## VPI/NASA/Senior Design Project 1990-91

## 1N-16-6R <br> An SEP Cargo Vehicle for <br> Mars Missions

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(MASA-CP-139GO5) SUNMASTER: AN SFP CARGO N92-21245
VEHICLE FOR MARS MISSIONS (virginig
Polytechnic Inst. and state univ.) 132 p
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June 1991
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This study examines options for an unmanned solar powered electric propulsion cargo vehicle for Mars missions. The six primary areas of investigation include: trajectory, propulsion system, power system, supporting structure, control system and launch consideration.

Optimization of the low thrust trajectory resulted in a total round trip mission time just under 4 years. The argon-propelled electrostatic ion thruster system consists of seventeen 5 N engines and uses a specific impulse of 10,300 seconds. At Earth, the system uses 12 engines to produce 60 N of thrust; at Mars, five engines are used, producing 25 N thrust. The thrust of the craft is varied between 60 N at Earth to 24 N at Mars due to reduced solar power available.

Solar power is collected by a Fresnel lens concentrator system using a multi-stacked cell. This system provides 3.5 MW to the propulsion system after losses. Support for the concentrator systems is provided by a three ring, hexagonal shaped, precision truss system constructed of aluminum coated, graphite-epoxy, thin-walled struts. The center of the ship is supported by a three-cell, orthogonal, tetrahedral truss.

Control and positioning of the craft are provided by a system consisting of three double gimballed control moment gyros (CMG). Four shuttle " C " launches will be used to transport the unassembled vehicle in modular units to low Earth orbit where it will be assembled using the Mobile Transporter of the Space Station Freedom.

## Acknowledgements

The design team would like to acknowledge the efforts of Julie Fraser, who, with her creative knowledge of software packages and graphic artistry and ability to work several days without sleep, helped in the compilation of the report.

## Table of Contents

Abstract ..... i
Acknowledgements ..... ii
Introduction ..... 1
1.0 Orbital Analysis ..... 4
1.1 Preliminary Delta-V and Propellant Mass Estimates ..... 4
1.2 Trajectory Optimization ..... 5
1.2.1 Introduction ..... 5
1.2.2 Propellant and Thrusting Considerations ..... 6
1.2.3 Time Considerations ..... 7
1.2.4 The Optimized Trajectory ..... 12
1.3 Conclusions ..... 13
2.0 Propulsion Systems ..... 16
2.1 Introduction ..... 17
2.2 Preliminary Concepts ..... 17
2.3 Thruster Optimization and Configuration ..... 20
2.4 Thruster Degradation and Lifetime ..... 27
2.5 Propellant Tank ..... 28
2.6 Power Processing ..... 31
2.7 Thermal Control ..... 33
2.8 Propulsion Module Structure ..... 34
3.0 Solar Power ..... 36
3.1 Introduction ..... 36
3.2 Solar Power Options ..... 37
3.2.1 Concentrator Array ..... 37
3.2.1.1 Array Parameters ..... 37
3.2.1.2 Wiring Configuration ..... 38
3.2.2 Flat Panel Array ..... 38
3.2.3 Array Specifications at Earth ..... 38
3.2.4 Array Heat Transfer ..... 39
3.2.5 Array Wiring Connections ..... 39
3.3 Selection of Solar Power System ..... 40
3.4 Selected Solar Power System Design ..... 41
3.4.1 Solar Cell Configuration ..... 41
3.4.2 Fresnel Lens Concentrator Array ..... 43
3.4.3 Selection of Array Parameters ..... 43
3.4.4 Array Wiring ..... 44
4.0 Structures ..... 52
4.1 Truss Systems ..... 53
4.1.1 Surut Elements ..... 53
4.1.2 Strut Assembly ..... 55
4.2 Solar Array Support ..... 55
4.3 Structural Analysis ..... 58
4.3.1 Strut Analysis ..... 59
4.3.2 Hexagonal Tetrahedral Truss Analysis ..... 61
4.3.3 Vibrational Analysis ..... 64
4.3.4 Central Truss Analysis ..... 66
4.4 Assembly Scenario and Time Estimates ..... 68
4.5 Materials Selection ..... 70
4.6 Notes on the Configuration ..... 73
5.0 SEP Vehicle Launch to LEO ..... 84
5.1 General Considerations ..... 84
5.2 Manifest ..... 86
5.3 Launch One ..... 86
5.4 Launch Two ..... 90
5.5 Launch Three ..... 90
5.6 Launch Four ..... 91
5.7 Repeat Launches from LEO ..... 91
6.0 Summary ..... 94
Appendix 1 ..... 98
Trajectory Calculations ..... 98
A-1.1 Program TIMEOPT.BAS ..... 99
A-1.2 Program SPIRAL ..... 100
A-1.3 Program ORBITB.BAS ..... 101
Appendix 2 ..... 102
Propulsion Calculations ..... 102
Appendix 3 ..... 104
Solar Array Calculations ..... 104
A-3.1 Concentrator/Module Specifications ..... 104
A-3.2 Concentrator Array Wiring Connections ..... 105
A-3.3 Flat Panel Array Specifications ..... 107
A-3.4 Flat Panel Array Wiring Connections ..... 108
Appendix 4 ..... 110
Control System ..... 110
Appendix 5 ..... 116
Communications and Navigation Systems ..... 116
A-5.1 The RF Subsystem ..... 116
A-5.2 The CD\&H Subsystem ..... 119
A-5.3 Navigation and Batteries ..... 120
A-5.4 Power Requirements ..... 120
A-5.5 Mass of System ..... 121

## List of Figures

1.1 Propellant Required (Out) vs. Trip Time ..... 8
1.2 Thrust Duration vs. Outward Trip Time ..... 8
1.3 Propellant Required (Ret) vs. Trip Time ..... 9
1.4 Thrust Duration vs. Return Trip Time ..... 9
1.5 Outward Heliocentric Transfer Orbit Trajectory ..... 11
1.6 Return Heliocentric Transfer Orbit Trajectory ..... 11
2.1 Propellant Comparison ..... 19
2.2 Power Provided ..... 21
2.3 Beam Current vs. Isp ..... 21
2.4 Beam Voltage vs. Isp ..... 23
2.5 Thruster Power vs. Isp ..... 23
2.6 Total Efficiency vs. Isp ..... 24
2.7 Thruster Mass Scaling ..... 24
2.8a Electrostatic Thruster ..... 26
2.8b Propulsion Module ..... 26
2.9 Cryogenic Propellant Tank ..... 29
2.10 Propellant Distribution ..... 29
2.11 Power Distribution Diagram ..... 32
3.1 Triple Stacked Solar Cell ..... 42
3.2 Fresnel Lens Module ..... 42
3.3 Diagram for Sizing Program ..... 45
3.4 Box Sizing ..... 45
3.5 Concentrator Wiring Breakdown ..... 47
3.6 Wiring Divisions of Sections \#1 and \#2 ..... 48
3.7 Wiring Divisions of Section \#3 ..... 48
4.1 Strut Configuration ..... 54
4.2 Node Attachment Illustration ..... 54
4.3 Hexagonal Truss ..... 57
4.4 Nodal Displacements for the Hexagonal Truss ..... 63
4.5 Hexagonal Truss Axial Forces ..... 63
4.6 Orthogonal Tetrahedral Truss Configuration ..... 67
4.7 Box Truss Axial Forces ..... 67
4.8 Nodal Displacements for the Box Truss ..... 69
4.9 Thermally Induced Tip Distortion ..... 69
4.10a Configuration ..... 75
4.10b Main Body Configuration ..... 76
4.11a Configuration ..... 77
4.11b Main Body Configuration ..... 78
4.11c Main Body Configuration ..... 79
4.12 Craft Mass Distribution ..... 82
5.1 Shuttle C Launch Capabilities ..... 85
5.2 Earth to LEO Transportation Launch Scenario ..... 87
5.3 Strut Member ..... 88
5.4 Layout of Struts in Honeycomb Shape ..... 88
A-4.1 Control Loop Diagram ..... 112

## List of Tables

1.1 Initial Delta-V Estimates ..... 5
1.2 Mission Architecture Summary ..... 12
2.1 Thruster Characteristics ..... 25
2.2 Mass Summary ..... 30
2.3 Characteristics of the Thermal Control System ..... 33
2.4 Propulsion System Overall Specifications and Mass Breakdown ..... 34
3.1 Comparison of Flat Array and Concentrator System ..... 40
3.2 Specific Wire Selection ..... 49
4.1 Hexagonal Support Truss Parameters ..... 58
4.2 Material Properties Comparison ..... 71
4.3 Material Properties of Titanium ..... 72
4.4 Solar Support Structure Parameters ..... 73
4.5 Ship Specifications and Mass Summary ..... 81
5.1 Launch to LEO Packaging Scenario ..... 93
A-5.1 Parameters of the Communications and Navigation Systems ..... 122
A-5.2 Communications and Navigation Systems Power Requirements ..... 123
A-5.3 Communications Power Requirement Totals for Possible Operations ..... 125
A-5.4 Communication and Navigation Systems Weight ..... 126

## Introduction

As the United States looks forward to the challenges of space exploration in the twenty-first century, many objectives and missions will become feasible. None of these is more likely to inspire the nation and the world as landing men and women on the surface of another planet: Mars. Even with projected technology, such manned missions will be long and costly. To minimize the number of man-hours in space required for such an objective as well as reduce the cost, the possibility of having several unmanned cargo missions precede the manned expeditions becomes attractive. Costs could be minimized by having cargo vessels employ a high specific impulse electric propulsion system. This type of propulsion system allows for a much higher payload to total mass ratio than is attainable with conventional propulsion. Although electric propulsion results in relatively long trip times, this consideration is not as great a concern as it would be for a manned mission. With this philosophy in mind, this preliminary study was prepared in accordance with a number of design parameters specified by the Space Exploration Initiative Office of the NASA Langley Research Center.

To fulfill its Mars mission, the cargo ship must meet several design requirements and goals. First, the vehicle is an unmanned cargo vessel employing argon-propelled, ion thrusters.

The electric propulsion system will be powered by solar arrays that will convert the sun's energy directly into electric current. Because of the very large size of this solar powered craft, it will have to be assembled in low Earth orbit near Space Station Freedom (SSF). After leaving low Earth orbit, the craft will deliver a 61 mT payload to a low Mars orbit and return to Earth, and it must be designed to survive three such missions without extensive overhaul. Several specific design requirements were also provided. The LEO orbit of SSF is taken to be a 400 km circular orbit at an angle of inclination of $28.5^{\circ}$. The destination orbit upon arrival at Mars is
specified as a 500 km circular orbit at an inclination of $70^{\circ}$. Finally, Earth-to-Orbit launches will be performed by the Shuttle-C. The payload specifics for Shuttle-C are $61 \mathrm{mT}, 7.6 \times 27$ meter volume for cargo launches.

Several key areas of interest and analysis require special attention during this study. The first of these areas addressed is the analysis of the low thrust trajectories required by the mission. These trajectories must be optimized to require a minimum of propellant and thrusting time, as well as result in acceptable refit time, loiter at Mars, and overall mission durations.

Next, the selection and optimization of the propulsion system is considered in Section Two. Important parameters considered here are the optimized specific impulse, power level, and thrust level for the craft. In addition to the selection and optimization process, the problem of thruster degradation over the course of the mission is discussed. Also included in this discussion are the power processing system, propellant storage tank and propellant system, and the problem of rejecting the appropriate heat loss to space, as well as several other topics.

The propulsion system will be powered by a photovoltaic system that will convert sun light into electrical energy. Section Three begins with a discussion of various options for collecting this solar energy. The two primary options considered were a flat array of solar cells, and a Fresnel lens concentrator array. For the option selected, a discussion of heat transfer and wiring requirements is presented, as well as the general array specifications.

This large solar array must be supported by a truss structure able to withstand the anticipated loads. An analysis of this structure is presented in Section Four. This analysis includes the configuration, a discussion of the forces carried by the truss members, as well as the allowable freedom of movement of the overall structure.

Finally, the logistical problem of launching and assembling the components of the craft need to be considered. With the high cost involved in lifting heavy payloads to orbit, as well
as the difficulty and cost of assembly in orbit, it is very desirable to optimize both the launch and construction processes. The details of the solution of this obstacle can be found in Section 5.

Appendices 1 through 3 provided detailed calculations for the trajectory, the propulsion system, and the solar array assembly, respectively.

The attitude and direction of the craft are controlled and adjusted by a system of Control Moment Gyros (CMGs). This system is examined in Appendix 4, and includes a discussion of the control loop.

Appendix 5 discusses the navigation and communication systems of the spacecraft.
With this framework in mind, this preliminary study was organized into the areas of study outlined above.

## SECTION 1

## Orbital Analysis

## Nomenclature

| HTO | Heliocentric transfer orbit |
| :--- | :--- |
| IMLEO | Initial mass in low Earth orbit |
| Isp | Specific impulse |
| LEO | Low Earth orbit |
| mT | Metric tons |
| SSF | Space Station Freedom |

### 1.1 Preliminary Delta-V and Propellent Mass Estimates

The most important calculation for any orbital transfer is the estimate for the total change in velocity (delta-V) required to complete the mission. This delta-V represents the total impulse that the propulsion system must provide, and is therefore critical in estimating the mass of propellant required.

This mission will require a number of orbital transfer maneuvers. In the vicinity of Earth, the craft must initiate a plane change to get from the orbital inclination of SSF to roughly that of the ecliptic, as well as escape from the Earth's gravity. Upon Earth escape, the craft must inject itself into some heliocentric transfer orbit (HTO) that will allow it to reach the orbital radius of Mars. Once there, the craft must allow itself to be captured, then execute another plane change to reach the destination orbit. This process would essentially be reversed for the return trip. Initial estimates of the delta-V's for the trajectory are given in Table 1.1. These values were estimated using computer models that assumed constant tangential thrust.

Table 1.1 Initial Delta-V Estimates

| Maneuver | Delta-V (km/s) |
| :--- | :--- |
| Spiral Earth | 6.5 |
| HTO | 6.0 |
| Spiral Mars | 3.0 |

With the delta-V break down, it is possible to roughly estimate the initial mass of propellant required for the entire mission for a given "dry mass" (no propellant or cargo) for the ship. This is done by using the delta-V for each maneuver and equation 1.1:

$$
\begin{equation*}
\Delta V=I_{\boldsymbol{p}} G \ln \frac{M_{p}}{M_{0}} \tag{1.1}
\end{equation*}
$$

to determine the mass ratio $M_{p} / M_{0}$ for each maneuver, taking the total mass of the previous maneuver to be the payload mass of the next maneuver. Program PMEFDMAI.BAS (Propellant Mass Estimate From Dry Mass And Isp) cycles through various dry masses and specific impulses and estimates the mass of propellant required as a function of these parameters (Appendix 1 ).

### 1.2 Trajectory Optimization

### 1.2.1 Introduction

Allowing the thrust vector to change not only its magnitude, but its direction was the next logical step in our analysis. This optimization step required the use of a low-thrust, orbital calculation program published by NASA Ames in 1973. QUICKTOP, and its companion
program CHEB3, are both FORTRAN source codes that compute the optimum low-thrust interplanetary trajectories with variable thrust vector programs.

The QUICKTOP input file contains numerous variables that may be specified. Through proper use of this data file, it is possible to simulate a solar powered-electric propulsion departure from orbit to any given destination planet. The initial and destination orbits can be specified, as well as many other parameters and options. From this data, QUICKTOP generates a complete analysis of the trajectory. The program contains provisions to account for the loss of power available for propulsion due to increasing distance from the sun.

The output variables of primary importance in this study were propellant required (and thus, IMLEO), thruster firing time, loiter time at Mars, refit time at Earth, and overall mission times. Loiter and refit times were addressed using the TIMEOPT code which is described below (see Appendix 1 for details).

