A New method to determine the Chemical Composition of the Cosmic Rays beyond 10¹⁵eV

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abstract

The chemical composition of primary cosmic rays beyond 10^{15} eV could not be measured by the direct method. Here I would like to propose more sensitive method to determine the chemical composition. The idea has been checked by the simulation and compared with the exsisting data concerning on $N_e^{-N}u^{and}N_e^{-N}\gamma$.

The simulation will be also cmpared with the experimental results coming from ANI experiment in a near future.