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# Line Positions and Intensities for the

 $v_1 + v_2$  and  $v_2 + v_3$  Bands of  $160_3$ 

V. Malathy Devi Department of Physics College of William and Mary Williamsburg, Virginia 23185 U.S.A.

J.-M. Flaud and C. Camy-Peyret Laboratoire de Physique Moléculaire et Atmosphérique Tour 13, 3<sup>e</sup> étage Université Pierre et Marie Curie et C.N.R.S. 4 place Jussieu 75252 Paris cedex 05 France

> C. P. Rinsland and M. A. H. Smith Atmospheric Sciences Division NASA Langley Research Center Hampton, Virginia 23665-5225 U.S.A.

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Running title:  $v_1 + v_2$  and  $v_2 + v_3$  Bands of  ${}^{16}O_3$ 

Mail Correspondence to:

Dr. V. Malathy Devi Department of Physics College of William and Mary Williamsburg, Virginia 23185

## ABSTRACT

Using  $0.005-cm^{-1}$  resolution Fourier transform spectra of  ${}^{16}O_3$ , generated by electric discharge from a >99.98% pure sample of  ${}^{16}O_2$ , an extensive analysis of the  $v_1 + v_2$  and the  $v_2 + v_3$  bands in the 5.7  $\mu$ m region was performed. The rotational energy levels of the upper (110) and (011) vibrational states of  ${}^{16}O_3$  were reproduced within their experimental uncertainties using a Hamiltonian which takes explicitly into account the Coriolis-type interaction occurring between the rotational energy levels of both states. Improved vibrational energies and rotational and coupling constants were also derived for the (110) and (011) states. Precise transition moment constants for these two bands were deduced from analysis of 220 measured line intensities. Finally, a complete list of line positions, intensities, and lower state energies for both bands has been generated.

## INTRODUCTION

Because of its atmospheric importance, the  ${}^{16}O_3$  molecule and the isotopic species  ${}^{16}O^{18}O^{16}O$  and  ${}^{16}O^{18}O$  have been the subject of numerous recent spectroscopic studies (<u>1-6</u> and references therein). Almost all of the recent works were concerned with the pure rotation spectrum and the 10 µm absorption region which are widely used for remote sensing measurements of this molecule in the terrestrial atmosphere. However, absorption by atmospheric ozone is clearly visible in many other spectral regions as well, and we present in this paper our work concerning the 5.7 µm region of  ${}^{16}O_3$ . The prominent bands appearing in this spectral region are the two combination bands  $v_1 + v_2$  and  $v_2 + v_3$ . These bands were analyzed in previous laboratory studies (<u>7-9</u>), but at a resolution lower than that achieved in the present work. This work also reports the first experimental measurements of individual line intensities in these bands.

## EXPERIMENTAL DETAILS AND ANALYSIS

The ozone samples were prepared using >99.98% pure  ${}^{16}O_2$ . The technique of silent electric discharge was used in generating the ozone. The absorption cell used in these measurements was a 50-cm long 5.08-cm diameter Pyrex tube with Teflon valves. Potassium chloride windows were wedged 5-10 mrad on the cell to prevent channeling arising from multiple reflections. The two spectra used in this analysis were recorded at room temperature with ~10.6 and ~23.7 Torr of  ${}^{16}O_3$  in the absorption cell. The sample pressures and temperatures were monitored continuously during the ~1 hour scan time for each spectrum using a 0- to 100-Torr Barocel pressure gauge and a thermocouple kept in contact with the cell body. Standards in the  $v_3$  band of  ${}^{12}c{}^{16}O_2$ , reported

by Guelachvili (<u>10</u>) were multiplied by the factor 0.99999979 (average of the two factors reported by Brown et al. (<u>11</u>)) and used to calibrate the wavelength scale of the spectra. For additional experimental details see Flaud et al. (4) and Rinsland et al. (5).

Using the results reported in Ref. (9) it was rather easy to assign transitions with low and medium J and K<sub>a</sub> values (up to J = 45 and K<sub>a</sub> = 11). Following the initial assignments, a first calculation was performed using a Hamiltonian taking into account the Coriolis-type interaction between the rotational levels of the (110) and (011) vibrational states. In this way, an improved set of vibrational energies and rotational and coupling constants was obtained which allowed us to extend the assignments to new lines with higher J and K<sub>a</sub> values. The process was repeated until almost all of the  $v_1 + v_2$ and  $v_2 + v_3$  lines observable in the laboratory spectra were assigned. The rotational quantum numbers were extended up to  $J \leq 57$ , K<sub>a</sub>  $\leq 16$  for  $v_1 + v_2$  and  $J \leq 55$ , K<sub>a</sub>  $\leq 15$  for  $v_2 + v_3$ , thus greatly increasing the previous sets of observed quantum numbers (9). The experimental rotational energy levels of the (110) and (011) upper vibrational states were then obtained by adding the ground state energy levels calculated from the ground state rotational constants given in Ref. (3) to the measured line positions.

