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CONTROL OF LONG ORBITIMG TETHERS
(NASA-CR-170720)

August 1981

Smithsonian Institution
 Astrophysical Observatory Cambridge, Massachusetts 02138

> The Smithsonian Astrophysical Observatory and the Harvard College Observatory are members of the Center for Astrophysics

# STUDY OF <br> CERTAIN TETHER SAFETY ISSUES AND <br> THE USE OF TETHERS FOR PAYLOAD ORBITAL TRANSFER <br> CONTIMUATION OF <br> investigation of electrodymamic stabilization and CONTROL OF LONG CRBITIMG TETHERS 

Contract NAS8-33691

# Semi-Annual Report <br> For the period 1 March 1981 thru 31 August 1981 

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This report covers work done during the period 1 March through 31 August 1981 on two extensions of MASA Contract MAS8-33691. Section I, describes work done at SAO on the study of tether safety issues. Section II, describes the work done at MIT studying the use of tethers for payload orbital transfer.

This eection presents a mamary of work done at gro on the study of tether safety iseves. Detailed results are presented in the monthly reports for the period March - July.1981. The effort during this period has been directed toward understanding the behavior of the tether after a failure at various distances from the stouttle and after joming of the reel mechanim during deployment. Analytic expressions derived under simplifying aseuptions have been used to estimate the anount of recoil of the wire after a break as a function of the systen paraneters.

Since the first experiment with the Shuttle mey be an electrodynanics experiment with a short tether, we have used the following test case for the studies. We assume a 100 metric ton Shuttle in orbit at 220 kam with a 300 kg subeatellite deployed upward on a 10 km tether 2 m diameter. The wire is represented by discrete masses at 1 km intervals. Initial conditions have been computed such that the systerr is in equilibrium. A break in the wire is similated by omitting the mass points representing the subsatellite and the portion of the wire beyond the break in the integration of the equations of motion. Puns bave been done using 1, 2, and 5 of the masses representing the wire, plus the mass representing the Shuttle.

The case with only one wire mass can be described quite well without the use of numerical integration. From the tension in the wire, and the other parameters of the system we can calculate the velocity with which the mass point will recoil. Assuming that loss of tension between the wire mass point and the Shuttle occurs in a relatively short time, the initial position and velocity (consisting of the orbital velocity and the recoil velocity toward the Shuttle) can be used to calculate the new orbital elements for the wire mass point which now orbits as a free particle. From these orbital elements, the closest approach to the shuttle and the time of cloeest approach and be calculated. A small conputer progran has been written using the analytic expressions in order to do a parametric study for breaks at various positions along tethers of various lengths. A table of results obtained is given in the monthly report for June,1981. Two features are aqparent from the table. First, the amount of recoil is approximately imversely proportional to the length of the broken piece of wire. Second, the results scale with the length of the wire (assuming the tension is proportional to the length of the wire). That is, if a 1 km piece of wire broken from a 10 km wire recoils 11 meters, we can multiply all the mubers by 10 and get the result that a 10 km piece broken from a 100 km wire recoils about 110 meters.

The expreasion for the recoil velocity is interusing because it turns out to be independent of the length of the wire megment represented by the mass point. The expression is obtained by setting the kinetic energy of the wire masa equal to the energy stored in the stretched piece of wire. The mass of the wire point is the product of the wire density. cross section, and length. The atiffness of the wire segnent is the procirt of the elasticity, croes aection, and the inverse of the length. Performing the algebra gives the expression

$$
v_{r}=T / A \sqrt{p E}
$$

where $v_{r}$ is the recoil velocity, $A$ is the wire cross section, is the density, and $E$ is the elasticity. The recoil can be reduced by making the wire thicker, or using materials with a higher density or stiffness. This formula was derived treating the wire as a single lump. Multi-mass runs are required to study the behavior of the wire in more detail.

The fact that the length of the wire aegment does not appear in the expression for the recoil velocity suggests the possibility that all parts of the wire recoil with the seme velocity. When a break occurs, the loss of tension will propegate down the wire at the speed of sound which is given by the expression $\sqrt{E / P}$. For the parameters used in the simulation this velocity is about $5.3 \mathrm{~km} /$ second. A run has been done with 5 wire mass points plus the Shuttle to see the behavior of the wire in more detail after a break. Detailed results are presented in the monthly report for July, 1981. The tansion as a function of time is shown in Figure 2a. It takes about 1 second for tha loss of tension to propagate down the 5 km wire. After the initial loss of tension, the various sections go in and out of tension as seen in the latter part of the plot. All sections of the wire acquire a velocity toward the Shuttle of about $30 \mathrm{ca} / \mathrm{sec}$ except for the point closest to the Shuttle. The behavior of the last print is anomalous, presumably because of boundary effects. The wire appears to move more or less as a unit after the break.

