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TRACKING AND DATA ACQUISITION SYSTEM (TDAS) FOR THE 1990'S

VOLUME VI

TDAS NAVIGATION SYSTEM ARCHITECTURE

DRAFT FINAL REPORT

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Prepared Under: Contract NAS5-26546

Prepared for: NASA Goddard Space Flight Center Greenbelt, Maryland 20771



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PREFACE

This report documents the technical results obtained from navigation system definition studies performed under Tasks 5 and 15 of Contract NAS5-26546, "Tracking Data Acquisition System (TDAS) Study". Task 5 (TDAS System Architecture) and Task 15 (Alternative User Navigation Techniques with TDAS) were part of a two year pre-Phase A concept definition study for TDAS as the proposed successor to TDRSS, currently under development.

SCOPE OF WORK

The TDAS study covers a fifteen year planning period, (1995-2005). Potential missions to be flown in this time frame include free flyers, support vehicles and a space station/platform. Much of the TDAS requirement will be to support low earth orbit (LEO) missions in terms of communications, navigation and TT&C. Additional requirements could stem from user mission activities in higher (e.g., synchronous) orbits, and in support of inter-orbital transfers of materials and men for maintenance and repair in space, or for retrieval of platforms and experiments.

This report is Volume VI of nine volumes constituting the final report for the TDAS pre-Phase A study. Volume titles, largely self-explanatory, are given below:

- Volume I Executive Summary
- Volume II TDAS User Community Characteristics
- Volume III TDAS Communications Mission Model
- Volume IV TDAS Space Segment Architecture
- Volume V TDAS Ground Segment Architecture and Operations Concept

- Volume VI TDAS Navigation System Architecture
- Volume VII TDAS Space Technology Assessment
- Volume VIII TDAS Frequency Planning
- Volume IX TDAS Cost Summaries

OVERVIEW OF VOLUME VI

The navigation architecture aspect of this study involves examination of TDAS-based tracking alternatives for providing user orbit and time determination (OD/TD). Two-way range and doppler tracking as implemented in TDRSS is also an alternative for ground-based navigation support with TDAS. However, the primary focus here will be on one-way range and doppler tracking methods, specifically:

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- Forward Link Beacon Tracking (FLBT) with on-board processing of independent navigation signal transmissions broadcast continuously by TDAS spacecraft,
- Forward Link Scheduled Tracking (FLST) with on-board processing of navigation data received during scheduled TDAS forward link service intervals, and
- <u>Return Link Scheduled Tracking (RLST)</u> with ground-based processing of user-generated navigation data during scheduled TDAS return link service intervals.

This study addresses system configurations and requirements to support each method and assesses the potential navigation performance as a function of user orbit, TDAS constellation options and other parameters. Results are then compared with accuracy requirements in the TDAS mission model. Impacts of the above alternatives on both TDAS and users are evaluated and key issues/ tradeoffs are identified. The study also considers TDAS satellite tracking for two options (BRTS and VLBI)* to identify system configuration impacts of the TDAS constellation options and compare potential tracking accuracy performance.

^{*} BRTS - Bilateration Ranging Transponder System VLBI - Very Long Baseline Interferometry.

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Results of the study are summarized in Section 6. The major conclusions are:

- TDAS Beacon Tracking (FLBT) will satisfy all users in the TDAS mission model with position accuracy requirements down to 10 m.
- Scheduled tracking alternatives (FLST, RLST) can also meet the accuracy requirements except at low altitudes where performance is sensitive to:
 - Drag Uncertainty
 - Frequency of Tracking and/or
 - Frequency of Navigation Data Uploads (RLST/only)
- A two or three satellite TDAS constellation impacts performance as follows:
 - Selecting two satellites leads to tradeoff between coverage and accuracy. Increased satellite spacing improves coverage, but a point is reached where performance in high inclination orbits begins to degrade (130° spacing appears better than 162°).
 - Selecting three satellites provides full coverage and up to a 2:1 advantage in navigation accuracy over two satellites.
- Projected TDAS tracking accuracy requirements (25 m-position and 2.5 mm/sec-velocity) can be met with VLBI tracking but not with a minimal* BRTS configuration.

* 2 Bilateration Ranging Transponder System (BRTS) sites per TDAS spacecraft.

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SECTION 1

INTRODUCTION*

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1.1 BACKGROUND

By the 1990's, NASA missions will require extending TDRSS capabilities for user tracking and orbit determination in terms of greater speed, accuracy and throughput efficiency. Requirements are also expected to increase for on-board, real-time navigation data (orbit, time and attitude) to support mission operations, e.g., data annotation, antenna pointing, rendezvous etc.

A review of technology trends has identified various techniques (existing or under investigation) which could potentially support future orbit and/ or time determination needs using ground-based or quasi-autonomous methods. These can be broadly categorized according to the measurement technology employed:

- RF Signal Detection (range, range-rate, range-difference)
- LASER Reflection (range)
- Optical (angle)

A summary description of the various techniques is included in Volume II of the TDAS Study Report [1].

With TDRSS, RF methods are the basis for providing user tracking services. The focus in this volume will be on extensions to these techniques to support TDAS user navigation and TDAS satellite tracking.

The primary technique for TDRS and user orbit determination (OD) via TDRSS utilizes ground-derived two-way range and/or doppler data and requires a

^{*} A glossary of acronyms is provided in Appendix F.

coherent forward and return link during scheduled tracking intervals (see [2] and Figure 1-1a). This technique can also support user time determination (TD), since clock calibration parameters (bias, drift) may be estimated simultaneously in the OD process.* An alternate tracking technique is also offered, which employs one-way duppler data measured at the ground from user transmissions during scheduled MA return link service (see [2] and Figure 1-1b). User clock calibration cannot be performed with this method (only oscillator calibration). An advantage, however, is that return links are more plentiful and easier to schedule than a coherent twoway link (20 vs 2).

The improved coverage and faster, centralized data collection available with TDRSS should substantially increase the throughput efficiency, of ground-based OD/TD with these techniques. However, achieving the capacity for multi-user support with rapid turnaround or near real-time requirements will depend on computational enhancements at the TDRSS/OSCF. A current goal with hardware upgrades in process and proposed software developments is to achieve, by the late 1980s, a 10 minute turnaround between the end of a tracking pass and an updated orbit computation. [4]

An on-board OD/TD capability would provide timely navigation data and also relieve TDRSS ground requirements for routine mission support. A one-way technique has been developed in which a TDRSS user would extract doppler data during scheduled MA forward link service for on-board OD only (see [5,6] and Figure 1-1c). Ground tests are planned for 1984 to demonstrate the approach and associated user equipment enhancements: doppler extractor, stable oscillator and navigation computer (hardware/software). [7] Further developments will depend on test results and user initiatives.

Clock calibration is also possible with separate two-way measurements once the orbit is determined. [3]

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With respect to future navigation performance, inquiries were made, as part of the TDAS study $[1b]^*$, to determine potential user position and time accuracy requirements. Aside from Topex-type missions^{*}, the most stringent position accuracy requirement is in the 10 m (1 σ) range. This pertains to advanced resource observation type missions and a proposed space station. The most stringent time accuracy requirement (relative to UTC) is 1 µsec (1 σ), which applies to several classes of missions. ŧ.

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1,2 TDAS NAVIGATION ARCHITECTURE GOALS

In accordance with the preceding discussion, the goals of the navigation architecture study are as follows:

- Reduce ground requirements for routine two-way tracking support
- Support user on-board orbit and time determination
- Provide users timely access to ground-derived navigation data, and,
- Meet navigation accuracy requirements (\geq 10 m) of TDAS mission model.

1.3 TASK ASSIGNMENT

The TDRSS two-way tracking technique is a candidate for ground-based support of user navigation in the TDAS era. This study will examine alternative <u>one-way</u> tracking techniques that also support user orbit <u>and</u> time determination with TDAS. The specific alternatives to be studied are:

^{*} Appendix A provides a summary of the TDAS mission model and navigation requirements.

- Forward Link Beacon Tracking (FLBT) based on low power signals continuously broadcast by TDAS satellites to permit range and doppler measurements for on-board use.
- Forward Link Scheduled Tracking (FLST) based on ground originated signals during scheduled contact periods which permit range and doppler measurements for on-board use.
- <u>Return Link Scheduled Tracking (RLST)</u> based on user generated signals during scheduled contact periods for ground-based range and doppler measurements.

The study shall include a system level definition and requirements assessment for each configuration. An accuracy analysis based on applicable error models shall be performed for comparison of navigation performance among alternatives and against requirements in the TDAS mission model. Impacts on both TDAS and users and key tradeoffs will be identified and compared.

1.4 OVERVIEW OF METHODOLOGY

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The methodology followed in pursuing this task is summarized in Figure 1-2. The first step was to define the tracking configuration in terms of functions and system elements needed to support each alternative. The second step was to consider signal processing aspects beginning with a definition of the tracking signal in each case, then a link performance analysis to identify EIRP and/or G/T requirements and finally estimation of the measurement accuracy for range and doppler (range-rate) data types.

The third step was to address TDAS satellite tracking to identify p_{s} ssible system configuration impacts of TDAS constellation alternatives and to determine representative estimates of potential ephemeris accuracy. The analysis focused on two options:

ORIGINAL PAGE IS OF POOR QUALITY PERFORMANCE SUMMARY CONCLUSIONS REQIS. SUPPARY STUDY DATA EVALUATION - SLIDING BATCH FIGURE 1-2: USER ONE-WAY NAVIGATION ALTERNATIVES STUDY - SEQUENTIAL - INCLEMENDE IRACKING ALGORITIMS - ALTITUDE TDAS CONSTELLA-T10NS **IRACKIHG** SCHEDULES - 2 S/C - 3 S/C USER KAVIGATION ACCURACY ANALYSIS USER ł 2 - FRONTSIDE -BACKSIDE TRACKING CONFIG-URATIONS ACCURACY ANALYSIS - BRIS TDAS TRACKING ARALYSIS - VLBI **OVERVIEW** L INK PERFORMANCE TRACKING DATA HEAS ACC. (R, Ř) BEACON SIGHAL DEFINITION TRACKING SIGRAL ASSESSHENT E STANFORD TELECOMMUNICATIONS INC. **⊡** TUAS GROUKD Seghent IDĀS SPACE Sechent FUNCT TONAL OVERVJEN USER TRACKIHG CONFIGURATION DEFINITION SYSTEH ELEMENTS USERS - FLBT FLSI RLST VI-1-6

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- <u>BRTS Tracking</u> the primary technique for TDRS tracking [8], and
- <u>VLBI Tracking</u> proposed as a BRTS enhancement or eventual replacement. [9]

The fourth step was to evaluate potential user navigation performance in terms of OD/TD accuracy for the three tracking alternatives. Cases for evaluation were defined to assess both two and three satellite TDAS constellations, several classes of user orbits, different tracking schedules and two algorithms for tracking data processing (sequential and sliding batch).

The last step was to summarize the study data, compare navigation performance results with requirements defined in the TDAS mission model and then integrate the key findings in terms of a proposed system architecture to support user navigation functions.

1.5 OVERVIEW OF THE REPORT

Results of the efforts outlined above are presented in succeeding sections and in serveral appendices which contain much of the supporting detail. Section 2 describes the functions and system elements defining the configuration for each tracking alternative. Section 3 and Appendix B cover the signal processing analysis aspects. Section 4 and Appendix C consider the TDAS satellite tracking analysis. Section 5 and Appendices D and E present results of the navigation performance evaluation. Section 6 contains the summary data and study conclusions. 11、11日日には1日日の日の日の日

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VI-1-7

SECTION 2

USER TRACKING CONFIGURATION DEFINITION*

The proposed TDAS architecture design [1a,d] extends TDRSS data relay capabilities with enhanced SMA services for multiple access users and more SA channels (KSA + WSA + LSA) for single access users. This section defines the system configurations for supporting one-way user tracking (FLBT, FLST or RLST) assuming (for description purposes) that it is provided at S-Band.

In practice, however, the scheduled tracking alternatives (FLST or RLST) could also be accomodated via the SA services, as necessary. On the other hand, beacon tracking (FLBT) must be a generally available broadcast service and thus is not compatible with the SA services. While FLBT operation in a totally independent band is a possible option, frequency spectrum and user/system equipment considerations tend to favor S-Band reuse.

The following two subsections present the system functions and major elements needed for supporting each alternative. A concluding subsection compares tracking signals and data handling interface requirements for the user, space and ground segments.

2.1 FORWARD LINK TRACKING (FLBT/FLST)

An overview of the major functions involved with the two forward link tracking alternatives is shown in Figure 2-1. Tracking signals are uplinked from the ground at K_u -band, relayed by one or more TDAS satellites and received at S-band by the user for on-board processing. The fundamental differences between beacon and scheduled tracking relate primarily to the signal availability and format (discussed in Section 3).

* A glossary of acronyms is given in Appendix F.

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The beacon signal is assumed to be generated by a TDAS ground terminal for broadcast by each TDAS satellite using a single element of the 60 element SMA antenna array. Users would receive continuous tracking signals while within a beacon antenna's $\pm 13^{\circ}$ field of view. The corresponding upper limit on user altitude for 100% coverage exceeds 3100 km for all TDAS constellations considered.*

In the scheduled mode, signals for tracking are available only during an allocated contact period as part of normal SMA service. Since each TDAS can support two SMA forward channels [1d], up to four simultaneous users theoretically could be supported with a two satellite TDAS constellation depending on channel scheduling policy.

A block diagram of essential elements in the forward link tracking configuration is shown in Figure 2-2. For FLBT support, a separate (beacon) signal generator is employed at the TDAS ground station and a dedicated K_{u} to-S Band repeater is needed in each TDAS satellite [1d]. For either FLBT or FLST, the user requires a stable frequency standard, range and doppler extractors, a decoder for (ancillary) navigation data and a computing facility for on-board OD/TD processing. Also, the required ancillary data (time reference information, TDAS ephemerides, link calibrations, etc.) is generated by the ground segment and incorporated in the tracking signal.

2.2 RETURN LINK TRACKING (RLST)

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An overview of the major functions involved with this alternative is shown in Figure 2-3. During scheduled return link intervals, user S-band transmissions for tracking are relayed to the ground (at K_U -band) by the assigned TDAS satellite. Range and range-rate (doppler) data are measured at the ground and ancillary data (timing^{**}, position estimate, etc.) is decoded. The tracking data is processed to determine the user's orbit and clock parameters (e.g., bias, drift).

VI-2-3

^{*} Lower altitude coverage is governed by the zone of exclusion size for a given constellation (See [1d] for details).

^{**} This includes the user's current time and a sync word for range ambiguity resolution at the ground.



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To support users with on-board navigation data requirements, the ground also generates orbit and time prediction data for subsequent upload to the user during a scheduled forward link service interval. Such prediction data could comprise algorithm coefficients for Keplerian arc and/or polynominal curve fitting models [10,11] and the applicable time intervals. A user receives the ground-derived navigation data and computes orbit and time prediction between uploads. The interval between uploads is an important factor considered in the navigation performance evaluation (see Section 5).

A block diagram of the essential elements in the return link tracking configuration is shown in Figure 2-4. For the ground segment, the same types of equipment used for two-way tracking support and user orbit verification could also be used for RLST functions. A ground computing capability would also be necessary for user OD/TD and prediction data generation with adequate turnaround. The user requires a stable frequency standard, means for ancillary data generation and formatting, and a computing facility for navigation data interpretation and propagation between uploads. There would be no impact on the TDAS satellite configuration with this option.

2.3 TRACKING SIGNAL AND DATA HANDLING INTERFACES

The multiple beam antenna and switch enhancements for TDAS spacecraft provide the capability for simultaneous, direct transmissions between the space segment and several ground stations [1d,e]. This provides possibilities for <u>direct</u> control of user spacecraft by the mission control centers (MCC) instead of interfacing through the Network Control Center (NCC) and White Sands (WSN) terminal as in TDRSS (see Figure 1-1).

Figure 2-5 illustrates options for tracking signal and navigation data flow with each of the one-way alternatives. Since the beacon signal for FLBT is a general resource, it is assumed to originate at WSN, the assumed control point for TDAS spacecraft [le]. Navigation data computed on-board can be received by a TDAS ground terminal at the NCC directly and by MCCs

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with direct space/ground access. Additional interfacility transfer of data can occur to support MCCs without direct space/ground access or for general coordination and/or verification functions. The latter includes WSN which has control of TDAS spacecraft facilities (e.g., SA antenna/tele-scope pointing).

With FLST (see Figure 2-5b) the user tracking signal is imbedded in the normal uplink data communication traffic so it can emanate from either the NCC or a cognizant MCC. Navigation data computed on-board can be distributed in the same manner discussed above for FLBT.

With RLST (see Figure 2-5c) the user tracking signal is imbedded in the normal downlink data communication traffic, so it can be received by either the NCC or cognizant MCC. Ground processing for user OD/TD can occur at the Orbit Support Computer Facility (OSCF) or at the MCC with subsequent interfacility data transfer as noted above. Since the two-way tracking with ground-based processing is also a TDAS alternative, Figure 2 6 is included for comparison with the RLST data flow.



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SECTION 3

TRACKING SIGNAL CHARACTERISTICS

The one-way alternatives for user navigation are distinguished by tracking signal generation, transmission, and/or processing aspects. This section presents: a tracking signal definition for each case (including a candidate beacon signal); results of link performance analyses to assess EIRP and G/T tradeoffs; and estimates of the measurement accuracy for range and range-rate (doppler) tracking data.

3.1 FORWARD LINK TRACKING (FLBT/FLST)

3.1.1 Tracking Signal Definition

On-board navigation by user spacecraft, as defined, requires that the received forward link signals provide the means to make range and rangerate measurements. Table 3-1 details the possible signal characteristics for two different modes: Forward Link Beacon Tracking (FLBT) and Forward Link Scheduled Tracking (FLST). Users may perform doppler only tracking of the signals to obtain range-rate values, but the signals are modulated by pseudonoise (PN) codes to also allow estimation of range delay. A chip rate of 3.0778 Mcps is assumed here in keeping with current TDRSS practice, although higher chip rates are perhaps possible. The employed PN codes are unique to each user in the case of FLST and unique to each TDAS satellite in the case of FLBT. The signals are also modulated by data, providing the user with, as a minimum, the TDAS ephemeris and the timing information necessary for clock synchronization purposes. It is assumed that, in addition to onboard navigation capabilities, user spacecraft will also possess Viterbi decoders to permit convolutional coding of the forward link data and a consequent 5 dB coding gain. Current technology has already placed such decoders on VLSI chips.

