range of five meters. A second receiver with a 1024 elements over $60^{\circ}$ and a 3 meter range is also docuemented. Included is a discussion of design criteria, circuit operation, schematics, experimental results and calibration procedures.

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## PART 1

## INTRODUCTION

An autonomous robot vehicle has been suggested ${ }^{l}$ as a component of a scientific triad of orbiter, rover and penetrators devised to continue the investigation of the planet Mars. "The global reconnaissance carried out by the orbiters, the network science and limited characterization of otherwise inaccessible environments carried out by the penetrators, and the detailed study of local environments conducted by the rover," ${ }^{1}$ comprise the three elements of a post-Viking, pre-sample return mission.

Speaking specifically of rover operations, early mission studies ${ }^{1}$ envisioned a two-year, 100 kM survey of the Martian surface. These studies indicated that in order to survey such a significant portion of the planet's surface while stili enjoying the option of intensive investigation of selected scientifically interesting sites, the machine would be obliged to travel at relatively high speeds, in excess of 1.5 meters/min, between sites. Round trip radio communication between earthbound mission control and a Martian rover (142 million mile orbit) would involve from ten to 40 minutes. Clearly, with such a loop time, real-time human control could not achieve the desired speed. However, an autonomous vehicle, capable of independently avoiding obstacles, could proceed rapidly, maneuvering through its hazardous environment with only intermittent human guidance. Such an independent operation demands a sophisticated terrain-sensing and decision-making facility on the part of the machine.

Information about surface topography, obtained for the purpose of obstacle negotiation would, as envisaged, be gathered from a triple hierarchy of sensors; orbiters overhead, stereo cameras/laser rangers aboard the rover, and a short-range hazard detection system. At the highest level, the orbiter, equipped with a radio altimeter and high resolution imagers, would provide the operators at JPL mission control with the requisite data to identify macro-obstacles and science goals for the next 24 hours. Coupling this information with the most recent stereo images taken by the rover's cameras, a list of traverse coordinates linking the desired science goals by a best perceivable path can be generated and transmitted to the rover.

Upon receipt of the transverse link coordinates and the location of science sequences selected by JPL, the vehicle, utilizing onboard logic and hazard avoidance equipment, begins its advance across the Martian terrain. Operating in a 50 minute "halt-sense-think-travel" cycle the machine employs its mid-range ( $30-40$ meter) sensors, stereo cameras and laser rangefinder to perform the required obstacle detection and path planning during the 25 minute "halt" portion of the cycle. In the "travel" segment the rover, travelling at a rate of 1.5 meters/min, then pursues the course plotted in the mid-range analysis, depending solely on its short-range hazard detection system. Functioning in the 1-5 meter range, the short-range system is responsible for rapidly recognizing and evading those obstacles ( $\pm 25 \mathrm{~cm}$. step, $\pm 30 \mathrm{deg}$. slopes) which may have eluded the mid-range examination. It is the short-range problem; terrain sensing, obstacle identification and path selection,


RENSSELAER AUTONOMOUS ROVING VEHICLE

$$
\begin{aligned}
& \mathrm{OR}_{\mathrm{G}_{\mathrm{A}}} \\
& \mathrm{OE}_{P O O R^{2}} \mathrm{p}_{A_{\mathrm{GE}}} \text { IS } \\
&
\end{aligned}
$$



ELEVATION SCANNING MAST - FRONT VIEW

which has been investigated at RPI in recent years.
To this end a four-wheeled, remotely-controlled vehicle has been constructed, complete with terrain sensors, control electronics, rough terrain capability and radio links to an off-board computer, to serve as a test bed and demonstration tool for hardware and algorithms, Fig. 1. Research has been conducted in the areas of packaging/deployment, vehicle stability and performance, terrain modelling, navigation and path selection as well as obstacle detection employing the laser triangulation ranger described herein.

Evaluation of the robot's performance carried out in the laboratory and at an outdoor test site involves placing the rover in an environment consisting of many arbitrary obstacles (rocks, walls, craters, etc.). The machine, equipped with a directional gyro, tachometers and terrain sensor is charged with the task of negotiating the obstacle field under autonomous control to an operator-designated goal (specified as an $x-y$ target or vector heading).

The terrain sensor, located in the front of the vehicle and embodied in the laser/detector mast, Fig. 2, gathers direction and range information from terrain points in the vehicle's vicinjty. Employing the principle of triangulation in the Laser Direction and Ranging system (ladar) scans about the vehicle firing the laser at predetermined angles, Fig. 3. At each firing the light energy is reflected from the terrain back to an array of photodiodes which resolve the angle of return. The information from this process, when sampled at a suitable spatial frequency and coupled with the vehicle roll and pitch state,


> B -ANGLE OF LASER RAY
> $\propto$-ANGLE OF LASER RETURN

$$
Z=\frac{H_{\beta} \tan \beta-H_{\alpha} \tan \alpha}{\tan \beta-\tan \alpha}
$$

$$
\mathrm{R}=\frac{\tan \alpha \tan \beta\left(\mathrm{H}_{\beta}-\mathrm{H}_{\alpha}\right)}{\tan \beta-\tan \alpha}
$$

FIGURE 3 RELATIONSHIP BETWEEN $(\alpha, \beta),(Z, R)$
enables an off-board computer to perform the necessary obstacle detection and path planning.

Aboard the RPI Mars Roving Vehicle (MRV), the short-range terrain sensing has been accomplished through application of a pulsed laser/photodiode triangulation technique, but not without first considering other alternatives. Radar, for instance, used extensively in direction and range finding, offers itself as a well developed technology but suffers from a lack of resolution in the short range and large physical size of its antennae. Sonar, with shorter wavelengths, has the needed range accuracy, but can be dismissed because of the tenuous Martian atmosphere. Mechanical feelers, although suitable as the last line of defense in the form of bumpers, cannot achieve both the range needed for rapid movement and ruggedness for long service. A laser rangefinder measuring the time of flight of terrain echoes can, by appropriate averaging of many returns, estimate the range of terrain points to an accuracy of several centimeters through the use of sophisticated picosecond circuitry. Yet, though such precision may be acceptable for mid-range slope and obstacle estimation, the need for finer discrimination and higher speed in the short-range case points toward the pulsed laser triangulation scheme.

The sensitivity of a triangulation scheme increases rapidly with shorter distance and computer simulations ${ }^{2}$ have demonstrated that for a laser-to-receiver distance of one meter, and an angular resolution of one degree in both laser and detector, step obstacles of 25 cm and slopes exceeding $\pm 30$ degrees can be reliably located in the one-three


FIGURE 4 DISCRETE DETECTORS ABOARD PROPOSED 1984 ROVER
meter range. Also the triangulation system, by avoiding the averaging demanded in the time-of-flight ranger and unencumbered by the computational load of the stereo correlation approach, permits the rapid sampling over the wide field of view imposed by the short-range problem. Furthermore, because of the active nature of the laser ranger, it can be utilized at night without the artificial illumination requirement of the passive stereo camera sensor, a feature which increases available traveling time while minimizing power expenditure. Another advantage of the triangulation scheme is that laser sources and photo-receptive receivers can be easily proliferated around the machine fixed in appropriate directions, thereby eliminating the need for mechanical scanning, Fig. 4. For these reasons and others, including low cost, silicon reliability and conceptual ease, the terrain sensors in service aboard the RPI MRV have been of the pulsed laser triangulation variety. The first terrain sensor employed aboard the RPI MRV was the single-laser/single-detector system depicted in Figs. 1, 5. The laser and detecțor were both fixed in elevation aboard a vertical mast which oscillated through $140^{\circ}$ of azinuth. In principle only laser reflections within the detector cone of vision were considered as evidence of passible terrain. If in any azimuth a missing retum occurred, an entry was made into the path selection terrain map indicating the presence of a hazard in that azimuth at a range of 1.5 meters.

By adjusting the field of view of the detector, by means of an iris in the photo-sensitive image plane, step obstacles of 25 cm , positive and negative, as well as slopes exceeding $\pm 10^{\circ}$, could be


BY CONTROLLING DETECTOR FIELD OF VIEW STEP OBSTACLES OF $\pm 25 \mathrm{CM}$. (WHEEL RADIUS ) CAN BE PERCEIVED
however this restricts slope capability to $\pm 10$ deg.


FIGURE 5 SINGLE LASER -SINGLE DETECTOR SYSTEM
distinguished. Yet the vehicle was designed to climb $\pm 30^{\circ}$ slopes. Furthermore since there was no indication of range, the range was estimated conservatively at 1.5 meters thereby foregoing a good deal of maneuvering room, particularly in the case of negative obstacles. The inability of the single-laser/single-detector scheme to detect both step obstacles and $\pm 30^{\circ}$ slopes as well as its range ambiguity restricted its realm of operation to relatively level terrain $\left( \pm 10^{\circ}\right.$ roll and pitch) populated by discrete obstacles several vehicle widths apart.

Although the single-laser/single-detector system was limited in terms of scene interpretation by its geometry it did demonstrate the feasibility of the hardware, namely that a pulsed (10 watt, 200 ns ) laser and a silicon photodiode could function reliably over the wide range of reflectivities, ranges and ambient light conditions encountered in its environment. Given the confirmed viability of the laser/photodiode combination it remained to improve the perceptual performance of the terrain sensor to include the capability for detection of step obstacles embedded in $\pm 30^{\circ}$ hillsides.

The multi-laser/multi-detector's concept was to increase the number of laser and detector rays so that the density of their intersections would be sufficient to distinguish a 25 cm step obstacle anywhere within the volume defined by worst-case slope constraints. Figure 6 illustrates how the $\pm 30^{\circ}$ slope limit of the rover defines the volume within which passable terrain must be sampled $\left( \pm 60^{\circ}\right.$ with respect to vehicle) and the distribution of the source/sensor intersections

-60 DEG. w.r.t. VEHICLE

+60 DEG W.R.T. VEHICLE


FIGURE 6 VEHICLE PITCH, SENSOR FIELD OF VIEW, RESOLUTION
generated in the elevation scanning scheme. The field of view of laser and detectors required to encompass this volume can be estimated from this illustration to be approximately $45^{\circ}$ and $60^{\circ}$ respectively. The angular resolution of the laser and detector can be estimated from the elevation cross-section of the 25 cm obstacle at the maximum range of four meters. Simply,

```
tan}\Delta\beta=25\textrm{cm}/4.0\mathrm{ meters }->\Delta\beta~4
```

To sample at twice this frequency with an extra $100 \%$ margin of error, a resolution of $1-2$ degrees was set as a goal for both laser and detector. Subsequent computer simulations ${ }^{2}$ using $1^{\circ}$ resolution and $40^{\circ}$ field of view for both laser and detector have since confirmed the validity of this choice.

The hardware needed to scan the laser/detector through azimuth and elevation was designed ${ }^{3,4,5}$ in the $1977-78$ academic year and is shown in Figs. 2 and 7. The system allows 1024 laser shots fired in a $32 \times 32$ pattern with a resolution of $1.05^{\circ}$ in a $90^{\circ}$ elevation field and $2.8^{\circ}$ resolution in a $360^{\circ}$ azimuth scan.