Because of the gradient of the gravitational and centripital accelerations there is a stretching force acting on the wire. In a run done with only two wire mass points, the two masses moved toward each other in the initial contraction after the break until the stretching forces halted the motion and brought the masses back into tension. The masses bounced back and forth from each other in a cyclic fashion during the run. The run with five masses described previously exhibits the ame behavior in a more complicated fashion. Basically, the wire contracts after the break to approximately the natural length of the wire, and then the distances between mass points oacillate around the natural length with different sections going in and out of tension. The wire as a whole moves toward the Shuttle with the recoil velocity given by the equation presented earlier.

Without atmogheric dreg, corialis focces canse the wire to move forvard when the wire recolis (for upmard deployment). the railal ve in-plame bebevior for the cun with Eive wire mave pointe plus the Elouttie is shown in Fiqure 20. The in-pime motion in greatiy exagarated in the plot. pum have been dom for 2 wire mase pointe with, and without drag to ree which effect doninater in this particuiar cose. The motion without drag was ebout .9 matars fommerd after 60 ceconds, and about 2.8 meters to the coar when drag mase egilied; indicating that drag dominates by a few factors in the case.

Two suns have been done to study the bebavior of the wire in the cace where the reel jame during daploymant. Integrating the motion of only the aubatellite and the stuttle with the wire considered masulens gives a tension which incresses to a maximu value and then falls to zero as the subatellite recoils toward the shuttle. The bieak etrength of the 2 kevlar wire was not exceeded with a $20 \mathrm{~m} / \mathrm{sec}$ deployment velocity and a 300 kg abeatellite. In analytic calculation mows a cloeest approach of 5.5 km to the Shuttle with the reel jam occurring at 10 km from the shuttle. The rm was repeated adding 4 wire mass points. Details of this run are given ir the monthly report for July, 1981. Tension variations from 114 to 230 kg are geen along the wire as a reault of the oscillations caused by the reel jaming. After the recoil, the sections of wire remain out of tension. Eremination of the velocities of the various mass points shows the recoil velocity to be roughly proportional to the distance from the Shuttle. The wire is therefore continuing to contrert on itgelf in contrast to the behavior geen in a break where the wire recoils with a nearly uniform velocity. More details of the reel jamming case will be presented in the next monthly report.

The figures from Monthly Report No. 4 are shown in the following pages af a summary of the study of the cases of a broken wire and jaming of the reel durirg deployment.

Figure 1. Redial ve in-plane component(cin) after a break 5 kil from the Shuttle plotted every second for the first 18 mecondis.

Pigure 2. Simulation of a break 5 km from the Shuttle plotted every .1 seconds. Part a) is the tension(dynes) ve time (enc). Part b) is the radial vi in-plane component (cin).

Figure 3. Notion of the wire after a break 5 km from the Shuttle. In part a) the initial value of the radial component for esch mass has been subtracted from the subsequent values and the curves have been separated from each other by 100 cm. Part b) is the modified radial component vs the in-plane component.

Figure 4. Motion in the radial direction after a break 5 km from the Shuttle. The values plotted were obtained by first subtracting the initial value for each mass, then subtracting the values for a reference mass from all the masses, and finally separating the curves by 35 cm . In part a) the sixth mass is used as the origin, and in part b) the second mass is used as origin.

Figure 5. Behavior of the wire after a break at 5 km with the first .8 seconds excluded from the plots. In part a) the value at .8 seconds is subtracted from es. point, then the values for the second mass are subtracted from each curve, and finally the curves are separated by 2 cm . The value of the sixth component which is ancmalous is essentially ignored by holding it fixed at the value of the separation constant. Part b) shows the tension vs time after . 8 seconds.

Figure 6. Tension vs time after a reel jam with the subsatellite at 10 km being deployed at $20 \mathrm{~m} / \mathrm{sec}$. In part a) the mass of the tether is neglected. Part b) shows the tension for each wire segment with wire masses every 2 km along the wire.



8.

9.

10.


MASS SYMBOL



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14.

15.

Figure 6b

## ACKNOWLEDGEMENT

The anti,lors of this report are Mr. D.A. Arnold and Dr. M. Martinez-Sanchez.

## II. The Use of Tethers for Payload Orbital Transfer

## Introduction

During this first reporting period the work has centered on our first task namely, the selection of concepts for orbit transfer missions which show promising improvements by the use of tethers.

