## MODAL IDENTITIES FOR MULTIBODY ELASTIC SPACECRAFT -an aid to selecting modes for simulation

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## ABSTRACT

This paper answers the question: Which set of modes furnishes a higher fidelity math model of dynamics of multibody, deformable apacecraft-hinges-free or hinges-locked vehicle modes? Two sets of general, discretized, linear equations of motion of a spacecraft with an arbitrary number of deformable appendages, each arciculated directly to the core body, are obtained using the above two families of modes. By a comparison of these equations, ten sets of modal identities are constructed which involve modal momenta coefficients and frequencies associated with both classes of modes. The sums of infinite series that appear in the identities are obtained in terms of mass, and first and second moments of inertia of the appendages, core body, and vehicle by using certain basic identities concerning appendage modes. Applying the above identities to a four-body spacecraft, the hinges-locked vehicle modes are found to yield a higher fidelity model than hinges-free modes, because the latter modes have nonconverging modal coefficients--a characteristic proved and illustrated in the paper.

## I. INTRODUCTION

The use of appendage modes for simulating dynamics and control of multibody flexible spacecraft is widespread, in as much as they are eminently suitable for both small angle (linear) and large angle (nonlinear) dynamics. To win this benefit, however, a simulation engineer must retain a sufficient number of these modes for each appendage so that the simulation program has acceptable fidelity. When there are a large number of appendages in a spacecraft, and/or an appendage has a large mass and moment of inertia relative to those of the rigid core body of the spacecraft, the total number of appendage modes for a high accuracy model may become unacceptably great (Reference 1), possibly diminishing the utility of the appendage modes for simulation. Furthermore, control systems for a multibody spacecraft are most easily designed by considering one axis of one body at a time, because different bodies generally serve different purposes and so the control systems' intrinsic features are generally quite different. Having designed them so, to ensure they all perform as desired in the mutual presence and in the presence of flexibility, a compact mathematical model of the entire spacecraft's dynamics is desired so that the control designs can be refined fast and economically about all axes. For this purpose, the linear, small angle models of spacecraft flexible dynamics are just right, and so the engineer could beneficially employ the vehicle modes of the spacecraft. Hughes ${ }^{2}$ conceived of two families of vehicle mades for multibody spacecraft: "hinges-free" and "hinges-locked" vehicle modes (although he does not use this terminology). By definition, hinges-

Eree modes are obtained by leaving all hinges free, that is, unlocked and unforced, so that the associated natural vehicle modes may contain motion of the articulated bodies relative to the inboard bodies. Conversely, in the hinges-locked modes, the relative motion of the articulated bodies is, by definition, zero, and some force or torque is applied at the hinges to keep the motion 30. In Reference 3, these vehicle modes are formulated, and their zero linear and angular momentum properties, the orchogonality condicions, and the associated modal momenta coefficients are theorized.

A critical question whose answer is sought in this paper is: Between the hinges-free and hingeslocked vehicle modes, which one furnishes a higher fidelity dynamic model, retaining the same number of modes in the simulation? To this end, a multibody spacecraft is considered in chis paper that consists of a rigid core body, and N flexible appendages, each articulated directly to the core body. Three sets of discrete motion equations of this spacecraft are obtained from a continuum set by using appendage modes, hinges-free vehicle modes, and hinges-locked vehicle modes. To compare the last two Eamilies of modes, modal identities are devised that express the sum of contribution of all infinite number of modes in terms of first and second moments of inertia of the articulated bodies, the core body, and the vehicle, following Hughes ${ }^{4}$. The analysis is amply illustrated, and definitive conclusions are summarized at the end of the paper. Although for concreteness, the paper considers a multibody spacecraft with level-i articulated bodies (the terminology of $\mathrm{Ho}^{5}$ ), it will be clear that the conclusions dram apply to a wider range of multibody spacecraft.

## II. FORMULATION OF CONTINUUM EQUATIONS OF MOTION

Fig. 1 portrays an $N+1$-body spacecraft that consists of a three-axis stabilized core rigid body $B_{0}$, and deformable bodies $E_{1}, \ldots, E_{N}$, each articulated directly to the core body. The motion equations will be developed with respect to the


Figure 1. An N+1-Body Spacecraft With N Articulated, Deformable Appendages

[^0]reference point $O$ in Fig. 1 which is neither the mass center $\oplus_{0}$ of the body $B_{0}$, nor the mass center $\oplus$ of the entire vehicle $V$. This generality in the formulation is warranted because the NASTRAN modal data corresponding to such multibody spacecraft are often with respect to an arbitrary reference node 0 , and the mass centers are generally nodeless empty points. The mass of each body is denoted $m_{p}(p=0,1, \ldots, N)$; the mass of all N articulated bodies together, $m_{e}$; and the mass of antire spacecraft, $m$; clearly, $m=m_{0}+m_{e}$. The first moment of mass of $B_{0}$ relative to 0 is $\underline{c}_{0}$, and those for the hinged bodies ( $j=1, \ldots, N$ ), measured from the respective hinges $0_{j}$, are denoted $\underline{c}_{j}$. Similar to $g_{p}(p=0,1, \ldots, N)$, the vector $\stackrel{s}{s}_{-0}$ emanates from 0 and $\underline{r}_{j}$ from $O_{j}$. Note that the subscript $p$ covers all bodies, while $j$ covers only the articulated badies. The rectors b. ( $j=1, \ldots, N$ ) originating from 0 locate the hinges $O_{j}$ of the hinged bodies $E_{j}$. The first moment of inertia of the entire spacecraft, then, is
$\underline{c}=\underline{c}_{0}+\sum_{j}\left(m_{j} \underline{b} j+\underline{c}_{0} j \underline{c} \underline{j}\right) \underline{\Delta} \underline{c}_{0}+\underline{c}_{e} ; \sum_{j}=\sum_{j=1}^{N}$
where the matrix $C_{0 j}$ transforms the $E_{j}-f i x e d$ vector $c$ to a $B_{0}-f i x e d$ vector, and the vectors ${ }^{-}$; are expressed in the $B_{0}$-fixed frame. Next, $J_{0}$ denotes the inertia matrix of the body $B_{0}$ about the reference point 0 , while $J j$ is the inertia matrix of the hinged body $E j$ in its own frame about the hinge 0 . . The inertia matrix of $E$. expressed at the reference point 0 in the $B_{0}$-fixed frame is denoted $\mathrm{J}_{\mathrm{j}}^{\mathrm{J}}$ and
\[

$$
\begin{gather*}
\underline{J}_{j}^{0}=\underline{c}_{0 j} \underline{J}_{j} \underline{c}_{j 0}-\left(m_{j} \underline{b}_{j}^{x} \underline{b}_{j}^{x}+\underline{b}_{j}^{x}\left(\underline{-}_{0} j \underline{c}_{j}\right)^{x}+\right. \\
\left.\left(\underline{c}_{j} \underline{c}_{j}\right)^{x} \underline{b}_{j}^{x}\right) \tag{2}
\end{gather*}
$$
\]

where (*) means the $3 \times 3$ skew-symmetrix matrix associated with the vector (•). The inertia matrix $J$ of the entire renicie at the poinc 0 will then be

$$
\begin{equation*}
\underline{J}=\underline{J}_{0}+\left[\underline{J}_{j}^{0} \triangleq \underline{J}_{0}+\underline{J}_{e}^{0}\right. \tag{3}
\end{equation*}
$$

Anticipating our later needs, the cross inertia matrix $J_{0}$ between the bodies $B_{0}$ and $E_{j}$ expressed in the $\bar{E}_{j}^{0 j}$ fixed frame equals

$$
\begin{equation*}
{\underset{\sim}{J}}_{j} \triangleq{\underset{J}{j}}_{j}-\left(\underline{c}_{j 0} \underline{b}_{j}\right)^{x} \underline{c}_{j}^{x} \tag{4}
\end{equation*}
$$