The two primary components of the mast controller are the elevation and azimuth encoders, Fig. 8. As the eight-sided mirror rotates the elevation encoder outputs 2048 pulses per revolution (every $0.18^{\circ}$ ). Likewise the azimuth encoder produces 256 pulses (every $1.4^{\circ}$ ) in its revolution. In the controller, the counts from these two encoders are compared to a firing pattern stored in the azimuth and elevation


FIGURE 7 LASER DIRECTION AND RANGING MAST


PROM (Programmable Read Only Memory). When a match occurs the laser is fired and the detector output is loaded into the telemetry FIFO First-In-Eirst-Out buffer) along with two five-bit words, $\beta \#$ and $\theta \#$, which identify at which elevation and azimuth angle the laser was fired. This process can be repeated at a 10 kHz rate allowing a complete azimuth scan to be completed in as little as 1.65 seconds. With the laser scanner designed it remained to select a detector compatible with the elevation scanning scheme.

## DETECTOR SELECTION

### 2.1 Detector Criteria

In any triangulation scheme the location of an unknown point is determined by measurement of angles from two known points. In the laser ranger used aboard the RPI Martian Rover the unknown points are the laser spots on the terrain and the known points are the laser atop the rotating mast and the detector situated below. The angle from the laser, $\beta$, is measured by the mirror shaft encoder and constrains the terrain point to lie along the indicated laser ray, Fig. 3. The function of the detector is to measure $\alpha$, the elevation angle of the intersecting image ray. Three schemes for resolving $\alpha$ were considered: discrete receivers, linear arrays and two-dimensional arrays. Below is a discussion of the considerations involved in the detector selection.

Practical factors such as cost, size, weight and safety being equal, four criteria pertinent to this analysis have been isolated: resolution, speed, scope and sensitivity.

The resolution of the receiver is its ability to discriminate differences in the position of the laser reflection upon the ground. Measured in terms of angular accuracy, it is obvious that an increased number of photo-receptive elements within a given angle serves to enhance the resolution of the detector. A large number of detectors also affords the researcher the flexibility necessary to examine a wide variety of possible detector geometries. Computer simulations ${ }^{2}$ have. indicated that a detector resolution of $1^{\circ}-2^{\circ}$ is adequate to perceive
the 25 cm threshold obstacle.
Also of prime concern in detector selection is the speed at which it can function. Since the vehicle travels in a continuous fashion, it is imperative that the laser/detector system operate at a rate sufficient to ensure that any terrain has been scanned several times before the machine rolls over it. Specifically, if the vehicle velocity is $1 \mathrm{kM} / \mathrm{hr}$ (approximately $\frac{1}{4}$ meters/sec), the sensor's range 4 meters, and the time between successive scans of the terrain is 4 sec , then the terrain passing beneath the vehicle will have been sampled four times. As a further example, at the mission study speed of 1.5 meters/min, to scan the terrain ten times, a detector with a range of three meters has as long as 12 seconds to complete its azimuth rotation. In general,

$$
\frac{\text { Sensor Range }}{\text { - Vehicle Velocity }}=\# \text { times terrain scanned. }
$$

Given the rate at which the terrain must be scanned, the relationship between the mast/mirror rotation and the desired angular resolution of the laser determine the detector's operating speed. Assuming an n-sided mirror spinning continuously through $360^{\circ}$ in both azimuth and elevation, the time between laser shots and scan time can be related, Fig. 9.

$$
\begin{aligned}
\omega_{\text {mast }} & =2 \pi / \Delta t_{\text {scan }} \\
\Delta \theta & =\omega_{\text {mast }}\left|\frac{2 \pi}{N} / \omega_{\text {mirror }}\right|
\end{aligned}
$$



FIGURE 9 DETECTOR SPEED, SCAN TIME; andANGULAR RESOLUTION

$$
\Delta \beta=\Delta t_{\operatorname{det}} \omega_{B}=\Delta t_{\operatorname{det}} 2 \omega_{\text {mirror }}
$$


where
$t_{\text {det }}=\begin{aligned} & \text { time allotted for detector operation, also time between } \\ & \text { closest consecutive laser firings }\end{aligned}$
$\Delta \beta=$ minimum elevation angle between successive laser shots (rad)
$\Delta \theta=$ azimuth angle within which elevation sweep occurs (rad)
$N=$ number of faces on polygonal mirror
$\Delta t_{\text {scan }}=$ time between successive scans of terrain
For example, the 4 sec . scan, with a laser resolution of $1^{\circ}$ in elevation, an azimuth error of $\pm 1^{\circ}\left(\Delta \theta=2^{\circ}\right)$, and using an eight-sided mirror, constrains the receiver to $246 \mu s e c(\sim 4 \mathrm{kHz}$ ) or faster. The dependency of $\Delta t_{\text {det }}, \Delta \beta$ and vehicle velocity illustrates the trade-offs between spatial and temporal data density inherent in the rotating mast-mirror scheme. If it were possible to scan the laser at electronic frequencies, instead of employing rotating mechanical components, then the tradeoffs would be remarkably improved. Investigation into other means of laser deflection (such as acousto-optical crystals (Glen Herb $\left.{ }^{\prime} 73\right)^{5}$ ) showed that the combination of large deflection ( $50^{\circ}$ ) and high speeds was not feasible.

Another criterion applying to the detector is its scope.
What is meant here is the volume, defined by the detector's field of view and range, within which the laser spot on the terrain can be seen.

Terrain which falls outside this envelope cannot be measured and must be assumed to be hazardous. Ideally the scope or envelope would encompass all passible terrain, slopes of $\pm 60^{\circ}$ and a range long enough to perceive hazards outside the vehicle's turning radius ( $\sim 2$ meters). This implies that the detector's scope should be $60^{\circ}$ in elevation, $180^{\circ}$ in azimuth and in excess of 3 meters in range. The detector must respond reliably to any laser spot which falls within its volume of perception. The detector must be sensitive to the wavelength of the laser ( 904 nm. ) as well as the low light level of the terrain reflection. The amount of energy that can be expected to retum can be calculated by considering the amount of energy released by the laser, account for path losses and then arrive at the energy impinging on the detector.

The laser (Meshach '78) generates 100 watts for 40 ns. The collimating lens transmittance and mirror reflectance are assumed to be $85 \%$ and $90 \%$ respectively. Ground reflectance is assumed to be as low as $10 \%$. Assuming further that the energy reflected from the ground radiates in all directions equally (a reasonable assumption considering the diffuse nature of the terrain the rover should encounter), the ratio of the energy entering the detector to that radiated away is:

Area Lens Aperture
Area of Hemisphere with Radius $=$ Range

If the transmission of the detector lens and filter are included

```
Energy \(=\) (100 watts) (40ns) (lens trans) (mirror reflect)
    def
    (ground reflect) \(\left(\frac{r^{2}}{2 R^{2}}\right.\) ) (filter trans) (lens trans)
    (100 watts) ( 40 ns ) (.85) (.90) (.10) \(\frac{1}{2}\left(\frac{\mathrm{r}}{\mathrm{R}}\right)^{2} .8\) (.85)
\(=10^{-7}\left(\frac{x}{R}\right)^{2}\) joules
```

Assuming the ratio of the lens radius (on the order of millimeters) and the range (1-5 meters) can be approxinated by $1 / 1000$, then the amount of energy passing through the detector's lens could be presumed to be
0.1 picojoules or 2.5 山watts peak

Now that the criteria for the detector have been enumerated it is possible to evaluate the three proposed receiver candidates, i.e., discrete arrays, linear arrays and area arrays.

### 2.2 Discrete

A discrete receiver approach calls for $n$ individual detectors, each with its own lens, photosensor and amplifier. Each detector would have both variable attitude and field of view. A single detector of this sort is presently functioning reliably aboard the RPI MRV.

The resolution of this system is limited by the number of receivers that can be mounted aboard the rotating mast. Since only 25 to 30 discrete detectors could realistically be mounted, the potential resolution is limited. Although the detector cones could be overlapped to increase the effective resolution potential errors due to vibration and misalignment appear likely.

The speed of the discrete detector's alternative is the


FIGURE 10. DISCRETE DETECTOR SCHEME
fastest of any of the three approaches. Since each detector measures returning power and has individual amplifiers, its speed, constrained only by the rise and fall time of a single detector, can approach the megahertz range.

The scope of the discrete detector scheme is the most versatile. Since each detector can be pointed in any desired direction any conceivable detector geometry can be generated. Range measurements taken under laboratory conditions show that this laser/detector combination has a reliable range in excess of five meters.

Sensitivity of the discrete approach is confirmed by the successful operation in rough terrain environment of the current single-laser/single-detector system. It has operated reliably throughout its desired range and in bright sunlight.

With its high speed, flexible scope and proven sensitivity a discrete detector approach could have readily been achieved by replicating the single-laser/single-detector receiver. Yet this system's disadvantages, including low resolution, large size and very difficult alignment problems, did not seem to make it the most attractive proposal.

### 2.3 Area Arrays

Two-dimensional detectors, as opposed to discrete and linear arrays, have also been suggested for use as receivers. Devices such as vidicons and solid state matrix arrays offer very high resolution with typically 100 elements per dimension, and the sensitivity, particularly of silicon-intensified-vidicons, is quite good. Yet for most
area arrays, excluding the CID random-access imagers, the speed of area arrays is limited by the need to sample the entire frame at rates near $30-60 \mathrm{~Hz}$. This disadvantage could be circumvented if the laser were scanned rapidly in a direction perpendicular to the line between the laser and receiver. Specifically if the laser is located above the receiver it should be scanned rapidly in azimuth while the elevation angle is held constant, Fig. 11. Then the image of the laser would be a single-valued function of the horizontal dimension of the array. A striking advantage of this scheme is that only the mirror must be rotated, not the receiver, obviating the need for slip-rings. Furthermore the measurement of the azimuth angle is derfved directly from the horizontal position in the array, thereby eliminating the need of an azimuth encoder. Yet although the resolution is high, the speed not a strong disadvantage and the mechanical scanning greatly simplified, matrix arrays not only have a limited fleld of view in elevation but in azimuth as well. To achieve a $180^{\circ}$ azimuth field of view in front of the vehicle might require three or four cameras. Each camera would have to be aligned separately and devoted exclusively to short-range bazard detection. Although it is a simple matter to proliferate cameras aboard the vehicle, and it may be advantageous to have redundant cameras, the costs of the cameras exceed the project budget and it was felt that a single dedicated unit contained on the mast fits more with the spirit of the proposal for a short-range hazard detection system.


FIGURE 11. 2-DIMENSIONAL AREA ARRAY IMPLEMENTATION

### 2.4 Linear Arrays

A detector incorporating a linear array of precisely manufactured photo-diodes behind a single lens avoids the severe alignment problems of $n$ discrete receivers. With a single lens, only the pointing angle of the camera and the distance from lens to array need be adjusted. Fig. 12

The resolution of the linear array is a function of the number of elements along its length, and the field of view chosen. Linear arrays are available with from 16 to 1728 elements, which with a field of view of $60^{\circ}$ gives a resolution of $3.75^{\circ}$ and 2 min 5 sec respectively.