A preliminary look at the original MPPF concept (a Shuttle-based tether system flown to and from orbit in each Shuttle flight) indicated only marginal benefits. The alternative concept was then evolved of a free-flying tether system, deployed in a first flight, then left in orbit for docking with successive payload-carrying Shuttle flights. Significant payload increases were found if this device is used for LEO-GEO transfers, in a mode where the payload carries a propulsion stage (IUS or Centaur) for first $\Delta V$ assist and for the circularization $\Delta V$ in $G E O$.

A more ambitious system was also examined in which a second tethered system in GEO takes the role of providing the circularization step. Particular attention was paid to the requirement that the transfer duration should be a rational fraction of one day, to frovide additional ercounters in case cf rendezvous failure. It was found that for most combinations of parameters, the length of the upper tether ranged around $10,000 \mathrm{Km}$, while the one in LEO was about $1,000 \mathrm{Km}$. Progressive addition of $\Delta \mathrm{V}$ capabilities at the two ends of the transfer can, of course, reduce these lengths. Operational limitations based on minimum perigee height and maximum tether weight were studied.

A preliminary conclusion of this work has been that operating modes which combine tether assist with substantial rocket-derived velocity increments can be of j mportance in increasing the payload deliverable to high and geosynchronous orbits, or for escape. This capability may be crucial for certain missions, such as the Galileo mission, which are at the edge of the Shuttle's capabilities.

## Single tether concepts for orbit transfer.

Use of a tether facility as a permanent facility of the Shuttle does not appear justified for missions that fall within the operational envelope of the orbiter with its integral OMS tanks. This is because, even though the tether allows deployment of the payload from a lower Shuttle orbit (typically an elliptic one), the payload cannot be increased due to other constraints, such as payload bay structural integrity and c.g. location. The only savings are then in the use of less OMS fuel, but these cannot balance the loss of revenue from the payload displaced by the tether itself. An example is shown in Table 1: a 47 cm tether allows payload to be placed in a $500 / 500 \mathrm{Km}$ orbit from a Shuttle in a $185 / 453 \mathrm{Km}$ orbit, with an OMS fuel savings of $\$ 33,000$. However, the mass and length of the tether facility displaces payload worth $\$ 2.80 \mathrm{M}$. Similar results are shown for a polar orbit.

There are some possible scenarios where a Shuttle based tether could be cost-effective. These refer to low Earth orbits high enough (particularly for polar orbits) that payload is limited by OMS fuel capacity, including extension kits. A trade-off study is planned to determine how far the operating envelope can be extended by a permanent Shuttle tether.

## TABLE 1.

## COST TO LON ENERGY MISSION**

|  | Space Telescope | Polar Orbit |
| :---: | :---: | :---: |
|  | Orbit 500km/28.8 ${ }^{\circ}$ | 1000 knd $97^{\circ}$ |
| Weight of Payload (kg) | 11,000 | 3,000 |
| Length of Payload (m) | 13.1 | 9.0 |
| Diameter of Payload (m) | 4.26 |  |
| Cost to current Shuttle (\$M) | 20.20 | 23.07 |
| Cost to Shuttle + Orbiter based tether system (\$M) | 23.00 | 29.8 |
| Lost revenue from displaced payload (\$M) | -2.80 | -6.73 |
| OMS fuel savings (\$M) | (0.033) | (0.083) |
| Benefit of using tether system (SM) | -2.77 | -6.647 |
| ** |  |  |
| 1) Cost per Shuttle flight $=\$ 27.3$ at ETR |  |  |
| \$46.9 at WTR |  |  |
| 2) Elliptic shuttle orbit + tether tra perigee altitude $=185 \mathrm{~km}$ | nsfer |  |

## TABTE 2.

## PAYLOAD BENEFIT FOR GEOSYNCHRONOUS ORBIT TRONSFER*

Tether length
(km)
Payload Weight (kg)
Payload increase ..... ( 3 )
0 ..... 2465
100 3122 ..... 18
200 3675 ..... 39
300 4326 ..... 63
400 5100 ..... 93

* Calculation conditions:

1. SHUHITE + Two stage ..... IUS
Stage ..... 1 ..... 2
Isp (sec) $291.9 \quad 289.7$
f stru. .....  946 ..... 933
WI prop. (kg) 9707 ..... (2722)
2. Parking orbit: 300/300 km
3. Tether system dock with shuttle in parking orbit.