#### THEORETICAL CONSIDERATIONS AND RESULTS

A. Line Positions

According to the energy level scheme of ozone, the (110) and (011) vibrational states of this molecule can be treated as a diad of interacting states and the corresponding Hamiltonian matrix is given in Table I. The v-diagonal part consists of Watson-type Hamiltonians (one for each vibrational state) whereas the off-diagonal part contains operators which take into account the

Coriolis-type interaction affecting the rotational levels of the two interacting states. The experimental energy levels were reproduced quite satisfactorily with this Hamiltonian matrix since a standard deviation of  $0.43 \times 10^{-3}$  cm<sup>-1</sup> close to the measurement accuracy was obtained. The vibrational energies, rotational and coupling constants derived from the fit are gathered in Table II and, as noticed in previous works (2, 3, 9), it was not possible to determine simultaneously all the Coriolis coupling constants. Consequently, the  $h_{(011,110)}^{cc}$  coefficient of the operator  $iJ_y$  was fixed to the value previously used (3) for the (001) and (100) vibrational states of  $16_{0_3}$ .

From the diagonalization of the Hamiltonian matrix, it was easy to obtain the vibration-rotation wavefunctions which can be written as:

$$| \mathbf{v}_1 \mathbf{v}_2 \mathbf{v}_3 \quad \mathbf{J} \quad \mathbf{K}_a \mathbf{K}_c \rangle = \sum_{\mathbf{v}=(110),(011)} \sum_{\mathbf{K}} C_{\mathbf{K}}^{\mathbf{v}} | \mathbf{v} \rangle | \mathbf{J} \quad \mathbf{K} \quad \gamma \rangle$$
(1)

where the  $\int K \gamma >$ 's are Wang functions:

$$| J K \gamma \rangle = \frac{1}{\sqrt{2}} (| J K \rangle + \gamma | J - K \rangle) \quad K > 0, \gamma = \pm 1$$

$$| J 0 + \rangle = | J 0 \rangle.$$
<sup>(2)</sup>

From the wavefunction of a given vibrational level it is possible to calculate the mixing coefficient, % (v) on the state  $|v\rangle$  of this level, which is given by

$$\%$$
 (v) =  $\sum_{K} | C_{K}^{v} |^{2}$ .

(3)

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This coefficient indicates whether the level is affected by a resonance or not. We have plotted in Fig. 1 the mixing coefficients of some (011) rotational levels and the following observations can be made:

- a. There is a smooth behavior of the mixing coefficients with J for the series of levels with  $K_a = 12, 13, 14, 15, and 16$ .
- b. Contrary to the above series, the  $K_a = 5$  series is not regular. The large value of the mixing coefficient at J = 39 results from the accidental proximity of the resonating (110) [39 2 38] and the (011) [39 5 34] rovibrational levels.

B. Line Intensities

Using the equivalent width method, about 220 line intensities were measured with an average relative uncertainly of 10%. Although the two spectra described previously were used for the band analysis, only the higher pressure (~23.7 Torr) spectrum was used for intensity measurements. This is because many lines were too weak to be measured accurately in the lower abundance spectrum. Because the exact amount of ozone in the cell was not known, we decided to calibrate the intensities of the  $v_1 + v_2$  and  $v_2 + v_3$  bands of ozone with respect to intensities measured in the 10 µm region which contains the  $v_3$  and  $v_1$  bands (<u>3</u>). Consequently, we have measured from the same spectrum 95  $v_1$  line intensities which were calibrated using the results of Ref. (<u>3</u>). In this way, the intensities in the 5.7 µm region ( $v_1 + v_2$  and  $v_2 + v_3$  bands) are consistent with the intensities in the 10 µm region ( $v_1$ and  $v_3$  bands). The  $v_1 + v_2$  and  $v_2 + v_3$  experimental relative intensities were then reproduced using the following transformed transition moment operators (see Ref. <u>12</u> for detailed explanations):

