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THE MAXIMUM DEPTH OF SHOWER WITH  $E_0 > 10^{17}$  eV ON AVERAGE CHARACTERISTICS OF EAS DIFFERENT COMPONENTS

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## ABSTRACT

EAS development model independent method of the determination of a maximum depth of shower  $(X_m)$  is considered.  $X_m$  values obtained on various EAS parameters are in a good agreement.

<u>1. Introduction.</u> Investigations of the shower maximum depth  $X_m$  are carried out at various arrays and by different methods but the significant scattering of the obtained data is still available (Table 1). A reason of most of discrepancies is mainly due to methodical difficulties associated with the transition from the observed EAS parameters  $P=P(X_m)$ to  $X_m$ . Thereby one had to use the theoretical conceptions on EAS development difficult to test experimentally.

2. Method. We considered one more method of X<sub>m</sub> determination on the experimental data obtained at the Yakutsk EAS array. By computer calculations such parameters were found which are the functions of  $P=P(X_m/X)$  or P=P(X-Xm) type in wide limits of initial conditions: X=1020.sec 0. The calculations were carried out at vare ious Eo and 0 on two, quite different models of EAS development. The first model corresponded to scaling [10], the second one - to scaling at  $E < 10^{14}$  eV

Table 1

Parameter	Eo	X <sub>m</sub>	Work
na	1,6.10 <sup>17</sup>	660±30	[1]
T ½ (Q)	1,4.10 <sup>17</sup>	700±15	[2]
	2.10 <sup>17</sup>	681 <u>+</u> 20	[3]
	1,2.10 <sup>17</sup>	706 <u>+</u> 36	[4]
	10 <sup>17</sup>	620 <u>+</u> 20	[5]
	10 <sup>17</sup>	545 <u>+</u> 20	[5]
	10 <sup>17</sup>	500 <u>+</u> 20	[5]
	2•10 <sup>17</sup>	680 <b>±</b> 20	[6]
LDF(Q)	2.1017	627 <b>±</b> 20	[7]
	1,5.10 <sup>17</sup>	600±50	[8]
Ψ(μ)	3.1017	684 <b>±</b> 30	[9]
lg(Sc/Sp)	3.1017	750±30	[9]
$LDF(S_c)$	3•10 <sup>17</sup>	609 <b>±</b> 3	[9]

[10] and to  $n_s \sim E^{0,25}$ at  $E \ge 10^{14}$  eV. The cross-sections in inelastic processes on both models changed with energy according to [10]. The index of the LDF of electrons ne at the distance interval R=200-600  $\mu$  from the  $n_e$ shower core ( $g_e \sim R^{-n_e}$ and ratios of densities of the EAS Cerenkov light to electrons  $lg(Q/S_e)$  and of electrons to muons lg( ge / gm ) at R=300 m were considered. The above parameters are satisfactorily measured at the Yakutsk EAS array (ge=gs-gn).

<u>3. Results.</u> Calculation results at  $E_0 =$  $10^{17}-10^{18}$  eV and  $\Theta =$ 16, 32 and 40° are shown in Figs.1-3. From Fig.1 it is seen that  $n_e$  is unambigiously associated with X<sub>m</sub>/X independently







Fig.2

with  $X_m/X$  independently of E<sub>0</sub>,  $\Theta$  and characteristics of nuclear interactions. We use this peculiarity of electron LDF to find  $X_m$ :

$$X_{m} = \left(\frac{ne - 2.11}{1.7}\right) \cdot X, g/cm^{2}.$$
 (1)

The obtained  $X_m$ are given in Table 2. The parameter lg(Q/ $g_e$ ) which is the function of X- $X_m$  possesses the analogous feature (Fig.2).

$$X_{m} = X - 423(1g \frac{Q}{ge} - 0, 88).$$
 (2)