Based on propulsion system optimization, a specific impulse of 10,300 seconds was selected. This corresponded to a required power of 3.5 megawatts. These parameters were held constant throughout trajectory calculations.

Mission trajectories were generated by inputing a large variety of heliocentric travel times. The return trip from Mars to Earth and the outward trip from Earth to Mars were analyzed independently. The input data were manipulated so that the payload each way was zero and the IMLEO was estimated. The resulting negative value for payload represented a first estimate for the propellant required for that trajectory. After selecting an optimum trajectory, the solution was iterated until the output was consistent.

### 1.2.2 Propellant and Thrusting Considerations

For each of the input HTO travel times that converged, QUICKTOP produced the required
propellant, thrust time, as well as spiral escape and capture times. Twenty outward and twenty return trajectories were created and compared on this basis. The results are shown in Figures 1.1 through 1.4 which indicate propellant mass required and thrusting time as a function of travel time. Figures 1.1 and 1.2 refer to the outward trajectory, while Figures 1.3 and 1.4 refer to the return trip. Figure 1.1 presents the propellant mass required for the outward trajectory. As can be seen, propellant requirements tend to increase as HTO time decreases. Some peaks and valleys occur, indicating that there is no absolute relationship involved or single thrust program that will result in the optimum trajectory. Figure 1.4 shows the variation of the required thrusting time for the retum trip and indicated a minimum of about 5500 hr . The results obtained allowed us to select the final trajectory and mission times. Shorter trip times result in generally increased propellant requirements. Lower trip times often did not converge to solutions at all.

### 1.2.3 Time Considerations

Having compared a wide range of trajectories on the basis of HTO travel time, propellant required, and thrusting time, the synodic considerations of meeting with the target planet were examined. The escape time, capture time, HTO travel time, and the change in orbital true anomaly resulting from the HTO transfer, generated by QUICKTOP, were input into TIMEOPT, a program that calculates the required phasing of the planets for successful rendezvous. In addition, TIMEOPT also computes the necessary loiter time at Mars and refit time at Earth to match these phasing ("launch window") requirements. This resulted in approximately 400 possible round trip trajectories. These round trip trajectories were then evaluated in terms of the following criteria:

1. Total mission time. The total time for three consecutive missions is not allowed to exceed $12+$ years. This was the minimum time calculated for three missions; other possibilities were 18+ and 23+ years. These times were considered too long to support a reasonable mission.


Time (days)
Figure 1.1: Propellant Required (Out) vs. Trip Time

## Thrust Duration vs. Outward Trip Time



Figure 1.2: Thrust Duration vs. Outward Trip Time

## Propellant Required (Ret) vs. Trip Time



Figure 1.3: Propellant Required (Ret.) vs. Trip Time


Figure 1.4: Thrust Duration vs. Inward Trip Time
2. Loiter time at Mars. This time was not to exceed 70 days or fall below 20 days. A minimum delay at Mars is desirable, since the craft cannot be serviced or overhauled realistically in Mars orbit. A delay of at least 20 days was specified to allow for unexpected delays in off-loading cargo or unanticipated trajectory changes. Shorter loiter times could result in the missing of a launch window should the craft be delayed. This would cause a very extended, costly stay at the red planet. These times were arbitrarily selected and could easily be changed in the optimization process.
3. Refit time at Earth. A maximum, or at least very long, time for refit at Earth was desired. This would obviously allow more time for loading of cargo and refit/repair operations. This would also maximize the fraction of time that the craft was in Earth orbit or Earth's vicinity, where unexpected problems or malfunctions are more easily corrected.
4. Propellant and Thrusting Time Considerations. Plots 1.1 through 1.4 were used to evaluate the propellant consumption and thrusting time required for each round trip trajectory. Minimums were sought in each of these parameters; low propellant mass translates into low IMLEO, and short thruster firing time results in reduced degradation of engines (and lower redundancy requirements) and less extensive (and costly) refit time to replace degraded thrusters. Some thruster refit will invariably be necessary, however. The result of this analysis was a matrix of data concerning the possible round trip trajectories. Many of the candidates were eliminated by the time requirements listed above. A few of the remaining trajectories could be eliminated because of extensive propellant or thrusting time requirements. This still left about a dozen or so trajectories of which there was no clear choice.

At this point, it was decided that, given relatively similar thrusting times and propellant requirements, the longest Earth refit time possible should be selected, within the bounds of the parameters listed above. The resulting trajectory is summarized in Table 1.2; and illustrated graphically in Figures 1.5 and 1.6.

## Outward Low-Thrust Trajectory



$$
+D_{\text {csition }}
$$

Thrust Vestor

Figure 1.5: - Outward Heliocentric Transfer Orbit Trajectory


Figure 1.6: Return Heliocentric Transfer Orbit Trajectory

Table 1.2 Mission Architecture Summary

| Phase | Phase Time <br> (days) | Earth Position <br> (rads) | Mars Position <br> (rads) | Mission Time <br> (days) |
| :---: | :---: | :---: | :---: | :---: |
| Start | 0 | -2.62 | 0.183 | 0 |
| Earth escape | 152 | 0.00 | 1.57 | 152 |
| Out HTO | 650 | 11.19 | 7.52 | 802 |
| Mars capture | 108 | 13.04 | 8.51 | 910 |
| Loiter @ Mars | 60 | 14.08 | 9.06 | 970 |
| Mars escape | 38 | 14.73 | 9.41 | 1008 |
| Return HTO | 260 | 19.21 | 11.79 | 1268 |
| Earth Capture | 84 | 20.65 | 12.55 | 1352 |

Planet positions defined from arbitrary reference position of Earth at escape defined as 0.00 radians.

Refit time until next launch window $=210$ days $=7$ months
Total mission duration (3 round trips) $=12.2$ years

### 1.2.4 The Optimized Trajectory

The selected outward trajectory requires 910 days from the beginning of the Earth spiral to the completion of Mars capture. The propellant required for this trajectory is $13,258 \mathrm{~kg}$ (expressed as a negative payload in the optimization process). Thruster time for this trajectory is 8766 hrs . The IMLEO was taken to be roughly $130,000 \mathrm{~kg}$. The Earth escape spiral time is 152 days, the heliocentric transfer time is 650 days and the Mars capture time is 108 days.

The position and orientation of the craft as a function of the time in the HTO for the outward trip is graphically represented in Figure 1.5. For this study, the $z$ coordinate of motion (out of the ecliptic plane) was ignored because it was small relative to the x and y components. The data is given in radius-angle format. The angle (THETA) was very important, since the
change in this angle for the HTO trajectory was required to compute loiter and refit times.
For the outward trip, the craft coasts for about 100 days after Earth escape, then thrusts for about 80 days in a gradually rotating direction and decreasing magnitude. The craft then coasts outside of Mars' orbit, then rendezvous after crossing its HTO aphelion. Figure 1.5 indicates the magnitude and orientation of the thrust vector during the HTO trajectory.

Figure 1.6 describes the return trajectory. Interestingly, the return trajectory consists of an acceleration and a deceleration rather than a single coast-thrust-coast program. This probably is necessary because the return trip was a relatively short trajectory (relative to the other return trajectories), whereas the outward trajectory was one of the longer HTO times considered.

A few additional comments concerning Figures 1.5 and 1.6 are in order. These two plots represent the heliocentric phase of the outward and inward trajectories, respectively. In Figure 1.6, the time between each of the position data points is approximately 20.5 days. The maximum acceleration required in this HTO is about $3.6 \mathrm{AU} / \mathrm{yr}^{2}$, or about $0.00054 \mathrm{~m} / \mathrm{sec}^{2}$.

In Figure 1.6 (return trajectory), the time interval between position data points in about 11.6 days. The maximum acceleration required in this HTO is about $5.6 \mathrm{AU} / \mathrm{yr}^{2}$ or 0.00084 $\mathrm{m} / \mathrm{sec}^{2}$. Both of these accelerations are well within the capabilities of the electric propulsion system.

### 1.3 Conclusions

In conclusion, a general comment concerning the optimization of low-thrust trajectories is in order. For a wide range of trip times selected, propellant masses and thrusting times required did not vary significantly. Only several extreme trajectory cases could be eliminated because of excessive propellant or thrusting time requirements. Hence, the mission trajectory is driven primarily by the desired loiter, refit, and total mission time parameters.

This variation is possible because the nature of low thrust propulsion makes trajectory adjustments relatively easy compared to high thrust systems. Since the delta-V is applied gradually over a long thrusting arc, small corrections or changes can be made at relatively little cost in propellant. This characteristic of low thrust propulsion is very desirable; it enables the same system to be used for very different missions without a drastic increase in propellant mass. This gives a wide margin for error and unexpected delays in the mission execution. If a malfunction occurs or the craft is unexpectedly delayed, it is very likely that a suitable trajectory could be found that would fulfill the mission, or at least salvage the craft.

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## SECTION 2

## Propulsion Systems

## Nomenclature

A - area<br>C - exhaust velocity<br>D - diameter of screen grid<br>d - grid spacing<br>F,F' - variables in Isp optimization<br>$\mathrm{J}_{\mathrm{b}}$ - beam current<br>$\mathrm{J}_{\mathrm{n}}$ - neutralizer current<br>$k$ - thermal conductivity<br>$\mathrm{m}_{\mathrm{L}}$ - engine and propellant mass<br>$m_{0}$ - total craft mass<br>$\mathrm{m}_{0}$ - mass of propellant atom<br>MW - molecular weight<br>$P_{b}$ - beam power<br>$P_{n}$ - neutralizer power<br>$P_{\text {tox }}$ - total power<br>$\mathrm{q}_{\text {vap }}$ - heat of vaporization<br>$\mathrm{S}_{\mathrm{d}}$ - chamber specific loss level<br>T-thrust<br>$\mathrm{V}_{\mathrm{b}}$ - beam voltage<br>$\eta_{i}$ - ionization efficiency<br>$\eta_{\text {. }}$ - electrical efficiency<br>$\eta_{p}$ - propellant efficiency<br>$\eta_{\text {loo }}$ - total efficiency<br>$\Gamma^{\prime \prime} / \Gamma^{\prime}$ - ratio double to single ions<br>$\delta$ - neutralizer voltage<br>$\gamma$ - thrust correction factor<br>$\Delta \mathbf{v}$ - velocity increment<br>$\alpha$ - specific power

### 2.1 Introduction

The primary difference with this mission as opposed to many other space missions is the requirement of electric propulsion as the sole means of propulsion. The key jusdification for the electric propulsion is its high specific impulse and ability to provide very high payload to craft mass ratio. These factors may significantly reduce the transportation system mass and cost and make electric propulsion reasonable for long-term interplanetary missions.

### 2.2 Preliminary Concepts

There are three general types of electric thrusters: electrothermal, electromagnetic, and electrostatic. Resistojets and arcjets are two types of electrothermal thrusters. Each of these engines electrically heats the propellant and expands it through a nozzle to produce thrust. These engines are characterized by a relatively low (less than $1,500 \mathrm{sec}$.) specific impulse. Electromagnetic engines use an electromagnetic field to accelerate particles and produce thrust. The electrostatic engine accelerates ions through a voltage difference to produce thrust [5].

The electrostatic ion thruster was chosen for two critical reasons. First, it can reach very high specific impulses. Then, the ion thruster has the most fully characterized technology. Much research has been done on ion thrusters, and it is the most likely candidate for use in a deep space mission in the post 2010 time frame. An important feature of these thrusters is that they are scaleable, which means that they may be increased in size and power from smaller thrusters which presently exist. This may allow a significant reduction of lab research and experimentation.

Once the ion thruster is chosen; the task of deciding on a propellant still remains. As speciried in NASA's criteria, the recommended choice of propellant is argon; however, to be complete, analysis of this recommendation is necessary. From recent evaluations, the only two
propellants under serious consideration for such a long term mission are argon and xenon, but, to be complete, mercury will also be included. One case study gave estimates of the benefits and drawbacks of each propellant for a specific system (Figure 2.1). Because of the higher specific impulse that can be provided by an argon thruster compared to a xenon thruster the argon will consume more power, per unit thrust, than a xenon (or mercury) engine. This, as one might assume, leads to a larger dry system mass for argon due to the increased power requirement. However, higher specific impulse has an opposite effect on the propellant mass, causing that the latter decreases with an increase in specific impulse, which gives preference to argon. The propellant mass saved by using argon more than compensates for that needed by the increased power supply and therefore may give better total system performance.[1]

An evaluation was made for the characteristics of our system with both argon and xenon, and because of the length of the mission, it can be assumed that the relatively large propellant masses are the main contributing factor in choosing the propellant. In this case argon gives lower propellant mass and therefore a lower overall mass. Xenon is unable to achieve the same low propellant masses at the higher range of specific impulses because the beam voltages at high specific impulse may be beyond the technology of cathode construction. For example, for xenon to run at an specific impulse of 10,300 seconds it would require a beam voltage of just over 6000 V. Xenon also has the additional burden of being very expensive due to its limited supply. Because this mission would require large propellant mass, this cost must play an important role in the selection process. In contrast to xenon, argon is inexpensive, abundant, and easy to handle. Based on these characteristics the original choice of argon seems justified.

Propellant Comparison

Fligure 2.10

Serioe 1 Serios 2 Serios 3

### 2.3 Thruster Optimization and Configuration

From information received from trajectory analysis it was decided that an initial total thrust of 60 Newtons is reasonable to allow an adequate trip time. The power decrease that will occur on the trip to Mars is estimated by Figure 2.2, which shows that the power at Mars will be approximately $40 \%$ of the original power at Earth. This will cut the thrust at Mars to nearly 25 Newtons. This thrust decrease during the trip to Mars has been taken into account when estimating trip time.

A 5 Newton thruster was selected as a suitable thruster size. When comparing thrusters of various sizes while keeping the total thrust constant, the craft characteristics (e.g. mass, power) remained essentially unchanged. Therefore, the choice of the 5 Newton thruster was based on a compromise between flexibility and unnecessary complication. This selection would give 12 operating thrusters at Earth and 5 at Mars. Obviously, if more thrusters were used, each with a smaller individual thrust, then the flexibility or throttling capabilities would increase. However, a large number of thrusters in operation may over-complicate the power processing and distribution systems to such an extent that the extra mass needed would become cumbersome. Therefore, twelve engines were chosen for the initial configuration to allow for enough flexibility without over-complication of the power processing systems. For the case of engine malfunction a redundancy of 5 engines is assumed, bringing the total number of thrusters to 17 . This, obviously, does not increase the power requirements but only adds engine mass to the ship. The 17 thrusters also allow for a flexible thruster firing configuration which allows thrusters to be cut off one at a time, thereby saving on throttling demands.

Specific impulse optimization of the 5 Newton thruster was accomplished through the use of two methods. The first method made use of the trajectory program which we wrote, through


Figure 2.2: Power Provided


Figure 2.3: Beam Current vs. Isp
$\mathrm{T}=5 \mathrm{~N}$
which a range of specific impulses was run in an attempt to find the lowest initial craft mass, while still insuring a reasonable spiral out escape time. The lowest initial craft mass was found to be about 143 metric tons which corresponded to an optimum specific impulse of 10300 seconds. The second method also optimized by minimizing initial craft mass ( $m_{0}$ ), but accomplished it through a process of solving equations for optimum specific impulse by a Newtonian iterative process [8].

$$
\begin{aligned}
& F=1-\alpha / 2-e^{-\alpha}-m_{1} / m_{0}\left(e^{\alpha}-1-\alpha / 2\right) \\
& F=-0.5+e^{-\alpha}-m_{/} / m_{0}\left(e^{\alpha}-0.5\right) \\
& \alpha^{j+1}=\left(\alpha-F / F^{\prime}\right)^{j} \\
& C=\Delta v / \alpha \\
& I_{\text {op }}=C / 9.807
\end{aligned}
$$

The algorithm was iterated through by updating the initial guess for the alpha value ( $\alpha_{0}=1-m_{\sqrt{ }} / m_{0}$ ) until the desired accuracy of Isp was obtained. This algorithm gives an optimum Isp of 10,289 seconds, which so closely matched our first method that the number is assumed correct. Later investigations through the use of a more complex trajectory analysis proved to show a decrease in total mass to 131 mT (for the assumed performance of the power supply), however, an Isp optimization proved to be too time consuming, therefore the previously calculated Isp was kept. Once the Isp was chosen the parameters of the thruster could be found as shown in Appendix 2.1 [13]. Thruster beam current, beam voltage, total power required and efficiency were analyzed for a range of Isp and are shown in Figures 2.3 through 2.6. Total power required by the engines is 3.3 MWe at Earth and 1.32 MWe at Mars.

