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# NASA TECHNICAL MEMORANDUM 

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LOW DRAG ATTITUDE CONTROL FOR SKYLAB ORB! TAG LIFETIME EXTENSION

By John R. Glaese and Hans F. Kennel Systems Dynamics Laboratory

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In the fall of 1977 it was determined that Skylab had storted to tumble and th the original orbit lifetime predietions were much too optimistic. A decision had to be made whether to accept an early uncontrolled reentry with its inherent risks or try t attempt to control Skylab to a lower drag attitude in the hope that there was enough time do develop a Telcoperator Retrieval System, bring it up on the Space Shuttle anc then decide whether to boost Skylab to n higher longer life orbit or to reenter it in controlled fashion.

In the following the end-on-velocity (EOVV) control method is documented, whi was sucessfully applied for about half a year to keep Skylab in a low-drag attitude with the aid of the control moment gyros and a minimal expenditure of attitude contro gas.

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# LOW DRAG ATTITUDE CONTROL FOR SKYLAB ORBITAL LIFETIME EXIENSION 

INTRODUCIION


#### Abstract

On February 9, 1974, Skylab systems were confipured for a final power down and Skylab was denctivated in a passively stabilized pravity gradient (GG) attitude with the Mutiple Dockinf Adapter (MDA) up and the solar panels trailinf (Fig.1). Prediction of solar eycle 21 activity (the solar cycle predicted to bepin in 1977) indicated that this attitude would result in a potential storape period of 8 to 10 yr . However, in the fall of 1977 it was determined that Skylab had started to tumble randomly and was experiencing an inereased orbital decay rate. This was the result of the preater than predieted solar activity at the bopinning of solar cycle 21. This increased activity inereased the dirap forces on the vehiele. Skylab was now predicted to peenter the Earth's atmosphere in late 1978 or early 1979 unless something was done to roduce the drag forces actinp on it, It was necessary to make a decision to eikher accopt an early uncontrolled reentry (and with it the danger that relatively larpe pieces couid damape something or hurt somebody) or to attempt to actively control Skylab in a lower drag attitudo thereby extenclinp its orbital lifetime until a Space Shuttle mission could effect a boost or deorbit maneuver with Skylab.


In order to verify whet options could be accomplished with the onboard Skylab systems, in Murch 1978 a team of NASA engineers went to the Bermuda ground station to establish commonications. The resulting: data indicated no discemible depradation of the Skylab systems during its four years of orbital storage. The knowledge that Skylnb was in an unstable tumble prompted investigation into schemes wheh might extend the orbital lifetime of Skylab.

The first option investigated was to use the on-board thruster attitude control system (TACS) to maintain a quasistable tumble. However, it was soon determined that this option would not extend the lifetime sufficiently to correspond to the operational readiness of the Space Shuttle for a possible reboost or deorbit mission. The only alternative was to reactivate and continuously controi the Skylab in a minimum drap: attitude. In order to accomplish this, the low-draf, end-on-velocityvector (EOVV) attitude control scheme was developed in record time (the authors wore piven the task on Mnreh 20 and the scheme was flown on July 11). Skylab remained in the low-drag attitude until Jamuary 25, 1979, when the vehicle was commancied to its oripinal desipn attitude [solar inertial (SI)], which was a high-drag; attitude (See Skylab EOVV Time table for futher details).


Figure 1. Skylab.
With the active control of Skylab in the low-drag attitude, it was decided to accelerate the development of an orbital retrieval system (leleopertor Retrieval System) that might be accommodated on an early flight of the Space Shuttle, thus incroasing the chances of rendezvousing with Skylab. The rate of orbital decay, however, continued to increase due to the increased solar activity. Skylao's on-board systems also showed signs of deterioration, and there were stronf indications that the schedule of the Space Shuttle would slip. For these reasons, the effort for a Skylab recovery was terminated in December 1978, and Skylab was placed in the SI attitude in January 1979. In this attitude the ground maintenance was minimum and efforts could be concentrated on a method for a conrolled reentry of Skylab [1].

More detail about the reactivation mission can be found in Reference 2. The following section gives the development of the EOVV control scheme.

## ATTITUDE AND POINTING CONTROL SYSTEM (APCS)

The control of the Skylab attitude to the attitude reference was done exactly as in the original mission [3]. However, only the pointing control system (RCS) of the APCS was used; the experiment pointing control system (EPCS) was disabled.

The major parts of the APCS were the rate gyros, the Acquisition Sun Sensors (ACO SS), the Star Tracker (it had failed during the

## SKYLAB EOVV TIMETABLE

3/20/78 Authors made auare that a low-tras attitude momentum manarement method was needed.
4/ 7/73 Concent review by MSFC middle manarement.
4/26/78 Final equations for 2-CMg sovv to IBM.
5/22/78 Documentation on ATMLC software chance requirements for 2-CMG EOVV oneration completed by IBM.
6/8/78 CMO snin up in cared position (7:00 am CST, cafed to $H+\left[\begin{array}{llll}1.92 & 0.03 & 0.361)\end{array}\right.$
6/9/78 7 am CDT. fotation about sunline to place $x$ axis IUf.8 am CDT. CNG control in SI attitude.
12 moon CDT. Loss of control due to inadverdent servopower cut-off to CMill3 (fallity switoh selectorintroduces additional command)1 pm CDT. Rerained control with CMG's after a orbits.
6/11/78 8:27 UT. Entering EOWV A attitudc.
6/28/78 Loss of BOVV attitude control due to larpe anfular momentum and cmi saturation.
7/5/78 Re-estabilishment of EJVV A attitude.
7/9/78 Loss of all nower introduces loss of attitude control.
7/19/78 Reorientation maneuver to find attitude.
7/25/78 feeestablishment of EOV A attitude.
11/4/78 Maneuver from EOVV A to EOVV E at.titude
12/:0/78 NASA EO Dress aenference (J.Yardley/ri.Aller) whereSkylab is teclared unsavarle.
1/25/79 Keorientation maneuver to SI attitude (hold mode toallow time for the desion of the torque equilibriumcontrol methad for reentry).
oripinal mission), the Apollo Telescope Mount Digital Computer (ATMDC), the Workshop Comouter Interface Unit (WCIU), three double-gimbaled Control Moment Gyros (CMGs), and cold-cras (comprossed nitrogen) Thruster Attitude Comtrol System (TACS).

Six control modes were addressable: (1) STANDBY, (2) SOLAR INERTIAL (SI), (3) EXPERIMENT POINTING, (4) ATIITUDE HOLD/CMG, (5) ATIIT UDE HOLD/TACS, (6) ZLV (for $z$ axis along the local vertical), EOVV control was programmed to be a substate of the ZL' ${ }^{\mathbf{V}}$ mode. The basic ZLV attitude was with the positive $z$ axis along the ldeal vertical, pointing up, and the positive $x$ axis in the orbital plane, pointing in the direction of the velocity vector. Any anpular offset form the basic ZLLV attitude (offset identified by the quaternion qal) could be commanded by a set of three Euler anfles ( $x$ 's) with a $y, z, x$, rotation sequence. None of the original APCS capabilities were eliminated by the addition of the EOVV control method.

## CMC CONTROL SYSTEM

The CMG control system was composed of three orthogonally moynted, double gimbaled CMGS with angular momentum magnitude $H$ of 3050 Nms ( $2280 \mathrm{ft}-\mathrm{lb}-\mathrm{sec}$ ) as shown in Figure 2. The CMG control law utilized three normalized torque commands and the CMG momentum status to generate the proper CMG gimbal rate commands [4]. The (MG control law consisted of three parts: CMG steering law, rotation law, and pimbal stop avoidance logic. There also were some other routines for specialized situations like caging the CMGs to a desired momentum state [5].

The CMG control law had the cap bility to operate with either three or two CMG's for redundancy. Since $1, \mathrm{MG}$ No. 1 hat failed during the original Skylab mission the CMG control law was always in the two-CMG option.

## EOVV MOMENTUM CONTROL

The problem for EOVV was to determine variable reference attitudes such that, on the averape, the angular momentum was contained within the two-CMG capability (allowing the CMGs to hold the prescribed attitude reference), while the average reference attitude was consistent with the desired low aerodynamic drag.

Minimizing the drag on Skylab required that the least frontal area was presented to the wind while at the same time holding a GG torque equilibrium (at the altitudes of concern the GG torques were still very dominant and they were therefore used exclusively for momentum control;


Figure 2. CMG mounting arrangement.
for details on the aerodynamic torques, [1]). Keeping the minimum principal moment of inertia axis parallel to the wind direction fulfilled this requirement. To have the necessary electrical power from the solar cells as well as strapdown update information from the ACQ SS, the Skylab was rolled through the angle $n_{x n}$ (APPENDIX F) about the minimum principal moment-of-inertia axis (principal $x$ axis) such that once per orbit the sun line passed nominnlly through the center of the ACQ SS.

There were two attitudes which satisfied these requirements: One with the MDA forward (EOVV A) and one with the MDA backward (EOVV B). In either case, the MDA had to be pitched down by a varying amount (depending on the solar elevation angle $\beta=-\eta_{x}$ above the orbital plane) to align the principal $x$ axis with the orbit tangent.

Control of the angular momentum was split into the control of the momantum component perpendicular to the mrbital plane (POP control) and the component in the orbital plane (?JP acrirol), Since IOP control had some effect on POP control, IOP control in ifsated first.

## IOP Momentum Control

When Skylab was originally designed, it was desirable to minimize the GG torques about the minimum principal-moment-of-inertia axis as much as possible since the momentum management scheme [6] was least efficient about this axis. For EOVV control tias meant that there were basically no large GG torques available about this axis and, furthermore, there would be no change in torques when the principal y or $z$ axes were $\pm 45$ degrees from the orbital plane (tracking of the sun by rolling about the x axis would make this a frequent oceurrence). Therefore this first order effect had to bo abandoned. The actually used momenturn control scheme assumed that the Skylab moments-of-incrtia were cyolindrical, with an average moment-of-inertia difference of AI .

The problem was soived by using a second order effect. First, a large cyclic pop torque was generated by "nodding" (pitehing) the Skylab in the orbital plane. Orelie nodding was required to avoid a continuous momentum build-up in the PO' direction, 'The cyclic POP torque was the tilted as required (diftasently for each half cyele or quarter orbit) to generate a controllable component in the orbital plane. The second order effect stems from the fact that the IOP torque is proportional to the nodding angle times the tilting angle.

The effectiveness of the IOP momentum control did not depend on the frequency of the nodding, However, other considerations entered: (1) the lower the frequency, the larger the por momentum swing, and (2) the higher the frequency, the larger the maneuver momentum that has to be exchanged botween the vehicle and the CMG system. Since only a limited momentum volume was available with two CMGs and their associated gimbal stop problems, a nodding frequency of twice orbital frequency was chosen as a viable compromise ( $s \equiv \sin$ )

$$
\begin{equation*}
\eta_{y n}=-\eta_{y m} s\left(2 \Omega_{0} t-\eta_{0}\right) \tag{1}
\end{equation*}
$$

Therefore, the IOP momentum control calculations were done every quarter orbit. This had the added advantage that the resolution from the nearly inertial 0 sy/stem to the rotating $L$ system happened in 90 degree intervals allowing indexing of some of the saved momentum samples rather than requiring a full-fledged resolution.

To minimize transients at the quarter orbit sample points, the tiling angle was also sinusoidal with twice orbital frequency and its amplitude was the only changing quantity:

$$
\begin{equation*}
n_{z n}=-n_{z m} s\left(2 \Omega_{0} t-n_{0}\right) \tag{2}
\end{equation*}
$$

where

$$
\begin{aligned}
& n_{z m}=K_{1 \mid z}\left(e_{z o}-e_{a}+e_{\mathbf{r}} / 4\right) \\
& K_{n z}=(15 H) /\left(48 \eta_{y m} \Omega_{o} \Lambda I\right) \\
& e_{z o}=\text { momentum component to be desaturated } \\
& e_{\mathbf{a}}=\text { amplitude } \\
& e_{\mathbf{r}}=\text { ramp per orbit } .
\end{aligned}
$$

The phasing of the nodding and tilting angles was the same so that the amplitude was reached half-way batween sample points.

The sample points were chosen so that one of the samples (sample 2) occurred at the time when the sun was perpendicular to the solar panels. This happened before orbital noon for EOVV A and after noon for EOVV B (the difference of about 11 degrees between the geometric and the principal $x$ axes is the reason). In addition the nodding rate, being added to the orbital rate, was phased so that it slowed Skylab down when the solar panels were perpendicular to the sun and therefore maximized the power from the solar panels [giving lise to the minus signs in equation (1) and equation (2)].

The tilting angle amplitude was calculated so that the IOP momentum component, which could be affected during the next quarter orbit (it was along the direction of the connecting line between the present sample point and the next one), would be driven to the desired value. The desired value was basically zero, but any constant torque in the $x$ system caused only a cyclic momentum change (with normalized ampitude $e_{a}$ ) over one orbit and should not be compensated for. Therefore, the momentum attributable to a constant $L$ system torque was subtracted out of the momentum $e_{z o}$ to be desaturated over the next quarter orbit.

To recognize a cyclic as well as a ramp momentum change, four past momentum samples were saved. The samples were also used to generate strapdown update information once an orbit (flow charts in APPENDIX H).

