

Appendix F

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LLNL - Contributions to MPD Thrusters for SEI

MPD Thruster Technology Workshop
NASA, Washington, D.C.

LH075075

dwin
E. Bickford Hooper

May 16, 1991

LLNL CAN CONTRIBUTE TO MPD THRUSTER DEVELOPMENT FOR SEI



Near term:

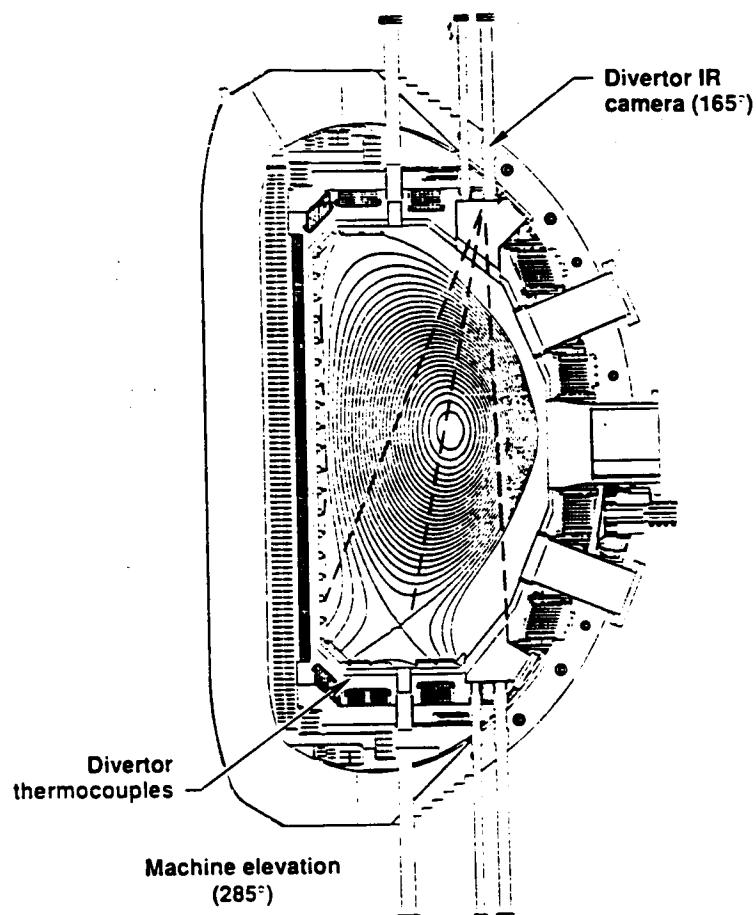
- Modeling of MHD characteristics using the TRAC code, which has been benchmarked against the RACE experiment at LLNL
- Application of tokamak "divertor" physics
 - Modeling of atomic - plasma interactions (gas penetration, ionization, excitation, radiation) using the Brahm and Degas codes
 - Measurements of MHD and atomic effects
 - Modeling of erosion/sputtering and redeposition of refractory materials
- Remote measurements of density, temperature, magnetic field using fusion diagnostics

These contributions can best be made in collaboration with ongoing experiments

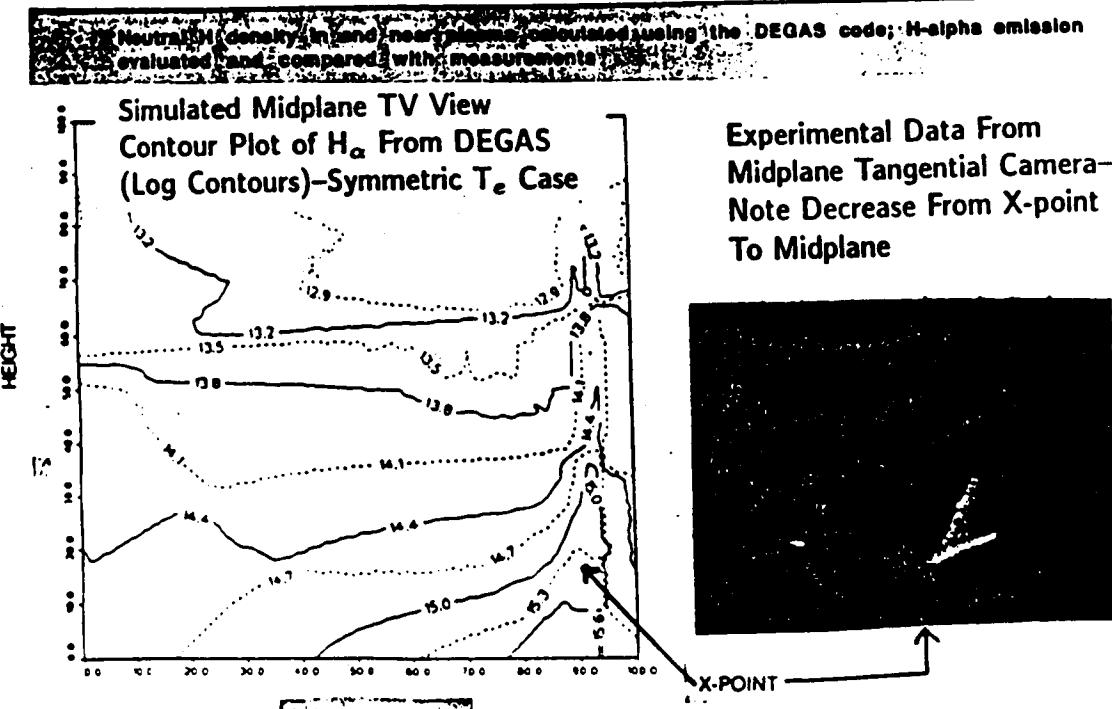
Long term:

- High power tests for lifetime validation using the MFTF-B facility

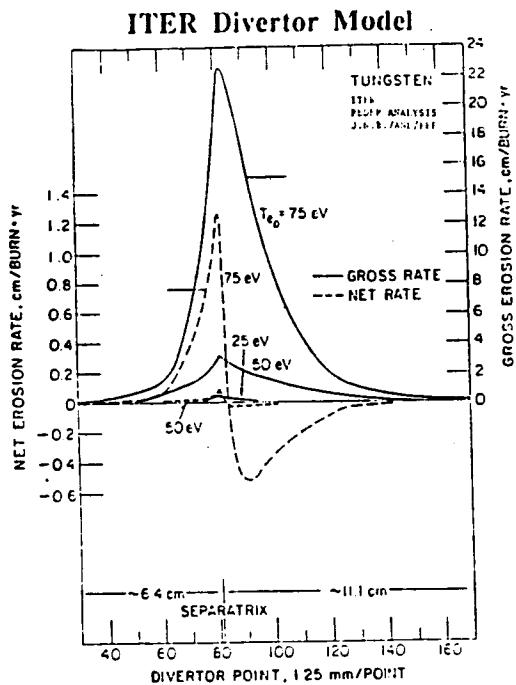
Divertor IR camera



COMPARISON OF TANGENTIAL H_α DATA WITH DEGAS



The MFE community has developed considerable expertise in plasma-induced erosion/redeposition

