NASA-CR-167535 1. Kumar Krishen ED6

7 ARL-TR-81-41

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2. MATHEMATICAL MODELING AND SAR SIMULATION MULTIFUNCTION SAR TECHNOLOGY EFFORTS

Final Report under Contract NAS9-16171

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5. NAS9-1617/

6 10 September 1981

Final Report

1 August 1980 - 31 July 1981

(NASA-CR-167535) MATHEMATICAL MODELING AND SAR SIMULATION MULTIFUNCTION SAR TECHNOLOGY EFFORTS Final Report, 1 Aug. 1980 - 31 Jul. 1981 (Texas Univ. at Arlington.) 60 p HC A04/MF A01 CSCL 171 G3/32

N82-19408

Unclas U9230

Prepared for:

NATIONAL AERONAUTICS AND SPACE ADMINISTRATION LYNDON B. JOHNSON SPACE CENTER HOUSTON, TX 77058





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SECURITY CLASSIFICATION OF THIS PAGE (When Data Entered)

REPORT DOCUMENTATION PAGE		READ INSTRUCTIONS BEFORE COMPLETING FORM		
1. REPORT NUMBER	2. GOVT ACCESSION NJ.	3. RECIPIENT'S CATALOG NUMBER		
MATHEMATICAL MODELING AND SAR SIMULATION MULTIFUNCTION SAR TECHNOLOGY EFFORTS		5. TYPE OF REPORT & PERIOD COVERED final report 1 August 1980 - 31 July 1981 6. PERFORMING ORG. REPORT NUMBER		
		ARL-TR-81-41		
7. AUTHOR(*) Carroll R. Griffin James M. Estes		NAS9-16171		
9. PERFORMING ORGANIZATION NAME AND ADDRESS Applied Research Laboratories The University of Texas at Austin Austin, Texas 78712		10. PROGRAM ELEMENT, PROJECT, TASK AREA & WORK UNIT NUMBERS		
11. CONTROLLING OFFICE NAME AND ADDRESS		12. REPORT DATE		
National Aeronautics and Space Adm Lyndon B. Johnson Space Center Houston, TX 77058	inistration	10 September 1981 13. NUMBER OF PAGES 57		
14. MONITORING AGENCY NAME & ADDRESS(II different	t from Controlling Office)	15. SECURITY CLASS. (of this report)		
		UNCLASSIFIED		
		15a. DECLASSIFICATION/DOWNGRADING SCHEDULE		
17. DISTRIBUTION STATEMENT (of the abatract entered in Block 20, if different from Report)				
IA. SUPPLEMENTARY NOTES				
19. KEY WORDS (Continue on reverse side if necessary and synthetic aperture radar orbiting platform SAR simulation				
The orbital SAR (synthetic aper developed by ARL:UT was used in sevadvanced SAR development. This represent operational radar designed by of antenna polarization effects, and different wavelengths. Avenues for application to the development of a schemes are indicated.	ture radar) simul veral simulation port details effo the NASA Jet Prop nd simulation of r improvements in	efforts directed toward orts toward simulating a pulsion Lab, simulation SAR images at several a the OSS and its		

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TABLE OF CONTENTS

			<u>Page</u>
LIST	OF F	IGURES	v
LIST	OF T	ABLES	vii
I.	INTR	ODUCTION	1
II.	ASAR	SIMULATION EFFORTS	3
	A.	Discussion	3
	В.	Dynamics Simulation	3
	c.	Radar Simulation	5
	D.	Antenna Simulation	8
	E.	Terrain Specifications	8
III.	OTHE	R SIMULATION INVESTIGATIONS	13
	A.	Antenna Polarization Effect	13
	В.	Frequency Diversity Effects	25
IV.	TECH	NICAL EFFORT SUMMARY	31

LIST OF FIGURES

Figure	<u>Title</u>	Page
1	Simulation Architecture	14
2	Imagery from ARL:UT Synthetic Aperture Radar (SAR) Math Model	15
3	Polarization Ellipse	17
4	APQ-102 Antenna Simulation Patterns	21
5	OSS Simulation of Six-Scatterer Array Using SEASAT System with APQ-102 Antenna Pattern in Azimuth, No Pulse Compression	22
6	OSS Simulation of Six-Scatterer Array Using SEASAT System with Unit Antenna Pattern in Azimuth, No Pulse Compression	23
7	Ambiguity Terrain	24
8	Results from SAR Simulation Program	26
9	Test Scene, Wavelength Diversity	28
10	ARL:UT High Resolution Display System	30
11	The Geometry of the Multiple Beam SAR	32

LIST OF TABLES

<u>Table</u>	<u>Title</u>	Page
I	RADAR SPECIFICATIONS RADAR ID: ASAR 1	7
II	ANTENNA SPECIFICATIONS	9
III	TERRAIN SPECIFICATIONS	10
IV	PARTIAL RESULTS OF FREQUENCY DIVERSITY STUDY	29

I. INTRODUCTION

The contract was initially oriented toward simulating the radar parameters of the ASAR (Advanced Synthetic Aperture Radar) system under development jointly by NASA/JPL and NASA/JSC. The former had been tasked to develop a system using a technique known as "burst mode" with wavelength diversity in pulse bursts of transmissions. NASA/JSC was responsible for developing the antenna system required for handling the wide band of frequencies involved. The ARL:UT supporting effort was to simulate the system design and particularly the "burst mode", to provide insight into the operation of the algorithms and anticipated results. The simulation developed by ARL:UT for an orbiting SAR platform had to be modified to support the aircraft based ASAR.

After three months of effort toward these objectives, a decision was made to stop further work on the ASAR system and instead concentrate on certain key SAR technology areas. Among these are multipolarization antennas, multifrequency radar, multibeam squint mode SAR (wide swath), calibration techniques, and turst mode implementation. Of these, ARL:UT addressed the first three in varying degree.

II. ASAR SIMULATION EFFORTS

A. Discussion

The ARL:UT orbital SAR simulation (OSS) had to be modified to simulate an aircraft flying in earth's atmosphere at a relatively low velocity and altitude compared to the orbiting platform. In addition, the radar parameters had to be obtained from the syster designers at NASA/JPL and the antenna parameters from the project personnel at NASA/JSC. The system was to have exhibited two unique design features: extremely broad wavelength coverage, and data processing of synthetic data arrays formed by bursts of pulses at several different radar frequencies.

The purposes of the simulation were to investigate and demonstrate the validity of the system design, in particular the so-called "burst mode" operation.

Two sets of problems were encountered in this effort. The first had to do with adapting the existing simulation, of a space-based orbiting radar, to the aircraft-based design. This problem set was rather easily solved, but attempts to define the design parameters of the radar system were relatively unsuccessful initially, and two or three months elapsed before a preliminary set of parameters for use in the simulation was agreed upon and obtained.

B. Dynamics Simulation

In order to utilize the existing routines in the simulation, the proposed aircraft platform (a CV 990) was considered like an orbiting platform in a low earth gravity environment, so that

it could maintain orbit at its usual altitude and velocity. The magnitude of the velocity vector is

$$V_s = a(1-e^2\cos^2 E)^{1/2} \cdot \dot{E}$$
 , (1)

with the altitude specified by the orbit radius from earth's center

$$r_{s} = a(1 - e \cos E) \qquad . \tag{2}$$

The gravitational constant of earth μ_p has been measured at 398,601 km³/sec². In Eq. (2) e is the eccentricity of the orbit, and E is the eccentric anomaly. If e is zero, the orbit is circular. The parameter a is the semimajor axis of the orbit ellipse, or the radius of the orbit with e=0.