 $v_1 + v_2$  (B-type Band)

 $(000)(110) \mu'_{Z} = (-0.5054_{6} \pm 0.0025) \times 10^{-2} \phi_{x}$ 

+ 
$$(-0.113_2 \pm 0.021) \times 10^{-4} \{i\phi_y, J_z\}$$
  
+  $(0.514_9 \pm 0.010) \times 10^{-4} \{\phi_z, iJ_y\}$  + ... (4)

 $v_2 + v_3$  (A-type Band)

$$(000)(011) \ \mu_{Z} = (-0.9081_3 \pm 0.0050) \times 10^{-2} \ \phi_{Z}$$

+ 
$$(0.93_4 \pm 0.21) \times 10^{-5} \frac{1}{2} [\{\phi_x, iJ_y\} - \{i\phi_y, J_x\}]$$
  
+  $(0.93_3 \pm 0.43) \times 10^{-5} \frac{1}{2} [\{\phi_x, iJ_y\} + \{i\phi_y, J_x\}] + \dots$   
(5)

with  $\{A, B\} = AB + BA$ 

 $\phi_{\alpha} = \phi_{z\alpha}$  (direction cosine)  $J_{\alpha} =$  component of J along the molecular axis  $\alpha$ .

All of the results are given in Debye.

The statistical analysis of the results is as follows:

 $0 < \frac{\delta I}{I_{obs.}} < 10\%$  69.1% of the lines  $10 < \frac{\delta I}{I_{obs.}} < 20\%$  24.8% of the lines  $20 < \frac{\delta I}{I_{obs.}} < 30\%$  6.1% of the lines.

where  $\delta I = |I_{obs.} - I_{calc.}|$ ,  $I_{obs.}$  is the observed, and  $I_{calc.}$  is the calculated intensity.

These results are very satisfactory since 69% of the intensities are reproduced within the experimental uncertainty of 10%. The following points of the analysis are of interest:

1) Since the line intensities are proportional to the square of the dipole moment matrix elements, only the relative signs of the transition moment constants are determined. A method to obtain their absolute signs (signs with respect to the permanent dipole moment of the molecule) would be to study the hot bands  $v_1 - v_2$  and  $v_3 - v_2$  as was done for the water molecule (13).

2) The constants, coefficients of  $\phi_x$  and  $\phi_z$ , are of the same order of magnitude which implies that the  $v_1 + v_2$  band is almost as intense as the  $v_2 + v_3$  band, which is not very common. Usually, for ozone, in a given spectral region the A-type band is much stronger than the B-type band (see the 10  $\mu$ m region for example).

3) The rotational correcting terms are not negligible particularly for the B-type band. In fact, these terms allow the calculation to reproduce quite satisfactorily the intensity peculiarities which were noted in Ref. (9).

Finally, using the vibrational energies and the rotational and coupling constants given in Table II as well as the transition moment constants given in Eq. (4) and (5), we have generated a list of 7110 line positions and intensities for the  $v_1 + v_2$  and  $v_2 + v_3$  bands of ozone. The calculations were performed<sup>1</sup> using an intensity cut-off of  $0.1 \times 10^{-24} \text{ cm}^{-1}/\text{molecule cm}^{-2}$  at 296 K, maximum values of 70 for J and 21 for  $K_a$ , and a maximum lower state energy of 3000 cm<sup>-1</sup>. Total band intensities of  $0.237 \times 10^{-19}$  and  $0.536 \times 10^{-19} \text{ cm}^{-1}/\text{molecule cm}^{-2}$  at 296 K were obtained respectively for

<sup>1</sup> A partition function Z(296) = 3473 was used in the intensity calculations.

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the  $v_1 + v_2$  and  $v_2 + v_3$  bands of  ${}^{16}O_3$ . These results agree within 10% with the low resolution integrated band intensities reported in Ref. (7), which is quite satisfactory.