POP Momentum Control
The torques associated with a rotation about the orbit normal are much stronger than the ones associnted with IOP control, Hence, the momentum sampling for POP control has to be done as frequently as pos* sible. However, the transients should have had a chance to settle before the next POP sample is taken. Twelve samples per orbit satisfied both requirements. The desaturation gain faPPENDIX D.b)

$$
\begin{equation*}
K_{y o}=H /\left(3 \rho_{0}^{2} \Delta I T_{\mathrm{des}}\right) \tag{3}
\end{equation*}
$$

is calculated so that a step attitude change of

$$
\begin{equation*}
\eta_{\mathrm{yc}}=\mathrm{K}_{\mathrm{yc}} \Delta \mathrm{o}_{\mathrm{y}} \tag{4}
\end{equation*}
$$

desaturates the desired $\Delta e_{y}$ in one desaturation interval, To further reduce transients, the calculated POP angle (which would have eliminated the momentum offset within the next desaturation interval if the angle were applied fully during the interval) was ramped in so that the angle was nehieved at the end of the interval. Since this only reduced the momentum offset by half, the angle was ramped-out during the following interval for a full momentum offset elimination. The ramp due to the newly calculated POP angle was simply superimposed on the ramp-down from the provious POP angle (Fig. 3). The attitude command is, therefore, given by ( N signifies the present and $\mathrm{N}-1$ the past value)

$$
\begin{align*}
\left(\eta_{y c}\right)_{N} & =K_{y c}\left[\left(\Delta e_{y}\right)_{N}-0.5\left(\Delta e_{y}\right)_{N-1}\right]  \tag{5}\\
& =K_{y c}\left(\Delta e_{y}\right)_{N}-0.5\left(\eta_{y c}\right)_{N-1}
\end{align*}
$$

The ramps connecting the ${ }^{7} \mathrm{yc}^{\text {'s }}$ are generated by


Figure 3. POP command superposition.

$$
\begin{equation*}
\left(\eta_{\mathbf{y}}\right)_{\mathbf{N}}=\left(\eta_{\mathbf{y}}\right)_{\mathbf{N}-1}+\Delta \eta_{\mathbf{y}} \Delta \mathbf{t}, \tag{6}
\end{equation*}
$$

where

$$
\Delta \eta_{y}=\left[\left(\eta_{y c}\right)_{N}-\left(\eta_{y c}\right)_{N-1}\right] / \mathrm{I}_{\mathrm{des}}
$$

This method resulted in a constant hang-off when necessary: The angle change due to the old angle being ramped-out was compensated by the ramp-in of the new angle (in flight, constant hang-offs were common due to strapdown and navigation errors and they were not detrimental, since the momentum control kept the vehicle at the truely desired attitude). A block diagram of the EOVV orbital y momentum control scheme is shown in Figure 4.

## EOVV STRAPDOWN UPDATE

Strapdown updates about the vehicle x and y axes were always furnished by the $A C Q S S$. To do that, the roll angle about the principal $\mathbf{x}$ axis was changed by large amounts to compensate for the large beta angle changes (a slow change) and relatively fast smaller corrections were applied to compensate for the nodding and the tisting angles. The overall effect was that the vehicle $z$ axis nominally traced a cone about orbital north. The difference between where the sun-line was at the closest approach to the ACQ SS center and where it was supposed to be according to the strapdown information gave the strapdown $x$ and $y$ information.

Since there was no other sensor available, it was more difficult to gain strapdown update information about the sun-line. The selected momentum control method, fortunately, had the feature that, due to the nodding angle, a misaligment between the ideal orbital plane and the indicated orbital plane generated an IOP momentum ramp.

The actual strapdown update was done by changing the reference quaterion (APPENDIX H.6)

$$
\begin{equation*}
\mathbf{Q}_{\mathbf{V I}}=\overline{\overline{\overline{\Delta Q}}} \mathbf{Q}_{\mathrm{VI}}, \tag{7}
\end{equation*}
$$

where the double-bar operator is defined in APPENDIX E and

$$
\begin{equation*}
\Delta Q=\left[\Delta Q_{1}, \Delta n_{2}, \Delta Q_{3}, \Delta Q_{4}\right]=\left[\Delta Q, \Delta Q_{4}\right], \tag{8}
\end{equation*}
$$



Figure 4. Orbital X -monentum control.
with

$$
\begin{equation*}
\Delta Q=0.5\left(\underline{s} \times \underline{v}+\mu_{2} \underline{s}\right) \tag{0}
\end{equation*}
$$

and

$$
\begin{equation*}
\Delta Q_{4}=\sqrt{1-\Delta Q \cdot \Delta Q} . \tag{10}
\end{equation*}
$$

The cross product in equation (9) is the $A C Q S S$ update and the last term is the IOP ramp update, where $v$ is a unit vector in the sun direction as calculated by the ATMDC and $\underline{s}$ is the measured sun direction unit vector. $\mu_{z}$ is the angle about the measured sun direction (the gain $K_{\mu z}$ is developed in APPENDIX D.e.):

$$
\begin{equation*}
\mu_{z}=-K_{\mu_{Z}}\left(e_{T L N 1} * e_{T L N 1 P}\right) \tag{11}
\end{equation*}
$$

where $K_{\mu z}$ is a gain and ( $e_{\text {TLN } 1}-e_{T_{\mu N 1 P}}$ ) is the angular momentum change (ramp) per orbit ( $\mu_{z}$ is calculated at sample point 1 ) modified by the ground commanded ramp bias crb since

$$
\begin{equation*}
\mathbf{e}_{\text {TLN } 1 P}=e_{\text {TLN } 1 P}+c_{\mathrm{rb}} \tag{12}
\end{equation*}
$$

is calculated right after the $\mu_{z}$ ealculation and, therefore, is used for the next $\mu_{z}$ calculation ( $\mathrm{e}_{\text {ILNNIp }}$ is modified at every sample point to account for the momentum changes commanded by $\eta_{z m}$ ).

## EOVV OPERATION AND PERFORMANCE

The original EOVV equation considered EOVV A only. In EOVV A CMG No. 2 received more solar radiation when the sun was north of the orbital plane (positive beta angle) and less when it was south of the orbital plane (negative beta angle). For large negative beta angles the CMG No. 2 bearing temperatures became critically low. As a consequence, the EOVV A equations had to be modified during EOVV operation to allow an EOVV B attitude during extended perinds of large negative beta angles (Fig, 5),


Figure 5. Sun elevation angle beta (April 1978/August 1979).
Constant $\eta_{y c}$ angle hang-offs (caused by strapdown, navigation, and other errors) required constant POP momentum hang offs to generate the necessary commands. Since the range of acceptable POP momentum component was rather limited ( $\pm 0.4 \mathrm{H}$ from the nominal; the nominal $\mathrm{e}_{\mathrm{TN}}$ being 2.5 H in EOVV A and 0.3 H in EOVV B) the nominal POP momentum had to be changed to accept large angle hang-offs ( 3 deg of POP angle hang-off required 0.1 H POP momentum hang-off). This change in nominal momentum was made from the ground at the beginning of the EOVV operation and later was automated on-board to guard against ground inattention, ground system failures, and long telemetry coverage gaps (Fig. 6.). Momentum excursions outside the specified range caused loss of attitude due to CMG saturation on one occasion June 28, 1978, and it was therefore very important to keep the POP momentum bounded.

Strapdown updating about the sun-line was done with information derived from the IOP momentum ramp. Unfortunately, the evaluation of the IOP momentum ramp yielded very noisy readings from one orbit to the next and could only be used with a very reduced gain $K_{\mu_{2}}$. This,

Figure 6. Orbital Y-momentum
in turn, could lead to large orbit plane misalignments to generate the required strapdown updates to keep up with the $\pm 50$ deg rocking of the true orbital plane (due to the precession when viewed with respect to the projection of the sun-line into the orbital plane). The large size of the maximum change per orbit ( 1.2 deg ) was not recognized at the beginning of the EOVV operation and was the cause for loss of attitude on June 28, 1978. After that the ideal strapdown update necessary to follow the rocking of the orbital plane was introduced open loop through the quantity called $e_{r b}$ [equation. (12)] and no more problems were experienced (APPENDIX C).

## APPENDIX A <br> COORDINATE SYSTEMS AND TRANSFORMATIONS

The coordinate systems which are pertinent to Skylab EOVV control are defined here. Each system has some special geometrical or physical feature which simplifies the solution of a particular problem.

The following coordinate systems are described: Principal, Orbital, Vehicle, Attitude Reference, Solar Inertial, and Z-Local Vortical. Each coordinate system consists of a set of mutually orthogonal axes exhibiting right-handedness.

An inertial (with respect to rotation only) coordinate system is a system which retains its orientation with respect to the celestial sphere, although the origin may be moving along any general curvilinear path in space. Similarly, a vehicle fixed system retains its onientation with respect to the vehicle.

## Orbital Coordinate System (O)

The Orhital Coordinate System ( $x_{0}, y_{0}, z_{0}$ ) is a precessing coordinate system with its origin at the Earth eenter of mass. The rate of precession about the Earth's north pole is approximately -5 degrees/day. The $z_{0}$ axis lies in the orbital plane, positive through the ascending node of the orbit. The $x_{0}$ axis also lies in the orbital plane 90 degrees ahead of the $z_{0}$ axis. Since the Skylab orbit was in the $x_{0} z_{0}$ plane at all times, the $y_{0}$ axis was parallel to the orbital angular momentum vector, completing the right-handed system (Fig. A-1).

Solar Inertial Coordinate System (I)
The Solar Inertial Coordinate System ( $x_{I}, y_{I}, z_{I}$ ) is only a pseudoinertiul system since it makes one revolution per year. It was used during the Skylab mission to point the instruments in the desired direction. The origin is coincident with the origin of the Vehicle Coordinate System origin. The $z_{I}$ axis is positive toward the center of the Sun. The $x_{I}$ axis lies at an angle $\nu_{z}$ from the orbital plane (this angle is calculated on-board such that the principal $x$ axis is in the orbital plane to minimize che build-up of angular momentum) and is positive toward the sunset terminavor.


Figure A-1. Coordinate system.

Z-Local Vertical Coordinate System (L)
The z-Local Vertical Coordinate System ( $\mathrm{x}_{\mathrm{L}}, \mathrm{y}_{\mathrm{L}}, \mathrm{z}_{\mathrm{L}}$ ) is a rotating system with its origin at the center of mass of Skylab (the rate of rotation is one revolution per orbit). The $X_{L}$ axis is positive in the direction of flight and lies in the orbital plane. The ${ }^{2}{ }_{L}$ axis is parallel to the local vertical direction and is positive outward, away from the Earth. The $y_{L}$ axis is parallel to the orbit normal and is positive toward orbital North.

## Vehicle Coordinate System (V)

The Vehicle Coordinate System ( $x_{V}, y_{V}, z_{V}$ ) is a vehicle-fixed system with its origin at the center of mass. The $X_{V}$ axis lies along the long axis of Skylab and is positive in the direction of the Multiple Docking Adapter (MDA). The $z_{V}$ axis is positive toward the Apollo Telescope Mount (ATM) and the $y_{V}$ axis completes the right-handed system.

## Yrincipal Axes Coordinate System (P)

The Principal Axes Coordinate System $\left(x_{p}, y_{p}, z_{p}\right)$ is a vehiclefixed system with its origin at the center of mass. The axes are along the principal moment-of-inertia axes, labeled such that the eigen angle between the $V$ and the $P$ system is minimized.

The following transforms are useful (the subscripts $1,2,3$ indicate rotation about $x, y, z$, respectively).

$$
\begin{aligned}
& {[L I]=\left[\Delta \eta_{t y}\right]_{2}\left[-\eta_{x}\right]_{1}\left[\cdot v_{z E}\right]_{3}} \\
& {[A L]=[K]^{T}\left[\eta_{1}\left[\eta_{z}\right]_{3}\left[\eta_{y}+\eta_{c y}\right]_{2}\right.} \\
& {[P V]=[K]}
\end{aligned}
$$

[VI] $=$ strapdown matrix (updated by $\underline{\underline{4}} \mathrm{~V}$, sunsensor data and momentum data)

## APRENDIX B

## ARAVI'IATIONAL TORQUE MODEL

The gravitational torques on a satellite produced by a large, spherical primary body are important contributors to its rotational dynamics. The force on a point mass $m$ exerted by the primary $M$ is

$$
\begin{equation*}
{\underset{i}{i}}^{F_{i}}-\frac{G M m}{\left|R_{i}\right|^{3}} \underline{R}_{i} ; G: 6.672 E-11 \mathrm{Nm}^{2} / \mathrm{kg}^{2} . \tag{B1}
\end{equation*}
$$

The satellite can be viewed as a collection of point masses. The net torque on this collention of masses about the origin of satellite coordinates is

$$
\begin{equation*}
{\underset{T}{G}}=-\sum_{i} \underline{x}_{i} \times \underline{E}_{i} \tag{D2}
\end{equation*}
$$

The vector ${\underset{m}{i}}$ is the position of $m_{i}$ relative to the origin which is at ${\underset{\sim}{0}}^{\mathbf{R}}$ relative to the primary center. Thus

$$
\begin{equation*}
\underline{\mathbf{R}}_{\mathrm{i}}=\underline{\mathbf{R}}_{\mathrm{o}}+\underline{\mathbf{r}}_{\mathrm{i}} \tag{B3}
\end{equation*}
$$

and

$$
\begin{equation*}
\underline{T}_{\mathrm{g}}=-\sum_{i} \underline{\underline{r}}_{\mathrm{i}} \times \mathrm{GMm}_{\mathrm{i}} \frac{\underline{R}_{0}+\underline{\underline{r}}_{\mathrm{i}}}{\left|\underline{R}_{0}+\underline{\mathbf{r}}_{\mathrm{i}}\right|^{3}} \tag{B4}
\end{equation*}
$$