By selecting the proper values for the six Kepler orbital elements, a, e, i, Ω , ω , and T_p , the orbit is specified. In Eq. (1), \dot{E} is the time derivative of the eccentric anomaly:

$$E - e \sin E = v(t-T_p)$$
 (3)

$$\frac{dE}{dt} (1 - e \cos E) = v \qquad . \tag{4}$$

Here v is the mean angular velocity of the platform

$$v = \sqrt{\mu_{\rm p}/a^3} \qquad . \tag{5}$$

In order to derive a usable value of $\mu_{\mbox{\scriptsize p}}$ we can simplify the operation by assuming a spherical earth and a circular orbit for the aircraft. We find that

$$r_g = a$$
 , $v_g = a \times \dot{E}$, $\dot{E} = v$; (6)

therefore

$$v_s = r_s v = a \left(\mu_p / a^3 \right)^{1/2} = \left(\mu_p / a \right)^{1/2}$$
 (7)

Assuming that the CV 990 will fly at 400 kt TAS (205 m/sec) at an altitude of 12,000 m (39,372 ft), we can solve for an imaginary value of $\mu_{\rm p}$ which will provide the proper angular rate.

$$\mu_{\rm p} = aV_{\rm s}^2$$
 (8)
= $(0.205)^2$ (12+6378.167)
= $268.2 \text{ km}^3/\text{sec}^2$

compared with the measured value for earth of

$$\mu_{\rm p} = 398,601 \, {\rm km}^3/{\rm sec}^2$$

All other planet earth parameters remain the same.

For the orbit specification, all parameters used for SEASAT, for example, can remain the same except the semimajor axis a, which becomes 6390.167 km (12+6378.167) and the orbit rotational rate n, which becomes

$$\frac{1}{2\pi} \frac{v_s}{r_s} = \frac{1}{2\pi} \left(\frac{0.205}{6390.167} \right) = 18.4 \times 10^{-4} \text{ deg/sec}$$

C. Radar Simulation

The proposed ASAR design was to have used linear FM (chirp) pulse compression, whereas the OSS is designed for binary phase coded pulse compression, which is particularly adapted to digital processing.

Some effort was spent in analyzing the pulse compression routine requirements to provide an analogous digital signal sampled at the Nyquist rate.

Following a specification review at JSC, a first attempt was made to assign radar specification values for the simulation. Table I contains most of these values.

The linear FM chirp modulation proposed for the ASAR has the following parameters.

$$f_0 = 10^9 \text{ Hz}$$
; $f_1 = 1.02 \times 10^9 \text{ Hz}$, $\frac{df}{dt} = 10^{12} \text{ Hz/sec}^2$
 $T = 20 \times 10^{-6} \text{ sec}$.

The equation for the phase as a function of time is

$$\phi = \phi_{0} + f_{0}t + \frac{1}{2}\frac{df}{dt}t^{2} , \quad \phi_{0} = 0 \text{ rad}$$

$$at \quad t = 0 , \quad \phi = \phi_{0} = 0$$

$$at \quad t = n , \quad \phi_{n} = \phi_{0} + f_{0}(n) + \frac{1}{2}\frac{df}{dt}(n)^{2}$$

$$at \quad t = n-1 , \quad \phi_{n-1} = \phi_{0} + f_{0}(n-1) + \frac{1}{2}\frac{df}{dt}(n-1)^{2} .$$
(9)

If
$$\Delta t = n - (n-1) = 1$$
, $\Delta \phi = \phi_n - \phi_{n-1} = f_0 + \frac{1}{2} \frac{df}{dt}$ (2n-1).

From the sampling theorem, sampling period = 1/2 period of highest frequency,

let
$$\Delta t = \frac{1}{2f_1} [n-(n-1)] = \frac{1}{2f_1} = \frac{1}{2.04 \times 10^9 \text{ Hz}} = 0.49 \text{ nsec}$$
, (10)

then

$$\phi_{n} = \phi_{o} + n\Delta t f_{o} + \frac{1}{2} \frac{df}{dt} (n\Delta t)^{2}$$
(11)

$$\phi_{n-1} = \phi_0 + (n-1)\Delta t f_0 + \frac{1}{2} \frac{df}{dt} (n-1)^2 \Delta t^2$$
 (12)

TABLE I

Radar Specifications Radar ID: ASAR 1

	λ	WL	Radar wavelength, m $\frac{L}{0.23513}$ $\frac{S}{0.093685}$	C 0.062457	X 0.037
	TB	TBW	Range time-bandwidth product	400	
		S/N	Received signal-to-noise ratio, dB	7.0	
	N _P	LRC	Sample length of range correlation 40,	800	
	n _s	LPI	Sample length across phase interval	1	
	ρr	RESR	Ground range resolution, m 25	$\sqrt{2}$	
	ρ _a	RESA	Azimuth resolution, m	2	
	s _r	SRR	Range sampling ratio	1	
		KPC	Phase code, $\frac{1}{2} = 10^{-3} \pi \left[0.98 + (2n-1) \ 0.24 \right]$	× 10 ⁻³] rad	ı
		KPCS	Phase code sequence, BCD	1	
	s _a	SRA	Azimuth sampling ratio	1	
t		KRFTNR	Name of range impulse response function,	BCD: cosin	ie ²
k		KWTFTNA	Name of aperture weighting function, BCD	Taylo	r
k		SHADFAC	Aperture weight factor	-30 d	В
	POR	PCOFFR	Patch-to-patch offset, range, m	0	
	POA	PCOFFA	Patch-to-patch offset, azimuth, m	0	
	SW	SW	Swath width, km	26	
		NA	Number of patches	1	
	LAT	STLAT	Map start latitude, rad	2π/10	
	r _t	TCR	Planetocentric distance to terrain center 0.63	, km 675192370 ×	104
	LTt	TCLAT	Terrain center latitude, rad	2π/10	
	LN_	TCLONG	Terrain center longitude, rad	π	

$$\Delta \phi_{n} = \phi_{n} - \phi_{n-1} = \Lambda t f_{o} + \frac{1}{2} \frac{df}{dt} \left[(n\Delta t^{2}) - (n\Delta t)^{2} + (2n-1)\Delta t^{2} \right]$$

$$= \Delta t f_{o} + \frac{1}{2} \frac{df}{dt} (2n-1)\Delta t^{2} \text{ cycles}$$

$$= 2\pi f_{o}\Delta t + \pi \frac{df}{dt} (2n-1)\Delta t^{2} \text{ rad}$$

$$= 2\pi (10^{6}) \left(0.49 \times 10^{-9} \right) + \pi \frac{20 \times 10^{6}}{20 \times 10^{-6}} (2n-1) \left(0.49 \times 10^{-9} \right)^{2}$$

$$= 10^{-3} \pi \left[0.98 + (2n-1) \left(0.24 \cdot 10^{-3} \right) \right] \text{rad}$$

$$= 0.18 \left[0.98 + (2n-1) \left(0.00024 \right) \right] \text{ deg}$$
(13)

and

$$N_p = \frac{T}{\Delta t} = (2.04 \times 10^9)(20 \times 10^{-6}) = 40,800 \text{ samples.}$$

D. Antenna Simulation

The antenna design was the responsibility of JSC. It posed a special problem because a single antenna for the 10 GHz bandwidth (2-12 GHz) would be very difficult to design and build. Secondly, its design largely depended on the JPL radar design, which was not finalized. Table II gives the parameters that were selected for use in the absence of any real antenna characteristics.