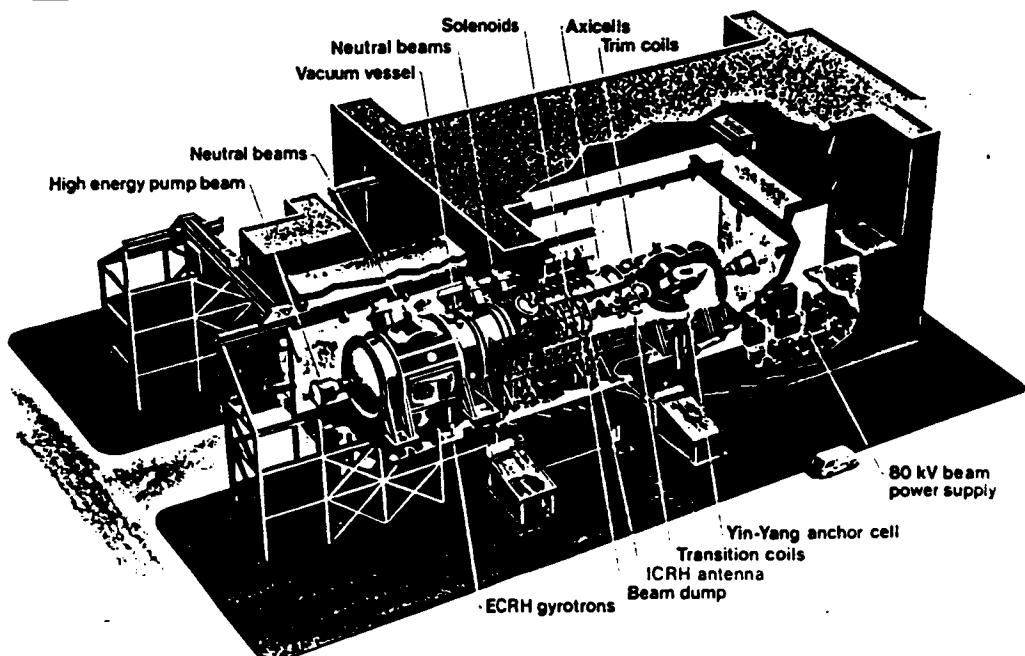


J. Brooks (ANL) Fusion Techn. 18 (1990) 239

KLW051391A

- Computer codes such as REDEP are used to predict net erosion including redeposition effects
- These calculations are benchmarked against measurements in tokamaks and off-line simulation facilities

THE MFTF-B



PROPOSED THRUSTER LIFETIME TEST FACILITY

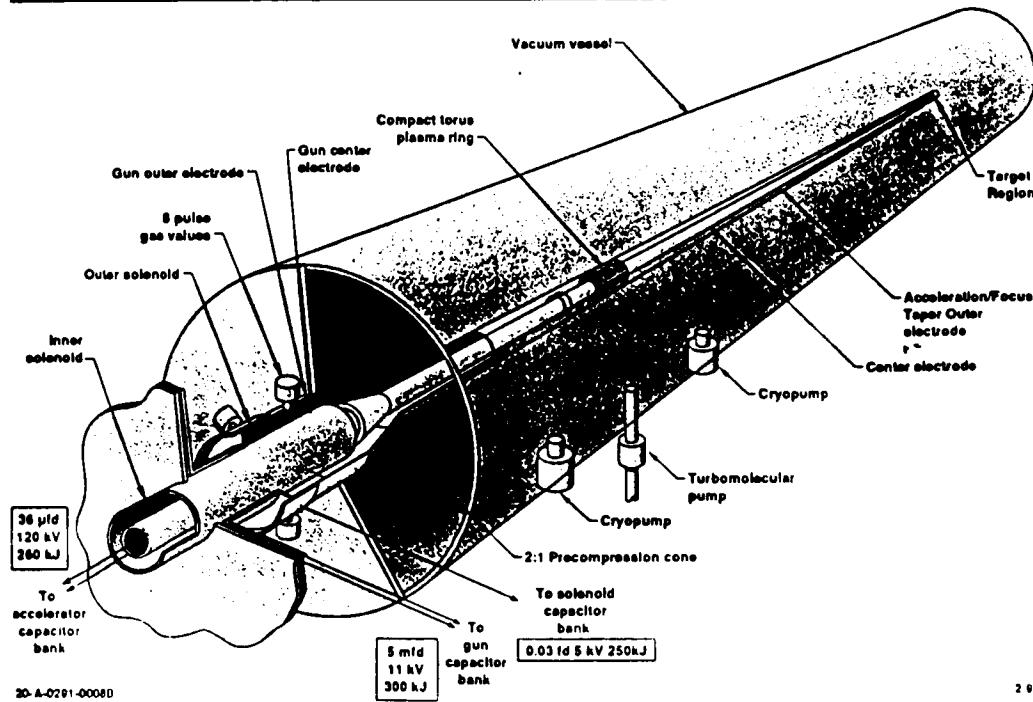
- MFTF-B: Size 35' diameter by 200' long
1000 m³ of cryopanels
11 kW of LHe cooling available for pumping
500 kW closed loop LN₂ system
250 MVA power line
- Example test conditions: mass flow = 0.4 g/s (thruster power = 1 MW at
 $v = 7 \times 10^4$ m/s)

Pumping speed 67×10^6 liters/s, D₂
 $67 \times 10^6 (4/A)^{1/2}$ liters/s,
 Mass A