To estimate thruster mass, a linear scaling based on thrust to mass ratio for smaller engines was done to achieve a mass of approximately 86 kilograms (Figure 2.7). Inspection of most


Figure 2.4: Beam Voltage vs. Isp $\mathrm{T}=5 \mathrm{~N}$


Figure 2.5: Thruster Power vs. Isp


Figure 2.6: Total Efficiency vs. Isp $\mathrm{T}=5 \mathrm{~N}$


Figure 2.7: Thruster Mass Scaling
recent information from NASA led us to a final estimate of 75 kg . The diameter was found using an equation for the thrust to area ratio:

$$
\begin{aligned}
& \mathrm{T} / \mathrm{A}=2 / 9 \varepsilon_{0} \mathrm{~g}^{4}(\mathrm{~m} / \mathrm{e} / \mathrm{d})^{2} \mathrm{Isp}{ }^{4} \\
& \text { where: } \\
& \varepsilon_{0}=8.85 \times 10_{-12} \mathrm{~F} / \mathrm{m} \text { and } \\
& \mathrm{D}=2(\mathrm{~A} / \mathrm{pi})^{0.5}
\end{aligned}
$$

The parameter d is the grid spacing which was taken as 0.01 meters. This value was chosen because it gives an acceptable beam current density for a long thruster lifetime. For a 5 N engine, a beam diameter of 130.6 cm was obtained. The diameter including thruster casing is about 150 cm . Overall thruster dimensions are given in Figure 2.8a and the propulsion module configuration can be seen in Figure 2.8b. The main characteristics of the thruster are presented in Table 2.1.

Table 2.1 Thruster Characteristics

Thrust
Specific Impulse
Propellant mass flow rate
Beam voltage
Beam current
Beam power
Discharge power
Neutralizer power
Total power
Electrical efficiency
Propellant efficiency
Total efficiency
Thruster mass

5 N
10,300 sec
$4.9 \times 10^{-5} \mathrm{~kg} / \mathrm{s}$ 2424 V
104.2 A
252.6 A
20.84 kw
1.69 kw

275 kw
0.9185
0.8619
0.714

75 kg


Figure 2.8 a: Electrostatic Thruster


Figure 2.8 b: Propulsion Module

### 2.4 Thruster Degradation and Lifetime

The lifetime of an ion thruster is limited by sputtering which occurs when surface ions are emitted by collision of incident ions. This causes erosion of discharge chamber components. The lifetime is limited specifically by the erosion of the screen grid. Lifetime may be defined as the time necessary for the screen to erode to half-thickness [4].

The erosion rate is highest at the center of the grid because of the high plasma density and the high ratio of doubly to singly charged ions near the centerline, but this problem may be alleviated by the use of the ring-cusp magnetic configuration, which lowers the concentration of double charged ions at the centerline [12]. An apparent solution to the wearing of the screen grid would be to just thicken it, but this causes unacceptable decreases in thruster efficiency. Some experiments have been done with adding small amounts of nitrogen to the discharge chamber when using xenon propellant. This has increased projected lifetimes from two to four times over what was achieved without using nitrogen [10]. The effect of adding nitrogen or some other life increasing agent to argon remains to be seen. As beam current density increases the lifetime of the thruster decreases. A xenon engine with a $0.00232 \mathrm{~A} / \mathrm{cm}^{2}$ beam current density was shown to have a lifetime in excess of 25,000 hours. When the density was increases 3.5 times the lifetime decreased to 14,000 hours [4]. This was without the introduction of nitrogen. For the 5 N thruster, beam current density of $0.00617 \mathrm{~A} / \mathrm{cm}^{2}$ is an acceptable value. If the nitrogen can increase the lifetime of a thruster by a factor of two, then a projected lifetime of 25,000 hours may easily be achieved. Since it is estimated from the amount of propellant needed for the round trip and the mass flow rate of the thrusters that the lifetime of the thrusters (without redundancy) must be about 14,500 hours, a lifetime of 25,000 hours would give several thousand hours above the necessary thrust time.

### 2.5 Propellant Tank

The propellant tank chosen is of cylindrical design with spherical ends (Figure 2.9). The tank has two shells. The outer is a vapor cooled, thin aluminum shield, while the inner serves as a pressure vessel. Trajectory estimates give a minimum propellant mass of $21,040 \mathrm{~kg}$. A $15 \%$ contingency and reserve for auxiliary propulsion is added to this mass giving a final propellant mass of $24,200 \mathrm{~kg}$. The argon propellant is to be stored cryogenically at a pressure of 1 atm and a temperature of 87 K . From these parameters the fluid density is found to be $1420 \mathrm{~kg} / \mathrm{m}^{3}$ which gives a tank volume of 17.04 cubic meters. An extra 0.46 cubic meters is added to allow for fluid expansion upon vaporization, giving a total volume of $17.5 \mathrm{~m}^{3}$.

The tank pressure vessel is composed of pre-approved cryogenic alloy, Ti-5 Al-2.5 Sn. The tank wall thickness calculated as a function of pressure, tank radius and allowable stress ( 0.15 mm ) is near or below the thickness that can be reliably manufactured. The tanks will be fabricated so that the thickness will vary from a nominal 0.4 mm to about 3 mm at the attachment points of the external supporting straps and at the locations of internal heaters and vapor acquisition devices. The tanks will be designed for space environment and will be filled in orbit. Because of the long propellant storage time in both near Earth and interplanetary space, the tank requires an adequate meteoroid/debris protection as well as thermal insulation. The latter is required to reduce boil-off during long waiting (no thrust) periods. The protective and insulating shield will consist of a surface coated polyamide face sheet, low density foam spacer with intermediate Kevlar reinforcement and multi-layer insulation (MLI) made of aluminized Mylar radiation films separated by Dacron net spacers. Assuming MLI thermal conductivity (k)


Figure 2.9: Cryogenic Propellant Tank


Figure 2.10 Propellant Distribution
of $0.0175 \mathrm{~mW} / \mathrm{m}-\mathrm{K}$ and 8 cm thick (d) insulation, the heat penetrating the MLI will be:

$$
Q=(k A / d)\left(T_{t s}-T_{v}\right)
$$

where
$A=$ tank surface
$T_{t s}=$ tank surface temperature
$\mathrm{T}_{\mathrm{v}}=$ vapor temp
Near LEO, $\mathrm{T}_{\mathrm{t}}$ will have a value of around 330 K , and Q is 1.8 W . Assuming that additional heat conducted through the straps supporting the tank, propellant lines and electrical lines is less than 2 W , the total thermal energy reaching propellant is less than 4 W . When the thrusters are not in operation, the vaporized argon will be re-liquified by a Stirling refrigerator.

When all 12 engines are thrusting near Earth, the propellant mass flow rate is $5.88 \times 10^{-4}$ $\mathrm{kg} / \mathrm{s}$ and the heat required to vaporize the liquid argon is

$$
\mathrm{Q}=\mathrm{m} \mathrm{q}_{\text {vap }}=5.88 \times 10^{-4} \times 160 \times 10^{3}=94 \mathrm{~W}
$$

This heat, only slightly reduced by the amount penetrating insulation, must be provided by the heating coils inside the internal tank shell. A very simplified diagram of propellant distribution is shown in Figure 2.10. The mass summary of the tank system is given in Table 2.2.

## Table 2.2 Mass Summary

Tank ..... 230 kg
Meteoroid/Debris Shield/Insulation ..... 210
Electrical heaters, sensors, controls ..... 90
Refrigeration unit ..... 130
Propellant lines, valves ..... 160

### 2.6 Power Processing

The power processing is an important part of the total SEP system. This system is responsible for conditioning the power generated by the solar arrays and distributing it among the craft systems. The thrusters require the majority of the power and therefore have a separate power distribution center contained within the engine system itself. The distribution center consists of one beam and discharge power filter for each thruster. These filters are used to modify any current fluctuations that might cause inefficiencies in thruster operation leading to engine degradation [9]. These filters are grouped together in sets of three in the thruster modularization scheme; the exception being the two end thrusters which are modularized separately. All processing units are located near the center of mass of the ship in order to keep transmission line length to a minimum. The propulsion power processing system consists of the low voltage discharge and neutralizer power supplies and of DC-DC converter units for the grid power supply which requires the majority of the power input. It is assumed that the power processing unit has an overall efficiency of 95 percent and operates at 600 K . After power processing, the voltage and current supplied to individual thrusters are 2424 volts and 130 Amp.

The remaining generated power is designated to all other on-board systems. Since this power is to be conditioned for the ships operating systems it must conditioned between the standard voltages of 100 to 200 V . The systems include navigation, guidance, control, and communication. A small amount of power must be rerouted to the thruster control mechanisms and then be used when needed for thruster shutdown or control in the event of a malfunction [9]. A block diagram of the power breakdown can be seen in Figure 2.11.


Figure 2.11: Power Distribution Diagram

### 2.7 Thermal Control

The heat generated by the engines and power processing units must be dissipated to space by a radiator. The latter is sized so as to be able to reject up to 11 percent of the input power, or about 370 kW . A radiator/heat pipe system was chosen as the simplest and yet most effective way of removing this heat into space. The heat pipe system will operate at a temperature of about 600 K and use stainless steel (or nickel) for pipe vessel and wicking material and potassium as a working fluid. The main design parameters of the radiators are the area and materials. Since a large amount of heat must be radiated, the main concern of the radiator design is to keep the mass down while maintaining high emissivity and conductance. The radiator will be constructed of titanium alloy providing good high temperature operation (see Appendix 2.2 for calculation). Two modularized radiator panels, each sized at $40 \mathrm{~m}^{2}$, will be mounted on the two opposite sides of the thruster module, facing directions perpendicular to the rotation axis of the solar arrays. Radiator modularization can be accomplished alongside the thruster modularization into one large component. In this way the radiator array can be broken down in such a way that there should be no problem fitting aboard the shuttle cargo bay. The main characteristics of the thermal control system are given in Table 2.3.

Table 2.3: Characteristics of the Thermal Control System

| Heat rate capacity | 370 kW |
| :--- | :--- |
| Heat pipe fluid | potassium |
| Radiator temperature | 600 K |
| Total radiator area | $80 \mathrm{~m}^{2}$ |
| Total mass | 1200 kg |

### 2.8 Propulsion Module Structure

The propellant module is 11.2 meters wide by 4.8 meters tall by 5.65 meters deep. The system is modularized in groups of three thrusters, except for the two outside modules, which only have one thruster. Each module also has the power processing unit, one seventh of the radiator, and its own structure. An overall schematic diagram of the configuration may be seen in Figure 2.8b. All systems will fit as is or equipped with a modular ability, such as the engine modules, within the cargo bay of the Shuttle $C$. The mass breakdown of the propulsion system is given in Table 2.4.

Table 2.4: Propulsion System Overall Specifications and Mass Breakdown

$$
\text { Number of engines (no redundancy) } 12
$$

Number of engines (including redundancy) ..... 17
Engine beam diameter ..... 130 cm
Specific impulse ..... $10,300 \mathrm{sec}$
Engine thrust ..... 5 N
Engine input power ..... 275 KW
Power processing efficiency ..... 0.95
Propulsion system efficiency ..... 0.714
Engine mass ..... 75 kg
Gimbal mass ..... 15
Power processing unit, per engine ..... 330
Thermal control system ..... 1200Interface module and housingstructure, lines, controls800
Propulsion system mass ..... 9140 kg

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## SECTION 3

## Solar Power

## Nomenclature

| AMO | Air mass zero |
| :--- | :--- |
| GaAs | Gallium Arsenide |
| GaAsP | Gallium Arsenic Phosphide |
| GaP | Gallium Phosphide |
| InAsP | Indium Arsenic Phosphide |
| InP | Indium Phosphide |
| Si | Silicon |

### 3.1 Introduction

In researching standard photovoltaic cells, it was discovered that predictions for new technology and advancements in these materials are very encouraging. Efficiencies which have previously been in the 10 to 15 percent range have skyrocketed to above 30 percent. Assuming technology remains on its current course of advancement, photovoltaic cell will reach outputs of over $300 \mathrm{~W} / \mathrm{kg}$. Solar cell material tends to degrade when left exposed to the space environment without any protection and thus efficiencies decrease. Concentrator arrangements provide this protection of the raw cell material from the hostile space environment, particularly in the Van Allen radiation belts.

Indium phosphide is a leading candidate for cell materials used in space applications due to its high efficiency and radiation resistivity. Although only achieving $19 \%$ efficiency in the laboratory, there is great reason to expect AMO efficiencies above $20 \%$ based on experience with gallium arsenide and silicon [1]. With respect to exposure to electron and proton fluence, InP out-performs GaAs and ultra thin Si [8], although both of these latter materials are considered to be good radiation resistant materials with slightly higher specific powers and efficiencies than
$\operatorname{In} P[1,9]$. Another impressive property of $\operatorname{In} P$ is its ability to essentially completely recover after exposure to a fluence of $10^{17}$ electrons by heating the cell at $115 \cdot \mathrm{C}[4]$. InP, therefore, possesses the ability for supplying significantly more end of life array output power than Si of GaAs in high radiation environments [8].

### 3.2 Solar Power Options

Flat array and concentrator systems were two photovoltaic options considered for the solar power supply. Both options utilize a InP cell material due to the properties and abilities mentioned above. The cell was also assumed to operate at $115^{\circ} \mathrm{C}$ in order to take advantage of indium phosphide's ability to complctely anneal at this temperature. The cell for the concentrator system, however, will be a multi-stacked cell using InP as only one of the cell layers since the structure will help to protect the cell.

### 3.2.1 Concentrator Array

The concentrator option was prepared using Fresnel lens concentrators, selected for its superior projected optical efficiency of $96 \%$ [1]. This efficiency compares well with other concentrator systems that are based on mirrors or internal reflection. The mirror concentrators of the TRW Cassegrainian system, for example, projected lower future efficiencies in the $80 \%$ range.

### 3.2.1.1 Array Parameters

Taking an operating temperature of $115^{\circ} \mathrm{C}$ into consideration, the module efficiency at AMO was assumed to be $27.7 \%$. The module has the dimensions $6 \mathrm{~cm} \times 6 \mathrm{~cm}$ and a specific mass ratio of $1 \mathrm{~kg} / \mathrm{m}^{2}$. As seen in the calculations appearing in Appendix 3.1, the specific power of the module was found to be $320.8 \mathrm{~W} / \mathrm{m}^{2}$. This specific power lead to the total array area of
$10,910 \mathrm{~m}$ and to the similar array mass of $10,910 \mathrm{~kg}$. This mass value includes the mass of the radiators required to keep the module temperature at a constant value. The dimensions of the hexagonal panels are given in Appendix 3.1, as well.