Table	2
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		*			
lg E <sub>o</sub>	17,20	17,55	18,02	18,42	18,94
X, g/cm <sup>2</sup>	1060	1060	1070	1080	1060
na	2,57	2,60	2,70	2,78	2,80
$X_{\underline{m}} \pm \Delta X_{\underline{m}}$	660 <u>+</u> 30	675 <u>+</u> 30	725 <u>+</u> 30	765 <u>+</u> 30	755 <u>+</u> 40
ne	3,18	3,22	3,26	3,31	3,31
$X_{\underline{m}} \pm \Delta X_{\underline{m}}$	665 <u>+</u> 60	690 <b>±</b> 60	720 <u>+</u> 60	760 <b>±</b> 60	750 <b>±</b> 60
lg(Q/ge)	1,94	1,91	1,79	1,71	1,59
$\mathbf{x}_{\mathbf{m}} \pm \Delta \mathbf{x}_{\mathbf{m}}$	615 <b>±</b> 30	625 <b>±</b> 30	685 <b>±</b> 40	730 <b>±</b> 40	760 <u>+</u> 45
$lg(\Phi/N_{e})$	5,37	5,31	5,19	5,19	5,0
x <sub>m</sub> ±∆x <sub>m</sub>	640 <u>+</u> 40	665 <b>±</b> 40	725 <b>±</b> 40	735 <b>±</b> 40	800 <b>±</b> 40
lg E	1	17,56	17,91	18,39	18,79
X, g/cm <sup>2</sup>	3400	1090	1080	1080	1080
1g(ge/gn)	-	0,76	0,82	0,99	1,04
$X_{m} \pm \Delta X_{m}$	34ap	670 <b>±</b> 20	670 <b>±</b> 20	715±20	725±25
lg(N <sub>e</sub> /N <sub>M</sub> )	-	1,17	1,28	1,47	1,54
$\mathbf{X}_{\mathbf{i}\mathbf{z}} \pm \Delta \mathbf{X}_{\mathbf{m}}$		665±35	6 <b>70±3</b> 5	720 <u>+</u> 35	735 <b>±</b> 35

As for  $\lg(\frac{9e}{9_M})$  the unambiguity condition due to the zenith angle is broken. Therefore the experimental and calculational data are needed to be compared at similar  $\Theta$ .

The averaging of data in Table 2 results in the following expression:

 $X_{m} = (700\pm35) + (66\pm6)(lg E_{0}-18), g/cm^{2}.$  (3)

Note that all the above parameters were experimentally obtained at fixed  $9_{\rm S}$  (300) and calculations were also carried out under such condition. If to make calculations at fixed E<sub>0</sub>, then X<sub>m</sub> value becomes ~ 50 g/cm<sup>2</sup> less. <u>4. Discussion.</u> The integral values  $\lg(\Phi/N_{\rm e})$  and  $\lg(N_{\rm e}/N_{\rm M})$  are close to the parameters  $\lg(Q/Q_{\rm e})$  and  $\lg(Q_{\rm e}/Q_{\rm M})$ . Here  $\phi$  is the total flux of the EAS Cerenkov light; N<sub>M</sub> and N<sub>e</sub> are the total numbers of muons and electrons. Their dependences on X-X<sub>m</sub> are analogous to ones presented in Figs.2 and 3.

Apply this method of the analysis of data to other arrays. One can do it without additional calculations with respect to the parameter  $\lg(9_c/9_{\mu})$  [9], since it is similar to our parameter. According to [9] at R=300 m, E<sub>0</sub>=1017 eV and  $\Theta \approx 15^{\circ}$  we have  $\lg(9_c/9_{\mu})=0.3$ . To recount from  $9_c$ measured at the Haverah Park array to  $9_e$  at R=300 m we use the ratio  $\rho_e / \rho_c \approx 1.8$  [11]. Then  $lg(\rho_e / \rho_{\mu}) \approx 0.58$  and from Fig.3 we find  $X_m \approx 620$  g/cm<sup>2</sup>, i.e. much higher than in Table 1. 5. Conclusion. The analysis of various EAS components based on the method of "model independent" parameters yields  $X_{m}=700\pm35$  g/cm<sup>2</sup> at E<sub>0</sub>=10<sup>18</sup> eV and at fixed  $\gamma_{s}(300)$ . References 1. Efimov, N.N. et al., (1983), Proc. 18-th ICRC, Bangalore, vol.6, 176. 2. Kalmykov, N.N. et al., (1979), Proc. 16-th ICRC, Kyoto, vol.9, 73. 3. Protheroe, R.J. and Turver, K.E. Nuovo Cimento, 51A, (1979), <u>277</u>. 4. Thornton, B. and Clay, R., (1980), Phys. Rev. Lett., vol.45, 1463. 5. Inoue, N. et al., (1981), Proc. 17-th ICRC, Paris, vol. 11, 270. 6. Hammond, P.T. et al., (1978), <u>Nuovo Cimento, 1C, 315</u>. 7. Dyakonov, M.N. et al., (1981), <u>Proc. 17-th</u> ICRC, <u>Paris</u>, vol.6, 106. 8. Hara, T. et al., (1981), Proc. 17-th ICRC, Paris, vol. 11, 277. 9. McComb, T.J.L. and Turver, K.E., (1982), Nucl. Phys., vol.8, 1119. 10. Hillas, A.M., (1979), Proc.16-th ICRC, Kyoto, vol.9, 13. 11. Kellermann, E.W. and Towers, L., (1969), Preprint, Univ.Leeds.

5

51