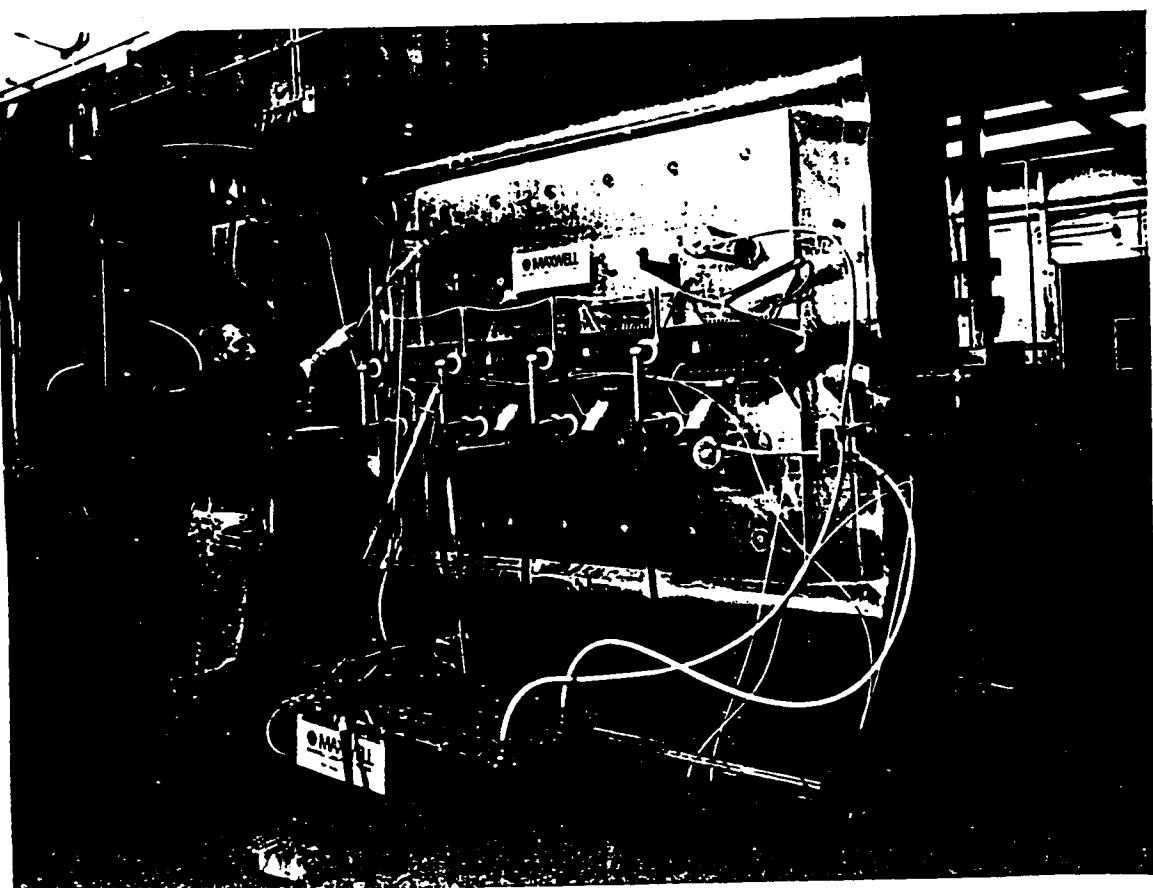
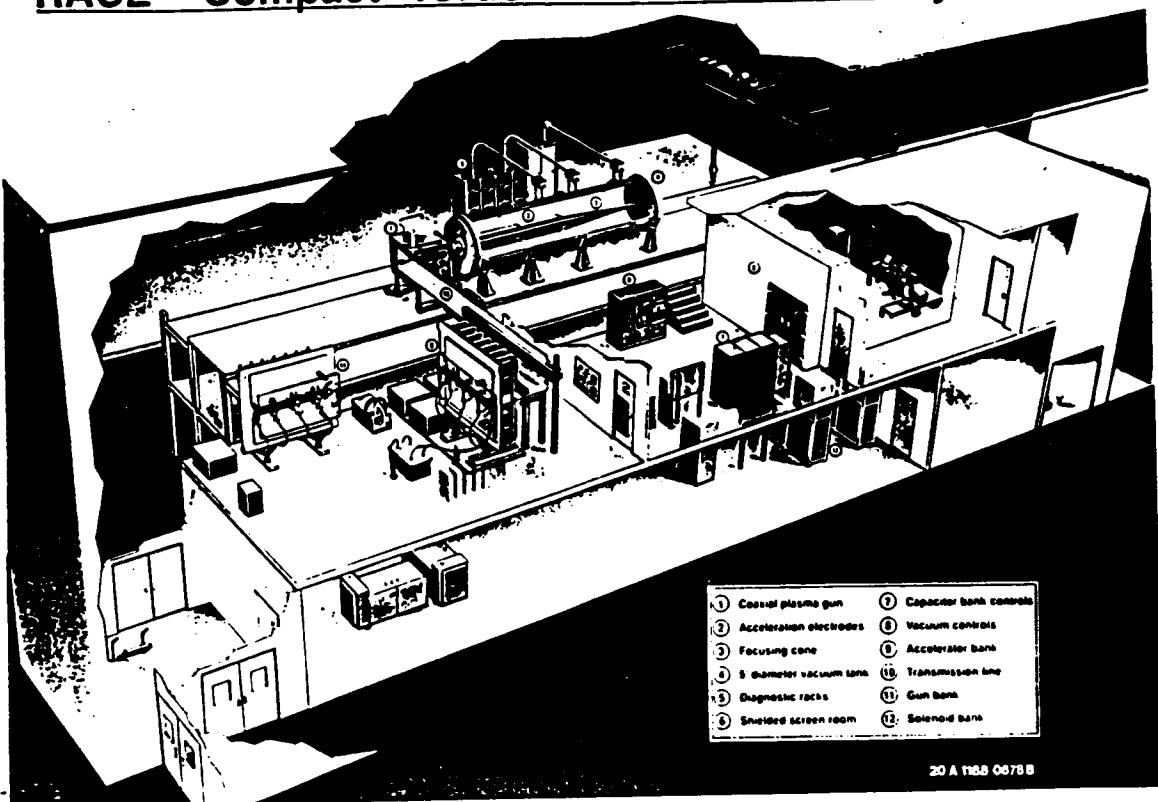
Equilibrium pressure