Forward Link Beacon Tracking: The ground-generated beacon signal is conceived to be transmitted to user spacecraft by a single element of the 60-element TDAS SMA antenna array. Due to the large beamwidth of a single

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FORMARD LINK TRACKING SIGNAL CHARACTERISTICS

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SIGNAL CHARACTERISTICS	BEACON (FLBT)	SCHEDULED (FLST)
TYPE	PN MODULATED BPSK DATA	 STAGGERED QUADRIPHASE PN (SQP (COMMAND & RANGE CHANNELS PEF TDRSS)
CODE(S)	 UNIQUE 1023-CHIP GOLD CODE 3.0778 x 10⁶ CHIPS/SEC 2.6 NS GRANULARITY (VIA INTERPOLATION TO 1/128 CHIP 	 SHORT CODE (COMMAND CHANNEL/ 1023 CHIP GOLD CODE) LONG CODE (RANGE CHANNEL/ 255 x 1023 CHIPS USED TO RESOLVE RANGE AMBIGUITY 3.0778 x 10⁶ CHIPS/SEC
DATA	 INCLUDES ANCILLARY DATA (E.G., TDAS EPHEMERIS, TIME REFERENCE AND PN EPOCH) < 1000 BITS/ FRAME CONVOLUTIONALLY-CODED (RATE ½) DATA RATE ~ 100 BPS DATA RATE ~ 100 BPS ASSEMBLED INTO 1000 BIT FRAMES (10 SECONDS) SYNC WORD IN EACH FRAME RE- BOLVES AMBIGUITY 	 INCLUDES ANCILLARY DATA (SIMILAR TO FLBT) CONVOLUTIONALLY-CODED (RATE ⁵)
DOPPLER COMPENSATION	NONE (NOT FEASIBLE FOR MUL- TIPLE USER SUPPORT, BUT COOR- DINATION WITH FLST ENABLES RE- DUCTION OF DOPPLER UNCERTAINTY FOR INITIAL BEACON ACQUISITION	 PROVIDED FOR SIGNAL ACQUISI- TION

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element, the beacon is available to all users within the TDAS field of view $(\pm 13^{\circ})$. In supporting more than one user simultaneously, however, adjustment of the transmit frequency to compensate for doppler shifting between TDAS and the user is not feasible. The consequent search in frequency required for a user to acquire the beacon signal may be significantly reduced by orbit predictions from prior tracking intervals. Coordination with other modes of tracking would provide a basis for such knowledge. Beacon signal acquisition is discussed further in Section 3.1.3.

The beacon signal is assumed to comprise PN modulated BPSK data transmitted at a data rate of approximately 100 bps. Such a low data rate is consistent with: the EIRP provided from a single TDAS SMA antenna element, low gain user antennas, and system noise temperatures. The data would be assembled into frames of roughly 1000 bits, each frame containing, for example:

- TDAS ephemeris parameters
 - Keplerian arc parameters
 - ephemeris reference time
 - age of data
- TDAS timing and link calibration parameters
 - PN code epoch
 - polynomial correction coefficients
 - age of data
- Frame overhead
 - frame synchronization word
 - error control field

Miscellaneous data

- almanac parameters for other TDAS satellites
- health of all TDAS satellites.

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The data is assumed to be synchronous with the PN code, allowing bit synchronization of the data to resolve range ambiguity, e.g., [12,13].

Forward Link Scheduled Tracking: This mode is discussed based on data for navigation derived from scheduled forward link SMA transmissions. The signal structure is assumed to be the same as in TDRSS. Separate, simultaneous command and range channel signals are transmitted using staggered quadriphase shift keying (SQPSK) as the modulation technique. The command channel contains all data and is asynchronously modulated by a short (1023-chip) PN code, thus permitting rapid acquisition; the range channel is acquired separately and is modulated by a long (256 x 1023-chip) PN code to allow the resolution of range ambiguity. Since data is contained only in the command channel, unbalanced SQPSK modulation is employed, maintaining the ratio of power in the command channel to that in the range channel at 10 dB. The command and range channel PN codes are time-synchronized so that acquisition of the command channel's short code assists in acquisition of the range channel's long code. Acquisition of both the carrier frequency and PN codes is further aided by doppler compensation so that the carrier frequency received by the user is within a predictable tolerance.

3.1.2 Link Performance

Assessment of the error contribution in range and range-rate measurements due to thermal noise depends on the received C/N₀, the ratio of signal power to noise spectral density. Table 3-2 shows the TDAS SMA forward link budget used in this study. The TDAS satellites act as relays from the ground to the user spacecraft, comprising a two-link communication system. Since the weaker of the two links constrains the overall communication performance, only the satellite-to-user link is considered here. Reasonable values are assumed for the polarization and miscellaneous system losses; the path loss is calculated assuming a satellite-to-user distance of approximately 41,000 km. The ratio C/N₀ for the forward link is thus parameterized by the user spacecraft's antenna gain and system noise temperature (G/T) and the EIRP emitted by the TDAS satellite.



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The maximum data rate that can be supported at a given bit error rate is also a function of the received C/N₀. If a desired 10^{-5} bit error rate, a 5 dB gain obtained from convolutional coding, and a 3 dB margin are assumed, Figure 3-1 shows the achievable data rates as a function of user G/T, parameterized by several representative TDAS power levels. For the beacon case, decoding 100 bps ancillary data implies a user G/T requirement > -32 dB/°K. This is not a very stringent requirement considering that a 0 dB antenna gain and a 1000° K receiver system noise temperature is easily attainable.

3.1.3 <u>Beacon Signal Acquisition</u>

The requirements for user acquisition of the beacon signal are affected by its frequency uncertainty due to doppler shifting. As an example, the Motorola 2nd generation TDRSS user transponder design [14] specifies a minimum C/N_0 of 33 dB-Hz for acquisition in 20 seconds with a probability of .9, assuming a frequency uncertainty of 1.5 kHz and a two channel search in time and frequency. Figure 3-2 shows the user G/T and TDAS EIRP required to surpass this threshold. With a single-element EIRP of 25.9 dBW for the beacon signal, the user G/T must be greater than approximately -27 dB/°K to meet the conditions of the Motorola design. The uncertainty in frequency and time for the beacon will lead to longer acquisition times, but this may be acceptable in the case of beacon acquisition.

To ameliorate the signal acquisition task, future user transponder designs might use more than two channels to effect the PN code acquisition. Figure 3-3 shows a 3-channel configuration. All three channels could be used in acquiring the beacon with a resultant decrease in acquisition time. When scheduled service begins, one channel may remain dedicated to tracking the beacon while the remaining two channels acquire the doppler compensated forward link scheduled service. Use of the three channels could be directed by the user's on-board computer in various handover scenarios, as suggested in Figure 3-4. Furthermore, the acquisition search in time and frequency could take advantage of past tracking data to optimally conduct the search.



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3.1.4 Tracking Data Measurement Accuracy

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From the analyses of Appendix B, measurement for or both FLBT and FLST modes were estimated and are shown in Table 3-3. The thermal noise calculations assume that the user G/T equals -27 dB/°K as a baseline value; code and carrier loop parameters are assumed to be in keeping with current design practice. It can be seen that the range error is dominated by systematic delay errors encountered in the ground-to-TDAS link. These are assumed to be roughly the same as specified for TDRSS, with a total systematic delay of 10 meters conjectured if a routine and thorough system calibration capability is implemented [15].

Range-rate error for FLBT is dominated by thermal noise, reflecting the relatively low power of the beacon signal. Increased beacon power and, of course, higher user G/T values would diminish the error. For FLST, phase noise introduced by the ground frequency standard and the voltage-controlled oscillators (VCO's) on-board the TDAS satellite are major components of the total range-rate error. The estimated phase noise contribution is derived from consideration of possible TDAS mixing frequency schemes, as detailed in Appendix B. The alternative of a Ka-band ground-to-TDAS link versus the Ku-band link used in TDRSS introduces additional phase noise into range-rate measurements due to the larger frequency multiplies involved, thus dominating the total range-rate error. Use of a crystal oscillator disciplined to the ground cesium frequency standard would, however, improve the ground contribution to phase noise. To estimate the impact, the analysis of Appendix B.3 was applied using the crystal oscillator phase noise spectral density for frequencies less than 1000 Hz and that of the cesium oscillator for frequencies greater than 1000 Hz as an approximation to the possible ground phase noise spectral density. This leads to a ground contribution of roughly .3 mm/s with a Ku-band pilot tone and roughly .6 mm/s with a Ka-band pilot tone. In comparison to the pessimistic values cited in Table 3-3 (1.8 and 2.0, respectively), the improvement is significant.

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TABLE 3-3

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SUMMARY OF FORWARD LINK TRACKING DATA MEASUREMENT ERRORS

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	<u></u>		·····		
NOTES	• USER G/T =-27 dB/°K • PN TRACKING LOOP BL = 1.0 Nz	• ESTIMATES BASED ON TDRSS USER'S GUIDE (STDN 101.2 [2])	 I-WAY DOPPLER IRACKING K_u-band ground-tdas Link 		
ERROR (MM/SEC) FLST		ł	0.2 2.0 (4.4)** 1.8 (2.0)**	3.2 (5.8)** -	
RHS RANGE-RATE Flbt	5.5	I	0.2 2.0 (ŝ.4)** 1.8 (2.0)**	6.1 (7.3)** -	
ROR (HETERS) FLST	1.0	10.5 9.0 (10.0)*	I	1.0 13.8 [10.0)*	
RMS RANGE ERI Flbt	4.9	10.5 9.0 (10.0)*	1	4.9 13.8 (10.0)*	CTCHTTO DE
ERROR CONTRIBUTOR	THERMAL Noise	TDAS S/C SYSTEMATIC GND DELAY GND (ESTIMATED)	PIIASE USER PIIASE TDAS NOISE S/C (ESTIMATED) GND	RSS RANDOM SYSTENATIC	ECTIMATED TATAL SVI

WING CALIBRANION LIDJ. 2 ** WITH K_a-band ground-tdas link.



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TABLE 3-4

RETURN LINK TRACKING SIGNAL CHARACTERISTICS

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SCHEDULED (RLST)	STAGGERED QUADRIPHASE PN (SQPN) (Per TDRSS DG1-mode 2/noncoherent)	 2047 CHIP GOLD CODES ON I AND Q CHANNELS (OFFSET > 20,000 CHIPS) 3.0778 x 10⁶ CHIPS/SEC 	 INCLUDES USER PN CODE EPOCH, CURRENT TIME REFERENCE AND SYNCH WORD FOR AMBIGUITY RESOLUTION CONVOLUTIONALLY-CODED (RATE 1/2)
SIGNAL CHARACTERISTICS	ТҮРЕ	CODE	DATA*

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* Modulo-2 Added Asynchronously to both PN Codes

3.2 RETURN LINK TRACKING (RLST)

3.2.1 <u>Tracking Signal Definition</u>

As a one-way navigation option, RLST involves the ground tracking of noncoherent transmissions by user spacecraft using the S-band multiple access return service. The signal structure is assumed to be consistent with TDRSS DG1 Mode 2 [2] and is shown in Table 3-3. As with forward link tracking, an ultra-stable oscillator is required on-board the user spacecraft to provide an accurate time and frequency standard. The user-unique 2047-chip Gold codes on the I and Q channels are not long enough to resolve range ambiguity; the data must therefore be synchronously modulated by the PN codes to allow bit and frame synchronization at the ground receiver to eliminate the ambiguity. To support one-way navigation, the data must contain the user's PN code epoch and timing information.

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3.2.2 Link Performance

In order to evaluate the measurement errors in Return Link Scheduled Tracking due to thermal noise, the link budget of Figure 3-5 was used to caluclate C/N₀. As in the calculations for forward link tracking, the stronger TDAS-to-ground link is neglected here. The budget assumes a G/T figure of -14.1 dB/°K per element of the 60-element TDAS SMA antenna array with the 60 elements providing a theoretical combiner gain of 17.7 dB. The received C/N₀ at the TDAS satellite (approximately equal to that received at the ground) is thus a function of the user EIRP. If a 10⁻⁵ bit error rate, a 5 dB convolutional coding gain and a 3 dB margin are assumed, Figure 3-5 shows the return link achievable data rate plotted against user EIRP. For example, a 1000 kbps user would require an EIRP \approx 0 dBW.

3.2.3 Tracking Data Measurement Accuracy

Table 3-5 summarizes the RLST measurement errors estimated using the techniques of Appendix B. Salient performance characteristics of the Harris wide dynamics demodulator were used to model the ground receiver's response

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SUMMARY OF RETURN LINK SCHEDULED TRACKING MEASUREMENT ERRORS

	_ ∝			
NOTES	 USER EIRP = 2 dBW WIDE DYNAMICS DEMODULATO 	• ESTIMATES BASED ON TDRSS USER'S GUIDE (STDN 101.2 [2])	 I-WAY DOPPLER TRACKING K_u-band ground-tdas link 	
E ERROR)			(4.6) ^{**} (8.6) ^{**}	(10.4)**
RMS RANGE-RAT (MM/SEC	3.5	I	.2 2.8 5.7	7.3 -
ERROR S)		*(0.01)		(10.0)*
RMS RANGE (Meter:	2.5	10.5 9.0	ł	2.5 13.8
		S/C GND	USER Tdas S/C GND	RANDOM Stematic
ERROR Contributor	THERMAL NOISE	TDAS SYSTEMATIC DELAY (ESTIMATED)	PHASE NOISE (ESTIMATED)	RSS SY

* ESTIMATED TOTAL SYSTEMATIC DELAY WITH CALIBRATION [15]

** WITH K_a-band tdas s/c-ground link.

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to thermal noise. As a baseline value, the user spacecraft is assumed to transmit an EIRP of 2 dBW, thus yielding a nominal C/N_0 of approximately 40.4 dBW at the demodulator. Systematic errors in range and range-rate measurements are estimated as for the forward link tracking modes. Again, TDAS systematic delays dominate the range error value.

Range-rate error is seen to depend most on phase noise introduced by the ground frequency standard. This result, however, pessimistically disregards the possible use in TDAS of a phase-locked crystal oscillator disciplined to the cesium frequency standard. If such a method were employed, the phase noise spectral density of the ground standard would be improved, reducing its phase noise contribution to range-rate errors. Calculations using the phase noise spectral density mentioned in Section 3.1.4 yield a ground contribution of .9 mm/s for a system using a K_u -band pilot tone and 1.5 mm/s with a K_a -band pilot tone. This compares to the quoted results of 5.7 mm/s and 8.6 mm/s, respectively. With this approximation, thermal noise becomes the dominate range-rate error source for RLST.

SECTION 4

TDAS TRACKING ANALYSIS

Accurate knowledge of TDAS satellite ephemerides is necessary for user navigation support and TDAS orbit maintenance functions. This section presents a preliminary analysis of TDAS satellite tracking to:

- Identify potential system configuration and operational impacts of TDAS constellation alternatives, and
- Estimate potential tracking accuracy based on BRTS and VLBI techniques.

An overview of the elements considered in pursuing this analysis is given . in Figure 4-1.

4.1 TRACKING CONFIGURATIONS

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Several constellation/network options identified in the TDAS study [1d,e], are illustrated in Figure 4-2. Satellite spacings as well as frontside/ backside deployments were chosen to assess possible coverage improvements and data distribution alternatives for servicing multiple earth stations in CONUS. Distributed, direct access links to/from CONUS are achieved by satellite crosslinks and/or multi-beam antenna enhancements on the TDAS satellites.

In terms of TDAS tracking, however, the various constellation options can be reduced to two situations: frontside and backside satellites. The following subsections discuss BRTS and VLBI tracking configurations to support either case.



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4.1.1 BRTS Tracking

Analogous to the TDRSS implementation, two-way range and range-rate tracking data (R, \dot{R}) are assumed to be acquired with Bilateration Ranging Transponders (BRT) operating at pairs of automated ground stations (one pair per TDAS). Tracking transmissions originate at White Sands (WSN) with replys from a co-located BRT and one at a geographically displaced site. A typical data rate would be 1-5 observations per minute for 5 minutes every hour, as expected for TDRSS. [8].

Figure 4-3 lists some geometrically compatible station pairs^{*} that cover four of the TDAS locations considered. (see also Figure 4-4) Two satellites are at TDRS locations and a third (at 100°W) is in full view of CONUS. The fourth is a backside satellite (at 98°E) which requires a satellitesatellite crosslink in the BRTS configuration if WHS is the base location for originating transmissions. This impacts tracking accuracy, since measurements are also affected by uncertainties in the frontside satellite orbit.

The last two BRTS configurations suggested for backside TDAS tracking in Figure 4-3 employ <u>two</u> backside stations to offset the crosslink impact. In the first approach, one station is configured to emulate WHS for tracking data generation and measurement, but this would add a significant overseas hardware and maintenance requirements. In the second approach, each two-way measurement made between WHS and a backside station pair (S'_2, S'_2) is differenced with a corresponding measurement (M_T) made between WHS and the backside TDAS. The effect is to cancel systematic uncertainties in the crosslink and frontside space/ground link. However, this also increases the noise error due to the additional measurement (M_T) , although increasing the data rate sufficiently may permit additional data smoothing to offset this.

* Political issues regarding potential overseas sites were not considered.

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4.1.2 VLBI Tracking

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The VLBI technique is under investigation as a possible BRTS enhancement [9] with potential for high accuracy TDRS tracking. The fundamental VLBI measurement data type is equivalent to the difference in range (ΔR) between a signal source and two receivers displaced along a known baseline. For a typical measurement, wideband RF signals are received at two sites, mixed down to baseband and recorded over a prescribed interval. Signal phase coherence and time synchronization are maintained at each site with an ultra-stable frequency standard. Subsequent crosscorrelation of the two signal streams determines the relative difference in time of arrival which equals (after converting to distance) the geometric range difference plus measurement errors (random and systematic). Investigators at GSFC using signals sampled from celestial radio sources by several VLBI stations report capabilities for ΔR measurement precision in the 0.3-3 cm range. [20]

For the TDAS application, tracking signals assumed to originate at WHS are relayed by the satellite constellation to automated VLBI ground stations. Received signals are time tagged and buffered and, on command, returned via TDAS to WHS for processing. A representative measurement data rate would be one processed pair of VLBI observations (ΔR) every hour per station set. Nominally, one set of three stations with adequate baseline geometry would be needed per TDAS satellite.

Figure 4-5 lists some geometrically-compatible 3 station sets that cover the 4 TDAS satellites considered (see also Figure 4-4). The station sets are grouped in terms of long and moderate baselines, the latter of interest in a CONUS-based VLBI network for frontside TDAS tracking. For the backside satellite, no significant impact on measurement accuracy due to the crosslink would be expected, since the equivalent time-of-arrival differencing in the signal processing tends to cancel common path uncertainties.



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Relative errors between stations (e.g., clock synchronization and baseline uncertainties) are a general concern, since they can have a significant impact on TDAS OD accuracy. A possible alternative to relieve stringent calibration requirements at each station is to periodically include VLBI differences, i.e., Δ VLBI measurements in the tracking process. For this, additional VLBI observations derived from a known celestial radio source are subtracted from TDAS VLBI observations made for the same station pair. The effect is to cancel station time synchronization errors and reduce effects of station survey errors on TDAS OD accuracy.

