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# THE NH BANDS AT $\lambda 3360$ 

G. W. Funke

Translation of "Die NH Banden bei $\lambda$ 3360." Zeitschrift fuer Physik, Vol. 96, 1935, pp. 787-798.

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# THE NH BANDS AT $\lambda 3360$ 

G.W. Funke

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A detailed analysis of the bands at $\lambda 3360$ and 3370 is given; their $\Lambda$ and spin splitting as well as a widening effect which can be observed at these 3 bands are ${ }_{3}$ described. The bands form the $0-0$ and $1-1$ transitions of a ${ }^{3}$ II and ${ }^{2} \Sigma$ system. The recordings were made at a dispersion of $0.42 \AA / \mathrm{mm}$.

The band spectrum of NH at $\lambda \lambda 3360$ and 3370 has been the object of repeated studi $3^{1}$ ). Hulthen and Nakamura found that they are the $0-0$ and $1-1$ bands of a $I I^{3}$ and ${ }^{3} \Sigma$ system, it has been proved to be difficult, however, to carry out a detailed analysis of the spectrum, partly because of the fact that the B-values in ${ }^{3}$ IJ and ${ }^{3} \Sigma$ are close and similar resulting in $Q$-branches forming dense line accumulations which are hard to dissolve. Additionally, the intensive 0-0 band of the second positive nitrogen group has an edge with a corner cutting toward violet at $\lambda$ 3371. Ineir lines superimpose over the already dense conglomerate of the NH bands thus increasing the difficulty of analyzing it.

## Experiments

In order to start with this system at ali, two conditions have to be fulfilled; the disturbing nitrogen spectrum has to be removed and the recordings have to be made with a spectrograph of as large as possible resolution. One way to remove the nitrogen spectrum is by using an illuminating gas ammonia flame or an oxygen ammonia flame as

[^0]light source; the obtained spectrum, however, is too dull, it cannot be received in a large spectrograph. This is why we used a vacuum light arc which can also be used as pressure arc. In order to find the most favorable conditions, experiments with various nitrogen-water mixtures were at first carried out without, however, receiving satisfactory results. It could be shown, however, that the nitrogen bands become weaker when the pressure was increased. Experiments with ammonia had better results, and if the gas flows and pressure is kept over 2 atm , the spectrum is completely clean. The increased pressure has, however, a new disadvantage; the line width is increased by the pressure. The plate best suited for the analysis was consequently exposed at a pressure of 1000 mm .

Based upon the above-mentioned experiments, the recording was then made in the following manner: a pressure arc with especially forceful water cooling was used as light source which could be loaded up to 15 amp at least. The light arc burned between two tungsten electrodes which do not have a disturbing atomic spectrum under these conditions. The voltage was 440 V . We might mention that this light arc has an extraordinary appearance inasmuch as it contracts to a thin thread when the pressure is increased; under the influence of the field of both eiectrodes this thread curves outward in a loop. If pressure or current are increased too much, this loop curves outward until it breaks and the arc becomes extinct. Under pressures and currents when the arc burned without becomine extinct, the thin lighting thread oscillated to and fro; this is why the complete arc was put on wheels, so that the picture of the thin light arc could be fixed to the slit during the exposure. Gas was introduced into the pressure arc from a tank with liquid ammonia; from the pressure arc it was led into water where the part of the ammonia which was not decomposed by the light arc was absorbed. In the experiments with a pressure lower than air pressure in the arc, a water jet pump was used to suck the gas from the arc. Simultaneously, it served as pump and absorption medium for $\mathrm{NH}_{3}$ and it is the simplest feasible setup in a case like this without disturbing OH bands.

In order to obtain the necessary resolution, the recordings were made in the 4 th order of our large concave grid with a dispersion of $0.42 \%$ in this range.

