

N91-11962

DEEP SPACE PROPAGATION EXPERIMENTS AT  $K_a$ -BAND

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ABSTRACT

This presentation discusses propagation experiments as essential components of the general plan to develop an operational deep space telecommunications and navigation capability at  $K_a$ -band (32-35 GHz) by the end of the 20th century. Significant benefits of  $K_a$ -band over the current deep space standard X-band (8.4 GHz) are an improvement of 4 to 10 dB in telemetry capacity and a similar increase in radio navigation accuracy. Propagation experiments are planned on the Mars Observer Mission in 1992 in preparation for the Cassini Mission to Saturn in 1996, which will use  $K_a$ -band in the search for gravity waves as well as to enhance telemetry and navigation at Saturn in 2002. Subsequent uses of  $K_a$ -band are planned for the Solar Probe Mission and the Mars Program.

# **DEEP SPACE PROPAGATION EXPERIMENTS AT K<sub>a</sub>-BAND**



**Presentation To  
Fourteenth NASA Propagation Experimenters Meeting  
Austin, Texas**

**by**

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**May 11, 1990**



# CONTENTS

<b>K<sub>a</sub>-BAND RATIONALE .....</b>	<b>2</b>
<b>DEEP SPACE DOWNLINK PERFORMANCE EVOLUTION .....</b>	<b>3</b>
<b>MISSIONS WITH K<sub>a</sub>-BAND LINKS .....</b>	<b>4</b>
<b>OVERVIEW K<sub>a</sub>-BAND SCHEDULE .....</b>	<b>5</b>
<b>MARS OBSERVER K<sub>a</sub>-BAND LINK EXPERIMENT .....</b>	<b>6</b>
<b>MARS OBSERVER MISSION SCENARIO .....</b>	<b>7</b>
<b>MARS OBSERVER SPACECRAFT CONFIGURATIONS .....</b>	<b>8</b>
<b>MARS OBSERVER K<sub>a</sub>-BAND BEACON PERFORMANCE .....</b>	<b>9</b>
<b>DOWNLINK BUDGETS FOR MARS OBSERVER .....</b>	<b>10</b>
<b>CASSINI TRAJECTORY .....</b>	<b>11</b>
<b>CASSINI SPACECRAFT .....</b>	<b>12</b>
<b>K<sub>a</sub>-BAND ADVANTAGES FOR CASSINI MISSION .....</b>	<b>13</b>
<b>CASSINI K<sub>a</sub>-BAND DOWNLINK SYSTEM ASSESSMENT .....</b>	<b>14</b>
<b>SUMMARY .....</b>	<b>15</b>



# **K<sub>a</sub>-BAND RATIONALE**

- **K<sub>a</sub>-BAND OFFERS SIGNIFICANT BENEFITS**
  - **TELEMETRY**
    - **EXPECTED x4 TO x10 IMPROVEMENT IN CHANNEL CAPACITY OVER X-BAND**
  - **NAVIGATION AND RADIO SCIENCE**
    - **REDUCED PLASMA SENSITIVITY ENABLES MORE PRECISE NAVIGATION AND HIGHER PRECISION GRAVITATIONAL EXPERIMENTS THAN X-BAND**
    - **500 MHz BANDWIDTH INCREASES RADIO METRIC ACCURACY AND FACILITATES ACCURATE RELATIVITY MEASUREMENTS**
  - **REDUCED FLIGHT RADIO SYSTEM COSTS**
    - **PERMITS REDUCTION IN ANTENNA SIZE FOR EQUIVALENT PERFORMANCE AT X-BAND**
    - **PERMITS REDUCTION IN POWER REQUIREMENTS FOR EQUIVALENT PERFORMANCE AT X-BAND**
  - **REDUCED DSN TRACKING TIME vs EQUIVALENT DATA VOLUME AT X-BAND**



# DEEP SPACE DOWNLINK PERFORMANCE EVOLUTION AT X-BAND AND K<sub>a</sub>-BAND 30 DEG ELEVATION, 90% WEATHER, GOLDSTONE, CA

COMPONENT	X-BAND (8.4 GHz)		K <sub>a</sub> -BAND (32 GHz)	
	1990	1995	1995	2000+
<b>SPACECRAFT</b>				
TRANSMITTER dBm	40.0	40.0	37.0	40.0
3.66 M ANT. GAIN dBi	48.3	48.3	59.3	59.3
POINTING LOSS dB	-1.0	-0.1	-1.0	-0.2
SPACE LOSS dB	-294.6	-294.6	-306.2	-306.2
<b>GROUND</b>				
ATMO. ATTEN. dB	-0.1	-0.1	-0.4	-0.4
70M ANT. GAIN dBi	74.2	74.4	84.0	85.8
POINTING LOSS dB	-0.3	-0.1	-0.3	-0.1
NOISE SPECTRUM dBm/Hz	-184.6	-185.9	-182.6	-182.6
SNR dB-Hz	51.1	* 53.7	55.0	** 60.8