E. Terrain Specifications

The terrain to be used initially was to be a model terrain used in previous work, called SINGLESCAT. It contains two discrete scatterers and a single homogeneous field. Table III describes the parameters for this model.

TABLE II

ANTENNA SPECIFICATIONS

*		KANT	Antenna identification, BCD	UNITPAT
*	$n_{\overline{B}}$	SNAC	Boresight nadir at map start, rad	20°,π/9
*	φ _B	SQUINT	Boresight squint at map start, rad	90°,π/2
*	φ az	ABW	Antenna azimuth angular coverage, rad	π
*	n _{e1}	EBW	Antenna elevation angular coverage, rad	π
*	61	PHCPB(3)	Phase center position, body axis xyz, m	0,0,0
8		ABRPY(3)	Antenna attitude, body axis rpy, rad	$0, -\frac{7\pi}{18}, \frac{\pi}{4}$
*		PDRPY(3)	Platform attitude rates, local orbital rpy, rad/sec	0,0,0

TARLE III

TERRAIN SPECIFICATIONS

KS N O	Terrain identification, BCD	SINGLESCAT
NDISC	Number of discretes in terrain model	2
NFLDS	Number of fields in terrain model	1
NSCATS	Total number of scatterers in terrain model	102
EMEAN	Mean echo strength	0.1
DISTAX	x-axis coverage, km	10
DISTY	y-axis coverage, km	10

A meeting was held at NASA/JPL in the third month of the contract, attended by representatives from ARL:UT and Environmental Research Institute of Michigan (ERIM), to evaluate the design objectives and status of the ASAR. The project had schedule and funding difficulties and, subsequently, NASA decided to cancel the entire project. Unfortunately, a substantial portion of the contract resources at ARL:UT had already been expended on the effort, and a simulation had not been run.

The project was reoriented with specific emphasis on three areas: antenna polarization effects, wavelength diversity effects, and implementation of a so-called "burst mode" wavelength diversity concept on a patch-by-patch basis.

III. OTHER SIMULATION INVESTIGATIONS

Figure 1 illustrates the architecture of the OSS simulation. Each intermediate step results in data on tape or disk bulk storage, which permits incremental completion of the entire simulated radar image. For the polarization and wavelength diversity experiments an existing terrain model was utilized, illustrated in Fig. 2. This figure illustrates the scene content of the test terrain as well as images derived from it.

A. Antenna Polarization Effect

A complete simulation would include the polarization sensitivity of individual scatterers and would incorporate the randomly polarized backscattering from the individual scatterers in fields of scatterers, particularly homogeneous fields. The general situation is that the backscattered wave is partially polarized, part of it completely polarized, and part of it unpolarized. Suppose we designate the degree of polarization as the ratio of the polarized power to the total power (following Kraus), or:

and denoting the Stokes parameters as so, s1, s2, s3,

$$d = \frac{\sqrt{s_1^2 + s_2^2 + s_3^2}}{s_0} \qquad 0 \le d \le 1 \qquad . \tag{16}$$

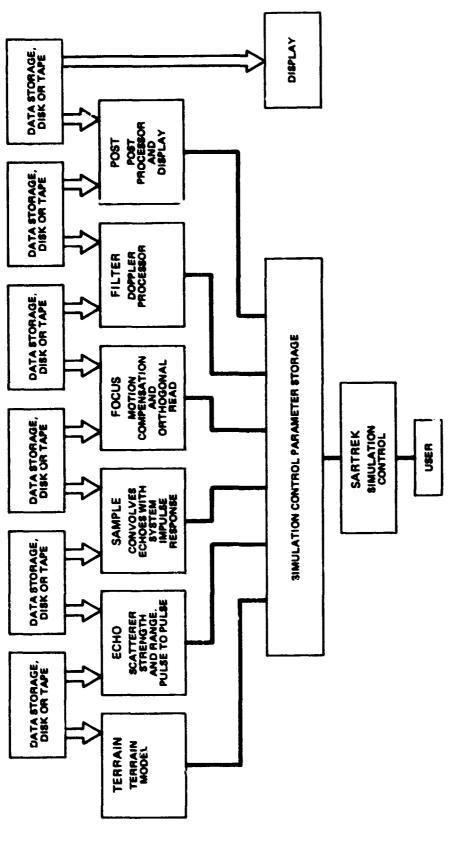


FIGURE 1
SIMULA: ION ARCHITECTURE

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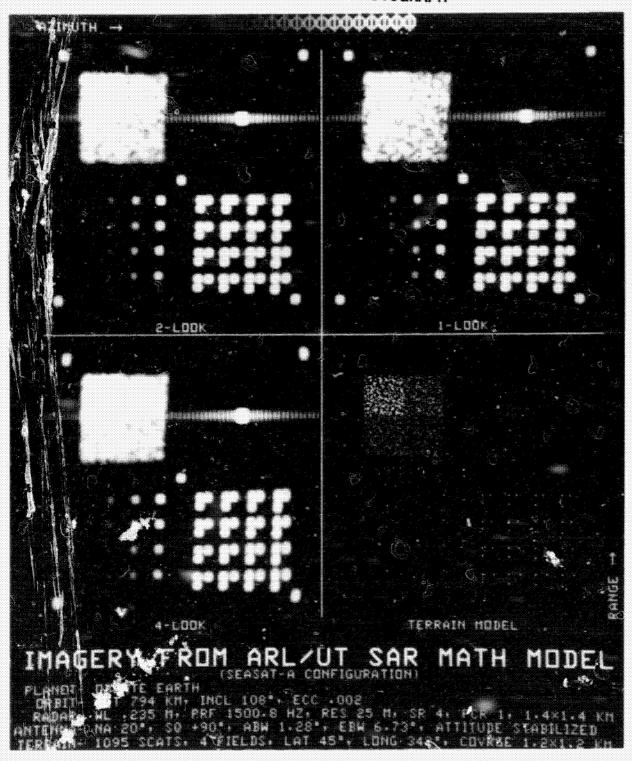


FIGURE 2
IMAGERY FROM ARL:UT SYNTHETIC APERTURE RADAR (SAR) MATH MODEL

These are normalized Stokes parameters, defined as follows, with reference to Fig. 3:

$$Ex = E \cos \sin \omega t \cos \tau - E \sin \varepsilon \cos \omega t \sin \tau$$
 (17)

Ey =
$$E_0 \cos \varepsilon \sin \omega t \sin \tau + E_0 \sin \varepsilon \cos \omega t \cos \tau$$
 (18)

If we now substitute the time dependent values of Ex and Ey we have

$$E_1 \sin(\omega t - \delta_1) = E_0(\cos \epsilon \sin \omega t \cos \tau - \sin \epsilon \cos \omega t \sin \tau)$$
 (19)

$$E_2 \sin(\omega t - \delta_2) = E_0 (\cos \epsilon \sin \omega t \sin \tau + \sin \epsilon \cos \omega t \cos \tau)$$
 (20)

If these are expanded, and sinut terms and cosut terms set equal (which eliminates the time dependence):

$$E_1 = E_0 \left(\cos^2 \epsilon \cos^2 \tau + \sin^2 \epsilon \sin^2 \tau\right)^{1/2}$$
 (21)

$$E_2 = E_0 \left(\cos^2 \varepsilon \sin^2 \tau + \sin^2 \varepsilon \cos^2 \tau\right)^{1/2} . \tag{22}$$

