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LWIR Detector Requirements for Low-Background Space Applications

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Detection of "cold" bodies (200 - 300 K) against space backgrounds has many important applications, both military and non-military. The detector performance and design characteristics required to support lowbackground applications are discussed, with particular emphasis on those characteristics required for space surveillance. The status of existing detector technologies under active development for these applications is also discussed. In order to play a role in future systems, new, potentially competing detector technologies such as multiple quantum well detectors must not only meet system-derived requirements, but also offer distinct performance or other advantages over these incumbent technologies.

# LWIR DETECTOR REQUIREMENTS FOR LOW-BACKGROUND SPACE APPLICATIONS

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# TOPICS

## APPLICATIONS OVERVIEW

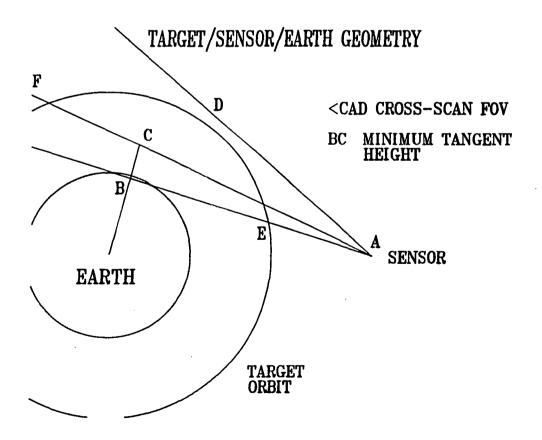
- OBJECTS OF INTEREST
- BACKGROUNDS
- RADIATION ENVIRONMENT
- SENSORS
- DETECTOR REQUIREMENTS

## APPLICABLE DETECTOR TECHNOLOGIES

- STATE OF THE ART
- TECHNOLOGY DEVELOPMENT DIRECTIONS
- REQUIREMENTS FOR NEW, COMPETING DETECTOR TECHNOLOGIES

## LOW-BACKGROUND LWIR APPLICATIONS

- STRATEGIC DEFENSE
  - SURVEILLANCE, ACQUISITION, TRACKING,
  - DISCRIMINATION, AND KILL ASSESSMENT ("SATKA")
  - WEAPON SYSTEM SUPPORT (FIRE CONTROL, HOMING, ETC.)
- **= OTHER MILITARY APPLICATIONS** 
  - RESIDENT SPACE OBJECT SURVEILLANCE
  - DETECTION OF NEWLY LAUNCHED OBJECTS
  - TREATY MONITORING
- NON-MILITARY APPLICATIONS
  - INFRARED ASTRONOMY
  - NEAR-EARTH PHENOMENOLOGY
  - SPACE 'JUNK' DETECTION AND TRACKING



## **OBJECTS OF INTEREST**

= EMISSIVITY-AREA PRODUCTS:

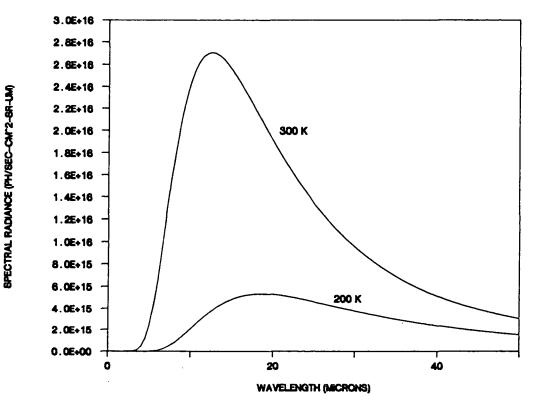
 $0.1 - 10 M^2$ 

**TEMPERATURE:** 

200 - 300 K

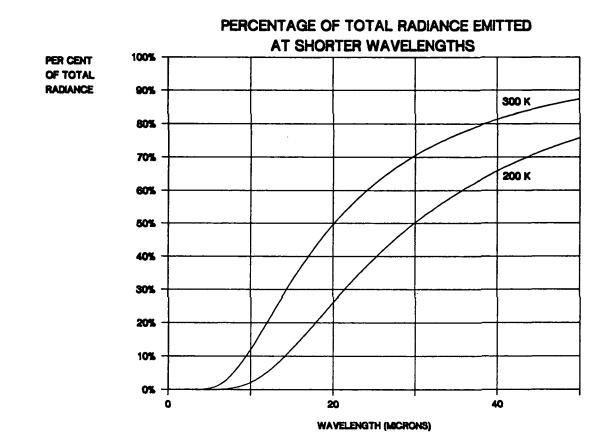
**RANGES**:

1000 - 8000 KM





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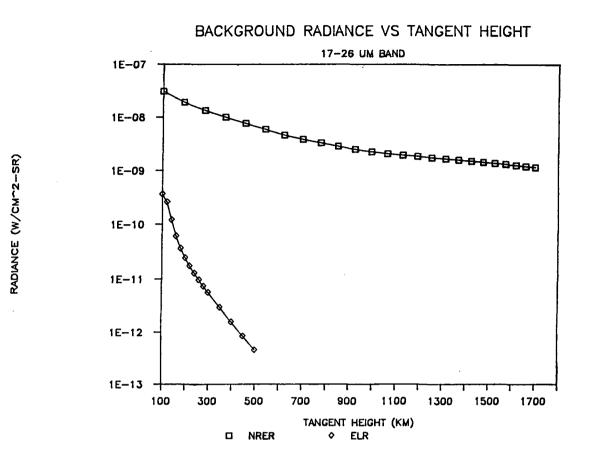


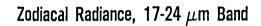
**IR BACKGROUNDS** 

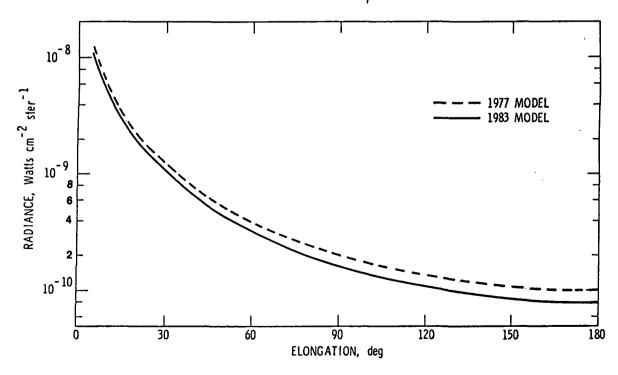
SPACE BACKGROUNDS

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- EARTH LIMB ("ELR")
- ZODIACAL
- CELESTIAL
- NON-REJECTED EARTH RADIANCE ("NRER")
- ENHANCED (NATURAL AND NUCLEAR)
- FOR TANGENT HEIGHTS > 100 KM AND REALISTIC ASSUMPTIONS ABOUT THE LEVEL OF LIKELY OPTICS CONTAMINATION, NRER WILL BE THE DOMINANT NON-ENHANCED BACKGROUND FOR NEAR-EARTH LINES OF SIGHT





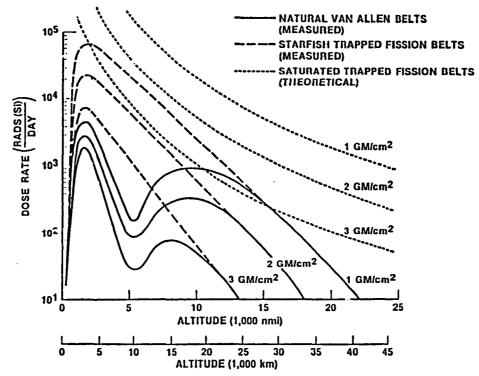


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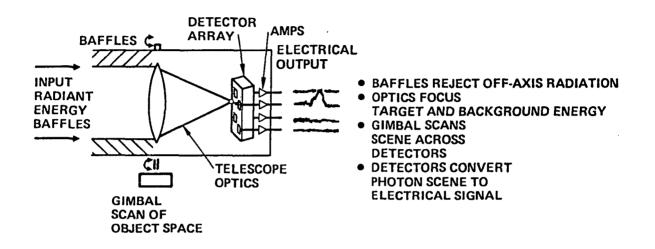
## **RADIATION ENVIRONMENT**