\* X-BAND IMPROVES BY 2.6 dB FROM NOISE REDUCTION (1.3 dB) AND 1.3 dB ANTENNA AND POINTING IMPROVEMENTS REQUIRED FOR K<sub>a</sub>-BAND

\*\* K<sub>a</sub>-BAND IMPROVES BY 5.8 dB FROM POWER INCREASE (3 dB) AND FURTHER TOLERANCE TIGHTENING ON ANTENNAS AND POINTING SYSTEMS



# MISSIONS WITH $K_a$ -BAND LINKS

## MARS OBSERVER – SEPT 16, 1992

- IMPLEMENTING 10 mWatt BEACON TO EVALUATE PERFORMANCE OF LINK RELATIVE TO X-BAND BASELINE

## CASSINI (SATURN) – APRIL 8, 1996

- DESIGNING A 5 Watt TELEMETRY LINK AS A MISSION ENHANCEMENT TO AUGMENT X-BAND BASELINE

## SOLAR PROBE – ABOUT THE YEAR 2000

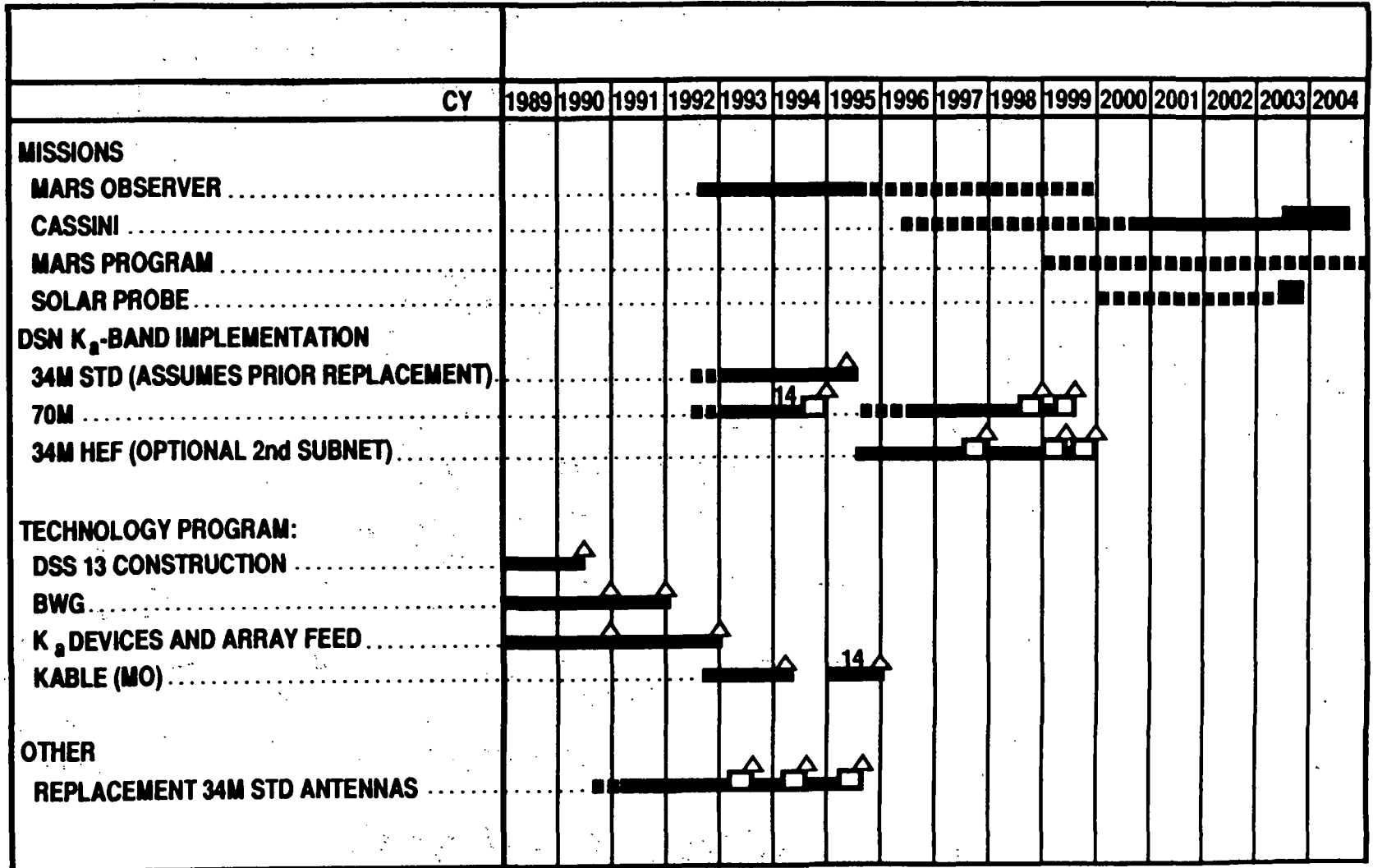
- PLANNING ON  $K_a$ -BAND AS THE BASELINE