The speed of the linear array receiver depends strongly on the number of elements involved. Arrays with 20 or less elements can provide separate outputs for each of their photodiodes in a single dual-in-line-package. These arrays are as fast as diode receivers since each photosensor is used in parallel with its own amplifier. When more than 20 elements are involved it becomes difficult to offer a separate output for each diode and some form of on-chip scanning or multiplexing is required. This either takes the form of analog multiplexors (C.I.D.s) or of analog shift registers as in the case of $C C D$ imagers. In both cases the time for array operation increases linearly with the number of elements. Typically the maximum scanning frequency is from 1 to 5 Mhz so that for 1024 elements the detector time is from 1 wsec to 200 usec. Using the expression relating detector time, $\Delta t_{\text {det }}$, to the time required for a scan, $\Delta t_{\text {scan }}$ :


FIGURE 12. LINEAR ARRAY IMPLEMENTATION

$$
\Delta t_{\text {scan }}=\frac{\pi^{2}}{\Delta \beta \Delta \theta} t_{\text {det }} \quad \begin{aligned}
& \Delta \beta=\text { resolution in } \beta-1^{0} \\
& \Delta \theta=\text { error in azimuth } \sim 2^{\circ}
\end{aligned}
$$

For $1 \mu s e c$ and 204 usec detector time, the 1024 element IInear array limits the scan time to greater than 16.5 or 3.3 sec respectively.

The scope of the linear array is a function of the Iens' focal length and is a trade-ofi between field of view and resolution. For a given number of elements, say 20 , the requirement for at least $1.5^{\circ}$ resolution limits the receiver's field of view to $30^{\circ}$. A linear array with 1024 elements enjoys such a high resolution that even at the maximum desired field of view of $60^{\circ}$ there is only 3 min 30 sec of arc between diodes.

The sensitivity of the linear array receiver also varies with the number of elements. For a 20 element array with parallel outputs the sensitivity is similar to that of the discrete approach with the important advantage of the single lens. In the discrete detector system each receiver has its own lens which, for 20-30 detectors must be relatively small. However in the linear array receiver each photodiode receives an image from the entire large lens providing substantial gain and thus reducing the amplification needed. The sensitivity of self-scanned linear arrays is more complicated to analyze. Although the longer arrays also benefit from the single lens they suffer from disadvantages peculiar to the scanning mechanism which permits their long length. First of all the switching noise and
capacitance of the multiplexors degrade the signal-to-noise ratio. Secondly, since the photodiodes must hold their photo charge until. scanned by the multiplexers, they are susceptible to leakage and dark currents. To hold their photo current the photodiodes must integrate the light intensity falling on them, thereby measuring the energy falling on them between samples. Yet the energy released by the laser ( 100 watts, $40 \mathrm{~ns}=4 \mu$ joules) when averaged over $t_{\text {det }}(204 \mathrm{sec})$ gives an effective power of $\frac{4 \text { 背 }}{204 \mu \mathrm{sec}}=20 \mathrm{mw}$. Hence the advantage of a high power ( 100 watt) laser exploited by the discrete and short linear schemes is dissipated by the integrating nature of the long self-scanned arrays.

In sumary, the three candidates for the multi-detector system, discrete, area and linear methoris, were compared in terms of their resolution, speed, scope and sensitivity. The discrete receivers, while having the highest speed, most versatile scope and proven sensitivity, were hampered by relatively low resolution and very difficult aligment. Area arrays such as vidicons and solid state matrices obviate the need for azimuth encoders, but the cost of several cameras to cover $180^{\circ}$ in azimuth exceeds the project budget.

Hence linear arrays with the alignment simplicity and increased sensitivity granted by the single lens appeared as a compromise between the complexity of the discrete detectors and the high cost of the multiple area arrays. Two linear arrays were purchased. A 20element array with parallel outputs promised proven sensitivity, 10 kHz speed and $1.5^{\circ}$ resolution over $30^{\circ}$ field of view. In addition, due to
the prospect of very high resolution a self-scanned 1024 element array was acquired in the hope that its relatively slow speed and questionable sensitivity could be overcome.

PART 3

20 ELEMENT RECEIVER

### 3.1 General Description

The 20-element detector system is divided into two main packages, Fig. 14. The detector head holds a $35 \mathrm{~mm} / \mathrm{f} / 2.0$ lens, the 20-element array, and 20 amplifiers in a $2 \frac{1}{4}$ inch diameter cylinder. The pointing angle of the detector head is adjustable by means of a worm gear ( $3^{\circ} /$ turn) through $90^{\circ}$. A 26 conductor ribbon cable connects the detector head, to the $2 \frac{1}{4} \times 3 \times 4 \frac{1}{2}$ ' signal conditioner and display unit. Contained in this second package are the threshold adjust/comparator board and the digital timing and display board, Fig. 15. The sensitivity of each of the photodiodes can be adjusted with one of the 20 threshold potentiometers; and the 20 LEDs (Iight Emitting Diodes) display which of the photodiodes have been illuminated. Digital detector output as well as power $(+24,15,5,-15)$ and Laser Fire Signal pass through a 16 conductor ribbon cable between the digital board and the slip rings.

The entire system offers $1.5^{\circ}$ resolution over $30^{\circ}$ field of view and is able to operate at well over 10 kHz

### 3.2 Optics

The lens, specified in terms of its focal length and f-number, determines the field of view, range and indirectly, the resolution of the detector system. Figure 16 illustrates how the focal length of the lens and the length of the photodiode array constrain the field of view. Given that the 20 elements of the linear array cover 20 mm and

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FIGURE 13. DETECTOR ON MAST



FIGURE 15. DETECTOR FUNCTIONAL BLOCK DIAGRAM
that the field of view of the receiver is desired to be $30^{\circ}$ one can calculate a suitable focal length

Detector lens focal length $=\frac{20 \mathrm{~mm}}{2} / \tan 15^{\circ}=37.32 \mathrm{~mm}$
35 mm is a readily available standard photographic focal length which allows a field of view of
$\tan \left[\frac{f . O_{0} v_{0}}{2}\right]=10 \mathrm{~mm} / 35 \mathrm{~mm}$

$$
\text { f.O.V. }=31.89 \text { degrees }
$$

Measurements taken with $35 \mathrm{~mm} / \mathrm{f} 2.0$ lens confirmed this analysis.

The range and sensitivity of the detector are directly related to the intensity of the laser image formed upon the photodiode array. The f-number of the lens, defined as the ratio of focal length to aperture diameter, should be chosen as small as possible to maximize the aperture to image size ratio. The lens that was purchased has a minimum $f$-number of 2.0 . At this setting the lens has an effective aperture of
f-number $=\frac{\text { focal length }}{\text { aperture }} \rightarrow$ diameter dameter $=\frac{35}{20}=17.5 \mathrm{~mm}$

For an effective area of

$$
\pi(9 \mathrm{~mm})^{2}=2.54 \mathrm{~cm}^{2}
$$

Since the lens' f-number can be varied from 2 to 16 , its light-gathering power can be varied over a range from 1 to 64 , a feature which is used to control the sensitivity of the receiver.

The effect of the focal length upon the resolution of the detector can be readily seen by dividing the 20 photodiodes into the


Lens Makers Equation : $\frac{1}{f}=\frac{1}{s}+\frac{1}{d}$

Magnification : $\frac{1}{0}=\frac{d}{s}$


FIGURE 16. GEOMETRIC OPTICS
field of view

$$
\frac{32 \text { degrees }}{20}=1.65^{\circ} \text { per element }
$$

The resolution of the detector is also dependent upon the f-number of the lens. From the "lens-makers-equation," 8 the distance of the image from the lens, $s$, can be related to the object distance, d.

$$
\frac{1}{p}=\frac{1}{s}+\frac{1}{d} \text { lens-makers-equation }
$$

This equation shows that a laser spot from 1 meter will be focused at 36.27 mm and a 5 meter laser return image at 35.25 mm . If the array is placed halfway between these extremes, at 35.76 mm , so that 1.6 meters is in perfect focus, it remains to be seen the degree of defocus present over the 1 to 5 meter depth of field.

As was said before, a point source on the lens axis at 1000 mm will form a point image at 35.25 mm . All the rays entering the lens aperture are bent to pass through this point. However the detector, placed at 35.76 mm , will not observe a point image but a blurred image termed a "circle of confusion." 8 The diameter of this circle can be calculated by similar triangles from the axial distances and the lens diameter.
$\frac{\text { diameter of circle of confusion }}{\text { distance array from focus pt }}=\frac{\text { lens diameter }}{\text { lens to focus pt }}$
diameter circle $=.25 \mathrm{~mm}$

Since the laser spot on the terrain is not a point source but has a diameter of approximately 5 cm of 5 meters the laser image formed without blurring effects is

$$
\frac{i}{5 \mathrm{~cm}}=\frac{35.25 \mathrm{mmn}}{5.0 \mathrm{~meter}} \rightarrow i=.35 \mathrm{~m} \mathrm{~m}
$$

Combining the blurring effects with the unblurred image the composite image can be expected to be 0.6 mm in diameter. Although the blurring effects are appreciable, considering that the photo receptive elements are on 1 mm centers, it is the diode array itself which constrains the detector's resolution.

### 3.3 Array and Preamplifier Board

Housed within the cylindrical detector head a single 2 inch diameter circuit board holds the 20 element linear array, biasing circuit and 20 operational amplifiers configured as current-to-voltage converters, Fig. 17.

The 20 element array purchased from Centronics, New Jersey, consists of 20 photodiodes each, $9 \mathrm{~mm} \times 4 \mathrm{~mm}$ on 1 mm centers in a single 22 pin dual-in-line-package. Located in the center of the analog board on the optical axis of the receiving lens, the photodiodes convert the incident laser radiation into current at the 20 individual diode outputs. The sensitivity of the photodiodes are $.43 \mathrm{amps} /$ watt incident.

The biasing and supply circuits provide filtered $\pm 15$ volts for the op-amps and 24 volts used to back-bias the photodiodes. The diodes in the supply lines were included to "idiot proof" the power inputs.

The current pulses from the photodiodes are transformed into negative going voltage waveforms by the 20 operational amplifiers


FIGURE 17a. DETECTOR PREAMPLIFIER BOARD


Rear of Detector Head


FIGURE 17b. DETECTOR HEAD LAYOUT


100mv/div

Detector Element \#19

Return with Laser spot 1 meter along polished cement floor, approx. 1.5 meters from Detector

Detector Element \#1
Return from 3.5 meters along floor

configured as inverting current to voltage converters. The $82 \mathrm{k} \Omega$ resistors were chosen as a compromise between the high gain and the needed amplifier stability. The expected signal from these converters can be estimated by recalling the amount of power entering lens (2.5 $\mu$ watts) multiplied by the array's photosensitivity (. 43 amps/watt) and the feedback resistance ( 82 kL ).
2.5 uwatts $\times .43 \mathrm{amps} /$ watt $\times 82 \mathrm{k} \Omega=88 \mathrm{mv}$.

Measured returns from poor reflectors at three meters exceeded 50 mv with an acceptable level of oscillation at the amplifier output, Fig. 18.

The Harris HA 4905 was selected because the quad operational amplifiers in a single dual-in-line-package were easy to service, offered high packing density, while still providing an 8 MHz gainbandwidth product. The five packages have been socketed for easy replacement. Finally the coupling capacitors between the amplifiers and the comparators block any d.c. offsets which may vary with temperature and device age. Overall, the analog board provides the conversion from the lens' spatial filter to the 20 voltage pulses in a simple, serviceable package while still meeting the physical constraints of the rotating mast.