- NATURAL ENVIRONMENT
  - WORST CASE DOSE RATE AT FPA, ASSUMING 0.5 INCH ALUMINUM SHIELDING:
    - 0.02 RAD(SI)/SEC (5 x 107 GAMMAS/CM2-SEC)
  - WORST CASE TOTAL DOSE AFTER 5 YEARS ON ORBIT: 3 x 10<sup>6</sup> RAD(Si)
- ENHANCED NUCLEAR ENVIRONMENT
  - TRANSIENT DOSE RATE DUE TO NUCLEAR DETONATIONS CAN BE ORDERS OF MAGNITUDE HIGHER
  - SUSTAINED DOSE RATE DUE TO SATURATED BELT CONDITION CAN BE 1 RAD(SI)/SEC, WORST CASE
  - WORST CASE TOTAL DOSE DUE TO SATURATED BELTS: , 10<sup>7</sup> ACCUMULATED OVER 10-300 DAYS

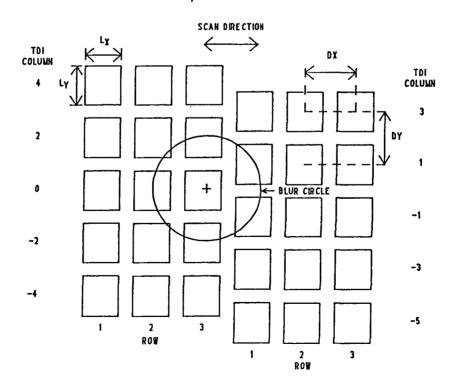
## EARTH'S VAN ALLEN BELTS VERSUS ALTITUDE AT 00



# SENSOR SYSTEM OPERATION

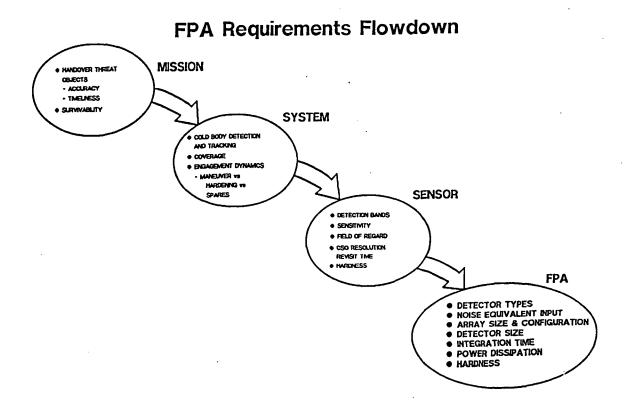


SAMPLE FPA DESIGN CONFIGURATION THREE TDI STAGES, TWO STAGGERED SUBARRAYS



## SIGNAL PROCESSING CHAIN

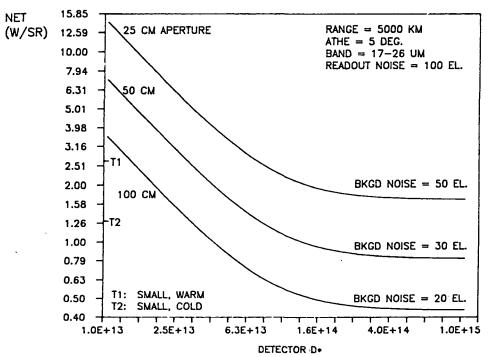
- Detector Bias/ Photocurrent Integration
- Charge-to-Voltage Conversion/ Buffer Amplifier
- First Stage Multiplexing
- Analog-to-Digital Conversion
- Gain & Offset Correction
- TDI
- BG Subtraction/ Clutter Suppression
- Matched Filtering/ Centroiding/ Track Formation ...