Of course, the $\Delta VLBI$ alternative would introduce additional equipment complexity at each station to also receive celestial source signals. Data handling requirements would also be increased. Equipment complexity and operational aspects are important issues especially for backside VLBI tracking since 3 remote stations per TDAS are involved compared to 1 (or perhaps 2) for BRTS.

4.2 TDAS TRACKING ACCURACY

Information about TDAS orbits is assumed to be provided for user OD/TD on a recurring basis. User navigation performance is then a function of TDAS orbit uncertainties in the interval between updates (i.e., prediction interval errors). To assess potential orbit prediction errors, an accuracy analysis was made of TDAS tracking with the BRTS and VLBI techniques. The following subsections discuss the error modelling approach and the major results. Additional details are given in Appendix C.

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4.2.1 <u>Error Modelling</u>

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In this study it was assumed that TDAS OD is performed for a prescribed epoch via batch processing of tracking observations taken over a given data arc (or tracking interval). The epoch solution propagated beyond the end of tracking (EOT) is the satellite orbit state in the prediction interval. Orbit prediction error covariances were evaluated using the ORAN program,

a linear covariance analysis tool [16, 17]. ORAN was configured to compute satellite orbit uncertainties versus time from epoch, given the following input data:

- Initial State Vector All TDAS were assumed to be in 5°, circular, geosynchronous orbits with epoch location at the nominal longitude indicated in Figures 4-3 and 4-5.
- <u>Tracking Station Locations</u> Stations pertaining to a given TDAS and tracking technique are identified in Figures 4-3, 4 and 5 with specific locations as tabulated in Appendix C.4.
- <u>Tracking Schedule</u> BRTS observations were assumed to occur at 1/min for 5 mins/hr and VLBI observations at 1/min to 1/hr*.
- <u>Tracking Error Model</u> Values are stated in Table 4-1. All local and dynamic errors were treated as systematic consider error sources in the analysis.

BRT5 measurement errors are consistent with values used for TDRS OD analyses [18] assuming comparable equipment and link quality. Station errors were improved by 3:1 assuming better site survey capabilities in the TDAS time-frame.

VLBI measurement errors were considered at two levels: a <u>baseline</u> model consistent with values used in studies of a Deep Space Network (DSN) tracking application to TDRSS [19] and a <u>reduced</u> model which reflects achieved capabilities with VLBI stations observing celestial radio sources [20]. Station errors were taken to be the same as BRTS for the <u>baseline</u> model and an order of magnitude better for the reduced model.

" OD performance was evaluated as a function of VLBI data rate. (See Appendix C.3).

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TDAS TRACKING ERROR MODEL ASSUMPTIONS

$\begin{array}{c c c c c c c c c c c c c c c c c c c $	STATIONS 3 M 3 M (X,Y,Z) 3 M	LBI (AR) BIAS 10 M/0 CM/S 3 M	BUREMENT NOISE 2 M/.13 CM/S 0.9 M	(BASELINE MODEL)	ERROR SOURCE BRTS	Ia VALUES
ULAL TROPOSPHERIC . IS REFRACTION .		STATIONS 3 M 3 M 3	LBI (AR) BIAS 10 m/0 cm/s 3 STATIONS 3 m 3 m 3	UREMENTNOISE2 M/.13 CM/S0.9RTS (R/ \mathring{R})BIAS10 M/0 CM/S3LBI (ΔR)BIAS10 M/0 CM/S3STATIONS3 M3 M3	UREMENT NOISE 2 M/.13 CM/S (BASELINE RTS (R/Ř) BIAS 2 M/.13 CM/S 0.9 LBI (AR) BIAS 10 M/0 CM/S 3 STATIONS 3 M 3 M 3 M 3 M	ERROR SOURCEBRTS(BASELINIUREMENTNOISE2 M/.I3 CM/S0.9RTS (R/Ř)NOISE2 M/.I3 CM/S0.9LBI (AR)BIAS10 M/0 CM/S3LBI (AR)STATIONS3 M3STATIONS3 M3 M3

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The tropospheric and solar pressure errors reflect modelling inaccuracies and are standard values used in TDRS OD studies [18,19], although some improvement may be possible. The gravitational errors are conservative since further improvements are likely by the TDAS timeframe.

4.2.2 Error Analysis Results

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TDRS studies indicate that the choice of tracking interval and reference epoch can significantly impact both definitive and predictive OD accuracy. A recent study [18] of TDRS tracking via BRTS evaluated orbit position/ velocity uncertainties over a 24 hour prediction interval* as a function of each parameter. Results show that a judicious combination can be selected in which the maximum uncertainty in the prediction interval is at or near a minimum.

In the TDAS study it was decided to also consider The prediction interval as a <u>selectable</u> parameter. The basic rationale is that a sequence of shorter prediction intervals with better accuracy may be traded off against more frequent OD processing for user navigation data updating. In both BRTS and VLBI analyses efforts were made to identify multiple combinations (epoch/tracking interval/prediction interval) which, taken together, yield better accuracy over a 24 hour period than any one alone.

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4.2.2.1 <u>BRTS Tracking</u>. TDAS position uncertainty was computed over a 24 hour prediction interval as a function of tracking interval and epoch for both frontside (100°W) and backside (98°E) cases. Figure (4-6) presents results plotted from the end of tracking (EOT) for several tracking intervals with a common epoch (satellite local noon) assumed in each case.^{**}

^{*} The 24 hour interval was considered since TDRS OD is planned to be performed only once per day.

^{**} Epoch times are stated as Greenwich Mean Time. The tracking configurations are based on Figure 4-3: frontside tracking - Configurations I and II; backside tracking (Option 1) - Configurations III and IV (T100 means TDAS @ 100°W is the relay to/from the backside TDAS @ 98°E).

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In the frontside results shown in Figure (4-6a), the peak uncertainty is due primarily to solar pressure modelling error, the leading error contributor. Consequently, the peak value turns out to be an oscillatory function of epoch with a 12 hour cycle over epoch time of day. As observed in Appendix C however, even the lowest peak values exceed, by at least 30%, the position uncertainty for the 18 hour tracking case in the 12 hour period after EOT (see Figure 4-6a). Also, because of the 12 hour cyclic property, approximately the same error profile occurs with epoch placed at local midnight. Thus, orbit predictions from two 18 hour tracking intervals with epochs at noon and midnight could be concatenated to produce a comparatively flat prediction error profile over 24 hours. Comparison of the maximum prediction uncertainties and primary contributors for continuous and concatenated cases is shown in Table 4-2.

Fc backside case shown in Figure 4-6b, the results are also dominated by solar pressure effects. However, they also reflect the frontside TDAS orbit uncertainty which was assumed in this analysis to be constant (at the 100 m level) corresponding to the concatenated 12 hour prediction results discussed above. No significant benefit is apparent from using a similar approach in this case based on 36 hour tracking, since the peak uncertainty over 24 hours is nearly the same as over 12 hours. A comparison of the peak values and primary error contributors can be made from the data in Table 4-2. Corresponding peak TDAS velocity errors are in the 8 mm/sec and 19 mm/sec range for the frontside and backside tracking cases, respectively.

The preliminary observation here is that the orbit prediction uncertainty over 24 hours for the backside TDAS (with tracking configuration Option 1^*) is apparently more than double that for the frontside TDAS. Certainly further analysis is warranted however, to evaluate possible improvements that alternative configurations and/or processing approaches may yield, e.g.,

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^{*} Tracking configuration options are defined in Figure 4-3.

TABLE 4-2

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MAXIMUM TDAS POSITION ERROR IN 24 HOUR PREDICTION INTERVAL WITH BRTS TRACKING

TDAS	TRACKING INTERVAL (HOURS)	PREDICTION INTERVAL(S) (HOURS)	MAXIMUM POSITION ERROR (M)	PRIMARY ERROR Contributors* (M)
FRONTSIDE	18	24	• 1/5	150 (sP) 80 (B) 40 (s)
100°W	18	12 + 12 **	110	80 (B) 60 (SP) 45 (S)
BACKSIDE	36	24	290	230 (s') 175 (sP) 25 (B)
98°E	36	12 + 12 **	280	220 (s') 165 (sP) 25 (B)
* LEGEND: S	P - SOLAR PRESSURI - STATION SURVEY	E; B - BRT BIAS; S - Y + T100 LOCATION (T	STATION SURVEY; Das 100°W).	

** CONCATENATED INTERVALS BASED ON TRACKING INTERVAL EPOCHS SPACED ~12 HOURS APART.

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- Backside tracking configuration Options 2 or 3* which utilize another backside station.
- Joint data processing for frontside and backside TDAS OD that accounts for mutual error source correlations that an independent solution (assumed here) does not.

4.2.2.2 <u>VLBI Tracking</u>. TDAS position uncertainty was computed over a 24 hour prediction interval as a function of the VLBI tracking interval and TDAS location for two error models defined earlier in Table 4-1. One is a conservative (baseline) error model and the other is a more optimistic (reduced) error model.

Figure 4-7 presents results for the baseline model plotted from the end of tracking (EOT) with a common epoch (satellite local noon) assumed in each case. The results in Figure 4-7b indicate that VLBI tracking performance is relatively insensitive to TDAS location, frontside or backside.

As shown in Appendix C, solar pressure modelling error is a major contributor for the longer tracking intervals (\geq 18 hours). Since this effect is also cyclic with epoch time of day, essentially the same error profiles result with epoch selected at local midnight. Thus, if the prediction interval segments for 18 and 36 hour tracking are concatenated appropriately, a significantly lower prediction uncertainty over 24 hours could be achieved than with any individual case alone.**

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With the shorter tracking and prediction intervals (≤ 12 hours), measurement and local error sources (noise, bias and station survey) are dominant elements. As illustrated in Figure 4-7a however, the prediction uncertainty increases significantly as the tracking interval decreases. While a slight improvement in prediction performance can be achieved by also estimating

^{*} Tracking configuration options are defined in Figure 4-3.

^{**}Results with VLBI bias estimation are given in Appendix C.3. These show that prediction segments for 18 and 12 hour tracking segments would give somewhat better accuracy that the 18 and 36 hour combination.

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the VLBI measurement bias, no really significant benefit is apparent from using shorter tracking and prediction intervals, when VLBI measurement and local errors are at the levels assumed in the baseline model. However, if a VLBI implementation can be realized with these errors at the level assumed in the reduced model, a different situation emerges.

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Figure 4-8 compares TDAS position uncertainties for short tracking and prediction intervals for both error models. In the latter case, the lowest position uncertainty over a given prediction interval occurs with 6 hour tracking. As observed in Appendix C, this is essentially the cross-over region between tropospheric and solar pressure effects with the latter becoming significant for longer tracking/prediction intervals. Thus, a sequence of short prediction intervals (e.g., 1 hour) could provide a position uncertainty in the range of 10 - 15m.

Figure 4-9 shows examples of TDAS position uncertainty from using multiple (concatenated) prediction segments over a 24 hour period. In the baseline case (Figure 4-9a)* the uncertainty ranges from 50 - 80m using 4 segments based on 18 and 36 hour tracking or 50 - 70m based on 18 and 12 hour tracking with VLBI measurement bias estimated. In the reduced case (Figure 4-9b)* the error ranges from 10 - 12m using 24 one hour segments based on 6 hour tracking intervals with epochs spaced one hour apart. A second curve based on 5% tropospheric error shows the improvement (to 5 - 8m over 24 hours) if better tropospheric error correction can be achieved. A comparison of the peak errors and primary error contributors in each case is given in Table 4-3. Maximum TDAS velocity errors corresponding to these position error results are at the 6 mm/sec and 1 mm/sec levels for the baseline and reduced error models, respectively.

^{*} The assumed VLBI data rate of 1/hr (one ΔR pair/hr) is more representative that the 1/min rate assumed in Figures 4-7 and 4-8. Nevertheless, the computed results for these tracking intervals are identical with either rate. Tables C-1 and C-2 in Appendix C show the impact on prediction uncertainty for other tracking intervals.



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TABLE 4-3

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MAXIMUM TDAS POSITION ERROR IN 24 HOUR PREDICTION INTERVAL WITH VLBI TRACKING

VLBI VLBI	TRACKING	PREDIC	CTION Val	MOLTION POSITION	PRIMARY Error
HODEL	IN IEKVALS (HOURS)	NO. Seghents	LENGTH (HOURS)	ERROR* (H)	CONTRIBUTORS**
BASELINE	18/36	4	6/3	87	72 (SP) 45 (VB) 25 (S)
BASELINE (NITH BIAS ESTIMATED)	18/12	ч	6/6	72	56 (SP) 45 (VN) 20 (S)
REDUCED	9	24	1	12	10 (TR) 6 (SP) 4 (S)
REDUCED N. 5% TROPO)	Q	24	H	œ	6 (SP) 4 (S) 3 (TR)

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SP – SOLAR PRESSURE S – STATION SURVEY VB – VLBI MEAS. BIAS VN – VLBI MEAS. NOISE TR – TROPOSHERE LEGEND:

TDAS @ 100°W; STNS: WHS-AGO, WHS-HAW; DATA RATE = 1/HR

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The foregoing results illustrate the significant improvement in TDAS OD accuracy that the VLBI technique could potentially provide. Moreover, this would apply to backside TDAS tracking as well, since crosslink range uncertainties are inherently cancelled out. It must be emphasized however, that achieving a 10m capability implies having measurement and station survey uncertainties at a level consistent with the reduced error model and using appropriate batch tracking/prediction interval combinations (or equivalent scheme) for OD processing. Thus, important areas for further study include:

- VLBI measurement precision (noise) with TDAS-based signals,
- Alternatives for VLBI tracking network calibration (time synchronization and baselines)
 - via **AVLBI** measurements (see discussion in Section 4.1.2) or
 - via terrestrial microwave links^{*} and a geodetic survey receiver (e.g., GPS [21]),
- OD processing options
 - Algorithm (batch vs sequential)
 - Update rate vs prediction interval.

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^{*} This is applicable to CONUS-based tracking configurations (e.g., a connected element interferometer network) currently under study for TDRS [9]. Some preliminary results for TDAS are given in Appendix C.2.3.

SECTION 5

USER NAVIGATION PERFERMANCE EVALUATION

User navigation via TDAS is based on computing orbits and time from range and range-rate measurements and known TDAS orbits. This section presents an evaluation of potential navigation performance in terms of OD/TD accuracy for the three one-way tracking alternatives defined in Section 2.

5.1 OVERVIEW OF APPROACH

Figure 5-1 gives an overview of the elements considered in the analysis. Cases for evaluation were defined to assess several types of user orbits and the impact of TDAS constellation options, different tracking schedules and two algorithms for tracking data processing (sequential and sliding batch).

To evaluate user OD/TD accuracy two error analysis programs were employed:

- <u>SEA Program [22]</u> for tracking based on sequential processing of measurement data, and
- <u>RDGTUS [23]</u> for tracking based on batch processing of measurement data.

Given a tracking schedule and nominal TDAS and user orbits each program computes OD/TD error covariances and sensitivities versus time with respect to measurement noise and other applicable error sources. The following subsections discuss the tracking configurations and methodology, the error modelling approach and the major results. Further detailed results are contained in Appendices D and E.

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5.2 TRACKING CONFIGURATIONS AND METHODOLOGY

TDAS constellation and user orbit types considered in the performance evaluation are defined next. Then, tracking data scheduling and processing options assumed for the three tracking alternatives (\$2.3T, FLST and RLST) are discussed.

5.2.1 TDAS Constellations

The various constellation/network options considered in the TDAS study were discussed in Section 4. For purposes of this evaluation the three constellations shown in Figure 5-2 were selected. Option 1 is analogous to TDRSS with two satellites spaced 130° apart which provides 85-100% coverage at altitudes down to 200 km.* Option 2 also uses two satellites, but with the maximum allowable spacing, 162°**, which yields 98-100% coverage. Option 3 has three satellites, two deployed as in Option 1 and a third on the backside, which together provide 100% coverage.

5.2.2 User Orbits

Six orbits were selected to compare navigation performance for various "orbit altitudes as shown in Figure 5-3. As is evident from Appendix A, (Table A-1), they are also indicative of the majority of potential user orbits in the TDAS mission model.

The low altitude (high drag) orbit types are of interest to determine whether more frequent tracking data, available with FLBT, is of significant benefit. The high and low inclination orbit types are of interest, since their coverage and geometrical properties can differ significantly. In the high inclination case, the % coverage is generally greater, but opportunities for (good) doppler tracking can be reduced. This occurs whenever the user orbit normal points sufficiently toward a TDAS, thereby

^{*} A detailed discussion of coverage characteristics as a function of satellite spacing is given in [ld].

^{**} This is to avoid earth occulation of the TDAS-TDAS crosslink.

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A	28°	200 KN
В	41	600
C	lt	1000
D	97.°	200
Ε	98°	600
F	4	1000

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reducing the range change over a pass. In the limit, there is no range change (zero doppler), yet full visibility.

5.2.3 <u>Tracking Data Processing and Scheduling Methodology</u>

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One-way range and range-rate (R,R) are the fundamental data types employed for user OD/TD with each tracking alternative. System configurations and technical considerations for deriving these data were covered in Sections 2 and 3. This discussion pertains to measurement scheduling and processing aspects that were assumed for the performance evaluation.

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Two algorithms for tracking data processing were considered: sequential and sliding batch. In the <u>sequential</u> case, estimated parameters (e.g., position & velocity; clock bias & drift) are updated after each measurement and propagated forward between measurements and/or tracking passes. For this analysis an extended Kalman filter was assumed to be employed. In the <u>sliding batch</u> case, selected parameters are estimated for a prescribed epoch via standard weighted least squares processing of tracking observations taken over a given data span (T). Current estimates are derived by propagating the epoch solution over a specified interval (P) beyond the end of tracking and whatever computation/data handling interval (C) is required. In other words, current estimates are predictions in the interval, C to C+P, beyond the end of tracking. For each successive batch the process is repeated with the epoch and tracking interval advanced by P.*

Table 5-1 lists the tracking algorithm and measurement scheduling options considered in the performance evaluation for each alternative. Only sequential processing was assumed for FLBT, since the sliding batch approach was felt to be inappropriate for on-board OD/TD in view of the tracking data volume and associated computation/data handling requirements. For FLST and RLST, two tracking schedule types were considered based on acquiring data every other orbit (Schedule I) and every orbit (Schedule II).

^{*} Since P>C, predictions for all batches extend at least 2C beyond the data span. In a sense, 2C is a measure of estimate "staleness".