Figures 2 to 4 show comparisons of the measured and calculated spectra in three different spectral regions. In all cases the observed features of the  $v_1 + v_2$  and the  $v_2 + v_3$  bands are reproduced very well by the calculations. The weak unassigned lines in Figure 2, marked by solid triangles, are believed to be transitions of the  $2v_2 + v_3 - v_2$  hot band. Similarly, the few unassigned absorptions marked by solid triangles in Figure 3 are probably lines of the  $2v_1 + v_2 - v_2$  hot band. These hot band lines are about 5% as strong as the strongest lines of the  $v_1 + v_2$  and the  $v_2 + v_3$  bands. Because of the weakness of the hot band lines in the present spectra and the absence of assignments for the  $2v_2 + v_3$  and the  $2v_1 + v_2$  bands, analysis of these features has not been attempted at this time. It should be emphasized that the absolute values of the intensities calculated here for the  $v_1 + v_2$  and  $v_2 + v_3$  bands have been calibrated against intensities from the 10 µm spectral region which have been obtained assuming the value  $\left(\frac{\partial \mu_z}{\partial q_3}\right)_e = -0.2662 D$  (see Ref. <u>3</u> for more details).

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#### TABLE I

## Hamiltonian matrices

v'= (011)	
Hermitic conj.	$\mathbf{v} = (000)$
	Watson-type Hamiltonian
Watson-type Hamiltonian H <sub>y</sub> , <sub>y</sub> ,	
	Hermitic conj. Watson-type

 $H_{vv} =$ 

 $\begin{aligned} &+ \left[ A^{v} - \frac{1}{2} \left( B^{v} + C^{v} \right) \right] J_{z}^{2} + \frac{1}{2} \left( B^{v} + C^{v} \right) J^{2} + \frac{1}{2} \left( B^{v} - C^{v} \right) J_{x,y}^{2} \\ &- \Delta_{k}^{v} J_{z}^{4} - \Delta_{J,k}^{v} J_{z}^{2} J^{2} - \Delta_{J}^{v} \left( J^{2} \right)^{2} - \delta_{k}^{v} \left( J_{z}^{2} , J_{x,y}^{2} \right) - 2\delta_{J}^{v} J_{x,y}^{2} J^{2} \\ &+ H_{k}^{v} J_{z}^{6} + H_{k,J}^{v} J_{z}^{4} J^{2} + H_{J,k}^{v} J_{z}^{2} \left( J^{2} \right)^{2} + H_{J}^{v} \left( J^{2} \right)^{3} \\ &+ h_{k}^{v} \left( J_{z}^{4} , J_{x,y}^{2} \right) + h_{k,J}^{v} \left( J_{z}^{2} , J_{x,y}^{2} \right) J^{2} + 2h_{J}^{v} J_{x,y}^{2} \left( J^{2} \right)^{2} \\ &+ L_{k}^{v} J_{z}^{6} + L_{k,K,J}^{v} J_{z}^{6} J^{2} + L_{k,J}^{v} J_{z}^{4} \left( J^{2} \right)^{2} + L_{k,J,J}^{v} J^{2} \left( J^{2} \right)^{3} + L_{J}^{v} \left( J^{2} \right)^{4} \\ &+ J_{k}^{v} \left( J_{z}^{6} , J_{x,y}^{2} \right) + I_{k,J}^{v} \left( J_{z}^{4} , J_{z}^{2} \right) J^{2} + J_{J,k}^{v} \left( J_{z}^{2} , J_{x,y}^{2} \right) J^{2} + 2 J_{J}^{v} J_{x,y}^{2} \left( J^{2} \right)^{3} \\ &+ P_{k}^{v} J_{z}^{10} + P_{k,K,J}^{v} J_{z}^{9} J^{2} + P_{k,K,J}^{v} J_{z}^{6} \left( J^{2} \right)^{2} + \dots \\ &+ P_{k}^{v} \left( J_{z}^{10} + P_{k,K,J}^{v} \right) J^{8} J^{2} + P_{k,K,J}^{v} J_{z}^{6} \left( J^{2} \right)^{2} + \dots \\ &+ P_{k}^{v} \left( J_{z}^{10} , J_{x,y}^{2} \right) + P_{k,K,J}^{v} \left( J_{z}^{6} , J_{x,y}^{2} \right) J^{2} + \dots \\ &+ P_{k}^{v} \left( J_{z}^{10} , J_{x,y}^{2} \right) + P_{k,K,J}^{v} \left( J_{z}^{10} , J_{x,y}^{2} \right) J^{2} + \dots \\ &+ P_{k}^{v} \left( J_{z}^{10} , J_{x,y}^{10} \right) + h_{v,v}^{v} \left( J_{z}^{10} , J_{x,y}^{2} \right) J^{2} + \dots \\ &+ P_{k}^{v} \left( J_{z}^{10} , J_{x,y}^{10} \right) + h_{v,v}^{v} \left( J_{z}^{10} , J_{x,y}^{2} \right) J^{2} + \dots \\ &+ h_{v,v}^{v} \left( J_{x,y}^{10} J_{y}^{10} + P_{v,v}^{v} \left( J_{z}^{10} J_{y}^{10} + h_{v,v}^{v} J_{y}^{10} J_{z}^{10} J$ 

with {A B} = AB + BA  $J_{xy}^{2} = J_{x}^{2} - J_{y}^{2}$   $J_{\pm} = J_{x} \mp iJ_{y}$ 