### 3.4 Comparator Card

The comparator card, Fig. 19, receives the 20 amplified photodiode signals from the analog board through the 26 cond cable. The pulses it receives are negative going with an amplitude of from


FIGURE 19. COMPARATOR AND THRESHOLD BOARD


FIGURE 19b. COMPARATOR BOARD LAYOUT AND TEST POINT LOCATION
ten to 500 mv , Fig. 18, Also superimposed on each line is noise generated primarily by the laser firing. The 390 pf capacitor and the 100 ohm resistor form a filter which is used to prevent the laser noise from triggering the comparators.

The comparators selected are HARRIS HA-4605. Again the quad comparators offer the advantage of high packing density, easy replacement and fast ( 130 ns ) fall time. The amplified photodiode pulses are fed into the positive inputs of the comparators and cause a lowgoing TTL-level at the comparator output whenever the pulse amplitude goes below the threshold set on the negative inputs.

The thresholds are adjusted by the voltage-dividing action of the 20 threshold potentiometers. The reason 20 individual thresholds are necessary is in part so that d.c. offsets can be nulled out. Also the individual adjustments allow for a correction of the lens roll-off or vignetting, whereby the image received on the edges of the field of view are less intense than images closer to the optical axis. By lowering the thresholds of the channels associated with the off-exis elements they can be made more sensitive than the center photo receptors, thereby fattening the lens roll-off. Furthermore the separate thresholds allow that the lower detectors, which normally operate at farther ranges, can be adjusted to be more sensitive than the higher elements. Thus each channel can be tuned to the expected signal level providing a large dynamic range over the array while minimizing the possibility of false returns. The plus and minus supplies to the potentiometers are regulated down from $\pm 15$ volts to increase the
discrimination of the threshold adjustment as well as isolate the thresholds from the $\pm 15$ volt supplies.

### 3.5 Digital Board

The digital board, Fig. 20, accepts the 20 digitized channels from the comparators along with the Laser Fire Signal. The Laser Fire Signal, generated by the Mast Controller, causes the laser to fire and is used by the receiver to synchronize itself with the incoming laser energy, Timing Diagram, Fig. 21. Upon receipt of the Laser Fire Signal the 20 j-k flip-flops are reset clearing the stored channels from the last laser shot. The two one-shots then form time window during which the $j-k$ flip-flops are allowed to change state if their clock inputs experience a falling transition. By this window mechanism only channels which have been illuminiaed during the window will be set, thereby discriminating against noise pulses which may occur at other times. In fact since the laser return trails the injected laser noise by several hundred nanoseconds (due to the relatively slow response of the photodiode, amp, comparator circuit), by delaying the beginning of the window the system effectively masks out the early laser noise pulses. Both the delay and width of window are independently variable with the ports $P 1$ and $P 2$ respectively.

After the window has closed one or more channels should have been latched. The question now arises of how to relay this information to the obstacle detection computer. Since the data path to the computer is only 16 bits wide this precludes the sending of all 20 bits in parallel. One could send five bits to indicate which element has


FIGURE 20a. DIGITAL TIMING AND DISPLAY BOARD


FIGURE 20b. DIGITAL BOARD LAYOUT


FIGURE 21. DETECTOR SYSTEM TIMING DIAGRAM
been illuminated, yet this would ignore the event of more than one return. The scheme implemented on the digital board generates five bits corresponding to the highest photodiode excited and another five bits representing the lowest illuminated element. The two five-bit words are then concatenated into a single ten-bit word to be sent in parallel to the Mast Controller FIFO (First-In-First-Out buffer).

The details of this encoding scheme can best be seen in the digital board schematic, Fig. 20. These are two sets of octal priority encoders (74148). The lower set has det $\# 20$ on its least significant input and the upper set has det \#l on its least significant input. Note that if no imputs are set then the output is uniquely $\alpha_{04}^{\# 1}=0 \alpha_{04}^{\#}=7$, a result which is used in the obstacle detection algorithms to indicate a missing return.
4.1 Pointing Angle, Focus, Aperture, Thresholds, Time-Window

Alignment of the detector involves only five variables: pointing angle, focus, aperture, comparator thresholds and timewindow. The pointing angle or attitude of the camera, defined as the angle between the mast and the center of receiver f.o.v., is adjusted by means of the worm gear connecting the camera to the mast. With 120 teeth on the gear a single turn of the worm results in a $3^{\circ}$ change in camera attitude. The desired pointing angle must be fully determined by future evaluation of its performance in $\pm 30$ pitch and roll, but has been estimated as $40^{\circ}$ in early level terrain experiments. The focus of the camera is determined by the lens to array position. Since the laser at 904 nm is outside the normal photographic range, the standard Pentax lens mount to image plane distance cannot be used. Instead the correct lens to array distance must be determined experimentally. From the lens-makers-equation $\frac{1}{f}=\frac{1}{s}+\frac{1}{d}$ the image distance, $s$, can be determined from the object distance, $d$. If the mean object distance is two meters then for a 35 mm lens the image distance is 35.62 mm , that is, if the lens is set to focus at two meters then the array should be placed at 35.62 mm from the idealized principle lens plane. But since the position of the idealized lens. plane is unknown, the plane to array distance must be adjusted indirectly through field of view measurements. A 19 mm image pl ane, measured from the centers of 1 st diode to the 20 th diode, when placed from 35.62 mm from
the lens plane should provide a $29.87^{\circ}$ field of view．Hence to align the lens focus，the lens should be set to two meters（using the infra－ red scale pointer）and the position of the array board adjusted until a $29.87^{\circ}$ f．o．v．is obtained．The field of view can be quickly determined by holding the receiver fixed and moving the laser while observing the amplified outputs of dets 20 and 1．Using the infra－red TV camera and monitor，measure the position of the laser spot when channels 1 and 20 are at their maximum．The field of view of the lens can be obtained from the laser positions．

The aperture of the lens should be made as small as possible while still allowing reliable returns over all surfaces encountered． The smaller size increases the depth－of－field，while decreasing the double returns possible with strong returns from ranges in poor focus．

The thresholds should be adjusted with the laser reflecting off a level surface of poor reflecting characteristics．In this arrangement the lower photodiodes（det $⿰ ⿰ 三 丨 ⿰ 丨 三 一 1)$ will be operating at longer ranges and hence receiving smaller signals．By turning the poten－ tiometers（clockwise increases sensitivity）the element－by－element sensitivity can be tailored to the varying ranges expected．With the laser spot in the center of the $n{ }^{\text {th }}$ photodiode，determined by observ－ ing amplified photodiode signal，the threshold should be set so that a return is possible off the poorest reflector with only minimum overlap between adjacent elements．

The time－window as explained earlier should occur only during expected transitions from comparators．By adjusting the delay poten－
tiometer, begin the window with the earliest detecter's return. The laser noise which may pass through the comparator threshold then occurs before the window and will be ignored. The window width should then be adjusted to accept all genuine detector transitions ( $-2 \mu \mathrm{sec}$ ).

Calibration of the detector can be performed in the same manner as the field of view was determined in the focusing procedure. By measuring the position of the laser spot which maximizes the return to each detector their respective elevation angles can be measured. The first step in the procedure is to level the mast. A bubble level was useful here. The next step is to disconnect the mast controller (at least Motor Speed and Rate Buffer Boards) so that the mirror can be hand turned and the drive to the laser opened. By using an external trigger to fire the laser ( $1 \sim 3 \mathrm{kHz}$ ) the spot can be observed with the silicon-intensified TV camera to move as the mirror is turned manually. As the laser spot is moved in and out along the floor, the location at which each diode's return is maximized can be marked on the floor. (Note, the surface of the floor should have constant reflectivity.) Since the height (Z) of the floor can be assumed to be zero these ranges can be converted to angles for each of the detector elements.

$$
\alpha(\alpha \#)=\tan ^{-1} R^{\#} / h_{\alpha}
$$

where $\alpha=$ elevation angle of $n^{\text {th }}$ detector;
$\alpha^{\#}=$ detector number 1 through 20 ;
$R=$ range associated with $n^{\text {th }}$ detector;
$h_{\alpha}=$ height of detector above floor.
This same procedure can be used to calibrate the laser.

Another method of calibration is to hold the laser spot fixed on a surface of constant reflectivity at a constant range while the detector is rotated. Figure 22 is a graph showing the magnitude of each channel's amplifier output as the receiver was rotated. These measurements were taken with a good reflector, the aperture wide open, and at a range of three meters. The variation in peak height from channel to channel is a result of device non-uniformity. The rounding of the individual peaks is indicative of both receiver defocus and a relatively large laser spot on the reflector. There is considerable overlap between adjacent diode response curves and the thresholds must be set carefully if only a single return is desired. The problem is greatly complicated when a multiplicity of ranges and reflectivities are considered. Yet with proper adjustment of the lens aperture and the 20 thresholds, returns can be restricted to single returns and double returns from adjacent elements. It should be noted that the overlap increases the effective resolution of the detector since the number of possible returns has been increased to 39,20 single returns and 19 double returns. The double (or triple) returns, provided for in the detector's encoders, can be exploited by the obstacle detection computer to effectively double (triple) the receiver's resolution.

In general the measurements taken thus far have shown that the angular separation between elements ( $\sim 1.5^{\circ}$ ) can vary somewhat with focus but remain constant over the length of the array without the pronounced pinching that occurs with receivers with wider fields of

figure 22. stgnal return vs. elevation angle of source
view. Since the separation can be assumed constant over the array the simplest calibration would only measure the angles between two elements, presumably det $\# 1$ and $\# 20$, and divide by the number of incremental separations between them. Thus if the focus has somehow been changed two quick measurements can re-establish the relation between detectnr number and angle, $\alpha$, required by the obstacle detection computer.

## PART 5

## 20 ELEMENT RECEIVER CONCLUSION

The 20 element receiver has been in operation since the fall of 1979. During this time its overall performance has been satisfactory meeting its design goals, yet some shortcomings have subsequently tecome apparent.

### 5.1 Secondary Returns

The most surprising problem was that of secondary returns. As can be seen in Fig. 23, secondary returns occur when the detector receives two retums, one from the primary laser spot and the second from the primary's reflection off a second surface. This occurs when specular reflectors, such as mirrors, polished metal or glass, either deflect the primary beam to a second surface or allow the viewing of the primary spot from a second vantage. The first attempt at a solution was to use diodes and amplifier feedthrough to the comparator's thresholds to select the channel with the maximum signal. This worked only marginally due to the low level signals involved and the fact that primary and secondary signals are often the same or even reversed with respect to their magnitudes. Although Mars may not have many specular surfaces in other applications any laser triangulation system will suffer from this malady, a problem which worsens with detector field of view.

### 5.2 Missing Returns

Another problem facing laser triangulation is that of missing


## Primary reflection specular



FIGURE 23. SECONDARY RETURNS TERRAIN AMBIGUITY


In this region only detector has access If laser and detector reversed, missing returns would only occur here.


Since on Mars obstacles with occlusions below them are rare, reversing laser and detector minimizes possibility of missing returns

FIGURE 24. MISSING RETURN PROBLEM
returns. In this case some object blocks the path of the laser spot between the terrain and the receiver. In our application this normally happens when the laser fires over an obstacle and strikes the terrain at a point shadowed from the detector by the same obstacle, Fig. 24. On a Martian rover the fact the surface is generally a function of two variables can be exploited to minimize the number of missing returns. In particular by mounting the detector above the laser a missing return could only occur if the laser spot were beneath an obstacle and its return blocked from above. Hence, in the peculiar environment of Mars, where overhanging objects are rare, by placing the detector above the laser and bringing the two as close as possible the problem of missing returns would be minimized.

