



# **Study of Auxiliary Propulsion Requirements** for Large Space Systems

**Volume 2 Final Report** 

**Boeing Aerospace Company** 

prepared for

### NATIONAL AERONAUTICS AND SPACE ADMINISTRATION

### **NASA Lewis Research Center Contract NAS3-23248**

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#### **ABSTRACT**

This document is the final report prepared under contract NAS3-23248, "Study of Auxiliary Propulsion Requirements for Large Space Systems." There are three technical tasks described herein. In Task 1 a range of single shuttle launched large space systems were identified and characterized including a NASTRAN and loading dynamics analysis. Task 2 consisted of an analysis of the disturbance environment, characterization of thrust level<br>and APS mass requirements, and a study of APS/LSS interactions. In the final task, state-of-the-art capabilities for chemical and ion propulsion were compared with the generated propulsion requirements to assess the state-of-the-art limitations and benefits of enhancing current technology.

#### KEY WORDS

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Attitude Control Auxiliary Propulsion Large Space Structures Shape Control Stationkeeping

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#### S**UMMA**RY

**-**"**T.,**eobjective **o**f contract NAS**3**-21**95**2, **"S**tudy **of** \_]**x**iliary **P**r**o**puls**io**n Requirements f**o**r Large Space Syst**e**ms," was to establish key A**P**S requirements and state-of-**t**he-art improvement benefits for a range of future spacecraft. The key issues raised or examined in the study included LEO deployment and o**p**eration, GBO stationkeeping duty cy**c**les, structural mode**l**ing, and th**e** technology a**rea**s t**o** impro**v**e. To **e**xamine these is**s**ue**s**, three technical task**s w**ere performed. These tasks a**r**e sh**ow**n bel**o**w:

Task i: **D**efinition of Advan**c**ed Structura**l** Con**c**e**p**ts

Task 2: E**s**t**a**bli**s**hment of A**P**S Re**q**uirements

Task 3: Assessment of Technology Improve. At Benefits

In Task !, six ve**h**icle classes were analy**z**ed including four large antenna structures and two s**p**ace platform desi**g**ns. N**AS**\_RAN analysis was condu**c**ted to deter**mi**ne the mode sha**p**es, **s**tructu**r**al frequen**c**ies, and to **v**e**r**ify a dyna**m**ic loading analysis. The loading anal**y**sis wa**s** condu**c**ted to dete**rm**ine **:** the e**f**fe**ct o**f pr**i**m**a**r**y thr**u**s**t g-l**oa**ding on the **mass p**r**op**ertie**s o**f **e**ach **LSS**.

**Ta**sk **2 was accomp**l**ish**ed **in fo**ur **step**s. **A** di**st**urba**nce a**na**lysis fo**r L**E**D d**epl**o**y**m**en**t **altit**udes and **GEO op**era**tio**na**l** a**ltit**ude**s w**as **fi**r**st** per**fo**r**m**ed. **The fo**r**c**e a**n**d **torq**ue re**q**u**ir**em**ents fo**r **e**a**ch L**E**S w**ere u**s**ed **t**o es**t**cb**lish** the thru**te**r re**q**u**i**reme**nts**. **Th**e**s**e re**qui**r**e**me**nt**s ind**i**ca**t**ed a need **for wi**de**ly seFa**ra**t**ed thru**st l**e**v**e**ls** be**tw**ee**n LEO** an**d G**\_**D and indicat**ed th**r**o**ttli**ng **g** r**eq**u**i**rem**e**n**t**s **fo**r **GEO o**pera**tion** o**f 2**:**1 t**o **6:1**. **In** the ne**xt s**ub**t**a**s**k **APS** ma**s**s **was c**h**a**ra**cte**r**iz**ed **fo**r mo**n**o**p**r**o**pe**ll**an**t,** b**ip**r**op**e**ll**an**t**, and **i**o**n** sy**st**em**s**. **It was f**ou**nd** th**at** du**ty cycl**e **play**e**d** the **maj**or ro**le** in de**te**rm**in**ing w**hich p**r**op**u**lsion** s**yst**em **was** th**e** mo**st v**iab**l**e **f**or **GEO** ope**r**a**ti**o**n**. **D**u**ty** c**ycl**es **of 3 h**o**u**r**s**/**o**rb**it** or gr**eat**er **favo**red **ion** sy**st**em**s**, wh**e**r**e**a**s s**hor**t**er du**ty cycl**e**s** re**q**u**i**red thrus**t l**eve**ls** wh**ich c**ou**l**d **n**ot be me**t with st**a**te-**o**f**-**th**e**-**a**rt** / pr**o**pul**s**i**o**n. A**s** a final a**na**ly**si**s, the interact**i**on\_ o**f** the pro**p**ul**s**ion sy**s**te**m w**ith the structure were examined. F**o**r th**r**u**s**t level**s** o**f** 7 to **3**0 Ne**w**ton**s** (depending upon the configur**a**t**i**on) **s**igni**f**icant de**f**ocu**s**ing o**f** amtenna sy**s**t**e**ms **w**as foun**d**.

In Ta**s**k **3** t**he** state-o**f-**the-**ar**t ca**p**abilitie**s f**or m**o**no**p**r**op**ellant, ': ipr**o**pellant, **a**nd ion **s**y**ste**m**s w**er**e** dete**r**mined. The**s**e **c**a**p**abilitie**s we**r**e** C c**c**mpsred **w**it**h** the requi**r**ement**s** generated in Ta**s**k **I** and limitation**s** t**o** There were four major limitations ', iden**t**i**f**ied and these limitation**s** a**r**e sh**o**wn below.

- **•** Mono**p**ro**pe**l**l**ant **I**s**p** limits m**is**sion lifetinm (<**5**-7 **y**eev**s) f**or centau**r** " G' de**l**ivery ca**pa**bilit**y**
	- **; B**i**pr**o**pell**a**n**t n**eed** lo**w**er thr**us**t ca**pa**bi**l**i**ty** (<**2**-**5 N**)

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These limitations were a key to assessing the enhan**c**ed technology benefits. A brief listing of those technology areas which would enhance or enable the LSS mis**s**ions identified **i**s shown below. \_'

- Increasing chemical system I<sub>Sp</sub> to > 300 seconds is mission enabling
- Minimum firing times of < .01 seconds yields mass advantage of jets over MMD's for 3-axis control
- Valve cycling of 2 x  $10^7$  cycles enables jet systems for 3-axis control
- Thruster levels **o**f .i to .4 N enhance ion prop**u**lsion f**o**r GBO operation
- Isp range for ion systems of 1000-2000 seconds optimum (using state-of-the-art PPU's)
- Ion power system mass must be red**u**ced for i**o**n systems to be competitive with shorter duty cycle, higher thrust engines

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## Int<u>roduction and</u>

With increasing fervor, plans to utilize the resources of space are being made wi**th**in NASA, DOD, and private industry. Many of these plans call for the use of Large Space Systems (LSS) to accomplish this wide variety of goals. These LSS will zequire new technology in analysis techniques and hardware to be enabled and utilized in the most cost effective fashion. To assess the propulsion technology requirements and recommend high leverage<br>advances in propulsion, a study was performed examining auxiliary in propulsion, a study was performed examining auxiliary<br>requirements for a range of single souttle launched LSS. This propulsion requirements for a range of single shuttle launched LSS. study considers auxiliary propulsion only &Id will supplement othe work examining prime propulsion requirements (Ref. 1).

This study is a more focused follow-on to contract NAS3-21952 (Ref. 2),<br>"Study of Electrical and Chemical Propulsion Systems for Auxiliary of Electrical and Chemical Propulsion Systems for Auxiliary Propulsion of Large Space Systems." The focus of this study was narrowed to examine only single shuttle launched LSS with two excep**t**ions - the Space Operations Center (SOC) and the Science and Applications Space Platform (SA**L**P). Also in this study, only advanced deployable LSS which had a heritage of ongoing preliminary design were examined. To add a final sharpening of focus, only well established propulsion options were examined and extrapolation of their capabilities was rooted in accepted scaling laws based on theory, test, and existing hardware. By establishing a narrow width for analysis, the dep**t**h of the analysis was enhanced. Details of the , effects of primary g-load on structure mass and the effect of a range of thrust levels on antenna performance were examined. Stationkeeping duty cycles and tolerance effects were studied and regions of operation for each propulsion system identified. The study was also able to make specific recommendations for auxiliary propulsion thrust level, I<sub>sp</sub>, minimum impulse bit, and cycle number for the range of **LS**S identified.

Several key assumptions were groundruled in the study. These assumptions are listed below. The first two assumptions and the propulsion option The first two assumptions and the propulsion option assumption have been previously discussed.

- \_ **•** Single shuttle launched (exception SOC, SASP)
- Advan**c**ed prel**i**minary design deployab**l**e LSS
- L**ED** (300-500 km) and GHD operat**i**on
- i0 year mission l**i**fe

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- NASA neutral atmospheric model assumed
- ( Only well established propulsion options examined (mono, bipropr ion)

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. No factors of conservatism were employed

For **t**he antenna systems the **L**E**D** al**t**itude range given i**s** the **assu**med deplo**y**ment alti**t**ude. Th**i**s range **c**orresponds **t**o the STS delivery ( ca**pa**bili**ty**. **Th**e**s**e **s**y**st**em**s ar**e th**en tr**a**ns**fe**r**red **to GED** fo**r th**e I0 **year** mi**ssi**on **o**pera**ti**on. Th**e** s**pac**e **pl**a**tf**orms examined are as**s**umed **t**o opera**t**e in **,** \_he LEO al**t**i**t**ude range shown.

The NASA neutral atmosphere was used as a basis of comparison for L**E**O torque and drag makeup calculations. This atmosphere is a worst case, long te**r**m density model and yields conservative \_]t realistic worst case results. Other models are discussed in Task 2. No contingency factors were used for propellant or thrust level calculations. It was felt that the application of such factors might vary with each mission, and because the purpose of the study was to illuminate trends in propulsion requirements, a solid basis of comparison was needed.

Several key issues arose in the **c**ourse of the study which drove the propulsion requirements and technology recommendations. These issues are listed below.

- Structural modeling
- LEO deployment and operation
- GBO operation duty cycles
- A**P**S system mass impacts

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The importance of having detailed and accurate structural models became • The importance of having detailed and accurate structural mode<br>
• very clear when the issue of thruster interaction was examined.<br>
• defocusing analysis is sensitive to section property and materials<br>
• assumptions which defocusing analysis is sensitive to section property and materials property<br>assumptions which must be modeled in NASTRAN to obtain the fundamental assumptions which must be modeled in NASTRAN to obtain the fundamental or frequencies and mode shapes. **S**m**a**ll differences in wall thickness or  $\frac{1}{2}$  section spacing can result in-significantly different interactions. The results presented in Task 2 are based on numerous iterations of mass results presented in Task 2 are based on numerous iterations of mass properties and section properties to match previous results of actual hardware tests (Ref. 3).

Operational issues at LEO and GEO are primary drivers for thrust level and Isp requirements. LED deployment drove stationkeeping thruster size a 1<br>propellant mass for even short stays at low altitude to such a degree the **;** propellant mass for even short stays at low altitude to such a degree tl**.** • L**ED** deploymen**t** seemed unadvisable f**o**r mos**t** IZS. GEO du**t**y cy**c**les **w**er\_ another key issue in the study because as duty cycle changed fr**o**m a few m**i**nutes a week **t**o a few hours per orbit, thrust requirements went from chemical capability to electric thruster capability. Longer duty cycles , would require au**t**onomous operation or high ground in the loop software c**o**sts.

**A**u**xi**lia**r**y pro**p**ul**sio**n **sy**s**t**em ma**ss c**a**n** be **3**0**-5**0% **o**f **t**he **to**tal sys**t**em m**ass** weing chemical systems for 10 year GEO missions. Reductions of this percentage by only 5 or 10% allow very large mass savings and hence lower \_ per**c**entage by onl**y 5** o**r** 10% all\_ very **l**arge mas**s** sav**i**ng**s** and hen**c**e lowe**r** : launch **c**o**sts**. To e**f**fec**t** the**s**e change**s**, sta**t**e-of-the-art limita**ti**on**s** in • **c**hemical **I**sp, power proc**ess**o**r** ma**ss**, and au**t**onomou**s o**pera**ti**on mu**st** be \_*.* over**c**ome as de**s**cr**i**bed in Ta**s**k **3**.

The program task flow is shown in Figure 1. Task 1 determined the relevant missions and spacecraft propertieswhich would be used to define propulsion requirements. The NASTRAN models and loads analysis gave us an insight into the variation of mass with primary thrust g-loading and were also used to determine the APS/L**S**S interactions in Task 2. Thrust requirements, impulse bit requirements, Isp effects, and hardware masses were determined in Task 2. These requirementswere comparedwith current capabilities and a set of limitations found in Task 3. The benefits in terms of enhanced mission capture and reduced APS mass were assessed in the final analysis of Task 3.



FIGURE 1. PROGRAM TASK FLOW

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#### 1.0 DEFINITION OF ADVANCED STRUCTURAL CONCEPTS

The objective of Task 1 was to select and characterize classes of large space systems. To accomplish this objective, three major subtasks were executed. The first subtask involved the selection of configurations which<br>met the study assumptions outlined in the introduction. These met the study assumptions<br>configurations were selec selected from previous studies, other relevant literature, and personal contact with potential users of deployable LSS. The second subtask was to perform a NASTRAN analysis of the selected configurations. To do this, detailed section properties were formulated using existing preliminary designs or in some cases derived section properties from known data such as structural frequency and **[**samber mass properties. NASTRAN models were then developed and a normal modes analysis The final characterization subtask estimated the effects of changing the primary transfer acceleration for LEO-GEO transfer of the deployed LSS. Loading equations were develoed which showed the sensitivity of system mass to changing acceleration levels. Key assumptions in Task 1 are outlined in Table 1 below.

#### Table i. Task 1 Assumptions

- 1.1 Selection of Configuration
	- Civilian GHO missions (exception SOC & SASP)
	- Single shuttle launched
	- Preliminary designs available

#### 1.2 NASTRAL Analysis

- : extending was used to the level of detail available
	- Section mass properties & g-load were given from the literature
	- ,,Analysis conducted for .15 g's
- **r** 1.3 G-Loading Effects<br> **1.3** G-Loading Effects<br> **1.3** G-Loading Effects

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- \_ Indiv**i**dual elements scaled
- Uniform mass dis**t**ribution for each element

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In Subtask 1.1 the focus of the study is shown to center around antenna systems that are deployed in LEO by a single shuttle and then transferred  $'$ to GEO for their operational life. The NASTRAN analysis was conducted for •15 g's because the section properties available showed **t**hat .i!\_g's strength produced the first modal frequency expected of .I Hz. This frequency was determined to be reasonable based on the data available in the preliminary designs and on the Boeing Company's previous experience<br>with other LSS. In the g-loading analysis critical elements such as In the g-loading analysis critical elements such as antenna support booms, solar array booms, and other truss structures were analyzed individually which resulted in a very detailed g-load sensitivity model. Differences in the overall scaling of mass with g-load in this study and others (Ref. i) were attributed to this detailed modeling.

The key issues which drove the selection and characterization of each **L**SS were threefold. First, the mission opportunities were made as broad as p**o**ssible. Mission selection included electronic mail, direct T**V** broadcast, mobile communications, forest fire detection and others. Configurations selected spanned all of these missions and are felt to be representative of \_**:** LS**S** f**o**r the 1**99**0 's. Conf**i**gu**r**ati**o**ns \_ere als**o** selected so as **to** span the Wide variations in center of pressure/center of gravity, inertia matrices, and area mass ratios were sought. Critical structure element definition was also a key issue because<br>it drove the NASTRAN modeling results and consequently the sought. Critical structure element definition was also a key issue be .<br>. The the NASTRAN modeling results and consequently thruster/structure interactions conclusions.

# F" i.i Selection of Configurations ,[**-**

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A set of six preliminary designs which pose a wide range of propulsion requirements was selected in this subtask. The designs are sufficiently detailed to allow good fidelity in defining A**P**S requirements and APS/LSS  $\therefore$  interactions. At the same time, the same time, the same time, the same time, the same used for a variety of  $\lambda$  purposes by simply changing the electronics and/or scaling the LSS either up or down in size. The classes and representative missions are in Table 2 • and discussed in more detail in Appendix A.

Each of the generic classes selected fits the study ground rules of a<br>single shuttle launch with the exception of the SOC which will be<br>considered as representative of a multiple shuttle launched LSS of the<br>1990's. The sca • single shuttle launch with the exception of the S0C which will be considered as representative of a multiple shuttle launched **L**SS of the The scaling of each LSS from the baseline also fits the ground rules of a single shuttle launchable **L**SS with an assumed orbi**t** transfer propulsion **t**o GBO. Figures 2, **3**, and 4 show the configura**t**ions selected \_ for anal**ys**is in **t**h**is** study.

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FIGURE 2. PHASED ARRAY AND WRAP RIB ANTENNAS

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# FIGURE 3. HOOP COLUMN ANTENNA AND<br>GEOPLATFORM DESIGNS

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FIGURE 4. SASP AND SOC SPACE PLATFORMS

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The Large Aperture Phased Array and SASP were sized using two already designed versions with the main difference in each design being the power requirement. Therefore, the main sizing change is in the solar arrays. The SOC is sized using two of its designed versions - the initial baseline and the operational basellne. The operational baseline is a growth of the initial baseline and has more elements, mass, and area. The other three generic classes will be sized using scaling outlined in Appendix A.

Four generic classes were chosen to investigate for their g-loading<br>characteristics. These are Large Aperture-Phased-Array-Antenna,-IMSS-These are Large Aperture Phased Array Antenna, LMSS -<br>200 Column, and the Geostationary Platform. Each of Wrap Rib, IMSS - Hoop Column, and the Geostationary Platform. these structures is designed with critical elements closely associated with flexible members and, therefore, most susceptable to changes in mass and packaging characteristics due to g-loading. Three g-loading designs will be determined for each of these four structures and sizes.

The SOC and SASP were not studied for g-loading effects because of their relatively rigid structure. Their critical elements are concentrated in a central and rigid mass area with only the solar arrays extending from flexible members. Solar array support structures may be stiffened, if necessary, without significant impact on mass or auxiliary propulsion requirements. An additional reason not to look at varying g-loading characteristics of tl**.**e two space platform designs is to avoid a proliferation of separate designs, sizes, and g-loading parameters. Such a large number of discrete cases would take away from the major thrust of this study. Table 3 shows the sizes selected for each generic class.

#### 1.2 NASTRAN Analysis

NASA Structural Analyzer (NASTRAN) models were developed to be used in the thruster/structure interactions study conducted in Task 2. To develop these models detailed mass property and section properties had to be determined. Once the mass, material, and dimensions of a given element was de**te**rmined, the va**ri**ous secti**o**n**s** could be l**i**nked tog**e**ther b**y** spec**i**fying the stiffness properties at the interfaces. In addition to the interface point, node points along truss work, columns, and other elements had to be specified. This proce**s**s was a tedious one and forced the study to consider only the antenna systems for NASTRAN modeling.

To avoid duplication and thus save time, general models of components common to more than one class of LSS (i.e., solar array**s**) were defined. These components are combined with componen**ts** unique to each class of structure. Figure 5 show**s** the gene**r**al model of jus**t** such a component - the solar array. When the mode**l**s we**r**e completed, we used a NASTRAN dynamic analy**s**i**s** to de**t**ermine the stru**ct**ural f**r**equencie**s** and mode shapes of each LSS design.



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Table **3**. Generi**c** Class Sizes

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# Large Aperture Phased Array &"

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The following is a description of the NASTRAN model of the Large Aperture Phased Array employed for the analysis. All components described are labeled in Figure **6**. The column, array as**t**rcma**sts**, and main astr**c**masts were modeled as triangular trusses made of graphite epoxy tubes with tie rods. The antenna rim was modeled as a graphite/epoxy tube in twelve The antenna rim was modeled as a graphite/epoxy tube in twelve segments. The lens staves and stabilizing lines were assumed to be graphite rods. Section properties for all members are shown in Table 4.

The three lens films in the antenna were assumed to be 2 mils  $(5.067 \times 10^{-5})$ m) thick each. Design stress for each lens was assumed to be  $20 \text{ n/m}^2$ . The ten**s**i**o**n in the three lens wa**s** modeled with rod elements between the rim and coltmm. Fach solary array was modeled as a ba**r** with rigid body elements across its width. The feed horn cluster and cylinder were modeled as a rigid body.

The weight breakdown for the entire structure is summarized in Lable 5. The weight of the lens (includes the weight of the phase shifters)  $\Delta s$  distributed as follows: 50% at twelve points around the rim, 45% at the center **o**f th**e** column, an**d** 2.5% **a**t each end **o**f the coltmm. The weight **o**f \_" the antenna rim wa**s** distr**i**buted evenly along it**s** length as **w**a**s** that **o**f th\_ .: column. The we**i**gh**t of** the antenna rim wa**s** distributed evenl**y** \_I**o**ng it**s** length as was that of the column. The weight of the main astromasts was also distributed evenly along their lengths. For the array astromasts the weight of the wiring and astromasts was divided in half and lumped at either end - half at the feed horn, half at the edge of the array. The  $\sim$  either end - half at the feed horn, half at the edge of the array. The array  $\sim$ , weight of the antenna feed, all the equipment inside the cylinder was lumped at the center of the cylinder. The arra**y** orientation equipment was assumed to be on the main structure, so its weight was also lumped at the center of the cylinder. The weight of the sola**r** arra**ys** was lumped at : either end of the panel assuming the following breakdown:

• 82% panel

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- 10% boom along center of panel
- 5% box **i**n which panel i**s** stored
- \_' **3**% sti**ff**ening rod **a**t f**a**r end of panel

Preten**s**ion loads of **6**.0 N in the stave**s** and 1**2**.0 N in the stabili**z**ing lines Free applied to prevent these members from going slack under a 0.06 g load applied at the antenna feed. Pretension loads in the staves and lens produced a differential stiffness matrix. This matrix was input to a :j,, **p**roduced **a** di**f**f**e**rent**i**al sti**ff**ne**ss** mat**r**ix. Thi**s** matr**i**x **w**as input to a NAS\_RAN Normal Modes Ana**ly**sis t**o** find the nat**u**r**a**l frequen**c**ies and mode : shapes of the st**r**uctu**r**e. The fi**r**st, second, an**d** thir**d** mode frequen**cies** are .0**93**, .ii**0**, and .1**6**0 **Hz**. Mode shape**s** fo**r** these three mode**s** are shown in **F**igure 7, **8** an**d** 9.

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FIGURE 6. LARGE APERTURE PHASED ARRAY ANTENNA -NASTRAN MODEL



PHASED ARRAY ANTENNA ASSUMED SECTION PROPERTIES TABLE 4.

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TABLE 5. MASS BREAKDOWN

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#### LMSS - Wrap Rib

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The following is a description of the NASTRAN model developed from the design of the Land Mobile Satellite System (LMSS) - Wrap Rib employed for the analysis. The components of the IMSS are labeled in Figure I0.

'l**h**e inbound and outbound UHF booms, boom housing, solar array mast, solar array booms, and s-band reflector boom were m**o**deled as triangular graphite/epoxy trusses. The UHF feed support and s-band feed boom were modeled as hollow graphite/epoxy tubes. The radial ribs in both reflector surfaces were modeled as hollow, collapsable, lenticular-shaped tubes of<br>graphite/epoxy. The circumferential rings in both reflector surfaces were The circumferential rings in both reflector surfaces were<br>te rods. The bus was modeled as a solid cylinder. The modeled as graphite rods. The bus was modeled as a solid cylinder. molybenum mesh surface of each reflector was modeled as a thin membrane. The mesh was assumed to have the stiffness of a membrane 2 mils (5.067E-5 m) thick made of Kapton  $(E = 1.0E9 \text{ N/m})$ . Although the actual reflectors will have pretension in the ribs and mesh to hold the mesh taut, there was insufficient design detail to incluse the effects of pretension at the time ; this analysis. The section properties of all members are shown in Table 6.

> Each solar array was modeled as a bar with rigid body elements across its width. The UHF feed was modeled in the same way.

The mass breakdown for the entire structure is summarized in Figure 11. The mass breakdown for the entire structure is summarized in Figure 11.<br>The mass of the UHF booms and cables was distributed evenly along the<br>booms. The mass of each reflector was lumped at the center of each<br>reflector. Th booms. The mass of each reflector was lumped at the center of each reflector. The mass of the UHF feed was lumped as follows: 1/4 m**a**ss at :**:**he top edge and 3/4 mass at the bottom edge. The mass of the s-band feed was lumped at a point 0.3 meters from the end of the bus (0.3 meters in the x-y plane as shown in Figure i0). The mass of the following were lumped in the plane as shown in Figure 10). The mass of the following were lumped in the bus: RF electronics, control equipment and sensors, 1/2 solar array mast and mechanism, batteries and power conditioner, and the bus structure, and mechanism, batteries and power conditioner, and the bus structure, exabling, T/C and cage. The mass of the outbound reaction wheels and<br>sensors were lumped at the center of the UHF reflector. The mass of the sensors were lumped at the center of the UHF reflector. The mass of the inbound reaction wheels and sensors and 1/2 the mass of the s-band ; reflector boom were lumped in the boom housing. The mass of the solar array booms and 1/2 the mass of the solar array mast was lumped at the end of the solar array mast. The remaining mass of the solar arrays was lumped part of the solar array mast. The remaining mass of the solar arrays when the following mass breakdown:<br>at either end of each panel assuming the following mass breakdown:

 $\frac{1}{2}$  82% panel 10% bo**o**m along center of pane.l **5**% bo**x** in **w**hich pane**l** i**s st**o**r**ed \_ 3% stiffening **r**od at far end of panel

A NAS\_RAN Normal Mode**s** Anal**ysis w**as employed **t**o f**i**nd the na**t**u**r**a**l** frequencies and mode shapes of the structure. The first, second, and third < mode **fr**equen**ci**e**s are** 0.**105**, **0**.i**ii**, and **0**.1**3**1 **H**z. Mod**e** shape**s for** th**e**se th**r**ee mode**s** are shown in F**i**gu**r**e**s** 12, 1**3** &\_d 1**4**.