The flux density in W/m^2 is the Poynting vector \overline{S} :

$$|\overline{S}| = S_x + S_y = \frac{E_1^2 + E_2^2}{Z_0} = \frac{E_0^2}{Z_0}$$
, (23)

where Z_0 is the free space impedance, 377 Ω per square measure, S_x is the Poynting vector for the E_0 wave component polarized to x, and S_y , for the one polarized to y.

$$S_{x} = \frac{E_{1}^{2}}{Z_{0}} = |\overline{S}|(\cos^{2}\varepsilon \cos^{2}\tau + \sin^{2}\varepsilon \sin^{2}\tau)$$

$$S_{y} = \frac{E_{2}^{2}}{Z_{0}} = |\overline{S}|(\cos^{2}\varepsilon \sin^{2}\tau + \sin^{2}\varepsilon \cos^{2}\tau) \qquad (24)$$

The Stokes parameters are now defined as

$$s_0 = \frac{S}{S} = 1$$

$$s_1 = \frac{S_x - S_y}{S} = \frac{\langle E_1^2 \rangle - \langle E_2^2 \rangle}{SZ} = \cos 2\varepsilon \cos 2\tau , \qquad (25)$$

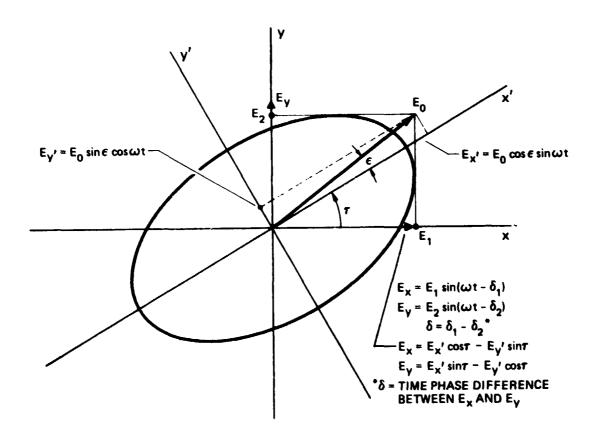


FIGURE 3
POLARIZATION ELLIPSE

ARL:UT AE-81-131 CRG - GA 9 - 16 - 81 where <> denotes time averaged value. Then

$$s_2 = \frac{2}{SZ} \langle E_1 | E_2 \cos \delta \rangle = \langle \cos 2\varepsilon \sin 2\tau \rangle$$

$$s_3 = \frac{2}{SZ} \langle E_1 | E_2 \cos \delta \rangle = \langle \sin 2\varepsilon \rangle$$

$$1 \geq s_1^2 + s_2^2 + s_3^2$$

These can be written in matrix form

$$S[s_{1}] = S\begin{bmatrix} s_{0} \\ s_{1} \\ s_{2} \end{bmatrix}$$

$$= S\begin{bmatrix} 1-d \\ 0 \\ 0 \\ 0 \end{bmatrix} + S\begin{bmatrix} d \\ d \cos 2\varepsilon \cos 2\tau \\ d \cos 2\varepsilon \sin 2\tau \\ d \sin 2\varepsilon \end{bmatrix}$$
(26)

for a partially polarized wave, where the first term of Eq. (27) is the polarized power density and the second is the unpolarized power density.

The antenna polarization response may also be expressed as a matrix, with $\mathbf{A}_{\underline{\mathbf{e}}}$ the effective aperture:

$$A_{e}[a_{1}] = A_{e}\begin{bmatrix} a_{0} \\ a_{1} \\ a_{2} \\ a_{3} \end{bmatrix} m^{2}$$
(28)

Thus the power out of the antenna is given by

$$W = \frac{1}{2} S A_{e} [a_{i}]_{t} [s_{i}]$$

$$= \frac{1}{2} S A_{e} \sum_{i=0}^{3} a_{i} s_{i} W/Hz$$
(29)

The simulation for the APQ-102 uses a linear polarized antenna, with a gain given by

$$G = \frac{4\pi A}{\lambda^2} \qquad , \tag{30}$$

and the magnitude of the gain function can be entered into the simulation via program ANTENA. Substituting the matrix representation for the ${\rm A}_{\rm e}$ we have

$$G[a_{\underline{i}}] = \frac{4\pi}{\lambda^2} A_{\underline{e}}[a_{\underline{i}}] = \frac{4\pi}{\lambda} A_{\underline{e}}\begin{bmatrix} a_{\underline{o}} \\ a_{\underline{1}} \\ a_{\underline{2}} \\ a_{\underline{3}} \end{bmatrix}$$
(31)

Combining the polarization matrix for a scatterer, we have

$$G[a_{1},s_{1}] = \frac{4\pi}{\lambda^{2}} \frac{|\sigma|}{2} [a_{0}a_{1}a_{2}a_{3}] \begin{bmatrix} s_{0} \\ s_{1} \\ s_{2} \\ s_{3} \end{bmatrix}$$
(32)

where the substitution of radar cross-section σ is made for the power density S, assuming that the backscattered power density is proportional to the cross section.

To implement this function in program ANTENA requires that the a_1 be specified, as well as the s_1 . The former may reduce to $a_0=1$, $a_1=0$, $a_2=0$, $a_3=0$, or it may depend on platform motion.

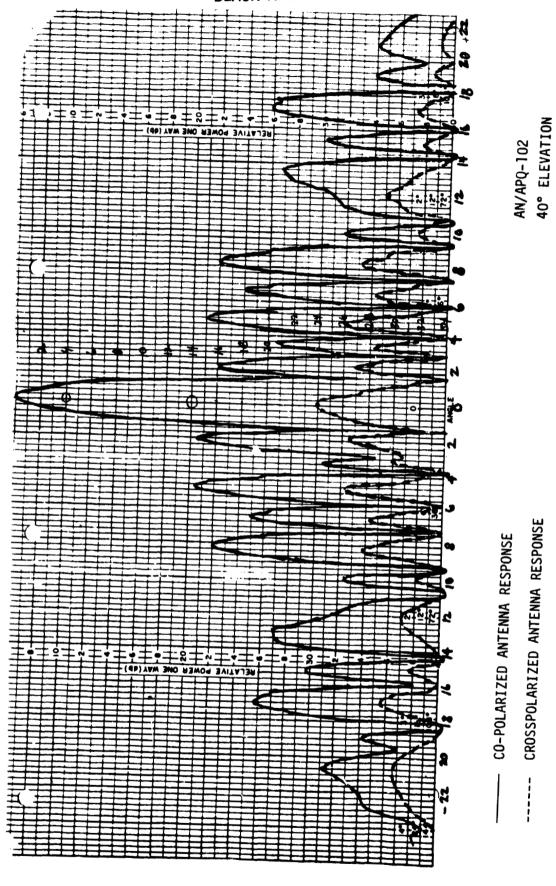
The specification of [s] on the other hand would be arbitrary or would depend on measured data for particular types of radar targets or clutter.

As a first step, it vas decided to investigate the effects of crosspolarized antennas for transmit and receive, and assume no polarization sensitivity for σ . Although the crosspolarized response of the receive antenna has been measured, it was not available in the azimuth plane, only in elevation. Based on the lobe structure and boresight gain of the elevation pattern, a crosspolarized azimuth response pattern was constructed and used to specify ANTENA. Figure 4 illustrates the pattern used.