TABLE 5-1: TRACKING ALGORITHM & MEASUREMENT SCHEDULING OPTIONS CONSIDERED

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	DAT	R, F				>	-
	DATA SPAN	I	24 6-24 ноиrs		6-24 HOURS		
	BEACON	CONTINUOUS FOR SELECTED TDAS	SELECTED TDAS		1		
		1	ONE 10 MIN. PASS PER TDAS EVERY ORBIT				CUOLT A TRAS
AGT	I KA	1	10 MIN. PASS Per TDAS Every other orbit				E RATE THDON
	DATA PROCESSING ALGORITHM	SEQUENTIAL	SEQUENTIAL	SL IDING BATCH	SEQUENTIAL	SL I DI NG BATCH	REMENTS AT 3/MINIT
	TRACKING ALTERNATIVE FORMARD LINK BEACON TRACKING (FLBT)		FORWARD	SCHEDULED TRACKING (FLST)	RETURN	SCHEDULED TRACKING (RLST)	(R,R) MEASUI

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《《新闻》》。 "我们就是我们是不是一个的时候,你们们不是一个人,不是一个人,这个人,不是一个人,我们是不是一个人,我们是不是一个人,我们们就是一个人,这些人,也能能让

For sliding batch uses, the tracking data span was considered as a parameter to be selected with possible values ranging from 6-24 hours.

5.3 TRACKING ERROR MODELING AND COMPUTATION

In this study user navigation was evaluated from OD/TD error covariance computations using the SEA and RDGTDS programs [22, 23]. Figure 5-4 illustrates some hypothetical error profiles which indicate the peak errors and time intervals of interest for the options considered:

- <u>Sequential Processing</u> For FLBT and FLST, user OD/TD accuracy was defined as the peak error occuring over 24 hours with tracking intervals scheduled according to Table 5-1. For RLST, this criterion also applies to ground-based operations, but not necessarily for on-board navigation. Since navigation data updates would be available only on a recurring basis, user OD/TD accuracy was defined instead as the peak error in a prediction-only mode between uploads.
- Sliding Batch Processing For FLST and RLST, user OD/TD accuracy was defined as the peak error occurring over 24 hours in a series of batch-derived prediction intervals (P), see Figure 5-4. This is defined by the number of batches/day for FLST and by the navigation data upload rate for RLST. In either case, the minimum interval for P is governed by the OD/TD computation/data handling time (C) from the end of any tracking span.*

To obtain user OD/TD uncertainties as a function of various parameters the SEA program was utilized for sequential processing cases and the RDGTDS program for batch processing. Error covariances and budgets were computed given the following input data (see Figure 5-5):

^{*} In Figure 5-4, C=10 minutes (for RLST) is based on anticipated groundsystem processing capabilities [4]; C=90 minutes (for FLST) is based on LSI-11 type on-board computing capabilities planned for evaluation in a TDRSS forward link doppler tracking experiment [7].



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- <u>User Location</u> A user was assumed to be in one of the orbits (A-F) indicated in Figure 5-2 with an initial location at O°N, O°E for the orbit selected.
- <u>TDAS Location</u> All TDAS were assumed to be in 5° circular, geosynchronous orbits with initial location at the nominal longitude indicated in Figure 5-1 for the selected constellation option (1, 2, or 3).
- <u>Tracking Schedule</u> (R,R) measurement times were set up to occur at the rate defined in Table 5-1.
- <u>Tracking Model Parameters</u> Values are stated in Table 5-2

TDAS users were assumed to estimate, as a minimum, 8 basic parameters: 3 position and 3 velocity states and two clock states, bias and drift. Unestimated parameters were treated as systematic (consider) error sources in the analysis. Since the analysis is linear, the results for any particular error source may be scaled up or down to note the impact of different a priori uncertainties.

Errors due to atmospheric drag were modeled with two interpretations of uncertainty in the drag coefficient (C_D). For the highest altitude orbits C_D was assumed not to be estimated, so 25% of the nominal value was taken as the consider error, a typical choice [24,33,34]. At altitudes where effects may be significant (e.g., orbits A,B,D,E) C_D was assumed to be estimated, but imperfectly. The residual component was treated as a consider error with a 1 σ value equal to 2.5% of the nominal C_D . In other words reducing the a priori uncertainty by estimating C_D was assumed only 90% effective.

Values for uncertainties in the other dynamic parameters (GM, gravitational harmonics and solar pressure coefficient) are undoubtedly conservative for the TDAS time frame. Sequential processing results are based on the GEM9 error model [35], resident in the SEA program. Sliding batch results were

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derived with the GEM9-GEM7 error model, the closest option resident in RDGTDS.

Constant values assumed for TDAS orbit uncertainty correspond to 1g errors of approximately 50 m in position and 5 mm/sec in velocity. Results given in Section 4.2 indicate that this accuracy level would still be fairly conservative based on VLBI tracking for TDAS. It would be optimistic by at least 2:1 for BRTS tracking*.

Errors in the tracking measurements (R,\tilde{R}) were defined in terms of random errors and system biases. Values used for the random errors and the range bias error are representative of the results given in Section 3 (see Table 3-3). The 1 mm/sec for range-rate bias was included only to observe potential sensitivity.

Frequency drift in the user's reference oscillator appears as a doppler rate error which affects range-rate measurement accuracy and as a clock bias acceleration error (\ddot{B}) which affects ranging accuracy. Oscillator drift was defined as a consider error with a 1 σ value of 10⁻¹⁰ parts/ day, a level consistent with currently existing quartz oscillator technology.[32]

Other modelling parameters involved in the analysis pertain to tracking data processor tuning for achieving "best" performance in some sense. In the <u>sequential</u> case, state noise parameters are used as an artifice for adjusting the Kalman filter gains to control the weight given to prior estimates. The objective is to achieve a balance between uncertainties introduced by new measurements and those caused by propagating prior estimates with an imperfect dynamical model. The values shown in Table 5-2 were found to give good results in various test runs^{**} and were adopted

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^{*} This assumes the minimal BRTS tracking configuration using two stations per TDAS (see Figure 4-3).

^{**} As implemented in the SEA program [22], state noise has no connection with physical disturbance phenomena acting on the user spacecraft or clock, although in theory, there could be.

as a baseline set for most cases considered. In the <u>sliding batch</u> case, the tracking interval choice can be made with the same objective in mind. Four durations given in Table 5-2 were considered in each case and the one yielding the lowest peak error in a 6 hour prediction interval from end of tracking was selected.

5.4 NAVIGATION PERFORMANCE RESULTS

TDAS user navigation performance was evaluated in terms of potential OD/TD accuracy for each of the one-way tracking alternatives. Results were obtained as a function of user orbit type, TDAS constellation option, tracking schedule and two processing algorithms. This section presents a summary and interpretation of the significant data. More detailed information can be found in Appendices D and E.

5.4.1 Evaluation Cases

Potential cases for evaluation were defined by the six user orbit types in Figure 5-2 and three constellation options in Figure 5-3. For each case, five subcases were defined based on processor type (sequential or sliding batch) and tracking schedule (I, II or beacon).* Altogether, 90 combinations were identified and 48 of these were evaluated as indicated in Table 5-3.

For each case, user position and time error profiles were computed for a 36 hour period (the last 12 hours in predict-only mode). From these the peak errors for assessing performance were determined as defined in Section 5.3 and Figure 5.4. Appendix D presents error profiles and corresponding tracking schedules for sequential processing. Appendix E has corresponding results for sliding batch processing. Results for all cases were based on the nominal model parameters given in Table 5-2 except for certain processor tuning parameters noted specifically on error profile plots in Appendices D and E.

Recall Table 5-1.

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TABLE 5-3: CASES CONSIDERED FOR NAVIGATION PERFORMANCE EVALUATION*

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USER		IDAS CONSTELLATI	NO
ORBIT	TWO FRONTSIDE (171°U/41°)	2 1 FRONT/1 BACK \$100°W/98°E)	3 2 FRONT/1 BACK (171°W/41°W/74°E)
28°, 200 KH	S(3)	S(3), B(2)	S(3)
в 28°, 600 кн			
С 28°, 1000 кн			
D 97°, 200 KH	S(3), B(2)	S(3), B(2)	S(3)
Е 98°, 600 кн	s(3), B(2)	S(3), B(2)	Stat
F 98°, 2000 KH	s(3)	S(3), B(2)	5 (3)
NO. CASES	30	30	30
* S() OR B() IND BASED ON SEQUEN	ICATES NO, OF C VTIAL (S) OR SLI	ASES COMPLETED DING BATCH (B) P	WITH TRACKING ROCESSING.

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5.4.2 <u>Sequential Data Processing Results</u>

Navigation performance data based on the error analysis results in Appendix D are presented in terms of user position and time accuracies as summarized in Figures 5-6 and 5-8. Corresponding position and time error budgets are given in Figures 5-7 and 5-9.

5.4.2.1 <u>User Position Accuracy</u>. Figure 5-6 illustrates that in low altitude orbits where drag is a factor, performance depends heavily on frequent data availability. Beacon tracking is clearly superior 'in this respect. At higher altitudes this is not the case and all three tracking alternatives can given comparable performance.

With respect to the various TDAS constellations, Option 3 gives typically better performance due to better geometric distribution of the tracking data available with three satellites. In the two satellite constellations, Option 1 with (130° spacing) is better in the high inclination orbits, while Option 2 (with 162° spacing) is better in low inclination orbits. Although Option 2 provides nearly full coverage (\geq 98%), the performance for high inclination users is sensitive to poor geometry conditions for doppler tracking. This occurs twice a day when the user orbit normal points in the general direction of each TDAS. Consequently a tradeoff exists between maximum TDAS spacing (162^A) for best coverage and lower spacing to achieve better OD performance.

Figure 5-7 indicates the major error components affecting OD performance. Those due to unestimated (consider) parameters are: gravitational harmonics (H), drag (D) and TDAS ephemerides (E). The noise component (N) is the random error in user position which is controllable by filter tuning parameters.^{*} Ideally the filter should be tuned to reduce the impact

By adjusting filter gains to partially deweigh prior estimates random errors arising from measurement and a priori uncertainties do not continuously decrease as the number of measurements increases. Rather a "noise floor" (not necessarily constant over time) is reached, which is controllable using velocity state noise tuning parameters.
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of the dominant consider parameter(s) without excessively increasing the noise component. Figure 5-7 suggests that the modelled filter is somewhat overtuned for the 28°, 200 km orbit, about right for the 97°, 200 km orbit and undertuned for the higher altitude orbits.

In the high inclination, high altitude orbits (98°, 600/1000 km) gravitational harmonic modelling uncertainty is typically the dominant error contributor. This is also the case at low altitude with beacon tracking since filter tuning acts more to suppress the drag contribution. At low altitudes with scheduled tracking, drag is the dominant contributor in all cases. TDAS ephemeris error is typically a secondary contributor. All of the other consider error sources identified in Table 5-2 are comparatively insignificant including user oscillator drift.

Also indicated in Figures 5-6 and 5-7 are user requirements extracted from the TDAS mission model (see Appendix A). Most would be met with beacon tracking and in many cases with scheduled tracking based on error analysis results using the assumed models. Meeting the more stringent requirements would imply some improvement in key error sources. For example, in the 98°, 600 km orbit with TDAS Constellation Option 1, the impact of gravitational harmonic modelling errors would have to decrease by at least 5:1 and TDAS ephemeris error by 2:1. Filter retuning would decrease the noise component so the net effect of all contributors would be ≤ 10 m. Such reductions could also lead to meeting the most stringent requirement (30 m) at a lower altitudes assuming no change in the modelled drag uncertainty.

5.4.2.2 User Time Accuracy

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Figure 5-8 indicates that TD performance is uniformly better with beacon tracking (\leq .25 µsec) than with scheduled tracking in all four orbits. Nevertheless, performance with the scheduled alternatives is sufficient to meet the most stringent user time requirement (1 µsec) identified in the TDAS mission model.

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Figure 5-9 shows the error contributors affecting TD performance. With FBLT, TDAS ephemeris errors tend to dominate. With FLST and RLST, in the high altitude orbits, gravitational harmonic errors are also a primary contributor. In the lower altitude orbits the TD errors due to the drag uncertainty increase but are still secondary. User oscillator drift is also a secondary source in all cases except for RLST if the interval between navigation data uploads begins to exceed 3 hours.

5.4.3 <u>Sliding Batch Data Processing Results</u>'

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Navigation performance data based on the error analysis results in Appendix E are presented in terms of user position and time accuracies as summarized in Figures 5-10 and 5-12. Corresponding position and time error budgets are given in Figures 5-11 and 5-12. FLST results are shown for batch intervals (P) of 1.5 and 3 hours.^{*} RLST results correspond to navigation data uploads at \sim 1, 2 or 3 orbit intervals.

5.4.3.1 <u>User Position Accuracy.</u> Figure 5-10 indicates that user position errors in low altitude orbits are substantially higher than the sequential provessing results. At higher altitudes however, OD performance is comparable. In either case no significant OD performance difference was observed with respect to the two satellite constellation options (1 and 2) studied.

Figure 5-11 indicates the major error contributors affecting OD performance. At low altitudes drag is dominant even though a tracking interval of only 6 hours was used. For the higher altitude orbits a longer tracking interval (18 hours) was found to give better performance; consequently user oscillator drift became a significant contributor along with gravitational harmonic error.

* See Figure 5-4. The assumed computation/data handling interval (C) was 1.5 hours for FLST and 10 minutes for RLST as discussed in Section 5.3.

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With respect to meeting user requirements these alternatives would not be good candidates in low altitude orbits. However, at the higher altitudes all but the most stringent requirement (10 m) could be met based on the error analysis results. Metting the 10M requirement would imply some improvement in key error sources, e.g., a decrease in gravitational harmonic errors by at least 4:1, TDAS ephemeris error by 1.5-2:1 and user oscillator drift by 5:1.

5.4.3.2 User Time Accuracy. Figure 5-12 indicates that TD performance is at best 2.5-3 μ sec. This is significantly lower than with sequential processing and would not meet the most stringent time requirement of 1 μ sec.

Figure 5-13 shows that the dominant error contributors are the same as for user position error -- drag at low altitudes and user oscillator drift at high altitudes.

5.4.4 Observations

A summary of user navigation performance based on the error analysis results is given in Table 5-4. In addition, the following observations can be made:

- <u>Sequential Processing</u>
 - At low altitudes (e.g. ≤400 km) beacon tracking (FLBT) provides significantly better orbit determination (OD) accuracy than scheduled alternatives (FLST, RLST). The latter are both highly sensitive to tracking frequency and RLST to NAV upload rate.

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- At higher altitudes all three alternatives provide compairable OD performance.

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- Performance projections based on error source improvements indicate that FLBT could satisfy all users in the TDAS mission model with position accuracy reqts. down to 10m. For this accuracy error sources requiring improvement include:
 - -- gravitational harmonic modeling (\geq 5:1 reduction in GEM-9 errors)
 - -- TDAS tracking accuracy (to 25m pos., 2.5 mm/s vel.)
- Time determination (TD) is better than 0.25 μsec with FLBT and better than 1 μsec with FLST and RLST.

Sliding Batch Processing

- Results for FLST and RLST indicate that compared to sequential processing:
 - -- OD performance tends to be worse for low altitude orbits (< 400 km) where drag is the dominant error source and comparable to or slightly better in higher orbits where user oscillator drift is also a major error source.
 - -- TD performance is worse (\geq 2.5 µsec) for all orbits considered and particularly sensitive to the tracking interval duration.

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TABLE 5-4

COMPARISON OF USER NAVIGATION PERFORMANCE (ERROR ANALYSIS)

SUMMARY OF USER NAVIGATION PERFORMANCE EVALUATION

·····	·····	USER ORRIT						
ACKING DATA PROCESSING ALGORITIM	EVALUATIJN ITENS	200 28°	USER ORBIT KM 97°	600/1000 KM 98°				
BEQUENTIAL {FLBT {FLST RLST	ORELT Accuracy	• FLBT IS SIGNIFICANTLY B • FLBT COULD MEET ALL REG MODEL DEPENDING ON CURS	• FLBT, FLST & RLST CAN GIVE COMPARABLE RESULTS WITH APPROPRIATE FILTER TUNING • ALL COULD MEET REQUIREMENTS > 30 M					
	TDAS CONSTELLATION® INPACT	ALL OPTIONS MEET ACCURACY Requirements Option J Offers up 10 211 Advantage	• PERFORMANCE DEGRADES SIG- NIFICANTLY FOR OPTION 2 DUE TO POOR GEOMETRY RECURRING TWICE PER DAY	• OPTION 3 HAS SLIGHT ACCURACY ADYAMTAGE WITH FLBI; OTHERWISE ALL OPTIONS ARE COM- PARABLE				
	LEADING ERROR CONTRIBUTORS	N. H Л. H. E. D		н, х. Е				
	TIME ACCURACY	 ● FLBT IS SUPERIOR FC: ALL CASES (NINIMUM ACCURACY ≤ 0.25 USEC) ● FLST & RLST (WITH UPLOAD RATE ≤ 3 HOURS) ALSO MEET HOST STRINGENT REQUIREMENT (1) 						
	TDAS CONSTELLATION" IMPACT	• ALL OPTIONS GIVE COMP	ALTERNATIVE					
<u></u>	LEADING ERROR CONTRIBUTORS	м, Е	N, E. H					
	ORBIT Accuracy	HEITHER FLST OR RLST HEI REQUIREMENTS	BOTH FLST & RLST COULD MEET REQUIREMENTS > 30 M					
SLIDING BATCH FLST RLST	LEADING ERROR CONTRIBUTORS	1	0, E, H					
	TINE ACCURACY	• MAXIMUM ACCURACY FOR						
	LEADING ERROR CONTRIBUTORS		0, E. H					
OPTION 1 (2	S/C. 100° APART); OP S: M = MEASUREMENT H = GRAY(ATIONA (ARJUNGEME)T IN	TICH 2 (2 S/C. 162° APART); OPTION 1 MOISE; E = TOAS ORBIT UNCERTAINTY; O L MARMONICS MODELLING EAROR. O = USE TABLE NOT NECESSARILY IN ORDER OF SI	(3 S/C). • RESIDUAL DRAG HODELLING ERROR. R OSCILLATOR DRIFT. GNIFICANCE.)	·				
Sti	STANFORD TELECOMM	UNICATIONS INC.						
	C - Q.	VT_	5_29					

SECTION 6

STUDY SUMMARY AND CONCLUSIONS

The TDAS navigation architecture study has focused on three TDAS-based one-way tracking alternatives for providing user orbit and time determination (OD/TD). This section presents a functional overview and requirements summary for each alternative, summaries of potential user navigation performance and TDAS tracking accuracy, and the conclusions reached from the study.

6.1 FUNCTIONAL OVERVIEWS AND REQUIREMENTS SUMMARY (2, 3)*

The proposed TDAS architecture design [1a] would extend TDRSS navigation and data relay capabilities to support:

- beacon tracking (FLBT) via an independent beacon signal, and
- scheduled one-way tracking (FLST, RLST) via enhanced multiple and single access (MA, SA) services.

A functional overview of each alternative is presented in Figure 6-1.

The beacon signal is assumed to be generated by a TDAS ground terminal for broadcast by each TDAS satellite using a single element of the S-band MA antenna array. Users would receive continuous tracking signals while within a beacon antenna's $\pm 13^{\circ}$ field of view. The corresponding upper limit on user altitude for 100% coverage exceeds 3100 km for all TDAS constellations considered. Lower altitude coverage is governed by the zone of exclusion (ZOE) size for a given constellation.