## MARS PROGRAM

- SIGNIFICANT USE OF  $K_a$ -BAND

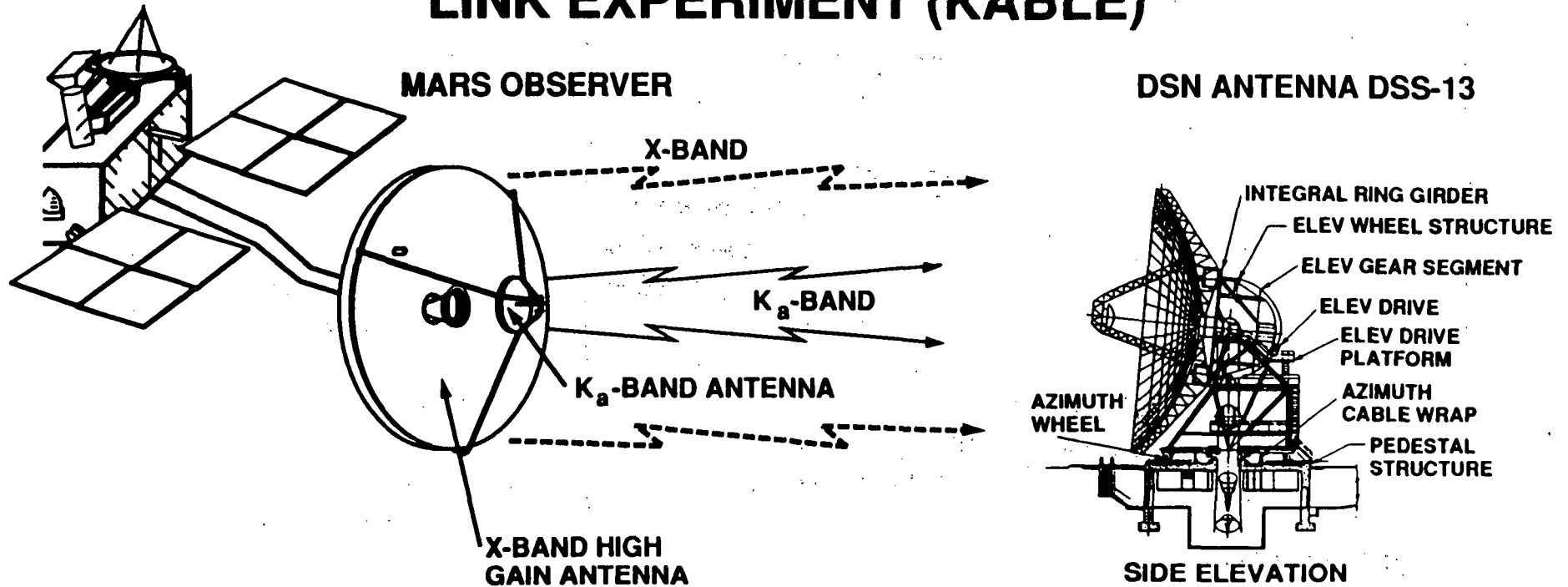


# OVERVIEW K<sub>a</sub>-BAND SCHEDULE



T6

# MARS OBSERVER $K_a$ -BAND LINK EXPERIMENT (KABLE)



92

## OBJECTIVE

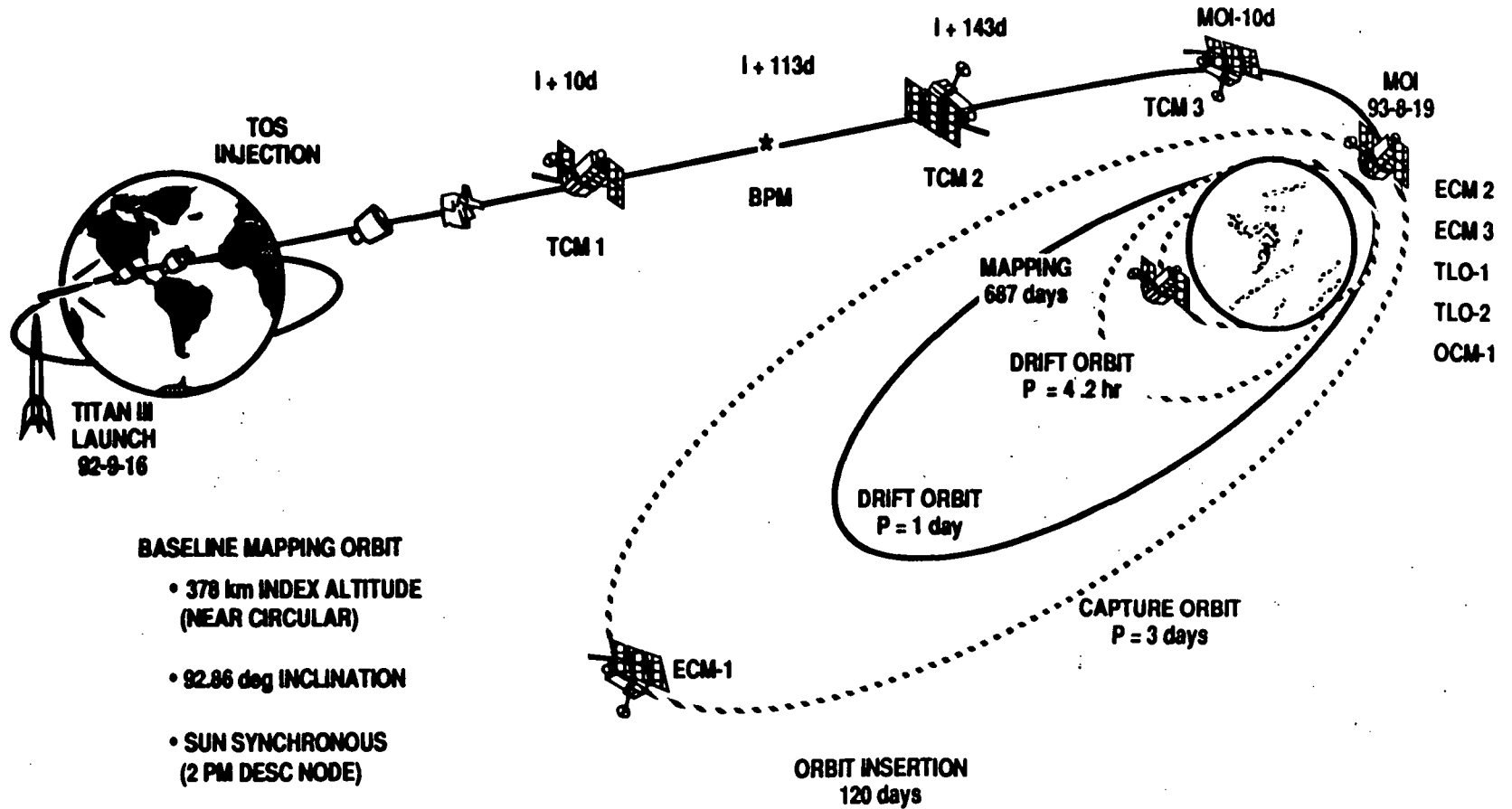
- EVALUATE  $K_a$ -BAND LINK PERFORMANCE RELATIVE TO X-BAND OVER THE SAME PATH AND OVER THE DURATION OF THE MISSION