The values for the patterns, digitized every 0.33° to a coverage of ±24°, were entered using an HP9810A with digitizer. The SEASAT radar parameters were used and six scatterers were entered into the terrain model for test purposes. Five of these were given a cross section of 1 m at the center and corners of the 1.4 x 1.4 km mapped area; the other was 26 dB greater at 20 m^2 . The imulation was exercised and the SARCON data were printed out--see the Appendix. Figure 5 shows the displayed results for the co-polarized antenna response. Post-processing was effected at 2 dB/gray shade, with the peak filter magnitude falling in gray shade 16. Figure 5(a) has gray shade 16 turned on and Fig. 5(b) has it off. To expedite processing no range pulse compression was used; hence range sidelobes are not present. Figure 6 is for the same scatterer array but uses a unit antenna pattern, i.e., it is isotropic. There is no appreciable difference since the entire imaged area falls well within the beam of the APO-102 1.3° 3 dB beamwidth for the orbiter altitude of 870 km.

To demonstrate antenna response, a realistic terrain scene without a prohibitively large scattering area was required and it was necessary to place scatterers at ambiguous (with respect to pulse repetition frequency (PRF)) response points. The total azimuth coverage imaged was only 1.4 km; if the terrain model covered only this area, then from an altitude of 870 km along 4.6 km of the orbit (the synthetic array length), only $\pm 0.25^{\circ}$ of the antenna patterns would be used. To create an effective but simple model of terrain, single scatterers were placed at ambiguous points as illustrated in Fig. 7. Only the single scatterer

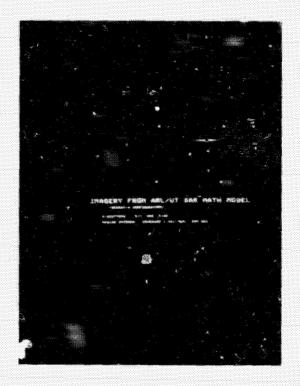
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APQ-102 ANTENNA SIMULATION PATTERNS

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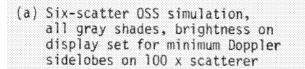


- (a) Six-scatter OSS simulation, all gray shades, brightness on display set for minimum Doppler sidelobes on 100 x scatterer
- (b) Gray shade 16 turned off, brightness increased to include all Doppler filter sidelobes

FIGURE 5 OSS SIMULATION OF S1X-SCATTERER ARRAY USING SEASAT SYSTEM WITH APQ-102 ANTENNA PATTERN IN AZIMUTH, NO PULSE COMPRESSION

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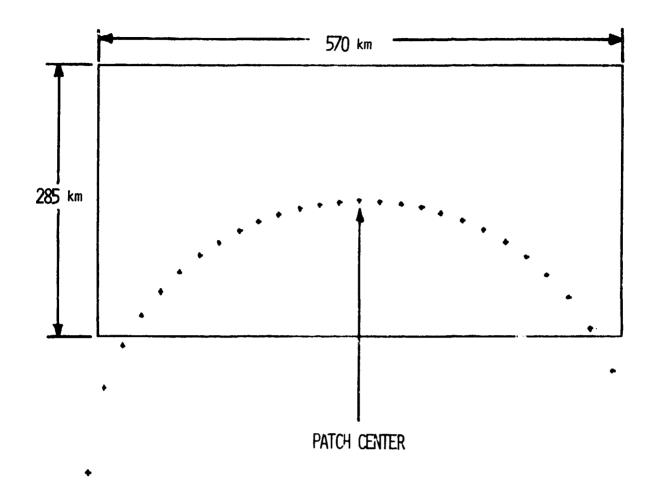






(b) Gray shade 16 turned off, brightness increased to include all Doppler filter sidelobes

FIGURE 6
OSS SIMULATION OF SIX-SCATTERER ARRAY USING SEASAT SYSTEM
WITH UNIT ANTENNA PATTERN IN AZIMUTH,
NO PULSE COMPRESSION



27 Discrete Scattemers in Model

FIGURE 7
AMBIGUITY TERRAIN

at the patch center was imaged; however, there were Doppler contributions from all the sidelobe scatterers. Two configurations of scatterers were used; the second one offset the ambiguous point scatterers in graduated steps 10-30 m from the ambiguity locations.

Figure 8 illustrates the results. Each photograph contains six images of the patch center; each of these is 1.4 x 0.4 km, with 25 m resolution. From the top, the first three strips are of the unoffset terrain model imaged with the isotropic, co-polarized, and crosspolarized patterns. All three patterns were normalized to 1.0 at the peak response. The 1.3t three from the top on the left are images of the offset scatterers for co-polarized, crosspolarized, and isotropic patterns. The top gray shade of the display is set down 20 dB from the peak response of the center pixel to show the weak sidelobe structure.

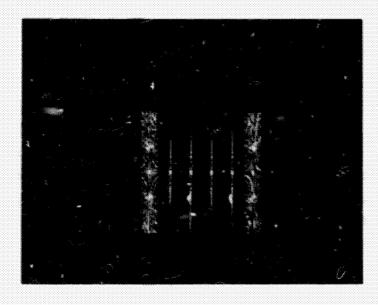
The right photo has the same order of images, but the top gray shade is 35 dB below the peak response, thus defining more clearly sidelobe response of the system configurations.

Analysis of these results shows a significant reduction in ambiguities between images with the isotropic antenna pattern and the APQ-102 patterns. The co-polarized pattern image shows no ambiguities and has lower sidelobes than the crosspolarized image. The crosspolarized pattern generates an ambiguity due to its high sidelobes. One must be careful in trying to interpret the power level of the ambiguities, because the energy returned outside the main beam is from only a few scatterers and may be lower than the energy returned from an actual continuous terrain area.

B. Frequency Diversity Effects

It was decided to investigate the image variations with wavelength on the scene of Fig. 2, which has been imaged at the SEASAT wavelength of 23.5 cm. Three other wavels other wavelength were selected: 1.8, 3.125, and 10.34 cm.

ORIGINAL PAGE BLACK AND WHITE PHOTOGRAPH



(b) Top gray shade of display set 35 dB from the imaged pixel at the patch center



(a) Top gray shade of display set 20 dB down from the imaged pixel at the patch center

FISURE B
RESULTS FROM SAR SIMULATION PROGRAM

To properly implement the simulation the variation of the cross section with wavelength should be included. Other wavelength dependent factors are the antenna response patterns, transmitted power, receiving system noise and, indeed, every factor in the radar range equation either implicitly or directly. In the interest of economy and reduced complexity, however, only the effects on the synthetic array processing were taken into account. The results are not very dramatic, as would be the case if, for example, the cross-section dependency were included. This could involve a simple adjustment since, generally speaking, radar cross section varies proportionally to the square of the wavelength. A useful set of relations might be:

$$\sigma_{L} = 1 \text{ m}^{2} = 0 \text{ dB m}^{2}$$

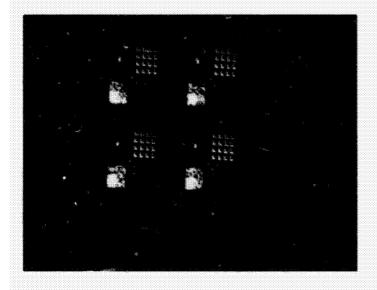
$$\sigma_{S} = \frac{\sigma_{L}(10.34)^{2}}{(23.5)^{2}} = 1.936 \times 10^{-1} \sigma_{L} = -7 \text{ dB m}^{2}$$

$$\sigma_{X} = \frac{\sigma_{L}(3.125)^{2}}{(23.5)^{2}} = 1.768 \times 10^{-2} \sigma_{L} = -17.5 \text{ dB m}^{2}$$

$$\sigma_{Ku} = \frac{\sigma_{L}(1.8)^{2}}{(23.5)^{2}} = 5.867 \times 10^{-3} \sigma_{L} = -24 \text{ dB m}^{2}$$

This is in general true only for perfectly conducting bodies, and is not the general case for clutter targets. For the first simulation effort, only the wavelength parameter was varied, and the resolution cell dimensions were forced to the same values by varying the number of pulses processed or, equivalently, the array length.