### 3.2.1.2 Wiring Configuration

The following wiring configuration for the concentrator system was limited by a maximum voltage of 200 V , a maximum current of 20 Amps and an open current cell voltage of 945 mV . The maximum number of cells connected in series was calculated as 211 (Appendix 3.2). From this value, the length of each series, 12 meters, and the resistance through each series wire of Aluminum gage \#7, $0.032 \Omega$, was determined. Along with the calculated current of 1.155 A , the power loss per series was found to be 0.0429 W , resulting in a total power loss of 80.2 kW . The total length and mass of the series and parallel Al gauge \#7 wires were estimated to be 286.5 m and $8,225 \mathrm{~kg}$, respectively.

### 3.2.2 Flat Panel Array

Next, a flat solar array option was developed. Since this array must survive three long round trip missions without large power losses, Indium Phosphide was chosen as the material for the solar cells for reasons mentioned earlier in the discussion on cell materials.

### 3.2.3 Array Specifications at Earth

An achievable cell efficiency of $\mathbf{2 1 . 7 \%}$ was assumed from a consensus of participants of the 1987 work-shop on InP at the NASA SPRAT conference [7]. Using this efficiency, the specific power (assumed at $25^{\circ} \mathrm{C}$ ) was calculated as $255 \mathrm{~W} / \mathrm{m}^{2}$. Assuming a steady state operating temperature of $115 \cdot \mathrm{C}$, the specific power and efficiency dropped to $198.5 \mathrm{~W} / \mathrm{m}^{2}$ and $17.6 \%$, respectively. These values along with the total array power needed ( 3.5 MW ) lead to a total
array area of $17,632 \mathrm{~m}^{2}$ and total weight of the array (not including supporting structure) of $14,987 \mathrm{~kg}$. The flat panel calculations and dimensions are listed in Appendix 3.3.

### 3.2.4 Array Heat Transfer

The steady state temperature used for the calculated values stated above was chosen for continuous annealing at the orbit around Earth. The annealing will help prevent long term power losses from the radiation encountered within the Earth orbits. Assuming a constant temperature and radiation heat transfers out into free space, heat transfer calculations (Appendix 3.4) were performed for two different scenarios. The first scenario assumed continuous annealing for the entire mission and the second scenario assumed continuous annealing only for near Earth trajectories. The assumption of continuous annealing over the entire mission was found to require a high temperature of $206 \cdot \mathrm{C}$ at Earth. A better solution was found by assuming $115 \cdot \mathrm{C}$ near Earth, leading to about $42 \cdot \mathrm{C}$ at Mars. This will allow for annealing in the radiation belts encountered while spiraling out of Earth's influence. Although this does not allow for annealing at Mars, the lower operating temperature will give better cell efficiency and specific power at Mars.

### 3.2.5 Array Wiring Connections

To obtain the high cell efficiency of $21.7 \%$, the cell voltage was assumed to equal 945 mV . For a $2 \times 2 \mathrm{~cm}$ cell, the power per cell, 0.120 W , lead to the power of each series wire equalling 25.3 W and to the current found in each series connection equalling 0.126 Amps . Using a similar wiring scheme to the one formulated for the concentrator system and an Aluminum gauge \#7 wire, the total wire mass was estimated to equal $11,100 \mathrm{~kg}$, bringing the mass of array and wiring to $26,087 \mathrm{~kg}$. Calculations are given in Appendix 3.5 .

### 3.3 Selection of Solar Power System

Initially, the concentrator system was chosen for several reasons including the higher specific power achievable and the added radiation protection which the lens provided for the solar cell. Through comparison of the concentrator system and the flat array option, justification for the use of concentrator systems over flat arrays can be seen.

Table 3.1 compares some assumed and calculated values for these two solar power options. The specific power at $115^{\circ} \mathrm{C}$ for the concentrator is about $62 \%$ greater than that achieved by the flat plate; therefore, the area of the concentrator array is $62 \%$ less than the flat plate area. The mass of the concentrator array is also much lower than that of the flat plate array. Supporting structure considerations also lean in favor of the concentrators which themselves provide a great deal more rigidity. For the supporting structure, the enlarged area of the flat system will increase the array support mass. Linearly scaling for characteristic length of array, a mass increase of about $17 \%$ for array support was found. Comparing the value of the flat panel and wiring mass to that of the concentrator and wiring mass, the flat array combined mass was found to be 6,952 kg more than the concentrator system value. Finally, the power losses calculated for the flat array are over 2.5 times greater than concentrator system losses.

Table 3.1 Comparison of Flat Array and Concentrator System

| Attribute | Flat Array | Concentrator |
| :--- | ---: | ---: |
| Specific Power at $115^{\circ}$ | $198.5 \mathrm{~W} / \mathrm{m}^{2}$ | $320.8 \mathrm{~W} / \mathrm{m}^{2}$ |
| Area of Array | $17,632 \mathrm{~m}^{2}$ | $10,910 \mathrm{~m}^{2}$ |
| Characteristic Length (d) | 93.0 m | 79.4 m |
| Masses: Array | $14,987 \mathrm{~kg}$ | $10,910 \mathrm{~kg}$ |
| Wiring | $11,100 \mathrm{~kg}$ | $8,225 \mathrm{~kg}$ |
| Power Loss of Wiring | $297,850 \mathrm{~W}$ | $80,170 \mathrm{~W}$ |

While the comparison above is preliminary and based on highly simplified assumptions, it does indicate that a concentrator system used in multi-megawatt applications may greatly outweigh the performance of a flat array power system.

### 3.4 Selected Solar Power System Design

### 3.4.1 Solar Cell Configuration

Combining the materials mentioned in the Introduction and stacking them in a configuration provides added efficiencies to the top cell. Lower cells absorb a different bandgap (wavelength) of light which adds efficiency aumbers to the top cell. Cells stacked in this configuration have efficiencies much higher than 30 percent with efficiencies under concentration predicted above 40 percent. Figure 3.1 shows the three solar cell stack selected for our array [5]. This cell will use a six-terminal wiring configuration to minimize current mismatch losses which may be caused by different degrees of radiation damage to the three solar cells. The GaAsP on GaP top cell which determines the performance of the triple stack is the most advanced wide bandgap cell currently under development. The state-of-the-art GaA s cell will serve as the middle cell. The InAsP bottom cell is the best developed, and, even when "immature" can only improve the total conversion efficiency by scavenging the low energy photons from the top two solar cells. A theoretical efficiency limit of this triple solar cell stack is $42.8 \%$ at AMO, and a practical system efficiency of $35.9 \%$ is predicted by Barnett, Trumble, et al [3]. In application to this mission, an efficiency of $27 \%$ is chosen to account for the high operating temperature of about $150 \cdot \mathrm{C}$ and the radiation damage occurring during the three Mars missions.


Figure 3.1: Triple Stacked Solar Cell (Barnett and Trumble)


Figure 3.2: Fresnel Lens Module

### 3.4.2 Fresnel Lens Concentrator Array

The solar array is made up of many small modules. Each small module consists of a Fresnel lens dome that concentrates the incident sunlight on the solar cell, a honeycomb supporting structure, a multistacked cell, and a back radiator surface to dissipate excess heat produced by the cell.

The Fresnel lens dome consists of annular prisms made of silicone that use refraction to concentrate light on the cell. Optical efficiencies of $92 \%$ have been achieved and with the addition of anti-reflective coating to the lens, the already mentioned projected optical efficiency of $96 \%$ can be obtained [1]. This dome system can take up to one degree tracking error and, if necessary, can be modified to allow for a tracking error of up to four degrees [1].

The honeycomb structure will be made of graphite-epoxy with a protective coating. This structure provides a simple way to create a rigid array with less weight penalty than other systems. A thin aluminum sheet covers the rear of the panel in order to radiate excess heat from the power producing cell. The mass estimate of the radiator is based on a $150 \cdot \mathrm{C}$ cell operating temperature (thickness $=0.0025 \mathrm{~cm}$ ). A detailed mass breakdown of the array is given in Appendix 3.9.

### 3.4.3 Selection of Array Parameters

The selection of array area parameters was determined from the power requirements of the propulsion system combined with the efficiency of the Fresnel lens system in converting sunlight into electric power. The thickness of the array was determined from strength parameters. The array was designed so as to allow a maximum of 2 degrees deflection between any point on the array and the central point of attachment to the ship's truss.

The honeycomb structure (Figure 3.2) provides essentially all of the array's resistance to
bending. The two parameters that can be varied are the thickness and the height of the honeycomb walls. Increasing the height of the walls will increase the thickness of the array and, therefore, increase the volume of the array, which means the array will increase in both strength and mass. These parameters were chosen so as to minimize mass while keeping a reasonably small array volume while also maintaining a maximum array angle of deflection of 2 degrees.

The array optimization program (see Appendix 3.6) plotted wall thickness against array mass. At each height value, wall thickness was chosen so as to maintain a deflection of 2 degrees at point A, as seen in Figure 3.3.

The plot generated by the optimization program was the basis for the choice of honeycomb substructure parameters. Looking at Figure 3.4 and realizing that increasing the height of the cell compromises the number of units which can be launched in the space shuttle, a wall height of 6.48 cm was chosen as a compromise, corresponding to a $6 \mathrm{~cm} \times 6 \mathrm{~cm} \times 6.48 \mathrm{~cm}$ module size. The honeycomb wall will be .015 cm thick, and using these parameters for the array, we have a maximum deflection at point $A$ of 1.89 degrees and at point $B$ of 1.69 degrees (Figure 3.3).

The total array area of both hexagons will cover $11,957.76 \mathrm{~m}^{2}$ and have a total mass of 6,792 kg with an added $15 \%$ contingency of $1,019 \mathrm{~kg}$ for a total mass of $7,811 \mathrm{~kg}$ estimated. The solar cells have the capability to provide $3,836,133$ Watts of power before transmission.

### 3.4.4 Array Wiring

The 3,321,600 Fresnel Lens concentrator modules are arranged into two hexagonally shaped arrays. The objective was to deliver the power produced by the cells to the power processing units aboard the ship. The following criteria were established in wiring of the solar cells. A maximum voltage of 200 Volts was established to prevent arcing in the array. The wires going through the array should not carry more than 20-25 Amperes in order to reduce high power


Figure 3.3: Diagram for Sizing Program


Figure 3.4: Box Sizing
losses and to prevent the creation of large magnetic fields within the panel. The wiring scheme was established systematically to ensure easier assembly of the solar arrays as the design evolves. To prevent high power losses and prohibitive wiring weights, the lengths of the high carrying wires was reduced as much as possible. The total length of wire was also kept to a minimum. Final wire design calculations appear in the Appendix 3.8.

Each of the two arrays is divided into three equal parallelograms, as seen in Figure 3.5. Each parallelogram is divided into 173 segments that deliver their power to an electrical load carrying cable that runs beneath the array and along a side of the parallelogram. Each segment is made of 16 series in a line. See Figures 3.6 and 3.7 for the design of these segments and series. The series contain 4 rows and 50 columns of modules that are linked together by one wire. This wire connects itself to wires running along both sides of its segment.

Table 3.2 shows the wires selected in order to minimize power losses and wiring mass. Standard sizes of copper and aluminum wire were considered. The selection was broken up into: wiring of modules into series, wiring of series into segments, wiring of segments to deliver power to the center of the arrays, and the wiring of the array to the engines' power processing units. The wiring program is set up such that 50 Watts of power loss is equivalent to 1 kg of wire mass and the wire that minimized this combination was chosen for each different wire current (see Appendix 3.7). Using this wiring system, transmission power losses equal 140,327.7 Watts and the wiring mass equals $1,696 \mathrm{~kg}$. With a $15 \%$ contingency of 254 kg , the wiring mass is estimated at $1,950 \mathrm{~kg}$. The total power delivered to the ship's power processing units, then, becomes $3,695,805 \mathrm{~W}$.


EACH MOOLLE: $0.06 \mathrm{M} \times 0.06 \mathrm{M}$
EACH SERIES: 5 RONS X 50 COLUMNS
EACH SEGENT: 16 SERIES LONG
EACH SECTION: 173 SEGMENS

Figure 3.5: Concentrator Wiring Breakdown

## SEGMENT DESIGN



SERIES DESIGN


Figure 3.6: Wiring Divisions of Sections \#1 and \#2


Figure 3.7: Wiring Divisions of Section \#3

Table 3.2: Specific Wire Selection

| Current (A) | Length (km) | Material Gauge \# | Power Loss <br> (W) | Mass (kg) |
| :---: | :---: | :---: | :---: | :---: |
| Modules Wired Into Series |  |  |  |  |
| 1.2286 | 216.0915 | Al 20 | 17810.12 | 302.53 |
| Series Wired Into Segments |  |  |  |  |
| 2.4572 | 6.5500 | Al 17 | 1079.69 | 18.34 |
| 3.6859 | 6.5500 | Al 15 | 1521.66 | 29.21 |
| 4.9145 | 6.5500 | Al 14 | 2151.48 | 36.81 |
| 6.1431 | 6.5500 | Al 13 | 2669.57 | 46.37 |
| 7.3717 | 6.5500 | Al 12 | 3043.31 | 58.49 |
| 8.6004 | 6.5500 | Al 11 | 3284.76 | 74.02 |
| 9.8290 | 6.5500 | Al 11 | 4290.30 | 74.02 |
| 11.0576 | 6.5500 | Al 10 | 4308.69 | 93.01 |
| 12.2862 | 6.5500 | Al 10 | 5319.37 | 93.01 |
| 13.5148 | 6.5500 | Al 9 | 5096.51 | 117.25 |
| 14.7435 | 6.5500 | Al 9 | 6065.27 | 117.25 |
| 15.9721 | 6.5500 | Al 9 | 7118.27 | 117.25 |
| 17.2007 | 6.5500 | Al 8 | 6550.14 | 148.03 |
| 184293 | 6.5500 | Al 8 | 7519.30 | 148.03 |
|  |  | Subtotal: | 77,828.52 | 1473.62 |
| Segments wired to center of hexagon |  |  |  |  |
| 19.658 | . 0033255 | Al 8 | 4.34 | . 07516 |
| 2x 19.658 | . 0033255 | Al 5 | 8.69 | . 15065 |
| 3x 19.658 | . 0033255 | Al 3 | 12.26 | . 23944 |
| 4x 19.658 | . 0033255 | Al 2 | 17.29 | . 30196 |
| 5x 19.658 | . 0033255 | Al 1 | 21.43 | . 37911 |
| 6x 19.658 | . 0033255 | Al 0 | 24.47 | 47888 |
| 7x 19.658 | . 0033255 | Al 0000 | 1.66 | . 96108 |


| Current (A) | Length (km) | Material Gauge \# | Power Loss (W) | Mass (kg) |
| :---: | :---: | :---: | :---: | :---: |
| 8x 19.658 | . 0033255 | Al 0000 | 2.17 | . 96108 |
| . | . | - | - | . |
| - | - | - | - | . |
| 173x 19.3658 | . 0033255 | Al 0000 | - | . 96108 |
|  |  | Subtotal | 59,148.32 | 162.12 |
| Hexagon to power processing units on ship |  |  |  |  |
| $\begin{array}{r} 784.8 \\ \text { (at } 2444 \mathrm{~V} \text { ) } \end{array}$ | . 20608 | Al 0000 | 3350.87 | 59.56 |

Total Transmission Power Loss $=140,327.71 \mathrm{~W}$
Total wire mass $=1,695.53 \mathrm{~kg}$
$+15 \%$ contingency $=1,950 \mathrm{~kg}$

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## SECTION 4

## Structures

| Nomenclature |  |
| :---: | :---: |
| $\mathrm{a}_{1}$ | acceleration of point 1 |
| $\mathrm{a}_{\text {uip }}$ | linear ship acceleration |
| A | cross sectional area |
| $\mathrm{A}_{\text {paned }}$ | fractional area of panel |
| $\mathrm{A}_{\text {eroum }}$ | equivalent strut area |
| D | plate stiffness |
| $\mathrm{D}_{\text {max }}$ | maximum diameter |
| E | Young's Modulus |
| f | natural frequency |
| H | support truss depth |
| I | area moment of inertia |
| I | initial I of the strut |
| L | strut length |
| L' | strut half-length |
| $\mathrm{m}_{\mathrm{ij}}$ | consistent mass matrix entry $\mathrm{i}, \mathrm{j}$ |
| $\mathrm{m}_{\text {inericul }}$ | mass of inertial load at pt 1 |
| $\mathrm{m}_{\text {lp }}$ | mass of fraction of solar panel |
| $\mathrm{m}_{\text {pode }}$ | node mass |
| $\mathrm{m}_{\text {sout }}$ | strut mass |
| $\mathrm{N}_{\mathrm{i}}, \mathrm{N}_{\mathrm{j}}$ | shape functions |
| $\mathrm{P}_{\mathrm{g}}$ | critical load |
| $\mathrm{r}_{1}$ | distance from c.o.m. to point 1 |
| $t$ | strut wall thickness |
| $\mathrm{t}_{\text {Lex }}$ | solar panel flat plate thickness |
| U | strain energy |
| w | mass fraction |
| $\alpha$ | moment of inertia exponent |
| $v$ | Poisson's ratio |
| $\rho$ | density |
| ■ | vibration rate |

### 4.1 Truss Systems

The supporting structures of the craft are crucial to the overall system. Since we are dealing with a very large solar array system, the main concern is that the supporting structure be lightweight, yet sturdy and controllable. Choices for such truss systems must be both adaptable and functional.