## tabis 3.

PAYLOAD benerfit for solar system exploration**


It appears, however, that a more efficient ayntem would be, in any case, one where the tether and its end platforms would be left deployed in epace, to be docked with the Shuttle each time. The Shuttle wuld transfer the payload (possibly with a transfer propulaion atage attached) to the tether lower platform; the payload would then move to the upper platform, be released, and the Shuttle would then detach and reenter. The problem of re-establishing the initial orbut for the tether has to be examined in more depth; some of the required propulsion could be provided by the Shuttle itself, but the fact that the platform stays in space opens the possibility of using substantial amounts of high specific impulse electric propulsion in the process. Thus, the system becomes a hybrid between the original TOTF (Tethered Orbital Transfer Facility) and the MTPF (Mechanized Tether Platform Facility) concepts, with the platform mass being a parameter to be optimized.

A preliminary examination of the performance gains for this ireeflying tether concept was made, and is shown in Tables 2 and 3. Particularly for transfer to geostationary orbits, large payload increases are shown to be possible with tether lengths not exceeding 400 km . Orbital perturbation during ascent (with the associated minimum perigee problem) and orbit reestablishment, as discussed, remain to be studied in more detail.

## Two-tether LEO-GEO systems.

The principal interest in orbital transfer relates to low-to-geosynchronous cases. We considered the possibility of performing such transfers with:ut any transfer propulsion, or with small $\Delta V$ 's at most. This
requires a tether to be attached to a low Earth orbiting platform for release into the transfer elipse, and another tether attached to a geosynchronous platform, to acquire the payload and circularize ite orbit.


FIGURE 1. Geometry for a two-tether system.

The length $H$ of the upper tether depends onj $y$ upon the period $p_{2}$ chosen for the orbit of the paylo (after applicaticn of an apogee velocity increment $\Delta V_{Q}$. This is because two elements of that orbit are prescribed, namely, the semimajor axis (by the period) and the angular momantum (by the requirement that the angular velocity at apogee must equal that in the geosynchronous orbit). These conditions can be expressed as

$$
\begin{equation*}
P_{2}=\frac{2 \pi}{\sqrt{\mu_{e}}}\left(\frac{R_{P_{2}}+R_{Q}}{2}\right) 3 / 2 \tag{1}
\end{equation*}
$$

(where $R_{P_{2}}$ is the perigee of the orbit and $\mu_{e}$ the gravitational constant of Earth) and

$$
\begin{equation*}
v^{2}\left(\text { apogee }_{2}\right)=\mu_{e}\left(\frac{2}{R_{Q}}-\frac{2}{R_{P_{2}}+R_{Q}}\right)=v_{G S}^{2}\left(\frac{R_{Q}}{R_{G S}}\right)^{2} \tag{2}
\end{equation*}
$$

whare the subscript GS refers to the yeosyncluronous orbit. Using
$V_{G S}^{2}=\mu_{e} / R_{G S}$, Eqs. (1) and (2) can be combined by elimination of $\mathcal{F}_{P_{2}}$, to give

$$
\begin{equation*}
\frac{2\left(R_{Q} / R_{G S}\right)}{2-\left(R_{Q} / R_{G S}\right)^{3}}=\frac{2}{R_{G S}}-\left(\frac{\sqrt{\mu_{e}}}{2 \pi} P_{2}\right)^{2 / 3} \tag{3}
\end{equation*}
$$

which can be solved for $R_{Q}$ once $P_{2}$ is prescribed. The tether lungth follows then from

$$
\begin{equation*}
H=R_{G S}-R_{Q} \tag{4}
\end{equation*}
$$

and the perigee $R_{P_{2}}$ from

$$
\begin{equation*}
R_{P_{2}}=\frac{R_{Q}^{4}}{2 R_{G S}^{3}-R_{Q}^{2}} \tag{5}
\end{equation*}
$$

The subscript 2 has been used so far to indicate conditions after application of $\Delta V_{Q}$. For the ascent orbit (before $\Delta V_{Q}$ ), the apogee velocity is

$$
\begin{equation*}
v\left(\text { apogee }_{1}\right)=v\left(\text { apogee }_{2}\right)-\Delta V_{Q}=V_{G S} \frac{R_{Q}}{R_{G S}}-\Delta V_{Q} \tag{6}
\end{equation*}
$$

and Eq. (2) can be modified to calculate the perigee $R_{P_{1}}$ of this ascent orbit:

$$
\begin{equation*}
\mu_{e}\left(\frac{2}{R_{Q}}-\frac{2}{R_{P_{1}}+R_{Q}}\right)=\left(v_{G S} \frac{R_{Q}}{R_{G S}}-\Delta v_{Q}\right)^{2} \tag{7}
\end{equation*}
$$