In genexal $\left|\underline{\underline{r}}_{\mathbf{i}}\right| \ll\left|\underline{\underline{R}}_{0}\right|$ and hence an expansion of $\underline{T}_{\mathrm{g}}$ keeping only low order terms becomes sufficient for most purposes. Now

$$
\begin{equation*}
\frac{\underline{R}_{o}+\underline{r}_{i}}{\left|R_{0}+r_{i}\right|^{3}}=\frac{\underline{R}_{0}}{\left|R_{0}\right|^{3}}+\frac{\underline{n}_{i}}{\left|R_{0}\right|^{3}}-3 \underline{R}_{0} \frac{\underline{R}_{0} \cdot \underline{r}_{i}}{\left|R_{0}\right|^{3}}+\text { h.o.t. } \tag{B5}
\end{equation*}
$$

Using the expansion of equation (B5) in (B4),

$$
\begin{align*}
\underline{\underline{r}}_{\mathrm{g}} & =-\underline{\mathbf{r}}_{\mathrm{cm}} \times \frac{\mathbf{G M m}}{\left|\underline{\underline{R}}_{0}\right|^{3}} \underline{\underline{R}}_{0}+3 \frac{G M}{\left|\underline{\underline{R}}_{0}\right|^{3}} \sum_{i} m_{i} \underline{\underline{r}}_{i} \times \underline{\underline{R}}_{0} \frac{\underline{\underline{R}}_{o} \cdot \underline{\underline{r}}_{i}}{\left|\mathbf{R}_{0}\right|^{3}} \\
& + \text { h.o.t. } \tag{B6}
\end{align*}
$$

We have used the definition $m_{m}{\underset{m}{c m}}=\sum_{i} m_{i}{\underset{i}{i}}, \quad$ Rearranging and grouping equation (B6) we sbtain

$$
\begin{align*}
& \text { +h.o.t. . } \tag{B7}
\end{align*}
$$

The term in parentheses in equation (B7) occurs in the definition of the moment of inertia dyadic (or tensor)

$$
\begin{equation*}
\underline{\underline{1}}=\sum_{i} m_{i}\left(\underline{r}_{i}^{2} \underset{=}{\underline{1}}-\underline{\underline{r}}_{i} \underline{x}_{i}\right) \tag{B8}
\end{equation*}
$$

Using this definition in equation (B7) and dropping the higher order terms yields the gravity gradient torque expression

$$
\begin{equation*}
\underline{T}_{\mathrm{gg}}=-\underline{\mathrm{r}}_{\mathrm{cm}} \times \frac{\mathrm{GMm}}{\left|\underline{R}_{0}\right|^{3}} \underline{R}_{0}+3 \frac{\mathrm{GM}}{\left|\underline{R}_{0}\right|^{3}} \frac{\mathrm{R}_{0}}{\left|\mathrm{R}_{0}\right|} \times \frac{\mathrm{I}}{=} \cdot \frac{\underline{R}_{0}}{\left|\underline{R}_{0}\right|} \tag{B9}
\end{equation*}
$$

As can be seen, if the origin is positioned at the center of mass the more familiar gravity gradient torque expression results:

$$
\begin{equation*}
\underline{\underline{T}}_{\mathrm{gg}}=3 \frac{\mathrm{GM}}{\left|\underline{R}_{0}\right|^{3}} \frac{\underline{R}_{0}}{\left|\underline{R}_{0}\right|} \times \underline{\underline{I}} \cdot \frac{\underline{R}_{0}}{\left|\underline{R}_{0}\right|} \tag{B10}
\end{equation*}
$$

For an orbiting body $m$, the orbital angular velocity magnitude is given by

$$
\begin{equation*}
\delta_{0}^{2}=G M /\left|\underline{R}_{0}\right|^{3} \tag{B11}
\end{equation*}
$$

Noting that $\underline{\underline{R}}_{0} / \underline{\underline{R}}_{0} \mid$ is a unit vector $\underline{u}_{R}$, we finally write

$$
\begin{equation*}
\underline{T}_{g g}=3 \Omega_{0}^{2} \underline{u}_{R} \times \underline{v} \cdot \underline{u}_{R} \tag{B12}
\end{equation*}
$$

The torque of equation (B12) is what is commonly referred to as the gravity gradient torque.

## APPENDIX C. <br> STRAPDOWN DRIFT ERROR AND MOTION OF I COORDINATE SYSTEM

The basic reference coordinate system used during the original Skylab manned mission was the so-called Solar Inertial (I) coordinate bystem (Appendix A gives the definition). As stated there, I is not truly an inertial coordinate system since its axes rotate with the sun and orbit regression. Its rotation rate varies with the solar elevation angle out of the orbit plane. By definition, the z axis points toward the sun and the unit vector $\underline{u}_{p 1}$ lies in the orbit plane pointed generally parallel to the vahicle velocity vector at orbital noon. The unit vector $\underline{u}_{p_{1}}$ is the direction of the $\times$ puincipal axis in vehicle coordinates $V$. Thus, when the $V$ and I systems are aligned, the vehicle $x$ principal axis is in the orbit plane so that ideally all gravity gradient torques are cyclic. The sun angle $\Gamma_{y}$ and orbit regression angle $\lambda_{y}$ are updated once per orbit at orbit midnight in the Skylab navigation calculations. In between midnights, the I system, as defined by vehicle on-board navigation, does not rotate. As a consequence of this, it is useful to redefine I such that the definition given previously is only satisfied at midnight. Let $I_{k}$ represent the true reforence position (we use the convention that true or physical parameters are labelled by a subscript $k$ ) of this system and I represents the estimate. This estimate is made using the vehicle angular rates as measured by the rate gyros.

$$
\underline{\omega}_{\mathrm{VI}}^{\mathbf{v}}=\underline{\omega}_{\mathrm{Vrk}}{ }^{\mathbf{V}}+\underline{\omega}_{\mathrm{D}}^{\mathrm{V}} .
$$

The angular rates $\underset{\sim}{\omega} V^{V},{ }^{\underline{\omega}} \mathrm{D}$ and $\underset{\sim}{\omega}$ VIk are all given in the vehicle system $V$. The rate $\underline{U}_{\mathrm{D}} \mathrm{V}$ is tl gyro drift rate and experience indicates it tends to be constant in the vehicle system $V$. To a good approximation the vehicle remains fixed relative to the local vertical system L. Thus

$$
\underline{\omega}_{\mathrm{VI}}{ }^{V}=\Omega_{\mathrm{o}} \underline{u}_{\mathrm{L} 2}{ }^{\mathbf{V}},
$$

where $\Omega_{0}$ is the orbital angular rate.

Since $\Omega_{0}, u_{L 2}{ }^{V}$ and $\underline{\omega}_{D} \mathbf{V}$ are constants, $\underline{\omega}_{V I k}^{V}$ must also be constant

$$
\underline{\omega}_{\text {VIL }}{ }^{V}=\Omega_{0} \underline{u}_{L 2}{ }^{V}-\underline{\omega}_{D}^{V} .
$$

We can integrate the rate $\underline{\omega}_{\text {oIk }}{ }^{V}$ using the differential equation

$$
\frac{\mathrm{d}}{\mathrm{dt}}\left[\mathrm{VI}_{k}\right]=-\tilde{\omega}^{\mathrm{VIk}}{ }^{\mathrm{V}}\left[\mathrm{VI}_{\mathrm{k}}\right] .
$$

This yields

$$
\left[\mathrm{VI}_{k}\right]=\exp \tilde{A}\left[\mathrm{VI}_{k}\right]_{o}
$$

with

$$
\exp \tilde{A}=\sum_{n=0}^{\infty} \frac{1}{n!} \tilde{A}^{n}=1-\sin |\underline{A}| \frac{\tilde{A}}{|\underline{A}|}+(1-\cos |\underline{A}|) \frac{\tilde{A} \tilde{A}}{|\underline{A}|^{2}},
$$

and

$$
\underline{A}=-\underline{\omega}_{\text {VIk }}{ }^{V}
$$

where

$$
\mathbf{T}=2 \pi / \Omega_{0},
$$

is the orbital period.

Let $\left[\mathrm{VI}_{\mathbf{k}}\right]_{0}$ is the $\left[\mathrm{VI}_{\mathbf{k}}\right]$ transformation at the beginning of the orbit.

$$
\underline{\theta}_{V I_{k}}=\underline{U}_{V I_{k}} \mathbf{T},
$$

and

$$
\underline{\theta}_{\mathrm{D}}=\underline{\omega}_{\mathrm{D}} T,
$$

then

$$
\underline{\theta}_{\mathrm{VI}_{\mathrm{k}}}=2 \pi \underline{u}_{\mathrm{L}_{2}}-\underline{\theta}_{\mathrm{D}}
$$

We shall now assume that $\left|0_{\mathrm{D}}\right| \ll 2 \pi$. In practice $\left.\right|_{0} \mid$ is at most $\therefore 0.05$ to 0.06 radians. Thus we can evaluate $\exp \left(-\theta_{V} V_{k}\right)$ keeping only 1 st order toms in ${ }^{1}{ }_{\mathrm{D}} \mathrm{V}$. First, wo find

$$
\underline{\theta}_{\mathrm{VI}_{\mathbf{k}}} \mid \times 2 \pi-\underline{\mathrm{u}}_{\mathrm{L}, 2} \cdot \underline{\underline{\theta}}_{\mathrm{D}}
$$

Using this result, wo find

$$
\frac{\underline{\theta}_{\mathrm{VI}_{\mathbf{k}}}}{\left|\underline{\theta}_{\mathrm{VI}_{\mathbf{k}}}\right|}=\underline{u}_{\mathrm{L}_{2} 2}-\frac{1}{2 \pi} \quad\left(\theta_{\mathrm{D}}-\underline{\theta}_{\mathrm{D}} \cdot \underline{\mathrm{u}}_{\mathrm{L} 2} \underline{u}_{\mathrm{L} 2}\right)
$$

From this we find

$$
\exp \left(-\underline{\theta}_{V_{k}}{ }_{k}\right)=\exp \left(\underline{u}_{L 2} \cdot \underline{\theta}_{\mathrm{D}} \underline{\underline{u}}_{\mathrm{L} 2}{ }^{V}\right)
$$

Finally

$$
\left[\mathrm{VI}_{\mathrm{k}}\right]=\exp \left(\underline{\theta}_{\mathrm{D}} \cdot \underline{u}_{\mathrm{L} 2} \ddot{u}_{\mathrm{L} 2}{ }^{\mathrm{v}}\right)\left[\mathrm{VI}_{\mathrm{k}}\right]_{0} .
$$

$$
\begin{align*}
& =\left[-\underline{0}_{\mathrm{D}} \cdot \underline{\mathrm{u}}_{\mathrm{L} 2}\right]_{\underline{u}_{\mathrm{L} 2}} \mathrm{~V}\left[\mathrm{VI}_{\mathrm{k}}\right]_{\mathrm{o}}  \tag{C1}\\
& { }^{N}\left(1+\underline{0}_{D} \cdot \underline{u}_{L 2} \tilde{u}_{L 2}{ }^{V}\right)\left[V I_{k}\right]_{o}
\end{align*}
$$

The result we have obtained in equation (C1) can be summarized by saying the strapdown error induced by gyro drift to the 1st approximation only accumulates along the orbit normal and tends to be purely cyclic along $\underline{u}_{\mathrm{L} 1}$ and $\underline{u}_{\mathrm{L}} 3^{\circ}$

Strapdown errors enter in also due to motion of the I coordinate system. These motions are not accounted for in the strapdown calculations and so must be correcind by updates from sun sensor data and momentum accumulation data as explained elsewhere. The solar inertial system motion is readily calculated from relative motion of the sun and from regression of the Skylab orbit. The sun moves approximately 1 leg per day while the orbit processes nearly 5.5 deg per day in the redrograde direction. Since there are about 16 orbits per day, these effects are usually small but occasienaify amount to more than 1 deg per orbit. To analyze these effects let us define the system $S$ such that the $z$ axis points to the Vernal Equinox and $y$ axis to the ecliptic north pole,

$$
[S E]=\left[\phi_{z}\right]_{3}=\left[\begin{array}{ccc}
\cos \phi_{z} & \sin \phi_{z} & 0 \\
-\sin \phi_{z} & \cos \phi_{z} & 0 \\
0 & 0 & 1
\end{array}\right] ; \phi_{z}=23.45^{\circ}
$$

Also, define $R$ such that $z_{R}$ points to the sun and $y_{R}$ to ecliptic north. From $R$ to $I$ is a rotation about $z_{R}$

$$
\begin{aligned}
{[\mathrm{IR}] } & =\left[\theta_{z}\right]_{3} \\
+[\mathrm{IE}] & =\left[\theta_{z}\right]_{3}\left[\Gamma_{y}\right]_{2}\left[\phi_{z}\right]_{3},
\end{aligned}
$$

where the numeric subscripts indicate the axis of the Euler angle rotation ( $1,2,3$ represent $x, y, z$ respectively).