Supporting truss structure for space applications has become an area of extensive study. Most of the research done on space applied truss structures and the construction scenarios for these systems has been directed for use by the Space Station Freedom (SSF). Much of it is extremely well suited for our purposes.

### 4.1.1 Strut Elements

A tetrahedral truss system made up of individual strut members and nodes, is widely studied for use on large, low mass structures, such as in the support of large reflector panels and for use on the Space Station Freedom [8]. A strut formed from two conical half-struts joined at their larger ends as seen in Figure 4.1 is suggested for space applications. The columns, end joints, and unions are designed for low-weight, high packing efficiency, single-assembly space truss applications [2]. The non-uniform cross sectional area allows for making the struts hollow, i.e. lighter, while not compromising the strut's strength and actually increasing its resistance to buckling.

This particular design aids in the prevention of buckling of individual members by tapering their cross sections. By increasing the thickness at the center of the half-conical strut to about 10 cm and tapering it to about 5 cm on the end, the resulting strut is sturdier [1]. Using strain energy principles, the critical buckling loads can be compared for a tapered and a uniform beam. Consider a uniform cross sectional beam that has an area moment of inertia, I, equal to


Figure 4.1: Strut Configuration


Figure 4.2: Node Attachment Illustration
the average of the tapered beam. The resulting critical buckling load for the tapered beam is 1.3 times higher than for the non-tapered beam. Thus, a typical strut member like that in Figure 4.1 can carry a higher load per unit mass than a uniform cross sectional beam.

### 4.1.2 Strut Assembly

Each half strut is connected by interlocking spring latches that are tapered to aid in alignment of the joint halves. The smaller end of the strut is threaded to receive one half of an internal spring loaded latching pawl that is designed for both quick attachment and quick release. One end of this joint is permanently attached to a nodal cluster similar to Figure 4.2, whereas the other half is attached to the strut itself [4]. Attachment of the strut to the node is also outlined on Figure 4.2. The strut end must be inserted into the node element and is then rotated to capture. It is then locked into place by rotating a locking collar 45 degrees [11]. The painted code, i.e. the black squares that line up when the strut is locked, was developed to ensure a quality connection and was suggested by astronauts while doing assembly tests at the Marshall Space Flight Center in their Neutral Buoyancy Simulator [5]. Assembly times of truss members, materials and masses will be discussed later in this section.

### 4.2 Solar Array Support

For the support of the solar arrays, we are faced with a number of specific criteria that will help determine the supporting structure. The power requirements of the ship establish the total area of solar panels that must be supported. Also, the solar cell system has a margin of sunlight detection error that dictates the maximum amount of out of plane deformation the structure can tolerate. In essence, this must be a near precision system with a maximum deflection of less than two degrees on its outer edge. For a large structure such as this, it is desirable to make it as stiff as possible [3]. That is, it must be resistant to vibration as well as
having the ability to support sufficient loads.
Several methods were considered for the support of the solar panels. Solar blankets were first considered for the flat solar arrays since this method was easy to deploy and had a very small mass per unit area. However, as the parameters of the ship, such as power required and choice of the solar cell evolved, this method became inadequate. The Fresnel lens system could not be "rolled up," as in a blanket, due to its rigid structure. Also, the required area was too large to be adequately supported against vibrations and deflections with the blanket method. Thus, the solar blanket was completely inadequate and not stable enough to justify its simplicity. Another method considered was to just wrap the solar cell membrane around a hexagonal frame. This also proved to be insufficient due to the nature of the Fresnel lens system. The Fresnel lens system cannot be wrapped around a frame which added very little support to the solar system against vibrations out of plane.

An alternate method of support was needed with better stability for a large surface area. Current research in precision reflector systems seemed well suited for our purposes. The underlying principle lies in that a hexagonal support structure made up of tetrahedral elements, as shown in Figure 4.3, is extremely stable. A hexagonal truss system has the tetrahedral, a 3dimensional triangle, as its primitive. A type of pyramid, the tetrahedral is the most basic 3dimensional, non-collapsible unit [7]. Thus, large arrays of tetrahedral elements, which can be easily arranged in the form of a hexagon, are currently being studied for large space structure applications [2].

The total solar panel area is assumed to be $11,600 \mathrm{~m}^{2}$. Each support structure must cover an area of $5,800 \mathrm{~m}^{2}$ which corresponds to a hexagon that is 48 m on a side. A hexagonal truss made up of tetrahedral elements is usually characterized by the number of "nings" it incorporates. The number of nodes, struts, and strut length size for a fixed area of $5,800 \mathrm{~m}^{2}$ for a number of


Figure 4.3: Hexagonal Truss
"rings" is given in Table 4.1. These numbers were found by a systematic counting system developed with the aid of a computer program that found the number and placement of the nodes for a given number of "rings" where a "ring" consists of a concentric array truss structure shells increasing the truss diameter. This was done in preparation for the structural analysis of several different numbers of rings. As seen in Table 4.1, the number of struts starts increasing dramatically as the number of rings increases. More importantly, with the exact number of nodes and struts known, supporting truss system masses have been determined for the solar panels using different numbers of rings (Table 4.1).

Table 4.1
Hexagonal Support Truss Parameters

| Rings | Nodes | Struts | Strut Length |
| :---: | :---: | :---: | :---: |
| 3 | 64 | 243 | 16 |
| 4 | 109 | 420 | 12 |
| 5 | 166 | 660 | 9.6 |
| 6 | 235 | 954 | 8 |
| 7 | 316 | 1302 | 6.9 |
| 8 | 409 | 1704 | 6 |

### 4.3 Structural Analysis

The structural analysis of this design consists of several parts. The stresses in individual members of all truss structure are of great importance, since we do not want to exceed the maximum allowable stress for the material. With the use of the finite element code, STAP, the box truss and the Solar Panel support structure were analyzed. STAP stands for Static Analysis Program and is a computer program that carries out a static linear elastic finite element analysis.

### 4.3.1 Strut Analysis

The half conical hollow strut members need to be modeled as a uniform truss element of uniform material and area. An equivalent area may be determined using the principle of conservation of strain energy. Taking advantage of the symmetry of the strut, we only have to model half of the strut. The equivalent area can be expressed as:

$$
U=\int \frac{E A}{2}\left(\frac{\partial u}{\partial x}\right)^{2} d x=\int \frac{E A_{\text {eq }}}{2}\left(\frac{\partial u}{\partial x}\right)^{2} d x
$$

Knowing that, for beam elements, $\mathrm{du} / \mathrm{dx}$ is a constant and the material Young's modulus is also constant, this relation becomes:

$$
A_{a q}=\frac{1}{L^{\prime}} \int A(x) d x
$$

Hence, the equivalent area is simply an average over the half strut length, L'. It is now necessary to find the initial cross sectional area as a function of x . Assuming that the wall thickness is constant and specifying a maximum strut diameter of 10 cm with a minimum diameter of 5 cm , this function becomes:

$$
A(x)=\pi\left(\frac{0.05 t}{L^{\prime}} x+0.05 t-t^{2}\right)
$$

Integrating from 0 to $\mathrm{L}^{\prime}$, this reduces to:

$$
A_{c q}=\pi\left(0.075 t-t^{2}\right)
$$

which is only a function of the wall thickness. Therefore, finding an equivalent area, using this method of energy conservation, becomes a simple task of setting the wall thickness. With the wall thickness and the strut length as parameters, the final strut element to be used can be
determined by examining the results of the STAP program.
Based on overall performance of the support structure for a number of rings, i.e. a number of strut lengths, the final sizes of the individual parts of the truss system were determined. The stress criteria, critical buckling loads, deflections, as well as overall mass, are the deciding factors in sizing. Although dynamic thermal loads will be the most crucial source of loads and deflections for this design, they are difficult to analyze with the resources available.

Inertial loads due to the ship's acceleration and rotation were evaluated in the STAP program. These loads are highest when the ship is spiraling out or in to Earth. Given a forward acceleration, $a_{\text {ship }}$, due to thrusting, of $5 \times 10^{-4} \mathrm{~m} / \mathrm{s}^{2}$, and an angular rotation rate, $\Phi$, of $1.13 \times 10^{-8}$ $\mathrm{rad} / \mathrm{sec}$, the local acceleration at any point along the ship is:

$$
a_{1}=a_{s h \varphi \varphi}+\omega^{2} r_{1}
$$

The radial distance, $r_{1}$, to a point on the ship is measured from the center of mass of the ship. The center of mass is 0.78 m offset toward the cargo pod from the center of the box truss connected to it. The resulting inertial forces are simply the mass of the component times the local acceleration given by the above equation and are applied as point forces.

The strain energy of a tapered element is given by:

$$
U=\frac{E}{2} \int I(x)\left(\frac{\partial^{2} v}{\partial x^{2}}\right)^{2} d x
$$

where $I(x)$ is the area moment of inertia. This is given as:

$$
I(x)=I_{0}\left[1+r\left(\frac{x}{L^{\prime}}\right)^{a}\right]
$$

where $\mathrm{I}_{0}$ is moment of inertia at the tip. $\mathrm{I}_{0}$ is $3.742 \times 10^{8} \mathrm{~m}^{4}$ for a strut wall thickness of 0.0008
m . The factors r and $\alpha$ are 7.1967 and 1.5663 respectively if symmetry is used and L ' is half the strut length. The stiffness matrix is found by invoking Castigliano's theorem and the buckling load, $P_{a r}$, may be found by solving the eigenvalue problem that results from work principles [13]. The geometric stiffness matrix, $\left[k_{z}\right]$, is the same as for a straight beam, thus giving the eigenvalue problem,

$$
\left|[k]-\lambda\left[k_{\boldsymbol{g}}\right]\right|=0
$$

where $\lambda$ is a function of $P_{c r}$. This calculation for a strut 16 m in length ( $L^{\prime}=8 \mathrm{~m}$ ) with $E$ equal to $1.723 \times 10^{11} \mathrm{~N} / \mathrm{m}^{2}$ yields a critical buckling load of 112.2 N . This method was also used to justify a tapered strut earlier.

### 4.3.2 Hexagonal Tetrahedral Truss Analysis

An exact analysis of the solar array support structure would incorporate plate elements as well as truss elements. Plate elements are needed to accurately model the added stiffness due to the solar panels. The STAP program considers only truss elements. Therefore, in order to account for the solar panel stiffness, mass as well as cross sectional area of the top truss elements are increased in the following manner.

An equivalent flat plate representing the solar panels has a thickness of 0.00566 m consisting of graphite epoxy. The cross sectional area of the top truss elements is then increased by $\left(w^{*} t_{\text {foce }}\right)$ or $2\left(w^{*} t_{\text {tace }}\right)$, where $t_{\text {foce }}$ is the plate thickness of 0.00566 m and $w$ is a fraction of the strut length less than $1 / 3$ the length. For example, a top strut, not on the edge, has an equivalent cross sectional area of:

$$
A_{e q}=A_{s \text { sraq }}+2 * w * t_{\text {foce }}
$$

Several values of $w$ were tested with STAP to show that in the worst case of a small $w$
(corresponding to the highest stresses), the stresses were considerably smaller than critical and displacements were on the order of $10^{-6} \mathrm{~m}$ maximum.

The inertial forces were adjusted by equally distributing the plate mass on the top truss nodes. This assumes that the panel will be attached to the truss at all top nodes. For example, consider a center top node. There are 9 struts attached to the node. Therefore, the mass for the inertial load at that node is:

$$
m_{\text {inertal }}=m_{\text {noch }}+9\left(\frac{1}{2}\right) m_{\text {strus }}+6\left(\frac{m_{\varphi}}{3}\right)
$$

where $\mathrm{m}_{\mathrm{tp}}$ is the mass of the solar panel divided by the number of equilateral triangles of the strut length that form the solar array hexagon. The masses for the inertial loads of the top nodes are calculated in a similar manner where the mass depends on the number of struts attached to the node and how many of them are top truss elements.

The results from the STAP program for the solar support truss with 3 rings are shown in Figures 4.4 and 4.5. The out of plane displacements of the nodes were no greater than $0.439 \times 10^{-6}$ m for the inertial loadings. The maximum stress is $0.854 \times 10^{3} \mathrm{~N} / \mathrm{m}^{2}$ and the maximum axial force is 0.0753 N . Both of these values fall well below their critical values. It may be expected that any additional loads due to thermal variations should not cause any problems with solar array panel deformation or buckling of the truss members.

The resulting length of the individual struts is quite long ( 16 m ). Therefore, the vibration of the truss structure must be examined. It is expected that the individual strut elements will vibrate at a lower frequency than the overall structure. Therefore, this is looked at first.


Figure 4.4: Nodal Displacements for the Hexagonal Truss


Figure 4.5: Hexagonal Truss Axial Forces

### 4.3.3 Vibrational Analysis

The tapered element will vibrate at rates given by the eigenvalue problem:

$$
\left|[k]-\omega^{2}[m]\right|=0
$$

where [ k ] is the same stiffness matrix defined in the critical buckling load analysis and [ m ] is the consistent mass matrix. The consistent mass matrix is found by

$$
m_{i j}=\rho \int A(x) N_{i}(x) N_{j}(x) d x
$$

where $N(x)$ are the shape functions of a beam and $A(x)$ is the cross sectional area as a function of $x$. If symmetry is used then half of the beam length can be used with the boundary conditions that the beam is pinned at one end and has zero slope at the other. This analysis results in $\boldsymbol{\infty}$ equal to $12.1 \mathrm{rad} / \mathrm{s}$, which corresponds to a frequency of 1.9 Hz . A non-tapered beam with the average radius of the tapered beam results in a frequency of:

$$
f=\left(\frac{\pi}{L}\right)^{2} \sqrt{\frac{I E}{\rho A}} \frac{1}{2 \pi}=\left(\frac{\pi}{16}\right)^{2} \sqrt{\frac{(1.325 e-7)(1.72 e I I)}{0.29}} \frac{1}{2 \pi}
$$

which equals 2.46 Hz . This last rough calculation was done as a check for the frequency analysis. Since the two numbers are of the same order of magnitude the frequency analysis using energy methods is accurate.

The natural frequency of the entire hexagonal truss structure may be approximated by a circular sandwich plate. The truss/panel system is considered dynamically as an equivalent flat circular sandwich plate [8]. This is a fairly good assumption for thin trusses and will be only an approximation for the 3 ring truss system [6].

