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Enthalpy Damping for the Steady Euler Equations

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SUMMARY. For inviscid steady flow problems where the enthalpy is constant at steady state, it has been proposed by Jameson, Schmidt, and Turkel to use the difference between the local enthalpy and the steady state enthalpy as a driving term to accelerate convergence of iterative schemes. This idea is analyzed here, both on the level of the partial differential equation and on the level of a particular finite difference scheme. It is shown that for the two-dimensional unsteady Euler equations, a hyperbolic system with eigenvalues on the imaginary axis, there is no enthalpy damping strategy which can move all the eigenvalues into the open left half plane. For the numerical scheme, however, the analysis shows and examples verify that enthalpy damping can be effective in accelerating convergence to steady state.

1. INTRODUCTION

Rapidly convergent numerical schemes for steady state solutions of the inviscid compressible fluid dynamics equations, the Euler equations, have been studied intensively in the last few years. Algorithms and codes have been developed by Jameson, Schmidt, and Turkel (non-multigrid) (Ref 1); Jameson (multigrid) (Ref 2), Rizzi (Ref 3); Pulliam (Ref. 4), Lerat, Sidès, and Daru (Ref 5), Johnson (Ref. 6); Ni (Ref. 7); Agarwal and Deese (Ref 8); Osher and Chakravarthy (Ref. 9), Mulder and van Leer (Ref 10), and by Buning and Steger (Ref 11), among others. One of the most successful and widely used algorithms is embodied in the FLO52 codes of A. Jameson. The algorithm used for the Euler equations in these codes is based on an explicit multistage time-stepping scheme, with (for steady state problems) acceleration devices such as local time step selection, implicit residual averaging, multigrid, and enthalpy damping. It is the purpose of this paper to provide an analysis of enthalpy damping for the two-dimensional Euler equations.

The original derivation of enthalpy damping by Jameson, Schmidt, and Turkel (Ref. 1) was based on reducing the unsteady Euler equations to the unsteady potential equation (a wave equation), introducing frictional damping to modify the equation to the telegraph equation, and translating the new equation back to the Euler equations. The next section begins by reviewing this derivation. It can be verified that enthalpy damping is effective in enhancing convergence for the Euler equations, and an example calculation is included.

The two-dimensional Euler equations with enthalpy damping are studied in section 3. The Euler equations are a system of partial differential equations of the form

$$w_t + \partial_x \mathcal{E}(w) + \partial_y \mathcal{F}(w) + \alpha \mathcal{H}(w) = 0 \quad (1.1)$$

where the enthalpy damping is embodied in the undifferentiated term \mathcal{H} , and α determines the strength of the damping term. We linearize (1.1) around a state w_0 with constant enthalpy, obtaining a linear variable coefficient system

$$w_t + \partial_x (A(w_0(x, y))w) + \partial_y (B(w_0(x, y))w) + \alpha C(w_0(x, y))w = 0 \quad (1.2)$$

where w is the perturbation from w_0 and $A = \partial \mathcal{E} / \partial w$, $B = \partial \mathcal{F} / \partial w$, $C = \partial \mathcal{H} / \partial w$. We study the frozen coefficient problem

$$w_t + A \partial_x w + B \partial_y w + \alpha C w = 0 \quad (1.3)$$

Fourier transforming (1.3) gives the system

$$\frac{d\hat{w}}{dt} + i(\omega_1 A + \omega_2 B)\hat{w} + \alpha C \hat{w} = 0 \quad (1.4)$$

where $w(x, y, t) = e^{i(\omega_1 x + \omega_2 y)} \hat{w}(t)$. For $\alpha = 0$, the system (1.3) is hyperbolic and all solutions \hat{w} of (1.4) are of the form $\hat{w} = \sum_j e^{\lambda_j t} \hat{w}_j(0)$ where λ is purely imaginary. For $\alpha > 0$ the natural conjecture is that the enthalpy damping helps by pulling the λ_j into the left half of the complex plane. We show that of the four eigenvalues λ_j for any pair (ω_1, ω_2) , one eigenvalue always remains on the imaginary axis, so the system (1.3) does *not* have the property

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that all its solutions decay to zero. The analysis also suggests a form of the enthalpy damping term for the energy equation more natural than that of Jameson, Schmidt, and Turkel (Ref. 1).

If the Euler equations at the partial differential equation level do not have decaying solutions when enthalpy damping is in effect, why do the numerical results show that enthalpy damping is useful? This question is addressed in section 4. To get a flavor of the idea, consider the system of ordinary differential equations

$$\frac{du}{dt} + Au = 0 \quad (1.5)$$

where the complex matrix A is assumed to be diagonalizable with all its eigenvalues on the imaginary axis, say at $i\lambda_1, i\lambda_2, \dots$, with $0 < \lambda_1 < \lambda_2 < \dots$. Solutions to (1.5) do not decay to zero as $t \rightarrow \infty$. However, if (1.5) is integrated with the implicit Euler method

$$u_{n+1} - u_n + \Delta t A u_{n+1} = 0 \quad (1.6)$$

the numerical solution u_n will go to zero as $n \rightarrow \infty$ because the eigenvalues of A are contained in the interior of the stability region of the method. Furthermore, the eigenvector associated with the eigenvalue $i\lambda_1$ will decay slowest because its damping factor is closest to 1. Thus if the matrix A is somehow modified so that the eigenvalue λ_1 moves to the left, such that the modified eigenvalue is farther from 1 in the complex plane than the original eigenvalue, and such that all the other eigenvalues are unchanged, then the numerical solution of the modified system will decay to zero faster than the numerical solution of the original system. In section 4 we attempt to show that roughly the same phenomenon occurs with the Euler equations and an explicit multistage integration scheme (i.e., the eigenvectors obstructing the convergence are associated with eigenvalues which are moved to the left when enthalpy damping is used). Thus, even though one of the eigenvalues does not move off the imaginary axis, the convergence rate can still improve. An attempt is made to quantify the improvement in convergence rate and to find an optimal damping parameter α .

The major omission of this work is the neglect of numerical dissipation. The numerical scheme is actually an approximation not to (1.1), but to

$$w_t + \partial_x \mathcal{E}(w) + \partial_y \mathcal{F}(w) + \alpha \mathcal{H}(w) + \epsilon \mathcal{D}(w) = 0 \quad (1.7)$$

where \mathcal{D} is a dissipation operator (e.g., a blend of second- and fourth-order dissipation terms). The inclusion of a dissipation term makes the analysis much more cumbersome, a point which will be touched on again. If an analysis including the dissipation term could be carried out, results more relevant to the performance of the computer codes could be obtained.

The author has had the benefit of several illuminating discussions with Eli Turkel in the course of this work. Many of the algebraic computations in sections 3 and 4 were carried out with *vaxima*, the VAX version of *Macsyma* (Ref. 12). Some of the computations would have been extremely difficult or impossible without *vaxima*.

2. DERIVATION OF ENTHALPY DAMPING

This section reviews the derivation of enthalpy damping via the unsteady potential equation and telegraph equation as originally sketched in Jameson, Schmidt, and Turkel (Ref. 1). The two-dimensional Euler equations in conservation form for a perfect gas may be written

$$\begin{aligned} \rho_t + (\rho u)_x + (\rho v)_y &= 0 \\ (\rho u)_t + (\rho u^2 + p)_x + (\rho uv)_y &= 0 \\ (\rho v)_t + (\rho uv)_x + (\rho v^2 + p)_y &= 0 \\ (\rho E)_t + (\rho uH)_x + (\rho vH)_y &= 0 \end{aligned} \quad (2.1)$$

where ρ is the density, u and v are Cartesian velocity components, E is total energy per unit mass, p is pressure, and H is the enthalpy, defined by

$$H = E + p/\rho$$

One can verify that

$$\frac{D}{Dt}H = \frac{1}{\rho} \frac{\partial p}{\partial t}$$

where D/Dt is the material derivative $D/Dt := \partial/\partial t + \mathbf{v} \cdot \nabla$. (The notation $a := b$ means a is defined as b) Hence in steady flow H is constant along streamlines, and if flow conditions are uniform at upstream infinity, then H is everywhere constant at steady state.