The angular velocity of 1 relative to $E$ is

$$
\underline{\omega}_{I E}^{I}=\dot{\theta}_{z}\left[\begin{array}{l}
0 \\
0 \\
1
\end{array}\right]+\dot{\Gamma}_{y}\left[\begin{array}{c}
\sin \theta_{z} \\
\cos \theta_{z} \\
0
\end{array}\right]
$$

The transformation [IE] can also be computed from onboard navigation parameters

$$
\begin{equation*}
[\text { IE }]=\left[\nu_{z}\right]_{3}\left[\eta_{x}\right]_{1}\left[\eta_{y}\right]_{2}\left[\lambda_{z}\right]_{3}\left[\lambda_{y}\right]_{2} \tag{C2}
\end{equation*}
$$

The angle $\lambda_{y}$ is the orbit regression angle and $\lambda_{y}$ is the nearly constant regression rate. The angles in equation (C2) can all be expressed in terms of $\lambda_{y}$ and $\Gamma_{y}$. The vector $\underline{u}_{p_{1}}$ is a row from the vehicle-toprincipal axes transformation [PV],

$$
\underline{u}_{P 1} V=[P V]^{T}\left[\begin{array}{l}
1 \\
0 \\
0
\end{array}\right]=\left[\begin{array}{l}
P V_{11} \\
P V_{12} \\
P V_{13}
\end{array}\right]
$$

For convenience, we let $[K]=[P V]$ so that

$$
\underline{u}_{P 1}^{V}=\left[\begin{array}{l}
K_{11} \\
K_{12} \\
k_{13}
\end{array}\right]
$$

The unit vector to the sun is $\underline{u}_{\mathrm{R}} 3$ and is

$$
\underline{\mathbf{u}}_{\mathrm{R} 3}^{\mathrm{E}}=\left[\begin{array}{c}
\sin \Gamma_{y} \cos \phi_{z} \\
\sin \Gamma_{\mathbf{y}} \sin \phi_{z} \\
\cos \Gamma_{y}
\end{array}\right]
$$

Similarly, the orbit normal is

$$
\underline{u}_{\mathbf{O} 2}^{\mathbf{E}}=\left[\begin{array}{c}
-\sin \lambda_{z} \cos \lambda_{y} \\
\cos \lambda_{z} \\
\sin \lambda_{z} \sin \lambda_{y}
\end{array}\right] .
$$

The angle $\eta_{y}$ is by definition the position of orbital noon, i.e., the position in orbit where the radius vector, the solar vectior, and the orbit normal are coplanar with $\underline{u}_{\mathrm{R} 3} \cdot \underline{u}_{\mathrm{L} 3} \geq 0$.

$$
\begin{gathered}
\eta_{y}=\tan ^{-1}\binom{u_{R 3 x}^{O}}{u_{R 3 z}^{O}} ; \text { where } \underline{u}_{R j}^{O}=[O E] \underline{u}_{R 3} E \\
{[O E]=\left[\lambda_{z}\right]_{3}\left[\lambda_{y}\right]_{2}} \\
\eta_{y}=\tan ^{-1}\left(\frac{\cos \lambda_{z} \cos \lambda_{y} \sin \Gamma_{y} \cos \phi_{z}+\sin \lambda_{z} \sin \Gamma_{y} \sin \phi_{z}-\cos \lambda_{z} \sin \lambda_{y} \cos \Gamma_{y}}{\sin \lambda_{y} \sin \Gamma_{y} \cos \phi_{z}+\cos \lambda_{y} \cos \Gamma_{y}}\right)
\end{gathered}
$$

Note: The $\tan ^{-1}$ technique used must be a 4 quadrant technique! The angle $\eta_{x}$ is the elevation angle of the sun out of the orbit plane with orbit south the positive hemisphere.

$$
\begin{aligned}
" x & =\sin ^{-1} u_{R 3 y}^{0} \\
& \left.=\sin ^{-1}\left[\sin x_{z} \cos y_{y} \sin y_{y} \cos \psi_{z}-\cos \right\}_{z} \sin r_{y} \sin \psi_{z}-\sin y_{z} \sin t_{y} \cos r_{y}\right]
\end{aligned}
$$

The remaining angle $\nu_{z}$ is obtained by solving

$$
\begin{equation*}
\underline{u}_{P 1} \cdot \underline{u}_{O 2}=0 \& \underline{u}_{P 1} \cdot \underline{u}_{L 1} \geq 0 \text { at orbit noon } \tag{C3}
\end{equation*}
$$

$$
+\nu_{z}=-\phi+\sin ^{-1}\left(\frac{K_{13} \tan \eta_{x}}{\sqrt{1-K_{13}^{2}}}\right) ; \phi=\tan ^{-1}\left(\frac{K_{12}}{K_{11}}\right)
$$

## (4 Quadrant $\tan ^{-1}$ )

With these intermediate angles and the values of $\lambda_{y}, \lambda_{z}, \Gamma_{y}, \phi_{z}$, we can compute [IE] from equation (C2) for any time. We can also compute $\theta_{z}$.
$\underline{u}_{o z}^{R^{\prime}}=\left[\begin{array}{c}-\cos \Gamma_{y} \cos \phi_{z} \cos \lambda_{y} \sin \lambda_{z}+\cos \Gamma_{y} \cos \lambda_{z} \sin \phi_{z}-\sin \Gamma_{y} \sin \lambda_{y} \sin \lambda_{z} \\ \sin \phi_{z} \cos \lambda_{y} \sin \lambda_{z}+\cos \lambda_{z} \cos \phi_{z} \\ -\sin \Gamma_{y} \cos \phi_{z} \cos \lambda_{y} \sin \lambda_{z}+\sin \Gamma_{y} \cos \lambda_{z} \sin \phi_{z}+\cos \Gamma_{y} \sin \lambda_{y} \sin \lambda_{z}\end{array}\right]$
From equation (C4) then we obtain

$$
A_{i}=-\sin ^{-1} \frac{K_{13} u_{02 z}^{R}}{\sqrt{\left(1-K_{13}^{2}\right)\left(1-u_{02 z}^{R z}\right)}}-\tan ^{-1} \frac{K_{11} u_{O 2 x}^{R}+K_{12} u_{O z y}^{R}}{K_{11} u_{O 2 y}^{R}-K_{12}{ }_{0}^{R}}
$$

We can now use equation (C1) to determine ${ }_{-1}^{I}$. Figure $C-1$ shows how the components of $\underline{\omega}_{\mathrm{IE}}^{\mathrm{I}}$ vary with time as the sun elevation angle (the so-called $\beta$ angle $\beta=-\eta_{x}$ ) goes through its maximum value in the northern orbital hemisphere. We had not realized prior to this that the solar inertial frame could move so fast and were unpleasantly surprised when EOVV attitude was lost due to strapdown drift and resultant momentum saturation. For more on this incident see Reference 2. After this incident, we developed a procedure to update the $z$ momentum bias in such a way that the necessary rotations would be supplied by the bias value rather than a momentum error. The parameter was called a ramp bias, $e_{r b}$ [equation (12)]. Since the $z$ update angle depends on $e_{r}-e_{r b}$, we could keep $e_{r}$ small by using $e_{r b}$ and since we now knew what to expect we could plan in advance what $e_{r b}$ updates were needed to keep $e_{p}$ small. Table C-1 and Figure C-1 show the $e_{r b}$ updates that should have been made to prevent the loss of momentum control we experienced. The knowledge and experience gained here allowed us to successfully pass through two similar peaks of opposite signs in November and January.


Figure C-1. Motion of I coordinate system (deg/orbit).

TABLE C-1. SCHEDULE OF ERB UPDATE TIMES

|  |  | 18119129 <br> 101123144 <br> 101117180 | 23 22 21 |
| :---: | :---: | :---: | :---: |
| UPDATE TINE | NEH ERD UALUE IN LSES | 181187180 | -21 |
| 16618810 | 2 | 181121:86 | 20 |
| 17010183 | 1 | 1821219 | 19 |
| 27217126 | 0 | 28216129 | 18 |
| 873116185 | -1 | 202110149 | 17 |
| 174183127 | - 2 | 102115132 | 16 |
| 1751417 |  | 102120123 | 15 |
| 175115123 | 4 | 10311126 | 14 |
| 17610134 | -8 | 103: 712 | 13 |
| 17617139 | -6 | 183112146 | 12 |
| 17618413 | -7 | 103119120 | 11 |
| 176119121 | -0 | 10412116 | 10 |
| 17780830 | -9 | 18418012 | 9 |
| 17714133 | - 10 | 184198150 | - |
| 17718135 | -18 | 105: 4144 | ? |
| 177112129 | -12 | 105:16:9 | 6 |
| 177115141 | -13 | 10615137 | 5 |
| 877118182 | -14 | 106121142 | 4 |
| 87712214 | - 18 | 10712714 | -3 |
| 1701115 | -16 | 100110153 | -2 |
| 17813151 | $=17$ | $190: 3127$ | 1 |
| 17816136 | - 18 | 151121133 | 0 |
| 178: 9122 | $=19$ | 19410140 | 1 |
| 17012216 | 20 |  |  |
| 170114155 | -21 |  |  |
| 170117143 | -22 | ASSUMES KMUZ 150.2 |  |
| 178120130 | -23 |  |  |
| 17812J180 | -24 |  |  |
| 17912147 | -25 |  |  |
| 17916130 | -26 |  |  |
| 179110113 | -27 |  |  |
| 179116136 | -20 |  |  |
| $100: 13188$ | -27 |  |  |
| 100119181 | - 26 |  |  |
| 10110147 | -25 |  |  |
| $181: 810$ | -24 |  |  |

## APPENDIX D. GAIN CALCULATIONS

## a. Calculation of IOP Desaturation Gain $K_{\eta z}$

The local vertical in principal axes components is ( $s=\sin , c=\cos$ )

$$
\underline{p}_{p}=\left[\begin{array}{ccc}
c z & s z & 0 \\
-s z & c z & 0 \\
0 & 0 & 1
\end{array}\right]\left[\begin{array}{ccc}
c y & 0 & -s y \\
0 & 1 & 0 \\
s y & 0 & c y
\end{array}\right]\left[\begin{array}{l}
0 \\
0 \\
1
\end{array}\right]=\left[\begin{array}{c}
-c z s y \\
s z s y \\
c y
\end{array}\right],
$$

where

$$
y=\eta_{y n}=-\eta_{y m} s\left(2 \Omega_{o} t\right), z=\eta_{z n}=-\eta_{z m} s\left(2 \Omega_{o} t\right)
$$

and the $\eta_{\mathrm{xn}}$ rotation is neglected (a cylindrical inertia distribution is assumed with $I_{y}=I_{z}=I$ and $\Lambda I=I-I_{x}$ ). The gravity gradient torque is (in P components)

$$
\underline{T}_{g g p}=K_{g g}\left[\begin{array}{cc}
0 \\
-r_{1} & r_{3} \\
r_{1} & r_{2}
\end{array}\right]=K_{g g}\left[\begin{array}{c}
0 \\
\text { czsycy } \\
- \text { czsyszsy }
\end{array}\right],
$$

where

$$
\mathrm{K}_{\mathrm{gg}}=3 \Omega_{\mathrm{o}}^{2} \Delta \mathrm{I}
$$

and in 0 components (where $Y=\Omega_{0} t+y$ )

$$
\underline{\mathrm{T}}_{\mathrm{ggo}}=\mathrm{K}_{\mathrm{gg}}\left[\begin{array}{ccc}
\mathrm{cY} & 0 & \mathrm{sY} \\
0 & 1 & 0 \\
-\mathrm{sY} & 0 & \mathrm{cY}
\end{array}\right]\left[\begin{array}{ccc}
\mathrm{cz} & -\mathrm{sz} & 0 \\
\mathrm{sz} & \mathrm{cz} & 0 \\
0 & 0 & 1
\end{array}\right]\left[\begin{array}{c}
0 \\
\text { czsycy } \\
- \text { szczsysy }
\end{array}\right]
$$

$$
=K_{g g} \operatorname{szczsy}\left[\begin{array}{c}
-c \Omega_{o} t \\
c y / t z \\
s \Omega_{o} t
\end{array}\right]
$$

For the further development only the IOP component is of interest, which is (assuming $y$ and $z$ small)

$$
\underline{T}_{\mathrm{ggz}}, \mathrm{x}=\mathrm{K}_{\mathrm{gg}} \eta_{\mathrm{ym}} \eta_{\mathrm{zm}} \mathrm{~s}^{2}\left(2 \Omega_{\mathrm{o}} \mathrm{t}\right)\left[\begin{array}{c}
\mathrm{s} \Omega_{\mathrm{o}} \mathrm{t} \\
-\mathrm{c} \mathrm{\Omega}_{o} t
\end{array}\right]
$$

Integrating over a quarter orbit yields

$$
\begin{aligned}
\underline{H}_{g g o z, x} & =4 K_{g g} \eta_{y m} \eta_{z m} \int_{0}^{\pi / 2 \Omega_{0}} d t\left(s^{2} \Omega_{o} t c^{2} \Omega_{o} t\right)\left[\begin{array}{c}
s \Omega_{o} t \\
-c \Omega_{o} t
\end{array}\right] \\
& =4 K_{g g} \eta_{y m} \eta_{z m}\left[\begin{array}{l}
\pi / 2 \Omega_{0} \\
\int_{0} d t \operatorname{s\Omega _{o}t}\left(c^{2} \Omega_{o} t-c^{4} \Omega_{o} t\right) \\
-\int_{0}^{\pi / 2 \Omega_{0}} d t c \Omega_{o} t\left(s^{2} \Omega_{o} t-s^{4} \Omega_{o} t\right. \\
0
\end{array}\right]
\end{aligned}
$$

Substitution yields

$$
\left(u=c \Omega_{0} t ; v=s \Omega_{0} t\right)
$$

${\underset{-g g o z ~}{x}}^{H_{0}}=\frac{4}{\Omega_{0}} K_{g g} \eta_{y m} \eta_{z m}\left[\begin{array}{l}\int_{0}^{1} d u\left(u^{2}-u^{4}\right) \\ -\int_{0}^{1} d v\left(v^{2}-v^{4}\right)\end{array}\right]=\frac{4}{\pi_{o}} K_{g g} \eta_{y m} \eta_{z m} \frac{2}{15}\left[\begin{array}{r}1 \\ -1\end{array}\right]$,

$$
\underline{H}_{\mathrm{ggoz}, \mathrm{x}}=\frac{24}{15} \Omega_{o} \Delta I \eta_{\mathrm{ym}} \eta_{z \mathrm{~m}}\left[\begin{array}{c}
1 \\
-1
\end{array}\right]
$$