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FIGURE 10. LMSS WRAP RIB - NASTRAN MODEL

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TABLE 6. LMSS - WRAP RIB ASSUMED SECTION PROPERTIES (CONTINUED)

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\*\* A = combined x-sectional areas of all three longerons for triangular trusses

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565  $LBS$ 740  $90$ 210 151  $\overline{171}$  $2571$ 500  $\mathbf{a}$ 367  $\overline{10}$ 288 850 6791 **WEIGHT**  ${\bf 50}$ 3040 KGS<br>336 166 386  $78$ liuó 256  $\overline{131}$  $\overline{4}$  55  $68$ 227  $\ddot{ }$ **TOTAL** 8. REACTION WHEELS & SENSORS - OUTBD 16. BUS STRUCT, CABLING, T/C, & CAGE 10. CONTROL EQUIP & SENSORS - BUS 1. UHF REFLECTOR & BOOM ATTACH 4. S-BAND BOOM & REFLECTOR 14. SOLAR PANEL MAST & MECH 15. BATTERIES & POWER COND' 9. REACTION WHEELS - INBD 2. Ulf BOOM-OUTBD & CABLE 3. UHF BOOM-INBD & CALLE 7. RF ELECTRONICS IN BUS 13. SOLAR PANELS 5. S-BAND FEED 6. UNF FEED LTEM  $\tilde{\mathbf{e}}$ CONFIGURATION 4A  $\sum_{i=1}^{n}$  $\mathbf{u} \cdot \mathbf{y}$ 

FIGURE 11. LMSS WRAP RIB WEIGHT STATEMENT

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FIGURE 12. LMSS WRAP RIB

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# FIGURE 13. LMSS WRAP RIB





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FIGURE 14. LMSS WRAP RIB

Based on the preliminary results, the finite element model will be modified as follows: the UHF feed support will be stiffened, and the mass of the UHF reflector ribds and mesh will be distributed to several grids on the reflector surface, rather than lumped at the center.

## LMSS - Hoop and Co!\_n\_**.**

The NASTRAN finite element model shown in Figure 15 consists of 282<br>crid-puints, and 555 elements and has 279 dynamic degrees-of-freedom. The grid-puints and 555 elements and has 279 dynamic degrees-of-freedom. The central column is a telescoping open lattice truss structure made of graphite epoxy tubes and cables. It has a hexagonal cross section and is modeled as 24 bar elements all with the same cross sectional area but with area m**o**ments which are a function of the radius of each segment. %**h**e hoop is modeled as a .076 m diameter, 1. mm wall hollow graphite/epoxy tube. To reduce the number of degrees-of-freedom, the hoop is divided into 24 segments instead of the 48 segments shown in **c**onfiguration drawings. The cables which support the hoop and the reflector surface are .00127 m diameter graphite rods (Celion fiber). To simplify the complex system of cables used to support and manage the reflector surface's shapes, the ; surface is modeled as a single gridwork of membrane elements bounded by graphite reinforc**e**ment ties (rod elements) to which the conical cable graphite reinforcement ties (rod elements)<br>arrays are attached. The molybdenum mesh<br>characterisites are assumed to be approximated<br>thick Kapton film (E = 1.0E9 N/m<sup>2</sup>). To provide<br>antenna, the cables arranged in a "bicyle characterisitcs are assumed to be approximated by a 2 mil  $(5.08E-5$  meter) thick Kapton film (E = 1.0E9 N/m<sup>2</sup>). To provide torsional stiffness to the : antenna, t**h**e cables arranged in a "bicyle spoke" configuration. Otherwise there would be no torsional stiffness until geometric nonlinearities become effective.

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Each of the four feed assemblies is modeled as a lumped mass located at its center and supported by a bar element sized to represent the characteristics of the tripod support shown in configuration drawings. Each solar array is modeled as a flexible bar with rigid elements across enter and supported by a bar element sized to represent the<br>characteristics of the tripod support shown in configuration drawings.<br>Each solar array is modeled as a flexible bar with rigid elements across<br>its width. The +Z graphite/epoxy tubes with .2 mm wall thickness. The -Z solar array support<br>booms are modeled as triangular graphite/epoxy truss beams. The section booms are modeled as triangular graphite/epoxy truss beams. properties of all structural elements are shown in Table 7.

,*:* .**.** The mass breakdown for the Hoop and Column LM**S**S is shown in Figure 16. The mass of the column is distributed uniformly along its length (the • circumferential and diagonal structural elements are assumed to be a small<br>• percentage of the total column mass). The hoop mass and reflector mass are<br>also uniformly distributed. The feed mass is divided between the fou percentage of the total column mass). The hoop mass and reflector mass are<br>also uniformly distributed. The feed mass is divided between the four The feed mass is divided between the four :: feeds and lumped at the cen**t**er **o**f ea**c**h. The mass distr**i**bution f**o**r each solar array is as follows:

82% panel

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- 10% boom along center of panel
- 5% bo**x** in which panel is stored
- 3% stiffner at outboard end of panel

The mass of the solar array support boom is divided between the inboard end<br>of each solar array and the bus to which it is attached. The mass of the of each solar array and the bus to which it is attached. The mass of the \_ s-band feed and reflector are lumped in the +Z bus.

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FIGURE 15. HOOP & COLUMN ANTENNA (LMSS) UNDEFORMED SHAPE

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TABLE 7. ASSUMED SECTION PROPERTIES OF HOOP & COLUMN LMSS

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TABLE 7. ASSUMED SECTION PROPERTIES OF HOOP & COLUMN LMSS (CONTINUED)

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\*\* A = COMBINED X-SECTIONAL AREA OF ALL LONGERONS

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 $\mathfrak{G}% _{M_{1},M_{2}}^{\alpha,\beta}(\mathfrak{g})= \mathfrak{G}_{M_{1},M_{2}}^{\alpha,\beta}(\mathfrak{g})$ 

Figure 16. Hoop & Column LMSS Mass Breakdown

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A NAS"RAN buckling analysis was performed to obtain the differential stiffness matrix resulting from pretension in the hoop cables and the reflector surface. The preloads were applied to both the forward and aft hoop cable arrays and to the outboard edge of the reflector surface. The hoop cable arrays and to the outboard edge of the reflector surface. tension required in each aft hoop cable to prevent the forward cables from going slack at an orbit transfer g-load of 0.15 g is calculated to be 42. N. A reflector surface tension force at each of the 24 hoop attachments is aribitrarily chosen to be I0.0 N. The resulting differential stiffness maxtrix was used in a NASTRAN normal modes analysis to determine mode shapes and natural frequencies. The first three modal frequencies are 0.114, 0.118 and 0.138 Hz and are associated with m**o**tion of the +Z solar arrays (Figures 17, 18 and 19). The first node which involves significant motion of the hoop/column is the ninth flexible mode whose frequency is 0.6 Hz. This mode is a torsion mode (Figure 20) where the column (including feeds and solar arrays) rotates about its axis and the hoop rotates in the opposite direction.

### **Experimental Geostationary Platform**

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The dynamic analysis of the Baseline Experimental Geostationary Platform 9BXGP) is described below. Where data was available, dimensions and merber sizes are those given by General Dynamics, Convair Division. The wrap-rib and peta antenna booms, and the array astromasts were modeled as "Convair deployable trusses.**"** The array support booms were modeled as triangular trusses (1/2 Convair deployable trusses). All trusses were assumed to have graphite/epoxy longerons and graphite tie rods between the longerons. The central core was modeled as a solid cylinder and was assumed to be a rigid<br>Character and experiments 301, 501, 502 and 604 are attached to<br>the core structure with "Convair space rails." Each space rail was modeled .- body. The radiator and experiments 301, **5**01, **5**02 and **6**04 are attached t**o** the core structure with "Convair space rails." Each space rail was modeled as a graphite/epoxy-honeyc**a**mb sandwich plate connected to a graphite/epoxy longeron with graphite tie rods to form a triangular cross section (see Figure 15). The section properties of all members are shown in Table 8.

The mass breakdown for the structure is summarized in Figure 21. The mass of the reflector booms was distributed evenly along each boom. The mass of each reflector was lumped at its center. The masses of the radiator, experiments 310, 501, 502 and 604, along with 1/2 the mass of each space ; rail was lumped at the outboard end of the space rails (501 and 502 are \_: attached to the same space rail). The mass payload 401 was lumped at a point approximately 5 meters from the center of the core. The mass of '\_ payload 123 was Itm\_ed on the space rail which supports payloads 501 and \* 502. The solar array mass was lumped at either end of each panel assuming<br> $\frac{1}{2}$  the following mass breakdown: the following mass breakdown:

82% Panel

- = 10% Astr**o**mast along center of panel
- **'**: **5**% Box in **w**hi**c**h panel i**s** st**o**red
	- 3% Stiffening red at outboard end of panel

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FIGURE 17. HOOP & COLUMN ANTENNA (LMSS) - 1st MODE, 0.114 HZ

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FIGURE 20. HOOP & COLUMN ANTENNA (LMSS) - 9th MODE, 0.601 HZ

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TABLE 8. GEOSTATIONARY PLATFORM ASSUMED SECTION PROPERTIES

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\*\* A = COMBINED X-SECTIONAL AREA OF ALL LONGERONS (AND PLATES FOR SPACE RAILS)

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# FIGURE 21. EXPERIMENTAL GEOSTATIONARY PLATFORM MASS BREAKDOWN

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One-half the mass of the array su\_Dport booms was lumped at t**h**e inboard end of each panel. The mass of the following was lumped in the core: UHF feed, peta feed, 1/2 each space rail, 1/2 solar array support booms, 1/2 packing boom, and the core structure itself. The other 1/2 of the packing boom mass was lumped at a point 5.7 meters above the core (at the top of the packing boom). The mass of the wrap-rib drive was lumped at the intersection of the horizontal and vertical wrap-rib trusses.

NASTRAN normal modes analysis was employed to find the natural frequencies and mode shapes of the structure. The model consisted of 22 lumped masses with a total of 87 dynamic degrees of freedon. The first, second, and third mode frequencies are 0.096, 0.133, 0.145 Hz. Mode shapes for these two frequencies are shown in Figure 22 and 23. The first two for these two frequencies are shown in Figure 22 and 23. modes are primarily bending of the solar array support booms. The third mode is a combination of solar -tray support boom, and peta and wrap-rib antenna boom bending.

# 1.**3** G-Loading and Mass Pr**o**perties

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Four generic clas**s**es were chosen to investigate for their g-loading characteristics. These are large Aperture Phase Array antenna, LMSS - Wrap Rib, LMSS Hoop-Column, and the Geostationary Platform. Each of these : structures is designed with critical el**e**l**n**ents closely assoc**i**ated with flexible members and, therefore, most su**s**cep**t**ible to changes in mass and packag**i**ng characteristics due to g-loading. Three g-loading designs were determined for e**a**ch of these four structures and sizes.

The SOC and the &**'**\_P will not be studied in detail for g-loading effects becau**s**e of their relatively rigid s**t**ructure. Their cri**ti**cal elements are concentra**t**ed in a central and ri?id mass area with only the sol**a**r arrays Solar array support structures may be stiffended, if necessary, without significant impact on mass or auxiliary pr**o**pulsion requiremen**ts**. An additional reason not to lo**o**k a**t** varying g-loading characteristic**s** of the two **s**p**a**ce **st**a**t**ion de**s**ign**s** wa**s t**o avoid a prolifera**t**i**c**n of separa**t**e de**s**i**g**n**s**, size**s** and g-load**i**ng p**a**r**a**met**e**r**s**. Such a large number of di**s**cre**t**e case**s** would take away fr**o**m the maj**o**r thrust of this study.

To define the range of g-loading that each of the classes will experience in transfer from L\_3 to GEO or the other orbits, it is **ne**cessary to groundrule an OTV design. The **s**pe**c**ific characteristic of the OTV that is important to correlate thrust level with g-loading is the burnout or dry m**a**s**s**. This m**a**s**s c**omb**in**ed **with the LSS** m**ass w**a**s** u**se**d **to** der**iv**e **th**e maxi**m**um **acc**elera**t**ion **th**a**t th**e **L**\_S \_**a**m**t en**dure dur**i**ng **t**ra**nsf**er. "**h**rus**t** l**ev**el **wi**ll b**e** trea**t**ed a**s a** p**a**ram**ete**r a**n**d u**s**ed o**nly t**o g**au**ge th**e** requ**i**reme**nts plac**ed on the O**TV** eng**i**ne **a**s **a si**de **issu**e. **In a**dd**ition t**o O**T**V m**a**ss**,** th**e size** o**f** th**e CTV is** n**ec**e**ssary t**o d**efi**ne **f**or a**n insi**g**ht int**o pa**c**k**in**g r**equir**eme**nts**.

: **T**h**e** O**TV c:**\_**sen f**or th**is** stud**y** (**Fig**ur**e 24**) **is capa**bl**e** o**f** tran**sp**or**ti**ng **a 620**0 **i**b pa**yl**o**a**d to **G**\_**D a**nd r**etu**r**n t**o L**EO f**or r**e**u**se** or **a 164**00 **I**b **payl**oad **with** an **e**x**p**e**nda**b**le** O**TV**. **T**h**r**u**st l**eve**ls** bet**w**ee**n 15**0 i**b a**nd 1**5**000 lb **c**a**n** be **achiev**ed a**s sh**o**wn in Fi**gur**e 24** de**p**end**in**g upon **the** degree o**f** no**zzle e**x**ten**s**i**on **a**nd pump **s**peed. A**n e**x**p**en**d**ab**le** (E**V will** be u**s**e**d f**or **the** ncm**ir**\_**l** la**r**ge \_**ace**



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FIGURE 22. EXPERIMENTAL GEOSTATIONARY PLATFORM - 1st MODE, 0.096 HZ

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**PERFORMANCE** 

1. THRUST:

L MAINSTAGE = 15,000LB' LB.

- b. PUMPED IDLE 3,750 LB. (NOZZLE EXTENDED)
	- = 750-1200 LB. (NOZZLE RETRACTED)
- c. TANK HEAD IDLE = 150-500 LB. ONLY AT FIXED INLET PRESSURE

DESIGN POINT MIXTURE RATIO = 6.0  $\mathbf{z}$ a. OTV MAXIMUM PERFORMANCE MIXTURE RATIO = 5.75

- DESIGN POINT I. 482.5 SEC. 3. a. OFF-DESIĜN 1.ª SEE FIGURE 3.5.2-5.
- 4. PUMPED IDLE  $I_n = 452$  SEC.  $\bullet$  F = 3,750 LB. (NOZZLE EXTENDED)  $-436$  SEC.  $\bullet$  F = 3,750 LB. (NOZZLE RETRACTED)
	- 412-419 SEC. @ F = 750-1,200 LB. (NOZZLE RETRACTED)

TANK HEAD IDLE  $I_s = 393-405$  SEC.  $\bullet F = 150-500$  LB. (NOZZLE RETRACTED) 5.

Figure 24. OTV ENGINE CHARACTERISTICS

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systems in this study since their weights exceed 6000 lb. A typical sequential mass statement for a GEO transfer mission is shown in Figure 25. For a given thrust value, the highest steady stage g-loads will occur when the OTV is nearly empty of fuel.

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G = \frac{T}{W_{\text{OTV}} + W_{\text{PL}}}
$$

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where  $T =$  steady thrust  $W_{\text{OTV}} = OIV$  weight<br> $W_{\text{PL}} = \text{payload we}$ = payload weight

Using the high value of thrust in the tank head idle mode  $(T = 500 \text{ lb})$ , an OTV weight of 6560 lb and a dynamic mangification factor of 2.0 to account for thrust start-up transients, the minimum g-load which each of the four baseline space systems must be sized to withstand is shown in Table 9.

|          | Satellite Class                     | W(1D) | q-load |
|----------|-------------------------------------|-------|--------|
| ر<br>- ا | Large Aperture Phased Array         | 9800  | .061   |
| νĒ       | Wrap Rib Antenna w/Offset Feed      | 9695  | .062   |
| Ş        | Hoop-Column Antenna                 | 10340 | .059   |
| ٠.       | Experimental Geostationary Platform | 11722 | .055   |

Table 9: Minimum Design Load Factors

The g-loading with respect to thrust for each of the four flexible designs is shown in Figure 26. The critical m**e**mbers of each design and the type of : load encountered are shown in Figures 27 through **30**. Th**e** location and direction of the primary prop**u**lsion load (T) is identified also. We chose

> Expressions for the loads in each of the critical el**e**ments identified were • derived f**o**r each st**r**u**ct**u**r**e. The **l**oads were derived as **f**un**cti**on**s** of g-load and pretension parameters. A discussion of each structure follows:

# \_;' Phased **Array** An**t**enna

Using the Educational TV sa**t**ellite as an example, expressions for the loads in critical elements due to g-loading are derived. Orbit transfer thrust • is **a**pplied **a**t th**e** feed horn cluster module.

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**M**AIN E**NG**. IS**P = 4**6**4**.6 AUX. PROP. ISP =  $220$ 

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**'**" **F**IGU**R**E25, EXPE**N**DABLEMO**D**EGE**O** DELIVERYMISSI**O**N**(OFF**-L**O**ADED**F**OR 65 K STS) SEQUENTIAL MASS STATEMENT

 $\alpha$ 

16000 OTV THRUST LINIT MA INSTAGE 14000 12000-10000-THRUST, LB. 8000-6000-4000-PUMPED TOLE (NOZZLE EXTENDED)  $\hat{\mathbf{v}}$ 2000-PUMPED IDLE (NOZZLE RETRACTED) TANK HEAD IDLE 8 4  $\overline{.6}$  $\overline{8}$  $1.0$ ACCELERATION, g. \$ EXP. GEO. PLATFORM + LMSS (HOOP-COLUMN) & EDUCATIONAL TV

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· LMSS (WRAP RIB)



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CRITICAL MEMBERS FOR PRIMARY PROPULSION LOADS

- 1. EXTENDIBLE S/A BOOMS (IF S/A'S ARE DEPLOYED) [BENDING]
- 2. DEPLOYABLE MASTS [AXIAL LOADS (BUCKLING)]

3. AFT STAVES [TENSION]

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4. LENS COMPRESSION RING [COMPRESSION]

FIGURE 27. THE CRITICAL MEMBERS OF EDUCATIONAL TELEVISION SATELLITE

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CRITICAL MEMBERS FOR PRIMARY PROPULSION LOADS

- 1. DEPLOYABLE BOOM [BENDING]
- 2. REFLECTOR RIBS [BENDING]
- 3. SOLAR ARRAY SUPPORT BOOM (IF DEPLOYED) [BENDING]

FIGURE 28. THE CRITICAL MEMBERS OF LMSS WRAP RIB SPACECRAFT

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LMSS HOOP COLUMN SPACECRAFT



CRITICAL MEMBERS FOR PRIMARY PROPULSION LOADS

- 1. TELESCOPING COLUMN [COMPRESSION (BUCKLING)]
- 2. HOOP

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- [COMPRESSION]
- 3. CONICAL CABLE ARRAYS [TENSION]
- 4. SOLAR PANEL SUPPORT BOOMS [BENDING]

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FIGURE 29. THE CRITICAL MEMBERS OF LMSS - HOOP COLUMN

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# EXPERIMENTAL GEOSTATIONARY PLATFORM - OPTION 4A VIEW LOOKING TOWARDS EARTH



CRITICAL MEMBERS FOR PRIMARY PROPULSION LOADS

- 1. "DEPLOYABLE TRUSS" STRUCTURE [COMPRESSION AND LOCAL BENDING]
- 2. P/L SUPPORT STRUCTURE [BENDING]
- 3. S/A SUPPORT MASTS [BENDING]

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FIGURE 30. THE CRITICAL MEMBERS OF GEOSTATIONARY PLATFORM



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#### Antenna Mast

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The cri**t**ical load in the antenna mas**t** i**s** compressi**o**n and occurs a**t** the a**f**t i" end of the mast since the total mass of the mast is supported at the aft stave attachment. Assuming a tension,  $T_1$ , in each of the 12 staves, the Assuming a tension,  $T_1$ , in each of the 12 staves, the preload in the mas**t** is:

 $F_{pt} = 12(T_1)\cos\theta$ 

where 8 is the angle between a stave and the mast. T**h**e g-load contribution to the load results from the mass of the mast plus the mass of a portion of . the lens which is attached to the mast. Assuming 25% of the lens mass is supported by the column:

$$
F_g = (M_{\text{max}} + .25M_{\text{lens}}) 9.8 g
$$

The total compression load is the sum of these two contributions:

$$
F_{\text{maxt}} = 12(T_1) \cos \theta + M_{\text{maxt}} + .25M_{\text{lens}}) 9.8 \text{ g} \text{ (Newtons)}
$$

**w**ith

Mma**s**t **= 9**.**9** kg **:**: Mlens**=** 207.2 kg  $\theta = 61.1$  deg.  $F_{\text{master}} = 5.79 T_1 + 604.4 g$ 

# ; Antenna Stave**s**

• The highest stave tension occurs in the aft staves which must support the<br>
rass of the mast in addition to providing pretension forces. The stave mass of the mast in addition to providing pretension forces. tension is the mast load distributed to the 12 staves.

 $F_{\text{staves}} = F_{\text{mast}}/12 \cos \theta$ <br>= T<sub>1</sub> + 104.2 g

 $= T_1 + 104.2$  g (Newtons)

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## Antenna Rim

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Stave tension  $(T_1)$  and lens tension  $(T_2)$  applied at each of the 12 attachments plus the g-load in the lense membrane results in radial loads which are reacted by c**o**mpression forces in the antenna rim.

The pretension radial load is:

 $F_R$  = 2T sin $\theta$  + T<sub>2</sub>

The lens under a q-load acts like a circular membrane under a pressure load. The running load in **t**he perimeter of the membrane is:

$$
N = (.328) (Ep2 Rt)1/3 (N/M)
$$

where

 $E = Young's modulus$ 

- $R =$  radius of membrane
- $t =$  membrane thickness

: and, th**e** equivalen**t "**pre**s**su**r**e**"** (p)is a function of the g-load (g):

 $P = (M_{1em} / \pi R^2)$  9.8 g

; The resulting radial force at **e**ach atta**ch**ment is:

 $F_R = 2 R_N (.75)$ 

 $\sim$  The  $\sim$  actor results from the ract that the antenna fens is supported by the ma**s**t at the **c**ent**e**r and th**e** membrane ten**s**ion **c**\_9\_ation **a**pplie**s** to a **:** membran**e** \_rted on**ly** at it**s** edg**e**. Ther**e**fore, **ap**pro**x**imat**e**l**y** 7**5**% o**f** t**he** : The .75 factor results from the factor the mast at the center and the membrane supported only at its edge lens weight contributes to tension.

" T**h**e g-loa**d o**n th**e** st**a**ve**s** also contribute**s** a **s**mall am**o**unt to th**e** r**a**dial rim l**o**ad.

 $F_R$  = M<sub>stave</sub> sin $\theta$  g

Once th**e t**otal r**ad**ial load a**t e**ach **po**in**t is** calcula**t**ed, the **co**mpressi**o**n l**o**a**d** in the rim el**e**ments i**s**:

 $F_{\text{rim}} = (F_R/2)$ .  $(1./\sin\theta/12)$ 

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$$
T_1
$$
 = lens save pretension (N)  
g = g-load (g's)

Ant**enna Staves Tension**

 $F_{\text{stawe}} = T_1 + 104.2g$  (Newtons)

Antenna Rim (Compression)

 $F_{\text{rim}}$  = 3.38 T<sub>1</sub> + 1.93 T<sub>2</sub> + 993.g<sup>2/3</sup> + 13.6g (Newtons)

where:

 $T_2$  = lens pretension at each attachment

**; Lens Support Struts (Co**mpression)

 $F_{\text{struct}} = 2.0 T_3 + 1204. g$ 

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 $T_3$  = stabilizing line pretension

Solar A**r**ray Bo**om (Bending)**

 $BM_{S/A} = 31420$ . g  $(N-m)$ 

• The effect of increased g-load**i**ng on the mass of the **E**ducational TV and Electronic Mail **w**ere determined. Although preliminary m**a**ss statements for **this sa**tellite **a**re **a**v**a**i**l**a**b**le, man**y as**smn**pti**on**s w**ere req**ui**red to est**a**bl**is**h **st**ruc**tural** parame**te**r**s (**s**iz**e, m**ate**r**ial**, **p**r**et**e**nsi**o**n l**oad**s, e**t**c**.**)** ne**ce**s**s**ar**y**  $\frac{1}{2}$  for these calculations.