In nonconservative form the Euler equations may be written

$$\begin{aligned} \rho_t + u\rho_x + v\rho_y + \rho(u_x + v_y) &= 0 \\ u_t + uu_x + vu_y + p_x/\rho &= 0 \\ v_t + uv_x + vv_y + p_y/\rho &= 0 \\ E_t + uE_x + vE_y + \frac{p}{\rho}(u_x + v_y) + \frac{1}{\rho}(up_x + vp_y) &= 0 \end{aligned} \quad (2.2)$$

The fourth equation here can be replaced by the transport equation for entropy

$$S_t + uS_x + vS_y = 0$$

The choice of the fourth equation affects the form of enthalpy damping that is derived

If the flow is barotropic (i.e., $p = p(\rho)$) one can show by differentiating the x -momentum equation with respect to y , the y -momentum equation with respect to x , and subtracting the latter from the former, that the vorticity $\omega = v_x - u_y$ satisfies

$$\omega_t + (\omega u)_x + (\omega v)_y = 0$$

With this and the use of the continuity equation one can show that

$$\frac{\partial(\omega/\rho)}{\partial t} + u\partial_x(\omega/\rho) + v\partial_y(\omega/\rho) = 0 \quad (2.3)$$

Thus if the fluid is irrotational (i.e., $\omega = 0$) at time $t = 0$ then it is irrotational for all $t > 0$. For an irrotational fluid a velocity potential ϕ may be introduced, where $\phi_x = u$ and $\phi_y = v$. Then the momentum equations may be integrated to give the unsteady Bernoulli equation

$$\frac{\partial\phi}{\partial t} + \frac{u^2 + v^2}{2} + \int \frac{dp}{\rho} = G(t) \quad (2.4)$$

where G is an arbitrary function of t . If the flow is homentropic (entropy everywhere constant), then we have $p = A\rho^\gamma$ and $\int dp/\rho = \gamma p/[(\gamma - 1)\rho]$. For an ideal gas we have $p = (\gamma - 1)\rho[E - (u^2 + v^2)/2]$ where γ is the ratio of specific heats. Hence

$$H = E + p/\rho = \frac{\gamma}{\gamma - 1} \frac{p}{\rho} + \frac{u^2 + v^2}{2}$$

and the unsteady Bernoulli equation becomes

$$\frac{\partial\phi}{\partial t} + H = G(t)$$

For exterior flow problems, if the conditions are uniform at infinity, this becomes

$$\frac{\partial\phi}{\partial t} + H = H_\infty \quad (2.5)$$

where H_∞ is the enthalpy at infinity

An alternate form of the unsteady Bernoulli equation (2.5) is

$$\phi_t + \frac{\phi_x^2 + \phi_y^2}{2} + \frac{\gamma}{\gamma - 1} A\rho^{\gamma-1} = H_\infty$$

To obtain the unsteady potential equation, this equation is differentiated with respect to t , x , and y , and solved for ρ_t , ρ_x , ρ_y . These expressions for ρ_t , ρ_x , and ρ_y are substituted into the mass conservation equation. Using the relation giving the speed of sound as $c^2 = \gamma p / \rho$, one obtains the unsteady potential equation

$$\phi_{tt} + 2u\phi_{xt} + 2v\phi_{yt} = (c^2 - u^2)\phi_{xx} - 2uv\phi_{xy} + (c^2 - v^2)\phi_{yy} \quad (2.6)$$

For constant u and v the change of coordinates $\xi = x - ut$, $\eta = y - vt$ transforms this to the wave equation

$$\phi_{tt} = c^2(\phi_{\xi\xi} + \phi_{\eta\eta})$$

A related equation (Ref. 13) is the *telegraph equation*

$$\phi_{tt} + \alpha\phi_t = c^2(\phi_{\xi\xi} + \phi_{\eta\eta})$$

For $\alpha > 0$ the solutions of this equation decay as $t \rightarrow \infty$. Transforming this equation back via $x = \xi + ut$, $y = \eta + vt$ we obtain the damped version of (2.6),

$$\phi_{tt} + 2u\phi_{xt} + 2v\phi_{yt} + \alpha\phi_t = (c^2 - u^2)\phi_{xx} - 2uv\phi_{xy} + (c^2 - v^2)\phi_{yy} \quad (2.7)$$

This is the equation one would have obtained from the mass conservation equation if the mass conservation equation had been

$$\rho_t + u\rho_x + v\rho_y + \rho(u_x + v_y) = \alpha\rho\phi_t$$

Now, ϕ is unknown, but the unsteady Bernoulli equation (2.5) allows us to replace ϕ_t by $H_\infty - H$, obtaining the modified mass conservation equation

$$\rho_t + u\rho_x + v\rho_y + \rho(u_x + v_y) + \alpha\rho(H - H_\infty) = 0 \quad (2.8)$$

Using this in place of the first equation of (2.2) and recombining into conservation form, one obtains the system of equations

$$\begin{aligned} \rho_t + (\rho u)_x + (\rho v)_y + \alpha\rho(H - H_\infty) &= 0 \\ (\rho u)_t + (\rho u^2 + p)_x + (\rho uv)_y + \alpha\rho u(H - H_\infty) &= 0 \\ (\rho v)_t + (\rho uv)_x + (\rho v^2 + p)_y + \alpha\rho v(H - H_\infty) &= 0 \\ (\rho E)_t + (\rho uH)_x + (\rho vH)_y + \alpha\rho E(H - H_\infty) &= 0 \end{aligned} \quad (2.9)$$

If the entropy equation is used as the fourth equation of (2.2), recombining into conservation form produces a different fourth equation in (2.9). Then (2.9) is replaced by

$$\begin{aligned} \rho_t + (\rho u)_x + (\rho v)_y + \alpha\rho(H - H_\infty) &= 0 \\ (\rho u)_t + (\rho u^2 + p)_x + (\rho uv)_y + \alpha\rho u(H - H_\infty) &= 0 \\ (\rho v)_t + (\rho uv)_x + (\rho v^2 + p)_y + \alpha\rho v(H - H_\infty) &= 0 \\ (\rho E)_t + (\rho uH)_x + (\rho vH)_y + \alpha\rho H(H - H_\infty) &= 0 \end{aligned} \quad (2.10)$$

In the next section we will analyze a general form of enthalpy damping which includes both (2.9) and (2.10).

In summary, for subsonic irrotational flow, systems (2.9) and (2.10) give rise (locally) to the telegraph equation while the system (2.1) gives rise (locally) to the wave equation. Furthermore, we may hope that steady states of (2.9) and (2.10) are steady states of (2.1), and vice versa. It is clear that any steady state of (2.1) is also a steady state of (2.9) and (2.10) (since steady states of (2.1) have constant enthalpy). It is not clear that steady states of (2.9) are steady states of (2.1). It is true, however, that steady solutions of (2.10) have enthalpy constant along streamlines, hence if flow conditions are uniform at upstream infinity then steady solutions of (2.10) have enthalpy constant everywhere, hence steady solutions of (2.10) are also steady solutions of (2.1). To see that steady solutions of (2.10) have enthalpy constant along streamlines, multiply the first equation of (2.10) by H and subtract the result from the fourth equation of (2.10). If the time derivatives are zero one finds that $uH_x + vH_y = 0$, hence H is constant along streamlines

In place of the fourth equation of (2.9) we can use

$$(\rho(E - H_\infty))_t + (\rho u(H - H_\infty))_x + (\rho v(H - H_\infty))_y + \alpha \rho(E - H_\infty)(H - H_\infty) = 0 \quad (2.11)$$

which is obtained by multiplying the first equation of (2.9) by H_∞ and subtracting from the last equation of (2.9).