As for the motion of the spacecraft, its mass center is assumed to perform some orbital motion, not coupled with its attitude motion under consideration. To develop motion equations, the local orbital frame is taken to be an inertial frame. The kinetic quantities of interest are: $\underline{V}_{0}(r)$. the percurbational velocity of the reference point 0 over the uniform orbital motion at time $t ; \omega_{0}(t)$, the inertial angular velocity of $B_{0} ; \underline{\Omega}_{j}(c)$, che angular velocity of each articulated body $E_{j}$ relative to $B_{0}$ at the hinge $0_{j}$; and $\underline{u}_{j}\left(\underline{r}_{j}, t\right)$, the deformation of $E ;$ at the location $\underline{f}_{j} E j$. These quantities are taken to be linear, first order, infinitesimal, 50 that
their products can be ignored in the analysis. The external forces and torques acting on the spacecraft are the force $\underline{f}_{0}(t)$ and the torque $g_{0}(t)$ acting on $B_{0}$ at 0 , and the Force $\frac{f}{j}_{j}(t)$ and the corque $g_{j}(t)$ on each $E_{j}$ at the hinge $0{ }_{j}$. The latter pair, ( $\underline{f}_{j}, g_{j}$ ) , includes a distributed force $\mathcal{L}_{j}\left(r_{j}, t\right)$ acting in the domain of the body $E_{j}$, anc if ${\underset{j}{j}}^{(\underline{r}} \underline{J}_{j}, t)$ is the only surface force acting on $E_{j}$, then

$$
\begin{align*}
& \underline{f}_{j}(t)=\dot{j} \underline{L}_{j}\left(\underline{r}_{j}, t\right) d A  \tag{5}\\
& \underline{g}_{j}(t)=\int_{j} \underline{r}_{j}^{x} f_{j}\left(\underline{r}_{j}, t\right) d A
\end{align*}
$$

where dA is an elemental area of $E_{j}$ and $\int_{j} \triangleq \int_{E}$. With the aid of the Dirac delta function and its derivative, the distributed force $Q_{j}(\underline{r}, t)$ also represents a distributed moment. Regarding the control forces and corques, those acting on $B_{0}$ are included in the quantities ${\underset{\sim}{f}}_{0}$ and $g_{0}$, whereas, if a control force or torque is produced in the interior domain of $E$; without acting against the core body $B_{0}$, then that is included in the pair ( $\underline{f}_{j}, \underline{g}_{j}$ ); however, if, for instance, the torque is produced by an electric motor which rests on $B_{0}$ at the interface $O_{\text {; }}$ and exerts on $E j$, then this is considered separately and denoted $\mathfrak{g}_{0}(t)(j=1, \ldots, N)$, for it produces a reaction torque $-g_{0}(r)$ which acts on $B_{0}$. The total force $f(t)$ and torque $f(t)$ that act on the vehicle are

where, of course, $g(t)$ does not include the control corque $\underline{g}_{0}(t)$ at the interface $0_{j}$.

The elastic spacecraft under consideration is relatively simple; it is straightforward to develop its linear, continumm motion equations following Hughes $4,6,7$. The equations governing the discrete variables $\underline{V}_{0},-_{0}, \frac{\Omega}{-} j \quad(j=1, \ldots, N)$ are:

$$
\begin{aligned}
& g_{j}(j=1, \ldots, N)
\end{aligned}
$$

where an overdot indicates differentiation with respect to time. To write the motion equation governing the deformation $\frac{1}{j}(\underline{r}, r)$ of the flexible body $E_{j}$, denote the related linear stiffness operator by $L_{j}$; the body $E_{j}$ is allowed to be anisotropic andor nonhomogeneous, and its mass
density is denoted $a_{j}\left(\underline{r}_{j}\right)$. The concinumm motion equation governing the detormation $\underline{u} j$ is chen:

$\left.\underline{r}_{j}^{x} \underline{\Omega}_{j}+\underline{u}_{j}\right]=\underline{b}_{j}\left(\underline{r}_{j}, t\right)(j=1, \ldots, N)$

The continum motion equations (7) and (8) are discretized in the next section.

## III. DISCRETIZATION OF CONTINUUM EQUATIONS OF MOTION

Three families of modes will be employed in this section for digcretization: (1) appendage modes, (2) hinges-free vehicle modes, and (3) hinges-locked vehicle modes. The use of appendage modes is standard; they are employed here in order to evaluate the infinite sums that appear in the hinges-free and hinges-locked modal identities in Section IV in terms of mass, and. first and second moments of inertia of the appendages, core body, and the vehicle.

## discretization by appendace modes

Following Hughes ${ }^{4}$, define the modal momenta coefficients $\underline{P}_{j a}$ and $\underline{H}_{j \sigma}$ concerning the appendage (cantilever) modes $\underline{U}_{j \sigma}^{a}\left(\underline{r}_{j}\right)$ of the articulated body $E_{j}$ :
$\underline{P}_{j \sigma} \stackrel{\Delta}{\equiv} \int_{j} \underline{U}_{j \sigma}^{a}\left(\underline{r}_{j}\right) d m \quad(j=1, \ldots, N ; \sigma=1, \ldots, \infty)$
$\underline{H}_{j \sigma} \triangleq \int_{j} \underline{r}_{j}^{\underline{U}} \underline{j}_{j \sigma}^{a}\left(\underline{r}_{j}\right) d m$
where $d m=$ elemental mass. The coefficient $\underset{\sim}{P} j$ o is associated with linear momentum and $\underline{H}_{j o w i t h}$ angular momentum of the mode at che hinge point $O_{j}$. The modal angular momentum coefficient relative to the reference point 0 (Fig. 1) is defined as

Then the continuum equations (7) and (8) discretize to




$(j=1, \ldots, N ; a=1, \ldots, \infty)$
where the superscript $T$ indicates transpose of the quantity; $Q_{j o}(t)$ is the modal coordinate and $Q_{j \sigma}^{C}$ is the frequency associated with the $\sigma^{j \sigma}$ th appendage mode $\underline{U}_{j o}^{a}\left(\underline{r}_{j}\right)$ of the body $E_{j}$;
and $Y_{j \sigma}^{a}(t)$ is the modal input to that mode:

$$
\begin{equation*}
\gamma_{j \sigma}^{a}(t)=\int_{j} \underline{u}_{j \sigma}^{a}\left(\underline{r}_{j}^{T}\right) \underline{f}_{j}\left(\underline{r}_{j}, t\right) d A \tag{12}
\end{equation*}
$$

Eqs. (1l) are a generalization of Eq. (35) of Hughes ${ }^{4}$, for the former include articulation motion of the appendages. Much more complex and general equations than Eqs. (11) are available in the vast literature on deformable multibody dynamics; see, for instance, the works of $\mathrm{HO}^{5}$, and Singh et al ${ }^{8}$ on spacecraft with arbitrary tree topology. Eqs. (11) nevercheless may boast of simplicicy which is eminently useful while designing the control systems for articulated bodies. More importancly though, Eqs. (ll) are derived here because in Section IV they will aid in developing modal identities. To facilitate this task, Eqs. (11) are abbreviated by using the definitions


$\left.\underline{P}_{j}^{T} \quad \underset{\underline{B}}{ } \quad \underline{P}_{j 1} \quad \underline{P}_{j 2} \cdots\right]$,
$\underline{P}_{A}^{0^{T}} \triangleq\left[\underline{C}_{01} \underline{P}_{1}^{T}, \cdots, \underline{C}_{0 N} \underline{p}_{N}^{T}\right]$
$\underline{H}_{j}^{0^{T}} \triangleq\left[\underline{H}_{j 1}^{0} \quad \underline{H}_{j 2}^{0} \cdots \mid\right.$,
$\underline{H}_{A}{ }^{0^{T}} \triangleq\left[\underline{H}_{1}^{0^{T}}, \cdots, \underline{H}_{N}^{0^{T}}\right]$
$\underline{H}_{\mathrm{H}}^{\mathrm{T}} \quad \stackrel{\Delta}{\underline{B}} \quad\left[\begin{array}{lll}\underline{\mathrm{H}}_{\mathrm{j} 1} & \underline{\mathrm{H}}_{\mathrm{j} 2} & \cdots\end{array}\right]$,

$g_{j}^{T} \triangleq\left[Q_{j 1} \quad Q_{j 2} \ldots\right], \quad g_{A}^{T} \quad\left[\begin{array}{l}Q_{1}^{T}\end{array} \cdots, g_{N}^{T}\right]$,
$\mathbf{g}_{\mathbf{A}}^{\mathbf{T}} \triangleq \quad\left[\begin{array}{llll}\mathbf{g} \\ \mathbf{g} & \cdots & \left.\mathbf{g}_{\mathbf{N}}^{\mathrm{T}}\right],\end{array}\right.$
$\underset{g_{0 A}}{T} \triangleq\left[\begin{array}{lll}\underline{g_{01}} & \cdots & \mathbf{g}_{0 N}^{T}\end{array}\right], \quad \underline{\Omega}_{A}^{T} \triangleq\left[\begin{array}{lll}\underline{\underline{\Omega}} \\ \underline{1} & \cdots & \underline{\Omega}_{N}^{T}\end{array}\right]$
$\left.\underline{\Omega}_{j}^{c} \quad \underline{\operatorname{diag}} \mid \Omega_{j 1}^{c} \Omega_{j 2}^{c} \ldots\right], \quad \underline{\Omega}_{c} \quad \underline{\underline{a}} \operatorname{diag}\left[\underline{\Omega}_{1}^{c} \ldots \underline{\Omega}_{N}^{c} \mid\right.$