Plate bending stiffness is defined as:

$$
D=\frac{E_{\text {face }} t_{\text {foct }} H^{2}}{2\left(1-v^{2}\right)}
$$

where,

$$
H=\sqrt{\frac{2}{3}} L_{\text {struat }}
$$

Truss Mass/Area may then be defined to be:

$$
\frac{M}{A}=\frac{\left(\text { struts) }\left(M_{\text {srrus }}\right)+(\text { nodes })\left(M_{\text {nock }}\right)\right.}{A}
$$

The resulting plate frequency is then calculated by:

$$
f=\frac{3.343}{\left(D_{\max }\right)^{2}} \sqrt{\frac{D}{M / A_{\text {trus }}+M / A_{\text {panal }}}}
$$

where $D_{\max }$ is the diameter of a circle having the same area as the solar panel surface [8]. That is,

$$
D_{\max }=\sqrt{\frac{5800 * 4}{\pi}}=85.9 m^{2}
$$

The parameters used for these equations are as follows:

$$
\begin{aligned}
& \mathrm{E}_{\mathrm{fcce}}=1.73 \times 10^{11}\left(\mathrm{~N} / \mathrm{m}^{2}\right) \text { (graphite epoxy) } \\
& \mathrm{t}_{\mathrm{fcec}}=5.66 \times 10^{-3}(\mathrm{~m}) \\
& v=0.3 \text { (Poisson's ratio) }
\end{aligned}
$$

The resulting fiequcicy for a 3 ring tuiss system is 42 Hz . As expected, this is higher than the natural frequency of the individual truss elements.

### 4.3.4 Central Truss Analysis

The orthogonal tetrahedral truss pictured in Figure 4.6 is analyzed in a very similar manner. The STAP program is quite accurate for this truss system. In an effort to keep the center of mass of the ship near the geometric center the propellant tank is placed in the center cell of the 3 cell truss. Since we want to minimize the distance to the solar panels from the center of mass of the ship, the strut length is 5.4 m . This is adequate for placing the propellant tank, with a length of 5.3 m , within a cell of the truss.

The critical loads for the orthogonal tetrahedral truss elements are found in the same manner as before. Using strain energy methods, we analyze the 5.4 m and 7.6 m strut elements. Once again, we assume a wall thickness of 0.0008 m , a maximum diameter of 10 cm and a minimum diameter of 5 cm . This analysis results in a critical buckling load of 984.13 N for the 5.4 m strut and 497.16 N for the 7.6 m strut. With these values known, we can examine the resulting axial forces to make sure no elements buckle. Once again, inertial forces due to the engine thrust are studied with STAP. The boundary conditions for this box trass include fixing the nodes that attach to the engine casing from displacement. This allows the STAP analysis to have a fixed reference, otherwise the applied forces would result in rigid body motion only. The propellant tank does not contribute inertial forces since it is attached to the four fixed nodes. An inertial force of $\left(61 \times 10^{3} \mathrm{~kg}\right)\left(5 \times 10^{-4} \mathrm{~m} / \mathrm{s}^{-4}\right)$, which equals 30.5 N , is distributed equally on the 4 nodes where the cargo attaches to the truss. This assumes that the initial acceleration is $5 \times 10^{-4}$ $\mathrm{m} / \mathrm{s}^{\mathbf{2}}$. We then transfer the moment and shear force created by the solar panels to the four nodes at each end of the truss. Given that the mass of a solar panel with support truss and wiring equals 5750 kg , the resulting inertial force equals 2.875 N . This force is translated and equally distributed to the four connecting end nodes. Finally the moment caused by this inertial force


Figure 4.6: Orthogonal Tetrahedral Truss Configuration


Figure 4.7: Box Truss Axial Forces
acting at the center of mass of the solar panel structure is transferred to the end nodes. This moment equals the inertial force, 2.875 N , times the moment arm of 51.5 m , giving a moment of 148.06 Nm . Coupled forces are applied to the four end nodes that are separated by 5.4 m . Therefore, a force of 13.7 N is applied to each of the four nodes.

The results from the STAP code are represented graphically in Figure 4.7 and 4.8. The maximum displacements do not exceed $0.305 \times 10^{-5} \mathrm{~m}$. The maximum stresses and axial forces do not exceed $0.736 \times 10^{5} \mathrm{~N} / \mathrm{m}^{2}$ and $0.1375 \times 10^{2} \mathrm{~N}$, respectively. Given that the yield stress is on the order of $50 \times 10^{6} \mathrm{~N} / \mathrm{m}^{2}$, these values are well below the critical buckling load and yield stress for the struts used. Once again a large margin of safety is available. This gap is necessary in order to insure a stable system under thermal and other loads not accounted for. Therefore, the truss is not decreased in strength even though this analysis shows that it more than meets the requirements.

### 4.4 Assembly Scenario and Time Estimates

A complex assembly such as this may be broken down into basic steps for the purpose of studying parameters such as assembly rates and tasks. These studies are being done at the NASA Langley Research Center where tetrahedral truss assembly is simulated in the Weightless Environment Training Facility [11]. In actual space studies, such as with the ACCESS shuttle mission, truss assembly studies were completed in space by astronauts with the aid of the Remote Manipulator System (RMS) [5].

The large structure of the spacecraft will require a significant amount of lead time so that the ship can be built in time for the launch window of our mission. From the NASA studies it was estimated that average in-space assembly rates of approximately 38 seconds per strut can be expected [5]. Therefore, a truss of consisting of 44 struts will require approximately 28 minutes


Figure 4.8: Nodal Displacements for the Box Truss


Figure 4.9: Thermaily Induced Tip Distortion
to assemble in space. According to predictions made by NASA as a result of the ACCESS shuttle experiment, a hexagonal truss system consisting of 789 struts would require approximately 6 hours to assemble with the aid of the Mobile Transporter system [12]. The Mobile Transporter is a mobile work station concept currently being developed by NASA. It is assumed that this technology will be well developed by our launch time. Current projections lead to an assembly time for a hexagonal shaped truss system of 3 rings, such as the one considered for our ship, of approximately 3 hours if completed with the aid of the Mobile Transporter [4]. The total predicted time for assembly of the central box truss and both solar array support trusses is a more conservative 36 hours. This estimate is higher than NASA predictions due to the length of the elements and the overall size of our structure.

### 4.5 Materials Selection

An obviously important step is to specify the materials to be used in this design, why they were chosen, and what the resulting masses are. The criteria for material selection results from an analysis of what the ship will have to withstand. A low coefficient of thermal expansion is desired due to the dynamic nature of the thermal loadings. These loads will vary considerably during spiral out and in to Earth. Also desired is low density with high strength to minimize the ship mass safely. Finally the material must be stable in the space environment, i.e. not subject to breaking down in radiation and able to withstand collisions with micro-meteorites. Degradation caused by radiation, thermal distortion, and minute particle bombardment is a major concern when dealing with polymer-based composites. However, recent studies of aluminum reinforced graphite-epoxy have shown that this material has better thermal conductivity and lower susceptibility to space environmental effects. In other words, aluminum reinforced graphite epoxy has the advantage of a lightweight epoxy without the disadvantage of high thermal
distortion. However, due to the uncertainty of the effects of atomic oxygen on this material, aluminum coated graphite epoxy would be a better material choice. This material is the prime candidate for use on the SSF. Figure 4.9 and Table 4.2 display the comparisons of graphite epoxy, and two forms of graphite reinforced aluminum, P100/6061, and P55/6061. The data clearly shows that P100/6061 has the least thermally induced tip distortion per unit weight. Also, this material has a significantly larger thermal conductivity and modulus of elasticity. It is assumed that the aluminum coating on the graphite epoxy will not change the stiffness characteristics of graphite epoxy since the coating is thin. Thermal properties will most likely be similar to the aluminum reinforced graphite epoxy.

Table 4.2
Material Properties Comparison

| Property | Graphite/Epoxy | P100/6061 | P55/6061 |
| :--- | :---: | :---: | :---: |
| $\mathrm{A}(11) 10^{-5} / \mathrm{K}$ | -0.080 | 0.086 | 0.307 |
| $\mathrm{~A}(22) 10^{-5} / \mathrm{K}$ | 3.67 | 2.30 | 2.40 |
| $\mathrm{~K}(11) \mathrm{W} / \mathrm{m} \mathrm{K}$ | 54.0 | 240.0 | 98.0 |
| $\mathrm{~K}(22) \mathrm{W} / \mathrm{m} \mathrm{K}$ | 0.7 | 193.0 | 98.0 |
| $\mathrm{NU}(12)$ | 0.21 | 0.4 | 0.27 |
| $\mathrm{NU}(21)$ | 0.010 | 0.031 | 0.041 |
| $\mathrm{E}(11) 10^{6} \mathrm{psi}$ | 24.97 | 51.0 | 30.8 |
| $\mathrm{E}(22) 10^{6} \mathrm{psi}$ | 1.17 | 4.0 | 4.7 |
| $\mathrm{E}(12) 10^{6} \mathrm{psi}$ | 0.62 | 2.0 | 1.9 |

A11 = Coefficient of Thermal Expansion (fiber direction)
A22 = Coefficient of Thermal Expansion (transverse direction)
K11 = Thermal Conductivity (fiber direction)
K22 $=$ Thermal Conductivity (transverse direction)
NU12 $=$ Poisson's Ratio
NU12 $=$ Poisson's Ratio
E11 = Modulus of Elasticity (fiber direction)
E22 = Míudulus of Elasicity (ransverse direction)
E12 $=$ Shear Modulus

Titanium and its alloys are rapidly becoming a leading space application material. In our situation they are considered for the fabrication of the nodal clusters. Pure titanium is as strong as steel, but $45 \%$ lighter. Although it is $60 \%$ heavier that aluminum, it is twice as strong [10]. A titanium alloy such as Ti-6A1-4 V, which is a titanium-aluminum alloy, has a lower density than pure titanium, slightly higher coefficient of thermal expansion, a slightly lower thermal conductivity, but a higher Young's modulus. For property values and comparison to other titanium alloys, see Table 4.3 [9].

Table 4.3
Material Properties of Titanium

| Property | Titanium (99\%) | Ti-6Al-4 V | Ti-5AI-2.5 Sn |
| :--- | :---: | :---: | :---: |
| Density lb/in ${ }^{3}$ | 0.163 | 0.160 | 0.161 |
| TEC $10^{-6} \mathrm{in} / \mathrm{in} \mathrm{F}$ | 4.7 | 5.0 | 5.2 |
| TC Btu/ft ${ }^{2} \mathrm{hr} \mathrm{F} \mathrm{ft}$ | $9.1-11.5$ | 4.2 | 4.5 |
| E $10^{6} \mathrm{psi}$ | 15 | 16.5 | 16.0 |

TEC $=$ Thermal Expansion Coefficient
TC = Thermal Conductivity
$\mathrm{E}=$ Young's Modulus

The resulting masses of the struts and nodes can now be determined. A hollow node, approximately 5 cm in diameter, can be estimated to be spherical with a thickness of about 1 cm . This gives a conservative nodal mass of 5.2 kg .

The mass of the struts is a function of their wall thickness. Specifying that the struts have a maximum diameter of 10 cm with a minimum diameter of 5 cm , their mass is given by:

$$
M_{\text {strut }}=\frac{2}{3} \rho \pi L\left(0.075 t-\frac{t^{2}}{2}\right)
$$

A half conical strut that is a total of 5.4 m long, made of aluminum coated graphite epoxy, with
a wall thickness of 0.8 mm , can be estimated to have a mass of 1.56 kg . Using the above formula, the strut mass may be determined for a number of cases as shown in Table 4.4. (Table 4.4 was calculated for a wall thickness of 0.8 mm .)

Table 4.4
Solar Support Structure Parameters

| Rings | Strut Length <br> $\mathbf{m}$ | Strut Mass <br> $\mathbf{k g}$ | Truss Mass <br> $\mathbf{k g}$ | Mass/Area <br> $\mathbf{k g} / \mathbf{m}^{2}$ |
| :---: | :---: | :---: | :---: | :---: |
| 3 | 16 | 4.60 | 1450.6 | 0.2501 |
| 4 | 12 | 3.48 | 2027.1 | 0.3495 |
| 5 | 9.6 | 2.78 | 2698.0 | 0.4652 |
| 6 | 8 | 2.32 | 3433.3 | 0.5919 |
| 7 | 6.9 | 1.99 | 4245.9 | 0.7321 |
| 8 | 6 | 1.74 | 5088.4 | 0.8773 |

It is apparent that the configuration and resulting mass of such a large hexagonal array is of great importance. Given that a aluminum coated graphite epoxy strut, 5.4 m in length, has a mass of approximately 1.56 kg and that a node has a mass of approximately 5.2 kg , the support structure mass is found and displayed in Table 4.4. The mass of the truss structure supporting the arrays is estimated at $2,901 \mathrm{~kg}$ with a $10 \%$ contingency of 290 kg to give a mass of 3,191 kg.

### 4.6 Notes on the Configuration

The development of a ship configuration is a trial and error process where ideas are brought up and either expanded to be incorporated into the configuration or discarded as inadequate. A crucial component of the vehicle was the solar array system and the supporting
structure.
Among the first choices for solar cell systems was a dynamic concentrator concept. This system would have required precise direction control and would have required the assembly of many individual cell packages. This lead to the first hexagonal shaped truss support. The hexagonal shape allowed for best packing efficiency and a very stable structure.

At the same time, flat solar cells such as silicon cells were considered. These solar arrays could be rectangular, but the area needed to produce the initial power requirements of about 3.5 MW was very large.

Once it was determined that the solar panel rotation only needed to be about one axis, the approximate points of rotation were determined. At this point, in an effort to keep the moments of inertia of the ship as small as possible for control purposes, these nodes were placed as close to the center of mass of the ship as possible.

The resulting layout is divided into four parts: the solar panel with its supporting structure, the connecting truss, the cargo hold and the engine support bay (Figures $4.10 \mathrm{a}-\mathrm{b}$ and $4.11 \mathrm{a}-\mathrm{c}$ ).

A $5,800 \mathrm{~m}^{2}$ solar array panel is located on two sides of the ship. The 48 m on a side hexagonal shaped arrays need to be attached in such a way that, when the ship is maneuvering, they can be rotated to an optimum position for gathering sunlight. Their need to be adjustable is why they are attached to the main body of the ship by rotational nodes. These rotational joints will carry the necessary moments and forces to rotate the panels.

The propellant tank was placed inside the inner most cell of the box truss. Since the tank has a maximum dimension of 5 m in length, a 5.4 m truss element is used so that it can be fitted in this way. This is done in an effort to keep the center of mass of the ship as close as possible to the geometric center.

Concentrator Array Configuration
centrator Aггay Configuration
Figure 4.10a: $\quad$ Configuration



Figure 4.11a:
Configuration


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Figure 4.11b:
Main Body Configuration


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Figure 4.11c:
Main Body Configuration

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The cargo pod contains the 61 mT payload and is oriented so that a minimum center truss length can be achieved while at the same time making it accessible for loading and unloading. The cylindrical hold is 27 m long and 7.8 m in diameter, as specified by the Shuttle ' C ' cargo hold.

The engine pod consists of 17 engines that were placed on the aft end of the ship in such a manner that it will "push" the ship to its destination.

The main body of the ship is mostly symmetric for easier control. The distance from the center of the ship to the rotational nodes must be minimized in order to minimize the forces at that node as well as the moment of inertial of the ship. The ship has developed into its current configuration due mostly to power requirements, solar cell choice, and ship control. It is an adaptable design that can be adjusted to meet a variety of mission requirement changes.

The general specifications and mass summary of the ship are given in Table 4.5. In addition, a craft mass distribution is displayed in Figure 4.12.