**Th**e mo**s**t **s**ignifica**n**t fa**c**tor in determining the effe**c**t of g-loading on s**p**ace**c**raft weight is the criteri**a** used to design it. If the **s**t**r**u**c**ture is a •" strengt**h** de**s**ign, i.e., de**s**igned to with**s**tand **a** given set o**f** load**s**, then an**y** ; in**c**r**e**ase in loading **w**ill re**s**ult ir re**s**izing t**he s**\_**r**u**c**ture with an a**p**propriate increa**s**e in weight. Most **sp**acecra**f**t de**s**ig**ns**, however, would be **"**: too **f**lexible i**f** de**s**igned enti**r**e**ly** u**s**ing **s**trength a**s a c**riter**i**on. **T**he  $r$  resulting resonant frequencies would fall in the control system bandwidth and cau**s**e sign**i**fic**a**nt **c**ont**r**ol and **al**igraent pro**b**l**e**m**s**. The**r**efor**e**, sti**ff**nes**s - cr**iteri**a** a**re** o**f**ten used to redu**c**e stru**c**tu**r**e/cont**r**ol int**e**r**a**ction and improv**e** satellite **p**er**f**orman**c**e. A stiffnes**s** designed **sp**acecraft, t**h**ere**f**ore, ha**s** mor**e** st**r**engt**h cap**abi**l**it**y** t**h**an i**s r**equ**ir**ed **f**o**r** the expected **l**oads and **c**ou**l**d **wi**thstand higher **g-**load**s w**ith no in**cr**ea**s**e in **w**eig**h**t. **T**he magnitu**d**e o**f** .. **g**-load whi**ch** would cause a r**e**si**z**ing o**f a p**ortion of **th**e st**r**uc**t**u**r**e i**s** m\_**n**\_**r**, u**ntil a sp**ec**ific** sa**t**e**ll**i**te i**s des**i**gned.

The assumption of a strength designed structure, therefore, provides an upper bound on the weight increase caused by increased g-loads. The second assumption **n**ade is that increased strength capability is accomplished by increasing member cross sectional aleas instead of using righer strength materials. For example, for the triangular expandable trusses, the areas For example, for the triangular expandable trusses, the areas of the three longitudinal elements are increased while the interconnecting elements remain constant. Increased bending capability of the solar array booms was also accomplished by increasing longitudinal element areas and not by changing the triangular dimensions. The third assumption is that the increase in structural weight of the load carrying elements of a member (axial el**em**ents in the previous example) is proportional to the increases \_n the member load. The fourth assumption is that the structure is designed for a g-load of 0.06 g. For the triangular expandable beams, it was assumed that half of the beam weight is in the longitudinal load carrying el**em**ents.

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Figures 31 and 32 show the effect of g-load on weight for the Phased Array Designs Satellite and identifies the incremental weight for the structural elements contributing to the increase. The main spacecraft body and an**t**enna feed c**o**mpr**i**se app**ro**xima**t**ely 70 pe::**c**en**t of** the **to**ta] weigh**t** and are assumed to be unaffected by increased g-loads.

## LMS**S (Hoop and Col**\_n)

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Expressions for element loads are based on the assumption that tension in the forward and aft hoop cables is defined by tension  $(T<sub>C</sub>)$  in the 48 aft : (**-**Z) **c**ables and that tensi**o**n in the r\_flector **s**urfa**c**e (**Ts**) is **app**lied at each of the hoo**p** attac**h**m**e**nt locations. Orbit t**r**ansfer thrust is a**p**plied in the -Z di**r**ection at the +Z Bus.

**:** Hoop (**c**om**p**res**s**i**o**n): Hoo**p** loads are the re**s**ults **o**f preten**s**ion and g**-**loads

 $F_{\text{homo}} = 11.41 \text{ T}_\text{C} + 10.64 \text{ T}_\text{S} + 1063.9 \text{ (Newtons)}$ 

Cabl**es** (Tension) : **T**ension in each of th**e 43** cabl**es** in ea**c**h o**f** th**e** five **c**oni**ca**l **c**abl**e** a**r**r**ays** (ring **i** i**s c**lose**s**t to th**e** column) ar**e** as **f**ol\_ows:

 $T_1$  (aft) **=** .32 $T_s$  + 1.07g  $T_1$  (fwd) **=** .46 $T_s$  $T_2$  =  $.08T_a + 4.97g$  $T_3$  = .15 $T_5$  + 15.43g  $T_A$  = .35 $T_S$  + 49.09g  $T_5$  (aft) =  $T_c + 91.88g$  $T_5$  (fwd) =  $.58T_5 + .78T_c$ 

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EFFECT OF G-LOAD ON MASS - ELECTRONIC MAIL SATE! ITE FIGURE 31.

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Solar Array Supports (Bending): For the +Z solar arrays, the critical  $\ddot{\phantom{a}}$ loads are at the inboard end of the support.

$$
BM_{S/A} = 5080 \text{ g} (N-M)
$$

 $\omega \rightarrow \infty$ 

Feed Assembly Supports (Bending): For each of the four feed arrays, the bending moment at the +Z Bus attachment is:

 $BM_{feed} = 7290.9$  (N-M)

The effect of g-loading on the mass of the LMSS (hoop and column) spacecraft was determined using the same assumptions as the Educational TV Satellite described in the previous section (i.e., a strength-designed spacecraft, structural capability increased through increases in  $\frac{1}{2}$  space-contional area which are reportional to the increase in perhor lead.  $\frac{1}{2}$  cross-sectional area which are proportional to the increase in member  $\frac{1}{2}$ etc.). Figures 33 and 34 show the effect of g-load on mass for the **L**MSS (Hoop and Column) spacecrafts and identifies the incremental mass for the structural elements contributing to the increase. The mass of both the  $+2$ and -Z Bus are assumed t**o** be unaffected by in**cr**eased g-load. Also the increase in solar**"**array and feed support mass contribute an insignificant amount to t**h**e total spacecraf**t** mass.

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EFFECT OF G-LOAD ON MASS - LMSS SPACECRAFT (HOOP AND COLUMN) FIGURE 33.  $\boldsymbol{\hat{}})$ 

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#### LMSS (Wrap Rib)

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Critical element loads are based on preloads in the antenna ribs and  $\overline{I}$ g-loads. Orbit transfer thrust is assumed to be applied to the bus.

Ribs (Bending): Pretension in the reflector ribs is a function of the bending stiffness (EI) and the deflected shape of each rib. Sufficient structural detail is not available at this point to determine the preload bending moment (BM<sub> $\alpha$ </sub>) at the root of each rib.

 $BM_{\text{rib}} = BM_{\text{O}} + 748 \text{·g}$  (N-M)

Boom (Bending & Compression) : With the orbit transfer thrust applied to the bus, the critical load in the expandable boom occurs just outside of the boom housing. The load is the result of the combination of compression in the long boom and bending which results form the offset reflector and short boom.

$$
BM_{\text{bosm}} = 1.17 \times 10^5 \text{g} \quad (N-M)
$$

$$
F_{\text{bosm}} = 5020 \text{ g} \quad (N)
$$

UHF Feed Support (Bending): Bending in the UHF Feed support structure is caused by g-loading.

$$
BM_{\text{feed}} = 4.0 \times 10^4 \, \text{g} \quad (N-M)
$$

Solar Array Support Boom (Bending): The bending moment at the base of the solar array boom is:

$$
BM_{S/A} = 2.15 \times 10^4 \, \text{g} \, \text{(N-M)}
$$

Figures 35 and 36 show the effects on mass of varying g-load for the two \_**;** wrap rib designs.

#### Geostationary Platform

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The only elements of consequence for the geoplatform were the solar arrays and wrap rib antenna. These elements have been described in the previous sections. Figure 37 shows the simple relations for these elements.

# Mass Property Summary<br>Table 10 summarizes

Table I0 summarizes the mass properties, g-loading effects, and nodal  $\approx$  frequencies for each ISS.

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EFFECT OF G-LOAD ON MASS - LMSS SPACECRAFT (WRAP RIB) FIGURE 35.  $\mathbf{r}$ 

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EFFECT OF G-LOAD ON MASS - BASELINE EXPERIMENTAL GEOSTATIONARY PLATFORM

FIGURE 37.

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计连续机 计对称处理

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医无线性 医心包 医生物 医白细胞 医前列腺 医心包 医中心性 计数据库 医中心反应 电子电子 医血管 医马克氏反应

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44、44、1021年初起, 44、1444年9月,18日的"迷茫"设备, 170.8、当44年(18日) \$P\$10、170.4国家当镇、参观、休息管理等, 190.7月

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TABLE 10. CHARACTERIZATION SUMMARY

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 $\begin{array}{c} \mathbf{1} \\ \mathbf{1} \\ \mathbf{1} \\ \mathbf{1} \end{array}$ 

 $\bar{1}$ 

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 $\begin{array}{c} 1 \\ 1 \\ 2 \\ 3 \\ 4 \end{array}$ 

 $\sum_{i=1}^{n}$ 

#### 2.0 ESTABLISHMENT OF APS REQUIREMENTS

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The objective of Task 2 was to establish the propulsion requirements due to environmental, operational and structural constraints. To accomplish this objective, four mjor subtasks were executed. The first subtask involved a disturbance environment-characterization and a stationkeeping and torquing<br>requirements determination at LEO and GEO. To fully define the requirements determination at LEO and GEO. thrust thruster and minimum bit requirements, the second subtasks established a set of criteria to place thrusters on each of the selected vehicles and used this criteria to locate and size the thrusters. With the preliminary thruster requirements and locations determined, the third subtask characterized APS mass and converged thrust/thruster requirements.<br>Both chemical and ion systems were used for APS scaling. In the final Both chemical and ion systems were used for APS scaling. subtask, the interactions between the auxiliary propulsion system and the structure were investigated. This critical subtask gave the first This critical subtask gave the indication that large space structure/propulsion interaction may be a driving issue in uhe design of the structure and integrated propulsion system. The key assumptions used in each subtask are shown below in Table ii.

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In subtask 2.1, conservation as**s**umptions were used for the atmospheric m**o**del, d**r**a**g c**oefficient and reflec**t**ivity. The atmospheric model was varied to understand effects of the model assumption on orbit decay, however, the N**A**gA neutral m**o**del was employed to derive propulsion requirements trends. Thi**s** mode**l** is a l**o**ng te**r**m worst case model and thus yields **c**onservative, but realistic answers. The drag coefficient wa**s** set at 2.5 which may be a l**i**tt**l**e **o**n the high **s**ide. Other st**u**d**i**e**s** ha**v**e u**s**ed drag coe**ff**i**ci**ent**s** at 2.0

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68 **D1**80-**2772**8-**2** to 2.7 which correspond to difference momentum exchange surface properties. Reflectivity was set at a relatively high value which made solar pressure torques somewhat exaggerated for certain lower reflective surfaces. Solar pressure did not add significantly to the torque requirements and only in a few cases were stationkeeping requirements effected, therefore small changes in this assumption would not affect the study conclusions.

The assumptions used to locate thrusters are decribed in more detail in Section 2.2. The two driving assumptions were maximum moment arms and 0 torque \_V requirement. Mom**e**nt arm was not as important as was thought beforehand, and, in fact actually drove minimum impulse bit requirements for pointing control lower than the state-of-the-art. Pure aV with no torque created a throttling requirement for thrusters which had to be located non-symmetrically around the CG. Subtask 2.3 generated scaling equations for monopropellant, biopropellant and ion (Hg, SEPS technology) thrusters. These scaling equations were based on existing hardware and, in the case of ion thrusters, accepted theory.

An assumption was made in subtask 2.4 that a 10% pr**o**ven loss in the antenna beam wa**s** the maximum structural information allowed. This assumption was • based on dis**c**ussions with the Boeing Space Antenna Systems group. For most i missions, a defocusing of the beam beyond 7-10% would be considered<br> $\therefore$  unacceptable. - unacc**e**ptable.

The driving key issues of Task 2 were threefold. First, the aerodynamic disturbances in LEO were so large for most of the antenna structures examined that operation even for short periods of time proved difficult. Examined that operation even for short periods of the proved difficult.<br>LEO deployment and checkout of the large antenna systems imposed much<br>higher thrust and large AV requirements versus the GEO operational higher thrust and large  $\Delta V$  requirements versus the GEO operational requirements, second, the stationkeeping strategy at GEO drove the requirements, second, the stationkeeping strategy at GBO drove state-of-the-art limitations depending on the duty cycle of thrusting. For short duty cycles  $(1 \text{ hour/orbit})$ , chemical system thrust levels were short duty cycles (<1 hour/orbit), chemical system thrust levels were .\_ required. The solar array duty cycles had to become very long (>3  $\mathcal{L}$  in the final key issue was the systems were viable. The final key issue was the allowed A**P**S mass fraction at G**E**O. If the cost of transportation is more is equired. The solar array duty cycles had to become very long ()3<br>hour/orbit) before ion systems were viable. The final key issue was the<br>allowed APS mass fraction at GDO. If the cost of transportation is more<br>effected b missions, then low Isp systems are a**c**ceptable. If the mass becomes \_ critical because of STS or **t**ransfer stage capability, then higher Isp'S are more desirable. It was shown that the delivery systems were challenged for \_ very low Imp (2**00**'s) examined, however, delivery c**o**sts w**e**re bey**o**n**d** the scope of the contract, and a firm notion of where  $I_{SD}$  should lie will be effected by such an analysis.

#### \_J;\_ 2.1 **Di**sturban**c**e \_lvironmen**t** Anal**ysis**

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This subtask consisted of **t**hree separate analysis which provided a detailed **.** e**x**amina**ti**on **of** d**ist**urban**c**e **t**orque**s**, LEO stat**io**nkee**p**ing f**o**rce**s** and e**ff**ects, and GBO sta**t**ionkeeping forces and e**f**fect**s**. LE**L**'di**st**urbances wire domina**t**ed : by aerodynami**c** influen**ce**s f**o**r both f**o**rces **a**nd L**;o**rques. GID disturbances we**r**e domina**t**ed by g**r**a**v**ita**ti**onal in**f**luen**c**e**s**. G\_d**v**i**ty** gra**d**ien**t to**rqu**es** an**d** solar/imla**r** g**r**avita**t**ional a**tt**r**a**ctior.s imposed th**e** mo**st s**ignifican**t •** pr**o**pulsion requirements f**o**r system**s** s**i**zed f**o**z GID operation. Solar **pr**e**ss**ure **f**or**c**e**s** an**d** torqu**es** wer**e of** lower magn**i**t**ude** a**t** C\_O **t**han gravitational influences and orders of magnitude lower than aerodynamics in LE**O**.

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#### 2.1.1 Disturbance Torque Analysis

Table 12 shows a summary of the disturbance torques relevant for this the theory of the team of the team of the t study. The first three disturbances were treated in detail in this study. Magnetic torques were shown in a previous analysis (Reference 2) to be of only minor importance for the composite antenna structures examined. A discussion of aerodynamic, gravity gradient and solar pressure follows:

#### Aerodynamic Disturbance

Aerodynamic force on the configuration was determined by the equation:

$$
F_{AERO} = C_D A \frac{V^2 \rho}{2}
$$

where

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 $C_D$  = Drag Coefficient

- A **=** Cross Sectional
- p **=** Atmospheric Density

 $\rho =$  Atmospheric Den<br>  $V =$  Orbit Velocity<br>  $\therefore$ <br>
Atmospheric density is the Atmospheric density is the most difficult parameter to accurately estimate.<br>Density is affected by two factors which relate to solar activity - the  $\mu$  behavior is affected by two factors which relate to solar activity - the solar activity - \_ geomagnetic index and the solar flux. Both of thes**e** factors vary with time due to changes in the solar activity cycle. A**c**tual measurements for 1979 through mid-1980 are shown in Figure 38. The solar cycle peaked in 1978 and will again reach a maximum around 1980.

: In these s**+.**udies, f**o**ur atmospheric models were considered in deriving the orbit decay data. See Figure 39. The nominal m**o**del is the U.S. Standard " Atmosphere, 1**9**76. The **o**the;: three m**o**d**e**l'\_were gen**e**rated v**i**a th**e** quick-lo**o**k density model in Appendix B for a latitude of 0. The NASA Neutral and Short Time Maximum Models use values suggested for space shuttle studies. The NASA Neutral Model is a high solar activity model, with a value of 230 for the mean 10.7 cm solar flux and a geomagnetic index (A) of 20.3. The Short Time Maximum Model uses a 10.7 cm solar flux of 230 and a geomagnetic These conditions would occur only for a time of 12 to 36 hours during an extremely large magnetic storm. The Minimum Model uses figures of 73.3 for the 10.7 cm \_olar flux and 10.9 for the ge**o**magnetic \_. Znd**ex**. The solar flux and ge**o**magn**e**tic inde**x** f**i**gure**s** ar**e** th**e 9**7.7 percentile figures for June 19**8**7 form the Marshall Space Fligh**t** Cen**t**er predictions.

> The **"**NASA neutr**a**l" **is** considered to be the wor**s**t long**-**term or **c**ont**i**nuou**s** case applicable to any reasonable resupply cycle or propellant loading analy**s**is.

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TABLE 12. DISTURBANCE TORQUE SUMMARY

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FIGURE 38. GEOMAGNETIC INDEX AND SOLAR FLUX MEASUREMENTS



FIGURE 39. ATMOSPHERE DENSITY MODELS

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#### Gravity Gradient

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Gravity follows an inverse square law, and, as a consequence, mass elements near the earth are attracted more strongly than those farther away. When these forces are summed over a body, the point of application; i.e., the center of gravity, is found to be a little closer to the earth than the center of mass. The separation is very \_nall with respect to a characteristic dimension of a structure but does lead to a gravity torque. General expressions can be found in References 5, 6, and 7. If the body axes are chosen as principal axes of inertia, the body axis torque expressions simplify to

 $-x^2 - 2p^3$   $x^2 - xy'$  cos v sinze  $T_y =$  "  $(I_{zz} - I_{xx})$  cos $\phi$ sin2 $\theta$  $T_z =$  "  $(I_{XX} - I_{yy})$  sin $\phi$ sin2 $\theta$ 

 $\frac{1}{2}$  In these expressions,  $\varphi$ ,  $\sigma_1$  and  $\omega$  are the roll, pitch and yaw Euler angles,  $\mu = \alpha n_{\text{earth}} = 3.500 \times 10^{5} \text{ Nm}$ /sec 2, and  $p = \text{radius of orbit.}$ 

#### Radiation Disturbances

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In the analysis of radiation disturbances for earth orbital missions, three sources of radiation require consideration. The primary disturbance is sources or radiation require consideration. Ine primary unsturate is the promotions and a plasma force from the solar wind. A secondary disturbance is earth illumination which can be ref-ceted sunlight or infrared emission from photons and a plasma force from the solar wind. A secondary . disturbance is earth illumination which can be ref:ected sunlight or infrared emission. Finally, small effects can resule from spacecraft asymmetrical radiation emission in the form of thermal hot spots or radio transmissions. This latter disturbance is many orders of magnitude lower than the other two and was not examined.

There are also three factors to be considered in the determination of forces from any radiation source. The quality of incident radiation by the intensity, spectrum, and direction is the first determinant. **E** and, is the geometry of the spacecraft including the shape of the surface and the location of the sun with respect to the spacecraft mass center. Finally, the optical properties of the surface upon which the radiation is incident or from which it is emitted must be considered. Table 13 summarizes the radiation sources and force determination factors to be used during this study.

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The two sources of direct solar radiation, photon pressure and the solar wind plasma force, are separated by four orders of magnitude. The solar wind is so much weaker than the photon radiation forces that its effect can safely be ignored.

Table 13. Radiation Disturbance Factors

B. Solar Wind B. Spectrum

Sources of Radiation **Force Determination Factors** 

I Direct Solar Radiation I Incident Radiation Properties

A. Photons A. Intensity

C. Direction

II Spacecraft Geometry II Spacecraft Geometry

A. Reflected Sunlight A. Surface Shape

B. Infrared Emission B. Location of Sun

- A. Thermal Hot Spots **A. Reflection**
- B. Radio or Power B. Emission Transmission
- 
- III Space Emission III Surface optical Properties
	-
	-
	- C. Absorbtion

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The sun provides essentially collimated radiation with a reasonably well defined intensity and spectrum. The solar photon radiation may be characterized by the solar constant  $I_0$  which is the rate of which energy at all wavelengths is received per unit area. The best estimate of this value is 1353  $\pm$  20 W/m<sup>2</sup> which when converted to force yields 4.513 x 10<sup>-6</sup> N/m<sup>2</sup>.  $\frac{1}{2}$  2000  $\frac{1}{2}$  2000  $\frac{1}{2}$  which when converted to force yields  $\frac{1}{2}$   $\frac{$ Because this constant \_as uits of force per unit area, it is often called a pressure. This terminology can be misleading as the pressure here is in reality a vector quantity not a scalar.

The solar constant follows an inverse square law which is important for interplan**u**tary flight, however f**or** ea**r**th **o**rbit mi**ss**i**o**ns the **on**ly **so**urce of distance variation is the eccentricity of the earth's orbit. The variation due to eccentricity changes the value by 3.5 percent and can, for the **p**urposes of this investigation, be ignored. Solar radiat**i**on, therefore, is : taken to be a constant of 1353 W/m 2 from a collimated **E**ource.

#### Earth Illumination (Albedc)

In addition to the direct solar radiation falling on a spacecraft, reflected radiation from the earth also exerts a pressure. The effect is a maximum at noon and tends to partially cancel the direct radiation forces. The earth and its atmosphere act as a diffuse reflector with the result that the albedo radiation is not collimated. This considerably complicates the determination of the resulting forces. Often these forces are ignored on the grounds that their omission will lead to conservative estimates of the total direct and reflected radiation effects. While this approach is often justified, large vehicles in relatively low orbits can experience significant relief from the albedo radiation and it may be important to include the effect.

Assuming the earth to be a perfectly diffuse reflector obeying Lambert's cosine law, the radiation emitted from an element of area  $dA_1$  is kUcos $\theta dA_1$  $N/m^2$  per unit solid angle in a direction inclined  $\theta$  to the surface normal. k can be identified as the albedo coefficient and U is the incident radiaticn. The radiation pressure at a distance r is

$$
V = \frac{kU\cos\theta \ dA_1}{\pi r^2}
$$

direct along r.

The normal and tangential components of radiation on an area dA, as snown in Figure 40 can be expresssed

$$
V_N = \frac{U}{\pi} \int \frac{k \cos \theta \cos \alpha}{r^2} dA_1
$$

$$
= \frac{-U}{\pi} \int \frac{k \cos \theta \vec{r} \cdot \hat{n}}{r^3} dA_1
$$

and

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$$
V_T = \frac{U}{\pi} \int \frac{k \cos \theta \sin \alpha}{r^2} dA_1
$$
  

$$
= \frac{U}{\pi} \int \frac{k \cos \theta \hat{n} x (\vec{r} \times \hat{n})}{r^3} dA_1
$$

where n is a unit vector from surface element  $dA_2$ .

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### FIGURE 40. RADIATION GEOMETRY

#### Earth Radiation

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The other source of disturbance from the earth and its atmosphere is a diffuse radiation with a spectrum approximated by the spectrum of a 288<sup>0</sup>K<br>black body. This temperature varies with the transparency of the This temperature varies with the transparency of the atmosphere from 218°K to 288°K with about 95 percent of the enitted radiation originating from the earth or the lower atmosphere. The radiation is not collimated and may be treated in the same way as the earth reflectance problem with the following result

$$
I_e = \frac{I_e^{\text{o}}}{\pi} \int_{Esss} \cos \psi \, ds/d^2
$$

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- $I_{\alpha}^{0}$  = global average emmission constant (243 W/m<sup>2</sup>)
- dS = element of differential area on the surface of the earth

 $d =$  distance from satellite to  $dS$ 

- $\psi$  = angle between the normal dS and d
- $Esss = earth surface as seen by satellite$

Figure 41 shows the relative values of solar radiation, earth reflectance If and earth radiation for a spherical satellite for a range or orbit radii.

#### Disturbance Torque Calculations

 $\dot{\zeta}$  . Before summarizing the complete disturbance torque analysis for each vehicle, an example of the calculations is shown below for the 12.5 kw SASP vehicle, an example of the calculations is shown below for the 12.5 kw SASP • configuration at 300 km. LSO altitude treated in this study ranged from 300 to 500 km. it is noted at the outset that torques at 300 km were an order of magnitude higher than 500 km due to increased atmospheric density. Torques at GH0 are 3 or 4 orders of magnitude lower than those at 300 km. This sensitivity proved to make radical differences in the thrust/thruster requirements for torque cancelation at LEO versus GEO. The same conclusion  $\zeta$  requirements for torque cancelation at LEO versus GEO. The same conclusion will be shown when stationkeeping is considered in the following sections.

while is included the smaller SASP (12.5 kw) configuration at 300 km<br>altitude is examined. Several assumptions were made to model the Several assumptions were made to model the<br>..., and gravity gradient disturbances. Some of aerodynamic, solar radiation, and gravity gradient disturbances. these assumptions came from Reference 8. Assumptions are as follows:



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- i. The positions of the forces on the satellite are described using a spherical coordinate system  $(\phi, \theta)$ .
- 2. The wind is considered to be a unit vector in the direction of the translational velocity, and the solar radiation is a unit vector exactly opposite to the wind. (This was varied for the analyses that follows. )
- 3. Solar radiation pressure =  $4.70 \times 10^{-6} \text{ N/m}^2$
- 4. Coefficient of drag  $(C<sub>D</sub>)$  = 2.5

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- 5. Gas density (Rho) =  $4.0 \times 10^{-11}$  kg/m<sup>3</sup> (NASA Neutral Model) for 300 km altitude
- 6. Solar radiation is either absorbed, reflected specularly, reflected diffusely or some combination. A combination of all three was used<br>with the absorption coefficient and the coefficient of diff  $\infty$ with the absorption coefficient and the coefficient of reflection, each having different values for the front and back ; surfaces of the satellite.