Eqs. (11) then take this concise form:

$\underline{c}^{x} \underline{\theta}_{0}+\underline{J} \underline{\dot{w}}_{0}+\underline{J}_{-0 A}^{T} \ddot{\theta}_{A}+\underline{H}_{A}^{0 T_{\infty}} \underline{Q}_{A}=g$



It is interesting to compare Eqs. (14) with Eqs. (43) of Hughes ${ }^{4}$. For even more compaction of these equations, the following matrices are introduced:

where 1 is a $3 \times 3$ identity matrix. Eqs. (14) chereby reduce to the following three matrix equations: one governing six overall degrees of freedom of the spacecraft, $g_{V}(t)$; the second governing $n_{a} x l_{\text {vector }} \Omega_{A}$ of $n_{a}$ relative angular velocities of $N$ articulated bodies; and the third governing the oxl vector $g_{A}$ of modal coordinates of appendage modes of all articulated bodies.
$\underline{M}_{V V} \ddot{q}_{V}+\underline{M}_{V A}^{T} \underline{\underline{g}}_{A}+\underline{T}_{A}^{0}{ }^{T} \underline{g}_{A}=\underline{u}_{V}(t)$
$\underline{\underline{M}}_{V A} \ddot{\underline{I}}_{V}+\underline{-}_{A} \ddot{\underline{I}}_{A}+\underline{\underline{H}}_{A}^{T} \underline{\underline{G}}_{A}=\underline{g}_{A}(t)+\underline{g}_{O A}(t)$

Modal identities associated with the modal momenta matrices $\mathbb{P}_{-A}^{0}$ and $\underline{H}_{A}$ are derived in Section IV.
discretization by hinces-free vehicle modes
In this technique, the continuum equations (7) and (8) are discretized all at once. For this purpose, the following modal expansion is postulated for the variables in Eqs. (7) and Eq. (8) (Reference 3):

$$
\begin{align*}
& \underline{v}_{0}(t)=\underline{\underline{q}}_{0}(t)+\Sigma \underline{x}_{0 v} \dot{\eta}_{v}(t), \\
& \underline{w}_{0}(t)=\dot{\underline{Q}}_{0}(t)+\Sigma \Phi_{0 v} \dot{\eta}_{v}(t), \\
& \underline{\underline{q}}_{j}(t)=\underline{\underline{\theta}}_{j}(t)+\sum \Phi_{j v} \dot{\eta}_{v}(t), \\
& \underline{u}_{j}\left(\underline{r}_{j}, t\right)=\sum \underline{u}_{j v}\left(\underline{r}_{j}\right) \eta_{v}(t), \sum=\sum_{\substack{v=1 \\
(j=1,2, \ldots, N}},
\end{align*}
$$

where $\underline{R}_{0}, \underline{\theta}_{0}$, and $\underline{\Theta}_{j}$ are the temporal coordinates for the rigid modes of the spacecraft; the total number of articulation degrees of freedom is $n_{a}$, so there are $n_{a}+6$ rigid modes in all. Furthermore, $\underline{R}_{0}$ is the translation of the reference point 0 , and $\underline{\theta}_{0}$ is the rotation of the spacecraft, both in rigid modes; similarly, $\underline{\theta}_{j}$ is the rotation of the hinged body $E_{j}$ relative $\mathrm{co}^{0} B_{0}$ at the hinge $0_{j}(j=1, \ldots, N)$ in arigid mode. The quantities $X_{0 v}, \Phi_{0 v}$, and $\Phi_{j v}(j=1, \ldots, N ; v=1, \ldots, \infty)$ are $v-t h$ modal coefficients contributing, respectively, to overall discrete motions ${\underset{\sim}{V}}_{0},{\underset{\sim}{w}}_{0}$, and $\underline{\Omega}_{j} ;$ and $\eta_{y}(t)$ is the associated modal coordinate. The eigenfunction $\underline{U}_{j}\left(\underline{r}_{j}\right)$ is that part of the hingesfree vehicle mode, denoted $\underline{w}_{v}(\underline{r})$, which defines the deformation of body $E_{j}$ in $v-t h$ mode. Although $\underline{U}_{\mathrm{j} v}\left(\underline{r}_{\mathrm{j}}\right)$ satisfies the condition of zero displacement and zero rotation at the hinge $O_{j}$, that is, $\underline{U}_{j v}\left(0_{j}\right)=\underline{0}$ and $\frac{1}{2} \underline{Z}^{x} \underline{U}_{\mathrm{j} v}\left(0_{j}\right)=\underline{0}$, it is not the same as the a-ch appendage mode $\underline{u}_{j}{ }^{\mathrm{a}}\left(\mathrm{r}_{\mathrm{j}}\right)$ used
before, because in the case of $\underline{U}_{j v}\left(\underline{r}_{j}\right)$, no torque acts, by definition, at the hinge $0_{j}$ to enforce the zero deformation and zero rotation condition, whereas in the case of $\underline{U}_{j a}^{a}\left(\underline{r}_{j}\right)$, the immobile support of the appendage enforces that condition. Because of the mobile support of the hinge $0_{j}$, the total motion $\underline{W}_{j v}\left(\underline{I}_{j}\right)$ of $E_{j}$ in an inertial frame is

$+\underline{C}_{0} j \underline{U}_{j v}\left(\underline{x}_{j}\right) \quad(j=1, \ldots, N ; v=1, \ldots, \infty)$
where the first two terms in the right side are because of the translation and rotation of the core body in the $v$-th mode, and the third term ${ }^{x}{ }_{j}{ }^{\phi}{ }_{j v}$ is caused by the relative rotation of $E_{j}$ at the free hinge $0_{j}$. The motion of the core body in $u$-th mode is given simply by

$$
\begin{equation*}
\underline{W}_{0 v}\left(\underline{r}_{0}\right)=\underline{x}_{0 v}-\underline{r}_{0}^{x} \Phi_{0 v} \tag{19}
\end{equation*}
$$

Thus a hinges-free vehicle mode ${\underset{W}{V}}^{(r)}$ repans entire spacecraft such that
$\underline{W}_{V}(\underline{r})= \begin{cases}\underline{W}_{0 V}\left(\underline{r}_{0}\right), & \text { if } \underline{r}=\underline{r}_{0} \\ \left.\underline{W}_{j v} \underline{\underline{r}}_{j}\right), & \text { if } \underline{r}=\underline{b}_{j}+\underline{C}_{0 j} \underline{r}_{j}\end{cases}$

$$
\begin{equation*}
(j=1, \ldots, n ; v=1, \ldots, \infty) \tag{20}
\end{equation*}
$$

Following Hughes ${ }^{2}$, the $6+n_{a}$ rigid modes of a spacecraft wich arciculated bodies are 1, - $\underline{r}^{\mathrm{x}}$, and $\underline{\underline{r}}_{\mathrm{j}}^{\mathrm{j}}(\mathrm{j}=1, \ldots, \mathrm{n})$. Not surprisingly, the elastic modes ${\underset{W}{v}}^{(\underline{r}}(\underline{a})(v=1, \ldots, \infty)$ are orthogonal to these rigid modes; that is:

$$
\begin{gathered}
\int_{V} \underline{W}_{v}(\underline{r}) d m=\underline{0}, \quad \int_{V} \underline{\tau}^{x} \underline{W}_{v}(\underline{r}) d m=\underline{0} ; \\
\int_{j} \underline{\underline{r}}_{j} \underline{W}_{j v}\left(\underline{r}_{j}\right) d m=\underline{0} \quad(21 . a, b, c)
\end{gathered}
$$

where $\int_{V}$ means the entire vehicle is the domain of integration. Eqs. (21) can be verified by substituting the expansion (17) in the continuum equations (7) with zero right sides. Indeed, Eqs. (2la,b) state that the linear and angulat momentum residing in a v-ch hinges-Eree vehicle mode are zero, whereas Eq. (21c) expresses a zero momentum-tike property of the articulated body $E_{j}$. These properties can be stated alternately by defining modal momenta coefficients ( $\mathrm{p}_{\mathrm{jv}}, \mathrm{h}_{\mathrm{j} v}$ ) for each articulated body and $\left(\mathrm{P}_{\mathrm{v}}, \mathrm{h}_{\mathrm{v}}^{0}\right)$ for all articulated bodies collectively:

$$
\begin{align*}
& \mathbf{P}_{\mathrm{jv}} \triangleq \int_{\mathrm{j}} \underline{U}_{\mathrm{j} v}\left(\underline{\underline{r}}_{\mathrm{j}}\right) d m \quad \underline{h}_{\mathrm{j} v} \triangleq \int_{\mathrm{j}} \underline{\underline{r}}_{\mathrm{j}}^{\mathrm{X}} \underline{U}_{\mathrm{j} v}\left(\underline{\underline{r}}_{\mathrm{j}}\right) \mathrm{dm} \tag{22}
\end{align*}
$$

where $h^{0}$ is defined relative to the reference point 0 . These may be compared with the definitions (9) and (10). The zero moment 3
properties (21) then transform to:

$$
\begin{aligned}
& \underline{c}^{x} \underline{x}_{0 v}+\underline{J} \underline{\theta}_{0 v}+\sum_{j} \underline{c}_{0 j} \underline{J}_{0 j}{ }^{\Phi} j v+\underline{h}_{v}^{0}=\underline{0}
\end{aligned}
$$

$$
\begin{align*}
& \text { ( } j=1, \ldots N ; v=1, \ldots, \infty \text { ) } \tag{23}
\end{align*}
$$

The eigenvalue problem which governs a hinges-free vehicle mode is given by

where $\omega_{v}$ is the frequency of $v$-th mode. The orthogonality conditions are obtained by performing the operation $\sum_{j} \int_{j} \underline{U}_{j \mu}^{T}\left(\underline{r}_{j}\right)(\bullet)$ dv over the eigenvalue problem (24) and recalling the properties (23). Here, dv is an elemental volume. One then obtains

$$
\begin{align*}
& \int_{V} \underline{W}_{\mu}^{T}(\underline{\varepsilon}) \underline{W}_{V}(\underline{r}) d m=\delta_{\mu \nu} \tag{25.a}
\end{align*}
$$

$$
\begin{align*}
& \text { (25.c) } \\
& \sum_{j} \int_{j} \underline{U}_{j \mu}^{T} L_{j} \underline{U}_{j v}\left(\underline{r}_{j}\right) d v=\omega_{v}^{2} \delta_{\mu v} \tag{25.d}
\end{align*}
$$

where $\delta_{\text {uy }}$ is the Kronecker delta.
Utilizing the modal expansion (17), the zero momentum modal properties (21) and (23), and the orthogonality properties (25), che continuum equations are discretized to these decoupled equations which separately govern the rigid and elastic modes of the spacecraft:

$$
\begin{align*}
& \ddot{\eta}_{v}+w_{v}^{2} \eta_{v}=\underline{x}_{0 v}^{T} f+\Phi_{0 v}^{T} g+\sum_{j} \Phi_{j v}^{T}\left(g_{0 j}+\underline{g}_{j}\right)+\gamma_{v}(t)  \tag{26.c}\\
& (v=1,2, \ldots, \infty) \tag{26,d}
\end{align*}
$$

where $Y_{v}(t)$ is the scalar input co $v-t h$ mode considering all articulated bodies collectively:

$$
\begin{equation*}
Y_{v}(t) \triangleq \sum_{j} \int_{j} \underline{U}_{j v}^{T}\left(\underline{r}_{j}\right) \underline{\underline{b}}_{j}\left(\underline{r}_{j}, t\right) d A \tag{27}
\end{equation*}
$$

wise, the eigenfunction $\underline{U}_{j \alpha}^{\mathrm{C}}\left(\underline{r}_{\mathrm{j}}\right)$ is the deformation of $E_{j}$ in the a-th mode, analogous to $\underline{U}_{j v}\left(\underline{r}_{j}\right)$ in the case of v-th hinges-free mode, except that now a force is exerted at the hinge $0_{j}$ to ascertain that $\underline{U}_{j a}^{c}\left(0_{j}\right)=\underline{0}$ and $\underline{\nabla}^{\mathrm{x}} \underline{U}_{\mathrm{ja}}^{\mathrm{c}}\left(\mathrm{O}_{\mathrm{j}}\right)=\underline{\underline{0}}$. The total motion of j -th appendage relative to the $B_{0}$-fixed frame in a-th vehicle mode is denoted $\underline{W}_{j a}^{c}\left(\underline{r}_{j}\right)$ and it equals [cf. Eq. (18)]
$\underline{w}_{j a}^{c}\left(\underline{r}_{j}\right) \triangleq \underline{x}_{0 a}^{c}-\left(\underline{b}_{j}+\underline{-}_{0} \underline{\underline{r}}_{j}\right)^{x} \Phi_{-0 \alpha}^{c}+\underline{-}_{0}{ }_{j} \underline{U}_{j a}^{c}\left(\underline{\underline{r}}_{j}\right)$
The a-th mode of the core body, $\underline{W}_{0 \alpha}^{c}\left(\underline{r}_{0}\right)$, on the other hand, will be

$$
\begin{equation*}
\underline{w}_{-0 a}^{c}\left(\underline{r}_{0}\right) \triangleq \underline{x}_{0 a}^{c}-\underline{\underline{r}}_{0}^{\mathbf{x} \Phi_{0 a}^{c}} \tag{33}
\end{equation*}
$$

Thus, like Eq. (20), the a-th mode ${\underset{a}{a}}_{c}^{c}(\underline{r})$ will be $\underline{w}_{0 a}^{c}\left(\underline{r}_{0}\right)$ or $\underline{W}_{j a}^{c}\left(\underline{r}_{j}\right)$ depending on the domain under consideration. Orthogonality of these modes with the six rigid modes $\underline{1}$ and $-r^{x}$, similar to Eqs. (21.a, 21.b), can be proved easily. To express these conditions in terms of hinges-locked modal momenta coefficients, define \{cf. Eq. (22)]:



Then, the above mentioned orthogonality is [cf. Eq. (23)]:

$$
\begin{align*}
& m{\underline{x_{0 a}}}_{c}-\underline{c}^{x} \underline{0}_{0 a}^{c}+\underline{P}_{a}^{c}=\underline{0} \\
& \underline{c}^{x} x_{0 a}^{c}+\underline{J}_{-\Phi_{0 a}^{c}}^{c}+\underline{h}_{a}^{0^{c}}=\underline{0} \tag{35}
\end{align*}
$$

The eigenvalue problem obeyed by the $a-$ th hingeslocked mode $\underline{w}_{j a}^{c}\left(\underline{r}_{j}\right)$ is
$\underline{L}_{j} \underline{U}_{j a}^{c}\left(\underline{r}_{j}\right)=\omega_{a}^{c^{2}} \underline{\underline{c}}{ }_{j 0} \underline{W}_{j a}^{c}\left(\underline{r}_{j}\right) \quad(j=1, \ldots, N)$
where $\omega_{a}^{c}$ is the associated hinges-locked frequency. "The orthogonality conditions between a-th and $\beta$-th modes are

$$
\begin{align*}
& \sum_{j} \int_{j} \underline{U}_{j a}^{c}{ }^{T} \underline{L}_{j} \underline{U}_{j B}^{c} d v=\omega_{y}^{c^{2}} \delta_{a B} \tag{37}
\end{align*}
$$

These are a bit more general than Eqs. (62) of Hughes ${ }^{4}$.

With the aid of the expansion (31), momental properties (35), and orthogonality properties (37), the original continuum equations (9) are discretized to


$$
(j=1, \ldots, N)
$$


Unlike the hinges-free set of discrete equations (28), the last two equations in (38) involve a new coupling term called "inertial modal angular momentum coefficient" $\underline{h}_{\text {ja }}^{\text {cI }}$ defined as


which is different from ${\underset{\sim}{\mathrm{h}}}_{\mathrm{ja}}^{\mathrm{c}}$ and ${\underset{\mathrm{h}}{ }}_{\mathrm{O}^{\mathrm{C}}}$ defined in (34). The disturbance input $Y_{a}^{c}(t)$ to each a-th mode equals [cf. Eq. (29)]
$r_{a}^{c}(t) \triangleq \int_{j} \int_{j} \underline{u}_{j a}^{c^{T}}\left(\underline{r}_{j}\right) \underline{f}\left(\underline{r}_{j}, t\right) d A$
To compact Eqs. (40), introduce

Then, recalling pertinent definitions from (13), (15), and (28), Eqs. (38) contract to
$\underline{M}_{V V}{\stackrel{\oplus}{q_{V R}}}+\underline{M}_{V A}^{T} \underline{\Omega}_{A} \quad=\underline{u}_{V}(t)$


The vector ${\underset{\sim}{n}}_{c}^{c}(t)$ and $\underline{y}_{c}(t)$, and the matrices $w_{c}, x_{0}^{c}, \phi_{0}^{c}$ are defined like their hinges-free
companions in (28).