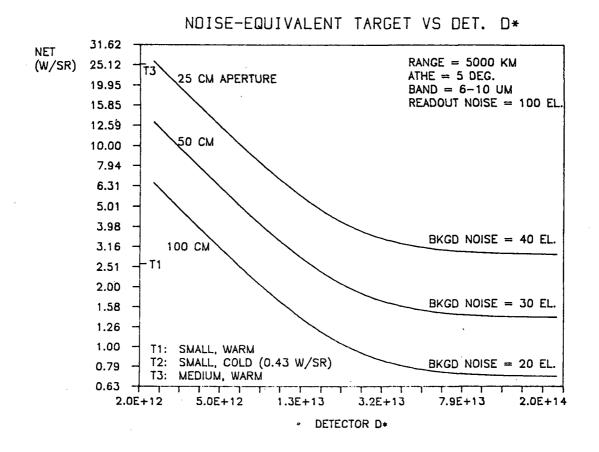
### SENSOR PERFORMANCE LIMITED BY NOISE

- DETECTOR NOISE
  - MAY BE REDUCED BY COOLING
- READOUT NOISE
  - DEPENDS STRONGLY ON COUPLING CIRCUIT DESIGN AND DEVICE CHARACTERISTICS
- IR BACKGROUND NOISE
  - RANDOM FLUCTUATIONS OF IN-FOV SOURCES
    - FUNDAMENTAL LIMIT ON SENSOR PERFORMANCE DUE TO NRER VIEWING NEAR EARTH, ZODIACAL RADIANCE VIEWING AWAY FROM EARTH
  - SPATIAL STRUCTURE ("CLUTTER")
  - OPTICS THERMAL EMISSION

CAN BE RENDERED NEGLIGIBLE BY COOLING



NOISE-EQUIVALENT TARGET VS DETECTOR D\*



### KEY DETECTOR REQUIREMENTS FOR LOW-BACKGROUND SPACE APPLICATIONS

SPECTRAL COVERAGE:

# SIGNIFICANT BROAD-BAND RESPONSE WITHIN 6-30 UM REGION, E.G., QUANTUM EFFICIENCY $\scriptstyle >$ 40% OVER A 5 UM SUB-BAND

SENSITIVITY:

D+ + 1E14 CM-HZ /W AT 20 UM

D+ > 5E13 CM-HZ /W AT 10 UM

= FREQUENCY RESPONSE:

BANDPASS > 10 KHZ, NO ANOMALIES

**DYNAMIC RANGE:** 

LINEAR RESPONSE FROM NOISE LEVEL TO 1E4 TIMES THE NOISE LEVEL

### KEY DETECTOR REQUIREMENTS FOR LOW-BACKGROUND SPACE APPLICATIONS (CONT.)

#### ■ POWER DISSIPATION:

POWER DISSIPATED ON FPA + 10 UW/DETECTOR

#### RADIATION HARDNESS:

TOTAL DOSE HARDNESS + 1E6 RAD(SI)

EFFECTIVE GAMMA AREA ( 1E-6 CM (100 UM DET.)

■ PIXEL SIZE:

50-150 UM (SQUARE)

#### ■ CONFIGURATION:

TWO-DIMENSIONAL MOSAIC ARRAYS, E.G., 10-20 X 50 DETECTORS PER CHIP

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## APPLICABLE TECHNOLOGIES

#### **DETECTORS**

	SPECTRAL CUTOFF (UM)	OPERATING TEMPERATURE (K)
StAs BC	26	12
SiGa BC	17	18
PV HgCdTe	10	40

#### READOUTS

- MATERIALS: SILICON, GERMANIUM, GAAs
- VERY LOW NOISE, RADIATION HARD DEVICES ARE UNDER DEVELOPMENT

## TECHNOLOGY ASSESSMENT

#### StAs BC

- MEETS PERFORMANCE REQUIREMENTS
- LOW OPERATING TEMPERATURE REQUIRES ADVANCED S-STAGE CRYOCOOLERS FOR SPACE-BASED SYSTEMS
- PRODUCELITY DEMONSTRATION PLANNED