Ca + C<sub>s</sub> + C<sub>d</sub> = 1.0<br>
For future analysis, we assumed 0 absorption, 3 diffuse, and .7 specular.<br>
The spherical coordinate system is shown in Figures 42 through 45.<br>
Aerodynamic drag and solar radiation calculations were The spherical uoordinate system is shown in Figures 42 through 45. Aerodynamic drag and solar radiation calculations were made for  $\phi$  values from 0 to 180 degrees in 15 degree increments and  $\theta$  values from 30 to 360 degrees in 30 degree increments. To calculate a sum of the forces, it was exaptitrarily decided to add together the forces occurring when the wind is<br>exactly opposite to the solar radiation, Figures 6 and 8. Later exactly opposite to the solar radiation, Figures 6 and 8. calculations considered a nominal position of the satellite in orbit. In this position, a sum of forces were made using the wind vector as a constant and varying the radiation vector, Figure 45. The spherical coordinate system is shown overlaid onto SASP in Figure  $46$ . This also shows the direction of the maximum force on SASP due to aerodynamic and radiation forces at 300 km altitude.

The following figures show the variation of torque with  $\phi$  and  $\theta$  at 300 km altitude. Figures 47, 48, and 49 illustrate the aerodynamic, gravity Figures  $47$ ,  $48$ , and  $49$  illustrate the aerodynamic, gravity \_' grad**i**ent and radiati**o**n pre**ssu**re t**o**rques at low altitude. Figure 50 sums these torques to show the RSS totals for the worst case axis. It is seen that aerodynamics d**o**minate the composite total. Appendix B shows a similar analysis at 500 km. At this altitude, aerodynamics is still dominant. **,** However, gravity gradient is now 10% of the total. Radiation torques have \_ decreased due to the increased radius and consequently decreasing earth reflection and albedo contribution.

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CRIGHTAL PAGE IS OF POOR QUALITY  $+Z$  $\phi = 40^{\circ}$ DIRECTION OF WIND FORCE WIND UNIT VECTOR  $+Y$  $\theta = 55^{\circ}$  $+X$ 

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FIGURE 43. REFERENCE AXES FOR SOLAR RADIATION CALCULATIONS

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WIND IS EXACTLY OPPOSITE TO SOLAR RADIATION --FIGURE 44. ORIENTATION USED FOR SUMMATION OF DISTURBANCE CALCULATIONS (PHI AND THETA REFER TO SOLAR RADIATION VECTOR)



WIND IS HELD CONSTANT & SOLAR RADIATION IS VARIED --FIGURE 45. POSSIBLE ORIENTATION FOR SUMMATION OF DISTURBANCE CALCULATIONS (PHI & THETA REFER TO SOLAR RADIATION VECTOR)

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FIGURE 46. SPHERICAL COORDINATE REFERENCE ANGLES FOR MAXIMUM FORCE ON SASP (12.5 KW) 300 KM ALTITUDE



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# LEGENO<br>0 THETA = 270 DEG<br>1 THETA = 270 DEG<br>2 THETA = 360 DEG

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## FIGURE 47. TORQUE UPON SASP (12.5 KW) DUE TO AERODYNAMIC DRAG, 300 KM ALTITUDE



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LEGEND<br>0 THETA = 30 DEG<br>1 THETA = 90 DEG<br>2 THETA = 210 DEG

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FIGURE 48. TORQUE UPON SASP (12.5 KW) DUE TO GRAVITY GRADIENT, 300 KM ALTITUDE

**COMMENTATION** 

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LEGENO<br>O THETA = 0 DEG<br>1 THETA = 60 DEG<br>2 THETA = 90 DEG

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FIGURE 49. TORQUE UPON SASP (12.5 KW) DUE TO RADIATION PRESSURE, 300 KM ALTITUDE



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**LEG**END <sup>0</sup>**<sup>1</sup>** <sup>T</sup>THETA **HE**TA:**<sup>=</sup> <sup>90</sup>**<sup>0</sup> DEG <sup>D</sup>**EG** <sup>2</sup> **THE**TA**<sup>=</sup>** <sup>18</sup>**0**D**EG**

FIGURE 50. TORQUE UPON SASP (12.5 KW) DUE TO SUM OF **A**ERODYNAMIC AND RADIATION AND GRAVITY GRADIENT TORQUE, 300 KM ALTITUDE

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After this initial calculation was made, a change in the spherical coordinate system was made to Cartesian coordinates and the more familiar roll (x-axis) or nominal velocity vector), pitch (y-axis) and yaw (z-axis or earth nadir). Nominal or operational attitudes were also compared with worst case or entations. The sum angle was varied within appropriate orbit constraints to find the worst case radiation oontribution in both cases. These orientations are summarized in Table J1.

Table 14. Nominal and Worst Case Orientations

Nominal Orientation

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- 300, 400, 500 km and GEO
- Varied roll, pitch, yaw I0 in all axes minimum bit requirements
- Used worst case sun angle
- Selected worst case attitude and torque for this 10 range

• , Worst Case Orientation

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- .\_'- **3**00, 400, **5**00 km and GSO
	- Varied roll, pitch, yaw 360 in all axes
	- Used wors**t** case sun angle
	- -**" S**elected w**o**rst case attitude, torque

A summary of the disturbance torques for each vehile size is shown in '- Tables 15-18. The mass properties for this set of tables were taken for a g-load of .15 g's. **A**dditional information for the other g-loads of .06 and 1.0 can be found between the g-loads. This resulted from a nonlinear and nonsymmetric addition of mass to structural m\_rs for increasing g-loads. As mass was added, inertia properties and CP-CG momentum arms were changed in a non**co**rrelating and vehi**c**le s**p**ec**i**f**ic** manne**r**.

Table 15 shows the LEO nominal orientation torques for each vehicle. Table<br>16 shows the same altitudes but in a worst case condition. The variation<br>16 of torques by altitude is shown to be around an order of magnitude from 16 shows the same altitudes but in a worst case condition. of torques by altitude is shown to be around an order of magnitude from 300 to 500 km. Variations of torques by class also vary by orders of magn**i**tude**s** due to **t**he c\_**,**\_nging **a**e**r**odynam**ic** and m**ass p**roperty va**r**iat**io**n**s**. The wrap rib and hoop/column designs have very high area/mass ratios and<br>large CP-OG offsets and are dominated by aero torques. An interesting large CP-CG offsets and are dominated by aero torques. result for the SOC designs was that the initial version had higher torques than the **op**er**a**tional **ve**rsion. The cr**oss p**roduct**s of** in**er**tia an**d** th**e** CP-OG **off**set for th**e** smaller ve**rs**ion **w**e**r**e greater than the more symmet**ri**c larger SOC design. This indicates that configuration optimization from a torque **Chack only in 1973**<br>OF POOR QUALITY

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\frac{1}{2}\n\end{array}$ 

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TABLE 15. DISTURBANCE TORQUES AT LEO FOR NOMINAL ORIENTATIONS

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 $\vdots$ 

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\cdot \\
\cdot\n\end{array}$ 

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 $\lambda$ .3300 .7070  $1.068$ <br> $-2.115$  $9.642$ <br> $9.371$ 7.432 9.862 RSS  $.2067$  $7.027 - 2.268$ .3342  $-.2380$  1.068  $-.6700$ <br> $-.9880$   $-2.115$   $-1.270$ 9.632<br>7.396  $9.559 - 9.562 - .0314$ TZ<sub></sub> 500 km  $.2637$ <br>-.5600 .6919  $-2.587$ <br>3.430  $\mathsf{r}$  $-6.931$  $-.3300$  $-5.011 - 8.622$ .4403  $\mathbf{r}$ .7510<br>1.439 1.825  $3.448$ <br> $6.482$ 30.24 14.57 23.24 RSS  $.2160$  $13.45 - 4.300$  $.7870$  3.448 -2.058<br>-3.273 -6.482 -3.710  $\frac{30.18}{7.984}$ .0802  $.809$ TZ 400 km  $-.7510-.5073$ <br> $.6046 - 1.439$ 22.20 1.751  $-6.276$ <br> $3.712$ M  $-13.24$ 22.15 1.242  $-15.36$ <br>14.95  $\mathbf{r}$  $2.612$ <br> $5.540$ 49.18 14.10<br>26.00 82.56 6.844  $122.4$ <br> $53.22$ **RSS**  $42.52 - 13.70$ .0178<br>.2258 .2995  $\frac{8.270}{14.60}$  $\frac{122.1}{9.603}$ 2.927  $\mathbf{L}$ 300 km  $2.612$   $2.033$ <br> $2.378 - 5.540$  $77.96 - 75.20$ 6.454  $3.240$  14.20<br>-13.49 -26.00  $-22.63$ <br> $37.80$  $\mathsf{r}$  $-42.82$ 4.830  $-61.57$ <br>48.18  $\mathbf{r}$ 50 x 33 m<br>9 PAYLOADS 12.5 kw INITIAL<br>OPER. 13 kw 120 m SIZE  $\epsilon$ 55 HOOP/ COL. WRAP RIB GEO PLT. CLASS LAPAA SASP  $50C$ 

OF POOR QUALITY

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DISTURBANCE TORQUES AT LEO FOR WORST CASE ORIENTATIONS TABLE 16.

WORST CASE DISTURBANCE TORQUES (N-M)

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poin**t o**f view c**o**uld significan**t**ly reduce momentum management and desaturation requirements.

Table 17 shows the GBO disturbance levels for each axis and the RSS **t**otal. These torques were used to size thrus**t**/**t**hruster requirements (worst case) and **t**o investigate the implications on minimum impulse bit (nominal orienta**t**ion) for poin**t**ing control. Gravi**t**y gradient was the dominant influence with solar pressure playing a significant role for the wrap rib and hoop/column designs. Table 18 shows a summary of the RSS torques for 400 km, 500 km and GEO. It is again noted that the difference between the LEO and GEO requirements is one to two orders **o**f magnitude. Also nominal and worst case torques are similar for some designs indicating some configuration optimization poten**t**ial.

#### 2.1.2 LEO Stationkeeping Requirements

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Using an operational scenario of LEO deployment and checkout, the **;** requiremer.\_s t**o** maintain orbit altitude in LEO must be considered. To analyze these requirements a Boeing proprietary simulation program called andlyze these requirements a boeing proprietary simulation program called<br>LTESOP (Long Term Earth Satellite Orbit Prediction Program) was used. This<br>program incorporates all significant perturbations in LEO. These divides all significant perturbations in LEO.<br>These perturbations are summarized below. perturbations are summarized below.<br>
• 4th order spherical harmoni

- 4th order spherical harmonic expansion for earth potential
- $\bullet$  Gravitational attracion of sun and moon
	- Atmospheric drag
	- Solar radiation pressure with shadowing from the earth

The AV and thrust level requirements are a function of the tolerance<br>allowed for recovery to the deplement altitude A range of altitude \_. allowed for recovery to the deployment altitude. **.**%range of altitude tolerances was considered from 0 (continuous thrusting) to 50 km tolerance. Nonlinear effects from density variation, solar/lunar gravity and earth triaxiality became pronounced for tolerances in excess of 10 km.

An additional complication in this analysis was the solar array angle to the "wind." Figure 51 shows that  $S/A$  angle has a major impact on  $\Delta V$ requirement for the SASP design. This angle is a function of orbit, body prequirement for the sasp design. This angle is a function of orbit, body orientation, and time of the year. An assumption was made that  $S/A$  angle would be treated as a constant at 20 degrees angle of attack (90 degrees = would be the constant as  $\mathbf{w}_i$  and  $\mathbf{w}_i$  and  $\mathbf{w}_i$  degrees angle of a constant  $\mathbf{w}_i$ flat t**o** the **w**ind). Th**is** shallow angle wa**s** th**o**ught to be a rea**s**onable :: average because the array may be "feathered" on the dark side and because a \_ flat a**r**ray would only occur around the terminator. It is suggested that a complete simulation of S/A incidence angle be conducted for future analysis to verify this assumption.

> The atmospheric model and orbit altitude as**s**umed also had a major impact in the calculation of  $\Delta V$ . Figure 52 illustrates these two effects for the operational SOC design. The small oscillations in the curves are due to **o**pe**r**at**io**nal S0C de**si**gn. **T**he small osc**i**llations in the cur**v**e**s** ar**e** due to the 2**8** day lunar cycle. A**s** stated previously, we as**s**umed the NASA neutr**al** atmosphere and examined altitude**s** from **3**00 to 500 km for each de**s**ign. It

TABLE 17. DISTURBANCE TORQUES AT GEO

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TABLE 18. DISTURBANCE TORQUE (N-M) RSS SUMMARY

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· NASA NEUTRAL ATMOSPHERE

· 500 KM INITIAL ALTITUDE



. NO STATIONKEEPING ASSUMED

Atmospheric Modeling Impacts

Initial Altitude Impacts

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FIGURE 52. IMPACT OF ATMOSPHERIC MODELING AND INITIAL ALTITUDE ON ORBIT DECAY

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became apparent early in this analysis that 300 km was an unreasonable altitude to maintain for any of the classes studied. Figure 53 illustrates the rapid decay of orbit altitude for representative classes. Based on this initial data 300 km was dropped from any further consideration.

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Altitude tolerance effects were shown to have a major impact on the  $\Delta V$ requirements for all vehicles. This effect is illustrated in Figure 54 for the operational SOC design. At lower altitudes, for this design, the increasing atmospheric density dominates the requirements and continuous thrusting yields a lower  $\Delta V$ . For higher altitudes the density effects are much smaller and higher tolerances yield more efficient operation and less<br>AV, For the other designs less sensitivity is seen to tolerance as shown For the other designs less sensitivity is seen to tolerance as shown<br>Des 19 and 19-1. A summary the LEO stationkeeping requirements in Tables 19 and  $19-1$ . A summary the LEO using a 10 km tolerance is shown in Figure 55. This figure shows that  $\Delta V$  for LEO stationkeeping is a linear function of area to mass. The for LEO stationkeeping is a linear function of area to mass. implications on the propellant mass required for L**E**D operation are illustrated in Table 19 for various I**s**p's and two altitudes. This data is compiled for 90 day LEO operation. Shorter time periods would, of course, **:** considerably lessen this requirement. The magnitude of the propellant requirements does indicate that even for short period of time 100's of kg<br>
of propellant would be required to maintain altitude. This requirement of propellant would be required to maintain altitude. must be traded against the benefit of LEO checkout in future studies.

The thrust requirements to maintain altitude were calculated for a range of duty cycles or thrusting times. The propellant requirements to maintain Example for 90 days were added to the structure. This requirement<br>The thrust requirements to maintain altitude were calculated for a range of<br>duty cycles or thrusting times. The propellant requirements to maintain<br>altitude of this increased mass on the thrust level requirement. Generally this<br>effect was insignificant as shown in the lower graph of Figure 56. This of this increased mass on the thrust level requirement. General<br>effect was insignificant as shown in the lower graph of Figure 56. figure shows the total thrust requirements for L**E**D stationkeeping. It indicates that for reasonably short periods, chemical propulsion is required to maintain altitude. Thrust times in hours must be used to lower the total thrust level to electric thruster range. Additional thrust level data for each class is contained in Appendix B.

#### • 2.1.3 Geosynchronous Stationkeeping

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An analysis of the geosynchronous stationkeeping requirements for the principal LSS designs was performed. The analysis used the area and mass prin**ci**pal I\_S des**i**gns **w**a**s** performed. The analy**s**i**s** used the area and mass numbers presented in Append**ix** A. These numbers give the total area and mas**s** for the structure. The propellant and thruster masses were not added in because they were determined in subtask 2C. When the APS mass is<br>determined, the stationkeeping requirements must be recalculated with the<br>total mass (structure + APS mass) and the resulting  $\Delta V$ 's used to adjust the<br>A dete**rm**ined, the stat**io**nkee**p**ing requirements must be recalculated **w**ith the total mas**s** (structure + A**P**S ma**s**s) and t/Je resulting aV's used to adjust the APS mass. This iterative process is performed until a converged value for<br>the APS mass is obtained. The  $\Delta V's$ , propellant mass and thrust the A**P**S m**ass is** obta**i**n**e**d. The \_**V**'**s**, **p**ropellant ma**ss** and thru**s**t requirements presented here are just the fir**s**t cut in dete**rm**ining the stationkeeping requirements. Final answers will differ by a **p**ercentage **f**rom I0 t**o** 20%. **Th**e tren**ds** shown in the preliminary data **wi**ll remain unch**a**nged**.**



200 KM ALTITUDE DECAY FIGURE 53.

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ORIGINAL PAGE IS FIGURE 54. EFFECTS OF LEO STATIONKEEPING FOR DIFFEKENT ALTITUDE TOLERANCES

AV REQUIREMENT PER 90 DAYS (M/SEC AT 400 KM ALTITUDE) TABLE 19.



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TABLE 19-1. VELOCITY INCREMENT PER 90 DAYS (M/SEC AT 500 KM ALTITUDE)

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LEO STATIONKEEPING AV AS A FUNCTION OF AREA TO MASS RATIO FIGURE 55.  $\left(\begin{matrix} 4 \end{matrix}\right)$ 

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FIGURE 56. LEO TOTAL THRUST LEVEL REQUIREMENTS

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Geostationary orbit stationkeeping requirements derive from three sources. A North/South or inclination perturbation results from a single source gravity perturbations from the sun and moon. East/West or longitudinal perturbations result from two sources. The first source is the earth's triaxiality (oblateness and equatorial ellipticity) and the second is from<br>solar pressure disturbances. These three disturbance effects were solar pressure disturbances. These three disturbance described in detail in Reference 9.

There are four standard methods of overcoming the solar pressure effect on a geostationary erbit. These methods are described in detail in Reference 9. Meth\_x] 1 is to continuously cancel the effect of solar pressure by thrusting in a direction toward the sun. This method is highly inefficient and results in aV's which are 20-40% higher than the other methods. Method 2 is to circularize the orbit each time the eccentricity becomes equal to some tolerance. This tolerance can be related to a longitudinal drift This tolerance can be related to a longitudinal drift tolerance and corrected with E/W thrusting. Method 3 is to rotate the line of apsides in such a manner that the solar pressure will cause the eccentricity to decrease to zero before increasing again. The final method maintains the eccentricity nearly equal to zero\* by frequently rotating the apsida] line in such a manner that the longitude of perigee is equal to the longitude of the sun. Methods 3 and 4 are more complicated than Method 2 and generally the difference in  $\Delta V$  for  $A/M$  ratios of .05 or greater between Methods 2, 3, and 4 is small. Method 2 is, therefore, the method we will assume when calculating APS mass. Figure 57 summarizes the GEO stationkeeping perturbations.

The results of the GEO stationkeeping  $\Delta V$  requirements are displayed in Tables 20, 21, and Figure 58. These tables illustrate the  $\Delta V$  and propellant mass of each LSS design. We have selected representative duty cycles for high thrust (p**=**.Gl=.24 hours/day) and low thrust (p**=**.4=9.6 hours/day) systems for c**c**\_parLng \_V requirements. The propellant requirements are based on the ,\_otal,\_V'sin Table 20 and are yearly or, more precisely, estimates **fo**r a 10-year missi**o**n were made when the t**o**tal APS mass estimates fed back to recalculate stationkeeping requirements.

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Figure 58 shows that solar pressure adds significant  $\Delta V$  for area to mass ratios greater than around a tenth. Another feature of stationkeeping is that N/S &V is independent of mass and area and is, therefore, a constant for a given eorrection frequency. Also, triaxiality contributes little to the E/W requirement. In comparing the high thrust table with the low<br>Fraction thrust table, we find that changing the duty cycle from 01 (high thrust) thrust table, we find that changing the duty cycle from .01 (high thrust) to .4 (low thrust) increases the  $\Delta V$  requirement by approximately 7%. The added \_V would be more than compensated for by an increased Isp f**o**r low

dece av would<br>thrust systems.<br>Table 21 show<br>listed in Table<br>require 10 ti Table 21 shows the first year propellant requirements using the total AV listed in Table 20. As a conservative estimate, a ten-year mission would require 10 times the numbers shown. For the Educational TV Satellite as**s**uming a**.**hydrazine (**Isp =** 200 sec**)** system, th**e** pr**o**pellant required for a : 10**-**year ml**ss**ion would be **36**% **of** the t**o**tal **s**ystem mas**s**. For **o**ther **c**las**s**es, the fraction ranges from 20 to 40%. Figure 59 summarizes the yearly pr**o**pellant requirement.

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• North/South (Inclination)

- . Gravity perturbations sun, moon
- $\Delta V$  = f(duty cycle)— $\rightarrow$  duty cycle = thrust time on/orbit time (24 hrs)

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• Thrust = f(duty cycle, mass, allowed  $\Delta$ lat, correction duration)



- East/West (Longitudinal)
- $Tr1a$ x $1a11ty 1.75$  m/s/year,  $f(1)$ ong)
- Solar pressure (SP) rotates the line of Apisides, induces eccentricity & E/W drift
- SP  $AV = f (AYM, duty cycle, method)$  $\bullet$
- Thrust =  $f(A/M)$ , duty cycle, long, correction duration, method)



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FIGURE 57. GEO STATIONKEEPING PERTURBATIONS

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- · Solar pressure method 2
- Duty Cycle  $.01 = 15$  minutes,  $.4 = 9.6$  hours



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40% DUTY CYCLE

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TABLE 21. PROPELLANT REQUIREMENTS FOR GEO

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# PROPELLANT REQUIREMENTS FOR GEOSYNCHROHOUS STATIONKEEPING 1. (ESTIMATED AMOUNT REQUIRED FOR FIRST YEAR)





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GEO STATIONKEEPING AV FIGURE 58.  $\mathbf{G}$ 

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· PROPELLANT FOR 1 YEAR OUTY CYCLE = 1%

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FIGURE 59. GEO STATIONKEEPING PROPELLANT REQUIREMENTS

GEO stationkeeping thrust requirements were calculated for the two duty cycles and for  $E/W$  and  $N/S$  requirements. Corrections to these numbers cycles and for E/W and N/S requirements. Corrections to these numbers which took into account the added mass of the APS system were made in subtask 2c. The numbers shown in Table 22 are, however, representative of the requirements. N/S requirements were seen to be higher in general than E/W requirements, howver, these requirements were not trivial for the systems considered and thrusters for N/S and E/W stationkeeping must be added to the structure. Some currently orbiting systems ignore the E/W<br>components. This table illustrates the total thrust requirements. table illustrates the total thrust requirements.<br>The following section. A Thrust/thruster requirements are addressed in the following section. summary of the GEO stationkeeping analysis is shown below:

- North/South requirements dominate up to A/M of .31
- East/West contributions 20-30% of total propellant for selected vehicles
- Correction frequencies of once/week required for future LSS - $.01^{\circ}$  tolerance
- : GEO stationkeeping propellant requirements are 25-30% of **:** payload mass for low Isp systems

#### 2.2 Thruster Location and Sizing

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To establish thrust/thruster requirements for each vehicle, the locations of the thrusters must be determined. Locating the thrusters is recognized as a complex process involving packaging, tank location, propellant line length, etc. In this study, however, the primary concerns were providing 3-axis control and meeting stationkeeping  $\Delta V$  requirements. Consequently, seven criteria for establishing thruster locations were groundruled**.** These " were:

- I. Maximum moment arms used.
- 2. N/S and E/W stationkeeping capability from nominal orientation.
- 3. Zero delta-V maneuvering capability.
- 4. Zero torque stationkeeping capability.
- 5. Minimal heat flu**x** and contamination from plume impingement.
- 6. No **t**hruster mounting on S/A surface or at the ends of S/A's.
- : 7. Minimi**z**e **t**he number **o**f thruster**s** used.

**T**O m**e**et th**e** f**ift**h **cr**i**t**er**ia**, **w**e ground ru**l**ed **that thr**u**st**er**s wi**ll be **cant**ed **at a** m**inim**um **of 45** d**eg**re**es away fr**om **criti**ca**l c**om**p**o**nents such** a**s s**o**l**a**r a**r**r**a**ys** on an**ten**na **su**r**f**a**c**es. Th**e** con**s**e**quenc**e o**f** th**e fi**r**st c**r**iteri**a **was** tha**t** mi**ni**mu**m i**m**p**u**ls**e b**its were** dr**iv**en be**l**o**w stat**e**-of**-**the-**a**rt** ca**pa**b**il**i**ty for s**om**e**

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IZS classes. Certain optimization of thruster may lessen this requir**e**ment. In general, however, the deployed LSS had to have long thruster moment arms // because of the CG location and large distances required to surround the CG with thrusters. The consequence of criteria 2-5 was that a minimum of 12 The consequence of criteria 2-5 was that a minimum of  $12$ thrusters were required for each **LS**S. In some cases up to 18 thruster locations were needed because of canting requirements. Figures 60-64 and associated Tables 23-28 show the locations and directions assigned for each thruster on each configuration. The SASP design presented problems meeting all location criteria. Three axis oontrol using thrusters could not be achieved because the **C**G location and panel positions did not allow the **CG** to be fully enc**o**mpassed. Thrusters m**o**unted on bo**o**ms or a redesign for this configuration was considered beyond the scope of this study. design was dropped at this point fr**o**m further consideration.