In the scheduled modes, signals for tracking are available only during an allocated contact period as part of normal MA or SA service. Since each

^{*} Numbers in brackets () refer to preceding sections with full details.

IGURE 6-1: ONE-WAY NAVIGATION ALTERNATIVES - FUNCTIONAL OVERV						
FORWARD LINK BEA	CON TRACKING (FLBT) FORWARD LINK SCHEDULED TRACKING (FLST)					
	TDAS TDAS					
5	R Q					
	26*					
	USAT					
TERM.	GXD TERM.					
	MAJOR FUNCTIONS					
ELEMENTS	FORWARD LINK BEACON TRACKING FORWARD LINK SCHEDULED TRACKING (FLBT) (FLST)					
TDAS	TRACKING SIGNAL & ANCILLARY DATA* GENERATION AND UPLINK					
GROUND SEGHENT	CONTINUOUS TRANSHISSION SCHEDULED SERVICE ONLY					
TDAS SATELLITES	BENT PIPE RELAY FOR SHA USERS ONLY (VIA BROADBEAM ANTENNA- SINGLE SMA ELEMENT) BENT PIPE RELAY FOR ALL SERVICES (SMA, KSA, WSA, LSA)					
USERS	RANGE/DOPPLER DATA MEASUREMENT					
	* UN-BOARD ORBIT/THE DETERMINATION					
	TDAS					
ETURN LINK SCHEDU	ED TRACKING (RLST)					
	USAT					
	TRACKING SIGNAL					
	TERM.					
ELEMENTS	MAJOR FUNCTIONS					
	TRACKING SIGNAL GENERATION AND TRANSMISSION DURING SCHEDULED RETURN LINK SERVICE					
USERS	 RECEIVE AND PROCESS GROUND-DERIVED DATA FOR ORBIT/CLOCK STATE PREDICTION UNTIL NEXT UPLOAD VIA SCHEDULED FORWARD LINK COMM. SERVICE 					
	RANGE/DOPPLER DATA MEASUREMENT					
TDAS	 USER ORBIT DETERMINATION AND CLOCK (OSC.) CALIBRATION 					
SEGMENT	 ORBIT PREDICTION DATA GENERATION AND TRANSMISSION TO USER AT SCHEDULED INTERVALS 					

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TDAS can support two S-Band MA forward channels and eight SA channels, up to 20 simultaneous users theoretically could be supported for FLST with a two satellite TDAS constellation depending on channel scheduling policy. Similarly, for RLST up to 20 simultaneous MA users theoretically could be accomodated, since each TDAS can support ten return S-Band MA channels.

Table 6-1 lists various system requirements for supporting each 1-way alternative. Table 6-2 summarizes characteristics of the tracking signals assumed in each case including estimates of metric tracking data (R, \hat{R}) accuracies based on transmission link performance analyses.

The multiple beam antenna and switch enhancements for TDAS spacecraft provide the capability for simultaneous, direct transmissions between the space segment and several ground stations [1d,e]. This provides possibilities for <u>direct</u> control of user spacecraft by the mission control centers (MCC) instead of interfacing through the White Sands (WSN) terminal and/or Network Control Center (NCC) as in TDRSS.

Figure 6-2 illustrates options for tracking signal and navigation data flow with each of the one-way alternatives. Since the beacon signal for FLBT is a general resource, it is assumed to originate at WSN, the assumed control point for TDAS spacecraft. Navigation data computed on-board can be received by a TDAS ground terminal at the NCC directly and by MCCs with direct space/ground access. Additional interfacility transfer of data can occur to support MCCs without direct space/ground access or for general coordination and/or verification functions. The latter includes WSN which has control of TDAS spacecraft facilities (e.g., SA antenna/telescope pointing).

With FLST (see Figure 6-2b) the user tracking signal is imbedded in the normal <u>uplink</u> data communication traffic so it can emanate from either the NCC or a cognizant MCC. Navigation data computed on-board can be distributed in the same manner discussed above for FLBT.

ELEMENT		SYSTEM REQUIREMENT	FLBT	FLST	RLST
	•	BEACON SIGNAL GENERATION	7		
	•	ANCILLARY NAV DATA	7	7	
GROUND		GENERATION & TRANSMISSION			
	•	USER OD/TD* COMPUTATION &			7
		NAV DATA UPLOAD			
	•	DEDICATED HARDWARE:			
SPACE		- TWTA & KU-TO-S BAND REPEATER	7		
		- SMA ANTENNA ELEMENT	7		
	•	DEDICATED RECEIVER CHANNEL	>		
	•	STABLE FREQUENCY STANDARD	7	7	7
	•	R,R EXTRACTOR	>	7	7
USEK	•	NAV DATA DECODER	7	7	7
	•	ANCILLARY DATA GENERATION			
	•	ON-BOARD COMPUTING FACILITY:			
		– FULL OD/TD [*] COMPUTATION	7	7	
		- ORBIT/TIME PROPAGATION ONLY			7

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		SCHEDNA ED (RLST)	 STAGGERED QUADIRPILASE PN (SQPN) (PER IDRSS DGI-HODE2/ KONCONERENT) 	 2047 CITP GOLD CODES ON 1 AND Q CINARELS (OFFSET 2 20,000 CITPS) 3,0778 x 10⁶ CITPS/SEC 	INCLUDES USER ANCILLARY DAYA (E.G., PN CUDE EPOCH, CURRENT TIME REFERENCE AND SYNCH MORD FOR AMBIGUILY RESOLUTION)	5) • CONVOLUTICNALLY-CODED (RATE 5)		2.5 M 10 M**	7 MYSEC	
-2	NAL CHARACTERISTICS	SCIEDULED (FLST)	 Slaggered Quadripliase ph (soft (command & range cliannels per torss) 	 SHORT CODE (COMMAND CHANNEL/ 1023 CHIP GOLD CODE) LONG CODE (RANGE CHANNEL/ 256 x 1023 CHIPS USED TO RESOLVE RANGE ANDIGUITY 3.0778 x 10⁶ CHIPS/SEC 	© INCLUDES ANCILLARY DATA (Similar to flet)	0 CONVOLUTIONALLY-CODED (RATE 5	• PROVIDED FOR SIGHAL, ACQUISI- ILON	- K 0	3 MV/SEC	51
TABLE - SUMMARY OF TRACKING SIG	SUMMARY OF TRACKING SIG	BEACON (FLBT)	PH HODULATED BPSK DATA	 UNIQUE 1023-CIILP GOLD CODE 3.0778 x 10⁶ ciltps/sec 2.6 hs granularity (via interpolation to 1/128 ciltp 	 INCLUDES ANCILLARY DATA (E.G., IDAS EPUEMERIS, TIME REFERENCE AND PN EPOCII) <u>2</u> [UNÜ BITS/FRAME 	 Cohvolutionally-Coded (rate 5) Data rate ~ 100 bps Assembled into 1000 bit frames (10 seconds) Sync mord in Each frame re- solves ambiguity 	 NONE (NOT FEASIBLE FOR MUL- TIPLE USER SUPPORT, BUT COOR- DINATION MITH FLST ENABLES RE- DUCTION OF DOPPLER UNCERTAINTY FOR INITIAL BEACON ACCUISITION 	5 M BU M.	6 NM/SEC	DED ASYKCHRONOUSLY TO PH CODE(S) Tematic delay (mith calibration) []
		SIGHAL CUMRACTERISTICS	IYPE	COIXE (S)		DAIA	DGPPLER Compensation	RACKING R (RANDON) IRACKING (SYSTEN) DATA	ACCURACY Ř (RAHDOH)	PROJECTED SYS

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With RLST (see Figure 6-2c) the user tracking signal is imbedded in the normal <u>downlink</u> data communication traffic, so it can be received by either the NCC or cognizant MCC. Ground processing for user OD/TD can occur at the Orbit Support Computing Facility (OSCF) or at the MCC with subsequent interfacility data transfer as noted above. Since two-way tracking with ground-based processing is also a TDAS alternative, Figure 6-3 provides a comparison with the RLST data flow.

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6.2 USER NAVIGATION PERFORMANCE SUMMARY

To assess the potential navigation performance with each of the one-way alternatives, user OD/TD accuracy was evaluated as a function of various parameters and compared with requirements in the TDAS mission model. Figure 6-4 gives an overview of the elements involved in the analysis.

<u>User Orbits and TDAS Constellations</u> (5.2) - Options considered are shown in Figure 6-5. Constellation Option 1 is analogous to TDRSS with two satellites spaced 130° apart which provides 85-100% coverage at altitudes down to 200 km. Option 2 also uses two satellites, but with the maximum allowable spacing, 162°, which yields 98-100% coverage. Option 3 has three satellites, two deployed as in Option 1 and a third on the backside, which together provide 100% coverage. .

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The low altitude (high drag) user orbit types are of interest to determine whether more frequent tracking data, available with FLBT, is of significant benefit. The high and low inclination orbit types are of interest, since their coverage and geometrical properties can differ significantly.

<u>Tracking Schedules/Processing Algorithms</u> (5.2) - For the FLBT mode, the schedule was assumed to be continuous for a selected TDAS with metric tracking data (R,R) processed sequentially. For the FLST and RLST modes, the tracking schedule impact was assessed based on tracking every orbit and every other orbit during a 10 min pass/TDAS. Also, navigation performance was evaluated for both sequential and sliding batch^{*} data processing.

<u>Error Modelling</u> - User OD/TD error was computed via covariance analysis programs [22,23] given nominal TDAS and user orbits, a tracking schedule and processing approach, and appropriate models of the tracking error sources. The latter comprised measurement errors and various systematic errors, e.g., gravitational harmonic modelling error, drag modelling error, TDAS ephemeris error, user oscillator drift affects, etc.

^{*} Details are given in Section 5.

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Table 6-3 lists the baseline modelling assumptions made for these particular error sources and their relative impact on achieving a 40 M OD accuracy with beacon tracking in the user orbits considered. In this case gravitational harmonic error dominates, so TDAS ephemeris error is only a secondary effect. The projected requirements indicate the gravitational model improvement required (in terms of GEM-9 errors) to achieve 10 M and 30 M accuracies in the noted orbits. At this level, the modeled TDAS ephemeris error would also be significant so a reduction of ~2:1 is indicated.

<u>Navigation Performance With Sequential Processing</u> (5.4.2) - Figure 6-6 shows the OD performance (max error over 24 hours) with beacon tracking for each TDAS constellation option along with user requirements identified in the TDAS mission model (see Appendix A). Higher errors tend to occur for Option 2 in near polar orbits particularly at low altitudes due to recurring poor geometry.*

<u>Error analysis</u> results correspond to the basic modelling assumptions and the <u>projected</u> results to the improved errors indicated in Table 6-3. Thus, beacon tracking is projected to satisfy all user requirements in the TDAS mission model down to 10 M.

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Figure 6-7 shows the navigation performance with FLST for two assumed tracking schedules and sequential processing. In contrast to beacon tracking, the performance in low orbits is drag dominated and substantially worse. At higher altitudes, results are comparable.

Figure 6-8 shows the navigation performance with RLST for the same tracking schedules and two different nav data upload rates - 1.5 hours and 3 hours. Again, performance is worse in low orbits (drag dominated), but is comparable at higher altitudes.

^{*} This occurs about twice a day when the orbit normal points toward a TDAS thereby limiting the range change over the pass (poor doppler character-istic).

TABLE 6-3

ERROR CONTRIBUTOR SUMMARY FOR BEACON TRACKING

(SEQUENTIAL DATA PROCESSING)

CONSENTS	≥ 10:1 1kPROVEMENT ANTICIPATED BY EARLY 1990'S [36]	OPTINIZE PROCESSOR TUNING PARAMETERS	 HINIHAL BRIS CONFIG. CANNOT MEET PROJECTED REQT. VLBI TRACKING MEETS PROJECTED REQT. 	CURRENILY AVAILABLE CAPABILITY
PROJECTED REQTS. FOR 10 m ACCY.* 30 m ACCY.**	20% GEH-9 Error	SAHE AS Hodel	25 m P05. 2.5 mm/s YEL.	SAHE AS Houel
ORBIT DETERHINATIOH IHPACT FOR 40 m ACCURACY	HAJOR	SECONDARY	SECONDARY	NEGL IGIBLE
MODELING Assumptions For Analysis	100% GE N- 9 Error	2.5% OF Nohinial CD (Residual Error)	50 m POS. 5 mm/s YEL.	10-10/DAY
ERROR CONTRIBUTOR	GRAVITATIONAL HARMONICS MODEL	DRAG COEFFICIENT (CD) (<600 Km ORBITS)	TDAS ORBIT ERNOR	USER OSC. DRIFT

[±]98°, 600 km 0rbit **98°, 200 km 0rbit

+ GEH-9 = GODDARD EARTH MODEL 9 (1979)

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TD accuracy results^{*} indicate that performance is uniformly better with beacon tracking (\leq .25 µsec) than with scheduled tracking in all orbits considered. Nevertheless, performance with the scheduled alternatives is sufficient to meet the most stringent user time requirement (1 µsec) identified in the TDAS mission model.

<u>Navigation Performance With Sliding Batch Processing</u> (5.4.3) - Results for FLST and RLST indicate that compared to sequential processing:

- OD performance tends to be worse for low altitude orbits (< 400 km) where drag is the dominant error source and comparable to or slightly better in higher orbits where user oscillator drift is also a major error source.
- TD performance is worse (\geq 2.5 µsec) for all orbits considered and particularly sensitive to the tracking interval duration.

* See Section 5.4.

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6.3 TDAS TRACKING ANALYSIS SUMMARY

Information about TDAS orbits is assumed to be provided for user OD/TD on a recurring basis. User navigation performance is then a function of TDAS orbit uncertainties in the interval between updates (i.e., prediction interval errors). To assess potential orbit prediction errors, an accuracy analysis was made of TDAS tracking with the BRTS and VLBI techniques. An overview of the elements considered in pursuing this analysis is given in Figure 6-9.

<u>Tracking Configurations</u> (4.1) - In terms of TDAS tracking, the various constellations reduce to two situations: frontside and backside satellites. Figure 6-10 illustrates configuration options for acquiring TDAS tracking data with each technique.

With <u>BRTS</u> tracking, two-way range and range-rate (R, \tilde{R}) data are acquired from transmissions originating at White Sands (WHS). These are returned by bilateration ranging transponders (BRTs) operating at pairs of automated ground stations (one pair per TDAS in the <u>minimal</u> configuration). For the backside TDAS, transmissions are also relayed via the crosslink which impacts tracking accuracy since measurements are affected by uncertainties in the frontside satellite orbit.

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With <u>VLBI</u> tracking, the fundamental measurement data type is the difference in range (ΔR) between a signal source and two receivers displaced along a known baseline. For the TDAS application, signals received at each automated station would be returned to WHS for processing. Typically, one processed pair of VLBI observations per hour per station set (three per TDAS) would be needed. Crosslink uncertainties are not a factor due to the inherent differencing of common path components.

<u>Error Modelling</u> (4.2.1) - TDAS OD was assumed to employ batch processing of tracking observations taken over a given tracking interval. Orbit prediction errors beyond the end of tracking were evaluated via a covariance

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analysis program [16] given a nominal orbit, tracking station locations, . tracking schedule and appropriate models of significant tracking error sources (measurement noise and bias, station survey, solar pressure and tropospheric uncertainties).

BRTS measurement errors were assumed consistent with TDRS OD analyses [18]. VLBI measurement errors were considered at two levels: a conservative <u>baseline</u> model consistent with values used in studies of a Deep Space Network (DSN) tracking application to TDRS [19] and a more optimistic <u>reduced</u> model which reflects achieved capabilities with VLBI stations observing celestial radio sources [20]. Station errors were taken to be the same as BRTS for the <u>baseline</u> model and an order of magnitude better for the <u>reduced</u> model.

<u>TDAS Tracking Accuracy</u> (4.2.2) - For evaluating performance, the first criterion considered was the peak error over a 24 hour continuous prediction interval from the end of tracking. Analysis results indicate, however, that since solar pressure is so dominant, OD accuracy can be significantly affected by the choice of reference epoch and duration of the tracking interval and prediction interval. Thus, in both BRTS and VLBI analyses efforts were made to identify multiple combinations (epoch/tracking interval/prediction interval) which, taken together, yield better accuracy over a 24 hour period than any one combination alone.

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Table 6-4 gives results for both frontside and backside TDAS OD accuracy via BRTS tracking using one 24 hour prediction interval and two concatenated 12 hour intervals. Multiple segments are effective for the frontside TDAS but not the backside which is >2.5:1 higher. The prediction error is due primarily to the impact of the frontside TDAS position uncertainty assumed to be 100 M.

Table 6-5 and Figure 6-11 present the results for TDAS OD accuracy with VLBI tracking for both baseline and reduced error models. The reduced

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TDAS	TRACKING Interval (Hours)	PREDICTION INTERVAL(S) (HOURS)	MAXIMUM POSITION ERROR (M)	PRIMARY ERFOR
FRONTSIDE	18	24	175	150 (sp) 80 (в) 40 (s)
100°¥	18	12 + 12**	110	80 (B) 60 (SP) 45 (S)
BACKSIDE	36	24	290	230 (s') 175 (sp) 25 (b)
BACKSIDE 98°E	36	12 + 12**	280	220 (s') 165 (sp) 25 (b)

LEGEND: SP - SOLAR PRESSURE; B - BRT BIAS; S - STATION SURVEY; S' - STATION SURVEY + TIGO LOCATION (TDAS 100"W).

** CONCATENATED INTERVALS BASED ON TRACKING INTERVAL EPOCHS SPACED -12 HOURS APART.

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TABLE 6-5

MAXIMUM TDAS POSITION ERROR IN 24 HOUR PREDICTION INTERVAL WITH YLBI TRACKING

VLBI ERROR MODEL	TRACKING INTERVALS (HOURS)	PREDIO INTEI NO. SEGMENTS	CTION RVAL LENGTH (HOURS)	HAXIMUN POSITION ERROR • (H)	PRIMARY ERROR CONTRIBUTORS** (M)
BASELINE	18/36	4	9/3	87	72 (SP) 45 (YB) 25 (S)
BASELINE (WITH BIAS ESTIMATED)	18/12	4	6/6	72	56 (SP) 45 (YN) 20 (S)
REDUCED	6	24	1	12	10 (TR) 6 (SP) 4 (S)
REDUCED (W. 51 TROPO)	6	24	1	ß	6 (SP) 4 (S) 3 (TR)

THAS @ 100"W; STNS: WHS-AGD, WHS-HAW; DATA RATE = 1/HR **

LEGENO: SP - SOLAR PRESSURE S - STATION SURVEY VB - VLBI MEAS. DIAS VN - VLBI MEAS. NOISE TR - TROPOSHERE

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model results illustrate dramatically the benefit of more accurate measurements since shorter tracking and prediction intervals can be used to offset the pervasive influence of solar pressure errors. Additional improvement to the 5-8 M level would be possible if better tropospheric correction $(\leq 5\%)$ can be achieved.