## Analysis

A complete ${ }^{3}$ II $-{ }^{3} \Sigma$ band should consist of 27 branches as is the case with the hydride of the following element in the same column of the periodic system, i.e. $\mathrm{PH}^{2)}$. It cannot be expected that all these branches exist for the NH molecule because the distance of the components in ${ }^{3}$ II is 10 or $20 \mathrm{~cm}^{-1}$ and it consequently belongs to Hund's coupling case. 9 main branches, one $P$ and one $R$ satellite were here observed. The other $S$ side branches which should be expected if ${ }^{3}$ II is close to case b) obviously are too weak to show on the plate. The $Q$ satellites cannot be separated from the compressed mass of other lines. The $P$ and $R$ satellites are too weak to be detected on the plates of the 4th order: they were measured on an overexposed plate of the lst order, 1.e., with a dispersion of about $1.7 \% / \mathrm{mm}$. It can possibly be assumed that these weak branches are caused by the nitrogen isotope $N^{15}$; this, however, is impossible since $\rho-1$ and $p-2$ are negative and $\left.r_{1}^{1}-:=\frac{0}{2}-1\right) v_{v}$ or $r_{r}^{i}-r=\left(g^{2}-1\right) r_{r} \quad$ would show the isotope lines at the red side of the main lines of $R_{2}$. The $P$ satellite is in the right position of $P_{2}$, with increasing $K$, the splitting, however, is decreased instead of increased, as it should be if it represented a rotation isotopic effect. The nuclear vibration effect is not considered at all since it is very small due to the equality of $\omega^{\prime}$ and $\omega$ ''.
$R_{1}, R_{2}, R_{3}$ and $Q_{1}$ form edges which have corner cutting toward red. $Q_{2}, Q_{3}$ and the $P$ branches, however, do not form edges. Based on these facts the peculiar structure of this band system can now be understood. The corner cutting at the short wavelength side of the maximum is due to the $Q_{2}$ and $Q_{3}$ branches and especially to the initially decreasing distance between the various lines in $Q_{3}$ from red toward shorter wavelengths until a distance minimum of $K=9$ is

[^1]reached; it then increases again. For the line distance within the $P$ branches the same is true as for $Q_{3}$.

The term diagram represented
in Fig. 1 cannot be deduced
in all details from facts known about NH; a comparison with other similar, known spectrums has to be used too. Already Mulliken stated that the ${ }^{3} \Sigma$ state has to be termed ${ }^{3} \Sigma$, and this means that all levels with even $K$ are negative. Levels with $J=K+1$ are termed $F_{1}$, those with


2798 $J=K$ are termed $F_{2}$ and those with $J=K-1$ are $F_{3}$. The numbering of the individual lines was carried out after $K$ in the final stage.

## The excited stage

Based on van Vleck's studies, Mulliken gave a general term formula for case b), but since not all functions in it are known for ${ }^{3}$ II and since the function referring to the $\Lambda$ splitting additionally does not conform well with theory in this case, we only use the formula

$$
F(K)=B_{v} K(K+1)+D_{v} K^{2}(K+1)^{2}+F_{v} K^{3}(K+1)^{3} .
$$

which represents the term without considering $\Lambda$ and spin spiliting. Calculating the constants, only values of ${ }^{3} \mathrm{II}_{1}$ were used and a mean value between $\Delta_{1} F_{a}$ and $\Delta_{1} F_{b}$ was calculated by the following equation: 3 )

$$
\begin{gathered}
\quad \frac{Q(K)-P(K)+R(K-1)-Q(K-1)}{2}=\Delta_{1} F(K) \\
=F(K)-F(K-1)=2 B_{v} K+4 D_{v} K^{3}+F_{v} K^{3}\left(6 K^{3}+2\right) .
\end{gathered}
$$