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The thrust per thruster was calculated using a general purpose thrust<br>requirements computer program. This program handles arbitrary requir**e**ments computer program. This program handles arbitrary configurations and thruster location/directions. The inputs to the program are torque and stationkeeping requirements, configuration OG location, • thruster locations, thruster direction cosines and thruster purpose (such<br>• as E/W or X torquing). Thrusters will generally be multiple in that they **i** as E/W or X torquing). Thrusters will generally be multiple in that they<br>**in the will provide both torque and delta-V.** To assure that requirements will be %, will provide both torque and delta-V. To assure that requirements will be met and stationkeeping and torquing will be decoupled, the program must solve the following simultaneous equations: =. met and sta**t**i**o**nkeeping and **to**rquing **wi**ll be dec**o**upled, the program must \_ solve the foll**o**wing simultaneous equations:

Eq. 1 -  $\Sigma$ Thrust(x,y or z) = Required Stationkeeping(x,y or z)  $\left\{\n \begin{array}{c}\n \text{Stat} \\
 \end{array}\n \right\}$ Eq. 2 & 3 - IThrust(x,y or z) **=** 0 5**t**at**l**onkee**p**ing

Eq.  $4,5,6 - \Sigma$ Torque(x, y and z) = 0

; and  $\mathbf{and}$ 

Eq. 1 -  $\Sigma$  Torque(x,y or z) = Required Torque(x,y or z) Eq. 2 &  $3 - \Sigma$ Torque( $x, y \text{ or } z$ ) = 0 Torquing Eq.  $4,5,6$  - Intust( $x, y$  and  $z$ ) = 0

If the**re** are four thrusters ava**i**lable for each **o**pera**t**ion, the simula**t**ion eq**u**a**t**ion**s** canno**t** be solved e**x**actly (i.e., si**x** equation**s**, four m\_**k**n**o**wns). :<br><sup>1</sup>/ The unknowns were solved by using a pseudo inverse operation:<br><sup>2</sup>/

$$
\overrightarrow{AX} = B
$$
  

$$
\overrightarrow{X} = (A^{T}A)^{-1} A^{T}B
$$

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A is th**e** N**M**M matrix **o**f c**o**efficients (N > M).

\_**.** X is the **t**hrus**t** requ**i**remen**t** vector o**f** order M.

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Figure 61 Configuration Drawing with Thruster Locations for<br>LAPAA - Electronic Mail and Educational TV

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TABLE 22. THRUST REQUIREMENTS FOR GEOSYNCHRONOUS STATIONKEEPING<sup>1</sup>

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TABLE 23. LAPAA THRUSTER COORDINATES - ELECTRONIC MAIL AND EDUCATIONAL TV

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\* Locations are relative to the arbitrary coordinate system used in<br>the configuration drawing - not to the CG

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#### TABLE 24. GEOSTATIONARY PLATFORM THRUSTER COORDINATES

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\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

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TABLE 25. LAND MOBILE SATELLITE SYSTEM WRAP RIB THRUSTER COORDINATES

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TABLE 26. LMSS HOOP COLUMN THRUSTER COORDINATES



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TABLE 27. SPACE OPERATIONS CENTER (INITIAL) THRUSTER COORDINATES

\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

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TABLE 28. SPACE OPERATIONS CENTER (OPERATIONAL) THRUSTER COORDINATES

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The pseudo inverse provides the optimal thrust levels based on decoupled stationkeeping and torquing. The program will not allow more than four thrusters to perform any one maneuver. This limitation did not cause a problem for any of the configurations and thruster locations considered.

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Using the requirements generated in substask 2a, the thrust/thruster requirements for each thruster were generated. Some thrusters had only disturbance torque cancelation requirements, while most thrusters had both stationkeeping and torquing requirements. If both requirements existed for a given thruster, the thruster was sized for the higher level. This was the stationkeeping requirement in all cases. Appendix C shows the detailed outputs of the thrust determination program. A summary of this data for<br>each design at .15 g-loading is shown in Table 29. Because of time and each design at .15 g-loading is shown in Table 29. funding constraints, the small wrap-rib (25m) and hoop column (60m) were not treated beyond this point. This table shows that LEO thrust/thruster requirements considerably exceeded the E/W requirements at GEO. This is also true for N/S requirements when longer duty cycles were amployed. It also shows that to lower thrust levels to electric propulsion levels **1.01-.1 N/thruster)** long duty cycles were required. This fact led to<br>
conclusions which will be discussed in more detail in Task 3.<br>
An additional conclusion can be reached from the data proported in Armordium \_. conclusions wh**z**ch will be discussed in more detail in Task 3.

\_; An additional conclusior can be reached from the data presented in Appendix Because of the unequal moment arms for stationkeeping thrusters,<br>tling may be indicated for stationkeeping with 0 torque. This throttling may be indicated for stationkeeping with 0 torque. {" consequence is illustrated in Figure 6**5**. \_**h**r**o**ttling rati**o**s of 2:1 t**o 6**:1 **,i** are required.

#### 2.3 APS Characterization

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In this subtask, the number of thrusters, thrust levels, and AV estimative in the subtask, the number of directory director and available of the requirements derived in previous subtasks were utilized to generate APS<br>- The first step to an and converged thrust level requirements. The f mass and converged thrust level requirements. T**h**e first step to characterizing A**P**S mass was the development of chemical and electric system scaling laws. These laws must have thrust level,  $I_{SD}$  and propellant mass estimates as inputs. Hardware scaling is performed by-derived scaling laws and then added along with the propellant estimate to the structure mass. A recalculation of thrust and propellant mass for the 10 year assumed mission r recalculati**o**n **o**f thrust and propellant ma**ss** f**o**r the i0 **y**ear a**s**sumed mi**ss**i**o**n life was then made. Interation was performed until a 1% differen**c**e between old and new system ma**s**s was obtained. This process is des**c**ribed in Figure ¢ **6**6**.**

> The scaling equations for ch**mi**cal systems were derived from curve fits of **•** e**x**isting hardwa**re**. **T**able **3**0 show**s** the data bas**e** f**o**r the mon**o** and biprop s**c**aling. Equations for chemical thruste**r** mas**s**, in**c**luding thruster, valves and moun**t**ing st**r**ucture are shown below:

> > Mono Thrust/Weight = 1.7 Thrust + 8.2 in kg Biprop Thrust/Weight **-** .722 T**h**ru**s**t + 13.1 in kg

TABLE 29. THRUST/THRUSTER (N) RANGE FOR STATIONKEEPING

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Table 30. Chemical Thruster SOA Summary

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网络警察大约学家 医前缀 人名法罗

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 $\label{eq:optimal} \mathcal{L}(\mathcal{D}^{\text{reg}}(\mathcal{A},\mathcal{I})) \geq \frac{1}{2} \sum_{i=1}^n \mathcal{L}^{\text{reg}}(\mathcal{A},\mathcal{I}) \geq \frac{1}{2} \sum_{i=1}^n \mathcal{L}^{\text{reg}}(\mathcal{A},\mathcal{I}) \geq \frac{1}{2} \sum_{i=1}^n \mathcal{L}^{\text{reg}}(\mathcal{A},\mathcal{I}) \geq \frac{1}{2} \sum_{i=1}^n \mathcal{L}^{\text{reg}}(\mathcal{A},\mathcal{I}) \geq \frac{1}{2}$ 

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## **CADITIONS**

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- THRUST LEVEL  $\mathbb{Q}$   $\mathbb{Q}$  , But  $\mathbb{Q}$  is vectored FOR EQUIVALENT MOMENT ARM  $\circledcirc$
- THRUST LEVEL  $\mathbb{Q}$  +  $\mathbb{Q}$  ,  $\mathbb{Q}$  is pulsed for EQUIVALENT MOMENT ARM  $\Theta$
- THRUST LEVEL 1 = 2 + USE MOMENTUM MANAGEMENT FOR COMPENSATION  $\odot$
- $\circled{0}$  thrust level  $\circled{0}$  >  $\circled{0}$

### **IMPACTS**

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- 
- 
- 
- **10 INTRODUCED POTENTIALLY UNDESIRABLE AV**<br>10 MAY EXCITE FLEXIBLE MODES<br>10 COMPLICATED ADDITIONAL PROPELLANT NEED FOR MMD DESATURATION<br>10 MUST HAVE THROTTLING IF SAME THRUSTERS ARE USED, THIS WILL ALSO MUST HAVE THROTTLING IF SAME THRUSTERS ARE USED, THIS WILL ALSO<br>ACCOUNT FOR CG DRAFT OVER LIFE TIME

## THROTTLING PATIOS

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THRUST RATIO REQUIREMENTS FIGURE 65.

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FIGURE 66. APS CHARACTERIZATION METHOD

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For the electric propulsion systems, the equations were derived as follows:  $, "$ 

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$$
M_{\text{T}} = K_1 D^n
$$
  
\n
$$
J_B = K_2 D^2
$$
  
\n
$$
\Gamma = K_3 J_B \text{ Isp}
$$
  
\n
$$
F = K_4 D_2 \text{ Isp}
$$

where

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 $M_t$  = thruster mass  $D = thrust diameter$  $J_B$  = beam current  $\overline{1}$ 

$$
M_T = K_5 (F/I_{sp})^{1/2}
$$

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 $K_1$  through  $K_5$  are constants

The scaling equations used for each APS system are:<br>  $T/W = 1.7T + 8.2$  ( $N_2H_d$ )

$$
T/W = 1.7T + 8.2 \t\t (N_2H_4)
$$
  
\n
$$
T/W = .722T + 13.1 \t\t (N_2O_4/MMH)
$$
  
\n
$$
thruster mass = (1/(T/W))T
$$
  
\n
$$
M_p = propellant mass = N_s (e^{\Delta V/GI_{sp}})^{-75} (Electric Ion Propulsion)
$$
  
\n
$$
M_p = propellant mass = M_s (e^{\Delta V/GI_{sp}} - 1)
$$
  
\n
$$
T_V = tank volume = M_p \neq v
$$
  
\n
$$
T_R = tank radius = 3\sqrt{\frac{3 \text{ Tv}}{4\pi}}
$$
  
\n
$$
T_A = tank area = 4\pi T_R^2
$$
  
\ntank mass = 5.62 T<sub>A</sub>

In addi**t**i**o**n**,** for electri**c** ion propulsion

$$
P = power = 9.8 (T_{max}) (I_{sp})/(2\eta_{sys} x 1000)
$$
  
Asolar array mass = 13.5 P

127<br>D180-27728-2  $\blacksquare$   $\triangle$  solar array area = 8.96 P

mass of power processing unit = 2.1 x 6.5 x P

where

T/W = thrust/thruster mass

 $T = thrust/thruster$ 

 $M_{\rm c}$  = satellite mass

The relevant scaling equations are summarized in Figure 67.

Tables 31 through 37 show the results of the APS scaling exercise in kg. The L**E**O propellant mass is done at 400 km for the Electronic Mail to 1885 kg for the operational SOC design using chemical systems. noted that a high price is paid in propellant mass for short-term operation in LHO, and LEO deployment is an issue deserving careful consideration. A **;** summary of GHO operation scaling is shown in Table 38. This table sh**o**ws a number of interesting conclusions. First, when considering chemical systems only, the percentage of APS mass to total system mass is relatively<br>constant. For mono-propellant systems, four of the five antenna systems<br>have APS percentaage of 24-29%. Bi-props are similar and range between<br>18constant. For mono-propellant systems, four of the five antenna systems have APS percentaage of 24-29%. Bi-props are similar and range between L- have APS percentaage of 24-29%. Bi-props are similar and range between  $\frac{10-22}{10}$ 18-23%. The difference between the 1% and 40% duty cycles makes only a small difference for chemical systems. However, electric systems did not even converge which indicated that power system mass and S/A area were<br>growing so fast that the added thrust requirement due to this added mass<br>and area drove the next iteration upward just as much as the first. At 40% t growing so fast that the added thrust requirement due to this added mass and area drove the next iteration upward just as much as the first. At 40% duty cycle, the converged percentage of APS mass were often 1/3, and in some cases, 1/5 of that for mono-chemical systems.

In summary, the GEO scaling study showed the following conclusions:

- \_; EP dominated by power system mass (thrust level)
- :. State-of-the-art electronic ion propulsin unfeasible at short duty cycles  $( $2-3$  hr.)$
- EP looks attractive at longer duty cycles

- For .4 duty cycle:

- · EP mass savings over mono (500-1500 kg) Avg = 85% savings
- · EP mass savings over bi-prop (300-1000 kg) Avg = 78% savings
- $\mathbf{F}^{\mathbf{F}}$  , and  $\mathbf{F}^{\mathbf{F}}$  are  $\mathbf{F}^{\mathbf{F}}$  and  $\mathbf{F}^{\mathbf{F}}$  are 78% satisfies satis • Chemical systems dominated by propellant requirements, una**f**fected by duty cy**c**le
- F**o**r .01 duty cycle:
	- Bi-pro**p** ma**ss** saving over m**o**no (1**5**0-500 k**g**) Avg **=** 30% saving**s**



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FIGURE 67. APS SCALING EQUATIONS

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TABLE 31. APS MASS SUMMARY FOR ELECTRONIC MAIL

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TABLE 32. APS MASS SUMMARY FOR EDUCATIONAL TV

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TABLE 33. APS MASS SUMMARY FOR WRAP RIB

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TABLE 34, APS MASS SUMMARY FOR HOOP COLUMN



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\cdot\n\end{array}$ 

 $\begin{array}{c} \bullet \\ \bullet \end{array}$ 

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 $\frac{1}{4}$ 

 $\sum_{i=1}^{n}$ 

 $\frac{1}{2}$ 

TABLE 35. APS MASS SUMMARY FOR GEOSTATIONARY PLATFORM

 $\sum_{i=1}^{n}$ 

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TABLE 36. APS MASS SUMMARY FOR SOC INITIAL

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TABLE 37. APS MASS SUMMARY FOR SOC OPERATIONAL

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TABLE 38. CONVERGED APS MASS REQUIREMENTS FOR GEO STATIONKEEPING

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The effect of iteration on the thrust level requiremerts is shown in Figure 68. This effect is very noticeable for the shorter duty cycles using electric propulsion due to the very large power penalty associated with high thrust levels. For chemical systems, the increase was between 30 and 40%.

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Another output of the scaling e**x**ercise was electric power requirements for electric systems. As discussed above, the 1% duty cycle was unrealistic for electric propulsion (EP). For example, at a duty cycle of .01, the increase of total LSS mass due tc the APS is about 60%. At a duty cycle of .4, the increase is only about 4%. The decrease is attributed to the lower thrust levels required, which leads to lower power requirements, and consequently, lower thruster mass, solar array mass, and power processing unit mass. It should be noted that the propellant mass increases with increasing duty cycle. This is expected because there are greater cosine losses with longer duty cycle. Therefore, ch\_nical APS, which do not require as much hardware mass due to power, have mass increases with increasing duty cycle. Since EP shows a considerable savings in mass at the longer duty cycle, an investigation of duty cycle effects on power system requirements was made. Table 39 shows the appr**o**ximate EP power requirements at .01 and .4 study cycles. The table shows that, in general, the .4 duty cycle requires one to two orders of mangitude less power. the .4 duty cycle requires one to two orders of mangitude less power. - Figure 69 shows the total power required using a 40% duty cycle as a function of satellite mass. For the power levels being considered, the f**o**llowing fact**o**rs are important.

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- ': **•** High voltage can cause arcing due to interactions with plasma = near the satellite at LEO (this effect can be considered negligible at GED).
- **•** High currents can generate intense magnetic fields which **:** interact with the earth's magnetic field causing disturbance ., torques (this may be very minor at GBO).
	- **•** High power requ**i**rements result in la**r**ge trans**n**is**s**i**o**n lines which will add mass to cabling systems (this mass may be very significant if power levels are high and transmission lines are long).
- \_, High power requirements lead to grounding problems which can necessitate the coating of the **L**SS with a conductive substance which changes structural characteristics.
- If batteries a**r**e used they will contribute a substantial portion of the power system mass. For instance, in the \_ propo**s**ed SOC de**s**ign, batter**i**e**s** have been **s**i**z**ed at **8**000 ib**s**. Battery mas**s** is la**r**gely a function of the de**p**th of d**i**scharge (DGD). The lower **t**he D**C**D the lower the b**a**ttery **w**eight required. DOD **is** de**f**ined as the percentage **of** nn**=**Lgy needed ver**s**us the total energy available fr**o**m the battery. (In general, a D**C**D of 50% is most de**s**irable).

j 13**8** D1**80**-2**77**2**8**-2



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**新闻的人的复数** 

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\vdots \\
\vdots\n\end{array}$ 

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CONVERGED MAXIMUM APS THRUST REQUIREMENTS FOR GEO STATIONKEEPING FIGURE 68.

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 $\sum_{i=1}^{n}$ 



### TABLE 39. POWER REQUIREMENTS FOR ELECTRICAL PROPULSION (KW)

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FIGURE 69. ELECTRIC ION PROPULSION POWER REQUIREMENTS

Two options exist for APS power supplies - solar arrays and batteries. If two batteries are used, the duty cycle and hence thrust level has little effect on the size of battery used. This is because the total energy for either duty cycle is about the same (energy = power  $x$  time). If solar arrays are used a longer duty cycle has a positive effect in lowering the required power level and the arrays may be smaller in size. A possible negative effect from longer duty cycles may result if the payload operations and APS power sources are the same. In most missions the peak power load is considerably below the average power required. Solar arrays sized for this peak power load may be used to supply power to the APS during off peak time. As the duty cycle increases, the likelihood of being able to draw on this "free" power source decreases. The extent of this penalty for longer duty cycles is very mission dependent and is noted here only as a qualitative observation.

### 2.4 APR/LSS Interactions

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Interaction of the propulsion system with the structure becomes a key issue for LSS, particularly those systems which must operate continuously. If **;** the antenna beam is defocused in any way**,** this translates into a power or • gain loss in the signal and a degradation of user service. The approach<br>taken to analyze this interaction is summarized below.

- Used NASTRAN mode shapes to simulate dynamic interactions.
- t Examined steady state and transient response.
- Modeled with and without APS mass for higher (LEO) thrust values.
- Thruster masses placed at thruster locations.<br>• Propellant, tank masses at  $\alpha$ .
	- " Propellant, tank masses at 0**3**.
	- Modeled without APS mass at lower (**Q**?.; thrust levels.

It was found from the LEO analysis that t**h**ruster mass had little effect on APS mass to preserve resources.

degradation; therefore the GEO thrust level analysis was made without the<br>APS mass to preserve resources.<br>The major areas of interest for this analysis were 1) the defocusing of the<br>antenna, and 2) the stresses in the stru The major areas of interest for this analysis were 1) the defocusing of the antenna, and 2) the stresses in the structural members. The structural response was computed with the NASTRAN finite element computer program, **;** using the c**o**mputer models developed in Task i. A modal transient analysis \_T was used to comp**u**te the forces in the structural members, the relative ] displacement**s**, and the relative accelerati**o**n**s of** the point**s** on the ; structure. T**h**e mode shapes found in Task 1 were studied to determine the number of modes to be used **f**or **t**his analysis. The main consideration in \_ choo**s**ing **t**he modes **w**a**s to** in**c**lude all modes whi**c**h had **s**ignifican**t n**\_dal deflec**t**ion a**t** any thruster loca**t**ion. It was decided that the f**i**rst ten flexible modes would provide sufficient accuracy since all higher modes

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were ]ocal modes of components and would therefore have little effect on the analysis. The frequency range for the first ten flexible modes was 0.090 to 0.255 Hz.

 $\sqrt{2}$ 

Displacements computed with NASTRAN were verified using the following equations.

Steady State Modal Response

$$
q_{SS} = 1/m\omega^2 \sum_{j} F_j \phi_j
$$

where

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 $q_{\rm cc}$  = steady state modal response (m)

 $m =$  generalized mass (kg)

 $\omega$  = frequency of response (rad/sec)

 $F =$  force of point  $j(N)$ 

 $\phi_i$  = mode shape at point j for response frequency,  $(m/m)$ 

assuming damping is small

 $q_{max}$  ~ 2  $q_{ss}$ 

maximum displacement due to load F**j**

 $\delta$   $j_{max}$  =  $\phi$ <sub>j</sub> q<sub>max</sub> (meters)

Displacements calculated with these equations compared very well with the<br>maximum displacements calculated using NASTRAN. (NASTRAN computes an<br>entire displacement versus time history, while the above equations only<br>compute maximum displacements calculated using NASTRAN. entire displacement versus time history, while the above equations only compute the maximum value of displacement). The displacement calculated by NASTRAN were used to compute the amount of defocusing caused by each stationkeeping maneuver.

The question of structural integrity proved to be a non-issue. Maximum g-loading from even the short duty cycle LBO stationkeeping thrust levels WASTRAN were used to compute the amount of defocusing caused by each<br>stationkeeping maneuver.<br>The question of structural integrity proved to be a non-issue. Maximum<br>g-loading from even the short duty cycle LEO stationkeepi show typical bending moment and force responses for the LMSS Wrap Rib configuration. Tables 40 and 41 show that the percent of stress exhibited was always less than 1% of the maximum allowable. This conclusion does not mean defocusing did not occur, only that the flex in the structure did not threaten the structural integrity.

> To analyze defocusing effects, four geometrical definitions of defocusing were utilized. These definitions have been previously employed on programs such as the Space Telescope. These four effects are illustrated in Figure 72. The motion of selected grid points was summed to find the various defocusing effects. This approach is illustrated in Figure 73.

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LMSS WRAP RIB - BENDING MOMENT IN SOLAR ARRAY BOOM

FIGURE 70.

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FIGURE 71. LMSS WRAP RIB - AXIAL FORCE IN SOLAR ARRAY BOOM

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Maximum Axial Force Thrust Level of 8.12 N at Each Thruster Location



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TABLE 41. LMSS WRAP RIB - MEMBER FORCES

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FIGURE 73. APS/LSS INTERACTIONS DETERMINATION

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Sensitivity to defocusing is a function of frequency and f/d (focal length/diameter). Power loss of 10% was considered the maximum acceptable<br>qoal. Figures 74 through 78 show the sensitivity for the chree effects Figures 74 through 78 show the sensitivity for the three effects analyzed.

Stationkeeping thrusts were initially applied to the LSS without adding APS mass (propellant, tank, and thruster masses). This enabled us to determine the structural response (stresses and displacements) of each LSS without biasing the results toward any particular propulsion system. The the results toward any particular propulsion system. thrust/thruster for the first load case was chosen as the highest thrust required at any thruster to perform all stationkeeping and torquing maneuvers at LEO. These thrust levels were applied to the four models to determine structural response. From the displacements, three defocusing parameters (decenter, despace and tilt) and the power loss associated with each were calculated. Power loss was taken from Figures 74 through 78, where the operating frequency (i.e. UHF, SBAND) and f/d were given in the configuration data. The results of load case one, LEO stationkeeping thrust levels with no APS mass added to the structures, were as follows:

- Large Aperture Phased Array Antenna (LAPAA) less than ten • percent power loss for all three defocusing parameters.
	- Geostationary Platform (GP) less than ten percent power loss :{ for all three defocusing parameters for both the UHF and peta antennae.
	- Wrap Rib Land Mobile Satellite System (LMSS) less than ten percent power loss due to decenter and despace, and ten to fifteen percent power loss due to tilt.
	- Hoop Column LMSS less than five percent power loss due to despace and tilt, and ten to fifteen percent power loss due to : de**c**enter.

For the second load case, the propellant system was assumed to be chemical bipropellant. Propellant and tank masses were placed as close to the c.g. as possible, while thruster masses were placed at each thruster location. : Dynamic characteristics of each structure with A**P**S mass were computed. To :: cut expenses, **o**nly the worst case direction of stationkeeping (x or y) found in load case one was computed for load case two. LEO stationkeeping thr**us**t levels used in load case **o**ne were appl**i**ed t**o** the structures and stru**c**tural response was **c**om**p**uted. The change in defocusing parameters from load case one to two varied from structure to structure:

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- LAPAA all three parameters increase.<br>
 GP all parameters increase for the antenna, decenter and tilt increase, w , **•** GP **-** all parameter**s** increase f**o**r the peta antenna; **f**or the UHF antenna, decanter and tilt increase, while despace decreases.
	- Wrap Rib LMSS decenter an**d** tilt in**c**rease, despa**c**e decreases.
	- Hoop Column LF**S**S **-** all defocu**si**ng parameters dec**r**eas**e**.