### 5.3 Field of View

The field of view ( $30^{\circ}$ ) of the 20 element detector presently limits the roll and pitch the vehicle can sustain. Although it is yet to be experimentally demonstrated it seems clear that slopes exceeding $\pm 25^{\circ}$ with respect to the vehicle would be outside the sensors' scope. Building a second receiver would alleviate this problem as would rotating the detector under computer control.

### 5.4 Reliability

The reliability of the device as presently constructed is also of some concern. The many small wires and parts under rapid acceleration or continued handling are a source of failures and difficult to locate. The use of printed circuit construction would be a
great improvement as well as an aid in developing a second detector if desired.

### 5.5 Resolution

The resolution of the device at $1.5^{\circ}$ has yet to be proven sufficient. Difficulties in the Mast Controller and the PRIME computer interface have prevented the evaluation of the perception system until late August of 1980. Should it prove necessary the multiple return encoding in the detector could allow through the use of double overlap returns a two-fold increase in resolution.

PART 6
1024 ELEMENT CHARGE COUPLED PHOTODIODE RECEIVER

### 6.1 General Description

The 1024 element receiver is divided into two main packages: 1) the detector head, ( $33 / 4^{\prime \prime} \times 33 / 4^{\prime \prime} \times 43 / 4^{\prime \prime}$ ), Fig. 25 , housing the Array Board, a $12.5 \mathrm{~mm} / \mathrm{f}: 1.9$ lens and filter, and 2) the $41 / 2 \times 61 / 2^{\prime \prime}$ Mother Board holding the digital timing and signal processing circuits. The detector head is mounted below the laser with the optical axis of the lens, the 1024 element linear array, and the laser ray all in the same vertical plane, Fig. 3. Sampling of each of the 1024 photodiodes is accomplished by means of a parallel transfer into dual charge coupled analog shift registers aboard the array chip, Fig. 26 , followed by sequential passage through a peak detection circuit, Fig. 27, on the Mother Board. Timing for the CCD registers including four phase clocks, transfer and reset strobes are generated on the Mother Board and relayed to the Array Board along a six-inch 16 cond, ribbon cable. Amplified video output then returns to the signal conditioner and peak detector along a coaxial cable. Adjustments include video sample rate ( $1-5 \mathrm{MHz}$ ) and peak threshold. A four-digit BCD display indicates which of the elements has been illuminated.

Overall the 1024 element change coupled receiver, in conjunction with 100 watt 200 ns laser pulser operating at 2 kHz offers 3 min 30 sec resolution within $60^{\circ}$ field of view and a three meter range.

### 6.2 Optics

The lens employed in the 1024 element receiver is a Cosmicar


Motherboard

FIG. 25. 1024 ELEMENT RECEIVER CONSTRUCTION


FIG. 26. 1024 CCPD array block diagram


FIG. 27. BLOCK DIAGRAM OF 1024 CCPD RECEIVER

B1218, with a 12.5 mm focal length and an f-number of 1.8. Designed for use with one-inch vidicon tubes its maximum field of view is given as $64^{\circ} 50^{\prime}$. The 16 mm length of 1024 CCPD when combined with the 12.5 mm lens should provide a field of view of $65^{\circ}$. The resolution of the receiver can be estimated as $65^{\circ} / 1024 \sim 3 \min 30 \mathrm{sec}$.

Another optical component in the 1024 element receiver is a bandpass filter used to limit the ambient light striking the array. Located between the lens and the array a Wratten 87C lowpass filter only transmits light with a wavelength exceeding 850 nanometers, Fig. 28. Since the response of a silicon photodiode begins to fall off rapidly around 950 nanometers the combination of the Wratten filter and photodiode response effectively forms a bandpass filter around the laser wavelength (904 nm).

### 6.3 1024 Element Charge Coupled Photodiode Array

The photosensitive window of the array is comprised of a 1024, $16 \mu \mathrm{~m} \times 16 \mu \mathrm{~m} \mathrm{p}-\mathrm{n}$ junction photodiodes in a row 16 mm long, Appendix B. Light incident on the sensing aperture generates photocurrent which is integrated and stored as a charge on the capacitance of the reverse biased photodiodes. Pulsing of the transfer gate then simultaneously transfers the accumulated photocharge into one of two 512 stage surface channel CCD analog shift registers. The odd diodes are switched into one register and the even into another. Readout is accomplished by clocking the shift registers so that the charge packets are delivered sequentially to two on-chip charge detection diodes. The outputs of the two change detectors are then multiplexed off-chip

fig. 28. relative response of silicon photodiode and wratten filter
to obtain a step-wise continuous video signal.
The speed of the device, limited by transfer inefficiency and thermal heating considerations, is given as 5 MHz . However in this ranging application the array is operated at 2 MHz or a 2 KHz line rate. This is because the laser diode, which must be run at maximum power (100 watts) and pulse width ( 200 ns), in order to be sensed by the receiver, can only be run at 2 KHz if it is not to exceed its rated duty factors (.04\%).

The feasibility of the receiver/laser combination was analyzed before the array was acquired assuming a $40 \mathrm{~ns}, 100$ watts laser and an array noise equivalent exposure as given of $10^{-3}$ watts $/ \mathrm{cm}^{2} .^{9}$ The calculations, similar to those in Part 2.1, indicated that the signal returning from three meters would be approximately 40 times the noise level and hence the array was purchased. Subsequently when the array was tested it was discovered that the signal, from a poor reflector, equalled the noise level at a two-meter range. The error in the analysis was the assumption that the laser could be collimated sufficiently to ensure that the laser spot image was concentrated on only a single element. To fall within a single element the laser spot size, $x$, at three meters, would have to have been

$$
\frac{x}{3 m}=\frac{16 \mu \mathrm{~m}}{12.5 \mathrm{~mm}} \rightarrow x=3.7 \mathrm{~mm}
$$

To collimate the .41 mm laser source to 3.7 mm at three meters would have taken a lens of more than 300 mm with an exit pupil much larger than the mirror face. The lens that was used has a focal length of 85 mm which at a range of three meters gives a spot size of
approximately 2 cm . A 2 cm spot has as much area as $25,4 \mathrm{~mm}$ spots and hence due to the lack of sufficient collimation the actual intensity on the array was $1 / 25$ th of the level expected, or 1.6 times noise level. It was necessary to increase the 40 n sec pulse to 200 ns and at the same time reduce operating speed from 10 KHz to 2 KHz to enable the 1024 element receiver to function in the three meter range.

A further consideration involved in the selection of this array was the anti-blooming gate feature whereby the integration period can be made shorter than the line read-out time. By pulsing the antiblooming gate, $\mathrm{V}_{\mathrm{AB}}$, the photodiodes are reset thereby limiting the integration period to the time between the trailing edge of $V_{A B}$ and the trailing edge of the transfer gate, $\phi_{T}, 1.5 \mu s e c$ minimum. By this mechanism the anti-blooming gate acts as an electronic shutter that when open accepts the pulsed laser energy but when closed reduces the ambient light noise by a factor of

$$
\frac{500 \mu \mathrm{sec}}{1.5 \mu \mathrm{sec}} \sim 300
$$

The anti-blooming gate also allows the array to be run synchronously with the laser. Without the $V_{A B}$ feature the charge accumulated since the last laser shot would have to be shifted out before the laser could be fired, effectively halving detector speed.

The array board is a $3 \times 3$ inch printed circuit board holding the 1024 CCPD array and its associated drivers, bias circuits and amplifiers, Fig. B1. The drivers (MH 0026), Fig. B2, accept the TTL level timing signals and level shift them to the necessary 15 volt swing. The
bias circuits generate the d.c. bias voltages needed; substrate, output drains, input and output gate biases. The amplifier is a CA 3127, Fig B3, configured as a common emitter current amplifier with a current mirror as base bias. 10

The mother board contains the digital control/timing circuitry and the analog signal processing circuit, Appendix B, Fig. B4. The timing begins with oscillator, Ul, whose frequency can be varied by PI from 1.5 to 8 MHz giving sample rates of 750 KHz to 4 MHz . This clock frequency can be monitored at pin Jl-2. The four-phase clocks, $\phi_{1}, \phi_{2}, \phi_{3}, \phi_{4}$, transfer pulse $\phi_{T}$, and the sampling pulse are all derived from this clock. Scan Start on pin J1-E can be used to sync up to the line rate.

The analog circuitry consists of sample and hold circuit and a d.c. restoration curcuit. The sample and hold circuit helps to isolate high frequency noise while the d.c. restoration circuit can be used to set the d.c. level of the video signal and forms a high pass filter against noise sources with wavelength longer than line time. The switch edge cancellation circuit is used to cancel clock noise coupled through the sample switch and the blanking circuit pulls the video to zero volts between the 1024 th sample pulse and the next transfer pulse.

## PART 7

ADDITIONS/MODIFICATIONS TO RETICON DEMONSTRATION BOARD

The array and demonstration board as purchased from Reticon Corporation was described in Part 6 and in Appendix B. Some modifications and additions were required however to incorporate the demonstration board into the elevation scanning mast controller. Among these adaptations were: increasing amplification of preamplifier; addition of resettable peak detectors; anti-blooming gate implementation; changes to timing to allow array to run synchronously with mast controller; interface to controller FIFOs; and a display of which elements were illuminated.

### 7.1 Preamp Aboard Array Board

An LM 318 high speed operational amplifier was inserted for U7, LM310 unity-gain buffer, on the array board, Fig. 29. Since the pin-out was almost identical this was chosen as the location to increase the gain before the sample-and-hold circuit. The increase in gain was possible because the old circuit was designed for one volt saturation signals whereas the laser returns are on the order ot ten to 100 millivolts. The gain allowed the signal level to exceed the switching noise introduced in the sample-n-hold circuit and allows easier detection of laser return.

### 7.2 Anti-Blooming Gate

As in the device data sheet the charge on the photodiodes can be eset by pulsirg $V_{A B}$ to +5 volts for at least 1 usec. The integration


FIG. 29. PREAMPLIFIER AEOARD ARRAY BOARD
period is then the time between the trailing edge of $\mathrm{V}_{A B}$ and the trailing edge of next $\phi_{T}$ pulse. In this application $V_{A B}$ is held high except when the laser is fired. This avoids a 1 usec time delay in firing the laser associated with holding $V_{A B}$ high. Instead when the laser fire signal arrives $V_{A B}$ is brought Lo and $\phi_{T}$ is applied simultaneously. In fact $V_{A B}$ is tied directly to $\phi_{T}$ on the array board. This effectively reduces all ambient sources by a factor of approximately transfer pulse width/line scan time $=1 \mu \mathrm{sec} / 500 \mu \mathrm{sec} \sim 500$.

### 7.3 Peak Detectors

The function of the peak detector is to find the maximum signal along the line scan. $U_{16}$ and $U_{17}$ on the mother board were replaced with LM311 comparators, Fig. 30. The first comparator ensures that the storage capacitor is changed to the maximum level encountered while the second comparator senses where the video signal exceeds the stored peak. When the peak is exceeded the latches monitoring the array counter store the count value at the time of the peak. The blanking circuit on the mother board is used to reset the peak storage capacitor. The resetting prepares the peak detector for the next line and also forms a high pass filter discriminating against. noise with periods larger than the lla. rate.