## iv. modal identities for multibody

 elastic spacecraft[^1]First, consider the discrete set (16) based on appendage modes. By matrix manipulations, the following equation governing $\underline{\Omega}_{A}$ can be conscructed readily:

$$
\begin{align*}
& \left.{\underset{-X}{A}}_{T}^{\left(\underline{O}_{\infty}\right.}+\underline{\Omega}_{C}^{2} / s^{2}\right)^{-1}{\underset{\sim}{P}}_{-A}^{0} \underline{M}_{V V}^{-1} \underline{U}_{V} \\
& -{\underset{X}{X}}_{A}^{T}\left(\underline{U}_{\Phi}+\underline{\Omega}_{C}^{2} / s^{2}\right)^{-1} \underline{\underline{r}}_{A}^{a}-\underline{M}_{V A} \underline{M}_{V V}^{-1} \underline{u}_{V}+\underline{g}_{H} \tag{43}
\end{align*}
$$

where $s$ is the Laplace variable; theoretically, $\mathcal{H}_{A} i s$ an $\infty x n_{a}$ matrix, and $\mathcal{U}_{0}$ an $\infty x \infty$ symmetric matrix; $\underline{g}$ is an $n_{a} \times n_{a}$ inertia matrix, and $g_{H}$ is the total hinge corque vector:

$$
\begin{align*}
& \underline{\underline{9}} \underline{\Xi}_{\underline{J}_{A}}-\underline{M}_{V A} \underline{M}_{V V}^{-1} \underline{M}_{V A}^{T}, g_{H} \triangleq \underline{g}_{A}+\underline{g}_{0 A} \tag{44}
\end{align*}
$$

In (44), $\underline{1}_{\infty}$ is $\infty \times \infty$ identity matrix. Thus for an equation of ${ }^{\mathbf{9}} \dot{\underline{i}}_{\mathrm{A}}(\mathrm{s})$, the coefficient of the hinge corque $g_{H}(s)$ is

$$
\begin{equation*}
\left[\underline{1}-\left.\mathscr{H}_{-A}^{T}\left(\underline{Q}_{0}+{\underset{-a}{C}}_{2}^{2} s^{2}\right)^{-1} \mathscr{H}_{A} \underline{g}^{-1}\right|^{-1}\right. \tag{45}
\end{equation*}
$$

Anticipating our later needs, now we shall prove that the lim $s+\infty$ of the matrix $[\bullet]$ in (45) equals

$$
\begin{equation*}
\underline{1}-\mathcal{H P}_{A}^{\mathrm{T}} \underline{U l}^{-1} \mathcal{H}_{A} \cdot g^{-1}=0 \tag{46}
\end{equation*}
$$

Applying the matrix inversion lemma to $\underline{q l}_{\infty}$, its inverse is found to be:

On the other hand, owing to the identities ( $D, E, F$ )" of Hughes ${ }^{4}$

Also, by definition of $\underline{M}_{V V}$ in (15), and by virtue of Eq. (1) and Eq. (3)
which reduces $\underline{q}^{-1}$ to

$$
\begin{equation*}
\underline{U}_{\infty}^{-1}=\underline{w}_{\infty}+\underline{T}_{A}^{0} \underline{\underline{M}}_{0}^{-1} \underline{T}_{A}^{0^{T}} \tag{50}
\end{equation*}
$$

A comparison of $\underline{U}_{\infty}$ with $\mathcal{U}_{\infty}^{-1}$ amazes. Continuing with the proof nevertheless, call upon the basic identities $(D, E, F)$ " of Hughes ${ }^{4}$ to derive the following new identities associated with the articulation degrees of freedom:

$$
\underline{H}_{A}^{T} \underline{\underline{H}}_{A}={\underset{-}{A}} \quad \underline{H}_{A}^{T}{\underset{\underline{P}}{A}}_{0}^{0}=\underline{C}_{A}^{0} \quad \ddot{\underline{H}}_{A}^{T}{\underset{H}{A}}_{0}^{0}=\underline{J}_{O A}
$$

$$
\begin{align*}
& \mathcal{H}_{A}^{T} \mathcal{H}_{A}=\underline{Q}-\underline{M}_{V A} \underline{M}_{V V}^{-1} \mathcal{M}_{0} \underline{M}_{V V}^{-1} \underline{M}_{V A}^{T}, \\
& \mathcal{H}_{A}^{T} \underline{Q}_{A}^{0}=\underline{M}_{V A} \underline{M}_{V V}^{-1} \underline{M}_{0} \tag{I}
\end{align*}
$$

(New identities derived in this paper will be labeled with Roman numerals as they are cited.) These identities and Eq. (50), in turn, lead to the identicy

$$
\begin{equation*}
\mathcal{H}_{A}^{\mathbf{T}} \underline{u}_{\infty}^{-1}{\underset{H}{A}}=\underline{g} \tag{II}
\end{equation*}
$$

which proves Eq. (46)
An equation analogous to Eq. (43) is obtained from the hinges-free discrete set (29). For that, recall the second expansion in (11). Then it can be shown that

$-\underline{M}_{V A} \underline{\underline{M}}_{V V}{ }^{-1} \underline{u}_{V}(s)$
$+\left[\underline{1}+\underline{9} \underline{\underline{p}}_{A}^{T}\left(\underline{1}+\underline{\omega}^{2} / s^{2}\right)^{-1} \underline{\underline{p}}_{A} \mid \underline{g}_{H}(s)\right.$
The coefficient of the hinge torque $s_{H}(s)$ in Eq. (51) equals the term (45). They both reduce to $\frac{1}{}$ for the $\lim s+0$, and fior the lim $s+\infty$ they yiel̃, in view of Eq. (46), the identity

$$
\begin{equation*}
\left(\underline{1}+\underline{9} \Phi_{A}^{T} \Phi_{A}\right)^{-1}=\underline{0} \tag{III}
\end{equation*}
$$

which proves a fortiori that, since the inertia matrix $\underline{\underline{g}}$ is positive definite and $\Phi_{A}^{T} \Phi_{A}$ nonnegative definite, the modal coefficients $\Phi_{j v}(\mathrm{j}=1, \ldots, \ldots$; $v=1, \ldots, \infty$ ) [Eq. (17)] constitute a nonconverging series.

The hinges-locked discrete set (42) furnishes this equation for ${\underset{A}{A}}^{A}$ :

$$
\left[\underline{1}-\underline{h}_{A}^{c I^{T}}\left(\underline{1}+\underline{\underline{w}}_{c}^{2} / s^{2}\right)^{-1} \underline{\underline{h}}_{A}^{c I} \underline{g}^{-1}\right] \underline{\underline{g}} \underline{\underline{g}}_{A}=\underline{g}_{H}-\underline{M}_{V A} \underline{M}_{V V}^{-1} \underline{\underline{u}}_{V}
$$

$$
\begin{equation*}
-\underline{h}_{A}^{c I^{T}}\left(\underline{1}+\underline{w}_{c}^{2} / s^{2}\right)^{-1}\left(\underline{x}_{0}^{c} \underline{f}+\underline{\Phi}_{0}^{c} \underline{g}+\underline{x}_{c}\right) \tag{52}
\end{equation*}
$$

The equality of the coefficient matrices of $g_{4}(s)$ in Eq. (51) and Eq. (52) delivers this identity in the s-domain:

$$
\begin{gather*}
{\left[\underline{1}-\underline{h}_{A}^{c I}\left(\underline{1}+\underline{w}_{c}^{2} / s^{2}\right)^{-1} \underline{h}_{A}^{c I^{-1}} \underline{g}^{-1}=\underline{1}+\right.} \\
\underline{g} \Phi_{A}^{T}\left(\underline{1}+\underline{w}^{2} / s^{2}\right)^{-1} \underline{s}_{A} \tag{IV}
\end{gather*}
$$

For the $\lim s+0$, the left side of (IV) degenerates to 1 as does its right side. On the other hand, taking its limit $s \rightarrow \infty$ and recognizing the identity (III) produce the identity

$$
\begin{equation*}
\left[\underline{1}-\underline{h}_{A}^{c I^{T}} \underline{n}_{A}^{c I_{-}-1}\right]=\underline{0} \tag{v}
\end{equation*}
$$