## ENHANCEMENT OPPORTUNITY

- DEMONSTRATE ABILITY TO RECEIVE TELEMETRY AT  $K_a$ -BAND (EARLY OUTER CRUISE PHASE OF MISSION)
- EVALUATE DOPPLER AND INTERFEROMETRIC NAVIGATION USING  $K_a$ -BAND AND X-BAND



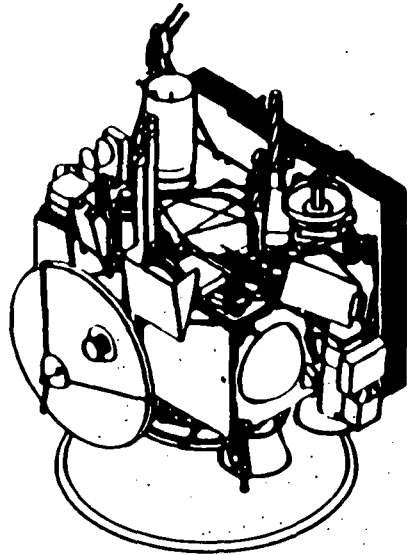
# MARS OBSERVER MISSION SCENARIO



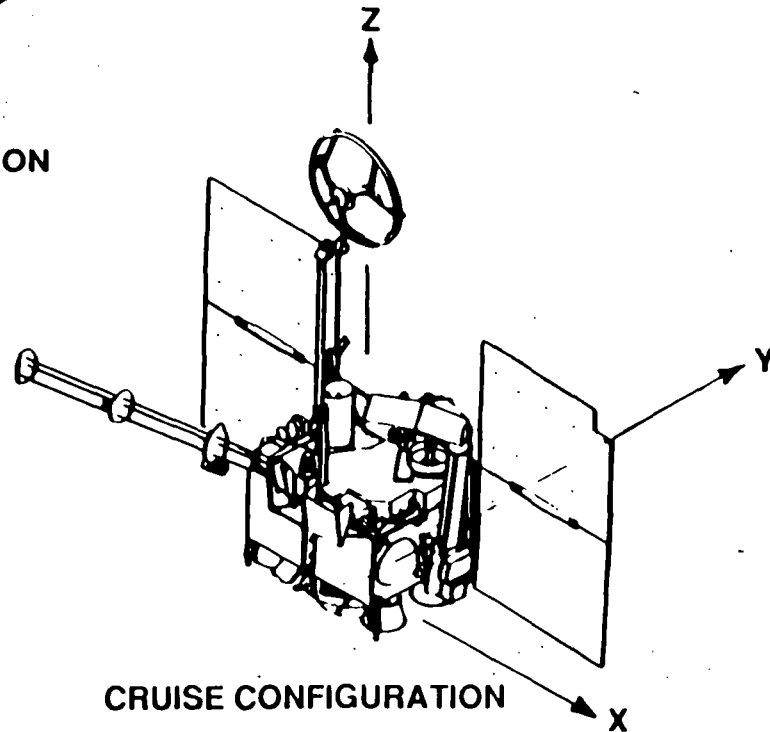
- BASELINE MAPPING ORBIT**
- 378 km INDEX ALTITUDE (NEAR CIRCULAR)
  - 92.86 deg INCLINATION
  - SUN SYNCHRONOUS (2 PM DESC NODE)
  - 117.5 m PERIOD
  - FROZEN (APSIDES)