# Table II. Vibrational energies, rotational and coupling constants

	Vibrational and Rotation	nal Constants				
	(110)	(011)				
E <sub>v</sub>	$1796.26187_0 \pm 0.000057$	$1726.52249_0 \pm 0.000075$				
AV	$3.6108121_{34} \pm 0.0000035$	$3.5523057_{11} \pm 0.0000053$				
Br	$0.44144616_{08} \pm 0.00000050$	0.43989711 <sub>75</sub> ± 0.00000055				
cv	0.39029870 <sub>66</sub> ± 0.00000043	0.38850339 <sub>96</sub> ± 0.00000046				
N K	$(0.239727_{92} \pm 0.000068) \times 10^{-3}$	$(0.22865_{41} \pm 0.00012) \times 10^{-3}$				
V KJ	$(-0.20752_{54} \pm 0.00091) \times 10^{-5}$	$(-0.1598_{22} \pm 0.0011) \times 10^{-5}$				
$\mathbf{v}_{\mathbf{v}}^{\mathbf{r}}$	$(0.46689_{33} \pm 0.00025) \times 10^{-6}$	$(0.45985_{71} \pm 0.00023) \times 10^{-6}$				
S V K	$(0.4321_{19} \pm 0.0042) \times 10^{-5}$	$(0.3495_{31} \pm 0.0010) \times 10^{-5}$				
S <sup>V</sup> J	$(0.6472_{74} \pm 0.0020) \times 10^{-7}$	$(0.7609_{92} \pm 0.0023) \times 10^{-7}$				
ł <sup>ν</sup> K	$(0.4996_{50} \pm 0.0048) \times 10^{-7}$	$(0.464_{48} \pm 0.010) \times 10^{-7}$				
I <sup>V</sup> KJ	$(-0.2103_{52} \pm 0.0059) \times 10^{-8}$	$(-0.1875_{97} \pm 0.0027) \times 10^{-8}$				
I <sup>V</sup> JK	$(-0.83_{76} \pm 0.13) \times 10^{-10}$	$(-0.311_{10} \pm 0.023) \times 10^{-10}$				
I <sup>V</sup> J	$(0.169_{22} \pm 0.079) \times 10^{-12}$	$(0.113_{62} \pm 0.059) \times 10^{-12}$				
V K	$(-0.27_{93} \pm 0.11) \times 10^{-8}$					
v . KJ	$(-0.33_{01} \pm 0.16) \times 10^{-10}$					
r J	$(0.91_{60} \pm 0.53) \times 10^{-13}$	$(0.199_{99} \pm 0.059) \times 10^{-12}$				
v K	$(-0.126_{35} \pm 0.010) \times 10^{-10}$	$(-0.107_{62} \pm 0.027) \times 10^{-10}$				
v KKJ	$0.2739 \times 10^{-12}$	$0.2739 \times 10^{-12}$				
, JJK	$-0.3251 \times 10^{-15}$	$-0.3251 \times 10^{-15}$				
v J	$0.2994 \times 10^{-17}$	$0.2994 \times 10^{-17}$				
v K	$0.1812 \times 10^{-12}$	$0.1812 \times 10^{-12}$				

for the (110) and (011) vibrational states of  ${}^{16}\mathrm{O}^{}_3$ 

	(110)	(011)				
PK	$0.333 \times 10^{-14}$	$0.333 \times 10^{-14}$				
Р <sup>V</sup> КККЈ	$0.1895 \times 10^{-15}$	$0.1895 \times 10^{-15}$				
	Coupling Co	nstants				
	$h^{c}_{(011)(110)} = (-0.10099_{7})$	$6 \pm 0.00016) \times 10^{-1}$				
	$h_{(011)(110)}^{,c} = -0.470$					
	$h''^{c}_{(011)(110)} = (0.855_{10} \pm$	0.046) × $10^{-6}$				
	$h_{(011)(110)}^{\prime,\prime,c} = (-0.1062_{48})$	$\pm$ 0.0063) × 10 <sup>-6</sup>				
	$h_{(011)(110)}^{h_{(011)}c} = (-0.59_{31} \pm$	0.23) $\times 10^{-8}$				
•	(/)					

Table II. Concluded

All the results are in  $cm^{-1}$  and the quoted errors are one standard deviation. The constants without errors were held fixed during the calculation; they come from Ref. (3).