It is suggested by Jameson, Schmidt, and Turkel (Ref. 1) that a forcing term $\alpha \rho H(H - H_\infty)$ in the fourth equation of (2.9) "can be destabilizing," and it is suggested that this term be replaced with $\alpha(H - H_\infty)$. The computer codes seen by this author use a forcing term $\alpha \rho(H - H_\infty)$ in (2.11); i.e., $E - H_\infty$ is replaced by 1 in the fourth term of (2.11). This *ad hoc* fix seems undesirable, if only on dimensional grounds. One of the purposes of this paper is to rationally derive a reasonable form for the forcing term in the energy equation, thereby clearing up some of the confusion that seems to exist.

Figure 2.1 shows a comparison of a Euler equation calculation with and without enthalpy damping. The code used was FLO52R (supplied by A. Jameson) with the multigrid and implicit residual averaging options turned off. The initial solution was a partially converged solution obtained by grid continuation (interpolation from a coarser grid). The enthalpy damping factor α was 0.05. The convergence rate improved from 0.9966 without enthalpy damping to 0.9954 with enthalpy damping. Jameson, Schmidt, and Turkel (Ref. 1) show an example of improvement in convergence rate from 0.9909 to 0.9863. The greater gain reported there may perhaps be due to the smaller number of grid points (128×32) or other mesh differences, differences in the boundary conditions, or differences in implementation of the numerical algorithm.

3 ANALYSIS OF ENTHALPY DAMPING FOR THE DIFFERENTIAL EQUATION

Consider the system of partial differential equations

$$w_t + \partial_x \mathcal{E}(w) + \partial_y \mathcal{F}(w) + \alpha \mathcal{H}(w) = 0 \quad (3.1)$$

where

$$\begin{aligned} w &= (\rho, \rho u, \rho v, \rho E)^T \\ \mathcal{E} &= (\rho u, \rho u^2 + p, \rho uv, \rho uH)^T \\ \mathcal{F} &= (\rho v, \rho uv, \rho v^2 + p, \rho vH)^T \\ \mathcal{H} &= (\rho(H - H_\infty), \rho u(H - H_\infty), \rho v(H - H_\infty), \rho \tilde{h}_4(H - H_\infty))^T \end{aligned} \quad (3.2)$$

Here we have taken a fairly general form of the fourth component of \mathcal{H} , with $\tilde{h}_4 = \tilde{h}_4(w)$. On dimensional grounds, \tilde{h}_4 should have the same units as E .

The first step is to linearize about a steady solution w_0 , and since we are interested in steady solutions with constant enthalpy, we will assume w_0 has constant enthalpy. The first-order terms in the equation for the variation $w_0 + w$ then give the linear variable coefficient equation for w

$$\partial_t w + \partial_x (A(w_0(x, y))w) + \partial_y (B(w_0(x, y))w) + \alpha C(w_0(x, y))w = 0 \quad (3.3)$$

where A, B, C are the Jacobian matrices of \mathcal{E}, \mathcal{F} , and \mathcal{H} , respectively. We will study this variable coefficient system by freezing the coefficients at an arbitrary (x_0, y_0) . This gives the constant coefficient problem

$$\partial_t w + A \partial_x w + B \partial_y w + \alpha C w = 0, \quad (3.4)$$

The justification for this reduction is the conjecture that if all the constant coefficient problems (3.4) have $w \rightarrow 0$ as $t \rightarrow \infty$, then the variable coefficient problem (3.3) has $w \rightarrow 0$ as $t \rightarrow \infty$, and that if the variable coefficient problem (3.3) has $w \rightarrow 0$ as $t \rightarrow \infty$ then the nonlinear problem is (locally) asymptotically stable at a steady state with constant enthalpy. This author is not aware of any theorems which rigorously quantify these conjectures.

We will analyze (3.4) by Fourier methods, take

$$w(x, y, t) = e^{i(\omega_1 x + \omega_2 y)} \hat{w}(t)$$

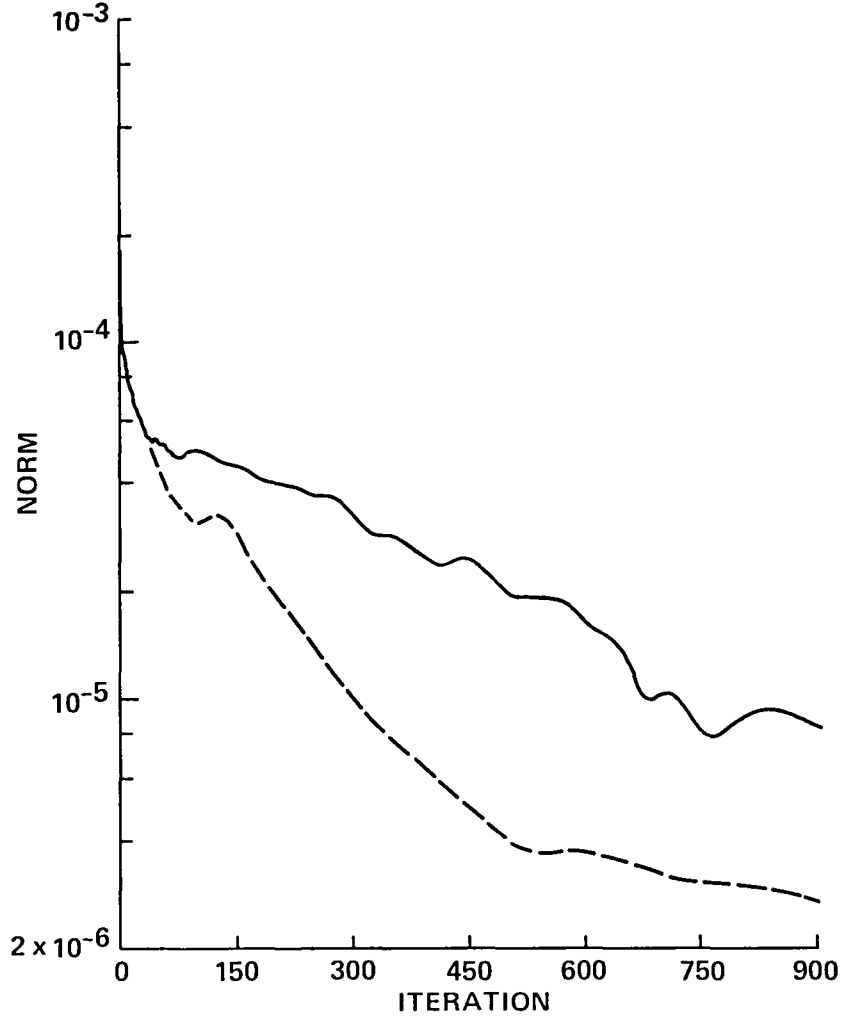


Figure 2.1 L_2 norm of residual in density equation for FLO52R without enthalpy damping (solid line) and with enthalpy damping (dashed line) NACA0012 airfoil, $M_\infty = 0.8$, angle of attack = 1.25° , 192×32 mesh, CFL=2, four-stage method, coefficients (1/4, 1/3, 1/2, 1), two dissipation evaluations. Iteration started with partially converged solution.

and substitute in (3.4) to obtain the ordinary differential equation

$$\frac{d\hat{w}}{dt} + [i(\omega_1 A + \omega_2 B) + \alpha C]\hat{w} = 0 \quad (3.5)$$

Here take $\hat{w} = e^{zt}w$ where z is complex and w is a constant four-vector. This gives

$$(zI + i(\omega_1 A + \omega_2 B) + \alpha C)w = 0 \quad (3.6)$$

If all the zeros of the polynomial $p(z) := \det(zI + i(\omega_1 A + \omega_2 B) + \alpha C)$ lie in the left half-plane $\{\Re(z) < 0\}$, then all solutions of (3.4) decay to 0 as $t \rightarrow \infty$.