Results of the TDAS tracking analysis lead to the following observations:

- Orbit determination error for backside TDAS satellite tracking via BRTS is significantly higher (\sim 2-3:1) than for frontside tracking based on two BRT sites per satellite.
- VLBI tracking offers the potential for significantly improved TDAS orbit accuracy (< 10m) for both front and backside satellites
 - This method is not sensitive to cross-link uncertainties, but tracking data from three stations would be required for each satellite
 - Realization of potential VLBI tracking accuracy requires:
 - -- Calibration of tracking biases and station location/ baseline uncertainties (e.g., using ΔVLBI data)
 - -- Judicious selection of epoch (time of day), tracking interval, and prediction interval combinations.

Preliminary results for VLBI tracking with Conus-based sites also indicate increased accuracy, although the shorter baselines may reduce the potential improvement compared to that with intercontinental baselines.

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6.4 CONCLUSIONS

Analysis of the study results leads to the following key findings:

- TDAS beacon tracking (FLBT) will satisfy all users in the TDAS mission model with position accuracy requirements down to 10 M.
- Scheduled tracking alternatives (FLST, RLST) can also meet the accuracy requirements except at low altitudes where performance is highly sensitive to:
 - Drag uncertainty
 - Frequency of tracking passes, and/or
 - Frequency of navigation data uploads (RLST/only).
- A two or three satellite TDAS constellation impacts performance as follows:
 - Selecting 2 satellites leads to a tradeoff between coverage and accuracy. Increased satellite spacing improves coverage, but a point is reached where performance in high inclination orbits begins to degrade (130° spacing appears better than 162°).
 - Selecting 3 satellites provides full coverage and up to a 2:1 advantage in navigation accuracy over two satellites.

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 Projected TDAS tracking accuracy requirements (25 M-POS. & 2.5 MM/SEC-VEL.) can be met with VLBI tracking but not with a minimal* BRTS configuration.

6.5 · RECOMMENDATION

Based on the study results beacon tracking is recommended as the prime approach for routine navigation support. Scheduled tracking alternatives, one-way and two-way, should also be considered for supporting user navigation functions, as proposed in Table 6-6.

* 2 BRTS sites/TDAS.

TABLE 6-6

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SUMMARY OF PROPOSED TDAS NAVIGATION FUNCTIONS

TDAS ARCHITECTURE WILL ACCOMODATE ALL ONE-WAY AND TWO-WAY TRACKING TECHNIQUES.

user Fl	R SUPPORT UNCTION	PROPOSED TECHNIQUE	RATIONALE
	ROUTINE	BEACON TRACK ING	- MAX. AVAILABILITY - MEETS REQUIREMENTS
NUTEATION	INITIAL ACQUISITION	SCHEDULED TRACKING (FORWARD LINK)	DOPPLER COMPENSATION AVAILABLE
	BACKUP	SCHEDULED TRACKING FWD./RTN./OR LINK/LINK/2-WAY	PROVIDE REDUNDANCY VIA COMM. CHANNELS
	VERIFICATION	SCHEDULED TRACKING (2-HAY)	- GROUND-BASED OPNS. - Quality control
COMM./NAV.	TDAS ANTENNA POINTING	SCHEDULED TRACKING (RETURN LINK) OR BEACON TRACKING	 GROUND-BASED OPNS. ACCURACY SUFFICIENT DIRECT DOWNLINK TO NCC

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APPENDIX A

TDAS MISSION MODEL

In Task 1 of the TDAS study [1b], a screened baseline of NASA plans was used to generate scenarios of experiments/missions for the 1990 - 2005 time period. An estimated flight schedule was established by first assigning <u>planned</u> experiments, then <u>candidate</u> experiments and finally <u>opportunity</u> experiments. Where planning data for the 1990s was unavailable, generic experiments/missions were developed based on trends established in the 1980's planning data. Table A-1 lists the adopted TDAS mission model and corresponding orbit data.*

Potential navigation accuracy requirements for the various experiments/missions were developed from user community survey data, subsequent conversations with designated points-of-contact and/or independent estimates. Table A-1 lists the position and time accuracy requirement? developed in each case. Except for the TOPEX mission, the most stringent position accuracy requirements is 10M. Time accuracy requirements were found to be no less than 1 µsec.

For purposes of comparing user navigation requirements with estimated navigation performance results, user orbits were divided into six categories (A-F). Tables A-2 and A-3 show the distribution by category and number of missions with a given requirement. Some missions are counted more than once, since the orbit range spans more than one category (see Table A-1).

More detailed information is available in [1b].

	T/	ABLE A-	-1		
TDAS	MISSION	MODEL	(1990	-	2005)

EXPERIMENT/MISSION		ALTITUDE (tm)	ORBIT* INCLINATION (deg)	CATEGORY	ACCURACY POSITION (#)	REQUIREMENT TIME (usec)
	SPACE TELESCOPE Advanced X-bay Astrophysics Facility	600 450	28.8	B	≥ 200 ≥ 1000	1
ļ	(AXAF)					***
	ASTROPHYSICS GENERIC (AG) = 1,2,3,4		28.5	8		100
	VERY LONG BASELINE RADIO Interferometer (VLBRI)	400-5000	45	, B1C		•
ASTROPHYSICS	X-RAY OBSERVATORY	300	28.5	Å	<u>></u> 200	<u>≥</u> 100
	COSMIC OPTICAL SYSTEM OF MODULAR IMAGING Collectors/100 meter thinned aperature Telescope	500	28	B	•	1
	LARGE OPTICAL/ULTRAVIOLET INTERFEROMETER	450	28.5	8	۳	· M
	ORBITING SUBMILLIMETER TELESCOPE	1000	SUN SYNC	F	4	TBD
	LARGE AMBIENT DEPLOYABLE IR TELESCOPE (LADIR)	400-700	28-50	8	H	<u> </u>
	SOLAR CYCLE AND DYNAMICS MISSION (SCADH)	575	28 or 98	B or E	<u>></u> 1000	1
TERRESTRIAL	SOLAR TERRESTRIAL GENERIC (SG) - 2	575	28	B	•	ta -
	SOLAR TERRESTRIAL GENERIC (5G) - 1,3	400	57	B	μ	
	TOPOGRAPHY EXPERIMENT (TOPEX)	1334	63.4	¢	3(0,1 ALT)	2 100
GLOBAL	ENVIRONMENTAL GENERIC (EG) - 2,5,6	1334	63	C	4	•
ENTIRUAMENT	OCEAN SYNTHETIC APERATURE RADAR (OSAR)	790	98	E	50	1
	ENVIRONMENTAL GENERIC (EG) = 1, J, 4	790	98	£	•	
	SOIL AND SNOW MOISTURE RESEARCH AND ASSESSMENT MISSION	400-700	60-98	BorE	100	
RESOURCE	OPERATIONAL EARTH RESOURCES SATELLITE/ Advanced Land Observing System (Oers/Alos)	705	99	E	10	N
OBSERVATION	RESOURCES GENERIC (RG) - 1.2.J	700	98	Ε		۲
	ADVANCED THERMAL MAPPING APPLICATIONS	620	97.8	٤	50	•
	HAGNETIC FIELD SURVEY B	300	97	D	30	
HETEGROLOGY	OPERATIONAL HETEOROLOGY SATELLITE (2)	830	98.7	F	500	> 100
	SHUTTLE - 1	185-1110	28.5-57	٨-٥	100	1 (FUTURE)
SPACE	SHUTTLE - 2	185-1110	70-104	0+F	•	*
	TELEOPERATOR MANEUVERING SYSTEM (THS)	185-1000	VARIOUS	A-F	M	TBD
TIMISPURIATION	ORDITAL TRANSFER VEHICLE (OTV)	LEO-GEO	VARIOUS	A=F	H	
	HEAVY LIFT LAUNCH VEHICLE (HLLY)	200-500	VARIOUS	A,B or D,E	м	*
	MANNED URBITAL TRANSFER VEHICLE (MOTV)	LEO-GEO	YARIOUS	A=F		
	SPACE STATION/PLATFORM (2)	400	28.5	6	10-30	1

* LEO - Low Altitude Earth Orbit

GED - Geostationary Earth Orbit

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** For Clock Synchronization Only, Needs Hydrogen MASER Stability.



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ORIGINAL PAGE 19 OF POOR QUALITY <u> 780 - 5</u> <u>2</u> 100 µs - 1 USER TIME ACCURACY REQUIREMENTS AND ORBIT DISTRIBUTION TED-3 1 µs - 3 h < 800> 100 µs - 1 ب ALTITUDE, h (km) 300 ≤ h ≤ 800 <u>></u> 100 µs - 1 TBD - 5 1 µs - 11 TBD - 4 l μs - 8 IN TDAS MISSION MODEL æ ш TABLE A-3 > 100 µs - 1 h <u>≤</u> 300 1 µs - 1 4 1µs - 2 4 TBD -**TBD** -Ο К INCLINATION I (DEG) I <u><</u> 70° $I < 70^{\circ}$ TELECOMMUNICATIONS INC.

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APPENDIX B

TRACKING DATA MEASUREMENT ERRORS: RANDOM COMPONENT

Random errors in the range and Doppler measurement data arise due to noise in the tracking link. This appendix estimates these errors for the three one-way tracking alternatives based on the signal definitions and assumptions in Section 3. Parametric results are presented where applicable to indicate the sensitivity to design parameters; e.g., TDAS EIRP and user G/T or user EIRP. The results are presented as follows:

- 8.1 Thermal Noise in Range Measurements
- 8.2 Thermal Noise in Range-Rate Measurements
- B.3 Phase Noise in Range-Rate Measurements

B.1 THERMAL NOISE IN RANGE MEASUREMENTS

The measurement of range in the systems considered here relies upon the acquisition and tracking of a pseudonoise (PN) ranging code. In all the one-way navigation modes, the receivers exploit the correlation properties of PN codes to produce an error signal which directs the timing correction. The principal figure of merit is the variance of the timing error as a function of the input ratio of signal power to noise density, C/N_0 .

B.1.1 Forward Link Tracking

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The Motorola 2nd generation TDRSS user transponder design [14] is assumed to apply to the users to be supported in the TDAS era. The design's code tracking loop is a digital implementation of a 2nd order type II tau-dither loop; a rate aided tracking loop configuration is also present, providing a 1st order tau-dither loop with frequency tracking from the carrier loop. Motorola expresses the variance of the tracking jitter due to thermal noise as

 $\sigma_{\rm T}^2 \leq \frac{{\sf B}_{\rm L}}{{\sf C}/{\sf N}_{\rm O}} \left[1 + 2 \frac{{\sf B}_{\rm IF}}{{\sf C}/{\sf N}_{\rm O}}\right] \qquad ({\tt chips})^2 \; , \label{eq:stars}$

where B_L is the closed loop bandwidth of the loop, either 1 Hz for the independent tracking mode or .125 Hz with rate aiding, and B_{IF} is the IF filter bandwidth of 4 kHz. This result can be seen to stem from Hartmann's analysis [25] of the tau-dither loop where the incoming PN coded signal is alternately correlated with early and late versions of a locally generated code replica.

Based on the forward link budget in Figure 3-1, Figure B-1 shows the RMS range error σ_R in meters as a function of the user G/T, parameterized by the TDAS EIRP. The mean square range error σ_R^2 is derived from the mean square timing error σ_T^2 by the simple conversion

$$\sigma_{\rm R}^2 = (c \ T_{\rm C} \ \sigma_{\rm T})^2 = (d_{\rm C} \ \sigma_{\rm T})^2 \quad ({\rm meters})^2,$$

where c is the speed of light and T_c is, again, the chip duration, assumed here to be 1/3.0778 Mcps. From these curves the forward link range errors due to thermal noise are extracted for a user with G/T = -27 dB/°K for both FLBT and FLST modes.

B.1.2 <u>Return Link Tracking</u>

For return link tracking, a ground receiver incorporating the Harris wide dynamics demodulator (WDD) is envisioned. The WDD uses a delay-locked loop to track the PN coded signal: early and later versions of the local PN code are correlated simultaneously with the input PN code sequence. An analysis of the exact implementation used in the WDD is not readily available, but work by Simon [26] provides results applicable to a delay-locked loop where the bandpass arm filters have bandwidths on the order of the data rate. A linear analysis for the mean-square timing jitter, assuming a large equivalent loop signal-to-noise ratio yields

$$\sigma_{\rm T}^2 = \frac{B_{\rm L}}{2 C/N_0} \cdot \frac{1}{S_{\rm L}} \quad (\rm chips)^2,$$

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where B_L is the single-sided tracking loop bandwidth and S_L is the squaring loss of the delay-locked loop. Simon shows that

$$S_{L} = \frac{D_{m}}{K_{D} + K_{L} \left[\frac{2B_{arm}/R_{s}}{R_{d}D_{m}}\right]}$$

where

$$R_{s} = data symbol rate = 1/T_{s},$$

$$B_{arm} = 2\text{-sided arm filter noise bandwidth}$$

$$R_{d} = symbol SNR = \frac{U}{N_{o}} T_{s},$$

$$D_{m} = \int_{-\infty}^{+\infty} S_{d}(f) |H(f)|^{2} df,$$

$$K_{D} = \frac{\int_{-\infty}^{+\infty} S_{d}(f) |H(f)|^{2} df}{\int_{-\infty}^{+\infty} S_{d}(f) |H(f)|^{2} df},$$

$$K_{L} = \frac{\int_{-\infty}^{+\infty} |H(f)|^{4} df}{\int_{-\infty}^{+\infty} |H(f)|^{2} df},$$

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 $S_d(f)$ = baseband power spectral density of the data, and H(f) = transfer function of the arm filter.

In keeping with the WDD parameters, it is assumed that $B_L = 4$ Hz in the tracking mode and that 1-pole arm filters are used. The symbol rate used for the calculation is assumed to be the maximum achievable data rate for a given user spacecraft EIRP, as detailed in Section 3.2.2 and shown in Figure 3-5. With the anticipated rate 1/2 convolutional coding, the symbol rate R_s is found as twice the achievable data rate. The tracking filters used in the WDD are chosen on the basis of symbol rate to optimally bandlimit the input noise. The 3 dB bandwidth f_c for the WDD is thus

 $f_{C} = \begin{cases} 8 \text{ kHz for all } R_{S} < 4 \text{ ksps} \\ 39 \text{ kHz for all } R_{S} < 26 \text{ ksps} \\ 300 \text{ kHz for all } R_{S} \leq 300 \text{ ksps} \end{cases}$

where $B_{1F} = \pi f_{C}$ for a 1-pole filter, $H(f) = \frac{1}{1+jf/f_{C}}$. The parameter K_{L} for a 1-pole filter is found as:

$$K_{L} = \frac{f_{c} \frac{\pi}{2}}{f_{c}\pi} = \frac{1}{2}$$

The other parameters in Simon's expression depend on the data modulation format as well as the arm filter transfer function. Assuming equiprobable independent transmitted symbols — an assumption not strictly valid with coding — the power density for the NRZ signalling format is given by

$$S_d(f) = T_s \frac{\sin^2(\pi f T_s)}{(\pi f T_s)^2}$$

for biphase (Manchester) coding, $B_d(f) = T_s \frac{\sin^4(\pi f T_s/2)}{(\pi f T_s/2)}$. It can be shown that

;

$$D_{\rm m} = \begin{cases} 1 - \frac{1}{2\pi {\rm R}^{\,\prime}} (1 - {\rm e}^{-2\pi {\rm R}^{\,\prime}}) & \text{for NRZ coding} \\ \\ 1 - \frac{1}{2\pi {\rm R}^{\,\prime}} (3 - 4 {\rm e}^{-\pi {\rm R}^{\,\prime}} + {\rm e}^{-2\pi {\rm R}^{\,\prime}}) & \text{for biphase coding} \end{cases}$$

$$K_{D} = \begin{cases} \frac{1 - \frac{[3 - (3 + 2\pi R')e^{-2\pi R'}]}{4\pi R'}}{D_{m}} & \text{for NRZ coding} \\ \frac{1 - \frac{[9 - 4(3 + \pi R')e^{-\pi R'} + (3 + 2\pi R')e^{-2\pi R'}]}{4\pi R'}}{D_{m}} & \text{for biphase coding} \end{cases}$$

where $R' = f_c T_s$.

Figure B-2 shows the above results expressed as an RMS range error in meters versus the user EIRP corresponding to the return link budget in Figure 3-5. The discontinuities appearing on the graph result from the change in arm filter bandwidth with data rate. A user EIRP of 2 dBW is assumed as a baseline value.

B.2 THERMAL NOISE IN RANGE-RATE MEASUREMENTS

Range-rate is estimated on the basis of a signal's observed doppler shift in frequency due to the relative motion between the transmitter and receiver. A carrier-tracking loop in the receiver provides the means to extract the phase ϕ (in radians) of the received signal and thus the received frequency, since $\phi = 2 \pi$ ft. By counting positive-going zerocrossings, N(t), over the averaging time T_{av}, an estimate \hat{R} of the range rate \hat{R} is derived as

$$\hat{\vec{R}} = \lambda \left[\frac{N(t+T_{av}) - N(t)}{T_{av}} \right] = \hat{\vec{R}} + \frac{\lambda}{2\pi} \frac{\delta \phi(t+T_{av}) - \delta \phi(t)}{T_{av}},$$

where $\delta\phi(t)$ represents noise in the phase measurement and λ is the carrier wavelength, assumed to be known a priori for one-way navigation. Assuming T_{av} is long enough so that the phase samples are independent and identically distributed, the RMS range-rate error is given by

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$$\sigma_{R}^{\circ} = \frac{\lambda}{2\pi T_{av}} \sqrt{\sigma_{\phi}^{2}(t+T_{av})^{+} \sigma_{\phi}^{2}(t)} = \frac{\lambda}{2\pi T_{av}} \sqrt{\gamma_{\sigma}^{2} \sigma_{\phi}^{2}(t+T_{av})^{+} \sigma_{\phi}^{2}(t)}$$

Here σ_{ϕ} is the RMS phase error in radians and one may identify $\sqrt{2}\sigma_{\phi}$ as the RMS doppler phase error. The performance of the carrier-tracking loop in the receiver thus determines the range-rate error due to thermal noise. In the results presented here, the doppler averaging time T_{av} is assumed to be 1 second.

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B.2.1 Forward Link Tracking

The Motorola 2nd generation TDRSS user transponder's carrier-tracking loop is a 2nd order type II modified Costas loop. For Costas loops in general, the mean square phase error is given by [27, 28]:

$$\sigma_{\phi}^2 = \frac{4BL}{C/N_0} \cdot \frac{1}{SL} (radians)^2,$$

where B_L is the loop noise bandwidth and S_L is the Costas loop squaring loss. The phase of a Costas loop, however, is twice the received signal phase, so the thermal noise phase jitter σ_{d} that is of concern here is

$$\sigma_{\phi}^2 = \frac{B_L}{C/N_0} \cdot \frac{1}{S_L}$$

For the Motorola design, $B_L = 40$ Hz and $S_L = .73$. With these parameters, Figure B-3 shows the RMS one-way range-rate error anticipated for forward link tracking assuming the link budget of Figure 3-1 and the various possible TDAS EIRP values. As before, a user G/T value of -27 dB/°K is used as a baseline.