The identity (IV) can be rearranged such that it reveals poles and zeros of the dynamics. For thaf, recognize that when $j= \pm j \omega_{v}\left(v=1, \ldots, \infty i j^{2}=-1\right) \quad$ the right side of (IV), which is also the coeificient of $g_{H}(s)$ in (51), is unbounded, so $\pm j w$, are the
poles of the spacecraft. Consejuently, for unboundedness to occur, the left side of (IV), expressed in terms of individual hinges-locked modes, yields the identity
$\operatorname{det}\left|\underline{1}-\sum_{a}\left(1-\omega_{a}^{c} / \omega_{V}^{2}\right)^{-1}{\underset{\sim}{h}}_{c I}^{h_{a}^{c I}} \underline{g}^{-1}\right|=0$
where $h_{\text {h }}^{\text {cI }}{ }^{T}$ is the $\alpha-$ th row of the matrix $\underline{h}_{A}^{c I}$. similarly, when $s= \pm j_{d}^{c}(a=1, \ldots)$, the matrix within (•l on the left side of (IV), which is the coefficient of $\underline{\underline{0}}{\underset{A}{i n}}^{i n}(52)$, is unbounded, which implies that $\pm j \omega_{a}^{c}$ are the zeros of the dynamics. Therefore, to realize unboundedness, the right side of (IV) bears forth
$\operatorname{det}\left[\underline{1}+\sum_{\mu}\left(1-\omega_{\mu}^{2} / \omega_{a}^{c^{2}}\right)^{-1} \underline{q} \underline{Q}_{\mu} \underline{Q}_{\mu}^{\mathrm{T}}\right]=0$
where $\stackrel{D}{\mu}^{T}$ is the $u$ th row of the matrix $\phi_{\text {A }}$ [Eq. (28)]. Knowing the poles and zeros, the identity (IV), keeping in mind its lim, has this alternate form [cf. (Y) of Hughes ${ }^{4}$ ]: s+m
$\left[\underline{1}-\underline{h}_{A}^{c I^{T}}{\underset{h}{A}}_{c I} g^{-1}\right]^{-1} \underset{a=1}{n_{f}}\left(s^{2}+w_{a}^{c}\right) /\left[\begin{array}{ll}n_{f} \\ \mu=1 & \left(s^{2}+\omega_{\mu}^{2}\right) \\ \mu=\end{array}\right]=$

(VIII)
where $n_{f}=$ total number of retained modes. Because of the identity ( $V$ ), however, this form seems to be less useful than the form (IV). Following Garg ${ }^{9}$, one can examine how far the identities (IV) or (VIII) are satisfied in the s-domain. The identities (VI) and (VII) are useful in several ways; for instance, known hinges-locked parameters can be used to determine hinges-free modal parameters, or vice versa, after Hughes and Garg ${ }^{10}$. Incidentally, the identities (VI) and (VII) are analogous to the identities ( $M$ ) and (Q) of Hughes ${ }^{4}$. As in Reference 4, ${ }^{\theta}$ under conditions of symmetry, these identities reduce to those concerned with individual articulation degrees of freedom. Owing co symmetry, since different sets of modes will contribute to different articulation degrees of freedom, the set $a(a=1, \ldots, \infty)$ may form $n_{3}$ subsets $a_{j}\left(j=1, \ldots, n_{2}\right)$ and each $a_{j}$ will span the range $1, \ldots, \infty$; the set $\mu\left(\mu=1, \ldots, j_{\infty}\right)$ fragmentates likewise. The identities (VI) and (VII) then simplify to

$\sum_{u_{2}=1}^{\infty}\left(\omega_{u_{2}}^{2} / \omega_{a_{\ell}}^{c^{2}}-1\right)^{-1}\left[\left.\underline{ \pm_{\mu} D_{\mu}^{T}}\right|_{\ell, k}=\delta_{\ell k}\left(Q, k=1, \ldots, n_{a}\right)\right.$
V. ILLUSTRATION OF MODAL IDENTITIES AND DISCUSSION

The identities will now be illustrated for a four-body deformable spacecraft shown in Fig. 2. It has two flexible solar arrays, $E_{1}$ and $E_{2}$, each having one articulation degree of freedom about $y_{1}$ - and $y_{2}$-axis, respectively, relative to the core body $B_{0}$, and a sensor having two rotational degrees of freedom about $x_{3}-$ and $y_{3}$-axis. These four articulation angles are denoted
${ }^{\theta}{ }_{2 y},{ }^{\theta} 2 y,{ }^{\theta} 3 x$, and ${ }^{\theta} 3 y$, and the spacecraft thus has ten rigid modes. Hinges-free and hinges-locked vehicle modal data for the spacecraft were obtained by using NASTRAN. From a detailed finite element model having 19,434 degrees of freedom and 3,239 nodes, 63 hinges-free and 67 hinges-locked elastic modes below 25 Hz were computed. Since the vehicle is essentially symmetric (the sensor causes a slight asymmetry), both symmetric and antisymmetric vehicle modes arise in transuerse bending and in-plane bending of the arrays, and the vehicle modes are categorized accordingly in Table 1 and Table 2. Fig. 3a confirms the prediction from the identity (III) that the hinges-free modal coefficients, in this case $\phi_{1 \mu y}(\mu=1, \ldots, 63)$ for the $y_{1}$-solar array, form a nonconverging series. In Fig. 3a, the largest modal coefficients $\phi_{1 \mu y}$ for $\mu=8,11,18,28, \ldots$ cor ${ }^{-}$ respond to those vehicle modes which predominately entail torsion of the $y_{1}$-array about $y_{1}$-axis (Table 1). In contrast, those contributing to $\theta_{3 y}$, namely, $\boldsymbol{j}_{\mu y}(\mu=1, \ldots, 63)$, form essentially a convergent series because the sensor is rigid, and symmetric transverse bending of the arrays (Table 1) or local high-frequency deformation of $\mathrm{B}_{0}$ at the sensor base produce $\phi_{3 \mu y}(\mu=1, \ldots, 63)$. The hinges-locked coupling coefficients hif for $y_{1}$-array and $h_{3 a y}^{c l}$ for $\theta_{3 y}$ rotation for the modes $a=1, \ldots, 67$ are displayed in Fig. 4. Unlike ${ }_{l_{\mu}}, h_{\text {lay }}^{c I}$ forms a converging series.

The identities (III) and (V) are the simplest, for they involve only modal coefficients, no frequencies. The identity (III) is illustrated in Fig. 5. The error indexes $e_{k k}^{H F}(k=1,3)$ ( $H F$ means hinges-free) are the corresponding diagonal elements of the ( $4 \times 4$ ) matrix $\left[\underline{1}+\left.\underline{q} \oplus_{A}^{\top} \oplus_{A}\right|^{-1}\right.$. In contrast to their zero ideal value, the asympoces


Figure 2. A Four-Body Deformable Spacecraft

Table 1. Hinges-Free Modes

| Characteristics <br> of the Mode |  |  | Affected <br> Rotational <br> Degrees of <br> Freedom |
| :--- | :--- | :--- | :--- |
| Transverse bending of <br> arrays and A-frames | Symmetric | $1,5,9,13,15,21, \ldots$ | $\theta_{o y r} \theta_{3 y}$ |
|  | Antisymmetric | $2,6,10,14,16,22, \ldots$ | $\theta_{\mathrm{oz}}$ |
| Torsion of arrays and <br> A-frames | Array 1 | $8,11,18,28, \ldots$ | $\theta_{1 y}$ |
|  | Array 2 | $7,12,19,29, \ldots$ | $\theta_{2 y}$ |
| In-plane bending of <br> the A-frames and <br> solar arrays | Symmetric | $3, \ldots$ | None |