### 7.4 Synchronization with Mast Controller and Dark Current Limiting

The demonstration boards run on their own internal clock asynchronous with anything external. As described in Appendix B there are two counters, the Array Counter and Integration Counter, which


FIG. 30. PEAK DETECTOR WITH BLANKING RESET


FIG. 31 TIMING DIAGRAM RETICON MODIFICATIONS
control, the time the blanking circuit is activated, and the time between transfer pulses. It is necessary that this system be modified to operate at a rate dictated by the mast controller. When the laser fire signal arrives the array must be ready to accept the reflected laser light and shift it out before the next laser pulse.

Before any laser pulse arrives the system is in the Stand By mode. In this mode the $4 \phi$ clocks continue running, thereby preventing the build-up of dark current in CCD shift registers. Also the $\phi_{\text {TRANSFER }}$ is disabled holding $V_{A B} H I$ and keeping $C C D$ registers free of charge. This ensures that the device will be prepared to integrate light as soon as possible after receipt of Fire Laser In.

This Stand By modification is realized by inserting Fire Laser In between the integration counter $U_{12}$ and $U_{8}$. Then until the arrival of FIRE LASER IN the system continues to cycle the four-phase clocks while $\phi_{T}$ and $V_{A B}$ remain $H I$. Unon the arrival of FIRELASER IN, the transfer $\phi_{T}$ and $V_{A B}$ are enabled, and the blanking generator releaṣed. Fire Laser Out (same as scan start) which is used to fire the laser, falls with $\phi_{T}$, and $V_{a}$, Allowing time for the array to begin integration before the laser is fired. After the transfer is complete the $4 \phi$ clocks shift the video-out to the signal processing circuits until the array counter counts 1024 and sets the blanking and resets the peak storage capacitor.

### 7.5 Latches Interface to Mast FIFOs

A pair of 74174 hex latches are used to monitor the array counter and are clocked by the peak detector output. At the end of the
line scan they contain the value of the array counter when the peak was sensed. This information is held for the FIFOs to absorb until the next laser pulse occurs and the blanking Flip-Flop is released.

### 7.6 Display

A display has been constructed giving the BCD count of the number of clock pulses at the time the peak was sensed. Constructed from 4-74143, counter and seven segment drivers, the counters are clocked by the clock out (J1-7) and latched by the output of the peak detector.

PART 8
1024 CCPD RECEIVER CONCLUSION

## $8.1360^{\circ}$ Vision Around Vehicle

As the CCPD receiver revolves on the mast it sweeps it's $60^{\circ}$ field of view $360^{\circ}$ around the vehicle. During that time 1024 shots ( 32 azinuth. 32 elevation) are taken with $500 \mu \mathrm{sec}$. alloted for each shot. With antiblooming integration control only $1.5 \mu \mathrm{sec}$. are used to gather light. In a 8 to 10 second scan the remaining time could be used to interlace video information so as to gather a grey level image that completely surrounds the vehicle. What would be needed would be a mechanical shutter removing the infrared filter during the grey level line samples and a video frame store for display. Also some form of slip ring to relay the video stream from the rotating mast. The camera should of course be mounted as high as possible to enable the system es atrvey behind the machine as well. Moving the laser and sensor closer together, made possible by the high resolution of the 1024 CCPD array, reduces the problem of secondary returns (sec. 5.1), and placing the detctor above the laser aids in the missing return case (sec. 5.2).

## 8. 2 Buried Channel Charge Coupled Array

The use of a buried channel CCD array such as that manufactured by Fairchild Camera would offer better performance since they are inherently a lower noise device ${ }^{11}$ when compared to surface channel arrays such as the Reticon CCPD product. Surface channel devices require a bias charge injected into their inputs to overcome the image smearing caused by partially filled interface states with random relaxation times. Injection of charge at the input as well as the reset required at the output ports introduce even more
noise into the video stream.
A buried channel device requires no bias charge hence no injection noise and the reset of the charge detector may be reset at the end of the line scan instead of after every pixel. This method of avoiding injection and reset noise was attempted with the Reticon array but proved unfeasible due to excessive smearing of the laser image along the array length.

Cooling of the device would also minimize noise since much of the noise is thermally generated and the array has been observed to run quite warm.

### 8.3 Improved Rangefinder

The ultra high resolution of the 1024 element CCPD array permitts the laser and receiver to be brought much closer together than possible with the 20 element sensor. In fact if the laser and detector were brought close enough to be reflected off the same mirror then neither the laser nor the receiver need rotate, Fig. 32. Of course the mirror used would have to be larger than the 8 -sided s-anner presently used.

With the laser and detector stationary no slip rings would be necessary. Furthermore the two components when so close are housed in a much smaller and practical package. Instead of a 2 meter mast roating about a vertical axis, a small, ( 6 inch), mirror mounted in gimbals would scan the laser ranger only where desired through a dust tight window.

The mirror could still rotata through $360^{\circ}$ in both azimuth and elevation but slip rings would be needed only for the elevation motor and encoder. An 8-sided mirror with 6 -inch faces would be prohibitively large. It seems preferable to employ a single faced mirror foregoing the $360^{\circ}$ scan and drive the gimbals with stepping motors. Then the time sacrificed


FIG. 32. IMPOVED RANGEFINDER
in using the one-sided mirror could be recovered by not scanning the rasger to the rear of the machine. In this case without $360^{\circ}$ azimuth scanning even the elevation slip rings could be eliminated.

- The question of scan time can be analyzed by assuming an angular resolution, dividing it into the field of view, and multiplying by the time required per motor step. The present scanner has a resolution of $0.35^{\circ}$ in elevation and $1.4^{\circ}$ in azimuth. Given that the field of view is desred to be $60^{\circ}$ in 32 elevation sweeps and $140^{\circ}$ in azimuth:

$$
\Delta t_{\text {scan }}=\left[32\left(\frac{60^{\circ}}{.35^{\circ}}\right)+\frac{140^{\circ}}{1.4^{\circ}}\right] \Delta t_{\text {step }}
$$

Stepping motors with a step frequency in excess of $1000 \mathrm{steps} / \mathrm{sec}$. are available Appendix B giving a $\Delta t_{\text {step }}$ of approximately lmsec., and a $\Delta t_{\text {scan }}$ of $\simeq 6$ seconds.

If inertial effects and others contributed a factor of two we still might assume a scan time of 12 seconds; which for a 3 meter range, 1/4 meter /sec. velocity, allows the terrain to be scanned once before the vehicle rolls over it. Or at the mission study speed of 1.5 meters/min. each portion of the surface would be sampled 10 times.

The lens used in the scheme would be of a long focal length, 100 mm . or more. It should be chosen so that it's field of view encompasses the laser ray from $\mathbf{- 1}$ meter to fime and would vary with laser/receiver separation. For example at a separation of 100 mm , a receiver field of view of $4.56^{\circ}$ would be suficient. A focal length of 200 mm . would be required to subtend $4.56^{\circ}$ with a 16 mm . array.

Another optical problem, in addition to trying to find a long focal length lens with a small f-number, is the small depth of field of such long focal length lenses. For example, with a 100 mm . lens, a 1 meter spot
would focus at 111 mm . and a 5 meter laser spot at 102 mm . A solution is possible since a given element can be expected to operate at the same range when the laser and receiver are fixed with respect to each other, Fig. 33. The array can be tilted to alleviate the depth of field allowing each element to function at or near it's optimum focal distance.


FIG. 33. OPTICS FOR IMPROVED RAyGFINDER

PART 9
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APPENDIX A


| HARFIS <br> SEMICONDUCTOR <br> A unnewon op mande oonergintions | HA-4900/4905 <br> Precision Quad Comparator |
| :---: | :---: |
| - FEA TURES | - DESCRIPTION |
| - FAST RESPONSE TIME <br> - LOW OFFSET VOLTAGE <br> - LOW OfFSET CURRENT <br> - SINGLE OR OUAL-VOLTAGE SUPPLY OPERATION <br> - selectable output logic levels <br> - active pull-up/Pull-DOWN OUTPUT CIRCUIT - NO EXTERNAL RESISTORS required <br> APPLICATIONS <br> - threjhold detector <br> - zero-crossing detectór <br> - WINOOW DETECTOR <br> - ANALOG INTERFACES FOR MICROPROCESSORS <br> - high STABILITY OSCILLATORS <br> - LOGIC SYSTEM INTERFACES | The HA-4900/4905 are monolithic, quad, precisi3n comparators offering fast response time, low offset voltage, low offset current, and virtualiy no channel-to-channel crosstalk for applications requiring accurate, high speed, signal level detection. These comparators can sense signais at ground level while being operated from either a single +5 volt supply (digital systems) or from dual supplies (analog networks) up to $\pm 15$ volts. The KA $-4900 / 4905$ contains a unique current driven output stage which can be connected to logic system supplies ( $\mathrm{V}_{\text {Logic }}{ }^{+}$and $\mathrm{V}_{\text {Logic-1 }}$ ) to make. the output levels directly compatible (no external components needed) with atiy standard logic or special system logic leveis. In cornbination analog/digital systems, the design employed in the HA-4900/4905 input and output stages prevents traubiesome ground coupling of signals beiween analog and digital portions of the system. <br> These comparators' combination of features makes them ideal components for signal detection and processing in data acquisition systems,- test equipment, and micropracessor/ analog signal interface networks. <br> Both devices are available in 16 pin dual-in-line ceramic packages. The HA-4900 operates from $-55^{\circ} \mathrm{C}$ to $+125^{\circ} \mathrm{C}$ and the HA-4905 operates over a $0^{\circ} \mathrm{C}$ to $+75^{\circ} \mathrm{C}$ tempersture range. |
| PIN OUT | SCHEMATIC |
| Top View <br> \|CAUTION: These devices are sensitive to eiectrostatic discharge. Users should follow IC Handling Procedures specified on pg. 1-4. |  |

[^0]
## ABSOLUTE'MAXIMUM RATINGS (Note 1)

| Voltrge Berween V+ and V- | 33 V - |
| :---: | :---: |
| Voltage Benween $V_{\text {Logic }}(+)$ and $V_{\text {Logic }}(-)$ | 18 V |
| Differential Input Voitage | \#15V |
| Peak Output Current | $\pm 50 \mathrm{~mA}$ |
| Internal Power Disipation(Note 7, 8) | 580 mW |
| < Storage Temperatura fiange | -650 ${ }^{\circ} \leq T_{A} \leq 150^{\circ} \mathrm{C}$ |