Note that C has rank one, since we have $\mathcal{N}(w) = (w_1, w_2, w_3, w_1 \tilde{h}_4)^T (H - H_\infty)$ and

$$C = \frac{\partial \mathcal{N}}{\partial w} \Big|_{H=H_\infty} = \begin{pmatrix} w_1 \\ w_2 \\ w_3 \\ w_1 \tilde{h}_4 \end{pmatrix} \cdot \nabla_w H$$

The zeros of the polynomial $p(z)$ are -1 times the eigenvalues of the matrix $i(\omega_1 A + \omega_2 B) + \alpha C$. We know that $i(\omega_1 A + \omega_2 B)$ is diagonalizable with eigenvalues $i\omega U$, $i\omega(U + c)$, and $i\omega(U - c)$, where $U = \omega_1 u + \omega_2 v$ and $\omega^2 = \omega_1^2 + \omega_2^2$. The following is a theorem about rank one perturbations of diagonalizable matrices with multiple eigenvalues

THEOREM. *Let R be a (complex) matrix with a double eigenvalue λ and two linearly independent eigenvectors associated with λ . Let S have rank one. Then $R + S$ has λ as an eigenvalue.*

PROOF Let v_1, v_2 be two linearly independent eigenvectors associated with λ . Since S has rank one, there is some nontrivial combination $av_1 + bv_2$ such that $S(av_1 + bv_2) = 0$. Then $(R + S)(av_1 + bv_2) = \lambda(av_1 + bv_2)$

Applying the theorem with $R = i(\omega_1 A + \omega_2 B)$ and $S = \alpha C$, we see that $i(\omega_1 A + \omega_2 B) + \alpha C$ has an eigenvalue $i\omega U$. This leads to the conclusion that enthalpy damping cannot produce a system (3.4) which has solutions that decay to 0 as $t \rightarrow \infty$. This result is very general; it applies to any form of enthalpy damping for which the enthalpy damping term is proportional to the difference between the local enthalpy and the freestream enthalpy. Indeed, if the enthalpy damping term in (3.2) is formulated as $\mathcal{K} = (H - H_\infty)(\beta_1(w), \beta_2(w), \beta_3(w), \beta_4(w))^T$, then the matrix C has rank one.

We are going to further study the system of equations with enthalpy damping, to investigate the behavior of the eigenvalues that do move off the imaginary axis and, to find a reasonable choice for \tilde{h}_4 . To do this, it is not convenient to work with the matrices A and B , for the entries of these matrices are complicated functions of (w_1, w_2, w_3, w_4) . It is easier to work with similarity transforms $N^{-1}AN$ and $N^{-1}BN$, where N is the matrix $\partial w / \partial \tilde{w}$ and $\tilde{w} = (\rho, u, v, p)$. This simply amounts to writing the Euler equations in primitive variables rather than in conservative variables. One has

$$N = \begin{pmatrix} 1 & & & \\ u & \rho & & \\ v & 0 & \rho & \\ (u^2 + v^2)/2 & \rho u & \rho v & 1/(\gamma - 1) \end{pmatrix}$$

$$N^{-1} = \begin{pmatrix} 1 & & & \\ -u/\rho & 1/\rho & & \\ -v/\rho & 0 & 1/\rho & \\ (\gamma - 1)(u^2 + v^2)/2 & (1 - \gamma)u & (1 - \gamma)v & \gamma - 1 \end{pmatrix}$$

The matrices $N^{-1}AN$ and $N^{-1}BN$ have been given in Warming and Beam (Ref. 14). They are

$$N^{-1}AN = \begin{pmatrix} u & \rho & 0 & 0 \\ 0 & u & 0 & 1/\rho \\ 0 & 0 & u & 0 \\ 0 & \rho c^2 & 0 & u \end{pmatrix}, \quad N^{-1}BN = \begin{pmatrix} v & 0 & \rho & 0 \\ 0 & v & 0 & 0 \\ 0 & 0 & v & 1/\rho \\ 0 & 0 & \rho c^2 & v \end{pmatrix}$$

The matrix $N^{-1}CN$ turns out to be

$$N^{-1}CN = \begin{pmatrix} c^2/(1 - \gamma) & \rho u & \rho v & \gamma/(\gamma - 1) \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ < \text{first row times } (1 - \gamma)(u^2 + v^2 - 2\tilde{h}_4)/2 > \end{pmatrix}$$

Since $p(z) = \det(zI + i(\omega_1 N^{-1}AN + \omega_2 N^{-1}BN) + \alpha N^{-1}CN)$, we can use the expressions for $N^{-1}AN$, $N^{-1}BN$, and $N^{-1}CN$ to compute $p(z)$. It turns out that with $U = \omega_1 u + \omega_2 v$ and $\omega^2 = \omega_1^2 + \omega_2^2$ we have $p(z) = p_0(z) + \alpha q(z)$ where

$$\begin{aligned} p_0(z) &= (z + i\omega U)^2 (z + i\omega(U + c))(z + i\omega(U - c)) \\ q(z) &= \gamma(\tilde{h}_4 - E)z^3 + i\omega U((2\gamma + 1)(\tilde{h}_4 - E) - c^2/\gamma)z^2 \\ &\quad \omega^2 \left[\frac{c^2}{\gamma^2} (2U^2 - c^2) + \frac{(\tilde{h}_4 - E)}{\gamma} (c^2 - (\gamma + 2)U^2) \right] z \\ &\quad - i\omega^3 U (U^2 - c^2) (\tilde{h}_4 - E - \frac{c^2}{\gamma}) \end{aligned} \tag{3.7}$$

The polynomial $p_0(z)$ is the polynomial one would obtain with no enthalpy damping; its roots are purely imaginary, which shows that the Euler equations without enthalpy damping are a hyperbolic system. (To show hyperbolicity, one also has to verify that there are linearly independent eigenvectors associated with the double zero at $z = -i\omega U$.) It can be verified that $q(-i\omega U) = 0$; this is a consequence of the theorem stated above.

We can obtain further information. First, take $U = cM$ and replace z by $c\omega\zeta$. Next, take $\tilde{h}_4 - E = kc^2/\gamma$. Note that for $k = 0$ we have $\tilde{h}_4 = E$ and for $k = 1$ we have $\tilde{h}_4 = H$. The polynomial $p(z)$ becomes

$$\begin{aligned} & \omega^4 c^4 [(\zeta + iM)^2(\zeta + i(M+1))(\zeta + i(M-1)) \\ & + \frac{\alpha c}{\omega} \{k\zeta^3 + \frac{i}{\gamma} M(2k\gamma + k - 1)\zeta^2 \\ & + \frac{1}{\gamma} ((2 - 2k - k\gamma)M^2 + k - 1)\zeta - \frac{i}{\gamma} (k-1)M(M^2 - 1)\}] \end{aligned}$$

This factors as

$$\begin{aligned} & \omega^4 c^4 (\zeta + iM) \{ \zeta^3 + 3iM\zeta^2 + (1 - 3M^2)\zeta + iM(1 - M^2) + \frac{\alpha c}{\omega} [k\zeta^2 + i\frac{M}{\gamma}(\gamma k + k - 1)\zeta + \frac{k-1}{\gamma}(1 - M^2)] \} \\ & =: \omega^4 c^4 (\zeta + iM) R(\zeta) \end{aligned} \quad (3.8)$$

(The polynomial $R(\zeta)$ is exactly the polynomial one would obtain by analyzing enthalpy damping applied to the Euler equations in one space dimension.) The plan now is to study the location of the zeros of the polynomial $R(\zeta)$. Assume $\alpha > 0$. Writing $R(\zeta) = S(\zeta) + (\alpha c/\omega)T(\zeta)$, we note that for ζ on the imaginary axis, $S(\zeta)$ is purely imaginary and $T(\zeta)$ is real, hence R cannot vanish on the imaginary axis unless S and T have a common zero on the imaginary axis. The zeros of S lie on the imaginary axis and are $-iM$, $-i(M+1)$, $-i(M-1)$. At these ζ , $T(\zeta)$ takes on the values $(k-1)/\gamma$, $-(M+1)(\gamma k - k + 1)/\gamma$, and $(M-1)(\gamma k - k + 1)/\gamma$, respectively. Thus if $M^2 \neq 1$ and $k \neq 1$, R has no zeros on the imaginary axis. Hence in this case we may choose any convenient value of M and use the algorithm given by J. J. H. Miller (Ref. 15) for deciding whether a polynomial has all its roots in the open left half-plane.