Table 2. Hinges-Locked Modes

| Characteristics of the Mode |  | Mode No. | Affected Rotational Degrees of Freedom ${ }^{\dagger}$ |
| :---: | :---: | :---: | :---: |
| Transverse bending of arrays and A-Frames | Symmetric | 1, 2, 6, 8, 12, 16 | $\theta_{\text {oy }}$ |
|  |  | 1,2,... | $\theta_{3 \gamma}$ |
|  | Antisymmetric | 3, 9, 13, 17, 20, 27, .. | $\theta_{\text {Oz }}$ |
| Torsion | Array 1 | $5,6,10,11,14,15,22,23, \ldots$ | $\theta_{17}$ |
| In-plane bending of arrays and A-frames | Antisymmetric | 7,18,21,.. | $\theta_{\text {ox }}$ |
|  |  | 18, $\ldots$ | $\theta_{3 y}$ |
| Vibrations of the spacecraft |  | $\begin{aligned} & 28,35,36,37,41,42,43,44 \\ & 47,48,49,53, \ldots \end{aligned}$ | $\theta_{\text {ox }}$ |
|  |  | $\begin{aligned} & 28,36,37,41,42,43,44,45, \\ & 47,48,49,53, \ldots \end{aligned}$ | $\theta_{\text {oy }}$ |
|  |  | $\begin{aligned} & 28,35,36,37,38,41,42,43, \\ & 44,47,48,53, \ldots \end{aligned}$ | $\theta_{\text {Oz }}$ |
| †Information about the interaction with $\theta_{2 y}$ and $\theta_{3 x}$ not available |  |  |  |



Figure 3. Hingea-Free Modal Coefficients of Articulation Motion of $y_{1}$ - Solar Array, $\left.\right|_{1_{\mu y}} \mid$, and of Sensor $B_{3},\left.\right|_{3_{\mu y}} \mid$, about $y_{3}$-Axis


Figure 5. Hinges-Free Identify III: Diminishing of Error Index with Hinges-Free Modes $\mu$
of these indexes are 0.0406 for $k=1$ ( ${ }_{1}$ rotation of the $y_{1}$-array) and 0.689 for $k=3 \quad\left(\theta_{3 x}\right.$, the $x$-rotation of the sensor). The error indexes diminish discretely at appropriate modes as predicted by Table 1. For instance, for $\theta_{1 y}$, the error index $e_{11}^{\mathrm{HF}}$ diminishes at the torsional modes $\mu=8,11,18,28, \ldots$ The index for the $y_{2}$-array motion ( $k=2$ ) is the same as that for $k=1$, except that it decreases instead at the adjacent torsional modes $\mu=7,12,19,29, \ldots$ (see Table 1). Surprisingly, the asymptote of the error index for ${ }^{6}{ }_{3 y}(k=4)$, not included in Fig. 5, hovers at 0.9976 instead of decreasing to the ideal value zero. Fig. 6 illustrates the hinges-locked identity $(v)$, rearranged as $\underline{h}_{A}^{c I^{T}}{\underset{A}{h}}^{c I} \underline{9}^{-1}=1$. For discussing this identity and the ones following, define a "completeness index C" which approaches unity for an error-free model [Reference 1]. [The completeness index for Fig. 5 is


Figure 6. Hinges-Locked Identify V: Growth of Completeness Index with Hinges-Locked Modes a
$\left(1-e_{k k}^{H F}\right) . \mid$ The growth of the diagonal elements (1,1) and (4,4) of the matrix $h_{A}^{c I^{T}} \underline{h}^{c I_{9}}{ }^{-1}$ are depicted in Fig. 6. Surprisingly, Alongide the error index $e_{11}^{H F}$ in Fig. Sa, $C_{11}^{H L}(H L$ means hingeslocked) approaches unity in just two hinges-locked torsional modes, 5 and 6 (Table 2), and its asymptotic value is 1.029. Furthermore, by contrast with the hinges-free completeness index $\mathrm{C}_{44}^{\mathrm{HF}}$ equal to 0.0024 (that is, the above mentioned error index $e_{44}^{\mathrm{HF}}$ of 0.9976 ), the hinges-locked completeness index $C_{44}^{H L}=0.9421$ in Fig. 6 b is remarkable; in fact, the first hinges-Iocked mode, 3 symmetric cransverse bending mode of the arrays (Table 2), contributes a mighty share, 0.9412 , to $\mathrm{C}_{44}^{\mathrm{HL}}$.

The identities which involve frequencies as well are now illustrated. First, consider the identity (VII) which is summed over all hingesfree modes ( $\mu=1, \ldots, 63$ ) for a specific $\omega_{a}^{c}$. When $\omega_{\mu}$ and $\omega_{a}^{c}$ are the same co several decimal places, it is difficult to verify this identity in this form. On the other hand, the identity (VIII) indicates that when $\omega_{\mu}$ and $\omega_{a}^{c}$ are truly the same, the corresponding poles and zeros cancel each other without affecting the articulation dynamics. A physical explanation of this is that when hinges-free and hinges-locked frequencies are truly equal, that particular mode does not contribute to the articulation motion, so such a mode may be deleted from the study. In numerical work, however, it is difficult to establish true equality between two real numbers. Besides, as will be seen shortly, for the example in hand, sometimes even though $\omega_{u}$ and $\omega_{a}^{c}$ are the same up to three or four decimal places, the minuscule difference between the two is still important for the verification of an identity. Consequently, the following results are obtained without truncating either modal set. Returning to the identity VII, one finds that when $a>9$, hinges-locked frequencies $w_{a}^{c}$ are so close to a corresponding hinges-free frequency $\omega_{\mu}$ that the determinant, instead of being zero, becomes an arbitrarily
large number. Among $a=1, \ldots, 9$, the identity (VII) is best satisfied with $a=7$ and next best with $a=2$, for wnich the determinants are, respectively, -0.00415 and 0.04732 (Table 3). The circumstances which produce these results are revealed by the identity ( X ). For a given $\omega_{\alpha}^{c}$, when all available hinges-free modes are added to calculate the ( $\ell, k$ ) element of the left side of ( $X$ ), it is denoted $C_{8 k, a s y}^{\mathrm{HF}}$ where HF "asy" means asymptotic value. Fig. 7 shows $C_{\ell l, \text { asy }}^{\mathrm{HF}}$ for $\ell=1,3$, and 4. The ideal value of this index is unity; however, when $\omega_{u_{z}} \neq \omega_{1_{2}}^{c}$ for some $u$ and $a$, this index assumes an arbicrarily large value, and for ploting purpases, such large numbers are replaced by 2 without altering their signs. In the left side of fig. 7.a, in the useful range 0 to 1 , the maximum value of $C_{11, \text { asy }}^{\mathrm{HF}}$ concerning the $y_{1}$-array rotation, $\theta_{1}$, is 0.451 for the hingeslocked mode $a=2--$ a symmetric transverse bending mode of the arrays (Table 2). On the orher hand, the first torsional hinges-locked mode having significant coupling with the rotation $\theta_{1 y}$ is $\alpha=5$ (Table 2 and Fig. 4), but $C_{11, a s y}^{\mathrm{HF}}$ corresponding to $\alpha=5$ is 0.16 , less than 0.451 for $~ a=2$. Although the index $C_{11, \text { asy }}^{H F}$ for a=5 should be, intuitively, greater than that for $a=2$, thig does not happen because the hinges-free frequency $\omega_{1}(0.25724 \mathrm{~Hz})$ is close to $\omega_{2}^{c}(0.25719 \mathrm{~Hz})$. To determine the contribution of the mode $\mu=1$, the growth of $C_{11}^{\mathrm{HF}}$ with successive addition of $\mu$ to the asymptotic value 0.451 for $a=2$ is shown in the right side of Fig. 7.a. $C_{11}^{H F}$ is found to escalate discretely at $\mu=1,8,11,18,28,29,35,41,42, \ldots$, which, except for $\mu=1$, involve torsion of the array l (Table 1). The contribution from the hinges-free mode $\mu=1$, a symmetric transverse bending mode of the arrays (Table 1) like $\alpha=2$ hinges-locked mode, is however, extraordinarily large: 937. Nevertheless, the bending mode $u=1$ is not pertinent to the articulation motion ${ }_{1 y}$, so $C_{11}^{\mathrm{HF}}=0.451$ for $a=2$ cannot be accepted, and, instead, $C_{11}^{\mathrm{HF}}=0.16$ for $a=5$, a torsional mode is accepted. Next consider the sensor motion ${ }^{8} 3 x^{\text {. }}$. The corresponding index, $C_{33, \text { asy }}^{\mathrm{HF}}$, shown on the left side of Fig. 7.b, is 1.0059 for $a=7$ (compare with $C_{11, a s y}^{\mathrm{HF}}$, and recall from Table 3 the value 0.00415 of the identity VII for $a=7$ ). The growth of $C_{33}^{\mathrm{HF}}$ versus $\mu$ for $a=7$ is displayed on the right side of Fig. 7.b, where it is observed to become unity at once when $\mu=4$. To understand this, note that both $\mu=4$ and $\alpha=7$ modes involve antisymmetric in-plane bending of the arrays--a motion which induces ${ }^{\theta} 3 x$ (see Table 1 and Table 2), and that $\omega_{4}=0.59541$ and $\omega_{7}^{c}=0.59538 \mathrm{~Hz}$. When the rotation $\theta_{3 x}$ of the rigid sensor is locked, the moment of inertia which must be turned by the antisymmetric in-plane bending, is increased, and that lowers the frequency commensurately. The ratio of the moment of inertia of the sensor and of the core body, both about $x_{0}$-axis, is 0.0717 . The decrement of $3.0 \mathrm{E}-5$ $\mathrm{H}_{2}$ noted above in the frequency $\omega_{4}$ is mathematically so precise that $C_{33}^{\mathrm{HF}}$ becomes unity at once when $a=7$. Moreover, although $\omega_{4}$ and $\omega_{7}^{c}$ are the
same up to three decimal places, the two modes cannot be truncated from the study of the verification of the identities VII and $X$. Next, consider ${ }^{\theta}$, rotation of the sensor--the rotation coupled with the transverse symmetric bending of the arrays (Table 1 and Table 2). The associated index, $\mathrm{C}_{44, \mathrm{AF}}^{\mathrm{HF}}$, versus a is shown in Fig. 7c. In the range 0 to 1 , the most it becomes is a startling low value: 0.07836 for $a=2$; for this $a$, the growth of $\mathrm{C}_{44}^{\mathrm{HF}}$ with $\mu$ indicates that 99.997 contribution arises from the first symmetric transverse bending mode $\mu=1$.