[^2]Table 1. Table of band lines. 0-0


Table 2. Table of banc lines. l-1

| $\boldsymbol{N}$ | $1 \cdot$ | 4 | $\mathrm{H}_{3}$ | $1 / 1$ | Q2 | R: | P1 | 4 | 14 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 |  |  |  |  |  |  |  |  |  |
| 2 |  |  | 20777,38 |  | 2:1675, 16 | 29767.16 |  |  |  |
| 3 |  |  | 805.07 |  | 672,23 | 7!6,19 | 99:060,93 |  | 29781.01 |
| 4 | 24064.0:3 |  | 8:39.61 | 295.18 .36 | 669.15 | 82t.06 | 631,62 |  | 814,78 |
| 5 | 20: 111 | 29676,97 | 8 860. 11 | 51.190 | cosp, is | 853.85 | 601.48 |  | 816.46 |
| 1 | 18:1) 5 | 6i83,68 | 888.10 | IMI.8s | 60, 6.81 | 883.37 | 171.67 |  | 475,34 |
| 7 | diticis | (i81). 40 | 9115:2 | 1/11.53] | (i62,30 | 1110,43 | 111.23 |  | 96.4,63 |
| 8 | 422,7\% | 66ic,98 | 019.15 | 117.31 | 669, 61 | 937,7.4 | 410.35 |  | 939.6! |
| 9 | 385.88 | 6:33.43 | \%18. 18 | 385.06 | 6i6i. 17 | (167.4,62 | 371,24 |  | !min,3! |
| 10 | 367.35 | 0, 09.61 | 9904,25 | 302.95 | cib3, 1 | 990,77 | 311.85 |  | 957,02 |
| 11 | 32.1 .71 | (inin. 00 | 30018,96 | 320.98 | 6.49,4 | 30016.10 | 816,71 | 20646.93 | 30013,08 |
| 12 | 2!2.24 | (6il,08 | 012.99 | 288.93 | 645,3 | $0.10,33$ | 28.50 .6 | 0.13,3.4 | U137.68 |
| 13 | 2atini | (ili, 615 | Oitios | 2\%\% | (illi, ic: | Mi:3, i: 4 | 2alw | (ixis, 1: | milisi |
| 11. | 227.14 | (1510.63 | 0, 7.911 | 29.30 | (365,69 | (18. 87 | 2:2, ivj | (13) 1 ¢ | 11:3, 318 |
| 15 | 1!\%.il | 63128 | 1 (18.(6i) | 113.40 | (\%20,81 | 106, 8 ? | 191,27 | (ies, 73 | 115.20 |
| 10 ] | 163.3.614 | $6: 27.77$ | 1:360 | 161.6. | 6:3.34 | 126,67 | 159.84 | 62e. 18 | 129.26 |
| 17 | 131.1 .7 | (i20.36 | 146.20 | 129.87 | 616.07 | 14,4,85 | 128,2\% | 615.195 | 14,3,8.4 |
| 18 | 099,3i | 6129 | 162,91 | 0178.88 | 608.08 | 161.68 | 096,63 | Gus.08 | 16i0, 8 \% |
| 19 | 067.03 | tam, 19 | 177.79 | (193, \% $^{1}$ | [194.16 | 176.77 | (6i.) ${ }^{\text {m }}$ | 6 019.419 | 176,301 |
| 20 | 0131.62 | -113.27 | 1:M1.:3 | 10:3. 11 | [s3), 71 | 1! 11.11 | 13:3,410 | [149, 71 | 190.11 |
| 21 | covered | Tverex | 202.31 | (1mi.01\% | 678.! | 211.101 | 101.015 | [178.92 | :01.60 |
| 29 | -x mix mit | C.11.16 | 211.11 | -x!mis.4 | , Wi2. 117 | :11.14 |  | furi.un | :11.11 |
| $\pm 1$ | : 26.6 , $1:$ |  | alsinion | ! 11.811 | 663.60: | :1s.6: | 111.70 | 661.010 | 2 Lt , 6 |
| 4 | :M1: $: 1$ | 611.78 | 23:36 | ¢101.21 | bisk:46; | 2.31,7i | ! 11.19 | 6331. 13 |  |
| : 5 | Rniz. 1 m | 6ximes | :2em, 18 | s.if. 149 | 6esem | 20.9.14 | M17, 101 | 6 m .11 | : $: 15,1: 8$ |
| 21 | 8:31, $3:$ | intich: | 2 El 12 | \%31.:3 |  | 2: 11.12 | 8:11.73 | (ali, imi | 20:10.610 |
| 9 | 7n, , in | 1: 11.1 | -13. $1:$ | ationis |  | -1:4, $1:$ | 7the, init |  | 211:37 |
| 2\%1 | 7in, ${ }^{\text {d }}$ | 121.31 | 202:36 | 368.24 |  | $203: 36$ | 769.411 |  | 210:11 |
| 40 | 715 689 60.49 | 1i1.:17 | 187.00 | 719.0 .1 079,19 |  | 187,06 | 7:0, 37 |  | 188,2:3 |
| 30 | (3)3, 49 |  |  | (79, 49 |  |  | 680,27 |  |  |

$\Delta_{1} F(K)$ is calculated from ixperimental data and then $B_{v}$ and $D_{v}$ are graphically determined hy wrawing $\Delta_{1}$ Ferjuxcatofunction of $K^{2}$. $B_{e}$ and $\alpha$ in equation $\dot{\nu}_{r}{ }^{2 n} B_{e}-\cdots(v+1 / 2)$ are also calculated and $\omega_{e}$ is obtained from Kzatzer's relstion. All constants are compiled in Table 3.

Tabie 3.