Once $R_{P_{1}}$ is so determined, the velocity at perigee can be expressed (from conservation of angular momentum) as

$$
\begin{equation*}
v_{P_{1}}=\frac{R_{Q}}{R_{P_{1}}} v\left(\text { apo }_{1}\right)=\frac{R_{Q}}{R_{P_{1}}}\left(v_{G S} \frac{R_{Q}}{R_{G S}}-\Delta v_{Q}\right) \tag{8}
\end{equation*}
$$

This velocity contains, in general, a propulsion-derived increment $\Delta V_{p}$, applied at or immediately after release. The velocity of the end of the low-Earth tether is therefore

$$
\begin{equation*}
v_{\text {tether }}=v_{p_{i}}-\Delta v_{p} \tag{9}
\end{equation*}
$$

and the orbital velocity $v_{c, L E}$ of the platform at $R_{L E}$ (or, more precisely, of the orbital center of the tether-payload system) is therefore

$$
\begin{equation*}
v_{C, L E}=\sqrt{\frac{\mu_{e}}{R_{L E}}}=\left(v_{P_{1}}-\Delta v_{P}\right) \frac{R_{L E}}{R_{P_{1}}} \tag{10}
\end{equation*}
$$

from which $R_{L E}$ can be calculated easily. Finally, the low-Earth tether length is

$$
\begin{equation*}
h=R_{P_{1}}-R_{L E} \tag{11}
\end{equation*}
$$

Fig. 2 and Table 4 show calculated results for the case of $\Delta v_{P}=\Delta v_{Q}=0$. If we impose the requirement that, in case of docking failure, the payload and the lower platform of the GEO tether should rendezvous again after an integer number of orbits, then the period $P_{2}$ ( $T$ in Fig. 2) must be a rational fraction $m / n$ of a day ( $m, n$ integers). Thus, appropriate values of $P_{2}$, for low $m$ and $n$, are
$1 / 3$ day $=8 \mathrm{hr} ., 3 / 8$ day $=9 \mathrm{hr} ., 2 / 5$ day $=9.6 \mathrm{hr} ., 1 / 2$ day $=12 \mathrm{hr}$.
As shown in Fig. 2, a period of $1 / 3$ day implies an upper tether length of over $10,000 \mathrm{Km}$, and a lower tether length of about 1200 Km from a low Eartu urbit at 1200 Km as well. Increasing the period to 1/2 day lowers the length of the upper tether to about 6000 Km but it also requires the low Earth orbit to be at some 9000 Km altitude, with a 1600 Km tether.

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FIGURE 2. TWO-TETHER SYSTEM CHARACTERISTICS WITH $\Delta V_{P}=\Delta V_{Q}=0$.