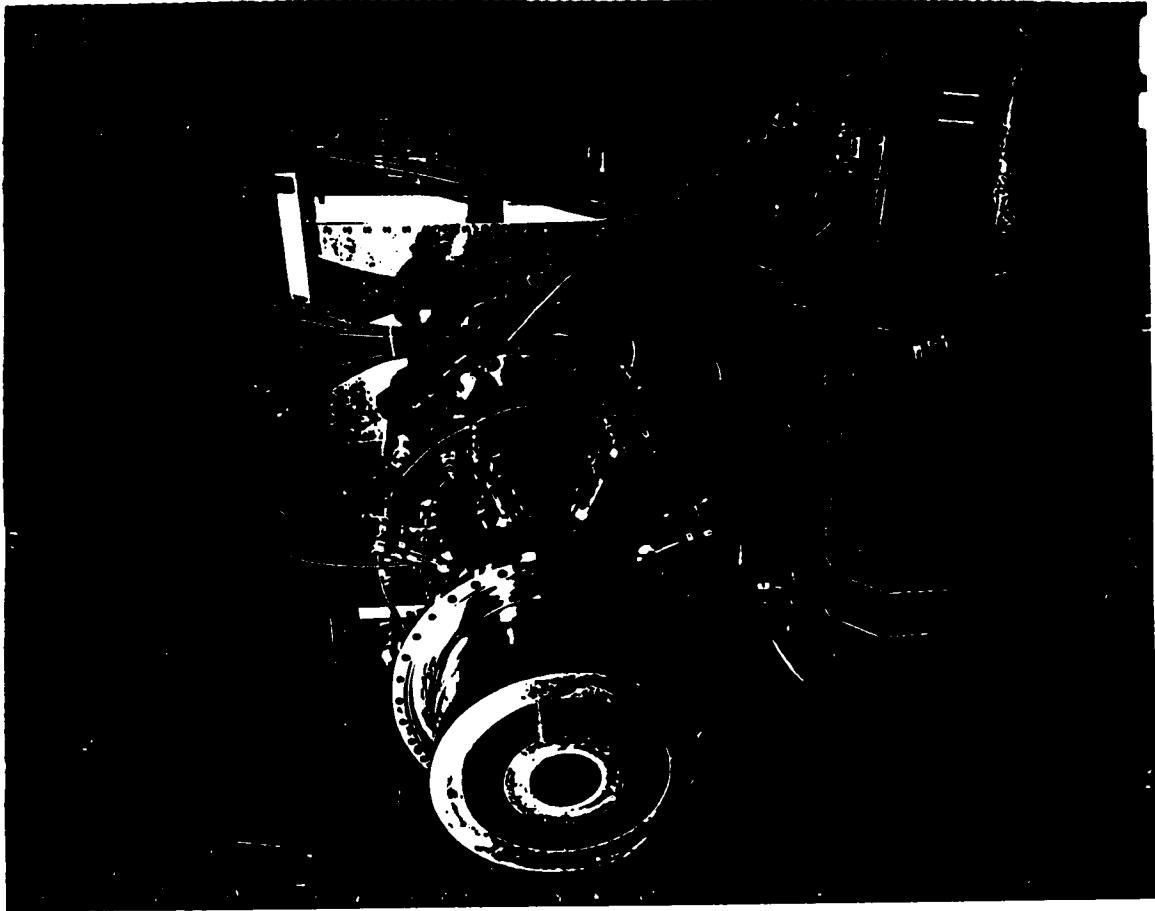
Hydrogen	5.2×10^{-5} torr
Argon	1.6×10^{-4} torr

The RACE experiment test the basic concepts of ring acceleration



RACE Compact Torus Accelerator Facility





RACE program summary

<u>Goals</u>	<u>Predictions</u>	<u>Results to Date</u>
Demonstrate ring formation	Magnetic energy Mass Length	2-40 kJ 5-500 microgram 70 cm
Demonstrate acceleration in linear coaxial system	Velocity Energy Efficiency $U_{kinetic}/U_{magnetic}$	$1-2 \times 10^8$ cm/sec Up to 100 kJ 0.4 5
Demonstrate ring focusing	$R_{focus}/R_0 \sim$	1/5

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Comparison of race data of plasma ring formation with the HAM 2D-MHD code

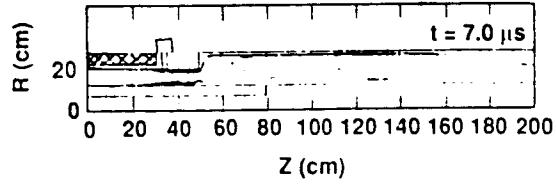
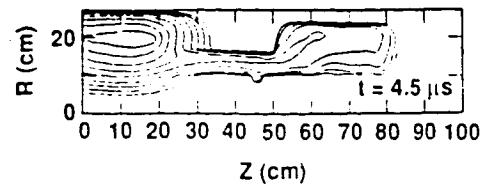
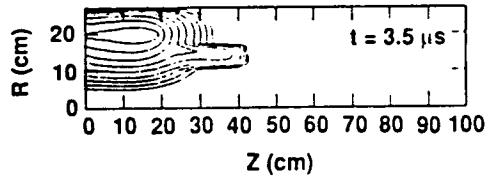
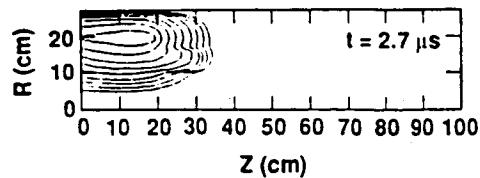


For these calculations HAM:

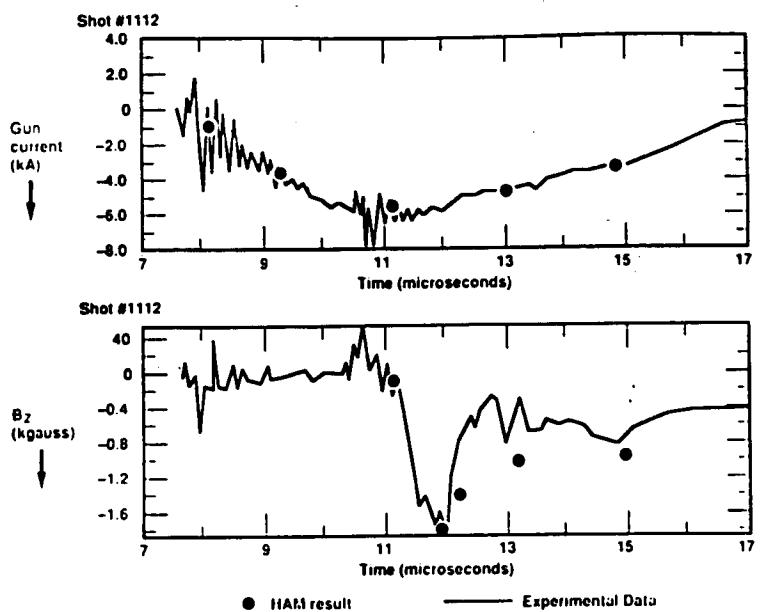
1. Calculates the initial poloidal field allowing for diffusion through conducting electrodes
2. Calculates the time-dependent gas density distribution from an injected puff of gas
3. Calculates gas breakdown and plasma ring formation using the gun capacitor bank parameters