The algorithm is as follows. For any function $p(z)$, define $p^*(z) = \overline{p(-\bar{z})}$. Given a polynomial $p(z)$, take a point ζ in the left half-plane for which $p(\zeta) \neq 0 \neq p^*(\zeta)$, and define

$$\check{p}(z) = \frac{p^*(\zeta)p(z) - p(\zeta)p^*(z)}{z - \zeta}$$

Then $\check{p}(z)$ is a polynomial of degree less than $p(z)$. The theorem is that $p(z)$ has all its zeros in the open left half-plane if and only if $|p^*(\zeta)| > |p(\zeta)|$ and \check{p} has all its zeros in the open left half-plane. Thus one can recursively apply this theorem, the degree of the polynomial decreasing at each step.

Consider the case $M^2 < 1$. In this case $M = 0$ is a convenient value to pick, and one can show (using Miller's algorithm) that R has all its zeros in $\{\Re(\zeta) < 0\}$ if and only if $k > 1$. Now consider the case $M^2 > 1$. In this case one can choose $M = 2$ (for example), the result is that R never has all its zeros in the open left half-plane. Now consider the case $M^2 = 1$. In this case R has a zero at $\zeta = 0$, and one can show that the other zeros of R are in the open left half-plane if and only if $k > 1$. In summary, the polynomial $R(\zeta)$ has all its zeros in the open left half-plane $\{\Re(\zeta) < 0\}$ if and only if $M^2 < 1$ and $k > 1$. Note that the condition $M^2 < 1$ is $(\omega_1 u + \omega_2 v)^2 < c^2(\omega_1^2 + \omega_2^2)$, which holds for all (ω_1, ω_2) if and only if $u^2 + v^2 < c^2$, i.e., the flow is subsonic. It is interesting to note further that $k = 0$ (unstable) is the case $\tilde{h}_4 = E$, the formulation of enthalpy damping we obtained in section 2 from the potential equation with enthalpy damping, while the case $k = 1$ (neutrally stable) is the same as $\tilde{h}_4 = H$, the formulation that "can lead to instability" according to Jameson, Schmidt, and Turkel (Ref. 1).

Thus enthalpy damping leads to a differential system which has a general solution that does not decay to zero. Of the four eigenvalues on the imaginary axis when there is no enthalpy damping, one stays on the imaginary axis when enthalpy damping is introduced and three move off the imaginary axis. The analysis suggests that the energy equation with enthalpy damping be written in the form

$$(\rho E)_t + (\rho u H)_x + (\rho v H)_y + \alpha \rho (E + kc^2/\gamma)(H - H_\infty) = 0$$

The eigenvalues that move off the imaginary axis move into the left half-plane provided that $k > 1$ and $u^2 + v^2 < c^2$

It is appropriate here to comment on the missing factor in the analysis, the dissipation term. Going back to (1.7), take a dissipation term of the form

$$D(w) = (\partial_x^4 + \partial_y^4)(w_1, w_2, w_3, \rho H)^T$$

The point here is that the fourth component is not w_4 , it is instead ρH . This is done because the system of equations with artificial dissipation still admits steady solutions with $H = H_\infty$. To see this, note that when the continuity equation with dissipation is multiplied by H_∞ and subtracted from the energy equation with dissipation we obtain

$$\begin{aligned} &(\rho(E - H_\infty))_t + (\rho u(H - H_\infty))_x + (\rho v(H - H_\infty))_y \\ &+ \alpha \rho(\tilde{h}_4 - H_\infty)(H - H_\infty) + \epsilon(\partial_x^4 + \partial_y^4)\rho(H - H_\infty) = 0 \end{aligned}$$

which evidently admits steady solutions with $H = H_\infty$. We would want to look at the polynomial

$$p(z) := \det(zI + i(\omega_1 A + \omega_2 B) + \alpha C + \epsilon(\omega_1^4 + \omega_2^4)D)$$

where $D = \partial D / \partial w$. If we do this and then form $N^{-1}DN$, we find

$$N^{-1}DN = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ (\gamma - 1)^2(u^2 + v^2)/2 & 0 & 0 & \gamma \end{pmatrix}$$

The fact that the fourth row here is not the fourth row of the identity matrix has the consequence that $p(z)$ with $\epsilon \neq 0$ is vastly more complicated than $p(z)$ with $\epsilon = 0$. This author has not been able to make any headway in analyzing the case $\epsilon \neq 0$ because of the formidable algebraic difficulties. In fact, even for the case of no enthalpy damping ($\alpha = 0$) this author is not aware of any proof of stability of the linearized system with dissipation, because the dissipation term is applied to a nonlinear function of w .

Some experiments were done with the code FLO52R in changing the form of the enthalpy damping term in the energy equation. The form $\tilde{h}_4 = E + kc^2/\gamma$ was used, and various values of k were tried. Only one line of the code was changed. The first case was flow about a NACA0012 airfoil at a freestream Mach number 0.8 and at an angle of attack of 1.25°. The code was run in multigrid mode with 40 cycles on a 48×8 mesh, then 40 cycles on a 96×16 mesh, and finally 100 cycles on a 192×32 mesh. A five-stage scheme with coefficients (1/4, 1/6, 3/8, 1/2, 1) was used, with two dissipation evaluations per time step. For the unmodified code, the convergence rate for 100 cycles on the 192×32 grid was 0.9557. A table of convergence rate vs k for the modified code is shown here.

k	1.00	1.25	1.50	1.75	2.00	2.50	3.00	4.00
convergence rate	.	0.9864	0.9565	0.9538	0.9553	0.9580	0.9630	0.9672	0.9788

Evidently the optimum k is approximately 1.5. For a subcritical case with a freestream Mach number of 0.63 and an angle of attack of 2°, with all other parameters the same, the unmodified code had convergence rate 0.9288. The modified code produced the following results:

k	1.00	1.50	2.00	2.50	3.00	3.50	4.00
convergence rate	. .	0.9451	0.9303	0.9282	0.9276	0.9273	0.9275	0.9302

Here the optimal k is apparently between 2.5 and 3.5. In both cases it is clearly difficult to improve on the original code. This author believes the excellent convergence behavior of the unmodified code may be explained as follows.

