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HADRONS REGISTRATION IN EMULSION CHAMBER WITH CARBON BLOCK

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NEC in X-ray emulsion chambers with carbon block, which are usually used in the "Pamir" experiment, was Monte-Carlo simulated.Going over from optical density to  $\Sigma E_r$  is discussed. Role NEC in the interpretation of energy spectra is analysed.

<u>1.Introduction.</u> As a result of the nuclear-electromagnetic cascade (NEC) in C-layer and Pb-layer of haderonic block, we get a spot on X-ray film of emulsion chamber and then we get several optical densities of the spot corresponding to several radii of a photometer diaphragm. General methodical problem of the hadronic block measurements is how to obtain a value of energy of the hadron -  $E_0$  (or energy of electromagnetic component of a cascade initiated by the hadron- $\Sigma E_{\sigma}$ ) from the data of the optical densities D.

2.Simulation Assumptions. Spots of the individual NEC were simulated. We took into account a chamber consisting of 6 cm Pb gamma block  $(0.35\lambda, 10.5 \text{ c.u.})$  and hadronic block having 65 cm C-generator  $(0.95\lambda, 2.7 \text{ c.u.})$  and 4 cm Pb-layer  $(0.23\lambda$ 7.0 c.u.). The calculation were divided into following parts: a.For purely proton spectrum with integral slope S=2, NEC in the chamber was Monte - Carlo simulated, useing scaling model b.For each particles of electromagnetic component at energy  $E_{g} > 0.05$  TeV the mean cascade function for electron density  $S(E_{g},r,t)$  at depths t was used. Target diagramme was constructed on the area  $300x300 \text{ µm}^2$  with cells  $12x12 \text{ µm}^2$  each. c.The electron density diagramme was transformed to the flux of light diagramme, assumed the characteristic curve D(r) = $4.0\{1-\exp[-3.25 S(r)]\}$  of X-ray film. At a position of diaphragm for X-ray film at depth 4 cm of Pb-layer was obtained.

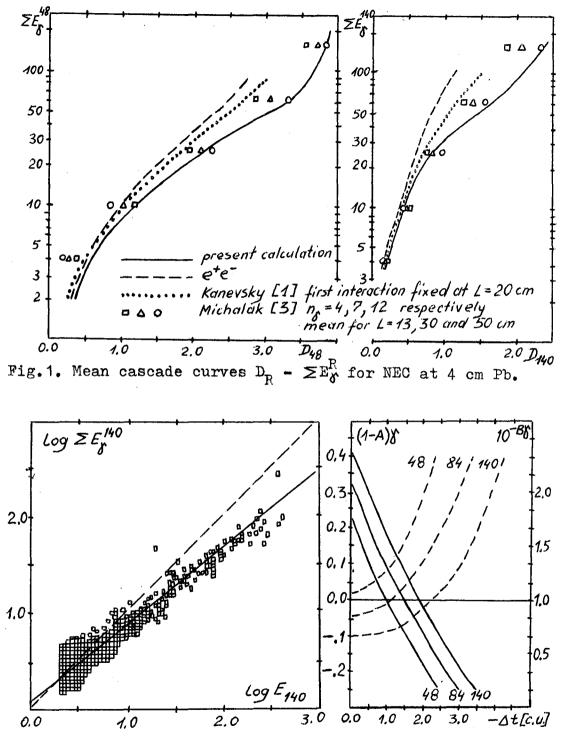
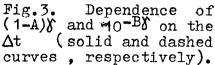
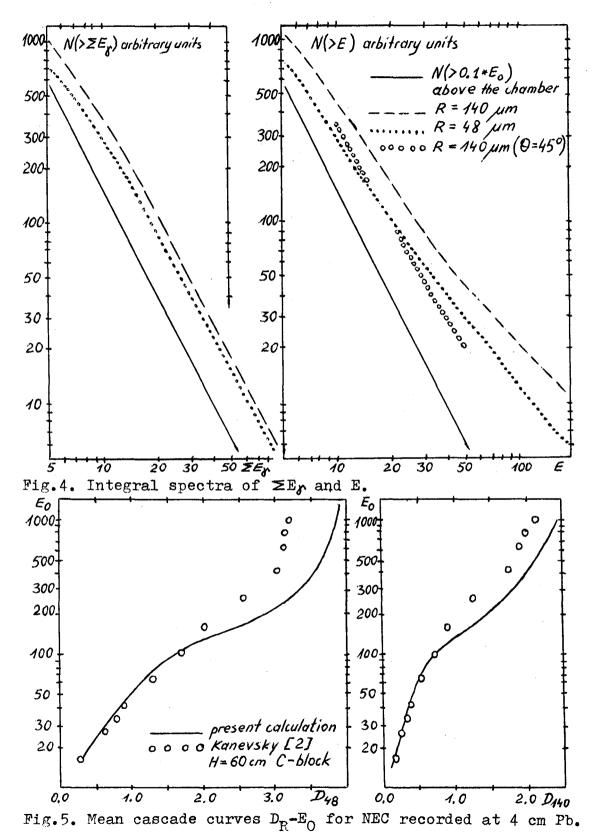