# MARS OBSERVER SPACECRAFT CONFIGURATIONS

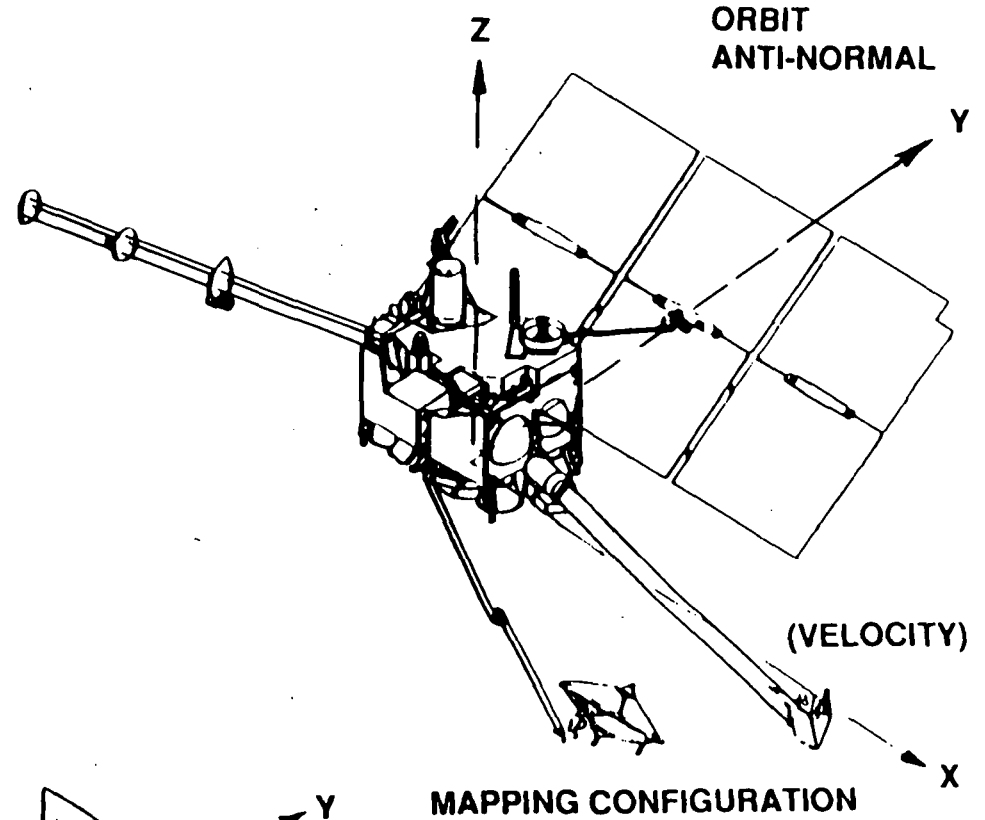
94



**LAUNCH CONFIGURATION**



**CRUISE CONFIGURATION**

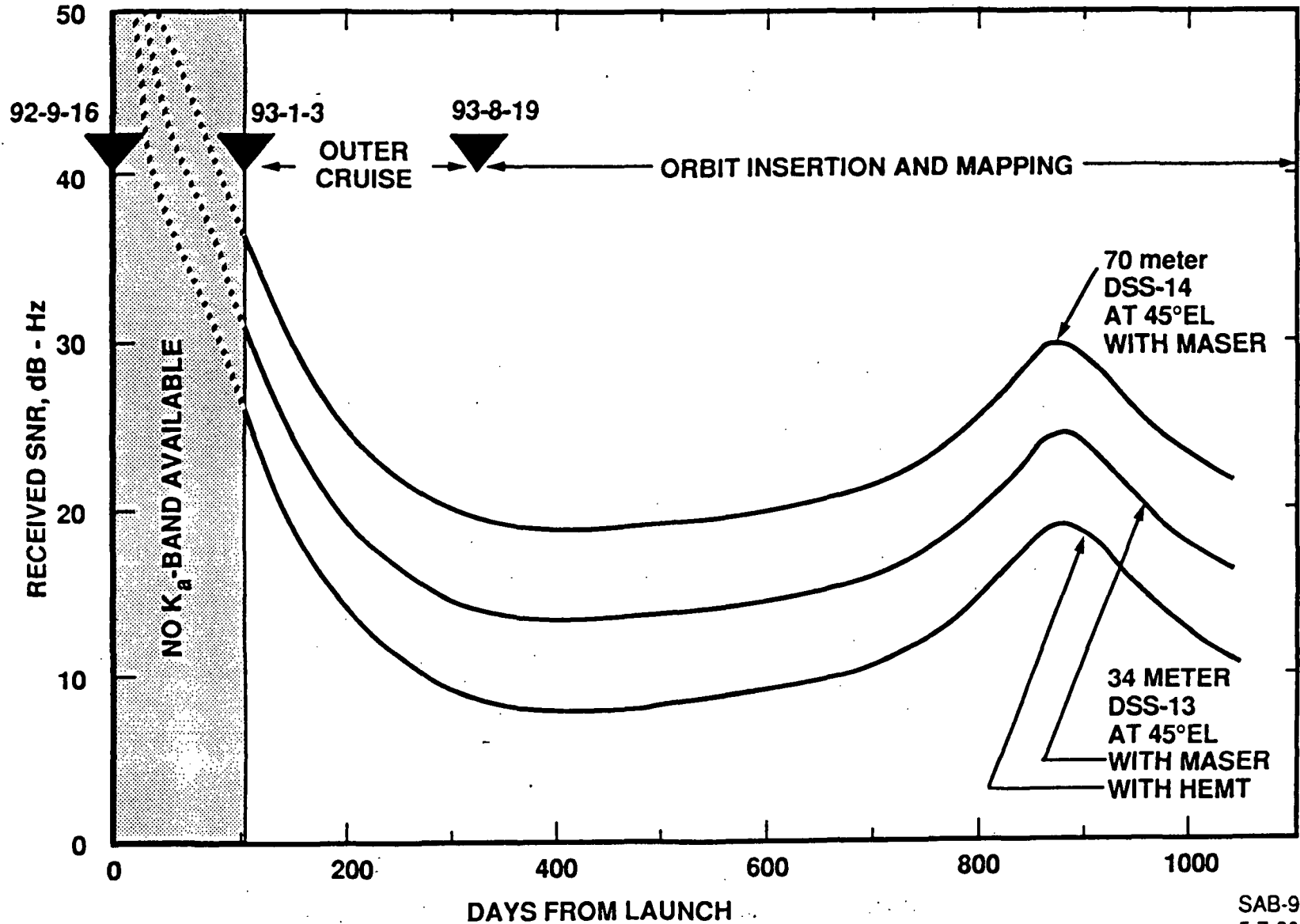


**MAPPING CONFIGURATION**



# MARS OBSERVER

## K<sub>a</sub>-BAND BEACON PERFORMANCE, 90% WEATHER AT GOLDSTONE, CA



56