Rotation to the effective axis system yields

$$
\begin{aligned}
& \underline{H}_{\mathrm{ggz}, \mathrm{x}}=\frac{24}{15} \Omega_{0} \Delta \mathrm{I} \eta_{\mathrm{ym}} \eta_{\mathrm{zm}} \frac{1}{\sqrt{2}}\left[\begin{array}{rr}
1 & 1 \\
-1 & 1
\end{array}\right]\left[\begin{array}{r}
1 \\
-1
\end{array}\right], \\
& \mathrm{H}_{\mathrm{ggz}}=0, \\
& \mathrm{H}_{\mathrm{ggx}}=-(24 \sqrt{2} / 15) \Omega \Delta \mathrm{I} \eta_{\mathrm{ym}} \eta_{\mathrm{mm}} \quad .
\end{aligned}
$$

Resolving actual momentum to be desaturated into the same axes

$$
\frac{1}{\sqrt{2}}\left[\begin{array}{c}
\mathrm{H}_{z}+\mathrm{H}_{x} \\
-\mathrm{H}_{z}+\mathrm{H}_{x}
\end{array}\right]=\frac{1}{\sqrt{2}}\left[\begin{array}{rr}
1 & 1 \\
-1 & 1
\end{array}\right]\left[\begin{array}{l}
\mathrm{H}_{z} \\
\mathrm{H}_{x}
\end{array}\right]
$$

Only the x component can be desaturated leading to (in units of the nominal angular momentum $H$ of one CMG):

$$
(H / \sqrt{2})\left(e_{x}-e_{z}\right)=(48 / 15 \sqrt{2}) \Omega_{0} \Delta I \eta_{y m} \eta_{z m} \text {, }
$$

or

$$
\eta_{\mathrm{zm}}=\mathrm{K}_{\mathrm{nz}}\left(\mathrm{e}_{\mathrm{x}}-\mathrm{e}_{\mathrm{z}}\right)
$$

with the gain being

$$
\mathrm{K}_{\eta \mathrm{z}}=(15 \mathrm{H}) /\left(48 \Omega_{\mathrm{o}} \Delta \mathrm{I} \eta_{\mathrm{ym}}\right)
$$

## b. Calculation of POP Gain $K_{y c}$

The gravity gradient orbital $y$ torque for a rotation $\eta_{y}$ about the orbital $y$ axis through $\eta_{y}$ is (APPENDIX B)

$$
\begin{aligned}
\mathbf{T}_{\mathbf{g g o y}} & =\mathbf{K}_{\mathbf{g g}}{ }^{\mathbf{s} \eta_{\mathbf{y}} \mathbf{e} \eta_{\mathbf{y}}} \\
& \approx \mathbf{K}_{\mathbf{g g}} \eta_{\mathbf{y}}
\end{aligned}
$$

If the angle $\eta_{y}$ is held constant over the desaturation period $T_{\text {dos }}$ (about $1 / 12$ of an orbit) we get an orbital $y$ momentum change of

$$
\Delta e_{g g o y}=K_{g g} I_{d e s} \eta_{y} / \mathrm{H}
$$

in units of the nominal CMG momentum $H$. For a given $\Lambda e_{o y}$ we need

$$
\eta_{\mathrm{yc}}=\mathrm{K}_{\mathrm{yc}} \Delta \mathrm{e}_{\mathrm{oy}}
$$

with

$$
\mathrm{K}_{\mathrm{yc}}=\mathrm{H} /\left(3 \Omega_{\mathrm{o}}^{2} \Delta \mathrm{IT} \mathrm{des}\right)
$$

$$
\text { c. Calculation of strapdown Update Gain } \mathrm{K}_{\mu \mathrm{z}}
$$

The components of the local vertical in the $P$ system are (neglecting the tilt angle $\eta_{z m} ; s=\sin , c=\cos$ )

$$
{\underset{\mathrm{y}}{\mathrm{p}}}=\left[\begin{array}{ccc}
\mathrm{cN} & 0 & -\mathrm{sN} \\
0 & 1 & 0 \\
\mathrm{sN} & 0 & \mathrm{cN}
\end{array}\right]\left[\begin{array}{ccc}
1 & z & 0 \\
-z & 1 & 0 \\
0 & 0 & 1
\end{array}\right]\left[\begin{array}{ccc}
\mathrm{cy} & 0 & \mathrm{sy} \\
0 & 1 & 0 \\
-\mathrm{sy} & 0 & \mathrm{cy}
\end{array}\right]\left[\begin{array}{l}
0 \\
0 \\
1
\end{array}\right]=\left[\begin{array}{r}
-\mathrm{sn} \\
-\mathrm{zsy} \\
\mathrm{cn}
\end{array}\right]
$$

with

$$
N=y+n ; y=\Omega_{0} t ; n=-\eta_{y m} s\left(2 \Omega_{0} t\right) ; s \mu_{z} \approx 2 ; c \mu_{z} \approx 1
$$

For $I=I_{y}=I_{z}$ and $\Delta I=I-I_{x}$ we get for the principal axis torque due to the gravity gradient in the inertially fixod 0 system (with $K_{g g}=3 \Omega_{0}^{2}{ }^{2} \mathrm{I}$ )

$$
\underline{\mathrm{T}}_{\mathrm{gg}}=\mathrm{K}_{\mathrm{gg}} \mathrm{r}_{1}\left[\begin{array}{ccc}
\mathrm{cN} & 0 & \mathrm{sN} \\
0 & 1 & 0 \\
-\mathrm{SN} & 0 & \mathrm{cN}
\end{array}\right]\left[\begin{array}{c}
0 \\
\mathrm{r}_{3} \\
\mathrm{r}_{2}
\end{array}\right]=\mathrm{K}_{\mathrm{gg}}{ }^{\mathrm{r}_{1}}\left[\begin{array}{l}
\mathrm{r}_{2} \mathrm{sN} \\
-\mathrm{r}_{3} \\
\mathrm{r}_{2} \mathrm{cN}
\end{array}\right]
$$

Only the torque about the a axis is of interest

$$
\mathrm{T}_{\mathrm{gg} z}=\mathrm{K}_{\mathrm{gg}}{ }^{z s n s y(\text { cyen-sysn })}
$$

With $n=K s 2 y ;$ sn $\approx K s 2 y ;$ on $\approx 1$ and $K=-\eta_{y m}$ we get

$$
\begin{aligned}
\mathrm{T}_{\mathrm{gg} 2} & =\mathrm{K}_{\mathrm{gg}} 2 \mathrm{Ks} 2 \mathrm{ysy}(\mathrm{cy-syKs} 2 \mathrm{y}) \\
& =0.25 \mathrm{~K} \mathrm{gg}^{2 \mathrm{~K}}[(1-\mathrm{K})-(1-\mathrm{K}) \mathrm{c} 4 \mathrm{y}=2 \mathrm{yc} 4 \mathrm{y}]
\end{aligned}
$$

Elimination of the eyclic terms leaves a constant bias torque of

$$
\mathrm{T}_{\mathrm{ggz}}=0.25 \mathrm{~K}_{\mathrm{gg}} 2 \mathrm{~K}(1-\mathrm{K})
$$

Integrating over an orbit ( $T_{\text {orbit }}=2 \pi / \Omega_{o}$ ) and substituting the value for $z, K_{g g}$, and $K$ gives

$$
\Delta H_{g g z b i a s}=-(3 / 2) s_{0} \Delta I \eta_{y m}\left(1+\eta_{y m}\right)
$$

## Inversion yields

$$
\mu_{z}=-K_{\mu z} e_{z \mathbf{r}},
$$

where

$$
K_{\mu z}=2 H /\left[3 \Omega_{0} \Delta I \eta_{y m}\left(1+\eta_{y m}\right)\right]
$$

and $e_{z 1}$ is the normalized $z$ momentum ramp,
Because of noise, a much smaller $K_{\mu z}$ value had to be used in actual operation (see EOVV OPERATION and PERFORMANCE),

## APPENDIX E

## QUATERNIONS - A BRIEF EXPOSITION

Complex numbers of the form $z=a+i b$ have proven to be a valuable concept in the study of many physical phenomena. A goneralization of this concept which proves useful in the study of rotational motion is the quaternion. Recall that the imaginary unit $i=\sqrt{-1}$. Let us define additional units $j$ and $k$ together with the product operation $n$ :

$$
\begin{align*}
& \text { io } i=j \circ j=k \circ k=-1 \\
& \text { io } j=-j \circ i=k  \tag{E1}\\
& \text { jok } \circ \mathrm{k}=-\mathrm{k} \circ \mathrm{j}=\mathrm{i}=\mathrm{j} \\
& \mathrm{k} \circ \mathrm{i}=-\mathrm{i} \circ \mathrm{k}=\mathrm{j}
\end{align*}
$$

We shall define a quaternion as any quantity of the form

$$
\begin{equation*}
Q=Q 4+i Q 1+j Q 2+k Q 3 \tag{E2}
\end{equation*}
$$

(See Note 1 at end of Appendix E.) By analogy to the complex number terminology Q4 is referred to as the real or scalar part of Q. It will also be convenient to think of the remaining part of $Q$ as the imaginary or vector part. The reason for this will become clear as we proceed. Let $R=R 4+i R 1+j R 2+k R 3$. The sum of quaternions $Q$ and $R$ is defined as

$$
\begin{align*}
S=Q+R= & (Q 4+R 4)+i(Q 1+R 1) \\
& +j(Q 2+R 2)+k(Q 3+R 3) \tag{E3}
\end{align*}
$$

From this definition we can see that the sum operation is commutative and associative. We can now give the complete definition of the product operation o:

$$
\begin{align*}
P=Q O R= & (Q 4 R 4-Q 1 R 1-Q 2 R 2=Q 3 R 3) \\
& +i(Q 4 R 1+Q 1 R 4+Q 2 R 3=Q 3 R 2) \\
& +j(Q 4 R 2+Q 2 R 4+Q 3 R 1-Q 1 R 3) \\
& +k(Q 4 R 3+Q 3 R 4+Q 1 R 2-Q 2 R 1) \tag{E4}
\end{align*}
$$

With this definition we can show that $o$ is associative and distributive but not commutative, i.e., $Q$ o $R \neq R o Q$. We shall call any quaternion havIng zero imaginary part a scalar and, obviously, the algebra of sealars is just the algebra of real numbers. Thus, multiplication of a quatomion by a scalar simply results in a quaternion whose elements are multiplied by that scalar according to definition E4. We can now define the difference eperation as

$$
\begin{equation*}
\mathrm{D}=\mathrm{Q}-\mathrm{R}=\mathrm{Q}+(-1) \text { o } \mathrm{R} \tag{1:5}
\end{equation*}
$$

For convenience, we shall always omit the o when multiplying a quaternion by a scalar so that ( -2 ) o $Q=-2$ Q.

By analogy to complex aggebra, let us define the conjugate quaternion to $\mathbb{Q}$. The conjugation operation will be denoted by ()*. Thus

$$
\begin{equation*}
\mathbf{Q}^{*}=\mathbf{Q} 4=\mathrm{i} Q 1=j Q 2=k Q 3 \tag{E6}
\end{equation*}
$$

So far all of our definitions have been extensions of those for complex numbers as can be seen by assuming $Q 2=\mathbb{Q} 3=R 2=R 3=0$. Thus the complex number system is a subset of the quaternions. It can be easily seen that

$$
\begin{equation*}
Q^{*} O Q=Q \circ Q^{*}=Q 1 Q 1+Q 2 Q 2+Q 3 Q 3+Q 4 Q 4 \tag{E7}
\end{equation*}
$$

Note that $Q^{*}$ o $Q$ is a pure real number or sealar. With this observation we can define the inverse:

$$
\begin{equation*}
Q^{-1}=\left(1 /\left(Q^{*} \circ Q\right)\right) Q^{*}=Q^{*} /\left(Q^{*} \circ Q\right) ; Q^{*} \circ Q \neq 0 \tag{E8}
\end{equation*}
$$

Finally then, we can define a division operation:

$$
\begin{equation*}
\mathbf{Q}: \mathbf{R}=\mathbf{Q} \circ \mathbf{R}^{-1} \tag{E9}
\end{equation*}
$$

We can see that $Q: Q=Q \circ Q^{-1}=Q^{-1} \circ Q=1$ so that $Q^{-1}$ satisfies the necessary properties of an inverse as long ns $Q \neq 0$. This will be useful later.