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List of Symbols

- v lower case Greek nu  $\mu$  - lower case Greek mu  $\delta$  - lower case Greek delta
- $\theta$  lower case Greek theta
- $\phi$  lower case Greek phi
- Δ upper case Greek delta
- $\partial$  symbol for partial derivative
- $\alpha$  lower case Greek alpha
- γ lower case Greek gamma

# Figure Captions

Figure 1.

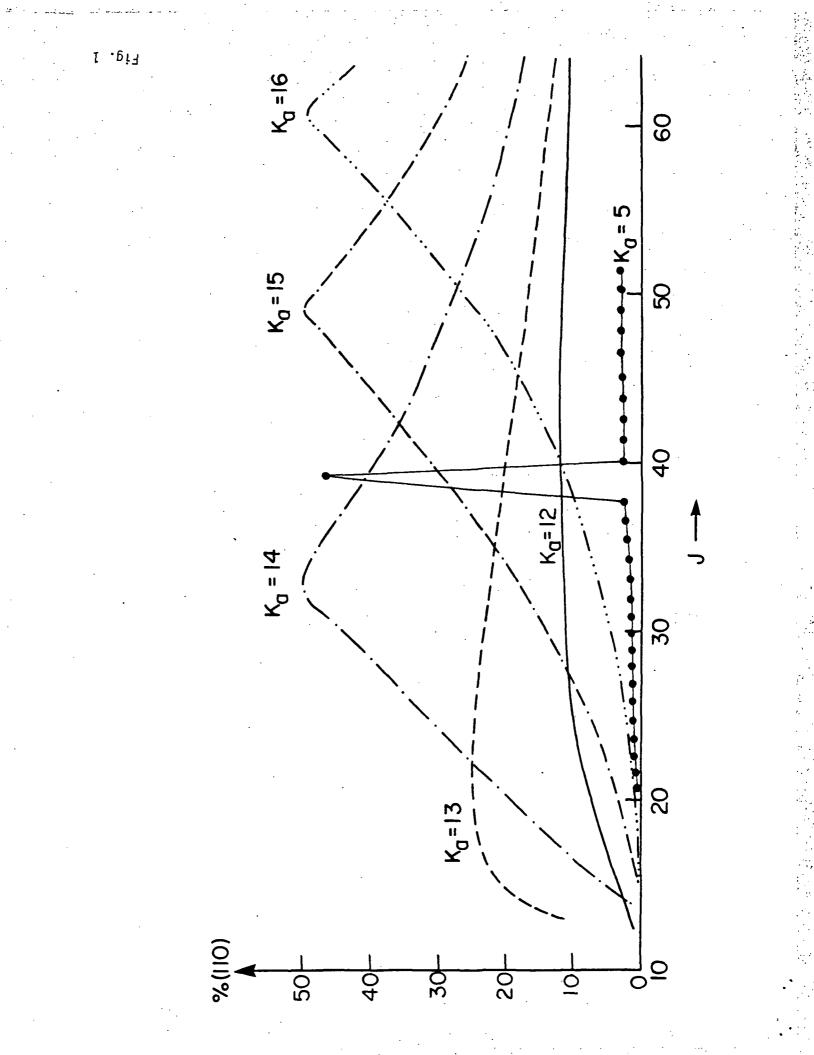
Mixing coefficients in % (110) for some rotational levels of the (011) state. For high K<sub>a</sub> values, the mixing coefficients vary smoothly with J; a crossing of levels occurs for  $K_a = 14$ , 15, and 16. For  $K_a = 5$ , the mixing is weak except for J = 39; the large mixing coefficient is caused by the accidental proximity of the (110) [39 2 38] and the (011) [39 5 34] rovibrational levels. Figure 2. Laboratory (solid line) and best-fit calculated (crosses) spectra in the region of the  $v_2 + v_3$  band around 1682 cm<sup>-1</sup> (lower panel). The laboratory spectrum was recorded with about 23.7 Torr of  $^{16}O_3$  in a 50-cm long absorption path at room temperature. Residuals (observed minus calculated) are shown in the upper panel. The  $v_1 + v_2$  line [39 2 38] + [40 5 35] located at 1681.838  $cm^{-1}$  appears only because it borrows its intensity from