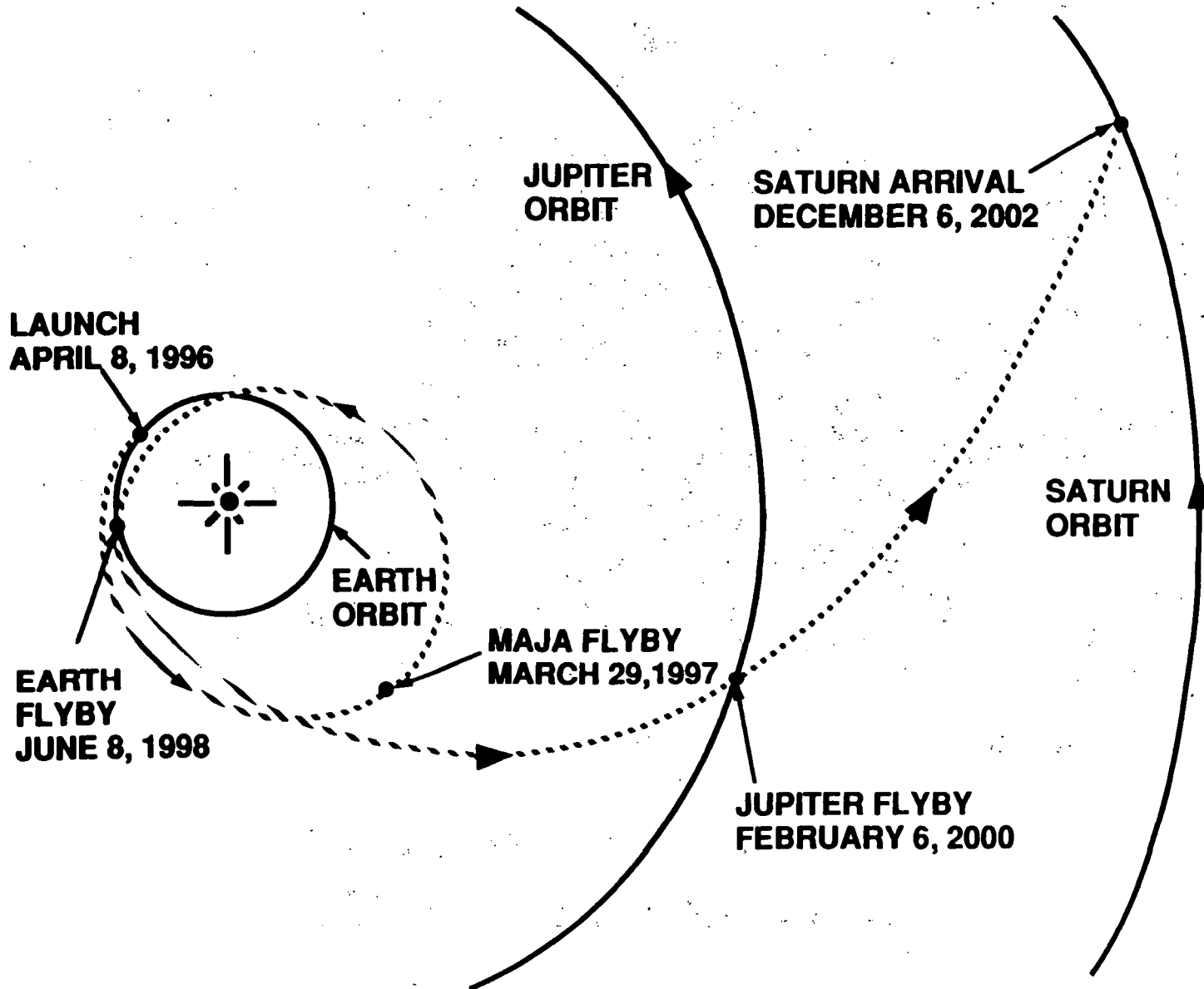


# DOWNLINK BUDGETS FOR MARS OBSERVER TO DSN 34M ANTENNA AT GOLDSTONE CA FOR 90% WEATHER AT 30 DEG ELEVATION

MISSION, DISTANCE AND DATE		MARS OBSERVER AT 2.45 AU, 1994	
SPACECRAFT SYSTEM		X-BAND	K <sub>a</sub> -BAND
FREQUENCY	(GHz)	8.42	33.7
XMTR. PWR.	(dBm)	43.0	10.0
ANT. GAIN	(dBi)	40.0	37.0
POINTING LOSS	(dB)	-0.8	-0.5
EIRP	(dBm)	82.2	46.5
SPACE LOSS		-282.4	-294.4
GROUND SYSTEM			
* ATMO. ATTEN.	(dB)	-0.1	-0.4
ANT. GAIN	(dBi)	67.8	78.0
POINTING LOSS	(dB)	-0.1	-0.2
* NOISE	(dBm/Hz)	-184.6	-182.6
SIGNAL/NOISE	(dB-Hz)	52.1	12.1