The verification of the identity (IX) for $\ell=k=1$ and 4 is considered in Fig. 3. Since this identity relaces to hinges-locked modal parameters, its left side is denoted $\mathrm{C}_{\mathrm{ek}}^{\mathrm{HL}}$. Earlier, the identity ( $V$ ) and Fig. 4 established that the hinges-locked coupling coefficients form a converging series. Therefore, the determinant identity (VI) and $\mathrm{C}_{\text {lk,asy }}^{\mathrm{HL}}$ in Fig. 8 do not become arbitrarily large numbers once $\mu \geq 28$. Indeed, only for $\mu=3,4,21,26,27$, is the index $\mathrm{C}_{\text {lk,asy }}^{\mathrm{HL}}$ unbounded, by concrast with the hingesfree index $C_{11, \text { asy }}^{\mathrm{HF}}$ in Fig. 7a which is unbounded for all $a \geq 9$. The index $C_{l l}^{\mathrm{HL}}$ depends on the selected hinges-free frequency ${\underset{2}{\mu}}_{\mu}$; for $a^{\prime}$ s having $\quad \omega_{a}^{c}>\omega_{\mu}$, the term $\left(1-\omega_{a}^{c} / \omega_{\mu}^{2}\right)$ becomes negative and these particular hinges-locked modes diminish the sum. Focusing first on $C_{11, a s y}^{\mathrm{HL}}$, surprisingly, it stabilizes early on to 1.05 when $\mu=7$ or $8-$ the first two hinges-free torsional modes. The ascent of $\mathrm{C}_{\mathrm{il}}^{\mathrm{HL}}$ to 1.05 for $\mu=8$ with hinges-locked modes $a(F i g .8 a$ ) indicates significant contributions from $a=5,6,10$, and 11--all torsional modes (Table 2); the contribution from higher corsional modes attenuates rapidly because of the fast convergence of $\mathrm{h}_{\text {lay }}^{\mathrm{cI}}$. As for the rotation $\theta_{3 y}$, the maximum value of $C_{44}^{\mathrm{HL}}$, asy, displayed in Fig. 8 b , in the range 0 to 1 is 0.959 when $\mu=5-$-the second hingesEree symmetric transverse bending mode of the arrays (Table 1). The growth pactern of $\mathrm{C}_{44}^{\mathrm{HL}}$ versus a for $\mu=5$, also shown in Fig. 8b, states that virtually the entire contribution arises from the first hinges-locked mode ( $\alpha=1$ ) involving symmetric transverse bending of the arrays.

## vi. SUMMINC UP

To draw conclusions about the relative merits of hinges-free and hinges-locked vehicle modes, Table 4 sumarizes the completeness indexes for the identities (III), (V), (IX), and (X). Evidently, the hinges-locked indexes are far closer to unity than the hinges-free indexes. The superiority of the hinges-locked vehicle modes to the hinges-free modes is established most persuasively by comparing the indexes for the articulation motion ${ }^{8}$ 3y of the sensor: $\mathrm{C}_{44, \mathrm{asy}}^{\mathrm{HF}}$ are 0.0024 and $0.0784-\mathrm{far}$ remote from unity, whereas $C_{44, a s y}^{\mathrm{HL}}$ are 0.9421 and $0.9593-$ almost unity. It must be understood, nevertheless, that the identities ( X ) and (IX) (or VII and VI)

Table 3. Identity VII: Variation of the Hinges-Free Determinant With Hinges-Locked Modes; Ideal Value $=0$

| $\alpha$ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $V I I$ | 0.97105 | 0.04732 | 0.27012 | 0.79457 | 0.68305 | 0.68009 | -0.00415 | 0.19551 | -0.32426 |

Minimum value of the determinant among those for $\alpha=10, \ldots, 63$, is 74.4 , and the maximum value is $\infty$ when $\omega_{\mu}=\omega \mathcal{\&}$ up to several decimal places


Figure 7. Identity $X$ : Asymptotic Values of the Hinges-Free (HF) Completeness Indexes Versus Hinges-Locked (HL) Mode a, and Growth of this Index Versus Hinges-Free Modes




Figure 8. Identity IX: Asymptotic Values of the Hinges-Locked Completeness Indexes Versus Hinges-Free Modes, and Growth of this Index with Hinges-Locked Modes

Table 4. A Summary of the Completeness Indexes for Hinges-Free and Hinges-Locked Vehicle Modes; Ideal Value $=1$

| HingesFree (HF) Indexes | $\begin{gathered} \text { Identity } \\ \text { III } \\ 1-\mathrm{HF}_{\mathrm{Rf}} \text {, asy } \end{gathered}$ | $\underset{X}{\text { Identity }}$ | HingesLocked (HL) Indexes | $\begin{array}{\|c} \text { Identity } \\ \mathrm{V} \end{array}$ | $\begin{aligned} & \text { Identity } \\ & \text { IX } \end{aligned}$ | Articulation Motion | Associated Mode of Deformation |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{CH}_{11, \mathrm{as} y}$ | 0.9594 | 0.160 | $\mathrm{CH}_{11, \mathrm{asy}}$ | 1.029 | 1.05 | $\theta_{19}$ | Torsion |
| $\mathrm{CHF}_{33, \mathrm{asy}}$ | 0.311 | 1.0059 | $\mathrm{CHL}_{33, \mathrm{asy}}$ | TBD $\dagger$ | TBD $\dagger$ | $\theta_{3 x}$ | Antisymmetric in-plane bending |
| $\mathrm{CH}_{44, \mathrm{asy}}$ | 0.0024 | 0.0784 | $\mathrm{CHL}_{44, \mathrm{asy}}$ | 0.9421 | 0.9593 | $\theta_{3 y}$ | Symmetric transverse bending |

represent two different situations: in the former, the hinges-free modes are employed to yield a bounded response at a hinges-locked frequency; and in the latrer, the hinges-locked modes are used to elicit an unbounded responge at a hingeg-free frequency. Therefore, a comparison of the indexes from these identities is slighty inappropriate perhaps; yet the conclusion from Table 2 seems inevitable that the hinges-locked vehicle modes yield a much more accurate model for simulation than the hinges-free vehicle modes do. This is caused by the nonconvergence of the hinges-free modal coefficients in contrast with the rapid convergence of the hinges-locked coupling coefficients--the attributes corroborated by the identities. Besides contrasting one family of modes with the other, the identities are clearly useful in sifting thraugh scores of finite-element generated modes to select a $k$ ew pertinent modes for an articulation degree of freedom in consideration. An important extension of the preceding work is to devise identities which involve modal coefficientg and frequencies of only one family of modes, hinges-free or hinges-locked, not both. Hughes has formulated such identities for an elastic body with no articulated members.

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[^1]:    Our principal concern is to compare hinges-free and hinges-locked modes for their accuracy in representing articulation motion. To accomplish this aim, an equation will be obtained from each of the above three sets of discrete equations which will be solely in terms of the articulation motion $\dot{\dot{g}}-A$ and stimuli. These three equations will then be compared to yield identities.