the  $v_2 + v_3$  line [39 5 34] + [40 5 35] at 1682.007 cm<sup>-1</sup>. Although the  $\Delta K_a = 3 v_1 + v_2$  line is blended by a weaker hot. band line, it can be seen that the intensities of the two resonating lines are almost equal because of a mixing coefficient of about 50% (see Figure 1). Several high K<sub>a</sub> lines of the  $v_2 + v_3$  band (marked with an asterisk) are also observable in this region. Unassigned absorptions in this region are believed to be from the  $2\nu_2 + \nu_3 - \nu_2$  hot band and are indicated by solid triangles.

Figure 3.

Laboratory (solid line) and best-fit calculated (crosses) spectra in the region of the  $v_1 + v_2$  band around 1832 cm<sup>-1</sup> (lower panel). The laboratory spectrum was recorded with about 23.7 Torr of  $160_3$  in a 50-cm absorption path at room temperature. Residuals (observed minus calculated) are shown in the upper panel. High J lines of the  $v_1 + v_2$  band (\*) are observed in this spectral region. Their absorption is well reproduced by the calculations. Weak unassigned lines in this region are believed to be from the  $2v_1 + v_2 - v_2$  hot band and are indicated by solid triangles. Laboratory (solid line) and best-fit calculated (crosses) spectra in the region of the  $v_1 + v_2$  band around 1860 cm<sup>-1</sup> (lower panel). The laboratory spectrum was recorded with about 23.7 Torr of  ${}^{16}O_3$  in a 50-cm long absorption path at room temperature. Residuals (observed minus calculated) are shown in the upper panel. The  ${}^{R}Q_{K_{a}^{"}=10}$  line series (\*) is clearly visible and is well reproduced by the calculations.

Figure 4.



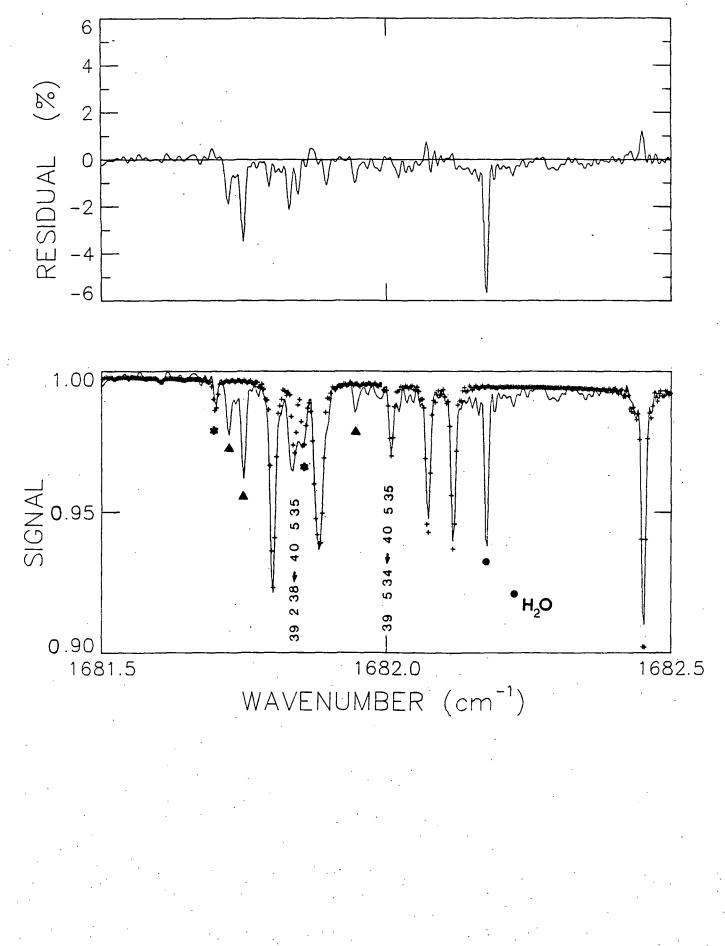
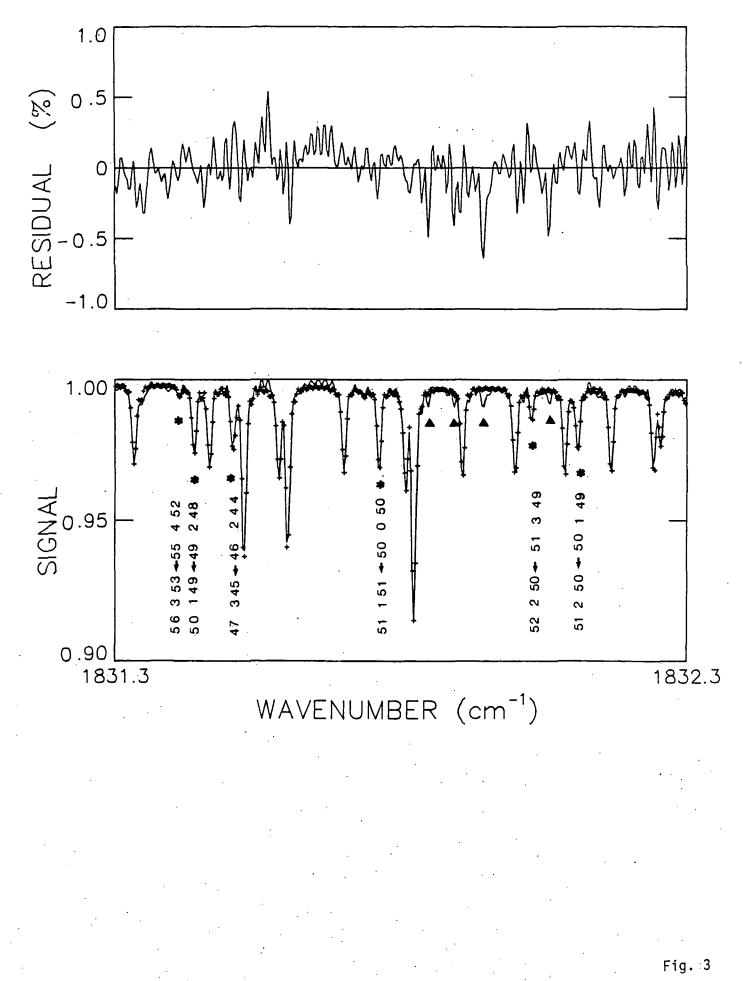