## ELECTRICAL CHARACTERISTICS

| PaAAMETER | TEMP. | $\begin{gathered} \mathrm{HA}-1900 \\ -55^{\circ} \mathrm{C} \text { to }+125{ }^{\circ} \mathrm{C} \end{gathered}$ |  |  | $\begin{aligned} & \mathrm{HA}-190 \mathrm{~S} \\ & 0^{\circ} \mathrm{C} \text { to }+75^{\circ} \mathrm{C} \end{aligned}$ |  |  | Uill, |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | min. | TYP. | max. | min. | TYP. | max. |  |
| input Chamacteristics |  |  |  |  |  |  |  |  |
| - Other Volime (Now 2 ) | 250C |  | 2.0 | 3.0 |  | 4.0 | 7.5 10.0 | $\mathrm{mbV}_{\mathrm{mbV}}$ |
| - OHmer Cume | $\begin{gathered} 250 \mathrm{C} \\ \text { Full } \end{gathered}$ |  | 10 | 25 35 |  | 25 | 50 70 |  |
| - Biancurmet (Nato 3) | $\underset{\text { Full }}{25 \varnothing C}$ |  | 50 | 15 150 |  | 100 | $\begin{aligned} & 150 \\ & 300 \end{aligned}$ | $\operatorname{cac}_{n A}$ |
| Input Sensitiviry (Note 4) | 250\% |  |  | 3.1 |  |  | 7.8 10.1 | mv |
| - Commen Mode Renge | Full | v- |  | $v+-2.4$ | $v$. |  | vi -2.4 | $v$ |
| transfer characteristics |  |  |  |  |  |  |  |  |
| Large Sipal Voituge Gsin | 2500 |  | 400K |  |  | 400 K |  | VN |
| Renponese Time (Tporlinote 5 ) | $25^{\circ} \mathrm{C}$ |  | 130 |  |  | 130 |  | ${ }^{\mathrm{m}}$ |
|  | $25^{\circ} \mathrm{C}$ |  | 180 |  |  | 130 |  | ns |
| OUTPUT Chagacteristics |  |  |  |  |  |  |  |  |
| - Oucpur Votiase imer |  |  |  |  |  |  |  |  |
|  | Fuil |  | 0.2 | 0.4 |  | 0.2 | 0.4 | $v$ |
|  | Fun | 35. | 4.2 |  | 3.5 | 4:2 |  | $v$ |
| Oumpt Curnot |  |  |  |  |  |  |  |  |
| ISink | full | 15 |  |  | 3.5 |  |  | ma |
| ISoure: | Fult | 10 |  |  | 3.0 |  |  | mA |
| POWER SUPFLY CHARACTERISTICS |  |  |  |  |  |  |  |  |
| - Suppir Cumenti ioriti | 2500 |  | 6.5 | 12 |  | 7 | 13 | mA |
| - Supent Comme Iprit | 250 C |  | 4 | ! |  | 5 | 7 | mA |
| - Supdy Currme, Ioritogic) | 2500 |  | 1.7 | 20 |  | 1.7 | 25 | $m A$ |
| Supaty Voltiep Range |  |  |  |  |  |  |  |  |
| $v$ * | Ful | +5.0. |  | +15.0 | +5.0 |  | +15.0 | $v$ |
| $v$ - | Full | -150 |  | 0 | -15.0 |  | 0 | $v$ |
| $V_{\text {Lopic }}$ (t) (Nate 71 | Fun | 0 |  | +15.0 | 0 |  | $+15.0$ | $v$ |
| $V_{\text {Lopel }}$ (Now 71 | Fow | -15.0 |  | 0 | -15.0 |  | 0 | $v$ |

-100\% tested for HA1-4900-8.



APPENDIX 3


Figure 10. Moded RC702A Evaluation Circuit Sehematic diagram (Array Board to be used with RC700A motherboard)

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NOTES: (UNLESS OTHERWISE SPECIFIED)

1. COMPONENT HEIGHT SHALL BE .SO MAX
2. LEAD PROTRUSION SHALL BE OG MAX
3. FOR SCHEMATIC SEE DWG OII-O275
4) INKSTAMP ASSY REV LETTER IN LOCATION SHOWN USING BLACK INK. COVER WITH ONE COAT OF CLEAR KRVLON NQIBO3.
3. TAPE HOLES FOR 22 PIN SOCKET, CIRCUIT SIDE, INSTALL SOCKET AT TOUCH-UP.
4. INKSTAMP APPROPRIATE DASH NR ACCORDING TO FOLLOWING TABLE:

| DEVICE | DASH NR |
| :---: | :---: |
| CCPD256 | -01 |
| CCPDO24 | -02 |
| ECPDIT28 | -03 |

7. BEND CI

OVER APPROX $45^{\circ}$ FROM VERTICAL IN DIRECTION TOWARD EDGE OF THE BOARD AND CIT.IB \& 19 TOWARD DI PRIOR TO SOLDERING




## RC700A/RC702A CCPD INTERFACE SYSTEM

## GENERAL

The RC700A/RC702A Interface system provides the timing, video processing, and blanking requirements for the Reticon CCPD series photodiode arrays. A simplified block diagram is shown in Figure 1. Each svstem consists of two boards; a $4.5 \times 6.5$ inch ( $11.4 \times 16.5 \mathrm{~cm}$ ) "motherboard" (RC700A) and a $3.0 \times 3.0$ inch ( $7.6 \times 7.6 \mathrm{~cm}$ ) satellite "array buard" with design to match each particular type of array. The two boards are interconnected, either directly or by means of a 16 -conductor 30 -inch (max) ribbon interconnect cable, to form a complete evaluation processing package for one of the Reticon 工harge Coupled Photo Diode (CCPD) arrays. A coaxial interconnect cable for the video signals is also required. The various satellite boards are identified by dash numbers in accordance with Table I.

## TABLE I

Board Description
RC702A-01
RC702A-02
RC702A-03
BOARD-ARRAY MATCH
Array Description
CCPD-256
CCPD-1024

## C15-C16 Capacitor Value

 1000 pf680 pf
none

The interface system, comprised of one RC-700A motherboard and one RC702A satellite board, is intended for use in evaluating the performance of one of the CCPD series of photodiode arrays. For some applications, the interface system may be used directly as a portion of a larger system such as a complete camera for a scanning controller.

- Accepts all CCPD series arrays.
- Operates at sampling rates from less than 100 KHz to 5 MHz .
- Sampled and held box-car video output.
- Output blanking provided between scans.
- Dynamic range more than 200:1.


## FUNCTIONAL DESCRIPTION

The system functions are separated into three major sections: (1) the sensor and its immediate interface, (2) the logic control, and (3) the analog signal processing circuit. Figure 2 is a block diagram which shows the organization of the circuits and their interrelations.

The logic control section includes the basic trigger generator, the four-phase clock circuits, and the transfer control circuits for the array, along with the


generation of the sampling and dc restoration pulses for the signal processing circuits.

The signal processing circuits comprise the preamplifier (physically located on the satellite board), the sampling switch and the sample-and-hold capacitor, the switch-edge-cancellation (SEC) capacitor, a simple filter, and the buffer amplifier for the output. The odd and even pixels from the array are summed at the preamplifier via the gain-balancing potentiometer. The combined samples are then resampled and held on the holding capacitor for the duration of a single pixel, then processed with dc restoration, filtering, blanking, and buffering.

In the resampling process, sampling switch energy induces extraneous charge on the capacitor. The switch-edge cancellation (SEC) trimmer capacitor is adjusted to minimize or cancel the extraneous charge by coupling charge from a negative-going version of the sampling pulse.

The de restoration circuit restores a ground reference to the sampled-and-held video during the retrace or blanking time of the array.

The simple RC low-pass filter is used to minimize_resimalmswitching_spikes_and to maintain minimum noise bandwidth consistent with the aray switching speed.

Normally, the filter is fixed, but in some situations it may be desirable to change the value of C23 to change the comer frequency of the filter to more closely match an unusual sampling rate.

## CIRCUIT DESCRIPTION

The array requires four-phase clocks, a transfer pulse, and complementary reset clocks. Please refer to the device data sheet for the detailed requirements. As specified in the data sheet, the clock waveforms have crossing-point and edge-control requirements, and the transfer pulse must occur at the proper time while the four-phase clocks are held stationary. The logic section provides the required sequencing and shaping, with all timing controlled by a basic trigger oscillator operating at twice the desired sample rate or four times the four-phase clock rate. The trigger oscillator is composed of U1, an edge triggered one-shot with regerative feedback. The frequency is controlled by $R_{1}, R_{2}$ and $C_{1}$. With values as given on the schematic, the trigger frequency may be varied from approximately 1.5 to 8 MHz , giving sample rates of 750 $\cdot \mathrm{KHz}$ to 4 MHz . By changing $C_{1}$ to 500 pi , the frequencies may be reduced to cover sample rates of 75 KHz to 1.25 MHz . With $\mathrm{C}_{1}$ deleted, the upper sample-rate limit becomes approximately 5 MHz . The trigger pulse may be monitored at connector pin $\mathrm{Jl}-\mathrm{Z}$.

The four-phase clocks receive basic sequencing by U19, a dual $D$ flip-flop connected in a twisted-ring conflguration. Each flip-flop produces comple-

|  |  | DRAWING NO. 045-0042 | C | SHEET | 3 | OF |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |

mentary square waves at one fourth the trigger period, with one trigger period of shift between the two flip flops to give the sequence $\varnothing_{1}, \varnothing_{2}, \varnothing_{3}, \varnothing_{4}, \varnothing_{1}, \cdots$

The $\emptyset_{1} / \varnothing_{3}$ flip flop also supplies the reset gates $\not \mathrm{RE}^{2}$ and $\not \varnothing_{R O}$, and the CCD receiving gate $\emptyset R G$. Clock shaping is desired for $\varnothing_{1}$ and $\varnothing_{3}$ to increase the overlap; for this purpose direct and delayed waveforms are combined in NOR U23 to give high-crossing clocks at the outputs of the translator U4 (RC702). Compensation for the lesser capacitances of the shorter arrays, CCPD1024 and CCPD256, is provided by additional capacitors C15 and Cl 6 at the arrey's clock input terminals. The values are as listed in Table I.

Clocks $\rrbracket_{2}, \varnothing_{4}$, $\emptyset_{\mathrm{R}} \mathrm{E}$ and $\star \not \mathrm{R}_{\mathrm{RO}}$ require no such clock shaping so that the appropriate generator outputs are directly translated through the MH0026 voltage translators and applied to the array. The clock swing must conform to value specified by CCPD Data Sheet.

The transfer pulse is required on ce per scan. It is sequenced at U21B by a combination of a transfer toggle from the transfer counter Ul5 and a clock pulse from the trigger divider U2A. Translator U3 (RC702A) converts the level to the range -4 to +5 volts for application to the array.

A dc restoration pulse and blanking pulse are similarly generated at U20B, translated by U13, and applied to switches Q3 and Q4.

The sampling pulse is generated once per signal sample by combination of the direct and divided trigger at U8-3. After translation, the inverted pulse is applied to sampler switch $Q_{1}$ and a complementary pulse for switch-edge cancelling coupled through SEC capacitor C12. The sampling-switch pulse is nominally 75 nanoseconds wide.

## ANALOG CIRCUIT

Odd and even outputs are balanced by R28 and applied to combiner-amplifier U6 followed by a unity gain buffer U7 (RC702A). The video signal from each output of the array is alternately reset at a high level of approximately +6 volts and released to carry the information at a lower level of approximately +2.5 volts. After combination and buffering, the saturated video signal has an amplitude of approximately 2.5 volts peak to peak, riding on a (dark) bias level of approximately +4.5 volts .

The sample-and-hold circuit consists of an FET switch, $Q_{1}$, hold capacitor $\mathrm{C}_{13}$, SEC capacitor C12, and buffer source follower Q2. The dc (ground level) restoration circuit is isolated by two unity-gain buffers, U16 and U17. The dc bias level for dark conditions is shifted to ground by clamping C20 during the

[^1]| SIR | DRAWING NO. | 045-0042 | C | SHEET 4 |
| :---: | :---: | :---: | :---: | :---: | :---: |

transfer period. At other times, charge does not change on C20 so that relative signal levels are transmitted with the proper ground reference. The reference level is adjustable at R25. Leakage may cause a shift in the reference, so C20 must have low leakage and for integration times longer than approximately 4 msec the value of C 20 should be increased to $0.1 \mu \mathrm{f}$ so that input current of U 17 will not upset the reference.