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FIGURE 74. POWER LOSS FOR UHF BAND (400 MHZ)

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FIGURE 75. POWER LOSS FOR L-BAND (1 GHZ)

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FIGURE 76. POWER LOSS FOR X-BAND (10 GHZ)

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FIGURE 77. POWER LOSS FOR S-BAND (3 GHZ)

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FIGURE 78. POWER LOSS DUE TO TILT (INDEPENDENT OF FREQUENCY)

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Time and budget prevented us from fully checking out the reasons for the varied responses to adding APS mass. However, we feel that a more detailed study of the dynamic characteristics cf each structure before and after the addition of APS mass could result in a systematic way of predicting the effect that the addition of a given APS mass distribution would have on the performance of the structure.

The third load case was done at GEO stationkeeping thrust levels, assuming the same APS mass distribution used in load case two. A nominal thrust level of 2.0 N thrust/thruster was used for all four models. This level is currently lower than the state-of-the-art for chemical bipropellant engines. It is also somewhat higher than those thrust levels available from electric systems. However, this thrust level is representative of GEO stationkeeping requirements. For this load case, all models except the Wrap Rib LMSS showed power losses of five percent or less associated with all three defocusing para**n**eters for both x and y stationkeeping directions. The Wrap Rib LMSS had less than five percent power loss due to decenter and despace, but five to ten percent power loss due to tilt y-stationkeeping and ten to fifteen percent power loss due to tilt during x-stationkeeping. The results of all three load cases are shown in Tables 42 through 45.

:<br>Tilt and decenter seem to be the limiting parameters in the performance of<br>the Wrap Rib LMSS at LEO thrust levels. Assuming that a power loss of more<br>than ten percent results in unacceptable performance, one can find the the Wrap Rib LMSS at LEO thrust levels. Assuming that a power loss of more than ten percent results in unacceptable performance, one can find the average decenter at ten percent power loss between Figures 74 and 75, then<br>linearly interpolate between 8.12 N/thruster and 2.0 N/thruster, load cases<br>2 and 3 respectively on a plot of decenter vs thrust level to determine linearly interpolate between 8.12 N/thruster and 2.0 N/thruster, load oases 2 and 3, respectively, on a plot of decenter vs thrust level to determine<br>
that a thrust of approximately 7.3 N/thruster would produce ten percent<br>
power loss due to decenter. Tilt, however, is a problem even at GEO thrust that a thrust of approximately 7.3 N/thruster would produce ten percent power loss due to decenter. Tilt, however, is a problem even at GEO thrust levels (2.0 N) and an estimation of how low the thrust/thruster needs to be for a power loss of ten percent or less due to tilt is difficult to make.

For the Hoop Column, decenter limits performance at LEO. Assuming a linear<br>relationship between thrust level and decenter, one can interpolate between relationship between thrust level and decenter, one can interpolate between<br>30.0 N/thruster and 2.0 N/thruster to determine that a thrust of  $30.0$  N/thruster and 2.0 N/thruster to determine that a thrust approximately 5.0 N/thruster would result in ten percent power loss due to decenter. See Table 44 and Figure 74,  $f/d = .55$ .

> In reviewing the results it is important to keep in mind that the power losses due to each defocusing parameter must be added to estimate the overall power loss of each antenna. Assuming that an overall power loss of ten percent or less is acceptable for these structures, the results can be

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- summarized as follows:<br>
Werall power loss of<br>
ten percent or less is<br>
summarized as follows:<br>
LAPAA 6.96 N, Wr.<br>
(LAPAA 6.96 N, Wr.<br>
Platform 7.2 N)<br>
acceptable, while<br>
Column LMSS may b • At the LEO stationkeeping thrust levels used in this analysis **(LAPAA 6.96 N, Wrap Rib 8.12 N, Hoop Column 30.0 N, and Geo** Platform 7.2 N), the performance of the LAPAA and GP seem to be acceptable, while the performance of the Wrap Rib LMSS and Hoop Column L\_3S may pr**ove** to be unac**c**eptabl**e**.
	- " At G\_D sta\_**io**nkeeping thrust level**s**, perf**o**rmance of all bu**t** the Wrap Rib 'LMS<sub>S</sub>' seems to be acceptable.

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Table 42

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Defocusing for Large Aperture Phased Array



 $* 5$  10% Power loss

Note: 0-5% power loss unless otherwise stated

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Table 43

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Defocusing for Wrap Rib Land Mobile Satellite System



 $A = 0$  – 5% power loss

 $B = 5 - 10%$  power loss

 $C = 10 - 15%$  power

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Defocusing for Hoop Column Land Mobile Satellite System Table 44

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5% power loss  $A = 0$ 

10% power loss  $\overline{5}$  $\frac{1}{\alpha}$ 

 $C =$  more than  $10\%$  power loss

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Table 45<br>Defocusing for Geostationary Platform

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 $* 5 - 10%$  Power loss

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Noie: 0-5% power loss unless otherwise stated

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Although time and budget prevented us from doing so, we feel the following would be worthy of further investigation.

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- A more detailed analysis of the dynamic characteristics of each structure before and after the addition of APS mass.
- Methods of strengthening structure to minimize defocusing effects.
- Consequinces of distributing APS mass on flexible structures.
- Utilizing distributed thrusters to perform slew maneuvers of flexible appendages.

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### 3.0 ASSESSM**E**NT OF TECHNOLOGY IMPROVEMENT BENEFITS

The objective of Task 3 was to identify state-of-the-art adequacy/deficiency and the benefits of increasing technology capabilities. To accomplish this objective, three major subtasks were executed.<br>first subtask was to characterize the state-of-the-art propo subtask was to characterize the state-of-the-art propulsion capabilities. This subtask was broadened from the more focused scaling exercise to include such systems as inert gas thrusters, resistojets, MPD thrusters,  $\mu_2$ / $\mu_2$  and others. The second subtask was to determine the  $\texttt{state-or-check}$  limitations. This was done in terms of delivery system capability (both STS and OTV), pointing control capability in terms of minimum bit and valve cycling requirements, and thrust level/I considerations. The third subtask was to assess the enhanced technology This closely paralleled the second subtask but a more indepth examination of **.**momentum manag\_**n**ent versus jet control, electric propulsion system mass, and thrust level duty cycles were considered. The key assumptions used in Task 3 are shown below in Table 46.

- Table 46. Task 3 Key Assumptions<br>
A) State-of-the-art characterization<br>
(a) State-of-the-art characterization<br>
(b) Ton propulsion -- SEPS technology A) State-of-the-art characterization
	- Ion propulsion -- SEPS technology
- \_ B) State-of-**t**he-art limitations
	- Maximum moment arms
	- Uniform \_'**.**hrust/thruster
	- • 30,000 kg STS capability /
	- 4,800 kg (mix) LEO to G**B**O transfer capability (Centaur G'
- **;**\_ C**)** Enhan**c**ed **t**echn**o**logy bene**f**its
- .\_ Le**ss** ma**ss** means le**s**s cos**t**
	- Shorter du**t**y cycle**s** mest desirable

## 3.1 State-of-the-Art Characterization<br>
and State-of-the-Art Characterization

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The systems considered for comparison with the propulsion requirements derived in Task 2 are shown in Table 47. In addition to the systems, **c**e**r**ta**i**n **syst**em**s** ex**ist w**h**ic**h a**re l**e**ss** chara**ct**er**i**zed in terms **of** scaling propertie**s** but have exper**i**mentally or theore**t**ically verified performance \_. \_'egime**s**. F**i**gure **79 s**h(y\_**s** the se**t o**f es**t**abli**s**hed and in-devel**o**pmen**t thc**u**st**er **t**echnol**ogi**e**s**.

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TABLE 47. SOA CHARACTERIZATION

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SYSTEMS PERFORMANCE COMPARISON



## ION COMPONENT SPECIFIC MASSES

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# **SOA Capability/Requirements Map**



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### 3.2 State-of-the-Art Limitations

**T**he limitations of the state-of-the-art capabilities fal] into two general categories - those which constitute mission disabling limitations and those which if eliminated would be mission enhancing. The limitations iuentified are in the following areas: transfer vehicle derivery thrust,  $A = 1.5p$ impacts on STS delivery payload, minimum firing time, thruster cycling requirements, and thrust level requirements.

### Transfer Vehicle Thrust

It was found that the deployed antenna systems were sized for .15 g's. In Task 1 it was also shown that significant penalties in structure mass existed for increased g-loading. Factors of 75% increase were found for 1.0 g's sizing. Figure 26 showed that a 2500-3000 lbf engine was required for deployed LSS transfer. Current primary thrust engines, such as the RL-10, must operate in a less efficient mode for such low thrusts. A  $LO<sub>2</sub>/LH<sub>2</sub>$  engine of 3000 lb<sub>f</sub> would be mission enabling for LEO deployment and transfer.

### STS/OTV Delivery Limitations

**, in Task 2. The capability of the shuttle/centaur g' to deliver a payload** to GEO is 4810 kg. Using this payload as a benchmark the factor of conservatism allowed for the total system was calculated. Table 48 conservatism allowed for the total system was calculated. illustrates that a monopropellant APS allows little room for system mass \_ growth. In one case, going to a bipropellant Is**p** of 300 seconds is mission enabling. Considering the normal rate of growth of preliminary design mass estimates, the flight ready versions of these systems will require 300 enabring. Considering<br>estimates, the fligh<br>second or greater I<sub>sp</sub>.

### Minimum Firing Time

To establish the feasibility and advisability of using jet systems to cancel disturbance torque and provide pointing control we have analyzed \_ limit cycling under the influence of disturbance torques. In previous q analys**i**s **o**f momentum device**s** versus jet sy**s**t\_**n**s it was con**c**luded that at LBO many **o**f t**h**e configuration**s** had momentum requir\_nents that **exc**eeded the state-of-the-art momentum capability. It will be shown in this section that the propellant requirements for a jet system to point under the large disturbance torques encountered at LEO are in some cases greater than the<br>structure mass itself. We have also shown that stationkeeping propellant<br>at the lower altitudes (300-400 km) is also weny great. In short LEO at the lower altitudes (300-400 km) is also very great. In short, LEO \_ de**r**equirement **p**loyment **s**andinheop**r**ent **e**r**a**tion in the**o**fLSSLS**S**sizema**y**andbe**o**rbitpreclaltitude. uded by the l**a**rge A**P**S

:i GEO o**p**eration i**s** mu**c**h mor**e** beni**g**n an**d** lend**s i**ts**e**l**f** to **ei**th**e**r mom**e**n**t**um m**a**nag**e**ment or jet control for pointing. Jet control use**s** li**m**it **c**ycling to maintain pointing accuracy. Figure 80 shows limit cycling for no distu**r**ban**c**e (Method I**)** and under the presen**c**e **of d**i**s**turban**c**e torques

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TABLE 48. STS/OTV DELIVERY LIMITATIONS

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- Total system masses compared to centaur G' capability (4810 KG)  $\bullet$
- $CONS = Fuel$   $ConserveVottism$  allowed  $\bullet$
- Chemical at 1% duty cycyle, ION at 40% duty cycle  $\bullet$

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Mono propellant limits mission capture for proposed delivery systems

**CONCLUSION** 

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Method 1 - No disturbance torque limit cycling (2 pulses)

Method 2 - Small disturbance torque limit cycle (2 pulses)

- Method 3 Single pulse limit cycling with a critical disturbance torque level  $(D_c)$
- Method 4 Single pulse limit cycling with a disturbance torque higher than  $D_{\mathbb{C}}$
- Method 5 Continuous thrusting against disturbance torques

General Comments:

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- . Method 3 uses the least propellant of any method.
- . Method 3, 4 and 5 have the same average propellant consumption dependancy relation which is a function of disturbance torque, I<sub>SD</sub>, and moment arm.
- . Method 5 has unrealistically low thrust level requirements.

Figure 80. Phase Plane Pointing Method Guide

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(Methods 2-5). For a certain combination cf disturbance level, impulse bit, moment arm, inertia and point accuracy, method 3 results. This is a single pulse limit cycle which utilizes the disturbance torque to provide the second impulse rather than rely on an opposing jet as in 1 or 2. Method 3 require the lowest amount of propellant of any of th**e**se shown. The propellant consumption rates are defined in Table 49.

Propellant consumption is minimized by utilizing the disturbance torque to give you a single pulse limit cycle of the longest duration. This occurs when the disturbance torque is equal to Dc as shown in Figure 81. An example map was calculated for a certain set of IZS sizing and impulse bit as shown in Figure 82.

To minimize propellant consumption one would like to match the thrust requirements imposed at  $D_c$ . Before examining these thrust requirements, we have defined the propellant requirements to see if  $RCS$  is advisable for pointing control. It was determined that for disturbance torques of  $10^{-3}$ N-M or less, just ab(]t any type of propulsion system will have an  $acc$ <sub>ext</sub> acceptable mass. For disturbances of  $10 - 100$  through  $10 - 3$  one is forced into ; higher ISDand longer moment a**r**ms to compete with momentum management. For

torques above 10<sup>-1</sup> N-M momentum management will be required.<br>
Reexamining Table 17 we can see that the two LMSS designs h<br>
torque levels that preclude the use of jets for pointing for<br>
high (~3000 or greater) I<sub>SP</sub> syste .\_ Reexamining **T**able 17 we can see tha**t** the two LMSS designs have disturbance torque levels that preclude t**h**e use of jets for pointing for all but very c high (~**3**000 **o**r greater) **I**s**p** systems. The two SOC designs examined als**o** require high I<sub>SD</sub> and long moment arms to compete with momentum management. The other designs, el(ctronic mail, educational TV, geoplatform and SASP, have much lower **to**rques and jet sys**t**ems **o**f all three **I**sp**'**S and can yield a significant mass advantage over momentum management devices.

### APS Requirement.

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> **Minimum fi** ing the and impulse bits have been calculated for the seven .<br>ISS. The results are tabulated in Table 50. Minimum impulse bit The r:sults are tabulated in Table 50. Minim'm impulse bit • requirements are dominated by attitude control limit cycling. It was shown <sup>4</sup>**-** -"#-I\_ abov**e** tha**t** minimum **p**r**o**pellan**t** con**s**umption f**o**r the ACS **(**a\_**t**\_**\_**de con**tr**ol sy**s**tem) **is** ach**i**eved by u**s**ing a **si**ngle pulse limi**t** cycling scenario. The equati**o**n\_ des**c**ribing this type **of** limit cy**c**ling ar**e** a**s f**ollow**s**:

$$
\vec{a} = \frac{1}{2} \qquad \vec{b} = \frac{1}{4} \frac{\dot{\theta}_d^2}{\theta_d^2}
$$
\n
$$
\vec{b} = \frac{\tau}{2T} (Fe - D_c)
$$

where  $D_c =$  critical disturbance torque

- $I =$  **moment** of **unertia**
- 0d **-** des**ired p**oin**ti**ng a**cc**u**r**a**cy**
- $\dot{\theta}_A$  = vehicle rotation rate

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Table 49. Pointing Propellant Usage Guide

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Definitions:

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- $\omega$ - average propellant consumption rate
- F - thrust level
- thrust time on  $\mathbf{t}$
- moment arm e.
- $\theta_d$  pointing accuracy requirement
- $\mathbf{I}$ - inertia
- $I_{SD}$  specific impulse
- $D \quad \Box$ - disturbance torque
- critical disturbance torque  $D_{\rm C}$
- ratio of  $D/F_{\theta}$ R.

Method 1 - n. disturbance torque limit cycling (2 pulse)

$$
\dot{\mathbf{w}} = \frac{(\mathbf{F}\mathbf{\tau})^2 \mathbf{e}}{4 \theta_d \mathbf{I} \mathbf{I}_{sp}} \qquad \mathbf{D} = 0
$$

Method 2 - smal! disturbance torque limit cycling (2 pulse)

$$
\dot{\omega} = \frac{D}{e I_{sp}} \left( \frac{1 + R + (1 - R) \sqrt{1 - D/D_C}}{1 + R - (1 - R) \sqrt{1 - D/D_C}} \right) \quad 0 < D < D_C
$$

Method 3, 4 - single pulse disturbance torque limit cycling

$$
\omega = D/(e I_{sp}) \qquad D_c \leq D < F_e
$$

Method 5 - continuous thrusting

 $\dot{\omega} = F/I_{sp}$  $D = F_e$ 

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- Used thruster locations (i.e., moment arms) identified in Task 2 Approach:
	- · Calculated 0,0,0 orientation disturbance torques & forced these to be D<sub>C</sub> (critical disturbance)
- $\bullet$  Pointing requirements set at .10

Results:



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**Summary** 

 $\sim$   $\bullet$ 

- $\tau$  = thruster pulse duration
- $=$  thrust level
- $e =$  moment  $arm$

The procedure used to solve for the minimum thrust time is to solve equation (1) for  $\theta_d$  then solve equation (2) for r. The minimum impulse bit is  $Fr.$  The values used in these equations were:

 $D_c$  = the nominal disturbance torque calculated for each LSS

 $I =$  the x, y, or z inertia for each LSS

 $\theta_{\rm d}$  = .1 degree

F = maxim\_**'**a thruster value for each LSS multiplied by a factor of 5

**!**

e = sum of x, y, or z moment arms for each **L**SS

By using the calculated disturbance torques from Task i, we force the nominal disturbance torque to be the critical disturbance torque for minimum proepllant consumption. A maximum thruster value is used for F, since we must size the system for the maximum thrust. The factor of 5 . provides a margin of conservatism.

the Table 50 show, that, on the whole, the minimum firing times and impulse<br>bits exceed the state-of-the-art. This implies that either the : bits exceed the state-of-the-art. This implies that either the state-of-the-art must be enhanced in this area, or alternative pointing scenarios are needed. Some of the alternatives include making the moment arms smaller, or using two pulse limit cycling. A detailed examination of each LSS, each mission requirement, and each propulsion option is required to adequately quantify the benefits of these options. The level of detail each LSS, each mission requirement, and each propulsion option is required to adequately quantify the benefits of these options. The level of detail '; required for this trade study was felt to be beyond the scope of the current study.

### ,, Reliability of Thruster Pulsing

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In addition to mass, reliability also has to be considered. The state-of-the-art pulse range life for chemical thrusters lies fr $\mathfrak{m}$  10<sup>5</sup> to 10<sup>6</sup> pulses. The worst case number of pulses for each sitellite for different firing times was calculated. These are the equations used to calculate the number of pulses.

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Single pulse -

$$
\dot{\theta}_{d} = \frac{F \cdot e}{2I}
$$
\n
$$
T = \frac{g}{\dot{\theta}_{d}}
$$
\n
$$
number \text{ of pulses} = \frac{mission \text{ time}}{T}
$$

Two pulse -

$$
\dot{\theta}_{d} = \frac{F e}{8I}
$$
\n
$$
T = \frac{\theta_{d}}{\dot{\theta}_{d}}
$$
\nnumber of pulses =  $\frac{mission time \times 2}{T}$ 

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- e moment arm
- $F thrust$ <br>  $e moment$ <br>  $\therefore$ <br>  $r firing$  $\tau$  - firing time
	- I inertia

 $I -$  inertia<br>  $\frac{\partial}{\partial t}$   $\theta$  - pointing accuracy

- $\dot{\theta}_d$  angular rate across the dead band
- T time of single cycle dead band crossing

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The results are tabulated in Table 51. For a firing time of .01 seconds, most of the satellites have pulse ranges that push the state-of-the-art. By pushing the state-of-the-art in firing times and lowering  $\tau$ , all of the pulse ranges will fall into an obtainable range.

### Thrust Level Limitations

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The limitations of thrust level fall into the category of mission enhancing. This is true with the exception of LEO operation. Chemical propulsion is required to meet the disturbance torque levels and<br>stationkeeping requirements in LEO. The thrust level limitations are stationkeeping requirements in LEO. sunmarized in Table 52.

# 3.3 Enhanced Technology Benefits

Having identified areas of deficiency in pointing control, APS scaling (c**o**mponent and propellant mass) and thrust level, the benefits of enhancing state-of-the-art capability were addressed. Three areas of enhancement are discussed in the following paragraphs. The first deals with pointing control enhancement. This study traded APS mass against momentum duscussed in the following paragraphs. The first deals with pointing<br>control enhancement. This study traded APS mass against momentum<br>management mass for 3-axis pointing control. The focal point in this area<br>was Isp and mi was Isp and minimum firing time. The second study addressed APS mass benefits from enhancing electric propulsion technology. System efficiency, PPU and S/A specific mass and thrust level availability were examined. A comparison of EP mass to chemical system mass was then done to show the degree of mass savings for certain improvement ranges in the system \_- parameters. In the third analysis a compilation of state-of-the-art characteristics for both existing and developmental systems (example MPD and pulsed plasma) was mapped against the limits placed on  $I_{SD}$  and thrust level by the requirements calculated for the range of configurations examined. The develo**p**ment of systems in the region indicated shows the technology enhancements needed and the benefits in terms of mission captive, minimized structural excitation, and system mass optimization.

### 3.3.1 Fointing Control Enhancement

3.3.1 Pointing Control Enhancement<br>Minimum firing times to allow single pulse limit cycling which produces a<br>minimum pointing propellant requirement were identified in Section 3.2.<br>The purpose of the pointing control enhan minimum pointing propellant requirement were identified in Section 3.2. The purpose **o**f th**e** pointing contr**o**l enhan**c**ement **e**x**e**rcise w**as** t**o** und**e**rstand :t the benefits **of** redu**c**ed firing times in terms **o**f **c**ompa**r**is**o**r between **"** momen**t**um managemen**t** m**ass** ver**s**u**s** pr**o**pellan**t** m**ass** f**or 3-**axis con**t**r**o**l. Specifically, the firing time to allow propellant mass to be reduced to | ' equal the momen**t**um managemen**t** m**a**s**s w**as calculate**d**.

> Moment**u**m managemen**t** systems **c**onsits of either reaction wheel or control moment gyros which "absorb" environmentally induced torques by e**i**ther **.**: spinning a**t** higher and higher rates (R**N**) or by changing the axis of large momen**t**um vec**t**or and inducing an H\_R **t**orque ((R**G**'s). Both systmns will , r**ea**ch **a** satur**a**ti**o**n point whe**re** the**y** can n**o** longer ab**so**rb **n**\_mentum in **a** <sup>J</sup> given axis and must be desaturated by some other torquing device. Torques which do not vary in a cyclic fashion will cause saturation and are called



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Table 51. Pointing Control Benefits from Increased Thruster Pulse Number

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- · 10 year mission
- · 3 axis jet control
- Chemical SOA capability  $1 \times 10^6$  pulses demonstrated
	- GEO operation



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- Mission capture at .01 s firing time  $3/7$  for  $1 \times 10^6$  pulses
	- Mission copture at .01 s firing time  $5/7$  for  $5 \times 10^6$  pulses
- $2 \times 10^7$  pulses captures all missions over desired firing time range

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Table 52. Thrust Level Limitations

# · Mono

Range available covers all requirements  $\bullet$ 

# · Bipropellant

- Minimum available thrust exceeds maximum requirement for 15 minute duty cycle  $\bullet$
- Minimum available thrust causes defocusing for flexible antenna systems  $\bullet$

# $\bullet$  100

- LEO thrust requirements cannot be met for reasonable duty cycles  $\bullet$
- GEO thrust requirements can be met for long duty cycles (2 5 hrs)  $\bullet$

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secular torques. Torques which reverse sign every 1/2 orbit can be handled quite nicely by M\_D's providing they are sized to handle the momentum buildup in half an orbit. The cyclic and secular components of torque for our vehicles are shown in the following table.

#

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Table 53. Nature of Disturbance Torques - Earth Oriented Vehicles



The assumptions used to size momentum management systems are listed below:

- Momentum Wheels Used (Small Torques)
- $\begin{array}{c|c}\n\bullet & 4 \text{ Wheels Needed for Single Point Failure} \\
\hline\n\end{array}$

• Sized for Worst Momentum Axis

- Cyclic Torques Behave As:  $\begin{array}{ccc}\n\vdots & \vdots & \vdots \\
\downarrow^{m} & \downarrow^{m} & \downarrow^{m} \\
\downarrow^{m} & \downarrow^{m} & \downarrow^{m} & \downarrow^{m}\n\end{array}$ Where Tmax is Maximum Nominal Disturbance
	- Scaling Equations are Based on Actual Hardware Data

The MMD system being considered is a momentum wheel system which needs to be sized to the maximum angular momentum needed. Equations relating mass to angular momentum were developed based ondata from Sperry Flight Systems ".L" as Shown in F**i**gu**r**e**s 83** and **8**4.

For the momentum wheels,

 $M$ **ass**<sub>MW</sub> = -2.8582 x 10<sup>-6</sup> H<sup>2</sup> + 3.5102 x 10<sup>-2</sup>H + 10.358

For the supporting electrical system,

 $Mass_{ES}$  = 2.4476 x 10<sup>-3</sup> H + 2.6631

 $MMD$  Mass = Mass<sub>MW</sub> + Mass<sub>ES</sub>



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-2.8582E-6  $\lambda^2$  + 3.5102E-2 + 1.0358E+1

 $\frac{1}{2}$  .