## The nornal stage

We use the same term representation as in ${ }^{3}$ II. The best constant values are here also obtained by using $F_{2}$; since we do not have a $\Lambda$ splitting for which a mean value has to be calculated, we now start with the differences

$$
\begin{gathered}
R_{2}(K-1)-P_{2}(K+1)=U_{2} F_{2}=F_{2}(K+1)-F_{2}(K-1) \\
=4 B_{v}(K+1)+8 D_{v}\left(K+1 ; F_{v}-12 F_{v}(K+2)^{5}\right.
\end{gathered}
$$

and calculate the constants graphically.

## Spin splitting

The existence of a marked spin splitting in ${ }^{3} \Sigma$ levels can te seen from Table 4; it shows that $\left.I_{2} F_{i}^{\prime \prime} K_{i}\right\rangle 1_{2} F_{a}^{\prime \prime} K_{i}>1_{2} F_{1}^{\prime \prime}\left(K_{)}\right.$. The process of this spiltting cannot be obtained by exclusively using the main branches for the calculation. We can, however, form ${ }^{\prime \prime} Q_{21}(K)-R_{2}(K)$ and ${ }^{P} Q_{23} \sqrt{K}-P_{2}(K)$ and then get $F_{2}-F_{1}$ or $F_{2}-F_{3}$. As can be seen from Table 5, $\mathrm{F}_{2}-\mathrm{F}_{1}$ increases and $\mathrm{F}_{2}-\mathrm{F}_{3}$ decreases as K increases. No more exact values can be expected considering that the lines are very weak, lying close to very strong ones and dispersion was rather small as mentioned above.

Table 4


Table 5

|  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\stackrel{\sim}{*}$ | $F_{2}-r_{1}$ | $\mathrm{F}_{1}-\mathrm{F}_{1}$ | 1 | $r_{2}-r_{1}$ | $F_{3}-F_{3}$ |
| 7 | 1,13 | 1,13 | 10 | 1,19 | 1,03 |
| 8 | 1.23 | 1,22 | 11 | 1,85 | 1,27 |
| 9 | 1.21 | 1,01 | 12 |  | 1,00 |

Table 6. The odd term values for ${ }^{3}$, if $F(K=3)=0$

| $\stackrel{H}{ }$ | Fs, | $F$ | $F_{1}$ | $r_{3}-r_{2}, F_{1}-F_{1}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 3 | 0 | 0 | 0 |  |  |
| $j$ | 23.15 | $\therefore 202.90$ | 292.72 | 0.28 | 0.15 |
| 7 | 714, | - 714.00 | 713,55 | 0.36 | 0.42 |
| 9 | 120.3 3 | 1201.25 | 1260.59 | 0.48 | 0.59 |
| 11 | 1032 \% | 1932.01 | 1931,30 | 0,53 | 0.71 |
| 13 | 273.3 | 298.84 | 272.216 | 0.71 | 0.68 |
| 1.5 | 3430 | 3629.95 | 3629,24 | 0.85 | 0.71 |
| $1:$ | +64.4. 4 | 4649.03 | 4648.13 | 0.94 | 0.90 |
| $1: 1$ | 576 | 586.12 | 574.11 | 1.00 | 1.01 |
| 21 | -190, | 7100.78 | 71001.62 | 1.06 | 1.16 |
| : | s397.2\% | 8326.19 | 83.4 .04 | 3,9 | 1.16 |
| 25 | 91410.0 | 9738.48 | 973i.62 | $1.2 \%$ | 1.16 |
| 29 | 11236, 2 | 11235.01 | 11233.85 | 1,27 | 1.15 |

It can be seen from the Table that $F_{2}-F_{1}=F_{2}-F_{3}$ at $K=7$ or 8 , i.e., we can assume that $F_{1}=F_{3}$ at $K=71 / 2$. We can subsequentiy obtain the second term value for $F_{3}, F_{2}$ and $F_{1}$ by substituting $F_{3}(3)=0$, $F_{2}(3)=0$ and $F_{1}(3)=0$ (see Table 6). We then form the differences $F_{3}(K)-F_{2}(K)$ and $F_{2}(K)-F_{1}(K)$ and draw a diagram of the changes of these differences with $K . F_{2}$ will be in the abscissa of this diagram, above it will be $F_{3}$ and below it $F_{1}$. Lastly we shift the
curves for $\mathrm{F}_{1}$ and $\mathrm{F}_{3}$ parallel to themselves so that $\mathrm{K}=7, \mathrm{~F}_{1}=\mathrm{F}_{3}=$ 1.18. It can be seen that the differences in Table 5 which were obtained with the satellites fit well into the curve; it has furthermore the general appearance postulated by Kramer' theory and it is in accordance with the respective curves for $\mathrm{O}_{2}$ and PH .