Figure 9 presents the four images of the test pattern, photographed with two different exposures. The most obvious effect of the wavelength changes is in the coherent speckle pattern for the homogeneous fields. The data were analyzed statistically, and only very slight changes in the average values were observed. Table IV provides these data for comparison. The simulations were run for the SEASAT radar and the isotropic antenna pattern. The ARL:UT high resolution display system was used to display and photograph the results (Fig. 10).



1-A f8, 1/2 sec exposure ASA 3000 Azimuth |

Range

1-8 +8, 1/4 sec exposure ASA 3000 Azimet

Clockwise From Lower Left: L, S, X, K_y Band TEST SCENE, WAVELENGTH DIVERSITY

FIGURE 9

TABLE IV
PARTIAL RESULTS OF FREQUENCY DIVERSITY STUDY

STATISTICAL DATA

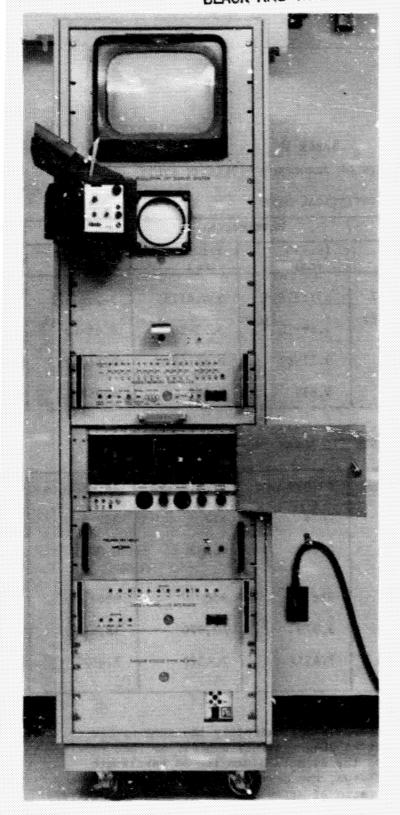
	Wavelength								
Filter Magnitude	23.5 (cm)	10.345 (cm)	3:125 (cm)	1.8 (cm)					
Meximum	336.63147	337.97644	334.47785	342.83377					
Minimum	6.86×10 ⁻¹²	3.69×10 ⁻¹⁴	5.13×10 ⁻¹³	8.95×10 ⁻¹⁵					
Mean	0.18181	0.17562	0.17848	0.180995					
Sigma	4.857	4.859	4.849	4.959					

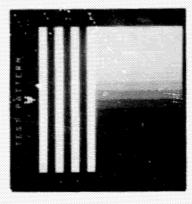
SYNTHETIC ARRAY PARAMETERS

Array No. 1	43,262 pts	43,039 pts	42,816 pts	42,816 pts	
Array Length, m	4602	2026	612	352.5	
Formation Time (sec)	0.617	0.271	0.082	0.047	
No. of Pulses	926	916	900	896	
PRF, Hz	1,500	3,370	10,970	18,948	
Platform Velocity (km/sec)	7. 4577	7.4577	7.4577	7.4577	

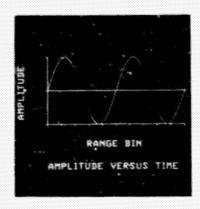
The PRF was increased to obtain approximately the same number of samples (pulses) in array lengths progressively shorter so that resolution would remain the same for the azimuth dimension as wavelength decreased. Had this not been done, the azimuth resolution would have drastically increased as the ratio of array length to wavelength increased.

ORIGINAL PAGE BLACK AND WHITE PHOTOGRAPH





CALIBRATION PATTERN



GRAPHICS



IMAGE CONTOURS FROM SELECTED GREY SHADES

FIGURE 10
ARL:UT HIGH RESOLUTION DISPLAY SYSTEM

IV. TECHNICAL EFFORT SUMMARY

Unfortunately, the dilution of the resources allocated to the simulation efforts by the ASAR project prevented completion of in-depth studies on any of the subjects undertaken. Initial analysis was commenced on a wide swath technique using squinted multiple beams. A major challenge will be the formation of the image from the output of multiple beams tracking the clutter along the velocity vector. Figure 11 illustrates the concept, which will be undertaken in a follow-on effort.

In summary, the ASAR work provided a method for adapting the OSS simulation to an aircraft platform, and to a linear FM (chirped) pulse radar. The analysis indicates the approach to be taken to fully characterize the effects of polarization of the backscattered energy and crosspolarization of the antenna system versus co-polarization of transmit and receive antennas.

Finally, some effects of frequency diversity in the generation of coherent speckle from homogeneous fields were synthesized with the simulation. These efforts point the way to more detailed and fruitful investigations in the future, using the OSS simulation programs.

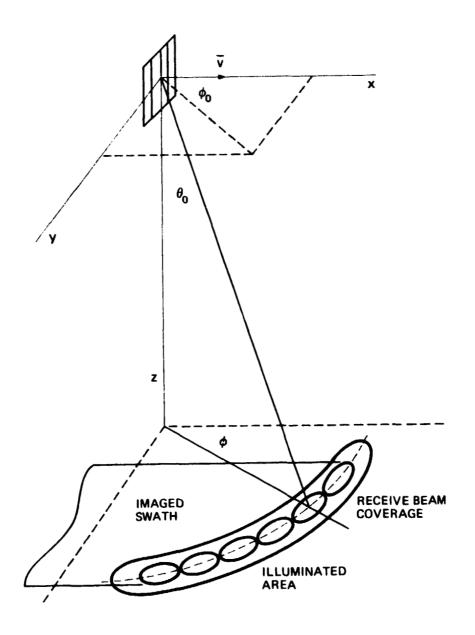


FIGURE 11
THE GEOMETRY OF THE MULTIPLE BEAM SAR

ARL:UT AE-81-137 CRG - GA 9 - 16 - 81 L-BAND WAVELENGTH DATA

INPUT KRRANCH
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.250000000000 + 02 .40000000000 E+01 TAYLOR UNITPAT BORESIGHT NADIR AT TO (DEG) ** .20000000000E+02 .900000000E+02 .900000000E+02 AZIMITH ANGULAR COVERAGE (DEG) ** .1280000000E+01 .6730000000E+01 PHASE CENTER, BODY AXIS X (M) ** 0. PHASE CENTER, BODY AXIS Z (M) ** 0. COORD SYS, BODY AXIS PITCH, (DEG) ** -.700000000E+02 .900000000E+02 PLAT ROLL RATE (DEG/S) ** 0. .63781670000E+04 .71865400000E+04 .10800000000E+03 .100000000000E+01 .62198398135E+02 **500000000E*02 .2350000000E+00 .1200000000E+01 -59376152657E-01 RANDOM ċ . • . • • . • . • • EARTH
-8182017996E-D1 GRAVITATIONAL CONSTANT (XM3/SEC2)
-41780745995E-02 TIME OF PRIME MERIDIAN PASSAGE (S) • • • • • • • • • • • • • • • • • • • • PATCH-TO-PATCH OFFSET. AZ (M) : • • . . . * . • INCLINATION (DEG) **
ARGUMENT OF PERIGEE (DEG)
ROTATIONAL RATE (DEG/S) * TESTPATERN NO OF DISCRETES **