 $\sum_{i=1}^{n}$ 

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Figure 83. Momentum Wheel Mass vs Impulse

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 $2.4476E-3 X + 2.6631$ 

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where  $H = angular momentum (N·m·s)$ 

The torques are cyclic, therefore

2Tmax T  $H = \frac{1}{\omega} = \frac{1}{\pi}$  where  $\frac{1}{\tau}$ for  $GED_{\tau} = 86,400$ so H = 2.75 x  $10^{4}$  T<sub>max</sub>

MMD mass was calculated for each satellite for one and four wheels. The results are in Table 54. The total M\_D sy**s**tem mass is the sum of the momentum wheel mass, the supporting electrical system mass, and the secular pitch torque propellant mass for a I0 year mission. The difference in mass of the MMD system with the secular pitch torque propellant at  $I_{5p}$  = 220 and the mass at Isp = 3000 is approximately 3% of the \_MD mass and can be considered negligible. Therefore the FLTD system mass is portrayed as unvarying with Isp.

**;** Figures 85 through 91 are graphs displaying both the APS propellant mass and the total MMD system mass. The trade off in mass between the two systems is clearly marked by the line representing the MMD system mass. The points where the lines in**t**ersec**t** represent critical firing times and are tabulated in Table 55. At these points, the MMD system mass equals the APS propellant mass. For smaller firing times than the cri**t**ical firing times (or at any point to the left of the MMD system mass line on the graph), the APS propellant mass is less than the R\_) sys**t**em mass.

### 3.3.2 Benefits from Enhanced Electric Propulsion Technology

The goal of this exercise was to identify the high leverage technology enhancements for ion electric propulsion in terms of total system mass and thrust level. The approach used for this analysis is described below:

- " Utilize developed software scaling, N/S Gec st\_ior\_**"** ".\_, (4 , thruster**s** used)
- **: @** Take EP scaling equations and vary:
	- Initial satellite mass 500 kg i00,000 kg
	- E**P** sy**s**tem 70-90%

%

%

- **P**PU and S/A specific mass (kg/kw) 13.5 5 kg/kw
- Th**r**ust level available .01 **5** N
- **•** Find G\_D stationke\_ping **s**ystem ma**s**s for given a**ss**tmi\_i**o**n**s**
- Compare EP mass t**o** chemical mass (220, **3**00, 500 sec)

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Table 54. Mass for Momentum Management Devices

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Figure 85. APS vs Momentum Management

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Figure 86. APS vs Momentum Management

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Figure 88. APS vs Momentum Management

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Figure 89. APS vs Momentum Management



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Figure 90. APS vs Momentum Management



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Figure 91. APS vs Momentum Management



Critical Firing Times for APS Mass to Equal MMD Mass Table 55.

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 $\bf{D}$  MMD system mass is always less than APS propellant mass for SOC Initial when  $\rm{I_{sp}\leq 300}$ 

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The results of the unenhanced state-of-the-art propulsion comparison is shown in Figure 92. Three points (A, B, and C) have been identified which wil] aid in the understanding of this figure. At point A, a .i N (t**h**rust/thruster) ion thruster system provides a net advantage over bipropellants for structure mass > 1000 kg. At point B a .05 N ion system has a net advantage over all chemical systems for structure mass > i000 kg. At point C, a factor of 2 mass advantage is seen for state-of-the-art ion thrusters for a 2500 kg structure with a thrust/thruster of around .03 N. It is also seen that as thrust/thruster requiremen**t**s approach .5 N all chemical systems have a mass advantage over electric ion systems. The operational scenario of stationkeeping thrusting is, therefore, closely tied to mass advantage. If short duty cycles are desired  $($   $\leq$  1 hour/orbit), relatively high ion thrust levels are required as <**,**own in Table 56.





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 $\label{eq:3.1} \frac{\partial \mathbf{r}}{\partial t} = \frac{1}{2} \left[ \begin{array}{cc} \mathbf{r} & \mathbf{r} \\ \mathbf{r} & \mathbf{r} \end{array} \right] \left[ \begin{array}{cc} \mathbf{r} & \mathbf{r} \\ \mathbf{r} & \mathbf{r} \end{array} \right] \left[ \begin{array}{cc} \mathbf{r} & \mathbf{r} \\ \mathbf{r} & \mathbf{r} \end{array} \right] \left[ \begin{array}{cc} \mathbf{r} & \mathbf{r} \\ \mathbf{r} & \mathbf{r} \end{array} \right] \left[ \begin{array}{cc$ 

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The conclusion of the sta**t**e-of-the-art examination is that if short duty cycles are desired, **t**echnology advances **t**o increase thrust densi**t**y for only small increases in **i**on system mass are required.

., Figure **93 s**how**s** th**e** appr**ox**ima**t**e regi**o**n**s of** the **s**ta**t**e-of-the-ar**t** capab**i**li**ty** • in chemica**l** and **e**lectri**c** sy**st**ems. Th**e** sy**st**ems in devel**o**pment, re**sisto**jets and s**tor**ed inert ga**s**se**s**, we**r**e n**o**t con**s**ide**r**ed in the A**P**S scaling **ex**er**ci**se. Overla**i**d on **t**h**e** capab**i**lit**i**es ma**p i**s th**e** recommended thrust and **Isp** regime f**or t**h**e c**lasse**s** studied. Due t**o** th**e** un**ce**r**t**ain**t**ies inh**e**r**e**n**t** in the forecasts based on preliminary design, a large region of crosshatching **ex**tends th**e** recommended region. A **f**undamental lower limi**t** in **Isp** f**o**r a I**0** year mission results from STS-Centaur G' mass delive<sub>Ly</sub> limitations for most of the classes analyzed. An upper limit in I<sub>SD</sub> is shown which indicates **o**f th**e c**lasse**s** anal**yz**ed. An upper lim**it** in **Isp** i**s** shown which indica**t**e**s** \_ pow**e**r sy**s**tem ma**ss (**in**c**luding the power source an**d** proce**ss**ing hardware**)** J be**c**ome**s** ck\_nan**t** over pr**o**pellan**t** ma**ss fo**r ion sy**s**tems. Thi**s** limi**t** varie**s** ; with thrus**t** lev**e**l requ**ir**ements and longer dut**y** cy**cl**es **of** 2**-5** hours/**o**rbi**t** ;**,** would raise thi**s** limit. In addit**i**on, lower P**P**U specific ma**ss w**ould , in**c**r**e**ase th**e Is**p limi**t t**o include **exi**st**i**ng **i**on thru**sters**.



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Figure 92. Propulsion Evaluat.on

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SOA Capability/Requirements Map Figure 93. **)** 

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Thrust level limitations vary g**r**eatly with LSS class and duty cycle. The lower limit range shown is for N/S stationkeeping with a long 9 hour/orbit duty cycle. Only the pulsed plasme class violates this range of limitations. Thruster lifetime limits for ion thrusters are also violated at these long duty cycles. For ion thrusters this limit is reached for a i0 yea**r** mission at duty cycles of only 5 hours/orbit. The region of structural interactions limits thrust/thruster to less than i0 N. This limit only applies to the large flexible antennas which must **op**erate during a stationkeeping maneuver.

In conclusion, Figure 93 shows that propulsion systems such as augmented N<sub>2</sub>H<sub>4</sub> and other forms of resistojets, low thrust bipropellants, and possibly low I<sub>Sp</sub> ion systems are in line with LSS requirements.

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APPENDIX A

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LSS Characterization

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### CONFIGURATION DESCRIPTION

The antenna is a series of three thin films which are stretched within compression beams and form a ground plane, input plane, and output plane for a bootlace lens. The lens is contained within a compression structure supported from a deployable mast with guy wires. This is supported to the feed horn cluster by space-extendable beams to form an antenna with its length approximately twice its diameter. The solar arrays form two paddles to be one-axis gimballed and sun oriented. They are sized for 65 kW in LEO, while the distribution, conditioning, and batteries sized for 50 kW at GEO. The lens portion will be closest to earth.

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# **OBJECTIVES**

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The LMSS is capable of relaying radio messages to land mobile units throughout the Continental United States. It is intended to service units such as ambulances, police cars, taxies -- in essence all radio dispatched vehicles. Its position in geosynchronous orbit avoids present radio interferences caused by tall buildings, hills, and other factors. By using a sophisticated relay satellite in space, the mobile ground stations can remain small, light, and relatively inexpensive and still provide high quality communications.



### CONFIGURATION DESCRIPTION

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Looking at the 55 meter offset wrap rib concept, the long boom points at the earth's center. The shorter, vertical boom at the right points up to the north supporting the antenna reflector. The large panel at the left is the ultra-highfrequency feed. It and the 55 meter diameter wire mesh reflector are angled to point at the center of the United States near Kansas City. Multiple beams emanating from the feed panel are arranged to cover all contigous 48 states, Alaska, Hawaii and parts of Canada. The solar arrays are sized for 10 kW.



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*GENERALIC CLASS* Land Mobile Satellite System (LMS) - wrap Rib II. Page 4 of 5  
\n*SCALING LANS* 25 Meter Diameter  
\nEffect of g-Load on Mass M<sub>0</sub> = 2041 kg 9<sub>0</sub> = .15 g  
\nBus, etc.: M<sub>0</sub>  
\nHub, etc.: M<sub>0</sub>  
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\frac{M_{Hub}}{N_0} = 0.0436
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Area and Support: M0\n
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\frac{M_{red}}{N_0} = 0.386 + 0.043 \frac{q}{9_0}
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\nRibs and Reflection:  $\frac{M_{ref1}}{N_0} = 0.0103 \frac{q}{9_0}$   
\nSolar Arrays: M<sub>0</sub>  
\n
$$
\frac{M_{S}}{N_0} = 0.0304 + 0.0048 \frac{q}{9_0}
$$
\n\n800ms: M<sub>0</sub>  
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$$
\frac{M_{S00m}}{N_0} = 0.0304 + 0.0112 (\frac{q}{9_0} - 2.6) + 0.00264 (\frac{q}{9_0} - 2.6)^2
$$
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$$
\frac{M_{S00m}}{N_0} = \frac{0.0152 + 0.0112 (\frac{q}{9_0} - 2.6) + 0.00264 (\frac{q}{9_0} - 2.6)^2}{1 - 0.0431 (\frac{q}{9_0} - 2.6)}
$$
\n\nfor  $\frac{q}{9_0} > 3.6$
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**GENERIC CLASS** Land Mobile Satellite System (LMSS) -- Wrap Rib  $II.$ Page 5 of 5 **SCALING LAWS** 55 Meter Diameter  $M_n$  = 3040 kg g<sub>o</sub> = .15 g Effect of g-Load on Mass  $\frac{M_{BUS}}{M_{\odot}} = 0.393$ Bus, etc.: CHOMAL DAGE 13 OF FOOR QUALITY  $\frac{M_{HUD}}{M_{\odot}}$  = 0.09 Hub, etc.:  $\frac{M_{\text{feed}}}{M_{\text{O}}}$  = 0.345 + 0.038  $\frac{q}{q_{\text{O}}}$ Feed and Support:  $\frac{M_{\text{Refl}}}{M_{\odot}}$  = 0.0329  $\frac{q}{q_{\odot}}$ Ribs and Reflector:  $\frac{M_{SA}}{M_{\odot}}$  = 0.049 + 0.0055  $\frac{q}{g_{\odot}}$ Solar Arrays:  $\frac{M_{\text{Boom}}}{M_{\text{A}}}$  = 0.045, for  $0 \leq \frac{g}{g_{\text{A}}} \leq 2.5$ Booms:  $\frac{M_{\text{B}}}{M_0} = \frac{0.0225 + 0.0160 \left(\frac{q}{g_0} - 1.5\right) + 0.0056 \left(\frac{q}{g_0} - 1.5\right)^2}{1 - 0.025 \left(\frac{q}{g_0} - 1.5\right)}$ for  $\frac{9}{9}$  > 2.5

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# **OBJECTIVES**

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The LMSS is capable of relaying radio messages to land mobile units throughout the Continental United States. It is intended to service units such as ambulances, police cars, taxies -- in essence all radio dispatched vehicles. Its position in geosynchronous orbit avoids present radio interferences caused by tall buildings, hills, and other factors. By using a sophisticated relay satellite in space, the mobile ground stations can remain small, light, and relatively inexpensive and still provide high quality communications.



### CONFIGURATION DESCRIPTION

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The 120 meter hoop column concept features independent power units one at either end. The central column points at the center of the United States near Kansas City. Each of the four feed panels at the upper left project a multiple beam pattern onto its assigned quadrant on the large, molybdenum-mesh reflector. There are uplink and downlink feeds for both the eastern and western halves of the country. The radio beams are arranged to cover all contigous 48 states, Alaska, Hawaii, and parts of Canada.

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Land Mobile Satellite System (LMSS) - Hoop Column III. Page 5 of 5 GENERIC CLASS 120 Meter Diameter SCALING LAWS  $M_0$  = 2909  $g_0$  = .15 g Effect of g-Load on Mass  $\frac{M_{+Z}}{M_{\odot}} = 0.563$ +Z Bus and Solar Array:  $\frac{M_{-Z}}{M_{\odot}}$  = 0.066 -Z Bus and Solar Array:  $\frac{M_{\text{Col}}}{M_0} = \frac{0.1981 + 0.0125 \frac{q}{g_0} + 0.00286 \left(\frac{q}{g_0}\right)^2}{1 - 0.0472 \frac{q}{g_0}}$ Column:  $\frac{M_{\text{Hoop}}}{M_{\Omega}}$  = 0.0184  $\frac{g}{g_{\Omega}}$  + 0.0861 Hoop:  $\frac{M_{\text{Refl}}}{M_{\Omega}}$  = 0.034  $\frac{g}{g_{\Omega}}$ CRIGHAL FAGE IS Reflector Surface: OF POOR QUALITY  $\frac{M_{Cables}}{M_0}$  = 0.0082  $\frac{q}{90}$ Cables:

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#### **REFERENCES**

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antenna structure provides greater rigidity and maintenance of surface accuracy which are needed for some applications. The solar arrays are supported by a<br>deployable-boom, and are sized for 8 kW. The remainder of the payloads are mounted on three rigid structures. The solar arrays will be closest to earth.

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GENERIC CLASS Baseline Experimental Geostationary Platform **IV** Page 3 of 3 SCALING LAWS  $M_0$  = 3737 kg  $g_0$  = .15 g Effect of g-Load on Mass All experiments and appendages are good for  $9/g_0 = 10.0$  except the 15 m wrap rib antenna assembly and the solar array assemblies. Bus and Constant Weight Structure:  $\frac{M_{\text{BUS}}}{M_{\text{A}}}$  = 0.8435, for  $\frac{g}{g_{\text{A}}} \leq 10.0$ Wrap Rib Antenna Assembly:  $\frac{M_{\text{Rib}}}{M_{\text{}}}= 0.064 + 0.0062 \frac{q}{g_{\Omega}}$ , for  $0 \leq \frac{q}{g_{\Omega}} \leq 8.9$  $\frac{M_{\text{R1b}}}{M_{\text{A}}}$  = 0.0611 + 0.0062  $\frac{q}{q_{\text{A}}}$  + 0.00294  $(\frac{q}{q_{\text{A}}}$  - 7.9), for  $\frac{q}{q_{\text{A}}}$  > 8.9 Solar Array Assembly:  $\frac{m_{SA}}{M_a}$  = 0.0864, for  $\frac{9}{90} \le 4.6$  $\frac{M_{SA}}{M_0} = \frac{0.0226+0.00233(\frac{9}{90} - 3.6) + 1.227 \times 10^4(\frac{9}{90} - 3.6)^2 + 0.048 + 0.003(\frac{9}{90} - 3.6)},$ for  $\frac{g}{g_0} > 4.6$ 

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# CONFIGURATION DESCRIPTION

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The First Order Platform consists of three stub arms attached directly to the Power System aft section. Attached to these arms are deployable, rotatable payload berthing systems to which payload elements may be connected. The deployment and/or rotation of the payload berthing systems will probably occur when they are being attached and the positions will not be commandable during flight. Power System subsystems will provide payload support. The solar arrays are sized for 25 kW. The vehicle orientation will be variable.

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GENERIC CLASS Science and Applications Space Platform (SASP) ν. Page 3 of 3  $\sum$ 

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# **OBJECTIVES**

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The SOC is a manned operations center supporting a crew of four which will be rotated every 90 days. It provides a location for construction, flight support, satellite servicing, science and applications research, space technology testing, spacecraft test and checkout, and for military derivatives. Some applications are: a payload storage facility used to check out and mate upper stages with their<br>payloads; the retrieval of "small" satellites to the SOC for servicing and then replacing them in their orbital slot; and servicing a communications platform with a manned servicing vehicle.



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![](_page_234_Figure_0.jpeg)

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![](_page_234_Figure_3.jpeg)

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FIGURE 24

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Table 2 WORST CASE DISTRIBUTION TORQUE ORIENTATION @ 300 KM (DEGREES)

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![](_page_239_Picture_72.jpeg)

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![](_page_239_Picture_73.jpeg)

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Table 4 WORST CASE DISTRIBUTION TORQUE ORIENTATION @ 500 km (DEGREES)

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![](_page_240_Picture_76.jpeg)

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WORST CASE DISTRIBUTION TORQUE ORIENTATION @ GEO km (DEGREES) Table 5

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Table 6 Worst Case Distribution Torque Orientation - RSS (Degrees)

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Table 2 LEO Worst Case Disturbance Torques

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Table 3 GEO Nominal and Worst Case Disturbance Torques

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![](_page_250_Figure_0.jpeg)

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![](_page_251_Figure_2.jpeg)

![](_page_251_Figure_3.jpeg)

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APPENDIX C

Thruster Requirements

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#### Location  $(m)^*$ Direction Thruster  $#$  $\boldsymbol{x}$  $\mathbf{Y}$  $\mathbf{z}$  $\pmb{\lambda}$  $\pmb{\mathsf{Y}}$  $\pmb{\mathsf{z}}$  $-4.800$  $0.000$  $0.000$  $\mathbf{1}$ 21.700  $1.000$  $0.000$  $\overline{\mathbf{a}}$  $-4.000$  $0.000$ 21.700  $0.000$  $-1.000$  $0.000$  $\bullet$ 4.800  $0.000$ 21.700  $0.000$ 1.000  $0.000$  $\blacktriangleleft$ 4.800  $0.000$ 21.700  $0.000$  $-1.000$  $0.000$  $\bullet$ 4.800  $0.000$ 21.700  $-1.000$  $0.000$  $0.000$  $\bullet$  $-4.800$  $0.000$  $0.000$ 21.700 1.000  $0.000$  $\overline{1}$  $0.000$ 2.000  $0.000$  $0.000$  $0.000$  $-1.000$ £  $0.000$ 2.000  $0.000$ 1.000  $0.000$  $0.000$  $0.000$ £  $-2.000$  $0.000$  $-1.000$  $0.000$  $0.000$  $10$  $0.000$  $0.000$  $-2.000$  $0.000$ 1,000  $0.000$  $11$  $0.000$ 2.000  $0.000$  $-0.707$  $-0.707$  $0.000$  $12$  $0.000$ 2.000  $0.000$ 0.707  $-0.707$  $0.000$  $13$  $0.000$  $-2.000$  $0.000$  $-0.707$ 0.707  $0.000$  $14$  $0.000$  $-2.000$ 0.000 0.707 0.707  $0.000$

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\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

Table Al LAPAA Thruster Coordinates

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### Thrust/Thruster Requirements (N)

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Electronic Mail (LAPAA), g-load = .15



Table A2 Stationkeeping Thrust Requirements - Electronic Mail

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Electronic Mail,  $g$ -load = .06



## Table A3 Stationkeeping Thrust Requirements Electronic Mail

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### Thrust/Thruster Requirements (N)

 $\begin{array}{l} \mathsf{OR}(\cap \mathcal{W}^{(n)}) \subset \mathbb{R}^n \setminus \{1, \ldots, n\} \\ \mathsf{Gr}(\mathsf{F}_{\mathcal{W}}(\mathsf{w}^{\mathcal{W}})) \subset \mathbb{R}^n \mathsf{F}_\mathsf{f} \end{array}$ 



 $\mathcal{C}$ 

ė

 $\ddot{\phantom{a}}$ 

 $\frac{1}{2}$ 

化二苯基甲基 医三氧

 $\begin{array}{c} 1 \\ 1 \\ 2 \end{array}$ 

D

 $\tilde{\zeta}$ 



### Table A4 Stationkeeping Thrust Requirements

Electronic Mail

 $\mathcal{L}_{\mathbf{a}}$  .

 $\frac{1}{2}$  .

 $\mathbb{Q}_{\ell}$  ,  $\mathbb{Q}_{\ell}$  ,  $\mathbb{Q}_{\ell}$ OF POOR QUALITY

 $\odot$ 

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### Thrust/Thruster Requirements (N)

 $\sqrt[n]{(4)}$ 

 $\hat{\mathbf{r}}$ 

İ.  $\bar{\epsilon}$ 

 $\begin{array}{c} \frac{1}{2} & \frac{1}{2} & \frac{1}{2} \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{2} \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{2} \end{array}$ 

 $\bar{\phantom{a}}$ 

 $\ddot{\cdot}$  $\frac{1}{2}$ 

 $\ddot{\phantom{a}}$ 

|詩『三天の方法』の1985年の創版である。

 $\bar{\mathbf{x}}$  $\frac{1}{\gamma}$  $\frac{1}{2}$ 

ディア・エント

 $\sim$ 

Electronic Mail,  $g$ -load = 1.0



# Table A5 Stationkeeping Thrust Requirements

Electronic Mail

 $\pmb{\ast}$ 

المنابي المساح

 $\frac{1}{4}$ 

 $\mathbf{C}^{\mathbf{m}}_{\mathbf{c}}$  ,  $\mathbf{C}^{\mathbf{m}}_{\mathbf{c}}$  ,  $\mathbf{C}^{\mathbf{m}}_{\mathbf{c}}$  $\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\left(\frac{1}{2}\right)$ OF  $\mathbb{R}^{n}$  ,  $\mathbb{R}^{n}$ 

 $(*)^{\cdot}$ 

 $\mathbf{C}$ 

### Thrust/Thruster Requirements (N)

 $\left( 1 \right)$ 

 $\ddot{\star}$ 

医血管体系 网络金属多元

人名英格兰教

Ż i<br>I

医血管反射 医第二次试验检尿病

医甲基萘胺 化重复重氮 医产

Electronic Mail,  $g$ -load = .06



# Table A6 Disturbance Torque Thruster Requirements Electronic Mail '

网络麻木之口 OF POOR COALITY

 $\vec{r}$ 

 $\mathbf{i}_\star$ 

#### Thrust/Thruster Requirements (N)

Electronic Mail, g-load = .15



# Table A7 Disturbance Torque Thruster Requirements Electronic Mail

 $\mathcal{L}^{\text{max}}_{\text{max}}$ 

 $I$  A.V

I

 $\bar{\textbf{r}}$ 

これに関すれたものでは、彼らのように、 かいかい 長野 プレイヤー かいしんぎ かんこちきにおい 死亡の状態を行い 解表にする プライ

28、新 446000

Ţ

 $C8$ D180-27728-2 ORIGINAL PAGE 11 OF POOR QUALITY

#### Thrust/Thruster Requirements (N)

 $\blacktriangle$ 

医无关节的 法有关人 化电压电压 医血管 医血管 医心脏 医阿尔伯氏征

李世隆

「大きいのです」

Electronic Mail,  $g$ -load = 1.0



Table A8 Disturbance Torque Thruster Requirements Electronic Mail

> $C9$ D180-27728-2

ORKEN FALL N OF POOR QUALITY  $\left(\begin{array}{c} \bigstar \end{array}\right)$ 

### Thrust/Thruster Requirements (N)

Educational TV (LAPAA),  $g$ -load = .15



1. 精神情

 $\frac{1}{2}$ 

医皮肤病的 医血管反应 医直肠内翻肠肠囊切开术

医牙的 医阿尔比亚氏菌属

 $\frac{3}{4}$ 

Table A9 Stationkeeping Thrust Requirements - Educational TV



**OPIORIE**<br>CRITACIO ACHANY

 $\mathbb{Z}^n \times \mathbb{R}^n$ 

j

### Thrust/Thruster Requirements (N)

Educational TV,  $g$ -load = .06



## Table A10 Stationkeeping Thrust Requirements Educational TV

 $\bar{\mathcal{L}}$  .  $\sim$   $\sim$ 

 $C11$ D180-27728-2

بيرها بالتسبب

医皮肤的 计前置 医异常分配试验器 医尿道  $\frac{1}{2}$ **Andrews** 

2.4. 经合同证券利用的处置率

 $\mathcal{L}$ 

# Thrust Thruster Requirements (N)

Educational TV,  $g$ -load = .15

**ORIGINAL** And Alle OF POOR COLLE



Table All Stationkeeping Thrust Requirements Educational TV

 $\frac{1}{2}$  )  $\frac{1}{2}$ 

 $\mathcal{L}$ 

 $\mathcal{L}(\blacklozenge)$ 

 $\left(\frac{1}{2}\right)^{4}$ 

### Thrust/Thruster Requirements (N)

 ${}^i\mathscr{L}$  .  $\infty$ 

ý.

ł.

 $\ddot{\cdot}$ 

医新闻管理 医阿尔伯氏病 医阿尔伯氏征

しんこうきん サライン・オフィング あみも アールもく アール

į.