The energy equation enthalpy damping term in the unmodified code comes from discretizing

$$\partial_t(\rho(E - H_\infty)) + \partial_x(\rho u(H - H_\infty)) + \partial_y(\rho v(H - H_\infty)) + \alpha \rho(H - H_\infty) = 0$$

i.e., this equation is used in place of the fourth equation of (2.9). In terms of the partial differential equation system, an equivalent equation is obtained by multiplying the first equation of (2.9) by H_∞ and adding to this, obtaining

$$\partial_t(\rho E) + \partial_x(\rho u H) + \partial_y(\rho v H) + \alpha \rho(H - H_\infty)(H_\infty + 1) = 0$$

In the notation of (3 2), $\tilde{h}_4 = H_\infty + 1$. With the definition of \tilde{h}_4 as $\tilde{h}_4 = E + kc^2/\gamma$ we would obtain $\tilde{h}_4 = H_\infty + 1$ if we had $k = \gamma(E_\infty - E + c_\infty^2/\gamma + 1)/c^2$. Now, the nondimensionalization in the FLO52 codes is such that $c_\infty^2 = \gamma$, hence at freestream conditions we would have $k = 2$. Thus $\tilde{h}_4 = E + 2c^2/\gamma$ is roughly the same as the unmodified code; also the previously noted experiments showed $k = 1.5$ in one case and $k \in (2.5, 3.5)$ in the other case were near-optimal. Thus $k = 2$ (nearly the original code) would be expected to give excellent results. It is also clear that the unmodified code is equivalent to a code with locally varying k ; this author has not experimented with other forms of locally varying enthalpy damping terms.

4. FINITE DIFFERENCE FORMULATION WITH ENTHALPY DAMPING

The aim of this section is to study the actual finite difference implementation of enthalpy damping. The aim is to understand why enthalpy damping can be effective in practice when on the level of the partial differential equation it does not produce a system with solutions which tend to zero.

The general formulation of enthalpy damping to be studied is as follows. First spatial differencing is applied, sending w^n to $\tilde{w}^n := \mathcal{F}(w^n)$. Then enthalpy damping is applied, via

$$w^{n+1} - \tilde{w}^n + \alpha \Delta t \mathcal{H}(w^{n+1}, \tilde{w}^n) = 0 \quad (4.1)$$

For example, the first equation could be taken to be

$$\rho_{i,j}^{n+1} - \tilde{\rho}_{i,j}^n + \alpha \Delta t \rho_{i,j}^{n+1} (\tilde{H}_{i,j}^n - H_\infty) = 0$$

Notice that in this formulation (the FLO52 formulation) the enthalpy damping term is in a semi-implicit form which is readily solved for $\rho_{i,j}^{n+1}$. If the enthalpy damping term were fully implicit, it would be nonlinear in w^{n+1} and difficult to solve.

We plan to study the stability of (4 1) about a solution w^* with $w^* = \mathcal{F}(w^*)$ and $\mathcal{H}(w^*, w^*) = 0$. Perturbing to $w^* + \epsilon w$, we find the perturbation equation is

$$w^{n+1} - \mathcal{A}w^n + \alpha \Delta t (\mathcal{B}_1 w^{n+1} + \mathcal{B}_2 \mathcal{A}w^n) = 0 \quad (4.2)$$

where

$$\begin{aligned} \mathcal{A} &= D\mathcal{F}(w^*) \quad (\text{the Jacobian matrix}) \\ \mathcal{B}_1 &= D_1 \mathcal{H}(w^*, w^*) \quad (\text{Jacobian with respect to first argument}) \\ \mathcal{B}_2 &= D_2 \mathcal{H}(w^*, w^*) \quad (\text{Jacobian with respect to second argument}) \end{aligned}$$

The equation (4 2) is equivalent to

$$(I + \alpha \Delta t \mathcal{B}_1) w^{n+1} = (I - \alpha \Delta t \mathcal{B}_2) \mathcal{A} w^n \quad (4.3)$$

Let us take

$$\mathcal{H}(w^{n+1}, \tilde{w}^n) = (\tilde{H}^n - H_\infty) \mathcal{H}'(w^{n+1}, \tilde{w}^n)$$

Then $\mathcal{B}_1 = 0$ and $\mathcal{B}_2 = \mathcal{H}'(w^*, w^*) \cdot \nabla_{\tilde{w}^n} (\tilde{H}^n - H_\infty)$. Furthermore, if the basic iteration scheme is an explicit multistage method with central spatial differencing, then $\mathcal{F}(w^n) = \mathcal{P}(\Delta t L) w^n$, where \mathcal{P} is a polynomial, $L = A\delta_x/2\Delta x + B\delta_y/2\Delta y$, and δ_x and δ_y are central space difference operators. Thus the matrix problem (4 3) becomes

$$w^{n+1} = (I - \alpha \Delta t \mathcal{B}_2) \mathcal{P}(\Delta t L) w^n$$

Fourier transforming in space produces the problem

$$\hat{w}^{n+1} = (I - \alpha \Delta t \mathcal{B}_2) \mathcal{P}(\Delta t \hat{L}) \hat{w}^n$$

where $\Delta t \hat{L} = i(\omega_1 A + \omega_2 B)$ and $\omega_1 = \Delta t \sin \theta / \Delta x$, $\omega_2 = \Delta t \sin \phi / \Delta y$, and θ, ϕ are the Fourier dual variables.

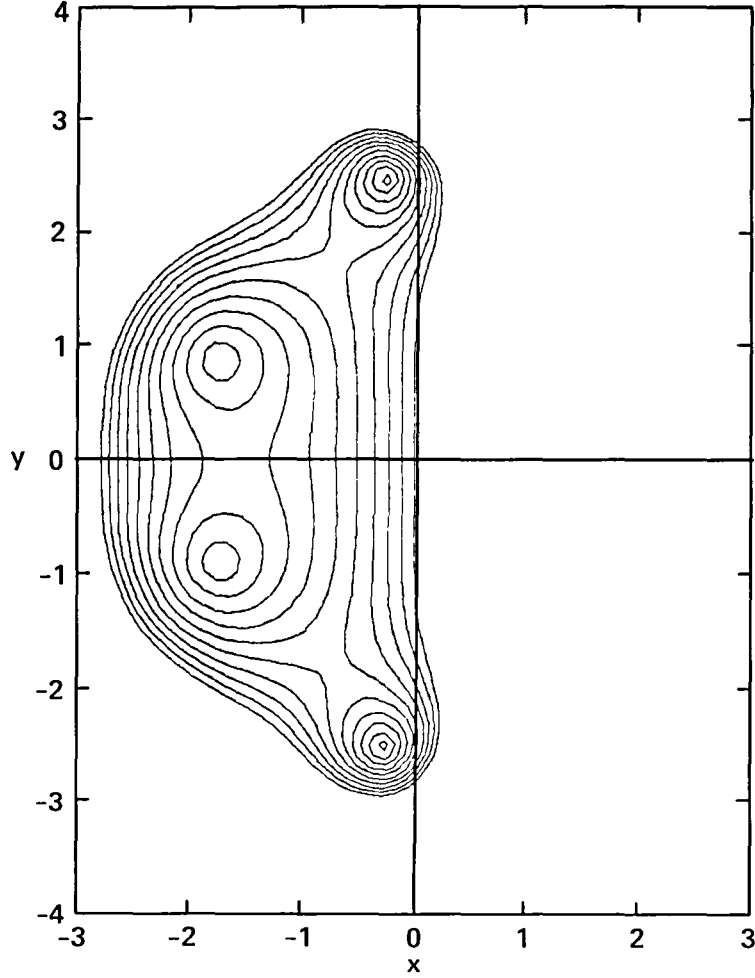


Figure 4.1. Contours of $|\mathcal{P}(z)|$ in the complex plane for classical fourth-order Runge-Kutta. Contour levels: 0.1, 0.2, 0.3, ..., 1.0.