SGC/LU/1994/12

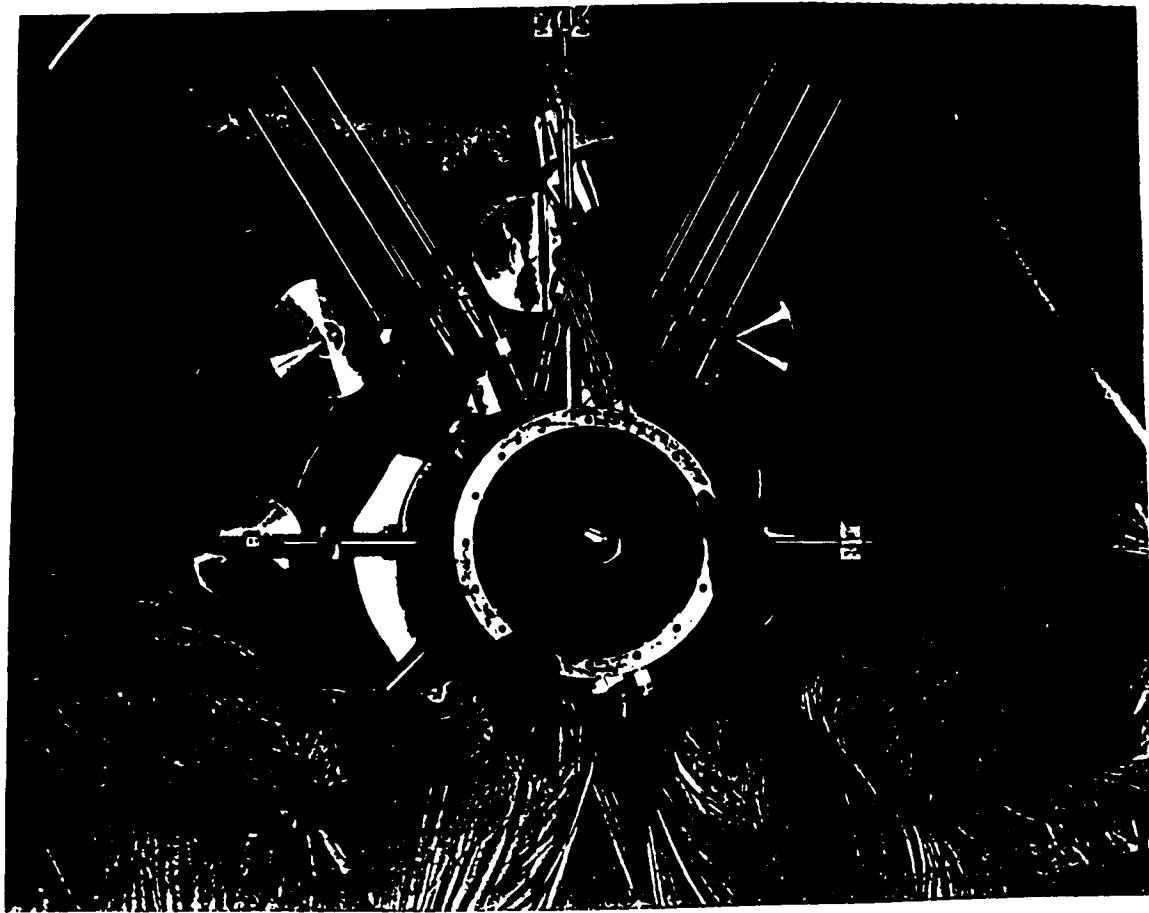
Flux contours for HAM simulation



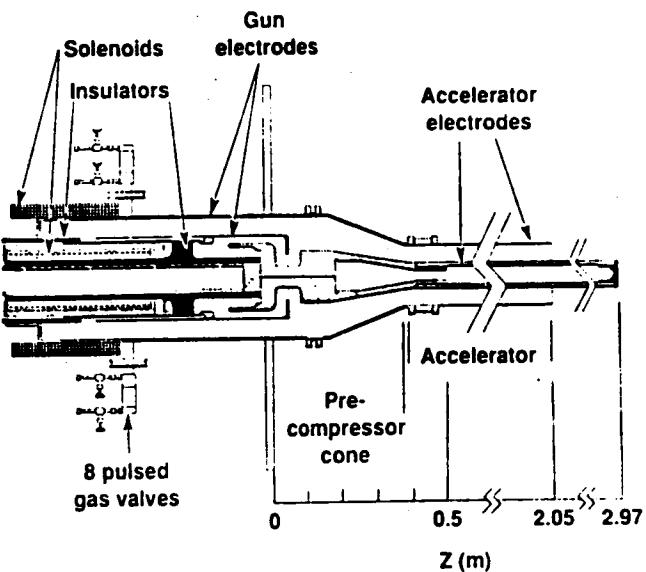
2D MHD simulations agree with the experimentally observed current



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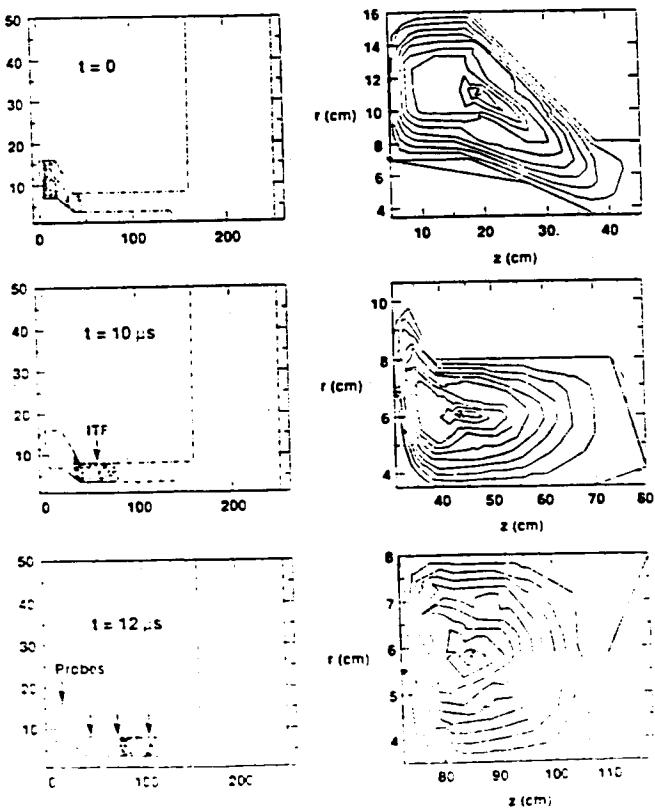