Fig. 2



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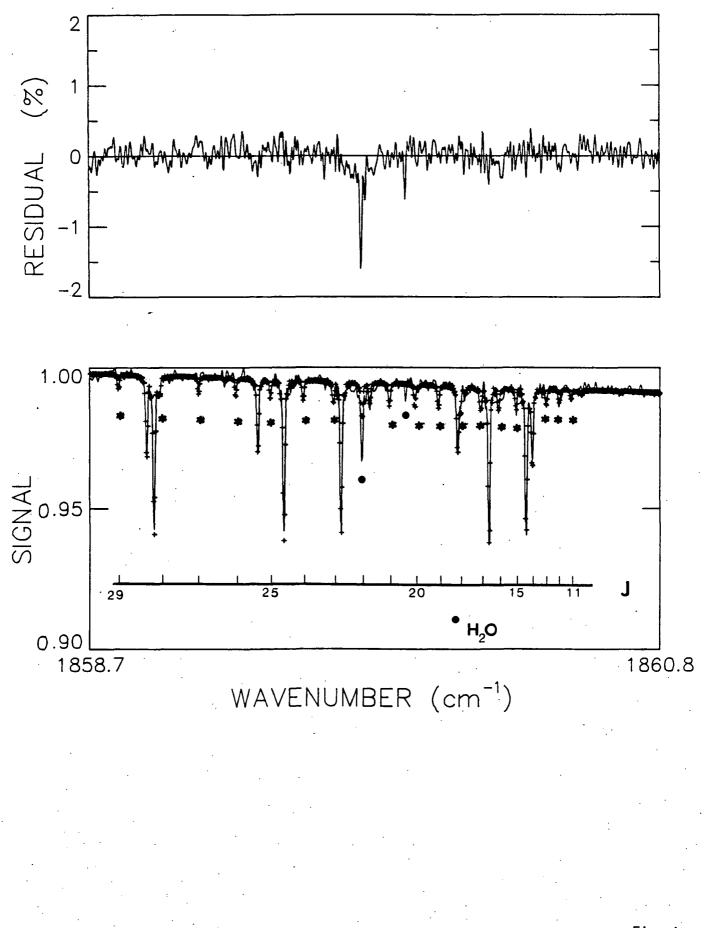


Fig. 4