## TABLE 4. TWO-TETHER SYETEM CHARACTERISTICS WITH $\Delta V_{P}=\Delta V_{Q}=0$.

| . $\mathrm{H}=$ | 7000.000 | $T=$ | 10.77674 | $H 1=$ | 1534.303 | HLEO $=$ | 6402.393 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| - $\mathrm{H}=$ | 7250.000 | $T=$ | 10.52514 | Hi= | 1561.823 | HLEO $=$ | 5900.962 |
| $\mathrm{H}=$ | 7500.000 | $\mathrm{T}=$ | 10.28i63 | H1 $=$ | 1537.997 | HLEO $=$ | 5419.849 |
| $H=$ | 7750.000 | $T=$ | 10.04582 | $\mathrm{H} 1=$ | 1513.000 | HLEO= | 4958.085 |
| $\mathrm{H}=$ | 8000.000 | $T=$ | 9,81.7351 | $\mathrm{H} 1=$ | 1486.986 | HLEO $=$ | 4514.776 |
| $\mathrm{H}=$ | 8250.000 | $\mathrm{T}=$ | 9.595901 | H1 = | 1460.100 | HLEO $=$ | 4089.070 |
| $H=$ | 8500.000 | $T=$ | 9.381143 | $H 1=$ | 1432.474 | HLEO= | 3680.165 |
| $\mathrm{H}=$ | 8750.000 | $T=$ | 9.172782 | $H 1=$ | 1404.236 | HLEO= | 3287.295 |
| $H=$ | 9000.000 | $T=$ | 8.970538 | H1 $=$ | 1375.486 | HLEO $=$ | 2909.765 |
| $\mathrm{H}=$ | 9250.000 | $\mathrm{T}=$ | 8.774145 | $\mathrm{H}=$ | 1346.331 | HLEO $=$ | 2546.801 |
| $H=$ | 9500.000 | $T=$ | 8.583347 | H1: | 1316.364 | HLEO= | 2193.010 |
| $\mathrm{H}=$ | 9750.000 | $T=$ | 3.397915 | H1: $=$ | 1237.193 | HiEEO $=$ | 1862.546 |
| $\mathrm{H}=$ | 10000.00 | $\mathrm{T}=$ | 8.217616 | $\mathrm{H1}=$ | 1257.319 | HIEO= | 1539.918 |
| $\mathrm{H}=$ | 10250.00 | $1=$ | 8.042239 | $\mathrm{HI}=$ | 1227.390 | HLEO= | 1229.577 |
| $\mathrm{H}=$ | 10500.00 | $\mathrm{T}=$ | 7.871587 | H1 $=$ | 1197.443 | HLEO= | 931.0140 |
| $H=$ | 10750.00 | $T=$ | 7.705462 | H1= | 1167.537 | HLEO: | 643.7385 |
| $\mathrm{H}=$ | 11000.00 | $T=$ | 7.543689 | $111=$ | 1137.726 | HLEO: | 367.2815 |
| $\mathrm{H}=$ | 11250.00 | $T=$ | 7.366090 | $\mathrm{H} 1=$ | 1108.058 | HLEO $=$ | 101.2060 |
| $\mathrm{H}=$ | 11500.00 | $\mathrm{T}=$ | 7.235506 | H1= | 1078.575 | HLEO $=$ | $-154.9056$ |
| $\mathrm{H}=$ | 11750.00 | $T=$ | 7.082776 | H $1=$ | 1049.318 | HLEO $=$ | -401.457\% |
| H: | 12000.00 | $T=$ | 6.936757 | H1= | 1020.325 | HLEO:- | -630.8290 |

NOTE: HI is equivalent to $h$, which has previously been used.

For a constant-stress tether (stress $=0$, density $=\rho$ ) extending from $r_{L E}$ to $R=r_{L E}+h$, the area distribution is easily found to be

$$
\begin{equation*}
A(r)=A_{L E} \exp \left[\frac{\mu_{e} \rho}{\sigma}\left(\frac{3}{2 r_{L E}}-\frac{r^{2}}{2 r_{L E}^{3}}-\frac{1}{r}\right)\right] \tag{12}
\end{equation*}
$$

where 'AE' the thickest section, is found from equilibrium at the higher end, where a satellite of mass $M_{\text {sat }}$ is attached:

$$
\begin{equation*}
M_{s a t}\left(\frac{\mu}{r_{I E}^{3}} R-\frac{\mu}{R^{2}}\right)=\sigma A_{L E} \exp \left[\frac{\mu_{e^{\rho}}}{\sigma}\left(\frac{3}{2 r_{L E}}-\frac{R^{2}}{2 r_{L E}^{3}}-\frac{1}{R}\right)\right] \tag{13}
\end{equation*}
$$

Most of the tether mass is concentrated near the lower end, where a good approximation to Eq. (12) can be obtained by series expansion of the exponent

$$
\begin{equation*}
A(r) \simeq A_{L E} \exp \left[-\frac{3}{2} \frac{\mu_{e} \rho}{\sigma r_{L E}}\left(\frac{r-r_{L E}}{r_{I E}}\right)^{2}\right] \tag{14}
\end{equation*}
$$

and this can be integrated to obtain an expression for the tether mass:

$$
\begin{equation*}
\frac{M}{M_{\text {sat }}}=\frac{2}{3} \sqrt{\frac{\pi}{2}} \frac{3+3 \delta+\delta^{2}}{(1+\delta)^{2}} \gamma e^{\gamma^{2 \frac{1+\delta / 3}{1+\delta}}} \operatorname{erf}(\gamma) \tag{15}
\end{equation*}
$$

where

$$
\begin{equation*}
\delta=\frac{h}{r_{L E}} \quad, \quad \gamma=\frac{h}{r_{L E}} \sqrt{\frac{3 \mu e^{\rho}}{2 \sigma r_{L E}}} \tag{16}
\end{equation*}
$$

For $\delta S 0.2, M / M_{\text {sat }}$ is seen tc depend mostly on the nondimensional group $\gamma$ (Eq. (16) ) For $\delta$ in the vicinity of $0.2(\mathrm{~h} \simeq 1350 \mathrm{~km})$, the variation is shown in Table 5 below.