RACE, the Ring ACcelerator Experiment, configuration during precompressor tests

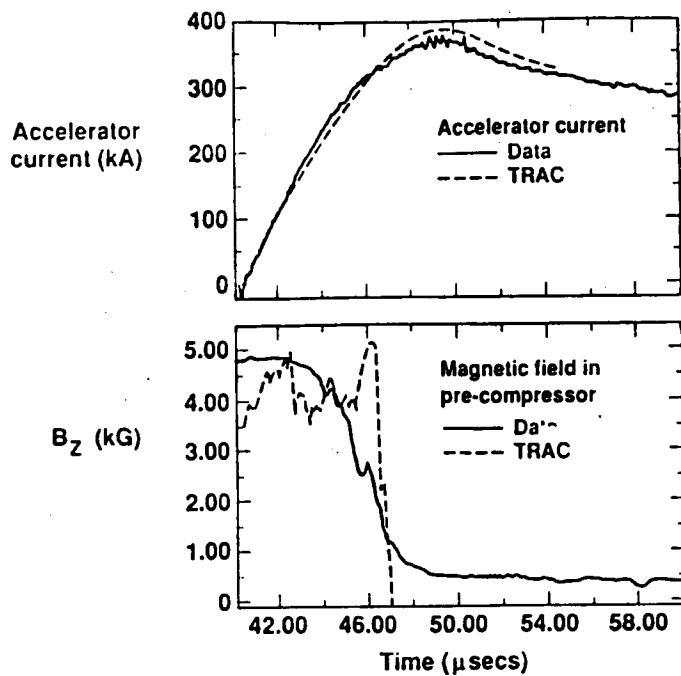


20-A-1090-0340C
Rev 2

TRAC (Two-dimentional Ring Acceleration Code) has been used to model the RACE pre-compressor

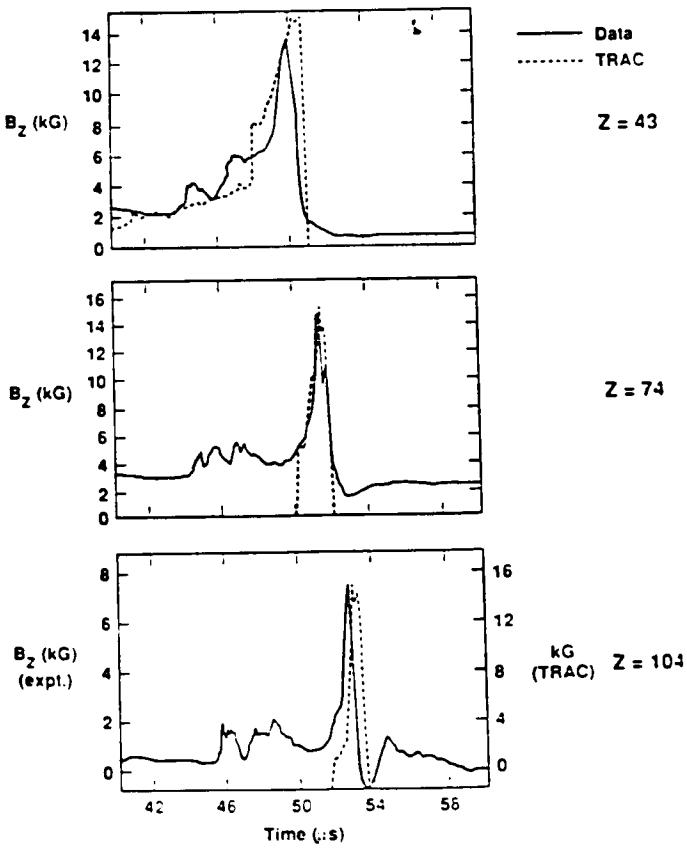


Comparison of trac with shot 5554 ($V_{ACC} = 80$ kV)



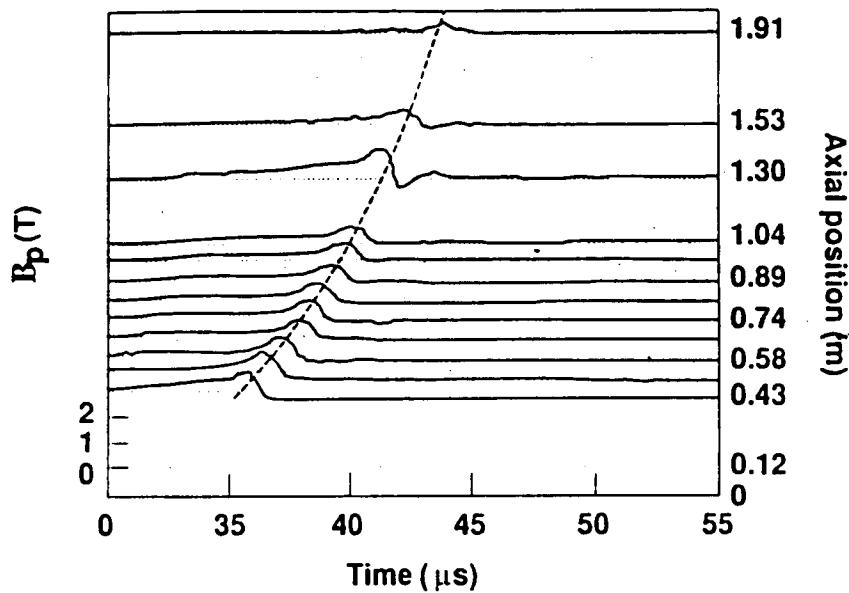
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mc1

Comparison with shot 5554 cont'd B_Z vs. t at different locations in straight section



20-A-0490-0124B
c

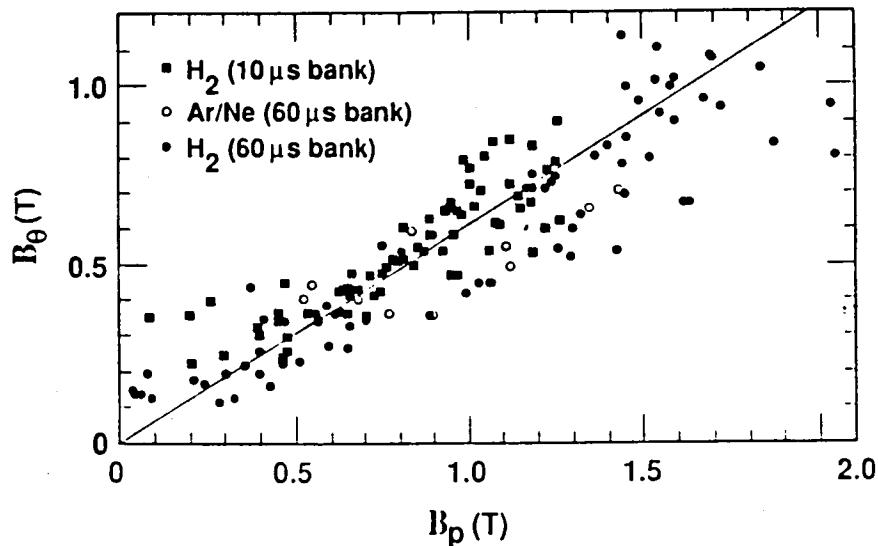
CT accelerates and is stable after precompression



(Vertical offsets of B_p probe signals proportional to axial location)

2G-A-1090-03408
Rev 2

CT in quasi-static pressure balance during compression in conical electrodes



(Accelerator field proportional to poloidal field at 0.43 m for three gun conditions, consistent with line predicted by TRAC code.)