A reasonable form (consistent with the work of section 3) for the implementation of the enthalpy damping step is

$$\begin{aligned}
 \rho^{n+1} - \widetilde{\rho}^n + \alpha \Delta t \rho^{n+1} (\widetilde{H}^n - H_\infty) &= 0 \\
 (\rho u)^{n+1} - (\widetilde{\rho u})^n + \alpha \Delta t (\rho u)^{n+1} (\widetilde{H}^n - H_\infty) &= 0 \\
 (\rho v)^{n+1} - (\widetilde{\rho v})^n + \alpha \Delta t (\rho v)^{n+1} (\widetilde{H}^n - H_\infty) &= 0 \\
 (\rho E)^{n+1} - (\widetilde{\rho E})^n + \alpha \Delta t [(\rho E)^{n+1} + k \widetilde{p}^n] (\widetilde{H}^n - H_\infty) &= 0
 \end{aligned} \tag{4.4}$$

The parameter k here is the same as in the previous section. Then our perturbation equation is

$$w^{n+1} = (I - \alpha \Delta t B_2) \mathcal{P}(\Delta t \widehat{L}) w^n$$

For the decay of perturbations about w^* , it is necessary and sufficient that all eigenvalues λ of $(I - \alpha \Delta t B_2) \mathcal{P}(\Delta t \widehat{L})$ satisfy $|\lambda| < 1$. Furthermore, the asymptotic rate of decay is given by the modulus of the largest eigenvalue. For $\alpha = 0$ (no enthalpy damping) we recover the usual CFL-like condition $\max_j |\mathcal{P}(\lambda_j)| < 1$, where λ_j are the eigenvalues of $\Delta t \widehat{L}$: $\lambda_1 = \lambda_2 = i\omega U$, $\lambda_3 = i\omega(U + c)$, $\lambda_4 = i\omega(U - c)$, where $\omega U := \omega_1 u + \omega_2 v$ and $\omega^2 := \omega_1^2 + \omega_2^2$.

For example, the explicit Euler method has $\mathcal{P}(\Delta t \widehat{L}) = I - \Delta t \widehat{L}$ (this is unstable); the classical fourth-order Runge-Kutta method has $\mathcal{P}(x) = 1 - x + x^2/2 - x^3/6 + x^4/24$, for which it is known that the stability condition $\max_j |\mathcal{P}(\lambda_j)| < 1$ becomes $\max_j |\lambda_j| < 2\sqrt{2}$. Contours of $|\mathcal{P}(z)|$ in the complex z -plane are shown in Figure 4.1 for the fourth-order Runge-Kutta method.

The question before us now is, does the modulus of the maximum eigenvalue decrease when $\alpha\Delta t$ is increased from 0? Also, what are good values for $\alpha\Delta t$ and k ? For \mathcal{H} as in (4.4) we have

$$\mathcal{B}_2 = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ \rho E + kp \end{pmatrix} \cdot \frac{1}{\rho} (-\gamma E + (\gamma - 1)(u^2 + v^2), (1 - \gamma)u, (1 - \gamma)v, \gamma)$$

Again we have a matrix of rank one, \mathcal{B}_2 , and a diagonalizable matrix $\mathcal{P}(\Delta t \hat{L})$ with a double eigenvalue at $\mathcal{P}(i\omega U)$, so the theorem of the previous section implies that $(I - \alpha\Delta t \mathcal{B}_2)\mathcal{P}(\Delta t \hat{L})$ has as one of its eigenvalues $\mathcal{P}(i\omega U)$. Thus if this eigenvalue is the eigenvalue of maximum modulus, then enthalpy damping will not improve the convergence rate of the iteration. On the other hand, if one of the other eigenvalues is obstructing the convergence, then enthalpy damping may help.

To help in studying this question, it is useful to diagonalize the matrix \hat{L} . The matrix T which diagonalizes $\omega_1 A + \omega_2 B$ is known analytically (see Warming and Beam (Ref 14)). It is

$$T = \begin{pmatrix} 1 & 0 & \rho/(\sqrt{2}c) & \rho/\sqrt{2}c \\ 0 & \omega_2/\omega & \omega_1/(\sqrt{2}\omega) & -\omega_1/(\sqrt{2}\omega) \\ 0 & -\omega_1/\omega & \omega_2/(\sqrt{2}\omega) & -\omega_2/(\sqrt{2}\omega) \\ 0 & 0 & \rho c/\sqrt{2} & \rho c/\sqrt{2} \end{pmatrix}$$

Then

$$\begin{aligned} T^{-1}(I - \alpha\Delta t \mathcal{B}_2)\mathcal{P}(\Delta t \hat{L})T &= (I - \alpha\Delta t T^{-1}\mathcal{B}_2 T)T^{-1}\mathcal{P}(\Delta t \hat{L})T \\ &= (I - \alpha\Delta t T^{-1}\mathcal{B}_2 T) \\ &\quad \cdot \text{diag}\{\mathcal{P}(i\omega U), \mathcal{P}(i\omega U), \mathcal{P}(i\omega[U + c]), \mathcal{P}(i\omega[U - c])\}. \end{aligned}$$

The characteristic polynomial of this matrix vanishes at $\mathcal{P}(i\omega U)$, and upon division by $z - \mathcal{P}(i\omega U)$ the characteristic polynomial becomes (writing $\sigma = \mathcal{P}(i\omega U)$, $\sigma_+ = \mathcal{P}(i\omega[U + c])$, $\sigma_- = \mathcal{P}(i\omega[U - c])$, and $M = U/c$),

$$\begin{aligned} &(z - \sigma)(z - \sigma_+)(z - \sigma_-) \\ &+ \frac{\alpha\Delta t c^2}{2\gamma} [z^2 \{2(k - 1)\sigma + (1 + M)(\gamma k - k + 1)\sigma_+ + (1 - M)(\gamma k - k + 1)\sigma_-\} - \\ &- z(\{\gamma k(1 + M) + (k - 1)(1 - M)\}\sigma\sigma_+ + \{\gamma k(1 - M) + (k - 1)(1 + M)\}\sigma\sigma_- \\ &\quad + 2(\gamma k - k + 1)\sigma_+\sigma_-) \\ &+ 2\gamma k\sigma\sigma_+\sigma_-] \end{aligned} \tag{4 5}$$

Note that α has units of time⁻¹velocity⁻² (cf. (4 4)), so $\alpha\Delta t c^2$ is dimensionless.

We will study the roots of the polynomial in (4 5) numerically. Fix a ‘‘Mach number’’ $M = U/c$. Define ν (‘‘CFL number’’) as $\max(|\omega U|, |\omega(U + c)|, |\omega(U - c)|)$. Then

$$\omega U = \begin{cases} \nu \text{sgn}(M) / \max(|1 + 1/M|, |1 - 1/M|), & \text{if } M \neq 0 \\ 0 & \text{if } M = 0 \end{cases} \tag{4 6}$$

Now for a given M we can graph the spectral radius vs ν by using the definition of ωU from (4 6) and then numerically finding the roots of the polynomial (4 5). We can further search for an optimal value of the dimensionless parameter $\alpha\Delta t c^2$ by using as optimality criterion the integral between the curves of spectral radius vs. ν for $\alpha\Delta t c^2 = 0$ and $\alpha\Delta t c^2 \neq 0$, and searching for $\alpha\Delta t c^2$ which maximizes this integral (Numerically, the composite trapezoidal rule was used for the integration.) Two typical graphs of spectral radius vs. ν without and with enthalpy damping are shown in Figure 4 2 Note that enthalpy damping has little or no effect at small CFL numbers, it is only for CFL numbers near the stability limit that enthalpy damping is significantly helpful For supersonic flow ($M > 1$), the spectral radius with enthalpy damping was observed to be greater than 1. In all subsonic cases investigated, the spectral radius with enthalpy damping was exactly the same as or slightly less than the spectral radius without enthalpy