Table 5. Approximate tether mass

| $\gamma$ | 0 | 0.1 | 0.2 | 0.4 | 0.6 | 0.6 | 1 | 1.2 | 1.6 | 2 | 2.5 | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| M/Msat | 0 | . 02397 | . 09748 | . 4172 | 1.0539 | 2.2143 | 4.3293 | R. 2.974 | 32.222 | 147.57 | 1365.3 | $1 \varepsilon .858$ |

For Kevlar, $\gamma \sim 10 \frac{h}{r_{L E}}$ (about 2 for $H=1200 \mathrm{Km}$ ). Thus, for
payloads of the crder of 10 Tonne, unreasonably heavy tether lines are obtained. However, the exponential nature of the $M(h)$ function indicates rapid mass reductions if the lower tether length can be reduced by application of moderately small $\Delta V ' s$.

For the mass of the upper tether, Eqs. (15), (16) still applies if the sign of $\delta$ is reversed; $r_{L E}$ is replaced by $r_{G S}$ and $h$ is replaced by $H$. The effective new value of $\gamma$ is then of the order of 1.1, which indicates a moderate tether mass, of the order of 8 times that of the payload itself. Thus, despite the much greater length of the GEO tether, it is the one in LEO that needs substantial reduction.

The effect of introducing both perigee and apogee firings ( $\Delta V_{p}$ and $\Delta V_{Q}$, respectively) was next investigated. The results for a wide range of paraneters are listed in Tables 6 through 9. For the cases of the $1 / 3$ and 1/2 day period, the results are also displayed graphically in Figs. 3 and 4. The effects are generally as follows:
(a) Increasing $\Delta V_{p}$ at constant per od increases the altitude of LEO, and decreases the lower tether length, $h$.
(b) Increasing $\Delta V_{Q}$ at constant period decreases the altitude of LEO. For low $\Delta V_{P}$, increases of $\Delta V_{Q}$ result in a shorter lower tether, but the reverse is true at high values of $\Delta V_{P}(2800 \mathrm{~m} / \mathrm{sec})$.
(c) As discussed before, the length $H$ of the upper tether is unaffected by either $\Delta V_{P}$ or $\Delta V_{Q}$, but is reduced if the period is allowed to increase.
(d) For each transfer time and each value of $\Delta V_{p}$, there is a maxi$\operatorname{mum} \Delta V_{Q}$ for which the lower Earth orbit becomes too low (a limit of 200 Km was assumed here). Similarly, for each $\Delta V_{Q}$, there is a minimum $\Delta V_{p}$ for the same reason.
(e) The length of the lower tether can be reduced to zero by increasing $\Delta V_{P}$ for each $\Delta V_{Q}$. The effect of $\Delta V_{Q}$ on $h$ is minor.

As an example of a combination which could be useful, we see in Fig. 4 that, from a 500 Km LE orbit, using a lower tether of length 390 Km and supplying a velocity increment $\Delta V_{p} \simeq 1500 \frac{\mathrm{~m}}{\mathrm{sec}}$ after release, a payload can be put into a transfer ellipse leading to capture by the lower and of a GEO tether of 5913 km length, if an apogee velocity increment $\Delta V_{Q} \simeq 725 \mathrm{~m} / \mathrm{sec}$ is applied prior to docking. If docking fails, anotiter attempt can be made after one day (two orbits of the payload). Notice that for this lower tether length, its mass can be of the order of the payload mass.
(

TABLE 7. TRANSFER TO GEO BY TETHER, WITH $\triangle V$ 's SUPPLIED BY PROPULSION
$\mathrm{T}=3 / 8 \mathrm{day} \quad$ Upper Tether Length $\mathrm{H}=8989 \mathrm{Km}$


KEY: Entries are: $h(\mathrm{Km}) /$ Altitude of LEO (Ka) $\quad \mathrm{X}=\mathrm{LEO}$ Altitude $<200 \mathrm{Km}$
$\mathrm{X}=\mathrm{L}$
h
negative

TABLE 8. TRANSFER TO GEO'BY TETHER, WITH $\triangle V^{\prime}$ 's SUPPLIED BY PROPULSION
$T=2 / 5$ day $\quad$ Upper Tether Length: $\mathbf{H}=8271 \mathrm{Km}$
(

TABLE 9. TRANSFER TO GEO BY TETHER, WITH $\triangle V$ 's SUPPLIED BY PROPULSION
$T=1 / 2$ day Upper Tether Length: $H=5913 \mathrm{Km}$


figure 3. $\Delta V_{P}$ (at release)
$\mathrm{m} / \mathrm{s} \in \mathrm{C}$


## . Orbital perturbations of the 1 wo platform

In the calculations so far we have implicitly assumed very heavy platforms both in LEO and in GEO. If the mas of the LEU platform is dominated by that of the Shuttle orbiter (docked to a light free-flying tether facility, the ratio M platform payload may not be very large (3:1 for a 10 tonne payload). The result may be an excessive lowering of the post-release Shuttle perigee. In this section we consider this effect, while still assuming a massive GEO platform.