B.2.2 <u>Return Link Tracking</u>

The wide dynamics demodulator's 1st order carrier-tracking loop, shown in Figure B-4a, has been analyzed by Weinberg [29], yielding the following expression for the received mean square phase error:

$$\sigma_{\phi}^{2} = \frac{B_{L}T_{S}}{\alpha \beta E_{b}/N_{O}} \left[\frac{1}{1 - \frac{2}{\beta}B_{L}T_{S}}\right] (radians)^{2}$$

Here

$$T_{s} = symbol duration,$$

$$\frac{E_{b}}{N_{0}} = symbol SNR = \frac{C}{N_{0}} T_{s},$$

$$\alpha = 1 - 2 Q(\sqrt{2 E_{b}/N_{0}}) + \sqrt{\frac{1}{\pi E_{b}/N_{0}}} e^{-E_{b}/N_{0}},$$



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$$\beta = 1 - Q(\sqrt{2 E_{b}/N_{o}}),$$

and Q(x) = complementary Gaussian distribution function

$$= \int_{\sqrt{2\pi}}^{1} \int_{x}^{\infty} e^{-u^2/2} du$$

In the WDD, the loop bandwidth is adjusted proportional to the symbol rate; for this calculation, it is assumed that $B_LT_S \approx .03$. If worst-case operation is assumed and the maximum achievable data rate is used for all values of user EIRP, then Figure 3-5 shows that E_b/N_0 is constant at 4.3 dB-Hz. Consequently, the carrier tracking error is constant, as summarized in Figure B-4b.

B.3 PHASE NOISE IN RANGE-RATE MEASUREMENTS

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Additive thermal noise introduced by the channel is not the sole source of random fluctuations that affect range-rate measurement precision. A system's transmitting and receiving oscillators, system mixing chains, AM to PM conversion in the channel, spurious vibrations, etc. all introduce phase noise components into the signal which degrade the carrier-tracking lcop performance of the receivers. To estimate the impact of phase noise on one-way navigation, consider first the doppler measurement process: the phase value ϕ extracted by the carrier-tracking loop is sampled at time t-T_{av} and at time t and then differenced. That difference $\Delta \phi$ is related to the range-rate R as $R = (\lambda/2\pi T_{av})\Delta \phi$. Since $\Delta \phi = \phi(t) - \phi(t - T_{av})$, the transfer function describing the operation is simply $F(f) = 1 - e^{-j2\pi fT_{av}}$. If the input to the doppler extractor has single-sided phase noise spectral density $S_{\phi}(f)$, then the spectral density of $\Delta \phi$ is given by

$$S_{\Delta\phi}(f) = |F(f)|^2 S_{\phi}(f) = 4 \sin^2 (\pi f T_{av}) S_{\phi}(f).$$

Since $\sigma_{\Delta\phi}^2 = 2 \int_0^{\infty} S_{\Delta\phi}$ (f) df, the input spectral phase noise density $S_{\phi}(f)$

contributed by the system can be seen to determine the range-rate mean square error due to phase noise.

In this analysis, only the effects of phase noise in the system oscillators and VCO's as propagated through the system mixing chain are considered. The key to the analysis, as discussed in [30], lies in characterizing a carrier-tracking loop by its closed loop transfer function H(f). A linear analysis of the carrier loop shows that the loop's VCO output tracks the carrier with noise contributions from two sources. One is the input phase jitter lowpass filtered by H(f), the other, the VCO's own phase jitter highpass filtered by 1-H(f). In other words, if the input phase noise spectral density is $S_{\theta}(f)$ and the phase noise spectral density of the VCO itself is $S_{\theta}'(f)$, then the coherent reference provided by the tracking loop has phase noise spectral density $S_{\theta}(f) = |H(f)|^2 S_{\theta}(f) + |1-H(f)|^2 S_{\theta}(f)$. The simple analysis explained here has been validated by extending it to model the two-way doppler tracking of TDRSS; the results are consistent with other estimates of the doppler measurement error due to phase noise.

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The forward and return links are modeled in terms of the ground and user oscillators, the tracking loops and their VCO's, and the frequency multiplies comprising the mixing chains. It is assumed that TDAS satellites will operate in the same fashion as do TDRSS satellites: a ground-generated pilot tone is tracked on-board the TDAS satellite to provide the reference tone for frequency-transition of the relayed signals. TDRSS uses a Ku-band space-to-ground link; in TDAS, both K_{u} - and K_{a} -band frequencies are being considered. Figure B-5 shows the models used here for the forward and return links, respectively. Both the user and ground station's oscillators are assumed to be 5 MHz standards; the frequency multiplies required to support return and forward link service and a Ku- or Ka-band space-to-ground link are shown in the diagrams. The Ku-band uplink, downlink, and pilot frequencies are taken from TDRSS; those for K_a -band are chosen arbitrarily. S-band forward link service centered at 2106.406 MHz (as in TDRSS) and an assumed IF frequency at the user spacecraft of 9.5 MHz yield a fixed multiply value at the user oscillator, as shown in Figure B-5a. S-band return link service centered at 2287.5 MHz similarly requires a fixed



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multiply value at the user oscillator in Figure B-5b. For return link tracking, an IF of 240 kHz into the ground carrier-tracking loop is assumed.

The generation of high-frequency signals from the ground and user frequency standards is pessimistically assumed to require frequency synthesizers and hence tracking loops and their VCO's. Noise from VCO's, of course, is also introduced at the pilot tracking loop on-board the TDAS satellite and at the ground and user receivers' tracking loops. For simplicity, all the tracking loops are assumed to be 2nd order with closed loop transfer function

$$H(w) = \frac{w_n^2 + j \sqrt{2} w_n w}{w_n^2 + j \sqrt{2} w_n w - w^2}$$

where w = $2\pi f$, w_n is the loop natural frequency, and the single-sided loop noise bandwidth $B_{L} \approx \frac{3w_{n}}{4\sqrt{2}}$ Each of the tracking loops may thus be parameterized by its loop bandwidth. The following values are chosen on the basis of present TDRSS practice, the Motorola 2nd generation TDRSS user transponder design, and arbitrary selection:

H <u>1</u> (f)	(ground frequency system):	BL = 700 Hz,
H2(f)	(TDAS frequency system):	BL = 262 Hz,
H3(f)	(user frequency system):	BL = 700 Hz,
H4(f)	(user tracking loop):	6L = 40 Hz,
$H_5(f)$	(ground tracking loop):	$B_{\rm L}$ = 70 Hz.

To complete the model, the single-sided phase noise spectral densities of the ground and user frequency standards and the VCO's are approximated as shown in Figure 8-6. As in TDRSS, the ground frequency standard is assumed to be an HP 5061A+004 cesium oscillator with its phase noise spectral density derived from the manufacturer's specifications [31]. The user frequency standard is assumed to be a precision quartz oscillator; its phase noise spectrum is estimated from the specifications for the FTS 1150 quartz standard [32]. Note that while the cesium oscillator provides superior long term stability, its short term phase fluctuations are more ł

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severe than those of the quartz oscillator. In all the figures, the spectral density is assumed to extend for offset frequencies beyond those shown in the same manner as immediately before the end of the depicted range. Only for the VCO spectral density will these high offset frequencies be significant since only the VCO output is highpass filtered by transfer functions of the form 1-H(f).

B.3.1 Forward Link Tracking

Figure B-7a explicitly shows the various filtering operations to which the phase processes ϕ_{CESIUM} , ϕ_{VCO} , and ϕ_{XTAL} in the forward link are subjected. The ground-to-TDAS delay τ_1 and TDAS-to-user delay τ_2 are accounted for by the appropriate phase shifts; τ_1 and τ_2 are approximated here as .14 seconds, corresponding to distances of 42000 kilometers. Inspection of the diagram allows the transfer functions associated with the ground, TDAS, and user segments to be identified separately. This yields the component single-sided phase noise spectral densities at the input to the user's doppler extractor:

$$\begin{split} s_{\phi_{\text{GND}}}(f) &= s_{\phi_{\text{CESIUM}}}(f) |H_4(f)H_1(f)e^{-jw(\tau_1 + \tau_2)}(K_2 - K_3H_2(f))|^2 \\ &- s_{\phi_{\text{VCO}}}(f) |(1 - H_1(f))H_4(f)e^{-jw(\tau_1 + \tau_2)}(K_2 - K_3H_2(f))|^2 \\ &= \left[s_{\phi_{\text{CESIUM}}}(f) |H_1(f)|^2 + s_{\phi_{\text{VCO}}}(f) |1 - H_1(f)|^2 \right] x \\ &|H_4(f)|^2 \left[\kappa_2^2 - 2K_2K_3\text{Re}\{H_2(f)\} + \kappa_3^2 |H_2(f)|^2 \right]; \\ s_{\phi_{\text{TDAS}}}(f) &= s_{\phi_{\text{VCO}}} |(1 - H_2(f))K_3 + H_4(f)|^2; \\ s_{\phi_{\text{USER}}}(f) &= \left[s_{\phi_{\text{XTAL}}}(f) |H_3(f)|^2 + s_{\phi_{\text{VCO}}}(f) |1 - H_3(f)|^2 \right] (419.38)^2 |H_4(f)|^2 \\ &+ s_{\phi_{\text{VCO}}}(f) |1 - H_4(f)|^2; \end{split}$$

where Re{•} denotes "the real part of {•}".



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The phase noise variances corresponding to each of these densities is obtained as simply $\sigma_{\phi_X}^2 = 2 \int_0^{\infty} S_{\phi_X}$ (f) df; the range-rate variance due to each component is similarly obtained as

$$\sigma_{R_X}^{\bullet 2} = 2 \int_0^{\infty} S_{\phi_X} (f) 4 \sin^2 (\pi f T_{av}) df$$

The results obtained from numerical integration are reported in Table 3-3.

B.3.2 RETURN LINK TRACKING

The necessary transfer functions are found from Figure B-7b in the same way as in the preceding section. As previously, the single-sided phase noise spectral density at the input to the doppler extractor may be broken into its component elements according to the phase noise sources:

$$\begin{split} S_{\phi_{\text{USER}}}(f) &= \left[S_{\phi_{\text{XTAL}}}(f) || \hat{H}_{3}(f)|^{2} + S_{\phi_{\text{VCO}}}(f) || 1 - H_{3}(f)|^{2} (457.5)^{2} \right] || H_{s}(f)|^{2}; \\ S_{\phi_{\text{TDAS}}}(f) &= S_{\phi_{\text{VCO}}}(f) || (1 - H_{2}(f)) K_{4} || H_{5}(f)|^{2}; \\ S_{\phi_{\text{GND}}}(f) &= \left[S_{\phi_{\text{CESIUM}}} || H_{1}(f)|^{2} + S_{\phi_{\text{VCO}}} || 1 - H_{1}(f)|^{2} \right] x \\ &\quad || H_{5}(f)|^{2} \left| \kappa_{4}^{2} || H_{2}(f)|^{2} + \kappa_{5}^{2} - \kappa_{5} \kappa_{4} [2 || Re\{H_{2}(f)\} \cos 2w_{\tau_{1}}] \right| \\ &\quad + 2 || Im\{H_{2}(f)\} \sin || 2w_{\tau_{1}}] \right| \\ &\quad + S_{\phi_{\text{VCO}}}(f) || 1 - H_{5}(f)|^{2}; \end{split}$$

where Im () denotes "the imaginary part of ()".

Again, the range-rate variance
$$\sigma_{R_X}^2$$
 due to each component is
 $\sigma_{R_X}^2 = 2\int_0^\infty S_{\phi_X}(f) 4 \sin^2(\pi fT_{aV})df.$

The results from numerical integration are shown in Table 3-5.

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APPENDIX C

TDAS TRACKING ANALYSIS - SUPPLEMENTARY RESULTS

C.1 INTRODUCTION

An evaluation of the potential tracking accuracy for TDAS satellites was made based on BRTS and VLBI tracking techniques. Section 4 presented a brief description of each technique, possible tracking configurations to support frontside and backside satellites, and the error analysis modeling approach and major results. This appendix provides supplementary results for BRTS tracking in Section C.2 and VLBI tracking in Section C.3. Relevant modelling data used in the analysis are given in Section C.4.

C.2 BRTS RESULTS

TDAS position errors in the prediction interval are presented in Figures C-1 through C-4 as a function of BRTS tracking epoch and tracking interval.

Figure C-1 (Frontside TDAS) illustrates the sensitivity of TDAS position errors in the prediction interval as a function of epoch time. In all cases, the minimum error bound is about 100 m, regardless of epoch time. However, the maximum error bound is cyclic with roughly a 12 hour period, and the peak to peak variation diminishes with longer tracking intervals. To minimize the position error in the 12 hour prediction interval, an epoch is chosen at the valley of the maximum error bound. Note that the valley epochs shift as the tracking interval changes. Of the four tracking intervals shown (18, 24, 30, and 36 hours), the 18 hour case has the lowest valleys (110 m), which occur near $18^{H}30^{M}$ GMT and $5^{H}30^{M}$ GMT. The 24, 30, and 36 hour cases all have valleys on the maximum error bound greater than 125 m, which occur at epochs different from the 18 hour case. For example, the valleys for the 24 hour case occur near $0^{H}30^{M}$ and $12^{H}30^{M}$ GMT, with a magnitude of 140 m. For the 24 hour prediction interval, the valleys in the maximum



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error bound are all above 140 m for all tracking intervals. Thus, for the frontside TDAS (at 100°W), the best choice is 18 hour tracking, 12 hour prediction, and epoch of $18^{\rm H}30^{\rm M}$, resulting in TDAS 1 σ position errors between 100 and 110 m.

Figure C-2 demonstrates the same relationships for the backside TDAS at 98° E, for which the best choice is 36 hour tracking, 12 hour prediction, and epoch of $12^{H}30^{M}$, resulting in TDAS 1 σ position errors between 210 and 250 m. However, note that the maximum error bound is relatively flat for the 36 hour tracking interval and remains below 300 m over all epochs. Thus, the epoch is not as critical for the 36 hour tracking case as it is for 18, 24, and 30 hour tracking cases in which the maximum errors fluctuate widely (exceeding 400 m).

Figure C-3 illustrates the TDAS position errors from the end of tracking to the end of prediction for a fixed epoch time. Figure C-3a is for TDAS at 100°W, with a tracking configuration using transponders at White Sands (WHS) and Santiago (AGO). WHS-WHS denotes a roundtrip path from WHS transmitterto TDAS to WHS ground transponder and return to TDAS to WHS receiver. Similarly, WHS-AGO denotes a roundtrip path from WHS transmitter to TDAS to AGO transponder and return to TDAS to WHS receiver. The TDAS position errors are plotted for 18, 24, 30, and 36 hour tracking intervals, all for 24 hour prediction intervals, based on an epoch time of 18^H30^M GMT. Again, note the best prediction interval occurs during the 12 hours after the end of the 18 hour tracking. Figure C-3b is for TDAS at 98°E with transponders on the 100°W TDAS (T100) and at Diego Garcia (DGA). T100-T100 represents a roundtrip path from T100 to TDAS (98°E) and return to T100.* T100-DGA represents a roundtrip path from T100 to TDAS to DGA transponder and return via TDAS to T100.* The best prediction interval for the backside TDAS is during the 12 hours after the end of the 36 hour tracking.

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For error analysis with existing ORAN capabilities the WHS-T100 link was not modelled. Instead, separate WHS-T100 path delay measurements were assumed to be subtracted from backside two-way BRTS measurements made at WHS. ORAN was set up to use T100 as the terminal location (analogous to WHS) with station errors equal to the position accuracy for frontside TDAS tracking.

Figure C-4 presents the TDAS maximum error contributors in the 12 hour prediction intervals for 18, 24, 30, and 36 hour tracking intervals. Note the position error scale is a log scale. Figure C-4A is for TDAS at 100°W, and shows solar pressure, VLBI measurement bias, and station survey errors as dominant, with magnitudes greater than 40 m. Figure C-4B is for TDAS at 98°E, and shows station errors and solar pressure uncertainties as the dominant error sources, with each contributing well above 100 m. Note that the station error reflects uncertainty due to the frontside TDAS (T100) tracking error assumed to be a constant 105 m (H = 26 m, C = 28 m, L = 98 m).

C.3 VLBI TRACKING

This section presents TDAS tracking accuracy results based on the VLBI tracking technique. Subsections C.3.1 and C.3.2 discuss results generated for the baseline and reduced error models defined in Section 4 (Table 4-1). Subsection C.3.3 presents some preliminary results for trontside TDAS tracking using VLBI stations located only in CONUS. These three subsections show the effects of tracking interval, epoch time, prediction interval, VLBI bias estimation, error sources, and data rate on TDAS uncertainty.

C.3.1 Baseline Error Model Results

A composite plot of TDAS 10 position uncertainty during the prediction interval is shown in Figure C-5 for several tracking intervals - 3, 6, 12, 18, 24, 30, 36, and 42 hrs, all using a fixed epoch time of $18^{\rm H}30^{\rm M}$ GMT. Figure C-5A shows the error profile without VLBI bias estimation. For the 3, 6, 12, 24, 30, and 42 hr tracking, the position uncertainties rise above 100 m after a few hours in the prediction interval. Only for the 18 and 36 hr tracking does the position uncertainty remain below 100 m for several hours in the prediction interval. As will be shown later, concatenation of several 18 and 36 hr tracking segments may provide 24 hours of prediction below the 100 m level.

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Figure C-5B shows the same cases as C-5A except VLBI bias is estimated. The best choices of prediction intervals are those following the 12 and 18 hr tracking. These may be concatenated to provide 24 hours of prediction below the 100 m level. Note that the position error is close to 200 m at the end of the 36 hr tracking interval.



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Figures C-6 and C-7 display the TDAS maximum error contributors for various prediction and tracking intervals with and without VLBI bias estimation. As mentioned above, the 18 and 36 hr tracking intervals are the best choices for either 6 or 12 hr prediction intervals. In Figure C-6A for 6 hr tracking note that VLBI bias station survey and noise errors are dominant, while for long tracking intervals (24, 30, and 36 hrs) solar pressure error is dominant. For intermediate tracking intervals (12, 18 hrs) VLBI bias, solar pressure, and station survey errors are dominant. As the prediction interval increases, however, solar pressure ultimately dominates. Note that in all of these cases, the gravity and tropospheric effects are relatively insignificant (less than 20 m).

Figure C-7 is based on VLBI tracking with bias estimation. As noted before, the 12 and 18 hr tracking intervals are the best choices for eitner 6 or 12 hr prediction intervals. With VLBI bias estimated, atkar related errors (station survey and tropospheric errors) are also reduced whereas noise and solar pressure error increase, particularly noise. The overall result, however, is a decrease in the total error.