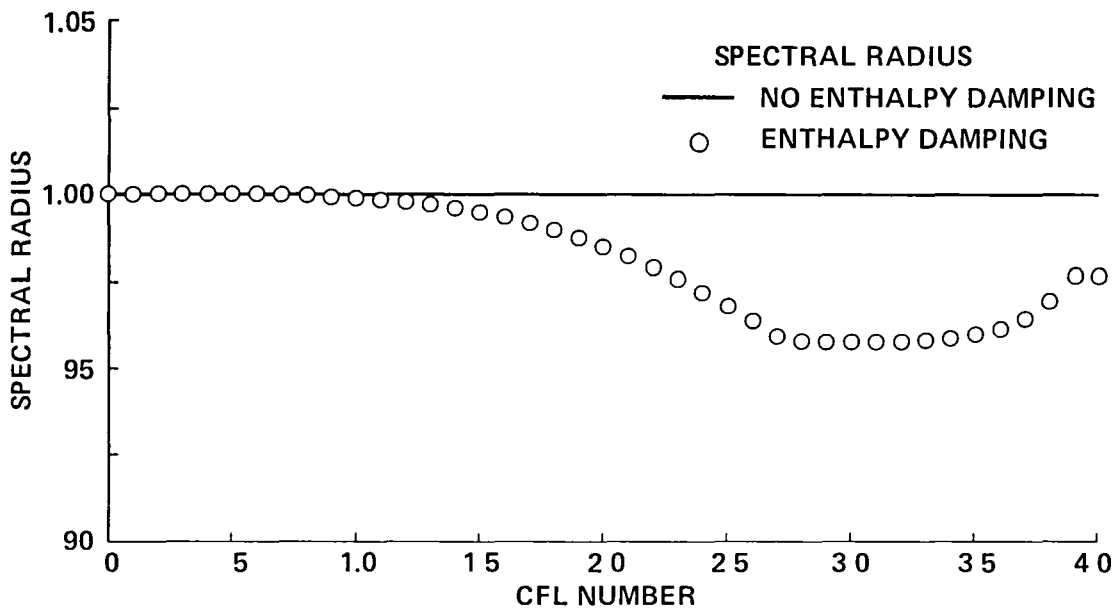


Figure 4 2(a). Spectral radius vs. CFL number for numerical scheme without and with enthalpy damping. Five-stage method with coefficients $(1/4, 1/6, 3/8, 1/2, 1)$, $M = 0.8$, $k = 2$, $\alpha\Delta t c^2 = 0.25$

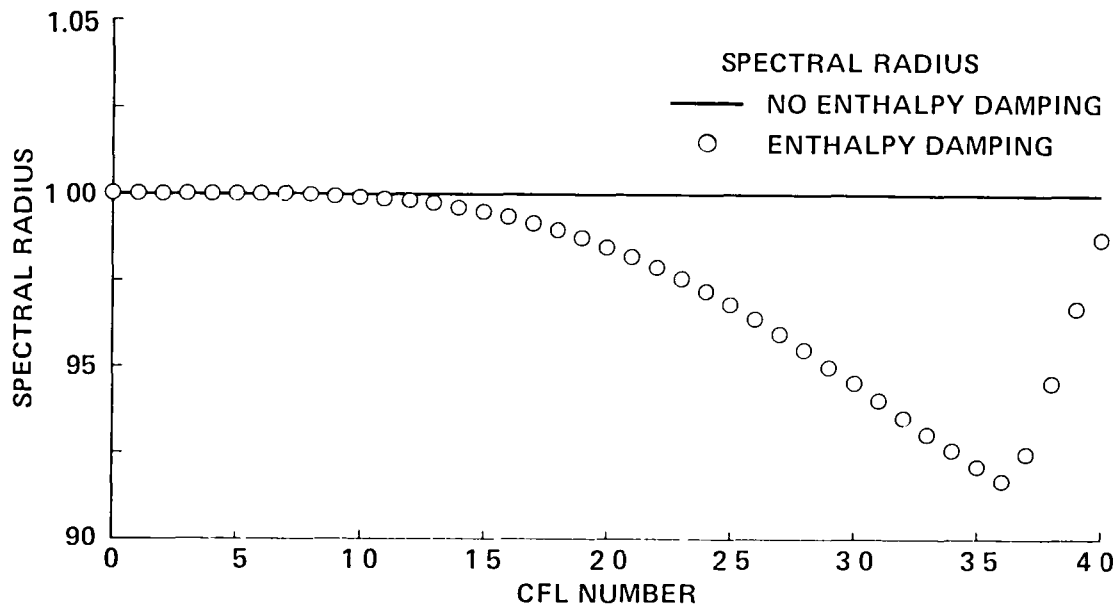


Figure 4 2(b). Same as (a) except $\alpha\Delta t c^2 = 0.5$.

damping, at least in the region of stability. The five-stage method with coefficients $(1/4, 1/6, 3/8, 1/2, 1)$ was studied, with $0 \leq \nu \leq 4$ (the stability limit). With $k = 2$ and two different Mach numbers, $M = 0.8$ and $M = 0.4$, the following results were obtained for the area between the curves as a function of $\alpha\Delta t c^2$.

$\alpha\Delta t c^2$	0.30	0.40	0.50	0.60	0.70	0.80	0.90	1.00
area ($M = 0.8$)	0.0840	0.0967	0.101	0.103	0.104	0.105	0.105	0.0685
area ($M = 0.4$)	0.0039	0.0039	0.0039	0.0039	0.0039	0.0039	0.0039	0.00277

Evidently there is not a sharp optimum value of $\alpha\Delta t c^2$, for in the first case any $\alpha\Delta t c^2$ between 0.5 and 0.9 gives approximately the same result, while in the second case any $\alpha\Delta t c^2 \in (0.3, 0.9)$ can be chosen. If the speed of sound is normalized to be γ at infinity (as in the FLO52 codes), then our good range for $\alpha\Delta t$ becomes $\alpha\Delta t \in (0.36, 0.64)$ in the first case, and $\alpha\Delta t \in (0.21, 0.64)$ in the second case.

5. CONCLUSIONS

The Euler equations with enthalpy damping can be studied as a system of partial differential equations. Linear stability analysis of the system around a state with constant enthalpy shows that not all solutions decay to zero. The analysis suggests that the enthalpy damping term in the energy equation be written as $\alpha\rho(E + kc^2/\gamma)(H - H_\infty)$. For this formulation, the linear stability analysis shows stability (neutral stability, i.e., no eigenvalues in the open right half plane) if $k \geq 1$ and the flow is subsonic.

When the equations are discretized via an explicit multistage method, the spectral radius can be decreased by enthalpy damping, even though it is again the case that one of the eigenvalues is unchanged. A numerical study with $k = 2$ revealed that $\alpha\Delta t c^2 \approx 0.5$ should give good results. The particular value of $\alpha\Delta t$ this leads to depends on the nondimensionalization of a given computer code.

This analysis has not taken the dissipation term into account, because the dissipation operator in the energy equation is applied to ρH , not ρE (so that the equations with dissipation will still have constant enthalpy at steady state), which makes the analysis much more complicated. All the analysis shows that enthalpy damping should be destabilizing for supersonic flow, but in practice enthalpy damping does not harm the convergence for the FLO52 codes in the transonic range. It is probably the case that the dissipation term is effective in stabilizing the computation, even in locally supersonic regions of flow.

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16 Abstract For inviscid steady flow problems where the enthalpy is constant at steady state, it has been proposed by Jameson, Schmidt, and Turkel to use the difference between the local enthalpy and the steady state enthalpy as a driving term to accelerate convergence of iterative schemes. This idea is analyzed here, both on the level of the partial differential equation and on the level of a particular finite difference scheme. It is shown that for the two-dimensional unsteady Euler equations, a hyperbolic system with eigenvalues on the imaginary axis, there is no enthalpy damping strategy which can move all the eigenvalues into the open left half plane. For the numerical scheme, however, the analysis shows and examples verify that enthalpy damping can be effective in accelerating convergence to steady state.			
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