#### SiGa BC

- REQUIRES DEVELOPMENT
- OPERATING TEMPERATURE NOT HIGH ENOUGH TO ALLEVIA TE CRYOCOOLER PROBLEM (3 STAGES STILL REQUIRED)

#### LWIR PV HgCdTe

- NDIVIDUAL DETECTORS WITHIN ARRAYS MEET REQUIREM ENTS
- OPERATING TEMPERATURE COULD BE SUPPORTED BY A 2- STAGE COOLER
- SEVERE NON-UNIFORMITY PROBLEM
- UNSUITABLE FOR SOME STRATEGIC DEFENSE SURVELLINA NCE MISSIONS
  - TRACKING COLD TARGETS
  - DISCREMENTION

### TECHNOLOGY DEVELOPMENT DIRECTIONS

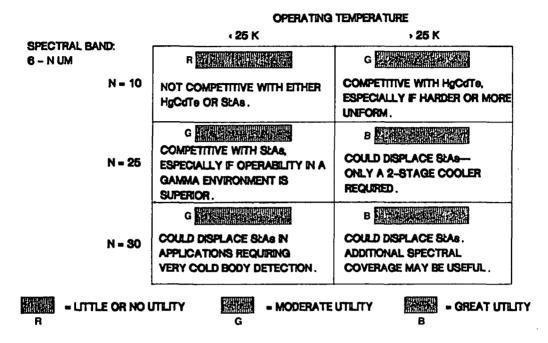
- GREATER SENSITIVITY
  - ULTRA-LOW NOISE READOUTS
  - IMPROVED ROA UNIFORMITY OF HgCdTe DETECTORS

#### GREATER TOTAL DOSE HARDNESS

- ULTRA-RAD HARD READOUTS
- IMPROVED HARDNESS OF HgCdTe DETECTORS
- **GREATER OPERABILITY IN GAMMA ENVIRONMENTS** 
  - DEVELOPMENT OF "INTRINSIC EVENT DISCRIMINATION" (IED) CONCEPTS
- **GREATER PRODUCIBILITY** 
  - PILOT LINE DEMONSTRATION FOR SLAS BC HYBRIDS IS PLANNED

### **REQUIREMENTS FOR NEW, COMPETING TECHNOLOGIES**

ASSUMPTION: NEW TECHNOLOGY HAS PERFORMANCE EQUIVALENT TO OR SUPERIOR TO THAT OF THE INCUMBENT TECHNOLOGY WITH WHICH IT COMPETES



### SUMMARY

- DETECTOR SENSITIVITY REQUIREMENTS FOR LOW BACKGROUND SPACE APPLICATIONS ARE STRINGENT AND ARE DRIVEN BY:
  - STRESSING MISSIONS, E.G., DIM TARGETS, LONG RANGES
  - LOW BACKGROUND NOISE LIMIT
  - AVAILABILITY OF LOW NOISE READOUTS
- DETECTOR RADIATION HARDNESS REQUIREMENTS ARE ALSO STRINGENT:
  - SPACE BASING MAKES HIGH TOTAL DOSE HARDNESS ESSENTIAL
  - STRATEGIC DEFENSE SURVIVAL AND OPERABLITY REQUREMENTS ARE EXTREMELY STRESSING
- NEW TECHNOLOGIES SUCH AS MOW DETECTORS CAN COMPETE WITH THE BETTER DEVELOPED EXISTING TECHNOLOGIES ONLY IF THEY:
  - MEET PERFORMANCE REQUIREMENTS
  - OFFER A SUBSTANTIAL ADVANTAGE OVER AN EXISTING TECHNOLOGY
    - $\mbox{\tt P}$  Higher operating temperatures (\* 25 K) with spectral coverage to 25-30 um
    - Higher Performance or producedity than hocito with comparable spectral coverage at a comparable operating temperature