Fig. $3_{\Sigma}^{2}$; $\operatorname{Spin}_{\mathrm{v}=0}$ splitting

IL must be stressed, however, that there can be some uncertainty about the position of the point of intersection of the $F_{1}$ and $F_{3}$ curves, due to the fact that the few lines of the satellites are analyzed correctly; an error of this kind, however, effects a parallel shift of the curves at the most and does not change their general appearance.

## A splitting

The material collected about $A$ in ${ }^{3}$ II stages is rather insignificant and does not agree very well with the theories. In case a) it should be expected that the splitting in ${ }^{3} I I_{0}$ is large and independent of $J$, smaller for ${ }^{3} I_{I}$ and proportional $J(J=1)$, and immeasurably small for ${ }^{3} \mathrm{II}_{2}$. In case $b$ ) the splitting proportional to $K(K+1)$ and equal to ${ }^{3} I I_{0},{ }^{3} I I_{1}$ and ${ }^{3} I I_{2}$ should be expected since $S$ is weakly coupled to $\Lambda$.

Examples of $A$ splitting are given by the first positive nitrogen group 4) and PH 5) for case a); both examples, however, do not follow the above-given diagram. The magnitude oi the splitting is at least the one to be expected for $N_{2}$. It is largest for ${ }^{3} I_{0}$, smaller for ${ }^{3} \mathrm{II}_{1}$ and smallest for ${ }^{3} \mathrm{II}_{2}$. Not even this is true for PH .

[^3]${ }^{3} \mathrm{II}_{0}$ has the smallest solitting, both the other levels have a silghtly larger one and ebout the same amorg themselves. For the study of this spifting, $N H$ is better suited since it here is significantly larger than in toth other molecules. It is largest for ${ }^{3} I_{I_{1}}$, smaller for ${ }^{3}{ }^{I} I_{2}$ and smallest for ${ }^{3}{ }^{1} I_{0}$, the difference, however, is as postulated by theory, not exceptionally large. It could be concluded consequently that NH conforms better with theory than the earlier mentioned nolecules. There are worse data, however, on the proportionality with $K(K+1)$. It is understandable that the curve follows a function $4 v_{a b}=c_{1}+c_{2} K+c_{3} K^{2}+\ldots$, but since the meaning of these constants is unknown there is no use in calculating their values. The curves for $v=0$ and $v=1$ show a very similar shape and it can immediately be seen that the spiltting is almost identical for ${ }^{3} \mathrm{II}_{2}$ and ${ }^{3} \mathrm{II}_{2}$ whereas the curve for ${ }^{3} \mathrm{II}_{0}$ is farther separated from the others, not to the extent, however, that it runs parallel to the curve for ${ }^{3} \mathrm{II}_{1}$.

A few other things should be mentioned. We here have a: ixample for $\Delta v_{a b}$ decreasing in its amount with increasing $K$, passing through zero and increasing again but with the opposite sign. Moving toward higher $K$ values, however, the point is reached where the increase stops and splitting even starts to decrease again. The same effect could be noticed for BaH ${ }^{6}$ ) for ${ }^{3}{ }^{1} I_{1,2}$ and in other molecules.


Fig. 3. $\Lambda$ splitting in ${ }^{3}$ II; $v=0$
6) S.M. Naude, Proc. Roy. Soc. London (A) 136, 114. 1932


Fig. 4. A splitting in ${ }^{3} \mathrm{II} ; \mathrm{v}=1$

The widening effect

As mentioned earlier, the recordings were made under rather high pressures so that the $\mathrm{N}_{2}$ bands should disappear: this, however, caused an increased widening of the lines with increased pressure which became prominent because of the high dispersion
used. This widening effect, however, does not disappear totally when pressure is decreased, it is practically constant in the pressure range 100 to 1000 mm . The widening can consequently not exclusively be explained as pressure effect (Lorentz shock absorber or Stark effect of surrounding molecules), it must rather also have a pressureindependent cause like, e.f., the Doppler effect. Since it is not very iorthwhile to study line widths in emission and especially in arcs with their hard to define relationships, no exact measurements of $3-\mathrm{db}$ band widths were carried out. In their magnitude, however, they a.e about 0.04 \& if pressure is lower than 1000 mm , and they seem to be the same for all lines.