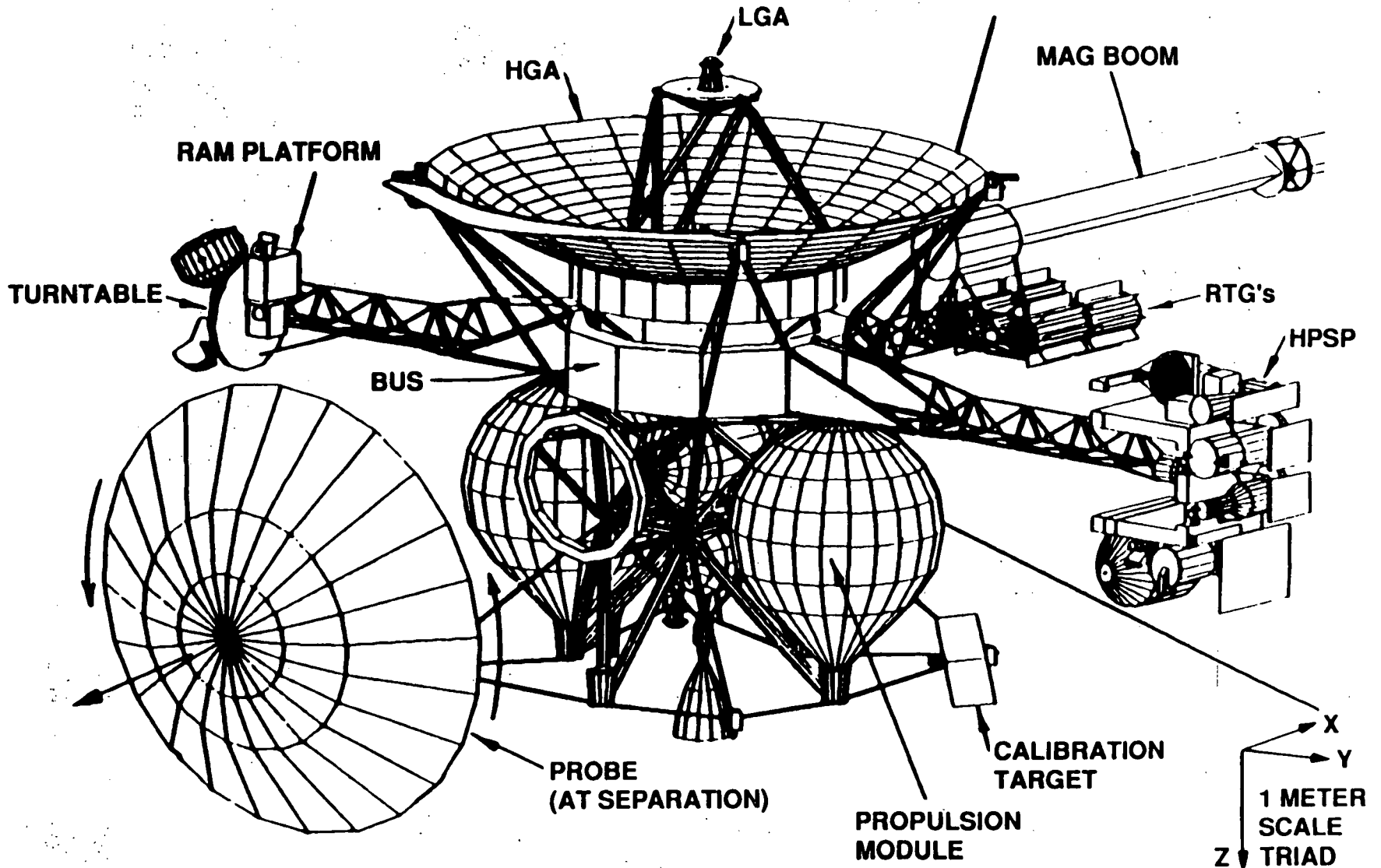
\* VARIES WITH ANTENNA ELEVATION ANGLE AND ATMOSPHERIC CONDITIONS.  
PROBABILITY IS 90% THAT VALUES WILL NOT BE EXCEEDED

# CASSINI INTERPLANETARY TRAJECTORY



97

# CASSINI SPACECRAFT DEPLOYED FRONT ISO VIEW



86



## **K<sub>a</sub>-BAND ADVANTAGES FOR THE CASSINI MISSION**

- **REDUCED SPACECRAFT DC POWER (16 Watts) FOR TELEMETRY**
- **ENHANCED TELEMETRY**
  - **HIGHER TELEMETRY RATES AT LESS POWER**
  - **AVOID TAPE RECORDER CAPACITY CONSTRAINT**
- **IMPROVED RADIO SCIENCE**
  - **REDUCED PLASMA NOISE**
  - **IMPROVED SMALL BODY MASS DETERMINATIONS**
  - **GRAVITATIONAL WAVE DETECTION**
- **IMPROVED NAVIGATIONAL ACCURACY**
  - **4 TO 6 dB REDUCED THERMAL NOISE FOR DOPPLER**
  - **WIDER BANDWIDTH FOR VLBI**



# CASSINI K<sub>a</sub>-BAND DOWNLINK SYSTEM ASSESSMENT

<u>PARAMETER</u>	<u>BASELINE X-BAND</u>	<u>K<sub>a</sub>-BAND</u>
● CLASS	A	A
● FREQUENCY (GHz)	8.4	32.0
● ANTENNA GAIN (DB)	48.3	59.3
● POINTING ACCURACY REQUIREMENT (MRAD, 3σ)	2.0	0.87
● RF POWER OUT (W)	10.6	5.0
● POWER REQUIREMENT DURING DOWNLINK (W)		
● TRANSMITTER DC POWER INPUT	40.0	20.0
● K <sub>a</sub> EXCITER	-	3.5
● SUN SENSOR	1.4	2.0
● TOTAL	41.4	25.5
● MAGNETIC TAPE RECORDER	HIGH DUTY CYCLE	LOWER DUTY CYCLE
● MASS (KG)		
● MASS DELTA (REDUNDANT EXC, PWR COND, TWT)	-	16.0
● SUN SENSOR	0.5	1.0
● TOTAL	0.5	17.0
● BENEFITS		
● ACHIEVABLE DATA VOLUME @ 10 AU ON 70M G & M WITH 2 RATES PER PASS (GBITS/DAY)	3.6	8.3
● RADIO SCIENCE	HIGH	HIGHER
● NAVIGATIONAL ACCURACY	HIGH	HIGHER

100





# SUMMARY

- **K<sub>a</sub>-BAND OFFERS SIGNIFICANT BENEFITS TO NASA FOR FUTURE MISSION SCIENCE RETURN**
- **A DEVELOPMENT ROADMAP IS IN PLACE AND BEING FOLLOWED**
- **PROPAGATION EXPERIMENTS WILL PLAY A KEY ROLE**