Fig.2. Scatter plot of  $\Sigma E_{\delta}$  versus E for R=140 µm and  $\Delta t=0$ . Dasched line is the dependence for the case  $\Delta t = -2.0$  c.u.





<u>3. Pesults and Discussion.</u> Mean cascade curves for  $D_R$  as a function of the energy  $\Sigma E_{\delta}^R$  (for radii of photometer diaphragm R=48 and 140 µm) are ploted in Figure 1. The difference between the densities given by NEC in chamber and the ones corresponding to electron cascade is significant at energies above 30 TeV. The present calculation shows the much stronger difference than the one pointed out by [1] (the difference in the C-block thick is negligible at high energy, see Fig.5)

In Figure 2 we presented scatter plot of  $\Sigma E_{\delta}$  and of energy E obtained useing the "average transition curves" for e<sup>+</sup>e<sup>-</sup> (recorded by diaphragm with R=140 µm). The energy  $\Sigma E_{\delta}$ depend on E as  $\Sigma E_{\delta} = 10^{B} E^{A}$ . Then the observed spectrum  $N(>E)=C \ 10^{-B\delta} E^{(1-A)\delta} E^{-\delta}$  should be flatten by power index of  $(1-A)\delta$  than the injection spectrum on  $\Sigma E_{\delta}$ :  $N(>\Sigma E_{\delta})=C(\Sigma E_{\delta})^{-\delta}$ . The quantities  $(1-A)\delta$  and  $10^{-B\delta}$  are presented in Figure 3 (at  $\Delta t=0$ ). Figure 4 shows the integral spectra of  $\Sigma E_{\delta}$  and E

Roughly, it is possible to estimate the energy spectra took into account the thick of carbon layer  $\Delta t$  penetrated by NEC. Nevertheless this manner is dangerous, because deepen in NEC development the difference is not so significant (see Figure 4, curve for  $\Theta = 45 \text{ deg}$ : t = 4.0/cos(45)Pb).

Mean cascade curves for optical density as a function of hadron energy  $E_0$  re ploted in Figure 5. The difference between present calculation and Kanevski [2] data are caused by the difference in transition  $\sum E_{\kappa} \rightarrow D$  mentioned above.

<u>4.Conclusion</u>. Useing  $D(\geq E_{\varsigma})$  corresponding to e<sup>+</sup>e<sup>-</sup> cascade or these calculated by [1,2] provide to flattenens of the observed hadron spectrum and decreases of attenuation m.f.p. (calculated from zenith angle distribution) at  $\geq E_{r} > 30$  TeV.

## References

 [1] Kanevsky B.L., in:Electron-Photon Cascade at High Energy Cosmic Rays, Moscow (1980), in Russian
[2] Kanevsky B.L., et al., 17 th ICRC, 5, 335, Paris (1981) and Moscow State University Report-in Russian
[3] Michalak W., Acta University of Lodz, <u>60</u>, 137, Lodz (1977)