ŗ, Ļ, Educational TV,  $g$ -load = 1.0

 $C^{m \times n_k}$  $\overbrace{\mathbf{O}_{\mathbf{K}}^{(n)}}^{\mathbf{L}}\mathbf{E}_{\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}^{\mathbf{L}}\mathbf{G}$   $\left(\frac{1}{2}\right)^4$ 

 $\frac{1}{2}^{\bullet}$ 



# Table A12 Stationkeeping Thrust Requirements Educational TV

 $\sim$   $\sim$ 

 $\left(\mathbf{\bar 1}\right)^{\dagger}$ 

 $\overline{\mathbf{I}}$ 

# Thrust/Thruster Requirements (N)

 $\mathbf{r}(\mathbf{t})$ 

 $\frac{1}{2}$ 

大人 计可转变 医布普氏菌素 人名阿尔弗勒 计分类系统 医麦克特氏反射线感染器 花板基

- 大阪時間の事に行ってある場合、最

4

 $\ddotsc$ 

Educational TV,  $g$ -load = .06



Table A13 Disturbanct Torque Thruster Requirements

Educational TV

 $C14$ D180-27728-2

 $C_{\mathcal{R}(\mathbb{C}^2)}$ UNGER<br>OF FOUR RELATY

 $\mathbf{r}^{\star}$ 

### Thrust/Thruster Requirements (N)

 $\frac{1}{\epsilon}$ 

 $\mathbf{I}$  $\bar{z}$ 

医囊突变性

 $\bar{\mathbf{r}}$  $\tilde{V}$  $\frac{1}{\sqrt{2}}$ 

> $\frac{1}{2}$  $\overline{\mathcal{I}}$

> > $\frac{1}{2}$

 $\frac{1}{2}$ 

Educational TV,  $g$ -load = .15

 $\mathcal{L}$ 



### Table A14 Disturbance Torque Thruster Requirements

Educational TV ·

 $C15$ D180-27728-2

 $\mathcal{F}^{(n)}$  and  $\mathcal{F}^{(n)}$  and  $\mathcal{F}^{(n)}$  and  $\mathcal{F}^{(n)}$ 

. . . . . . .

 $\bullet_{\mathcal{Z}}$ 

ORIGINAL JUGE 13 OF POOR QUALITY  $\sigma$ 

### Thrust/Thruster Requirements (N)

 $\sqrt{4}$ 

 $\label{eq:1} \mathbf{r}_1 = \mathbf{r}_2 + \mathbf{r}_3 + \mathbf{r}_4$ 

 $\bar{\psi}$  .  $\frac{1}{2}$ 

> $\blacksquare$  $\ddot{\phantom{0}}$  $\ddot{\phantom{a}}$  $\tilde{\mathcal{A}}$  $\hat{\mathbf{r}}$

 $\frac{1}{\sqrt{2}}$ 

不可能 医

 $\ddot{\phantom{0}}$ 

Educational TV,  $g$ -load = 1.0



Table A15 Disturbance Torque Thruster Requirements Educational TV

 $\sim$  ,  $\sim$   $\sim$ 

 $\sim$   $\sim$ 

 $C16$ D180-27728-2

 $\ddot{\phantom{0}}$ 

 $\sim$   $\sim$ 

#### ORIOWIEL FAGE IN OF FOUR QUALITY

 $\left( \blacklozenge \right)$ 

 $\ddot{\cdot}$ 

- 新のほぼおいても記憶器 シットル200プイダム ( head one search らえり 反映剤 れたり アイ

 $\frac{1}{\alpha}$ 

 $\ddot{\hat{\xi}}$ 

E.

 $\Delta Y$ 

#### Land Mobile Satellite System Wrap Rib



#### Table A16 LMSS Wrap Rib Thruster Coordinates

\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

 $\mathcal{L}^{\text{max}}$ 

### ORIGEIAL TURNER OF POOR QUALITY

 $\mathbf{r}$ 

Thrust/Thruster Requirements (N) Land Mobile Satellite System Wrap Rib  $g - load = .15$ 



### Table A17 Stationkeeping Thrust Requirements LMSS wrap Rib

 $\sim$   $\sim$ 

 $\ddot{\phantom{a}}$ 

 $\sum_{i=1}^{n}$ 

#### Thrust/Thruster Requirements (N)

 $\frac{1}{2}$ 

 $\ddot{\cdot}$ 

**ような、それにもしられるともありましょう。 アセップ もという こくこう そもち こくしょう マーマ 連発 こくちゃん マインド・システム きょうしょ** 

 $\tilde{\phantom{a}}$ 

 $\overline{\phantom{a}}$ 

LMSS Wrap Rib,  $g$ -load = .06

 $0.774$ OF POOK VOID LITY

 $\sum_{i=1}^{n}$ 

¥.

 $\bigcirc$ 



### Table A18 Stationkeeping Thrust Requirements LMSS Wrap Rib

 $\mathcal{L}^{\mathcal{L}}(\mathbf{X},\mathbf{X})$  . As

**Face State State** 

 $\sim 10$ 

 $C19$ 0180-27728-2

 $\sim$   $\sim$ 

والصداد الما

ORIGHA! PAGE IS OF POOR QUALITY  $\left(\frac{1}{2}\right)^{1}$ 

#### Th ust/Thruster Requirements (N)

 $\left(\stackrel{\bullet}{\rightarrow}\right)$ 

 $\hat{\boldsymbol{\beta}}$ 

 $\ddot{\cdot}$ 

これであるのでも、これになることが、その他があるのです。 よくこくぎょう かりか 大き 悪気

 $\frac{1}{\sigma}$ 

 $\pm$ 

 $\frac{1}{2}$ 

 $\mathcal{L}$  , where  $\mathcal{R}$  $\hat{\mathbf{A}}$ 

 $\mathbf{G}$ 

LMSS Wrap Rib,  $g-$ load = .15



# Table A19 Stationkeeping Thrust Requirements LMSS Wrap Rib

#### $C20$ D180-27728-2

كالمستحدث المتأ

 $\mathcal{L}^{\mathcal{L}}$  ,  $\mathcal{L}^{\mathcal{L}}$  ,  $\mathcal{L}^{\mathcal{L}}$ 

*Comment* 

 $\mathcal{A}$  is a set of  $\mathcal{A}$  , and  $\mathcal{A}$  is a set of  $\mathcal{A}$ 

 $\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}})$  and  $\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}})$  and  $\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}})$  and  $\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}})$ 

 $\frac{1}{2}$  and  $\frac{1}{2}$ 

 $\sim$   $\sim$ 

 $\mathbf{r}$ 

 $\epsilon$ 

ORIORIST -<br>OF Pinne Question

 $\left(\frac{1}{2}\right)^{2}$ 

 $\frac{\partial}{\partial x}$ 

### Thrust/Thruster Requirements (N)

 $\mathcal{L}(\mathbf{t})$ 

 $\ddot{\ddot{\tau}}$ 

2. 医神经细胞的 医牙骨的 医中间性皮炎 医中间性皮炎 医心理学 医血管 医前列腺 医神经 医心包 医心包 医中间性 医血管 医血管 医血管 医中间

LMSS Wrap Rib, g-load = 1.0



# Table A20 Stationkeeping Thrust Requirements LMSS Wrap Rib

## **COUNTAINSERS** OF POOR QUALITY

### Thrust/Thruster Requirements (N)

LMSS Wrap Rib,  $g-$ load = .06



Table A21 Disturbance Torque Thruster Requirements LMSS Wrap Rib  $\epsilon$ 

> $C22$ D180-27728-2

 $\ddot{\bullet}$ 

 $(1)$ 

### ORIGINAL PACE IS OF POOR QUALITY

 $\sum_{i=1}^{n}$ 

 $\mathbf{v}_1$ 

### Thrust/Inruster Requirements (N)

 $\sum_{i=1}^n\frac{1}{n_i} \sum_{i=1}^n\frac{1}{n_i} 

 $\ddot{\cdot}$ 

医氯化亚羟氨 人名布拉普基 医三角 医前缘的 的复数医多发条骨髓后隔部的 德鲁利特教学学校

The South States

 $\sqrt{6}$ 

LMSS Wrap Rib,  $g$ -load = .15



### Table A22 Disturbance Torque Thruster Requirements

I.MSS Wrap Rib

 $C23$ D180-27728-2

سا دارد.

ODDS Care and Call OF POOR QUALITY  $\left(\frac{1}{2}\right)^4$ 

ジ,

#### Thrust/Thruster Requirements (N)

LMSS Wrap Rib,  $g$ -load = 1.0



# Table A23 Disturbance Torque Thruster Requirements LMSS Wrap Rib'

 $C24$ D180-27728-2

 $\overline{a}$  .  $\overline{a}$ 

 $\mathbf{v}(\mathbf{v})$  ,  $\mathbf{v}(\mathbf{v},\mathbf{v})$  ,  $\mathbf{v}(\mathbf{v},\mathbf{v})$ 

 $\sim$   $\sim$ 

 $\sim$   $\sim$ 

f.

 $\sum_{i=1}^{n}$ 

**OWNER COM**<br>CORRECTORY

 $\left(\frac{1}{2}\right)^4$ 

 $\mathcal{L}$ 

 $\mathcal{L}$ 

#### LMSS Hoop Column

 $\mathscr{C}$ 

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 $\tilde{\mathcal{A}}$ 

·南方,传统的全型的复数,可以从进去了了科到各省的公司最近的现在分词增增和预期发

化乙酰苯胺 医阿拉伯氏病 医阿拉伯氏病

 $\bar{z}$ 



#### Table A24 LMSS Hoop Column Thruster Coordinates

\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

 $C<sub>25</sub>$ D180-27728-2

 $\epsilon$  .  $\sigma$ 

 $\sigma_{\rm{eff}}$  , and the second contract of the second contract of  $\sigma_{\rm{eff}}$ 

والمنتزع المنحار التوارد

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 $\alpha$ (四) (金沢紀) (金) OF FOOR CUALITY

Thrust/Thruster Requirements (N) LMSS Hoop Column,  $g-load = .15$ 



 $\ddot{\star}$ 

的复数化的复数 医内脏细胞增多

 $\bar{\phantom{a}}$  $\boldsymbol{\delta}$ 

日記入り 新参 下 事前の日本教授におけるため、優先の利益を

あらせもなれば…

医阿里氏病 医阿里氏

Table A25 Stationkeeping Thrust Requirements LMSS Hoop Column

 $\ddot{\phantom{a}}$ 

Thrust/Thruster Requirements (N) (1999) 1999 1999 1999 1999

 $I \times I$ 

÷

 $\mathcal{I}$ 

 $\bar{z}$  $\frac{1}{2}$ 

 $\ddot{\cdot}$ 

 $\gamma_{\rm k}$ ొక్ శక్ర  $\sum_{i=1}^{n}$ 

 $\mathbf{f}^{\star}$ 

 $\mathbf{r}$ 

 $\mathbf{v}_1$ 





Table A26 Stationkeeping Thrust Requirements LMSS Hoop Column

 $\ddotsc$ 

 $\overline{\phantom{a}}$ 

والمتعاطف والمتعارض

 $C27$ D180-27728-2

ومعاونه والمتعاونة والمتواطئ والمتني

STORY TO PACE 18 OF POOR QUALITY

### Thrust/Thruster Requirements (N)

LMSS Hoop Column, g-load = .15



Table A27 Stationkeeping Thrust Requirements LMSS Hoop Column

> $C28$ D180-27728-2

 $\mathbf i$ 

ORIGNWAL POUL 13 Thrust/Thruster Requirements  $\begin{array}{cc} \mathbb{Q}_F^F & P \in \Omega_R^{\infty} & \mathbb{Q}_R \cup \mathbb{Q}_F^{\infty} \end{array}$ LMSS Hoop Column, g-load = 1.0

ľ



 $\sim$   $\sim$ 

Table A28 Stationkeeping Thrust Requirements LMSS Hoop Column

 $\mathcal{L}^{\text{max}}_{\text{max}}$  , where  $\mathcal{L}^{\text{max}}_{\text{max}}$ 

 $\mathcal{A}=\mathcal{A}(\mathcal{A}(\mathcal{A}))$  , where  $\mathcal{A}$ 

 $\frac{1}{2}$ 

 $\overline{\mathcal{C}}$ 

 $\left(\frac{1}{2}\right)^{4}$ 

#### Thrust/Thruster Requirements (N)

 $(4)$ 

And Subject Market State State State State State のことができるから、事実として素晴らしくの意味です。何は表現でもある。

 $\frac{1}{2}$  $\frac{1}{3}$ 

LMSS Hoop Column,  $g$ -load = .06



Table A29 Disturbance Torque Thruster Requirements LMSS Hoop Column

> $C30$ D180-27728-2

 $\hat{\boldsymbol{\theta}}$ 

 $\begin{array}{ccc} \mathbb{C}_{\mathbb{Z}} & \mathbb{Z}^{(1)} \times \mathbb{Z}^{(2)} & \mathbb{Z}^2 \\ \mathsf{OE} & \mathsf{FGC}_{\mathbb{Z}} & \mathbb{Z} & \mathbb{Z}_4 \times \mathbb{Z} \\ \end{array}$ 

#### Thrust/Thruster Requirements (N)

LMSS Hoop Column,  $g-load = .15$ 

 $\mathbb{R}^2$ 



# Table A30 Disturbance Torque Thruster Requirements LMSS Hoop Column

 $0.71$ D18C-27728-2

 $\mathcal{L}$ 

 $\sim$ 

 $\sim$ 

 $\sim 10$   $\alpha$ 

 $\ddot{\phantom{a}}$ 

 $\cdot$  (

 $\frac{1}{2}$ 

 $\sim 100$  km s  $^{-1}$ 

í.
$\sim$   $\sim$ 

 $\omega$  is a second or  $\omega$ 

 $\ddot{\bullet}$ 

### CHARRE PACK IS OF POOR QUALITY

#### Thrust/Thruster Requirements (N)

LMSS Hoop Column, g-load = 1.0



#### Table A31 Disturbance Torque Thruster Requirements

LMSS Hoop Column

 $\left(\frac{1}{2}\right)^4$ 

#### $\mathcal{Q}^{(2)}$   $\cap$   $\mathcal{Q}^{(2)}$  . ी के ज OF  $PCC \leq 7$  $\gamma$

#### Geostationary Platform



 $\ddot{\cdot}$ 

4年の4歳に出来 サルチョウ

聖吉 ミ

あげき かいけいさい あから くる せきじょうそう うつう ああらくあき

 $\frac{1}{2}$ 

医不可逆性 医精神病毒素质

Table A32 Geostationary Platform Thruster Coordinates

\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

 $\mathcal{L}(\mathbf{x},\mathbf{y})$  . In the  $\mathcal{L}(\mathbf{x})$ 

 $\mathcal{A}$  is proportional.

 $C33$ D180-27710-2

## ORIENIAL 1:20E 1S OF POOR OUNCITY

## Thrust/Thruster Requirements (N) Geostationary Platform

 $g-load = .15$ 



## Table A33 Stationkeeping Thrust Requirements - Geostationary Platform

 $\ddot{\phantom{a}}$ 

 $\overline{a}$ 

 $\ddot{\star}$  $\frac{1}{2}$ 

医精神的 医热带细胞瘤 医二甲

からかけ かいかん かんこう きょうけいしょく ふみあん たんき

÷,

 $\ddot{\phantom{a}}$ 

 $\ddot{\star}$ 

 $\mathcal{L} = \bigcup_{i=1}^n \mathcal{L}_i$ 

ORIGNIAL PAGE IS OF POOR QUALITY

 $\blacktriangleright$ 

#### Thrust/Thruster Requirements (N)

أفرق أبياطهم

 $(4)$ 

1.45 → 第一/第一/第三/第三第三/第三/第三/第三

- その近代のアクセントによるところを生きるときに、全国の地域です。 夏

Geostationary Platform, g-load = .06



Table A34 Stationkeeping Thrust Requirements

Geostationary Platform

 $C - 4$ 

 $C35$ D180-27728-2

## CRIGWIAL PAGE IS OF POOR QUALITY

 $\mathcal{A}\mathcal{A}^{(0)}(t)=\mathcal{A}\mathcal{A}$ 

 $\left(\overline{\bullet}\right)$ 

 $\frac{1}{2}$ 

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 $\boldsymbol{\theta}$ 

#### Thrust/Thruster Requirements (N)

Geostationary Platform, g-load = .15

 $\left(\frac{1}{2}\right)^{1}$ 

 $\bm{\bm{\prime}}_t$ 



Table A35 Stationkeeping Thrust Requirements Geostationary Platform

> $C36$ D180-27728-2

 $\overline{1}$ 

OF FOUR QUALITY

 $\mathcal{N}$ 

#### Thrust/Thruster Requirements (N)

 $\sqrt[n]{\Delta}$ 

1.4、新闻的现在分词的"Sunday"的 不是不是我的,它的是我的是我的,我们就是我的,我们的是我们的,我们的人们不是不是。

 $\frac{1}{2}$ 

سيد

Geostationary Platform, g-load = 1.0



## Table A36 Stationkeeping Thrust Requirements

Geostationary Platform



 $\ddot{\phantom{a}}$ 

 $\sim 100$ 

ORICHANA, PRIDE ON OF POOR QUALITY  $\left(\bar{\bullet}\right)^{1}$ 

 $\mathbf{r}$ 

Thrust/Thruster Requirements (N) Geostationary Platform, g-load = .06

يبدأ أدراجهم

 $\frac{1}{2}$ 

1998年,1999年10月1日,1998年10月1日,1998年,1998年,1998年,1998年,1999年,1999年,1999年,1999年,1999年,1999年,1999年,1999年,

 $\frac{1}{2}$ 

 $\frac{1}{2}$ 

Ć



Table A37 Disturbance Torque Thruster Requirements Geostationary Platform

> $C38$ D180-27728-2

 $\delta \Sigma = 0.7$  and  $\delta \Sigma$ 

 $\mathcal{O}(\mathcal{O}(10^6))$  ,  $\mathcal{O}(\mathcal{O})$ 

 $\beta \rightarrow \beta$ 

 $\label{eq:1} \left\langle \left( \mathbf{v}^{\dagger} \cdot \mathbf{v}^{\dagger} \right) \right\rangle \left( \mathbf{v}^{\dagger} \cdot \mathbf{v}^{\dagger} \right) \left( \mathbf{v}^{\dagger} \cdot \mathbf{v}^{\dagger} \right) \mathbf{v}^{\dagger} \mathbf{v}^{\dagger} \mathbf{v}^{\dagger}$ 

 $\mathsf{C}$  $\mathcal{L}_{\mathbf{z},\mathbf{z}}$ 

### Thrust/Thruster Requirements (N)

Geostationary Plation, g-load = .15

¦.÷

2.4 . 东北警察公司新闻管理

- 1000年度

 $\frac{1}{10}$ 

 $\overline{\phantom{a}}$ 

 $\frac{1}{2}$ 

 $\ddot{\mathbf{z}}$ 



Table A38 Disturbance Torque Thruster Requirements Geostationary Platform

D,

#### ORIGANIA PARTICIPA OF POOR QUALITY

#### Thrust/Thruster Requirements (N)

Geostationary Platform,  $g$ -load = 1.0



Table A39 Disturbance Torque Thruster Requirements Geostationary Platform

 $\sim$   $^{\prime}$ 

والواوا والمالوني

 $C40$ D180-27728-2

 $\alpha$  ,  $\alpha$  ,  $\alpha$ 

 $\overline{\widetilde{Y}}$  .

D)

 $(\hat{\bullet})$ 

 $\bar{V}$ 

Children of Scott OF FOOR QUALITY

 $\left( \frac{1}{2} \right)$ 



#### Space Operations Center - Initial

 $\mathcal{U}^{(i)}$  , we

ني.

 $\ddot{\bullet}$ 

医淋巴性肌瘤 化碳基二聚合物 化异聚戊基己烯

 $\frac{1}{2}$ 

 $\frac{1}{2}$  $\frac{1}{2}$ 

 $\pmb{j}$ 

 $\frac{1}{2}$ 

## Table A40 SOC-Initial Thruster Coordinates

\* Locations are relative to the arbitrary coordinate system used in the configuration drawing - not to the CG

#### CRICKWL PARE IS OF POOR QUALITY

Thrust/Thruster Requirements (N)

aansa line

**A** 

 $\vec{r}$ 

医异常的 化反射器 网络非常

Space Operations Center - Initial



Table A41 Stationkeeping Thrust Requirements - SOC-Initial

 $\mathbf{i}$ 

 $Gr \sim$ OF POOR  $Q \in \mathbb{R}$ 

 $\ddot{\phantom{a}}$ 

 $\tilde{L}$ 

 $\mathbf{r}$ 

 $\langle \cdot, \cdot \rangle$ 

#### Thrust/Thruster Requirements (N)

 $\mathcal{A}=\mathcal{A}^{\mathcal{A}}$  , where  $\mathcal{A}^{\mathcal{A}}$ 

 $\bullet$ 

 $\frac{1}{\pi}$ 

 $\boldsymbol{\varepsilon}$  $\tilde{z}$ 

「日本の大きなのであることを、今回のは、はは、は、また、

 $\overline{\phantom{a}}$  $\pmb{i}$ 

 $\mathfrak{p}^{\prime}$  $\mathcal{L}_{\bullet}$ ٠.  $\star$ 

Space Operations Center - Initial



#### Table A42 Stationkeeping Thrust Requirements

SOC - Initial

 $\mathcal{L}$ 

## ORIGINAL PAGE 15

## Thrust/Thruster Requirements (N)

Space Operations Center - Initial



## Table A43 Disturbance Torque Thruster Requirements

 $SOC - Initial$ 

 $\bar{\mathcal{A}}$ 

 $\sim 10^{11}$  km s  $^{-1}$ 

 $\sim 100$  km  $^{-2}$ 

 $\sim$   $\sim$ 

 $\ddot{\phantom{a}}$ 

あたななり こうりょうきょう うちんとうしゃ かんしゅう

计字符 医重装

4

 $\mathbb{Z}_{\mathbb{Z}}$ 

# Chronic Country

 $\left(\overline{\bullet}\right)^{n}$ 

 $\mathbf{r}$ 

 $\bigcirc$ 

#### Space Operations Center-Operational

 $\cdot$ 

 $\begin{array}{c} 1 \\ 1 \\ 1 \end{array}$ 

 $\frac{1}{2}$ 

 $\ddagger$  $\cdot$  $\vec{r}$  $\mathcal{A}$  $\overline{\phantom{a}}$ 

 $\frac{1}{2}$ 

 $\bar{\beta}$ 

 $\hat{\mathcal{A}}$ 

 $\mathcal{A}$  $\ddot{\tilde{\mathbf{r}}}$ 

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 $\frac{1}{2}$ 

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 $\mathbf{I}$ 

 $\sim$   $\sim$   $\sim$ 

 $\sim$ 

 $\ddot{\phantom{a}}$ 



Table A44 SOC-Operational Thruster Coordinates

#### ORIGWIAL PACE IN OF POOR QUALITY

 $\mathcal{L}$ 

Thrust/Thruster Requirements (N)

مده

 $\mathcal{X}$ 

 $\ddot{\bullet}$ 

本文 化基本分析医基督教 未

 $\bar{\mathbf{A}}$ 

计可变 医生物

美国学院 医弗朗斯氏 化机械 医卵体气候 化甲烷

 $\mathbf{r}$  $\cdot$ 

#### Space Operations Center - Operational



#### Table A45 Stationkeeping Thrust Requirements SOC-Operational

CRIGHAM PAGE 15

 $\left(\frac{1}{2}\right)^{3}$ 

 $\sqrt{2}$ 

#### Thrust/Thruster Requirements (N)

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## Table A46 Stationkeeping Thrust Requirements Soc - Operational

 $\mathbf{r}$  and  $\mathbf{r}$  and  $\mathbf{r}$ 

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 $\ddot{\bullet}$ 

"我们的是我们在我们的人的人,我们的人的人,我们的人们的人,我们的人们的人,我们的人们的人,我们的人们的人,我们的人们的人们,我们的人们的人们,我们的人们的人们,

Ý,

 $\bar{\mathbf{r}}$ 

 $\overline{\phantom{a}}$ j.  $\bar{z}$ 

ţ

 $\mathcal{L}_{\mathcal{F}}$  and  $\mathcal{L}_{\mathcal{F}}$  are all  $\mathcal{L}_{\mathcal{F}}$  and  $\mathcal{L}_{\mathcal{F}}$  $\alpha$  , and  $\alpha$  , and  $\alpha$  , and there is a second  $\alpha$ 

**Contractor** 

## ORIGINAL PA ... IS OF POOR QUALITY

 $\sqrt{4}$ 

 $\mathcal{O}(\log n)$  .

## Thrust/Thruster Requirements (N)

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 $\ddot{\cdot}$ 

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 $\mathbb{R}^3$ 

 $\int_{-\infty}^{\infty} \frac{1}{\sqrt{2\pi}} \frac{d^2\mathcal{L}^2}{d\mathcal{L}^2} \, \mathrm{d} \mathcal{L}^2 \left( \frac{d^2\mathcal{L}^2}{d\mathcal{L}^2} \right) \, \mathrm{d} \mathcal{L}^2 \left( \frac{d\mathcal{L}^2}{d\mathcal{L}^2} \right) \, \mathrm{d} \mathcal{L}^2 \left( \frac{d\mathcal{L}^2}{d\mathcal{L}^2} \right) \, \mathrm{d} \mathcal{L}^2 \left( \frac{d\mathcal{L}^2}{d\mathcal{L$ 

 $\sim$  10  $\sigma$ 

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Table A47 Disturbance Torque Thruster Requirements SOC - Operational

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