* 12000000000E*01 Y*AXIS COVERAGE (KM) **

*6367519237UE*04 TERRAIN CENTER* LAT (DEG) • * SEMI-MAJOR AXIS (KM) CIFICATION SPECIFICARION ECIFICATION CIFICATION CIFICATION *140000000000000000 NO OF PATCHES .2000000000E-02 SPE --100 J000000E-03 .34191125864E+03 a SPE SPE S 794CIRCIIL LANET オとこよ 2 H A S 8 1 7 0 A R • . ċ • α Œ * SIGNAL-TO-NOISE RATTO (DR) **
SAMPLE LENGTH OF RANGE CORRELATION **
BINARY PHASE CODE **
GROUND RANGE RESOLUTION (M) **
RANGE SAMPLING RATIO (M) **
RANGE IMPULSE RESPONSE FUNCTION **
PATCH-TO-PATCH OFFSFT* RNG (M) ** ۵ * * :: z 0 œ * w * PHASE CENTER. BODY AXIS Y (W) ** COORD SYS. BODY AXIS ROLL. (UEG) ** COORD SYS. BODY AXIS YAW. (DEG) ** PLAT PITCH PATE (DEG/S) *** ⋖ -* :: ANTENNA I.D. **
BOHESTGHT SAUINT AT TO (DEG) **
ELEVATION ANGULAR COVERAGE (DEG) • ORHITER INITIALIZATION TIME (S) * RECIEVER/TRANSMITTER BW (MHZ) LONG OF ASCENDING NODE (DEG) . TEPRAIN T.D. ++
NO OF FIELDS ++
X-AXIS COVERAGE (KM) ++
TERRAIN CENTER, R (KM) ++
TERRAIN CENTER, B (KM) ++ * • . START LATITURE (DEG) . PLANET NAME ** ECCENTRICITY ** ROTATIONAL MATE (NEG/S) RANGE SWATH WIDTH (KM) MAP START LATITUDE (DE • ECCENTRICITY ** ORBIT T.D.

* SYNTHETIC ARRAY PARAMETERS *****	.78561515863E+03	.83463340843E-n]	.74576917166E+01			.539833022R1		.64851836682E+03		■ .22680692212E+02	.71766915525E+04	.43920144547E-02	.33848242890E+03	1
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ARDAW STATISTICS FOR 43262 POINTS.

MAXIMIM 336.6314724528. MINIMIM 6.861199744928E-12

MEAN .1814136300036. SIGMA 4.857006193401

EOF WRITTEN ON CIMAGE . AFTER RECORD 224

S-BAND WAVELENGTH DATA

INPUT KRHANCH
INPUT INTLINF, IFILING
INPUT INTLINF, IFILING
NPUT I/O DATA ON IGDATF
IITPUT IMAGE DATA ON CIMAGE
ITPUT IMAGE DATA ON CIMAGE
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ITPUT IMAGE DATA ON CIMAGE
INTERVAL IS SET AT SIA/I
O. OF FILTERS RECUIRED FOR FULL PRF COVERAGE IS 3180
RESUM RATIO OF 14 IS DERIVED FROM A SAMPLE POINT SAROH TOLERANCE OF .002
O. DF PRESUMMED FIT FILTERS IS 227

.63781670000E.04 .63781670000E.04 .39860100000E.06	. 11865400000E+04 .10800000000E+03 0.	-10345000000E+00 -10000000000E+01 -67208: 79602E+01 -67208: 79602E+01 -2500000000E+02 -4000000000E+01 TAYLOR	.20000000000E.02 .12800000000E.01 0. 0. 7000000000E.02	7] 1095 •12000000000€•01 •45000000000€•02
C I F 1 C A T 1 O N	IFICATION *	OPERATING WAVELENGTH (M) ** OPERATING WAVELENGTH (M) ** RANGE !!ME-BANDWIDTH PRODUCT ** AZO SAMPLE RATE (MHZ) ** SAMPLE LENGTH ACROSS PHASE INTERVAL HINARY PHASE CODE SEQUENCE ** AZIMUITH RESOLUTION (M) ** AZIMUITH SAMPLING RATIO (M) ** APERTUHE WEIGHT FUNCTION ** PATCH-TU-PATCH OFFSET* AZ (M) **	C I F I C A T I O N	C I F I C A T I O N NO OF DISCPETES
EARTH .895E-01 .4178074595E-02	8 1 1 S P E C 794C1RCULR 20003000000000000000000000000000000	SEASATAA • 15552207400E+02 • 10000000000E+03 B • 2500000000E+02 • 400000000E+01 • 2500000000E+01 • 4000000000E+01 • 45000000000E+01	UNITHAT	TESTPATERN 4 -120000000E+01 -6367519237UE+04 -34191125864E+03
PLANET NAME ECCENTRICITY ROTATIONAL RATE (DEG/S)	ORBIT I.D. ** ECCENTRICITY ** LONG OF ASCENDING NODE (DEG) ** TIME OF PERIGE PASSAGE (S ** ORBITER INITIALIZATION TIME(S) **	RADAR 1.D. ** RECIEVER/IRANSMITTER RW (MHZ) ** SIGNAL-IO-NOISE RATIO (DB) ** SAMPLE LFNGTH OF RANGE CORRELATION ** BINARY PHASE CODE ** GROUND RANGE RESOLUTION (M) ** RANGE IMPULSE RESPONSE FUNCTION ** PAICH-IO-PAICH OFFSET* RNS (M) ** RANGE SWATH WIDTH (KM) ** RANGE SWATH WIDTH (KM) **	ANTENNA I.D. ** BORESTGHT SOUTNI AT TO (DEG) ** ELEVATION ANGULAR COVERAGE (DEG) ** COORD SYS, RODY AXIS Y (P) ** COORD SYS, RODY AXIS YAW, (DEG) ** PLAT PITCH RATE (DEG/S) **	TERRAIN 1.D. ** * * * * T E R NO OF FIELDS ** X-AXIS COVERAGE (KM) ** TERRAIN CENTER. R (KM) ** TERRAIN CENTER. R (KM) **