We need some additional results and definitions. First we can show that

$$
\begin{equation*}
(\mathbf{Q} \circ \mathbf{R})^{*}=\mathbf{R}^{*} \circ \mathrm{Q}^{*} . \tag{E10}
\end{equation*}
$$

If the quaternion $Q=Q^{*}$, then $Q$ is necessarily a scalar. Also, if $\mathbf{Q}=$ $-Q^{*}, Q$ is purely imaginary or a vector quaternion. If $V$ is a vector quaternion, we shall designate this by an underline as is also used to designate a 3 -space vector, i.e., ( $i, j$, and $k$ will not be underlined)

$$
\begin{equation*}
\underline{V}=\mathrm{i} V 1+j \mathrm{~V} 2+\mathrm{kV} 3 \tag{E11}
\end{equation*}
$$

For compactness of our notation we shall let

$$
\begin{equation*}
\mathbf{Q}=\mathbf{Q} 4+\underline{\mathbf{Q}}, \tag{E12}
\end{equation*}
$$

where

$$
\underline{Q}=\mathrm{i} Q 1+\mathrm{j} Q 2+\mathrm{k} Q 3 .
$$

Thus

$$
\begin{equation*}
Q \circ R=(Q 4 Q 4-\underline{Q} \cdot \underline{R})+Q 4 \underline{R}+\mathbf{R} 4 \underline{Q}+\underline{Q} \times \underline{R} \quad . \tag{E13}
\end{equation*}
$$

The operations and $\times$ are defined as for 3 -space vectors so that

$$
\begin{equation*}
\underline{\mathrm{Q}} \cdot \underline{\mathrm{R}}=\mathrm{Q} 1 \mathrm{R} 1+\mathrm{Q} 2 \mathrm{R} 2+\mathrm{Q} 3 \mathrm{R} 3 \tag{E14}
\end{equation*}
$$

and

$$
\begin{align*}
\underline{Q} \times \underline{R}= & i(Q 2 R 3-Q 3 R 2)+j(Q 3 R 1-Q 1 R 3) \\
& +k(Q 1 R 3-Q 3 R 1) \tag{E15}
\end{align*}
$$

For vectors $\underline{A}$ and $\underline{B}$,

$$
\begin{equation*}
\underline{\mathbf{A} O \underline{B}=-\underline{A} \cdot \underline{B}+\underline{A} \times \underline{B} . . . . ~ . ~} \tag{E16}
\end{equation*}
$$

In general the product of quaternions mixes scalar and vector parts together so that this product is not very interesting in the study of rotational motion in 3 -space. The triple product

$$
\begin{equation*}
V^{\prime}=Q^{*} \circ V \circ Q \tag{E17}
\end{equation*}
$$

is more interesting since it does pueserve scalar and vector parts of $V$ without mixing them. This property is trivial for the scalar part of V and follows for the vector part since

$$
\begin{equation*}
\underline{V}^{\prime *}=\left(\mathbf{Q}^{*} \circ \underline{\mathrm{~V}} \circ \mathrm{Q}\right)^{*}=-\mathrm{Q}^{*} \circ \underline{\mathrm{~V}} \circ \mathrm{Q}=-\underline{\mathrm{V}}^{\prime} \tag{E18}
\end{equation*}
$$

Hence as noted the twiple quaternion product (E17) takes a scalar into a scalar and a vector into a vector for any quaternion Q. Furthermore, the length of the vector $|\underline{V}|=\sqrt{\mathrm{V} \cdot \underline{\mathrm{V}}}$ and

$$
\begin{align*}
\underline{\mathbf{V}}^{\prime} \cdot \underline{V}^{\prime} & =\underline{V}^{\prime} \circ \underline{V}^{\prime}{ }^{*}=\mathbb{Q}^{*} \circ \underline{\mathbf{V}} \circ \mathbb{Q} \circ\left(\mathbb{Q}^{*} \circ \underline{\mathbf{V}} \circ \mathbb{Q}^{*}\right. \\
& =\left(\mathbb{Q}^{*} \circ \mathbb{Q}^{2} \underline{\mathbf{V}} \circ \underline{V}^{*}\right. \tag{E19}
\end{align*}
$$

Equation (E19) indicates the triple product (E17) multiplies vector length by the factor $Q^{*} \circ Q$ which is a real number. We note that if $Q^{*} \circ Q=1$, vector length is preserved and the vector mapping $V \rightarrow V^{\prime}$ looks like a rotation operator. It is a linear operator in that $\bar{a} A+b B+a A^{\prime}+b B^{\prime}$. Restricting ourselves to normalized quaternions which preserve length, we see that E 17 is equivalent to a rotation of vector $V$ into $V^{\prime}$. Since we are looking only at normalized quaternions, we can withoūt loss of generality, represent $Q$ as

$$
\begin{equation*}
\mathbf{Q}=\cos \phi / 2+\sin \phi / 2 \underline{\mathbf{u}} \quad ; \quad \text { where } \underline{\mathbf{u}} \cdot \underline{\mathbf{u}}=1 \tag{E20}
\end{equation*}
$$

The triple product equation (E17) can be combined with equation (E20) to give

$$
\begin{equation*}
\underline{\mathbf{V}}^{\prime}=\cos \phi \underline{\mathbf{V}}-\sin \phi \underline{\mathbf{u}} \times \underline{\mathbf{V}}+(1-\cos \phi) \underline{\mathbf{u}} \underline{\mathbf{u}} \cdot \underline{\mathbf{V}} \tag{E21}
\end{equation*}
$$

Equation (E21) is the general form of the rotation of a vector V about axis $u$ through the angle $-\phi$. That this is true is seen by examining the rotation operation. Clearly any vector along the rotation axis $u$ is not changed by the rotation so that if the vector $V$ is broken into parts parallel to and normal to $u$ i.e.

$$
\begin{equation*}
\underline{v}=\underline{v}_{\|}+\underline{v}_{\perp}, \quad \text { where } \underline{v}_{\|}=\underline{v} \cdot \underline{u} \underline{u} \text { and } \underline{v}_{1} \cdot \underline{u}=0 \tag{E2.}
\end{equation*}
$$

We must also have $V^{\prime}=V_{\|}+V_{\perp}^{\prime}$; where $V_{\perp}^{\prime} \cdot \underline{u}=0$ and $V_{\|}=V_{\|}^{\prime}$. Since $V_{\perp}$ and $V_{\perp}^{\prime}$ are normal to $\underset{\sim}{\square}$, we can express ${\underset{\sim}{V}}^{\prime}$ as

$$
\begin{equation*}
\underline{v}_{\perp}^{\prime}=x \times \underline{v}_{\perp}+y \underset{\sim}{u} \times\left(\underline{u} \times \underline{v}_{1}\right) \tag{E23}
\end{equation*}
$$

Now $\underline{y}^{\prime} \cdot \underline{y}=\left|\underline{V_{1}}\right|^{2} \cos x=-\mathrm{y}\left|\underline{V}_{1}^{2}\right| \times y=-\cos$ ix and $\underline{V}_{1} \times \underline{V}_{1}^{\prime}=$ $\left|V_{1}\right|^{2} \frac{1}{\sin } x \frac{d}{u}$ and thus $x=\sin x$; where $x$ is the rotation angle. Combining, we obtain

$$
\begin{equation*}
{\underset{\perp}{V}}_{\prime}^{\prime}=\sin \alpha \underset{u}{u} \times V_{L} \times \cos \alpha \underline{u} \times\left(\underline{u} \times \underline{V}_{1}\right) \tag{E24}
\end{equation*}
$$

and

$$
\begin{equation*}
\underline{V}^{\prime}=\underline{\mathbf{u}} \underline{\mathbf{u}} \cdot \underline{V}+\sin \times \underline{\mathbf{u}} \times \underline{V}-\cos \times \underline{\mathbf{u}} \times(\underline{\mathbf{u}} \times \underline{v}) \tag{E25}
\end{equation*}
$$

The equivalence between equations (E21) and (E25) for $\alpha=-\phi$ is established. Thus, the mapping (E17) is equivalent to a rotation operator in a vector 3 -space. Since the coordinate directions are also vectors, we can rotate the coordinate system instead of the vector. Rotating the vector through - ${ }^{-\phi}$ yields the same components $V^{\prime}$ as rotating the coordinate axes through $\phi$. Thus we can look at $V^{\prime} \bar{a}$ as a new vector formed from $V$ by rotation and expressed in the old axis system or as the old
vector expressed in new axes rotated relative to the old. We have now demonstrated that any coordinate system rotation can be represented by a quaternion. Note that if $Q$ satisfies equation (E17), then so does -Q. Looking back at equation (E20) tells us that $-Q$ corresponds to $\phi+360$ degrees which represents the same attitude $\phi$ does.

We now look at the time variations of $Q$. Since $Q$ is constrained to be normalized, we necessarily have

$$
\begin{equation*}
d / \operatorname{dt}\left(Q^{*} \circ Q\right)=0=\dot{Q}^{*} \circ Q+Q^{*} \circ \dot{Q} \tag{E26}
\end{equation*}
$$

We see from equation (E26) that $Q^{*} \circ \mathbf{Q}=-(Q * \circ$ Q $) *$ and hence must be a vector. Let us define this vector by

$$
\begin{align*}
& Q^{*} \circ \dot{Q}=1 / 2 \underline{\varphi}  \tag{E27}\\
& \rightarrow \dot{Q}=1 / 2 Q \circ \underline{\varphi}\left(\text { since } Q \circ Q^{*}=Q^{*} \circ Q=1\right) . \tag{E28}
\end{align*}
$$

We shall see the reason for the $1 / 2$ factor later. When we evaluate the rate of change of a vector with time in two reference frames, we find

$$
\begin{equation*}
\underline{\underline{V}}=\mathbb{Q} \circ \dot{\underline{V}}^{\prime} \circ Q^{*}+\dot{Q} \circ \underline{V}^{\prime} \circ Q^{*}+Q \circ \underline{V}^{\prime} \circ \dot{Q}^{*} \tag{E29}
\end{equation*}
$$

Using equations (E28) in (E29),

$$
\begin{align*}
& \underline{\mathbf{V}}=Q \circ\left[\dot{\mathbf{V}}^{\prime}+1 / 2\left(\underline{\underline{u}} \circ \underline{V}^{\prime}-\underline{V}^{\prime} \circ \underline{( }\right)\right] \circ Q^{*}  \tag{E30}\\
& \dot{\mathbf{V}}=Q \circ\left[\dot{\mathbf{V}}^{\prime}+\underline{\underline{u}} \times \underline{V}^{\prime}\right] \circ Q^{*} . \tag{E31}
\end{align*}
$$

Now the reason for the factor $1 / 2$ becomes clear. It is so that we can identify u.. Equation (E31) is exactly like the corresponding equation for 3 -space vectors if we identify $\underline{a}$ as the angular velocity of the primed reference frame expressed in primed coordinates. This identifention follows from the fact that equation (E31) holds for an arbitrary vector $V$. Thus, $\underline{\omega}$ is identified as the relative angular velocity of the primed axes with respect to the unprimed.

The above discussion completes the basic development of our quaternion tools. We now turn to the problem of developing a more convenient notation. The most logical choice which comes to mind is a matrix representation. The quaternion $Q$ would logically become

$$
\mathbf{Q}=\left[\begin{array}{l}
\mathbf{Q} 1  \tag{E32}\\
\mathbf{Q} 2 \\
\mathbf{Q} 3 \\
\mathbf{Q 4}
\end{array}\right]
$$

Looking back to the definition equation (E4) of the quaternion product 0 , we see that for $P=Q \circ R$ we have

$$
P=\left[\begin{array}{rrrr}
Q 4 & -Q 3 & Q 2 & Q 1  \tag{E33}\\
Q 3 & Q 4 & -Q 1 & Q 2 \\
-Q 2 & Q 1 & Q 4 & Q 3 \\
-Q 1 & -Q 2 & -Q 3 & Q 4
\end{array}\right]\left[\begin{array}{l}
R 1 \\
R 2 \\
R 3 \\
R 4
\end{array}\right] \triangleq \hat{Q} R
$$

(See Notes at end of Appendix E.) Similarly, i. D $=A$ o B o C, then

$$
\begin{equation*}
D=\widetilde{A \circ B} C=\widetilde{\tilde{\tilde{A} B}} C=\hat{A} \hat{\tilde{B}} C \tag{E34}
\end{equation*}
$$

This rult shows that the set of matrices of the form $\underset{Q}{\mathbb{Q}}$ have the properties, of the quaternions and in fact comprise a maxtrix representation of quaternion algebra with matrix multiplication corresponding to o. We can also express the quatemion product in the alternate form

$$
P=\left[\begin{array}{rrrr}
R 4 & \mathrm{R} 3 & -\mathrm{R} 2 & \mathrm{R} 1  \tag{E35}\\
-\mathrm{R} 3 & \mathrm{R} 4 & \mathrm{R} 1 & \mathrm{R} 2 \\
\mathrm{R} 2 & -\mathrm{R} 1 & \mathrm{R} 4 & \mathrm{R} 3 \\
-\mathrm{R} 1 & -\mathrm{R} 2 & -\mathrm{R} 3 & \mathrm{R} 4
\end{array}\right]\left[\begin{array}{c}
\mathrm{Q} 1 \\
\mathrm{Q} 2 \\
\mathrm{Q} 3 \\
\mathrm{Q} 4
\end{array}\right] \triangleq \hat{\overline{\mathrm{R}}} \mathrm{Q}
$$

(See Notes at end of Appendix E.) Here the mapping also yields a matrix representation except the order of the factors must be reversed. Thus, we have

$$
\begin{align*}
& \mathbf{P}=\hat{\tilde{Q}} \mathbf{R}=\overline{\mathrm{R}} Q \\
& \mathbf{D}=\tilde{\mathrm{A}} \mathbf{B} \mathbf{B}=\overline{\mathrm{C}} \overleftarrow{\mathrm{~B}} A \tag{E36}
\end{align*}
$$

According to these definitions and results, we have

Thus, an interesting and sometimes usefu! result is that $A=0$ Let us now look at our previous work and make use of these new definitions:

Note that we now have a matrix formed form $Q$ which rotates coordinate axes and produces $V$ 'from $V$. Linear vector spaces are also represented by matrices. The $\overline{3}$ space vectors, $\underline{V}$ and $\underline{V}$ 'are related as

$$
\begin{equation*}
\underline{V}^{\prime}=\mathrm{M} \underline{\mathrm{~V}} . \tag{E39}
\end{equation*}
$$