The new geometrical arrangement for the lower tether is shown in Fig. 5. The orbital center is at $R_{c}$, given by (Ref. 1).


Figure 5. Geometry for a finite lower platform mass.

Ref. 1. Study of Certain Launching Techniques Using Loin Orbiting Tethers by Giuseppe Colombo. Final Report on grant NAG-8008, from the SAO to NASA, Feb. 1981.

$$
\begin{equation*}
R_{c}=\left(\frac{\sum r_{i} m_{i}}{\sum \cdot \dot{m}_{i} / r_{i}^{2}}\right)^{1 / 3} \tag{17}
\end{equation*}
$$

and is located at a distance $h^{\prime}=R_{P_{1}}-R_{c}$ from the transfer orbit perigee. Thus, $h^{\prime}$ replaces $h$ and $R_{c}$ roplaces $R_{L E}$ in our previous analysis (Eqs. (10, (11)) . The new $R_{L E}$ must be obtained from the explicit form of Eq. (17); for example, accounting only for two end masses $M_{1}, M_{2}$ (Fig. 5), we have

$$
\begin{equation*}
R_{c}=\frac{R_{L E O} M_{1}+R_{P_{1}} M_{2}}{M_{1} / R_{L E O}^{2}+M_{2} / R_{P_{1}}^{2}} \tag{18}
\end{equation*}
$$

which can be solved for $R_{\text {LEO }}$.
The perigee of the post-release platform orbit can be calculated from Eq. (6) of Ref. 1, which for our case reads

$$
\begin{equation*}
S_{2}=-R_{L E O}+2 /\left(2 / R_{L e O}-R_{L E O}^{2} / R_{c}^{3}\right) \tag{19}
\end{equation*}
$$

The effect of this modification is to require a longer lower tether and to make high $\Delta \mathrm{V}_{Q}$ values unfeasible 'negative perigee). As an example, Tables 10 and 11 show a comparison (for $1 / 3$ day period) of two cases, one with a massive LEO platform $M_{1}=5000$ Tonne for $M_{2}=10$ Tonne) and the other with a light LEO platform (the Orbiter, $M_{1}=80$ Tonne,. In the first case, where only a slight perturbation is introduced to the orbit, a tether length $h=998 \mathrm{Km}$ can be used from a 521 Km orbit, which becomes a 521/511 orbit after release. Velocity

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TABLE 10. PLATFC.EM IN LEO

M1 $=5,000,000 \mathrm{~kg}$ (platform)
$M 2=10,000 \mathrm{~kg}$ (satellite)
$P=1 / 3$ day
$\mathrm{H}=10,390 \mathrm{~km}$


Entries are $\quad h$ in $k m$

$$
\begin{aligned}
& \text { apogee altitude } / \text { perigee altitude (km) } \\
& \quad(\mathrm{km})
\end{aligned}
$$

TABLE 11. SHUTTLE IN LEO

$$
\begin{array}{rlr}
M 1=80,000 \mathrm{~kg} & \text { (shuttle) } \\
M 2=10,000 \mathrm{~kg} & \text { (satellite) } \\
P=1 / 3 \text { day } & H=10,390 \mathrm{~km}
\end{array}
$$



Entries are
$h$ in km
apogee altitude
(km)
perigee altitude (km)
increments $\Delta V_{P}=300 \mathrm{~m} / \mathrm{sec} . \quad \Delta V_{Q}=100 \mathrm{~m} / \mathrm{sec}$ are required. In the case with the light platform, the $\Delta V_{Q}=100$ is not allowable, and so, for $\Delta V_{p}=300 \mathrm{~m} / \mathrm{sec}$, only $\Delta \mathrm{V}_{\mathrm{Q}}=0$ is possible. The result is a longer tether ( 1155 Km ) and a higher orbit (1291/656).

## Work forecast.

In the following period, we anticipate progress on the scilowing points:
(1) Completion of a trade-off study on the feasil a extensions of the Shuttle missicn envelope by an on-board tether system.
(2) Construction of an operational map fc: the LEO-GEO two-tether system, including the bounds dictated by excessive tether mass and post-release perigee.
(3) Initial definition of propulsion systems for restoring platform orbits.