2G-A-1090-0340A
Rev 2

An Alternate Application of MPD Arc Sources: Plasma "Tethers" for Tapping the Solar Wind EMF for Power > 10 MW

Plasma plumes generated by MPD arc sources can extend of order 1000 km across the solar wind magnetic field. The electric field, $E = u_{\text{wind}} \times B$, gives a voltage drop along the plume, and currents are induced as in the AMPTE artificial comet experiments.

The available power is:

$$P = 2 M_p v_p v_A \quad M_p = \text{mass ejection rate} \quad v_p = \text{plume velocity}, \\ v_A = \text{Alfvén velocity}$$

An example:

$$M_p = 10 \text{ g/sec}, \quad v_p = 60 \text{ km/sec}, \quad v_A = 80 \text{ km/sec}, \quad P = 100 \text{ MW}$$

The power could drive thrusters with a specific impulse of about 3000 sec.

A lunar power station could extract large amounts of power since there is unlimited available mass. The energy extracted is about 10^{10} Joules/kg

*Plasma "Tethers" generate a bow shock
in the supersonic solar wind Plasma*

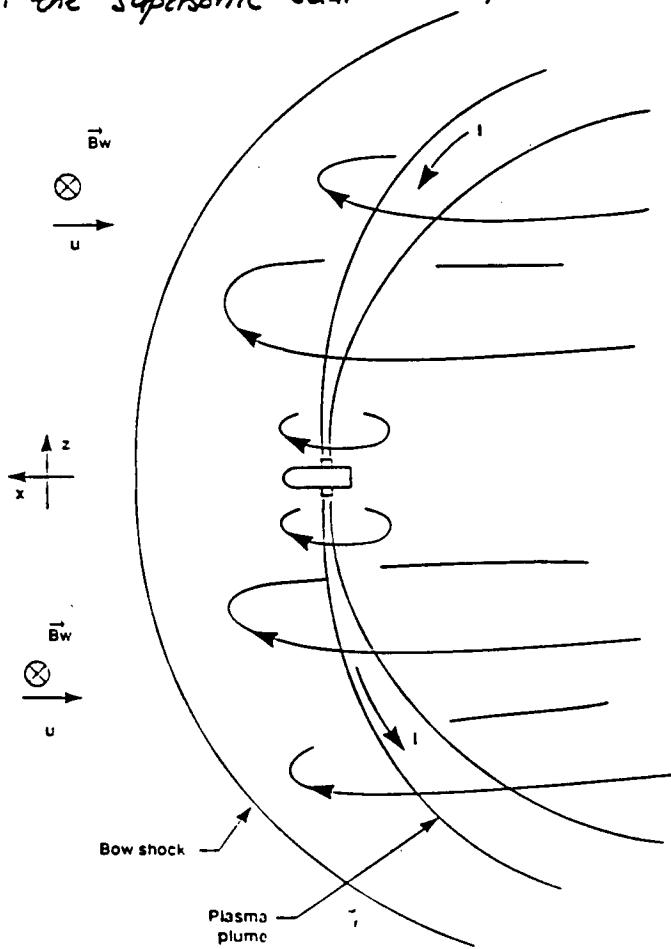


Diagram 1

The plume in cross-section

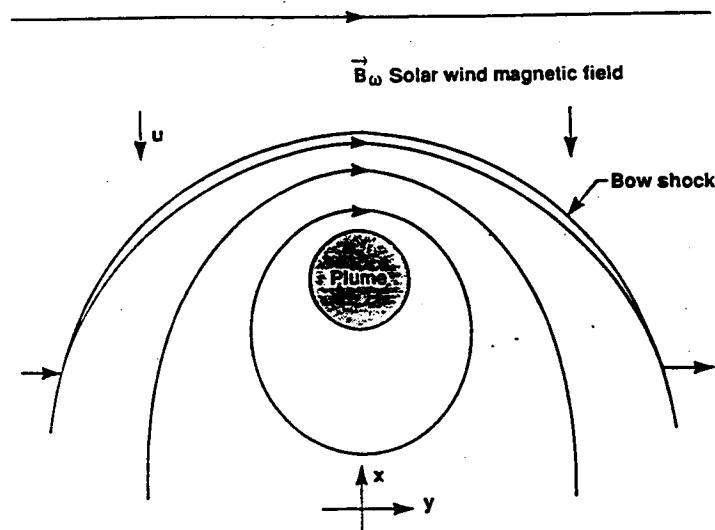


Figure 1

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A conceptual solution for self-sustaining,
plasma guns / thrusters