Table 4.5 Ship Specifications and Mass Summary
Cargo mass ..... $61,000 \mathrm{~kg}$
Array area
Array output power
Thruster specific impulse
Number of engines
Engine input power
Engine thrust
Solar array type
Propulsion system efficiency ..... 0.714Fresnel lens concentrator/multi-stacked cell
$10,800 \mathrm{~m}^{2}$
3.465 MW
10,300 sec
17 (including 5 redundant units)
275 kW
5 N
LEO to LMO trip duration960 days
Return trip duration
Round trip (LEO-LMO-LEO) duration ..... 382 days
3.8 years
Masses, kg
Cargo ..... 61,000
Propellant ..... 24,200
Tankage ..... 820
Propulsion system ..... 9,140
Solar arrays ..... 7,811
Wiring ..... 1,950
Truss structure supporting arrays ..... 3,191
Main body (trusswork, engine mounts, ..... 3,600
cargo bay, rotational $\alpha$-joints)
Power distribution and ..... 1,000
auxilliary power systems
Attitude control system, CMG ..... 2,000
Guidance/Navigation/Communications/ ..... 800
Data Systems
Total Mass ..... $115,510 \mathrm{~kg}$

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## SECTION 5

## SEP Vehicle Launch to LEO

## Nomenclature

LEO Low Earth Orbit<br>SSF Space Station Freedom<br>CMG<br>Control Moment Gyro

### 5.1 General Considerations

The transportation of the vehicle components and the cargo to LEO will be performed using four Shuttle C missions (Table 5.1). The Shuttle C will have a cargo launch capability of 7.6 m diameter by 27 m length with a mass of 61 metric tons. Its tanker launch capability is 4.6 m diameter by 25 m length with a mass limit of 71 metric tons (Figure 5.1).

Critical components in planning the launch to LEO are a) the economical packaging of the components for assembly in LEO and b) packaging of the components in such a way that their resultant center of mass will fall inside the prescribed envelope of the Shuttle Cargo bay. One goal of the process of planning the launch scenario is to minimize the number of shuttle launches to the Space Station. Another consideration is the minimization of assembly work (both mechanical and human) at SSF. Most of the system will be assembled on the ground leaving only some mechanical assembly to be performed in LEO.

Shuttle "C" Launch Capabilities


### 5.2 Manifest

The manifest of components to be launched to LEO is as follows:
$0 \quad$ Solar Array
o Supporting Truss Structure, Nodes, and Alpha Joints
o Propellant Tank
o Wiring for Solar Array
o Engines
o Control Systems
o Communications
o Power Conditioning Systems (for solar array and engines)

- Cargo

The order of the launches will be determined by the priority for assembly in space. The more complicated systems will be sent up first; the pre-assembled systems will be last.

### 5.3 Launch One

The first launch of the shuttle will contain the most complex components of the vehicle with regards to assembly in orbit: the support structure members and nodes. These members will be shipped unassembled and will be assembled in orbit at SSF using the mobile transporter. The remaining space in the first shuttle launch will be used to launch the solar array wiring from the arrays to the engines, which needs to be in place on the support structure before the solar array is attached, and some of the solar array. In addition, the control moment gyros will be launched at this time. Figure 5.2a shows a breakdown of launch one.

The individual strut members (Figure 5.3) will be arranged in a honeycomb box (Figure 5.4) pattern which will maximize packing efficiency. The struts for the support structure are in three different lengths: $16 \mathrm{~m}(4.6 \mathrm{~kg}), 7.6 \mathrm{~m}(2 \mathrm{~kg})$, and $5.4 \mathrm{~m}(1.565 \mathrm{~kg})$. The four-hundredeighty 16 m struts which will be packaged in a box 3 m by 1 m by 24 m . The eighteen 7.6 m struts will be assembled in a container 0.475 m by 0.01 m by 26.6 m . The container for the twenty-eight 5.4 m struts will be 0.625 m by 0.01 m by 21.6 m . These smaller bundles will be

EARTH TO LEO TRANSPORTATION
LAUNCH FOUR


Figure 5.2:


Figure 5.3: Strut Member


Figure 5.4: Layout of Struts in Honeycomb Shape
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placed alongside the larger box. The nodes will be placed in the remaining area next to the support structure package.

Within the additional area on the sides and bottom of the cargo bay surrounding the strut box, the nodes and wiring will be packaged. The spherical nodes will be lined up in a box 5 cm on a side with a length of 4 m . The cylindrical shaped wiring will line the remaining area in the bottom and sides of the cargo area.

After packing the support structure, nodes, and wiring, there remains 5 m of usable height left in the cargo bay for storage of the solar array which is the next priority for assembly in LEO. The solar array will be packed in modules and will be mechanically assembled at SSF.

In packaging the solar array for launch, it is desired to keep the array intact in large segments so that assembly in space is a minimum. The area of the arrays will be split into 105 rhomboid segments which will be 24 m on the long sides and 4.8 m wide. Each segment will be 6.48 cm high and will be separated by 0.02 cm of packaging material so that the glass lenses are not damaged. The remaining area of this first launch will hold 70 of these segments for a mass of 5.5 metric tons.

The six control moment gyros will be split into two sets of three and will be placed in the sides of the cargo pod on either side of the stored solar array. Each gyro is a cylinder 3 m in diameter and approximately 1 m long with a mass of 333 kg each. The total mass of the CMGs and RCS is 2,000 kilograms.

Because of the four $g$ acceleration, which the cargo will have to endure upon takeoff, packing materials will fill any empty spaces in the cargo bay so that the components will remain stationary and rigid.

### 5.4 Launch Two

The second launch will also be a cargo launch and will consist of the remainder of the solar array and the engine pod structure. In addition, the communications and control systems and additional miscellaneous hardware will be transported in this mission. Figure 5.2 b shows a breakdown of launch two.

The remaining 35 segments of the solar array will be packaged in the same cross sectional manner as in launch one and will be 227.5 cm high. The portion of the array to be launched at this time will have a mass of approximately 2.3 metric tons.

The engine pod structure is 5.65 m deep by 4.8 m high and has a mass of 9.14 metric tons. For launch packaging, the pod structure will be placed with its larger dimension ( 11.2 m ) along the length of the cargo bay. The large concentrated mass will be kept in place with springs in order to dampen any large motions it might experience in launching. The engine pod will be attached to the main structure using the robotic arm on SSF.

The communications system will be highly sensitive to jarring and will need careful protection during lifting to LEO. Upon arrival in LEO, the communications will be tested by the SSF crew.

The second launch will carry the power conditioning systems, propellant feed system, guidance, navigation and control systems, and the attachment components for the engines to be fitted to the main ship. This equipment will have a mass of 2 mT . Some of these systems and/or components will be assembled by astronauts.

### 5.5 Launch Three

The third launch will be a tanker launch and will contain the argon propellant tank. The propellant is launched in this later launch as it is self contained and storage time will be
minimized. The remaining area of the launch can be used to transport other materials to SSF. Figure 5.2c shows a breakdown of launch three.

The propellant tank is a spherical-ended cylinder with a length of 4.16 m and radius of 1.3 m with spherical ends. The tank and propellant have a mass of approximately 26 metric tons. The argon will be launched in a stronger storage tank, hereafter referred to as the "launch tank," as the tank for the mission is not strong enough to withstand the force of the launch from Earth when loaded with the argon. The launch tank will be packaged so as not to move in the cargo bay as this cargo will be sensitive to movement and force. The argon will then be transferred to the ship's tank in LEO, and the launch tank may then be reused. In LEO, the propellant tank will be joined to the ship using robotic assembly.

The remaining space in launch three can be used by NASA to transport equipment needed at Space Station Freedom for this or other projects. In addition, this space may be sold to private industry for transport of experiments or material to Space Station Freedom.

### 5.6 Launch Four

The fourth and final launch to LEO will consist of the cargo to be transported to Mars as specified. A breakdown of launch four may be seen in Figure 5.2d.

### 5.7 Repeat Launches from LEO

Upon returning to LEO from the Mars mission, some components which will have to be replaced are:
o Propellant Launch Tank, filled ( 26 metric tons)
o Engine Pod (if defective) ( 17.5 metric tons)
o Faulty Segments of Solar Array (if necessary)
o Any faulty components
These components may be launched from Earth using subsequent Shuttle C launches.

Replacement material for each subsequent mission should not require more than one Shuttle " C " launch. Replacement of components will be achieved in LEO using robotic arms and human assemblers.

Table 5.1 Launch to LEO Packaging Scenario

Launch One

| Component | Volume Used $\left(\mathrm{m}^{3}\right)$ | Estimated Mass (mT) |
| :---: | :---: | :---: |
| Struts | 75.0 | 3.2 |
| Nodes | 0.0073 | 0.5 |
| CMG, RCS | 7.07 | 2.5 |
| Solar Array | 525.0 | 5.6 |
| Wiring | 0.0016 | 0.16 |

Launch Two

| Component | Volume Used $\left(\mathrm{m}^{3}\right)$ | Estimated Mass (mT) |
| :---: | :---: | :---: |
| Communications | 50.0 | 0.8 |
| Thruster System | 252.0 | 9.14 |
| Solar Array | 262.0 | 2.8 |
| Miscellaneous | 50.0 | 1.5 |

Launch Three

| Component | Volume Used (m) | Estimated Mass (mT) |
| :---: | :---: | :---: |
| Propellant Tank | 35.0 | 0.820 |
| Propellant Launch Tank <br> (filled) | 35.0 | 27.0 |
| Non SEMMII | 300.0 | 40.0 |
| Power Distribution | 50.0 | 1.0 |

Launch Four

| Component | Volume Used $\left(\mathrm{m}^{3}\right)$ | Estimated Mass (mT) |
| :---: | :---: | :---: |
| Cargo | 1225.0 | 61.0 |

## SECTION 6

## Summary

The design of a large, unmanned solar powered electric propulsion cargo vehicle for use in Mars missions presents a number of interesting technical challenges. For each of the areas studied in this report, several solutions were considered. The overall configuration was arrived at through a rough optimization process that attempted to improve the performance of each subsystem, but also considered how each subsystem would fit into the overall mission architecture. The resulting configuration reflects the trade-offs that best suited the mission as a whole.

In the area of trajectory optimization, approximately 400 round trip trajectories were considered using the Quicktop trajectory calculation code. These trajectories were compared on the basis of required propellant, engine thrusting time, loiter time at Mars, refit time at Earth, and overall mission time.

A wide range of possible trajectories were found to consume similar amounts of propellant and require similar thrusting times for a given power input. Because of the relative insensitivity of the trajectory to these two design parameters, the trajectory was optimized primarily with respect to loiter and refit times required. It was desired to minimize the loiter time at Mars and maximize the refit time at Earth for the shortest overall mission (three round trips). The result of this analysis is a trajectory that requires 960 days to reach Mars (including spiral escape and capture), a Mars loiter time of 60 days, a return trip that requires 382 days, and a refit time at Earth of 210 days. Thus an overall mission time for three round trips is just over 12 years.

The electric propulsion system was the next topic considered. In order to achieve a high
payload to mass ratio, electrostatic ion thrusters were selected because of their characteristically high specific impulse. Argon was selected primarily on the recommendation of NASA Langley Research Center.

Mission time parameters lead to a desired initial thrust level at Earth of 60 N . After considering various sizes of thrusters, it was decided that this initial thrust could be achieved using an array of twelve 5 N engines. Five engines will be used near Mars for a maximum thrust of 24 N . Five additional engines were included for redundancy, bringing the total size of the thruster array to 17 engines, each with a projected lifetime of 25,000 hours. The optimum specific impulse was found to be 10,300 seconds.

The solar power collection system was selected after comparing varicus cell materials in terms of efficiency, radiation resistance, and annealing capabilities; a multistacked cell was chosen to provide 3.5 MW electrical power to our vehicle. To choose an array configuration, design point comparison calculations were carried out for a flat array and for a concentrator system. A Fresnel lens concentrator system was chosen and optimization of the concentrator module size was carried out. The concentrator system was selected with a total area of about $10,800 \mathrm{~m}^{2}$, divided equally between two hexagonal concentrator arrays.

A wiring scheme was developed so as to minimize both the weight and power loss associated with the solar panel wiring configuration.

The structural support of the Fresnel lens concentrator arrays had to be resistant to out of plane deformations as well as vibrations. The hexagon shaped, three ring truss structure was designed to be light weight, yet resistant to deflections. It consists of 234 half-conical, thinwalled struts with 64 nodes and will be made of aluminum coated graphite epoxy. One of these truss systems was needed to support each of the two solar arrays. Structural analysis with the STAP computer code verified the integrity of the solar array support structure under inertial loads
due to acceleration and rotation.
An orthogonal tetrahedral truss was analyzed for the central supporting structure of the ship. Inertial loads due to engine thrusting with cargo, propellant tank, and solar array attachments were considered. A three cell truss consisting of members 5.4 m and 7.6 m in length resulted in satisfactory stresses and deflections in members.

The control system will be used to a) direct the solar array panels so as to maximize the amount of sunlight captured by the solar cells and b) vary the ship's attitude and orientation during orbital maneuvering and thrusting. The control system will include control moment gyros (CMG) and reaction control system. The CMG system will consist of three double gimballed control moment gyros allowing for 100 percent momentum utilization about three primary axis, as well as 100 percent redundancy of the total control. The selected units will allow the Earth spiral out maneuver to be completed without a need for desaturization.

The transportation of the vehicle components and cargo to LEO will be performed using four Shuttle C missions. The first launch will contain the supporting structure struts and nodes, 70 percent of the solar arrays, and the control moment gyros for a total mass of about 13 metric tons. The second launch will contain the remaining portion of the solar array, the engine pod, the communications package, attachment components, and other miscellaneous items for a total mass of 24.2 metric tons. The third launch will contain the propellant tank at 20 metric tons. The fourth and final launch will carry the cargo at 61 metric tons. The final assembly (both mechanical and human) of the vehicle will take place in LEO.

For each additional LEO-LMO-LEO trip, one shuttle launch will be used to service or replace engines as needed, supply the propellant for the next trip, as well as provide any other miscellaneous components or repair parts.

The vehicle design described in this study represents a feasible, albeit preliminary,
configuration that could fulfill the unmanned cargo support role in a long-term mission architecture for a manned exploration of Mars in the 2010-2020 time frame.