Here $M$ is a $3 \times 3$ matrix which transforms components of $V$ to primed coordinates. Referring back to equation (E25) and replacing is by $-\phi$ wo see that

$$
\begin{equation*}
\mathrm{M}=\underline{\mathrm{u}} \underline{u}^{\mathrm{T}}-\sin \phi \tilde{\mathrm{u}}+\cos \phi\left(\mathbb{\pi}-\underline{\mathbf{u}} \underline{u}^{\mathrm{T}}\right) \tag{E40}
\end{equation*}
$$

(See Notes at end of Appendix E.) The matrix $\tilde{u}$ is the so-calded cross product matrix and just happens to be the upper $3 \times 3$ formed by dropping the final row and column of $\underline{\ddot{W}}$. The matrix $\mathbb{1}$ is the identity matrix of the appropriate size to fit the current application. We have already shown the equivalence between $M$ and $\mathbb{\widetilde { Q }} *$. Let us look more closely at the latter since it is $4 \times 4$. We can partition the double-tilde or doublembar matrices as

$$
\hat{Q}=\mathbf{Q 4 1 1}+\left[\begin{array}{c:c}
Q & \underline{Q}  \tag{E41}\\
\hdashline \mathbf{Q} & 0
\end{array}\right] ; \quad \bar{Q}=\mathbf{Q 4} 11+\left[\begin{array}{c:c}
-\tilde{Q} & \underline{Q} \\
\hdashline-Q^{T} & 0
\end{array}\right]
$$

Since $Q^{*} \circ Q=1, Q^{*}=Q^{-1}$ so that $Q^{*} Q=1$. Also, $Q^{*}=Q$ (transpose) so that

$$
\tilde{Q} * \overline{\bar{Q}}=\left[\begin{array}{cc:c}
Q 4^{2} 1-2 Q 4 \tilde{Q}+\hat{Q} \hat{Q}+\underline{Q} \underline{Q}^{T} & 0  \tag{E42}\\
\hdashline 0 & 1
\end{array}\right]
$$

Hence $M=Q 4^{2} 1=2 Q 4 Q+2 \underline{Q} \underline{Q}^{T}-Q^{2} \mathbf{Q}$. In expanded form

$$
M=\left[\begin{array}{l|l|l}
Q 1^{2}-Q 2^{2} Q 3^{2}+Q 4^{2} & 2(Q 1 Q 2+Q 3 Q 4) & 2(Q 1 Q 3+Q 2 Q 4)  \tag{E43}\\
2(Q 2 Q 1-Q 3 Q 4) & Q 1^{2}+Q 2^{2} Q 3^{2}+Q 4^{2} & 2(Q 2 Q 3+Q 1 Q 4) \\
2 Q 3 Q 1+Q 2 Q 4) & 2(Q 3 Q 2-Q 1 Q 4) & Q 1^{2} \cdot Q 2^{2}+Q 3^{2}+Q 4^{2}
\end{array}\right]
$$

From equation (E28)

$$
\dot{Q}=\frac{1}{2} \mathcal{Q} \circ \underline{\omega}=\frac{1}{2} \dot{\omega} Q=\frac{1}{2}\left[\begin{array}{cccc}
0 & \omega 3 & -\omega 2 & \omega 1  \tag{E44}\\
-\omega 3 & 0 & \omega 1 & \omega 2 \\
\omega 2 & -\omega 1 & 0 & \omega h^{\hat{2}} \\
-\omega 1 & -\omega 2 & -\omega 3 & 0
\end{array}\right]\left[\begin{array}{l}
Q 1 \\
Q 2 \\
Q 3 \\
Q 4
\end{array}\right]
$$

Equations (E43) and (E44) summarize the useful results from our discussion.

We are now ready to consider the question of successive rotations applied to a coordinate reference. A coordinate frame rotation is a rigid displacement of all the poits in the system with a fixed axis passing through the origin. Thus, it would seem that several successive rotations should displace every point except the origin. Let us now consider the coordinate frame as a rigid body and determine the most general displacement of it which keeps one point fixed. We must first explain what is
meant by a rigid body displacement. A rigid body displacement is one which preserves distances between every possible pair of points in the body. The displacement is mathematically represented as a vector function $f$. This function then has two basic properties:

1) $\underline{f}(0)=0$
2) $\left.\mid \underline{f}_{\underline{\mathbf{r}}}^{A}\right) \cdots \underline{f}\left(\underline{r}_{B}\right)\left|=\left|\underline{\underline{r}}_{A}-\underline{\underline{r}}_{\mathbf{R}}\right|\right.$.

To study this in more detail, we define two additional points ${\underset{\mathbf{r}}{1}}^{\text {and }}{\underset{\underline{r}}{2}}^{\underline{f}}$ together with their images under $\underset{\sim}{f}, \underline{f}\left(\underline{r}_{1}\right)$ and $\left.\underset{\underline{f}}{\underline{r_{2}}}\right)$. Let ${\underset{\sim}{f}}_{1}=\underset{\sim}{f}\left(\underline{r}_{1}\right)$ and $\left.\underline{f}_{2}=\underline{f}_{\underline{\underline{p}}}^{2}\right)$. Let us define unit vectors

$$
\begin{align*}
& \mathbf{i}=\underline{\mathbf{r}}_{1} /\left|\underline{\mathbf{r}}_{1}\right| ; \quad \underline{u}_{1}=\underline{\mathrm{f}}_{1} /\left|\underline{\mathrm{r}}_{1}\right| \\
& \mathbf{j}=\frac{\underline{\mathbf{r}}_{2}-\underline{\mathbf{r}}_{2} \cdot \underline{\mathbf{i}} \underline{\underline{i}}}{\left|\underline{\mathbf{r}}_{2}-\underline{\mathbf{r}}_{2} \cdot \underline{\mathbf{i}}\right|} ; \quad \underline{u}_{2}=\frac{\underline{\mathrm{e}}_{2}-\underline{\mathrm{f}}_{2} \cdot \underline{u}_{1} \underline{u}_{1}}{\left|\underline{\mathbf{f}}_{2}-\underline{\mathrm{f}}_{2} \cdot \underline{u}_{1} \underline{u}_{1}\right|}  \tag{E46}\\
& \mathbf{k}=\mathbf{i} \times \mathbf{j} \quad ; \quad \underline{u}_{3}=\underline{u}_{1} \times \underline{u}_{2}
\end{align*}
$$

The vectors $\mathrm{i}, \mathrm{j}, \mathrm{k}$ and $\underline{\mathrm{u}}_{1}, \underline{\mathrm{u}}_{2}, \underline{\mathrm{u}}_{3}$ each form orthonormal bases for 3 dimensional space. An arbitrary vector $\underline{r}$ can be expresses as

$$
\begin{equation*}
\underline{r}=x i+y j+z k \tag{E47}
\end{equation*}
$$

The corresponding $\underline{f}(\underline{r})=f_{x} \underline{u}_{1}+f_{y} \underline{u}_{2}+f_{z} \underline{u}_{3}$. Condition 2 of equation (E45) can only be satisfied if

$$
\begin{equation*}
\underline{f}(\underline{r})=x \underline{u}_{1}+y \underline{u}_{2} \pm z \underline{u}_{3} \tag{E48}
\end{equation*}
$$

Thus we can define two functions $f+$ and $f$ that both satisfy equation (E45) and app $\underline{r}_{1}$ into $\underline{f}_{1}$ and $\underline{\underline{r}}_{2}$ into $\underline{f}_{2}$. The function $\underline{f}$ - can be viewed as the reflection $(x, y, z)+(x, y,-i)$ followed by $f+$. We are only
interested in continuous transitions from an initial position to a final position and thus reflections must be eliminated since it is not possible to go from ( $x, y, z$ ) to ( $x, y,-z$ ) continuously without violating condition 2 of equation (E45). Thus continuous rigid displacements can only oceur in the form $\mathrm{f}+$. This function can be written out as

$$
\begin{align*}
& \underline{f}(\underline{r})=x \underline{u}_{1}+y \underline{u}_{2}+z \underline{u}_{3}=\alpha_{x} i+\alpha_{y} j+\alpha_{z} k  \tag{E49}\\
& \alpha_{x}=x \underline{u}_{1} \cdot i+y \underline{u}_{2} \cdot i+z \underline{u}_{3} \cdot i \\
& \alpha_{y}=x \underline{u}_{1} \cdot j+y \underline{u}_{2} \cdot j+z \underline{u}_{3} \cdot j  \tag{E50}\\
& \alpha_{z}=x \underline{u}_{1} \cdot k+y \underline{u}_{2} \cdot k+z \underline{u}_{3} \cdot k
\end{align*}
$$

Equation (E50) ean be rewritten in the matrix form

$$
\begin{equation*}
\underline{x}=M \underset{2}{r} \quad . \tag{E51}
\end{equation*}
$$

The vectors and $r$ are of the same length and since this must hold for all pairs a and $\underset{\mathrm{m}}{\mathrm{r}}$ we must have that

$$
\begin{equation*}
M^{T} \mathbf{M}=1 \tag{E52}
\end{equation*}
$$

Equation (E52) also implies all cigenvalues of $M$ are of unit magnitude. The eigenvalues and eigenvectors of $M$ may be complex so that if $M \underset{x}{x}=$ $\lambda x$, then $x^{T *} M^{T} M x=1=\lambda^{*} \lambda x^{T}{ }^{T} x$. For 3 -dimensional space $M$ must have at least one real oigenvalue. Since $M$ is real, its eigenvalues must occur in complex pairs. Therefore at least one eigenvalue of $M$ must be equal to 1 . The value -1 could not be acceptable since it would imply $\mathrm{m} u=-\mathrm{u}$ which would be a reflection and already ruled out. Thus, we have that det $\mathrm{M}=1$.

The matrix $M$ is now looking very much like a rotation since the eigenvector $\underline{u}$ is an eigenaxis. All we must do now is to determine the angle of rotation. Along with the eigenvector $u$, let us define unit vectors $v$ and $w$ such that $u, v, w$ is an orthonormal basis set. Also, we assume $\underset{\sim}{w}=\underset{\sim}{u} \times \underset{\sim}{v}$. With these definitions we can express the matrix $M$ as

$$
\begin{equation*}
M=M_{11} \underline{u} u^{T}+M_{12} \underline{u} \underline{v}^{T}+M_{21} \underline{v} \underline{u}^{T}+\ldots+M_{33} \underline{w} \underline{w}^{T} \tag{E53}
\end{equation*}
$$

From the fact that $M \underline{u}=\underline{u}$ and that $M^{T} M=1$ equation (E53) reduces to

$$
\begin{equation*}
\mathrm{M}=\underline{\mathrm{u}} \underline{\underline{u}}^{T}+\mathrm{p}\left(\underline{\mathbf{v}} \underline{\mathbf{v}}^{T}+\underline{w} \underline{w}^{T}\right)+q\left(\underline{\underline{v}} \underline{\underline{w}}^{T}-\underline{w} \underline{\underline{v}}^{T}\right) ; p^{2}+q^{2}=1 \tag{E54}
\end{equation*}
$$

We can now eliminate the vectors $\underline{v}$ and $w$ from this equation by use of the proper function of $\underline{u}$. Thus

$$
\begin{equation*}
M=\underline{u} \underline{u}^{T}+p\left(1-\underline{u} \underline{u}^{T}\right)-q \tilde{u} \tag{E55}
\end{equation*}
$$

This completes the proof that the matrix $M$ is a rotation matrix. This is now obvious from inspection of equation ( E 55 ) by comparing it to equation (E40) with $p=\cos \phi$ and $q=\sin \phi$. Thus the most general displacement of a rigid body (or transformation of a coordinate gystem) in which at least 1 point remains fixed is a rotation about, a fixed nxis i, e, the final orientation an be obtained from the origiral by a single rotation about the axis $u$ through the angle $\phi$ ( $u$ and $\phi$ are determined from $M$ ) even though the actual motion from initial to final may have been more complex.

What all the previous discussion boils down to is that the product of a pair of rotations is itself a rotation. Thus, if $\mathrm{M}_{1}$ and $\mathrm{M}_{2}$ are rotations about $\underline{u}_{1}$ and $\underline{u}_{2}$ respeetively, then $M_{1} M_{2}=M_{3}$ is also a rotation through some angle $\phi_{3}$ about some nxis $\underline{u}_{3}$. In fancier terms the set of rotations forms a group under matrix multiplication.