For the Doppler effect the formula

$$
\delta_{d}^{2}=\sqrt{\ln 2} \frac{2}{c} \sqrt{\frac{2 H T}{\mu}} .
$$

Is valid and if temperature is roughly set to be $4000^{\circ}$, a Doppler effect of about $0.02 \%$ is obtained. Consequently, the pressure independent widening of the lines can at least partly be regarded as due to Doppler effect. A light source like this one is actually not suited for maximum resolution of dense line accumulations, at
least not for iight molecules since in this case the light source itself and not the optical instrument is setting the limit for resolution. Incidentally, the possibility exists that the Doppler effect is not alone responsible for the widening.

In order to better study these effects, experiments were carried out to obtain NH in absorption where all conditions like temperature, pressure etc. can be better determined. ${ }^{3} \Sigma$ is an original state and absorption should be possible. The first experiments were carried out with a quartz spectrograph in a 1 m long iron tube heated to $750^{\circ}$ by gas flames and filled with flowing ammonia gas. They were unsuccessful, however. Likewise no results yielded the experiments with a vacuum oven at different pressures and temperatures up to $1750^{\circ} . \mathrm{N}_{2}$ and $\mathrm{H}_{2}$ cannot be split and NA cannot be formed by increasing the temperature since the dissociation heat for $N_{2}$ is too high for an oven and besides it is higher than for NH , but it might be feasible that after decomposition of NH an intermediate product $\mathrm{NH}+\mathrm{H}_{2}$ or $\mathrm{NH}+\mathrm{H}+\mathrm{H}$ is formed and that this might result in NH absorption. $\mathrm{NH}_{3}$ decomposes readily at $900^{\circ}$; the presence of NH is, however, not indicated by absorption at even this temperature.

As mentioned earlier, the NH spectrum is emitted by an illuminating gas flame, and the molecule consequently has to exist in it. An experiment was now arranged so that a 1.75 m long tube had slits.: which were cut out perpendicular to the longitudinal direction of the tube. A inixture of $\mathrm{NH}_{3}$ and illuminating gas was introduced into the tube; this mixture flowed through the slits and then was lit so that a series of ievel flames was produced. Through them a parallel light beam from a tungsten lamp was sent toward a quartz spectrograph. A single, very diffuse absorption line at $\lambda 3360.14$ resulted. At $\lambda 3360.10$ lies the edge of $Q_{1}$, and there is no doubt at all that it is this edge indeed. Since temperature is low in such a case, absorption of only the lowest energy levels can be expected. $R_{1}(0)$ and $R_{2}(0)$ start from level $K=0$; these lines, however, do not appear.

The Q lines must have the highest intensity, and since additionally many of them fall together in one point, it is not so strange that the $Q$ edge should appear first. By this experiment is has consequently been confirmed that ${ }^{3} \Sigma$ is the original state; more ines, however, should have been expected since -as shown above- the molecule is bound to be formed during ammonia decomposition and is thus also bound to absorb.

Lastly I would very much like to thank Prof. E. Hulthen, who initiated this study and who showed great interest in it. I thank Prof. Herzberg, Darmstadt for interesting discussions.

$$
\text { Stockholm, Universit, Physics Institute, July } 1935 .
$$


[^0]:    *Numbers in the margin indicate pagination in the foreign text.

    1) A. Fowler u. CL. Gregory, Phil. Trans. Roz. Soc. London (A) 218, 251, 1919: K. Gleu, ZS. F. Phys. 38, 176, 1926; E. Hulthen u. S. Nakamara, Nature 119, 235, 1927; R.W.B. Pearse, Phus. Rev. 37, 1712, 193:H. Batsch, Ann.d.Phys. 18, 1933; Ritzmann,Dissertation,Breslau, 1934
[^1]:    $\left.{ }^{2}\right)_{\text {R.W.B. Pearce, Proc. Roy. Soc. London (A) 129, 328, } 1930}$

[^2]:    3)This assumption is in a way arbitray,but it should be the most suitable one until a detailed theory on $\Lambda$ splitting in ${ }^{3}$ II is found.

[^3]:    4) S.M. Naude. Proc. Roy. Soc. London (A) 136. 114. 1932.
    5) Loc.cit. Zeitschrift fuer Physik, Bd. 96