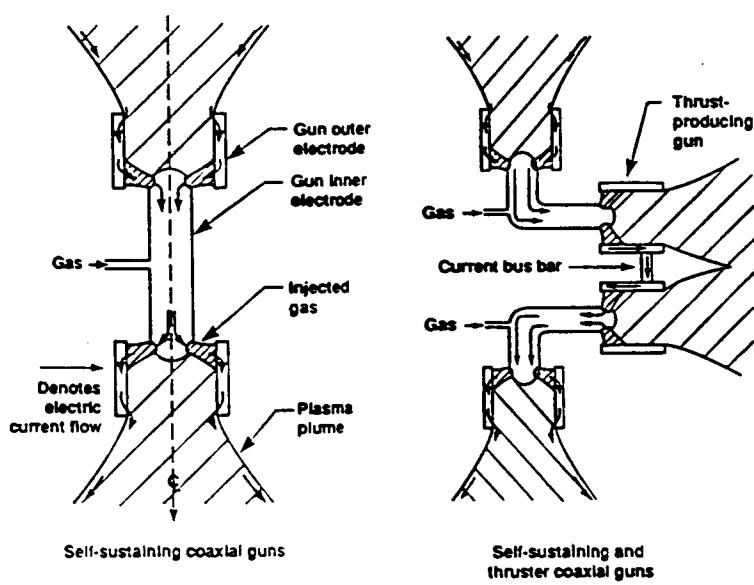
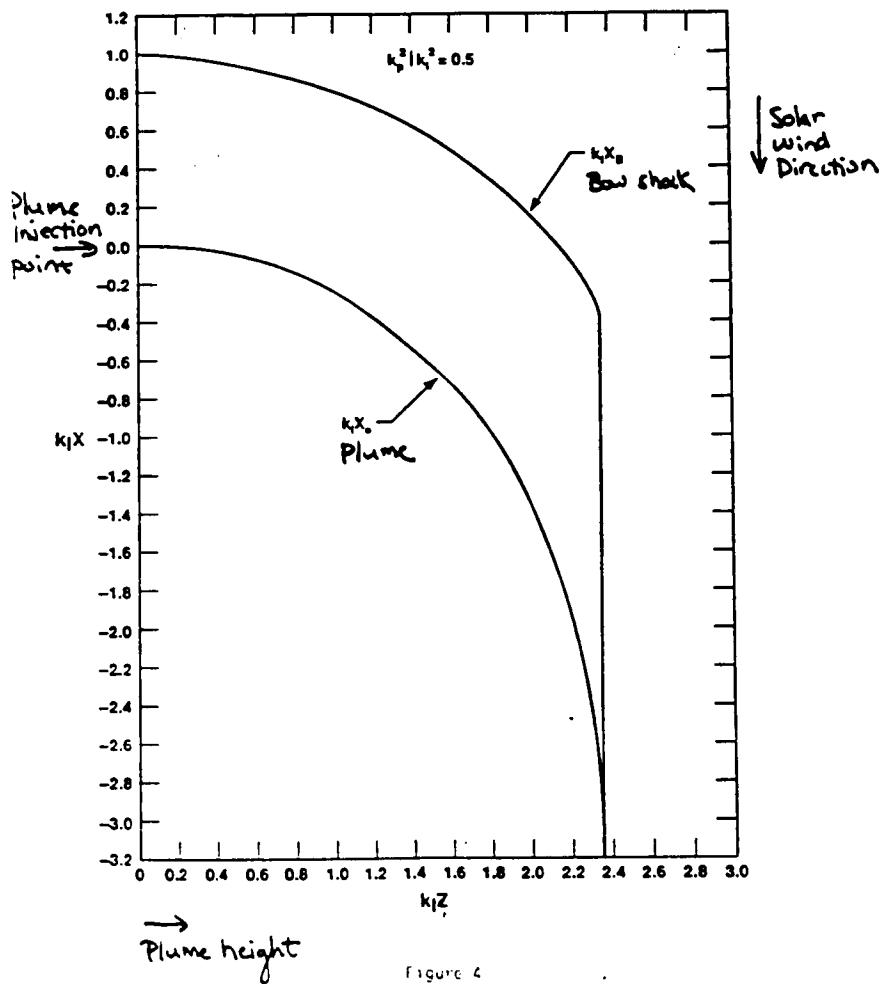


Figure 3

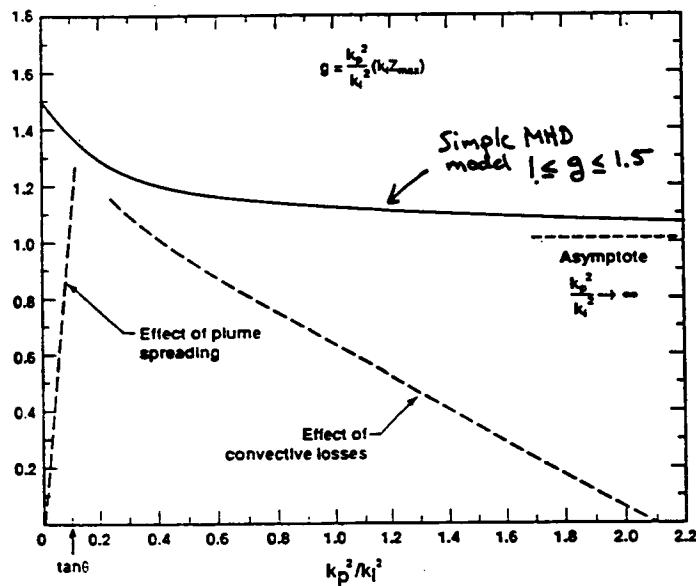
The plume dynamics can be calculated from a simple MHD model



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The Plume Power Extraction is
a function of the dimensionless ratio k_p^2/k_i^2

$$\dot{P} = 2 \dot{M}_p U_p U_A g$$



Θ = plume divergence angle

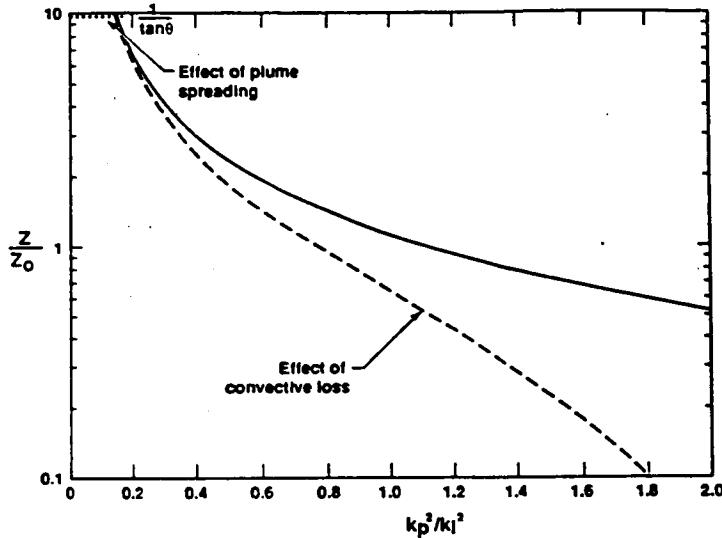
$$\frac{k_p^2}{k_i^2} = \frac{\mu_0}{12} \frac{I^2}{\dot{M}_p U_p}$$

Figure 5

60-X-0491-0021-4

The load impedance is a function of k_p^2/k_i^2 and is of order

$$Z \approx Z_0 = \frac{\mu_0 V_A}{4\pi} \quad V_A = \frac{\text{Solar wind}}{\text{Alfvén velocity}}$$



Choosing Z sets $k_p^2/k_i^2, g$

Figure 6

80-Y-0491-0021-5

Conclusion: LLNL has extensive expertise in physics and technology relevant to MPD thruster development

Areas in which we could contribute include:

Modeling of atomic physics, plasma surface interactions and 2D MHD flows

Results from ongoing high-power plasma accelerator experiments (RACE)

Plasma diagnostics

High pumping speed test stand for lifetime validation studies (MFTF-B)