The results of our previous discussions now suggest some new notation that may ad us in keeping up with the multiplicity of coordinate systems that must usually be dealt with in analysis of spacecraft rotational dynamies. To remain completely general, let us consider three coordinate frames $A, B, C$. We shet let the symbol [BA] represent the rotation matrix which traneforms vector expressed in the A frame to a vector expressed in the 13 frame.

$$
\begin{equation*}
\underline{v}^{(B)}=[B A] \underline{v}^{(A)} \tag{E56}
\end{equation*}
$$

For convenience we use the notation superscript (A) or (B) etc. to indicate which coordinate franes the vectors are being expressed in. If the superscripts are not specified, it means that the coordinate frame is implicit in the definition of the symbol or that it doesn't matter as long. as all vectors are in the same frame. We are here more interested in the rotations [BA] etc. There are three rotations between pairs: $\mathrm{AB}, \mathrm{BC}$, CA. From our previous work

$$
[C A]=[C B][B A] .
$$

Corresponding to equation (E57) is a quaternion relation of similar form. First, since $\underline{\underline{V}}^{(\mathrm{C})}=[\mathrm{CA}] \underline{\underline{V}}^{(\mathrm{A})}$, we have

$$
\begin{equation*}
\underline{\mathbf{v}}^{(\mathrm{C})}=Q_{\mathrm{CA}} * \underline{\underline{V}}^{(\mathrm{A})} \circ Q_{\mathrm{CA}} \tag{E58}
\end{equation*}
$$

Here, $\mathbb{Q}_{\mathrm{CA}}$ is the quaternion corresponding to [CA]. Thus analogous to equation (E57) we have

$$
\begin{equation*}
\underline{\mathbf{V}}^{(\mathrm{C})}=\mathcal{Q}_{\mathrm{CB}} * \circ \mathbb{Q}_{\mathrm{BA}} * \circ \underline{\mathrm{~V}}^{(\mathrm{A})} \circ \mathbb{Q}_{\mathrm{BA}} \circ \mathbb{Q}_{\mathrm{CB}} \tag{E59}
\end{equation*}
$$

Interestingly, we see that $\mathbb{Q}_{\mathrm{CA}}=\mathbb{Q}_{\mathrm{BA}} \circ \mathrm{Q}_{\mathrm{CB}}$ so that the factors occur in reverse order from equation (E57). However, if we use the doublobar operator we can multiply in the same order, i.e.

$$
\begin{equation*}
Q_{\mathrm{CA}}=Q_{\mathrm{CB}} Q_{\mathrm{BA}} \tag{E60}
\end{equation*}
$$

This result is the one which we wish to use analogous to equation (E57). Equations (E57) and (E60) have an casily remembered form and in fact behave as if multiplication cancelled the terms appearing on the inside. This makes it quite easy to construct chains of transformations to any desired system. In this notation we see that

$$
\begin{align*}
& {[B A]=[A B]^{T}=[A B]^{-1} ; \text { also }} \\
& \mathbb{Q}_{B A}=Q_{A B}{ }^{*}=Q_{A B}^{-1} \tag{E61}
\end{align*}
$$

Finally, there are some useful tricks with the new notation we have defined. Referring to equation (E28) and adding the subscripts we have defined, we have $\dot{Q}_{B A}=0.5 Q_{B A} \circ{ }_{B A}{ }^{(B)}$. The vector $\underline{\underline{\omega}}_{B A}{ }^{(B)}$ is the angular velocity of $B$ relative to $A$ with components in $B$. Consider the quaternion $\mathcal{Q}_{\mathrm{CB}}$.

$$
\begin{aligned}
& =\frac{1}{2} Q_{C B} \circ{\underset{C A}{ }}_{(C)}^{=}=\frac{1}{2} Q_{C B} \circ\left(Q_{C B}^{*} \circ \underline{@}_{B A}^{(B)} \circ Q_{C B}\right) \\
& =\frac{1}{2} Q_{C B} \circ \underline{C_{C A}}{ }^{(C)}-\frac{1}{2} \underline{u}_{\mathrm{BA}}{ }^{(B)} \circ Q_{\mathrm{CB}}
\end{aligned}
$$

or in matrix form

$$
\begin{equation*}
\dot{Q}_{\mathrm{CB}}=\frac{1}{2}\left(\frac{-}{4}(\mathrm{CA}) \mathrm{B}_{\mathrm{BA}}^{(\mathrm{B})}\right) Q_{\mathrm{CB}} \tag{E62}
\end{equation*}
$$

The utility of equation (E62) is most apparent when we use it to compute the attitude error of a spacecraft relative to a moving or moveable roference. Note that the components of $\omega_{C A}^{(C)}$ and $\underline{H B A}^{(B)}$ are expressed in different frames. Normally, ${ }^{4} \mathrm{CA}(\mathrm{C})$ would come from rate sensors which are body fixed while ${ }_{B A}{ }^{(B)}$ is a commanded maneuver rate which is naturally defined in the moveable reference. This equation then allows us to use both quantities directly without either being transformed.

It often becomes necessary to compute the quaternion corresponding to a given rotation matrix, i.e., find $Q$ given [BA]. We have developed n computer algorithm to do this.

1) Define matrix

$A$ is the given rotatea.
2) $\mathrm{S}^{\prime}=\mathrm{S}+\mathrm{S}^{\mathrm{T}}+(1-\operatorname{tr} \mathrm{A}) \mathrm{ll}$.
3) $I=\max S^{\prime}{ }_{i i}$ (index of largest element along diagonal of $S^{\prime}$ ).
4) $Q \mathbf{Q j}=\mathrm{S}_{\mathrm{Ij}} / 2 \sqrt{\mathrm{~S}_{\mathrm{II}}}$.
5) $Q^{\prime} j=Q j \operatorname{sgn} Q 4 ; \operatorname{sgn}=\left\{\begin{array}{l}-1 \text { for } Q 4<0 \\ +1 \text { for } Q 4 \geq 0\end{array}\right.$.

Another useful and perhaps obvious technique is the expression of the quaternion resulting from a sequence of Euler rotations (rotations about coordinate axes):

$$
\begin{aligned}
\boldsymbol{Q}_{\mathrm{BA}}= & \left(\mathrm{c} \frac{\phi_{1}}{2}+\underline{u}_{1} \mathrm{~s} \frac{\phi_{1}}{2}\right) \circ\left(\mathrm{c} \cdot \frac{\phi_{2}}{2}+\underline{u}_{2} \mathrm{~s} \frac{\phi_{2}}{2}\right) \circ \ldots \\
& \circ\left(\mathrm{c} \frac{\phi_{\mathrm{n}}}{\underline{2}}+\underline{u}_{n} \mathrm{~s} \frac{\phi_{\mathrm{n}}}{2}\right),
\end{aligned}
$$

where $s \triangle \sin$ and $\mathrm{c} \Lambda$ cos. The corresponding rotation is [BA] and is given by

$$
[B A]=\left[\phi_{n}\right]_{i_{n}} \cdots\left[\phi_{2}\right]_{\dot{j}_{2}}\left[\phi_{1}\right]_{i_{1}}
$$

The vectors $u$ can be any of the three coordinate axes $[1,0,0]{ }^{\mathbf{T}},[0,1,0]{ }^{\mathbf{T}}$ or $[0,0,1]^{T}$. If $u=[1,0,0]^{\prime}$, then $i_{1}=1$, etc. We have added the convention that a rotation bracket with a subscript is an Euler rotation about the indicated axis. As an example consider the quaternion formed when $\mathrm{i}_{1}=1, \mathrm{i}_{2}=2, \mathrm{i}_{3}=3$ :

$$
\begin{aligned}
\mathcal{Q}_{\mathrm{BA}} & =\left(\mathrm{c} \frac{\phi_{1}}{2}+\mathrm{is} \frac{\phi_{1}}{2}\right) \circ\left(\mathrm{c} \frac{\phi_{2}}{2}+\mathrm{js} \frac{\phi_{2}}{2}\right) \circ\left(\mathrm{c} \frac{\phi_{3}}{2}+\mathrm{ks} \frac{\phi_{3}}{2}\right) \\
& =\mathrm{c} \frac{\phi_{1}}{2} \mathrm{c} \frac{\phi_{2}}{2} \mathrm{c} \frac{\phi_{3}}{2}-\mathrm{s} \frac{\phi_{1}}{2} \mathrm{~s} \frac{\phi_{2}}{2} \mathrm{~s} \frac{\phi_{3}}{2}
\end{aligned}
$$

$$
\begin{aligned}
& +\mathrm{i}\left(\mathrm{~s} \frac{\phi_{1}}{2} \mathrm{c} \frac{\phi_{2}}{2} \mathrm{c} \frac{\phi_{3}}{2}+\mathrm{c} \frac{\phi_{1}}{2} \mathrm{~s} \frac{\phi_{2}}{2} \mathrm{~s} \frac{\phi_{3}}{2}\right) \\
& +\mathrm{j}\left(\mathrm{c} \frac{\phi_{1}}{2} \mathrm{~s} \frac{\phi_{2}}{2} \mathrm{c} \frac{\phi_{3}}{2}-\mathrm{s} \frac{\phi_{1}}{2} \mathrm{c} \frac{\phi_{2}}{2} \mathrm{~s} \frac{\phi_{3}}{2}\right) \\
& +\mathrm{k}\left(\mathrm{c} \frac{\phi_{1}}{2} \mathrm{c} \frac{\phi_{2}}{2} \mathrm{~s} \frac{\phi_{3}}{2}+\mathrm{s} \frac{\phi_{1}}{2} \mathrm{~s} \frac{\phi_{2}}{2} \mathrm{c} \frac{\phi_{3}}{2}\right)
\end{aligned}
$$

the corresponding [BA] is

$$
\left[\begin{array}{c:c:c}
c \phi_{2} \mathrm{c} \phi_{3} & \mathrm{~s} \phi_{1} \mathrm{~s} \phi_{2} \mathrm{c} \phi_{3}+\mathrm{c} \phi_{1} \mathrm{~s} \phi_{3} & -\mathrm{c} \phi_{1} \mathrm{~s} \phi_{2} \mathrm{c} \phi_{3}+\mathrm{s} \phi_{1} \mathrm{~s} \phi_{3} \\
-\mathrm{c} \phi_{2} \mathrm{~s} \phi_{3} & -\mathrm{s} \phi_{1} \mathrm{~s} \phi_{2} \mathrm{~s} \phi_{3}+\mathrm{c} \phi_{1} \mathrm{c} \phi_{3} & \mathrm{c} \phi_{1} \mathrm{~s} \phi_{2} \mathrm{~s} \phi_{3}+\mathrm{s} \phi_{1} \mathrm{c} \phi_{3} \\
\mathrm{~s} \phi_{2} & -\mathrm{s} \phi_{1} \mathrm{c} \phi_{2} & \mathrm{c} \\
& \mathrm{c} \phi_{2}
\end{array}\right]
$$

In this brief exposition, we have developed a number of useful quaternion results and notations. This by no means exhauste the possibilities, The available quaternion literature does not present the material in an easily applicable form and thus this short development is presented to fill that gap.

NOTLES: The following notes apply to the previous discussion:

1) We use Q4 rather than 0 for convenience. Since these quaternion equations will be adapted foz the computer and since 0 is not usually allowed as a subscript it becomes necessary to use something else. We desire to use $1,2,3$ for the vector components, hence $Q 4$ is the real part;
2) The symbol $\approx$ is called double tilde and the symbol $\approx$ is called double bar;
3) We shall define

$$
\tilde{Q}=\left[\begin{array}{ccc}
0 & -Q 3 & Q 2 \\
Q 3 & 0 & -Q 1 \\
-Q 2 & Q 1 & 0
\end{array}\right]
$$

which is the tilde or cross product matrix for the 3 -vector $\underline{Q}$.

## APPENDIX F, ROLL COMMAND $\eta_{\text {xn }}$

To use the acquisition sun sensor (ACQ SS) for a two axes strapdown update, the sun line has to nominnlly pass through the venter of the ACQ SS once per orbit, no matter what the tilting angle is. This can be done by rolling the vehicle about the principal xaxis by the angle ${ }^{\text {xn }}$ such that the vehicle $z$ axis always has an elevation angle $\eta_{x}$ above the orbital plane, equal to the elevation angle $\eta_{x}$ of the sun. The transformation from the $L$ system to. the $V$ system is ( $s=\sin , \mathrm{c}=\cos$ )

from which we dexive

$$
V L_{32}=-s \eta_{x}=K_{13} s \eta_{z n}+K_{23} \text { on } n_{x n} e \eta_{z n}-K_{33} s \eta_{x n} c \eta_{z n}
$$

or

$$
K_{33^{s \eta_{x n}}} / R-K_{23^{\mathrm{C}} \eta_{x n}} / \mathrm{R}: \eta_{x p}
$$

with

$$
\eta_{x p}:=\left(s \eta_{x}+K_{13} s \eta_{g m}\right)\left(\left(R e \eta_{q m}\right) \text { and } R=\sqrt{K_{23}{ }^{2}+K_{33}^{2}} .\right.
$$

This yields

$$
\eta_{\mathrm{xn}}=\arctan \left\{\mathrm{K}_{23} / \mathrm{K}_{33}\right\}+\arctan \left\{n_{x p} / \sqrt{1-n_{x p}^{2}}\right\}
$$

Inverse tangents are used since no other ARC functions are available on board. The values of the K's are given in APPENDIX G. They apply for the EOVV A orientation. For the EOVV $B$ orientation we have

$$
[K]_{B}=\left[\begin{array}{rrr}
-1 & 0 & 0 \\
0 & -1 & 0 \\
0 & 0 & 1
\end{array}\right]\left[{ }^{[K]} A\right.
$$

i.e., $K_{13}$ and $K_{23}$ change sign.

## APPENDIX G. DATA

Vehicle moments of inertin matrix $\left(\mathrm{kgm}^{2}\right)$

$$
I_{v}=\left[\begin{array}{rrr}
894828 & -63414 & -529360 \\
-63414 & 3763111 & -27295 \\
-529360 & -27295 & 3598005
\end{array}\right]
$$

Principal moments-of-inertia matrix ( $\mathrm{kgm}^{2}$ )

$$
I_{p}=\left[\begin{array}{rrr}
793332 & 0 & 0 \\
0 & 3767879 & 0 \\
0 & 0 & 3694732
\end{array}\right]
$$

Transformation from vehicle to principal coordinate system

$$
[K]=[P V]=\left[\begin{array}{rrr}
0.982357 & 0.022682 & 0.185638 \\
0.017288 & 0.977351 & -0.210913 \\
-0.186213 & 0.210401 & 0.959716
\end{array}\right] .
$$

## APPENDIX H. FLOWCHARTS

## MANEUVER REOUIREMENTS





## STRAPDOWN IJPDATE REQUIREMENTS



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