•	.78561515863E+03	.83449265311E-01	745769766778+01	117071385135+04	666	519742501095 + 00	.24095655046E+01	- BABA941290E+01					_	
SYNTHETIC APART PARAMETERS	I TRANSMISSION START TIME (S)	.2C260126547E.04 ARRAY INCLINATION (DEG)	.27166739882E+03 PLAIFORM VFLOCITY (KM/S)	916 PRF (HZ) **	D. OF AZIMUTH FILTERS	ANT HANGE SKATH KIDTH (KM) ++	ANT HANGE SAMPLE INTERVAL (M)	*86822675313E+03 RANGE PATCH CENTER (KM) **	NOIR ANGLE PATCH CENTER (DEG)	S INCIDENCE AT PATCH CENTER (DEG) **	PRITER MASS CENTER. P. (KM)	PATCH CENTER. LAT (DEG) ** .45000000000E+02 ORBITEH MASS CENTER. LAT (DEG) **	PRITEH MASS CENTER. LONG (DEG)	
SYNTHETIC ASKA	1 18	.2C260126547E.04 AF	** .27166739882E+03 PL	910		:		+86822675113E+03 RA	(DEG) ** .901/6516786E+02 NA	0E6) **11282673345E+03()	.63675192370E + 04 OF	***50000000000000000	* * 34191125864F+03 OF	15 .999995288504
• • • •	SYNTHETIC APPAY NO	ARRAY LENGTH (M)		NO OF PULSES	NO OF MANGE SAMPLES	PATCH CENTER RANGE SAMPLE NO	SLANT RANGE BFSOLUTION (M)	START DANGE (KM) ++	SQUINT ANGLE PATCH CENTER (LOS AZIMITH AT PATCH CENTER	PATCH CENTER. R (KM)	PATCH CENTER. LAT (DEG)	PATCH CENTER. LUNG (DEG) *	PATCH CENTER PIXEL POWER 15

ARRAY STATISTICS FOR 43039 POINTS. 4EAN .1756182193471. SIGMA 4.858919965889 Of WPITTEN ON CIMAGE . AFTER RECORD 224

X-BAND WAVELENGTH DATA

INPUT KRHANCH

INPUT KRHANCH

INPUT KRHANCH

INPUT 1/0 DATA ON IODATF

UTPUT 1/0 DATA ON CIMAGE

. 6 .63781670000E.04 . 39860100000E.06	**************************************	-3125000000000000000000000000000000000000	**************************************	71 1095 -12000000000E+01 -+50000000E+02
* * P L A N E T S P E C I F I C A T I O N * * * * * * * * EARTH EQUATORIAL RADIUS (KM) * * * * * * * * * * * * * * * * * * *	+ + 0 P H I T S P E C 1 F 1 C A T I D N + + + + + + + + + + + + + + + + + +	SEASAT-A OPERATING WAVELENGTH (M) ** SEASAT-A OPERATING WAVELENGTH (M) ** ** ** ** ** ** ** ** ** **		TESTPATERN NO OF DISCRETES TESTPATERN NO OF DISCRETES 12000000000000000000000000000000000
PLANET NAME ** ECCENTRICITY ** ROTATIONAL PATE (DEG/S) **	ORBIT 1.D. ** ECCENTRICITY ** LUNG OF ASCENDING NODE (DEG) TIME OF PERTOFE PASSAGE (S) * ORBITER INITIALIZATION TIME(S)	RADAR I.D. ** RECIEVER/TRANSMITTED RW (MHZ) ** SIGNAL-TD-NDISE RATIO (DR) ** SAMPLE LEWGTH OF RANGE CORRELATION BINAHY PHASE CODE ** GHIDUND RANGE RESOLUTION (W) ** RANGE SAMPLING PATIO IN) ** RANGE THPULCE RESPONSE FUNCTION ** PATCH-TO-WATCH OFFSET, RNG (M) ** RANGE SWATH WIDTH (MM) ** RANGE SWATH WIDTH (MM) **	ANTENNA 1.D BORESIGHT SOUINT AT TO (DEG) ELEVATION ANGILLAR COVERAGE (F PHASE CENTED. RODY AXIS Y (W) COORD SYS. RODY AXIS YAW. (CE PLAT PITCH BATE (DEG/S)	TERRAIN I.D NO OF FIF.DS K-AXIS COVERAGE (KW) TERRAIN CENTER, R (KM) TERRAIN CENTER, LONG (DEG)

•	.785615159636+03	.83441538944E-01	.74576947323E+01	.10970162745E+05	223	.53969587071E+00	.24093873367E+01	.85848170559E+03	•			.43405755848E • 02	.33849303127E+03	
SYNTHETIC ARRAY PARAMETERS	I TRANSMISSION START TIME (S)	.61201445587E+03 ARRAY INCLINATION (OFG) ++	. R205482445E+02 PLATFORM VELOCITY (KM/S)	900 PRF (HZ) ++). OF AZIMUTH FILTERS	ANT RANGE SHATE BICTE (KE)	ANT RANGE SAMPLE INTERVAL (M)	. PARS1426692F+03 RANGE PATCH CENTER (KM)	.90053327205E+02 NADIR ANGLE PATCH CENTER (DEG) ++	OS INCIDENCE AT PATCH CENTER (DEG)	HITER MASS CENTER, B (KM) **	PRITER MASS CENTER, LAT (DEG)	.34141125864E+03 ORBITER MASS CENTER. LONG (DEG) **	
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ARDAY STATISTICS FOR 42816 POINTS.
MAXIMUM 5.17358976068E-13
MFAH .!784891654;74. SIGMA 4.844835825272
EGF WPITFN ON CIMARE. AFTER RECORD 224

 $K_{\mathbf{u}}$ -BAND WAVELENGTH DATA

INPUT KBHANCH
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O. OF FILTERS REQUIRED FOR FULL PRF COVERAGE IS 3110
RESUM RATIO OF 13 IS DERIVED FROM A SAMPLE POINT ERRUH TOLERANCE OF .002

	-71865400000E+04 -10800000000E+03 0-59376152657E-01	-16000000000000000000000000000000000000	.128000000000E+01 0.12800000000E+01 070000000000E+02	71 1095 •12000000000000000000000000000000000000
C I F I C A T I O N EQUATORIAL RADIUS (KM)	ECIFICATION * * * * * * * * * * * * * * * * * * *	C I F I C A T I O N • • • • • • • • • • • • • • • • • •	ECIFICATION • • • • • • • • • • • • • • • • • • •	E C I F I C A T I O N
PLANET NAME ** * * * * * * * P L A N E T S P E. ECCENTRICITY ** ** ** ** ** ** ** ** ** ** ** ** **	ORBIT 1.D ECCENTRICITY LONG OF ASCFNDING NODE (DEG) TIME OF PARTSEE PASSAGE (S) ORBITER INITIALIZATION TIME (S) ORBITER INITIALIZATION TIME (S)	RADAR I.D RECIEVER/TRANSMITTER BW (MHZ) SEASAT-A SEASAT-A SIGNAL-TO-NOISE RATIO (DR) SIGNAL-TO-NOISE RATIO (DR) BINARY PHASE CODE BINARY PHASE CODE BINARY PHASE CODE BINARY PHASE CODE BANGE SWAPLING RATIO (M) COSINE-OF PATCH-TO-MATCH OFFSET RANGE SWATH WIDTH (KM)14000000000000000000000000000000000	ANTENNA 1.D. 04 UNITPAT UNITPAT BORESIGHT SCUINT AT TO (DEG) 00 000000000000000000000000000000000	TERRAIN I.D. 00 NO OF FIELDS 00 X-AXIS COVERAGE (KM) 00 TERRAIN CENTER, R (KM) 00 TERRAIN CENTER, R (KM) 00 TERRAIN CENTER, R (KM) 00 TERRAIN CENTER, LONG (DEG) 00

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ARRAY STATISTICS FOR 42016 POINTS. MAXIMUM 342.3376838039. MINIMUM 8.954632127766E-15 MEAN .1804995086503. SIGMA 4.959671117744 EOF WRITTEN ON CIMAGE . AFTER RECORD 224