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Figure C-8 shows the sensitivity of the TDAS position error profile to epoch time of day at the beginning of an 18 hr tracking interval with VLBI bias estimation. For each prediction interval (1, 3, or 6 hr), the minimum and maximum error bounds are plotted. The best epoch choices are $6^{H}30^{M}$ GMT (TDAS local midnight) and $18^{H}30^{M}$ GMT (TDAS local noon). The minimum error bound is in the 48-55 m range for all three prediction intervals. The maximum error bound increases dramatically from 1 hr prediction to 6 hr prediction when the epoch is not near $6^{H}30^{M}$ or $18^{H}30^{M}$. For these best epochs, there is little change in maximum error bound as the prediction interval increases from 1 to 6 hr.

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C.3.1.1 <u>Impact of TDAS Location</u>. Figure C-9 illustrates that TDAS position uncertainty is insensitive to TDAS geographic location. Four prediction profiles are plotted, based on TDAS longitudes of 100°W, 98°E, 171°W, and 41°W. Both the 18 and 36 hr tracking cases are shown. The small differences among each set of four profiles can be attributed to geometrical differences in the tracking station configurations.

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C.3.1.2 <u>Effect of VLBI Data Rate</u>. Table C-1 is a tabulation of minimum and maximum TDAS position uncertainties as a function of data rate, tracking interval, and prediction interval for an epoch of $18^{H}30^{M}$ GMT. For a 6 hr tracking interval, the effect of data rate is dramatic with the maximum error increasing by more than 2:1 as the data period increases from 60 sec to 3600 sec. For a 12 hr tracking interval, the factor is about 20%. For longer tracking intervals (18, 24, and 36 hours), there are virtually no differences in position errors for data periods of 60, 1800, or 3600 sec. This has significant operational importance since VLBI measurements data collection and processing need occur no more frequently than once per hour.

C.3.1.3 <u>Prediction Interval Concatenation</u>. Figures C-10 and C-11 demonstrate how it is possible to concatenate several prediction segments to achieve better TDAS position performance over a 24 hr period compared with a single segment. The 18 hr tracking segments are shown in Figure C-10A. Two start times are indicated - S/C local midnight and S/C local noon on 3/22. At the end of tracking of these two cases, the 9 hr prediction profiles may be used, i.e., on 3/23 the interval between $0^{H}30^{M}$ and $9^{H}30^{M}$ and the second interval between $12^{H}30^{M}$ and $21^{H}30^{M}$. Gaps from $9^{H}30^{M}$ to $12^{H}30^{M}$ and from $21^{H}30^{M}$ to $24^{H}30^{M}$ may be filled using prediction intervals based on 36 hr tracking starting at S/C local noon and S/C local midnight on 3/21 and 3/22, respectively (see Figure C-10B). When all these segments are concatenated in Figure C-11A, the result is a 24 hr prediction interval in which the TDAS position error remains below 85 m.

The same procedure may be used when the VLBI bias is estimated (see Figure C-11B). By concatenating multiple prediction segments following 18 and



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TABLE C-1 MIN/MAX POSITION UNCERTAINTY (M) IN PREDICTION INTERVAL VS. VLBI DATA RATE AND TRACKING INTERVAL*

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PREDICTION	TRACKING	VLI	BI DATA RATE (S	SEC)
(HOURS)	(HOURS)	60	1800	3600
	6	66/180	92/437	101/520
	12	49/110	52/127	53/136
6	18	53/63	53/65	53/67
	24	53/131	53/128	53/126
	36	71/92	71/90	71/89
	6	66/481	92/1108	101/1318
	12	49/180	52/210	53/227
12	18	53/131	53/130	53/129
-	24	53/219	53/217	71/134
	36	71/131	71/133	71/134

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WITH BASELINE ERROR MODEL

* TRACKING CONFIGURATION: TDAS @ 100°W; STNS: WHS-AGO, WHS-HAW EPOCH: 86/3/21, 18H30M

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12 hr tracking intervals, an error profile that remains below 70 m can be obtained. Also shown in Figures C-11A and B are the radical, cross track, and along track components of TDAS position error. In both cases, the along track component dominates.

C.3.2 Reduced Error Model Results

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TDAS position uncertainties computed from the baseline and reduced error models are compared in Figure C-12. Four tracking intervals were used -1, 3, 6, and 12 hrs. For the baseline model, longer tracking intervals produce better prediction profiles, as observed in Subsection C.3.1. However, for the reduced error model, shorter tracking intervals appear advantageous. A 6 hr tracking interval followed by a 1 hr prediction interval seems to be an optimal choice. As will be demonstrated later, 1 hr segments may be concatenated to form a 24 hr prediction interval with TDAS position errors between 10 and 12 m, a significant improvement over the baseline model.

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TDAS maximum position error contributors are shown in Figure C-13 for the 1, 3, 6, and 12 hr tracking intervals. Note that the tropospheric error contribution is significant here, but decreases, while the solar pressure contribution increases with tracking interval. A crossover between these two error sources occurs around a 6 hr tracking interval. All the other error contributors are relatively small in this region.

C.3.2.1 <u>Impact of Tracking Data Rate</u>. Table C-2 lists minimum and maximum TDAS position uncertainties as a function of data rate, tracking interval, and prediction interval. For 1 hr tracking, the 3600 sec period (1 measurement per hr) is inadequate for orbit estimation (number of measurements less than the number of unknowns being solved for). For 3 hr tracking the 1/hr rate produces a maximum error about twice that of the 1/min rate. For 6 or 12 hr tracking, changing the data rate from 1/min to 1/hr has little impact on the results.

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TABLE C-2 MIN/MAX POSITION UNCERTAINTY (M) IN PREDICTION INTERVAL VS. VLBI DATA RATE AND TRACKING INTERVAL*

PREDICTION	TRACKING	VLI	BI DATA RATE (S	EC)
(HOURS)	(HOURS)	60	1800	3600
	1	96/114	333/398	x
7	3	14/16	19/23	23/29
*	6	11.5/18.7	11.6/18.9	11.7/19.3
	12	11.6/15.1	11.4/14.7	11.2/14/4
	1	96/212	333/740	X
,	3	14/27	19/43	23/56
3	6	11.5/18.7	11.6/18.9	11.7/19.3
	12	11.6/28.2	11.4/27.6	11.2/27.1

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WITH REDUCED ERROR MODEL

* TRACKING CONFIGURATION: TDAS @ 100°W; STNS: WHS-AGO, WHS-HAW EPOCH: 86/3/21, 18^H30^M

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C.3.2.2 <u>Prediction Interval Concatenation</u>. Figure C-14 illustrates how a series of 1 hr prediction segments may be concatenated to form a 24 hr prediction profile. Each segment is based on 6 hours of tracking data. In (A), the dominant error source is the tropospheric error (15%),* and the profile is relatively flat over 24 hours with a mean TDAS position error near 11.5 m. In (B), the effect of a possible reduction in tropospheric error to 5% is shown with TDAS position error reduced to the 5-8 m level. Solar pressure is the dominant contributor as is evident from the 12 hr cycle.

C.3.3 CONUS-Based VLBI Tracking Results

Some preliminary results on TDAS position uncertainty based on CONUS tracking sites rather than global sites are presented in Figures C-15 and C-16. The CONUS sites used were GSFC, HSN (Houston), and SUN (Sunnyvale). Figure C-15A shows the prediction error profiles generated with the <u>baseline</u> model for 6-48 hr tracking intervals in 6 hr steps. As observed in Section C.3.1, longer intervals give better performance than shorter intervals with measurement and station survey errors at the baseline <u>model</u> levels. Figure C-15B indicates that the maximum TDAS position error during a 6 hr prediction interval is dominated by either solar pressure or VLBI bias for the tracking intervals considered.

Figure C-16 demonstrates how several prediction interval segments can be concatenated to form a 24 hour prediction error profile. In (A), with segments based on 24 and 48 hr tracking intervals, the TDAS position uncertainty ranges between 80 and 125 m. In (B), VLBI bias is estimated, and the TDAS position uncertainty decreases to the 50-75 m level.

C.4 MODELLING DATA

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Tables C-3 and C-4 list various modelling data used in the analysis.

A 15% uncertainty in tropospheric delay modelling was assumed in the VLBI error model (Table 4-1).

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	EL ASSUMPTIONS	1º VALUES	(BASELINE MODEL)	н 6.0	З н	З М	15%	10-7 10-7 of GEM5 - GEM	10%	
TABLE C-3	acking error mode	BRIS		2 H/.13 CH/s	10 M/0 CM/S	Ян				(SAME AS TDRS).
	TDAS TRI	r source		NOISE	BIAS	STATIONS (X, Y, Z)	TROPOSPHERIC Refraction	GM GRAV, POT,	SOLAR PRESSURE*	sht = .042 m²/kg (.ations inc.
• ·		ERRON		MEÂSUREMENT - brts (r/r)	- VLBI (AR)	LOCAL		DYNAMIC		* TDAS AREA/METO STANFORD TELECOMMUNIC

	TABLE C-4		
EXISTING AN	D POTENTIAL GROU	ND STATIONS	
STATIONS LOCATIONS	STATION CODE	LATITUDE* (DEG.)	LONGITUDE (E) (DEG.)
ASCENSION ISLAND	ACN	0*8-	345.7
SANTIAGO, CHILE	AGO	-33,2	289.3
ALICE SPRINGS, AUSTRALIA	ALS	-23.8	133.9
AMERICAN SAMDA	AMS	-14.3	, 189.3
JOHANESBURG, S. AFRICA	BUR	-25.9	27.7
DIEGO GARCIA	DGA	-7.0	72.0
EASTER ISLAND	EIL	-27.5	251.0
GODDARD SPACE FLIGHT CENTER	GSFC	39.0	283.2
GUAM	6WM	13,3]]tht.7
HAWAII	HAW	22.1	200.3
HOUSTON, TEXAS	NSH	29.0	265.0
MADRID, SPAIN	MAD	40.5	355,8
MERRITT ISLAND, FLORIDA	MIL	28.5	279.3
ORRORAL VALLEY, AUSTRALIA	ORR	-35.6	149.0
SUNNYVALE, CALIFORNIA	NNS	37.0	238.0
VANDENBURG, AFB, CALIFORNIA	VAN	34.7	239.5
WHITE SANDS, NEW MEXICO	SHW	32.5	253.4
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APPENDIX D

USER NAVIGATION PERFORMANCE RESULTS - SEQUENTIAL DATA PROCESSING

TDAS user navigation performance was evaluated in terms of potential OD/TD accuracy for the three one-way tracking alternatives defined in Section 2. The objective was to assess several user orbit types and the impact of two and three satellite TDAS constellations, the tracking schedule and two algorithms for tracking data processing (sequential and sliding batch). This section contains the detailed performance results based on <u>sequential</u> <u>processing</u> of tracking data*. The error modeling approach and major findings were presented in Section 5.

Cases for evaluation were defined originally for three TDAS constellation options (see Figure 5.2) and six user orbit types (see Figure 5.3). The results that follow cover 12 cases: four of the user orbits (A, D, E and F)^{**} with each of the three TDAS constellations.^{**} Each case is presented in a standard format:

- <u>Tracking Schedule</u> This shows tracking contact intervals over 24 hours for Schedules I and II and for beacon tracking (see Table 5-1). An accompanying user ground trace indicates coverage outages due to a zone of exclusion (if any).
- <u>User Position & Time Accuracy Profiles</u> These show the 1_σ errors corresponding to the 24 hour tracking schedules and a subsequent 6 or 12 hour prediction-only interval.

Summary tables (D-1 thru D-4) included at the end of Appendix D record the peak errors and identify the major error contributors for each tracking alternative.

Performance results for sliding batch processing are in Appendix E.

^{**} Individual cases are identified by a two element mnemonic; e.g., A2 indicates user orbit Type A and TDAS constellation Option 2. Evaluation of user orbit types, B (28°, 600 km) and C (28°, 1000 km) remains to be completed.



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	171/41/746	A3	50	44	177	150	419	>1000	>1000	232	547	>1000	>1030
	171/41	ā	64	117	290	207	449	>1000	>1000	327	607	>1000	>1000
97° 200 FW	100/98	D2	116	128	396	260	475	>1000	>1000	408	735	>1000	>1000
	171/41/74E	D3	31	11	248	209	427	>1000	>1000	265	532	>1000	>1000
	171/41	E	42	44	47	44	47	75	135	53	23	23	113
98°	100/98	. 23	41	48	47	47	50	65	126	23	23	5	117
	171/41/74E	8	31	55	44	38	65	06	161	48	48	69	136
	171/41	E	38	31	37	27	ୟ	£	11	33	34	54	92
98°	100/98	F2	45	54	49	61	19	19	19	37	38	65	87
2001	171/41/74E	E	31	32	38	27	32	44	83	53	34	44	85

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	171/41/74E	ę	z			#	-	8	=	2	•		
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97° 200 KH	100/98	D2	H, H, E, D	N.D	¥	н, D, N	D.H.N	3	1	N, O	H C) x	k at
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	171/ 4 1/74E	8	N,H,E [†]	2	3	1	2	*	2	E	2	1	•
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	ONFIGURATION		BEACON	SCHE	DULE		SCHED	ULE II			SCHED	OLE I	
liced	TDBC					NAV	UPLOAD	RATE (HOU	IRS)	NAV	UPLOAD 1	ATE (HOU	IRS)
ORBET	CONST	CASE		Ħ	I	1.5	е	9	12	1.5	m	9	12
	171/41	Al	125	360	690	555	865	1520	3020	870	1270	2100	3910
28° 200 kh	36/001	AZ	35	260	565	440	750	1405	2890	620	960	1685	3290
	171/41/74E	A3	140	265	565	455	760	1405	2870	525	825	1465	2935
	171/41	10	160	290	500	820	1320	2350	4525	520	755	1280	2540
97° 200 KH	100/98	02	06	250	410	325	525	975	2100	455	700	1240	2535
	171/41/74E	Eđ	120	445	825	435	0EL	1345	2760	455	685	1180	2375
	171/41	E	230	165	230	750	1300	2340	4150	580	810	1280	2285
98° 600 km	100/98	E2	115	210	250	230	245	560	1315	235	340	560	1275
	171/41/74E	Ē	155	330	275	215	400	935	2035	300	425	750	1580
	171/41	Е	260	325	460	505	835	1300	2655	440	720	1165	2275
98° 1000 km	100/98	F2	230	295	270	200	300	670	1675	230	280	230	1120
	171/41/74E	E	170	345	390	415	360	1300	2725	495	865	1445	2955

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	171/41	10	N,E	×	N,H	H,N	н•и	N,H,D	N,H,D,O	z	Z	z	N,0
97° 200 KH	100/98	02	z	N,H,E,D	z	Ν,Ε	и,E	R,E	N,0,E	2	•	N	H,0,H
	171/41/74E	03	N,E,H	z	H,N	И.Н	N,H	И,Н,О	N,K,O	я		3	И,0
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98° 600 km	100/98	E2	E.N.H	N.H	N.N	N.H.E	N,H	H.N.O.E	H.O.N.E	н, Е	N.0.E	И,О	0°N
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APPENDIX E

USER NAVIGATION PERFORMANCE RESULTS - SLIDING BATCH DATA PROCESSING

TDAS user navigation performance was evaluated for two of the one-way tracking alternatives (FLST and RLST) assuming sliding batch processing of tracking data. The error modeling approach and major findings were discussed in Section 5; this section presents further details of the results obtained.

Cases for evaluation were selected from among the TDAS constellation options and user orbit types defined in Figures 5-1 and 5-2. The results that follow cover half of the cases presented in Appendix D for sequential processing — the same user orbits types (A,D,E,F) and two of the TDAS satellite constellation options (1,2).

For each case considered, two types of user position and time accuracy profiles were generated:

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<u>Single Batch Profiles</u> - These were computed to assess the sensitivity of user OD/TD accuracy in the prediction mode to tracking interval selection. Durations of 6, 12, 18, and 24 hours were evaluated for each tracking schedule (I and II).* Figures E-1 and E-2 are examples. The tracking interval yielding the lowest peak error in the first 6 hours from end-of-tracking was selected for generating sliding batch profiles in each case.

^{*} Schedules I and II for a 24 hour tracking interval are identical to those given in Appendix D for corresponding cases. Scheduled contacts for shorter tracking intervals were simply deleted.

<u>Sliding Batch Profiles</u> - These are comprised of prediction interval segments corresponding to a sequence of single batch solutions, as defined earlier in Figure 5-4. Twenty-four hour profiles were constructed based on a batch interval (P) of 3 hours and a computation/data handling interval (C) of 1.5 hours. Figure E-3 for Case A1 is an example.*

Summary tables (E-1 thru E-4) included at the end of Appendix E record the peak errors and identify the major error contributors for each tracking alternative.

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^{*} The time-from-epoch scale on all plots has the same interpretation as in Appendix D, i.e., the time elapsed since the user's initial location, (0°N, 0°E). Individual batch epochs were adjusted backward or forward depending on tracking interval in order to fit a given prediction segment slot.

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<u>USER NAVIGATION PERFORMANCE SUMMARY - MAX CLOCK ERROR (¹SEC) (SLIDING BATCH DATA PROCESSING)</u> TABLE E-3

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	"171°/41°	10	3.9	4.8	4.3	5.3	2.9	3.9	4 .8	3.3	4.3	5.3
200 KH	100°/98°E	102	8.7	10.7	5.6	6.8	6_6	8.7	10.7	4 .5	5.6	6.8
	1/1°/41°		2.7	2.9	2.8	0"E	2.4	2.7	2.9	2.5	2.8	3.0
-96 600 KH	100°/98°E	E2	2.6	2.9	3.0	3.2	2.4	2.6	2.8	2.9	3.0	3.2
98° 1000 km	100°/98°E	5	2.7	2.9	3.0	3.3	2.3	2.7	2.9	2.7	3.0	3.3

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APPENDIX F

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GLOSSARY OF ACRONYMS

BRTS	Bilateration Ranging Transponder System
CONUS	Contiguous United States
DSN	Deep Space Network
EOT	End of Tracking
FLBT	Forward Link Beacon Tracking
FLST	Forward Link Scheduled Tracking
GPS	(NAVSTAR) Global Positioning System
KSA	Ku-Band Single Access
LSA	Laser Single Access
MA	Multiple Access
MCC	Mission Control Center
NASCOM	NASA Communication Network
NCC	Network Control Center
OD	Orbit Determination
OD/TD	Orbit and Time Determination
OSCF	Orbit Support Computing Facility (GSFC)
POCC	Project Operations Control Center
RLST	Return Link Scheduled Tracking
SA	Single Access
SMA	S-Band Multiple Access
TD	Time Determination
VLBI	Very Long Baseline Interferometry
WSA	W-Band Single Access
WSN	White Sands (NM)
∆VLBI	Differential VLBI (using two signal sources)

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APPENDIX G

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