## BIAS CITCUITS

The system operates with the array substrate at -5 volts, supplied by regulator VR1. Fixed voltages for $V_{O G}, V_{R D}$, and $V_{D D}$ in accordance with the device data sheet are obtained by means of resistive dividers and filters from the +15 volt supply. Biases for $\mathrm{V}_{\text {IG }}$ odd and even are balanced to equal values by adjustment of R16, and their levels (approx. 10 volts) set bv R15 to give a (dark) video output level from U5 pin 2 or 21 of $2.0 . \pm 0.1$ volts. Later, the balance adjustment is further trimmed to minimize the dark signal odd-even unbalance (see set-up procedure).
$\mathrm{V}_{\mathrm{AB}}$ is nominally connected to ground with the jumper $\mathrm{W}_{1}$ provided on the board (see assembly drawing).

If desired, this jumper may be removed and a voltage from -3 to +5 volts applied to connector J1-F. Such a change will vary the level of saturation and so change the maximum output; however, antiblooming properties are affected and may be lost if VAB is taken more negative than approximately -3 volts.

## COUNTERS

There are three counters in the system: 1.) an array counter set to match the length of the particular array, 2.) an integration counter which determines the scan period, and 3.) a transfer counter which controls the transfer and flyback times independent of the other counters. The control sequence is as given in Figure 3.

The terminal count of the transfer counter controls the start of both the array and integration counters. This same control stops the transfer counter: The array counter runs at twice the rate of the other two, and has a preset number matched to twice the length of the array in use plus 4 . This preset number is determined by jumper links on the RC700A board as shown in Figure 4. The selected number as chosen by links W8-W12 must metch twice the number of elements in the array. The clocking rate and number sequence of the links result in the weighting of Table II. Links are ins talled to add weights to give the number chosen. For example, for a CCPD1024, only link W12 is selected from this group. Link W3 is permanently installed to account for internal timing.
PA

## TABLE II

WEIGHTING FOR ARRAY COUNTER
(See Figure 4)
Link Designation
(Binary No.) Equivalent Element Count

| $W-12$ | $(2048)$ | 1024 |
| :--- | :--- | ---: |
| $W-11$ | $(1024)$ | 512 |
| $W-10$ | $(512)$ | 256 |
| $W-9$ | $(256)$ | 128 |
| $W-8$ | $(128)$ | 64 |

The integration counter determines the scan period and hence the integration time. The count is selected by rocker switches S1 and S2 on the edge of the RC700A boar with the appropriate binary weighting designated on the board. The count as selected by the weighting must be not less than the number of elements in the array in use plus seventeen counts. The RC700A board's counter has an inherent count of one. In addition, the transfer timer adds an additional count of fourteen Therefore, to calculate the proper switch settings this formula must be used for minimum count:

$$
\left[\begin{array}{l}
2 \\
2
\end{array} \text { setting (selects element count.) }\right]+\text { Inherent count }+ \text { Transfer Timer }+2
$$

For example: The minimum count for a CCPD 1024 array would be achieved as follows: $\left[\right.$ S $2-10$ closed ( $2^{10}$ )] $+[$ Inherent count (1)] + Transfer Timer (14)] $+\left[\right.$ S1-1 closed $\left.\left(2^{l}\right)\right]=1041$. Larger numbers are permissable. (Note: The smallest count position $\left(2^{\circ}\right)$ is actually permanently shorted so that this switch position is immaterial.

## TABLE III

INTEGRATION COUNTER SWITCH SETTINGS (For Minimum Count)

CCPD
256
1024
1728

S1
4 closed ( $2^{4}$ )
4 closed ( $2^{4}$ )
$4,6,7$ closed $\left(2^{4}, 2^{6}, 2^{7}\right)$

S2
8 closed ( $2^{8}$ )
10 closed ( $2^{10}$ )
9,10 closed $\left(2^{9}, 2^{10}\right)$

The termination of the integration count starts the transfer counter, initiates transfer of stored array data into the CCD output register, and stops the fourphase clocks for the duration of the transfer period. The end of the transfer count then reinitiates the entire sequence above. A timing diagram showing the interrelationships is given in Figure 5.


Complete schematic diagrams are attached for the RC700A and the RC702A. Locatio of adjustment is shown in the assembly drawings. For power supply connections see the schematics and assembly drawings.

## Specifications

Inputs 1. Power +15 volts at approximately 150 mad d.c. -15 volts at approximately 70 ma d.c. +5 voits at approximately 600 ma d.c. This terminal is normally at ground (see test)

Outputs 1. Video Output Terminal:
a) Approximately 2.4 volts for saturated light output:
b) Output impedance $<10$ ohm
2. Scan Start Sync
3. Trigger Out

A negative TTL level pulse synchronous with the start of scan is provided for system updating.
A negative TTL-level pulse synchronous with the internal trigger generator.

## SETUP PROCEDURE

There are six controls: 1.) Clock Frequency Control R1, 2.) Dark Level Adjust R25, 3.) Sec Trimmer Capacitor Cl2, *4.) $V_{I C}$ Odd/Even Balance R16, *5.) $V_{\text {IG }}$ level R15, *6.) Gàin Balance Potentiometer R28.
*(Note: The items so marked are found on the RC702A array board.)
The light source must be uniform and must be perpendicular to the array, because any parallax will cause odd-even patterns and will interfere with the adjustment procedure.

Select the array and adjust the integration time with the long counter as described under "counters" in the Logic Control section.

When the array is selected, the corresponding clock line capacitors must be selected. The proper capacitance value for each array is listed in Table 1.

|  | SIZE <br> A | drawing no. $045-0042$ | C | SHEET | 7 OF |
| :---: | :---: | :---: | :---: | :---: | :---: |






Fig. 3 Control Sequence






## APPENDIX C

## LASER DIODE FOWER SUPPLY

## Laser diode

## Laser Diode Laboratory, Inc. Metuchen, New Jersey 08840 (201) 549-7700

Model
peak power
number of diodes
emitting area
peak forward current
duty factor
wavelength
spectral width
rise time

```
LD-167/0.04% duty factor
100-105 watts
5
16 x 16 mils
75 amps
    .04% (selected)
904 nm
3.5 nm
0.5 ms
200 ns
750
LDL-8
170
190
$175.00
```

max. pulse width
max. operating temp.
package type
beam spread 50\%
100\%
price

## Pulser

```
Power Technology, Inc.
Mablevale, Arkansas }7220
(501) 568-1995
```

Model \#
pulse current
pulse width
rise time
fall time
pulse rate (max.)
current input a
input voltage
size
price
IL75C

IL75C
40-80 amps
30-200 nsec
8-24 nsec
$10-28 \mathrm{nsec}$
10 KHz
300 ma
26-70 volts
$11 / 4 \times 21 / 2$
\$1,250.00

```
CAY110 Capsule Slip Ring
Airflyte Electronics Company
New Hook Road
Bayonne, New Jersey 07002
(201) 436-2230
Mechanical
```

1. rings
2. brushes
3. speed
4. lifetime
5. rotation

Electrical

1. current
2. voltage
3. resistance
4. 24 rings
gold alloy, 24
gold alloy spring tapered
500 rpm
$10^{8}$ revoluts.
bi directional

1 amp
150 VDC
5 million

```
Signals
1 laser fire
1 mirror motor drive
1 elevation encoder CW pulse
1 elevation encoder zero ref.
8 detector output
\(\overline{12}\)
```



Power Technology Incorporate<br>Address correspondence to: Box 4403 tittle Rock, Arkansas 72224<br>Plant location: 7925 Mabelvale Cutoff Mabelvale, Arkansas $72103 \quad$ 501-568.15

## ILC SERIES Laser Diode Pulsers


3) $40 \%$ to $110 \%$ of any soscified current between 1 and 9 amps

May axceedmaximum dury factor of commercialiy avaitable diodes Consult diode manufacturers specification sheat before opersting Aate specified is for 50 na pulse width. For 100 ns divide by 2 , for 200 ns divide by 4 . Pulse rate fimited to hele prevent darnage to laser diode
2k PPS or 200ns on ILIO0C only
and maximum puls width. Most applications require less than 25 mA . (External power supply madels anly)
Requires exteral clock for hiph rates.
$\mathrm{OR}_{R_{1}}$
OF


## Power Technology Incorporate

Address correspondence to: Box 4403 Litile Rock, Arkansas 72204
Pient location: 7925 Mabelvais Cutoff Mabeivale, Arkamas 72103

## Specifications Applicable to all Mode

Pulse Voltage - Suitable for one arode, 2 to 6 diode stacks on special order.
Pulse Current Stabilitv. $\pm \mathbf{1} / \mathbf{2 \%}$ for $\pm 10 \%$ battery voltage variation.
Puise Current Metering - 50 ohm coax connector. Amr phenol 27.9
0.1 voit per 1.04 for nominal pulse currents of one through 9 amps.
0.1 volt per 10A for nominal pulse currents of 10 through 40 mmps .
0.1 volt per 20A for nominal pulse currents of 41 through 100 amps .
Metering jack must be terminated in 50 ohm non-inductive resistor for correct pulse current reading. 51 ohm $5 \%$ composition resistor is generally suitable. Termination is necessary oniy while metering.
Temperature Compensation - Pulse current is reduced by approximately $10 \%$ for each $25^{\circ} \mathrm{C}$. below $27^{\circ} \mathrm{C}$.
Internal Clock Stability - $\pm 2 \%$
Internal Clock Programing - Internal ctock enabled by a jumper, clock rate set by one external resistor or potentionieter.
Internal Clock Limitations - Operation with battery voltage less than 9 volts may be unsatisfactory. Internal clock requires separate 9 to $\mathbf{2 8}$ voit clock power supply on pulsers designed for external power supply.
External Triggering - Terminals are provided. May be trig-
gered by TTL circuitry. Approximately 3 volt pulse 500 ohm required. Risotimg should be less than lus.
Sync. Output - Use External Trigger Leads. Isolate IK min.
Operating Voltage Range of Internal Power Suppiy - 6at low pulse rates, $14-24 \mathrm{~V}$ at maximum pulse rates, 9 V irnum for operation with internal clock.

Variable Pulse Width - Available as option. Consult fact Controls - Pulse current adjust potentiometer located jacent to connector.
Connectors . Power connector located on end opp diode. Mating connector provided for power. Pulse cur metering. Amphenol 27-9, located adjacent to p connector.
 $100 \mathrm{gr} ., 11 /{ }^{\prime \prime} \times 41 / 2^{\prime \prime}-125 \mathrm{gr}$.

WARNING - PERSONAL SAFETY PRECAUTIONS
Lasar Radiation. Diode lasers in operation emit invi sible infrared radiation which may be harmful to tt human eye. Appropriate safery precautions must $t$ taken to avoid the possibility of eye damage.
Electrical Shock. Operating voitages applied to som of these products present a shock hazard. Appropria precautions should be taken.


Power Supply Iñputs


Power Supply Outputs

LASER POWER SUPPLY




[^0]:    Users should follow IC Handling Procedures specified on pg. i-4

[^1]:    *(Called $\varnothing 1$ and $\varnothing 2$ respectively on the 702 A board.)