## APPENDIX 1

## Trajectory Calculations

## Nomenclature

a Orbital parameter
$\mathrm{A}_{1}$ Constant tangential thrust
E Orbital energy
e Eccentricity
EA Eccentric anomaly
G Gravitational acceleration of Earth Gravitational parameter of a planet
Mp Dry mass of craft
$\mathrm{M}_{\mathrm{o}} \quad$ initial stage mass (pmefdmai.bas)
$\phi \quad$ Flight path angle
p Curvature radius of orbit
$r$ Mean orbital radius
$r_{0}$ Initial orbital radius
$r_{2} \quad$ Radius at apoapse
$r_{p} \quad$ Radius at periapse
s Arc length of an orbit
TOF Time of flight
$v$ Orbital velocity
$\mathrm{v}_{0} \quad$ Initial orbital velocity
$v_{\text {. }} \quad$ Hyperbolic excess velocity
A-1.1 Program TIMEOPT.BAS
Progratn TIMtopt.BAS
This program computes the lolter time at mars and reflt lime at earth
from the gulckrop hiputs of time and change in true anomaly angle. OPEN "SYNOD DATT" FOR INPUT AS "1
L.PRINT"'TIME:OUT"."TIME:BACK","LOITER","REFIT","MD (years)"

 IF PEGA)2•••1 OR I'EGB>2•PI THEN GOTO IF ABS (PEGA-PEGB) $\quad$. 55 THEN GOTO 670
'PRINT ABS(PEGA-PEGB) GOTO 196
EG=PES+WE'T6
$P E 7=T E 6+W E^{\prime} T 7$
$P M 7=P M G+W N \cdot T 7$
PM7 $=$ FM $6+W \mathrm{~W}$



o PRINT "hto Etom"PERIMS.TETI TTZ
M cap SP"PES.IMS.T3.T1+T2+T3
PRINT "M loiter".PE4.PMA.TA.T1+T2
PRINT "M esc SP".PES.PMS.TE.T1 $+\mathrm{T} 2+\mathrm{T} 3+\mathrm{T} 4+\mathrm{TE}$


If PM7>2•PI OR PM7)2•H THEN GOTO 920
-
$F 2 \mathrm{~g}=\mathrm{PMO}-\mathrm{PE} \mathfrak{\mathrm { g }}$


 inT
56 PRINT:PRINT
60 LPRINT T1 $2+T 3 . T 5+T 6+T 7 . T 4 . T 8 \cdot(3 \cdot(T 1+T 2+T 3+T 4+T 5+T 6+T 7)+2 \cdot T 8) / 365$ 070 GOTO 395
PRINT"Change in true anomoly. earth to mars $={ }^{\prime \prime} ;$ A $^{\prime}:^{*}$ - rads
10 CLS:PRINT"SOLAR-ELECTRIC CARGO FERRY MISSION ARCHITECTURE (ORBITAL MECIIANICS
PRINT

INPUT"Earth escape spiral time":T1
INPUT"HTO Earth to Mars time":T2
User must frput times and true anomoly changes below
$\mathrm{Pl}=3.14159$
$\mathrm{WE}=.61721$
KEY OFF:COLOR 5.0 .3
PI= 3.1415 .3
$W M=9.150999 \mathrm{E}-63$
152
490
124
0
39
300
86
7.85
4.25
spiral capture / mars
loiter time (INITIAL G
spiral escape / mars
splral capture / earth
INPUT"True anomoly change. Earth to Mars HTO. deg":AA2:A2 $=$ AA2 ${ }^{\prime \prime}$ P1/18日
INPUT"Mars spiral capture llme";T3
INPUT"Mars spiral escape time":T5
INPUT"HTO Mars to fiarth time":T6
INPUT"True anomoly change. Mars to Earth HTO. deg":AA6:AG=AAG•Pl/I80
INPUT"Earth spiral capture time":T7 INPUT"Earth spiral capture time":T7
$7 \mathrm{~T} 4=\mathfrak{G}$

PRINT"For input phase times, see below (Dhasetime column):PRINT
PE $0=-$ WE
$P M G=A 2-W N^{\circ}(T 1+T 2)$
$P E 1=P E Q+W E \cdot T 1$
$P M 1=P M D+W M \cdot T 1$
$P E 2=P E 1+W E \cdot T 2$
$P M 2=P M 1+W M \cdot T 2$
$P E 3=P E 2+W E \cdot T 3$
$P M 3=P M 2+W M \cdot T 3$
60
$\mathrm{PE} 4=\mathrm{PE} 3+W E \cdot T 4$
$\mathrm{PM} 4=\mathrm{PM}: 1+\mathrm{WM}^{\prime} \cdot \mathrm{T}_{4}$

$$
\begin{aligned}
& \text { (2n) }
\end{aligned}
$$

$$
\begin{aligned}
& \text { 1-1+1 }
\end{aligned}
$$

$$
\begin{aligned}
& 1 \times 12
\end{aligned}
$$


PREQ = (ENGP 'NUMENG/ETAPPU + FIXP)
foxa Required Power. In kilowatls:
 A engne numbers.
SHIPM - FIXMAS
SHIPM - FIXMAS + POWM + TENGM
PROPM - SHIPM

THRUST - ENGT NUMENG
ZTIA
A1 - TIRUSITTOTM
Intul acceleration:
TI.SC - BISECT(ZETA.CROAT.VO)


 orcelerration during spiral owt. SESC - Dulance Iravelled before escuping
RISC = radius at which craft atains eccape velocity. XNESC - No of
Ixape velocity.


RESC = RESC/6.37114SD6
SLSC $=$ SESCIO 37814SD

Proplient mass used dunng spral out.
11DC - TESC 3600.0210
DL:IV=-10D0•C•DLOG(1-PROPU/TOIM)

A-1.2 Program SPIRAL



 CONV $=16 \mathrm{~L}-19$. Converston from eV to $\mathrm{J}_{\text {sules. }}$ -
MION: $=0.6327 \mathrm{D}-26$ Propellant atome mass in tulograms
IONEFY $=0.8$ - Ionuation effikiency.
-
 DO $40 \mathrm{~K}=15$

ENGT -10 K
Thust per engene $(\mathbb{N})^{*}$
WRITE: S)


 $15 P=3000$
$150301 .=1.18$
$15 P=15 P+\$ 000$


[^0]
## APPENDIX 2

## Propulsion Calculations

For $T=5 \mathrm{~N}$ and $\mathrm{Isp}=10,300 \mathrm{~s}$ :

$$
\mathrm{T}=\mathrm{m}^{\prime} u_{0}=\mathrm{m} \prime \mathrm{gIsp}
$$

Mass flow rate per thruster
$\mathrm{m}^{\prime}=\mathrm{T} / \mathrm{gIsp}=4.9 \times 10^{-5} \mathrm{~kg} / \mathrm{s}$
where:

$$
\mathrm{g}=9.81 \mathrm{~m} / \mathrm{s}
$$

Propellant current
$\mathrm{J}_{0}=\mathrm{m} \mathrm{m}^{\prime} / \mathrm{m}_{0}=119.4 \mathrm{~A}$
where:
$\mathrm{e}=1.6 \times 10^{-19} \mathrm{C}$
$\mathrm{m}_{0}=\mathrm{MW} \times\left(1.66 \times 10^{-27} \mathrm{~kg} / 1 \mathrm{AU}\right)$
and $\mathrm{MW}_{\text {arga }}=39.944 \mathrm{AU}$
$\mathrm{m}_{0}=6.63 \times 10^{-26} \mathrm{~kg}$
Beam Voltage
$\mathrm{V}_{\mathrm{b}}=\left(\mathrm{m}_{\mathrm{o}} \mathrm{Isp}^{2} \mathrm{~g}^{2}\right) /\left(2 \mathrm{e} \eta_{\mathrm{i}}\right) \times\left[\left(1+2 \Gamma^{\prime \prime} / \Gamma^{\prime}\right) /\left(1+\Gamma^{\prime \prime} / \Gamma^{\prime}\right)\right]$ $=2424.2 \mathrm{~V}$
where:
$\Gamma^{\prime \prime} / \Gamma^{\prime}=0.1$
and $\quad \eta_{i}=0.8$
Beam Current

$$
\mathrm{J}_{\mathrm{b}}=\eta_{\mathrm{i}} \mathrm{~J}_{\mathrm{o}}\left[\left(1+2 \Gamma^{\prime \prime} / \Gamma^{\prime}\right) /\left(1+\Gamma^{\prime \prime} / \Gamma^{\prime}\right)\right]=104.2 \mathrm{~A}
$$

Beam Power

$$
\mathrm{P}_{\mathrm{b}}=\mathrm{J}_{\mathrm{b}} \mathrm{~V}_{\mathrm{b}}=252.6 \mathrm{~kW}
$$

Discharge Power

$$
\mathrm{P}_{\mathrm{d}}=\mathrm{J}_{\mathrm{b}} \mathrm{~S}_{\mathrm{d} \mathrm{p}}=20.84 \mathrm{~kW}
$$

where:

$$
S_{\phi p}=200 \mathrm{~J} / \mathrm{C}
$$

Neutralizer Power

$$
\mathrm{P}_{\mathrm{a}}=\xi\left(\mathrm{J}_{\mathrm{b}}+\mathrm{J}_{\mathrm{n}}\right)=1.69 \mathrm{~kW}
$$

where:

$$
\begin{array}{ll} 
& \xi=16 \mathrm{~V} \\
\text { and } \quad J_{\mathrm{a}}=1.5 \mathrm{~A}
\end{array}
$$

Total Power

$$
P_{c o t}=P_{b}+P_{d}+P_{n}=275 \mathrm{~kW}
$$

Electrical Efficiency

$$
\eta_{\mathrm{e}}=\mathrm{P}_{\mathrm{b}} / \mathrm{P}_{\mathrm{lot}}=0.9185
$$

Propellant Efficiency
$\eta_{\mathrm{p}}=\mathrm{J}_{\mathrm{b}} /\left(\mathrm{J}_{\mathrm{o}}+\mathrm{J}_{n}\right)=0.8619$
Total Efficiency
$\eta_{\text {toe }}=\eta_{e} \eta_{p} \gamma=0.714$
where:

$$
\gamma=0.95
$$

Radiator, Titanium
emissivity $\varepsilon=0.9$
$\mathrm{T}=600 \mathrm{~K}$
$A=q /\left(\varepsilon \sigma\left(T^{4}-T^{4}{ }_{\text {space }}\right)\right)$
$=370,000 /\left(0.9 * 5.67 \times 10^{-8}\left(600^{4}-4^{4}\right)\right.$
$=55.95 \mathrm{~m}^{2}$
In order to account for (a) the possibility of occasional partial "view - blockage" by the solar array structure and (b) possible degradation of the emissivity, the radiator area is increased to 80 $\mathrm{m}^{2}$.

## APPENDIX 3

## Solar Array Calculations

## A-3.1 Concentrator/Module Specifications

optical efficiency of module $-\eta_{\text {opt }}=96 \%$ packing efficiency of modules $-\eta_{p}=97 \%$ mismatch efficiency of modules - $\eta_{m}=93 \%$ solar intensity for AMO -I = $1372 \mathrm{~W} / \mathrm{m}^{2}$
efficiency of cell at AMO -- $\eta=27.7 \%$ at $T=115$ •
specific area ratio -- $\mathrm{M}_{4 \mathrm{~s}}=1 \mathrm{~kg} / \mathrm{m}^{2}$
power required -- $\mathrm{P}=3.5 \mathrm{MW}$

$$
\begin{aligned}
\mathrm{D} & =\eta_{\text {ope }} \times \eta_{\mathrm{p}} \times \eta_{\mathrm{m}}=0.866 \\
\mathrm{P}_{\mathrm{sm}} & =\frac{\eta_{\mathrm{T}} \times I \times \mathrm{I}}{\mathrm{M}_{\mathrm{s}}}=\frac{(.27)\left(1372 \mathrm{~W} / \mathrm{m}^{2}\right)(0.866)}{1 \mathrm{~kg} / \mathrm{m}^{2}} \\
& =320.8 \mathrm{~W} / \mathrm{kg} \\
\mathrm{P}_{\mathrm{s}} & =320.8 \mathrm{~W} / \mathrm{m}^{2} \text { at } 115 \cdot \mathrm{C}
\end{aligned}
$$

Area of total array -- $\mathrm{A}_{\mathrm{a}}=3.5 \mathrm{MW} / 320 \mathrm{~W} / \mathrm{m}^{2}$
Mass of total array -- $M=A_{s} \times M_{s e}=10,910 \mathrm{~m}^{2} \times 1 \mathrm{~kg} / \mathrm{m}^{2}$
Using the hexagonal design:

$$
\begin{aligned}
& s=d /\left(2 \times \cos 30^{\circ}\right)=0.57735 \times d \\
& b=\left(d \times \tan 30^{\circ}\right) / 2=0.28868 \times d
\end{aligned}
$$

$$
\begin{aligned}
\mathrm{A}_{\Omega} \Omega & =\mathrm{sd}+\mathrm{bd} \\
& =0.57735 \mathrm{~d}^{2}+\left(0.28868 \mathrm{~d}^{2}\right. \\
\mathrm{A}_{\mathrm{a}} & =1.73206 \mathrm{~d}^{2}
\end{aligned}
$$

Therefore:

$$
\begin{aligned}
& \mathrm{d}=10,910 / 1.73206=79.4 \mathrm{~m} \\
& \mathrm{~s}=45.8 \mathrm{~m} \\
& \mathrm{~b}=22.9 \mathrm{~m}
\end{aligned}
$$

## A-3.2 Concentrator Array Wiring Connections

$\mathrm{V}_{\text {max }}=200 \mathrm{~V}$
$\mathrm{V}_{\mathrm{oc}}=945 \mathrm{mV}$
Area of cell - $A_{c}=6 \mathrm{~cm} \times 6 \mathrm{~cm}=0.0036 \mathrm{~m}^{2}$
Number of cells per $\mathrm{m}^{2}-\mathrm{N}=1 / 0.0036=277.8$
Max number of cells per series $-\mathrm{m}_{\mathrm{z}}=\mathrm{V}_{\mathrm{max}} / \mathrm{V}_{\mathrm{cc}}=211$
Max number of resistance per series - $\mathrm{m}_{\mathrm{R}}=210$
Design number of modules per series -- $n_{1}=200$
Resistance of Al Gage \#7- $\mathrm{R}_{8}=2.68 \Omega / \mathrm{km}$
At Earth:
Power per cell -- $P_{c}=P_{s t} / N$

$$
=1.155 \mathrm{~W}
$$

Power per series -- $P_{s}=n_{s} \times P_{c}$

$$
=231.0 \mathrm{~W}
$$

Length of series connections -- $\mathrm{L}=(0.06 \mathrm{~m})(200)$

$$
=12 \mathrm{~m}
$$

Resistance of series -- $R=(0.012 \mathrm{~km} \times 2.68 \Omega / \mathrm{km})$

$$
=0.032 \Omega
$$

Current through each series wire $-\mathrm{I}=\mathrm{P}_{8} / \mathrm{V}=231 \mathrm{~W} / 200 \mathrm{~V}$

$$
=1.155 \mathrm{~A}
$$

Power loss per series -- $P_{L S}=I^{2} R=(1.155 A)^{2}(0.032 \Omega)$

$$
=0.0429 \mathrm{~W}
$$

Number of series in parallel $-n_{p}=P / P_{s}=15,153$
Power loss in series connections -- $P_{T S}=P_{\text {LS }} \times \eta_{p}$ $=650.10 \mathrm{~W}$

Length of series wire at $0.084 \mathrm{~A}=209,538.6 \mathrm{~m}$
Length of parallel wire length at $19.635 \mathrm{~A}=76,963.25 \mathrm{~m}$
Using Al Gage \#7 at $28.71 \mathrm{~kg} / \mathrm{km}$,
Estimated total wiring mass $=8,225 \mathrm{~kg}$
Power loss in series connections -- $\mathrm{P}_{\mathrm{TS}}=\mathrm{I}^{2} \mathrm{R}$

$$
\begin{aligned}
& =(19.64)^{2} \times 76.963 \times 2.68 \\
& =79,520 \mathrm{~W}
\end{aligned}
$$

Estimated total power loss -- $P_{T L}=P_{T P}+P_{T S}$
$=79,250+650$
$=80,170 \mathrm{~W}$

## A-3.3 Flat Panel Array Specifications

optical efficiency of module $--\eta_{\mathrm{opx}}=96 \%$
packing efficiency of modules -- $\eta_{p}=94 \%$ mismatch efficiency of modules $-\eta_{m}=95 \%$ solar intensity for AMO -I $=1372 \mathrm{~W} / \mathrm{m}^{2}$ efficiency of cell at AMO -- $\eta=21.7 \%$ at $T=25 \cdot \mathrm{C}$ specific area ratio -- $\mathrm{M}_{\mathrm{se}}=0.85 \mathrm{~kg} / \mathrm{m}^{2}$
power required -- $\mathrm{P}=3.5 \mathrm{MW}$
temperature dependency factor -- TD $=-0.63 \mathrm{~W} / \mathrm{m}^{2}-\cdot \mathrm{C}$

$$
\begin{aligned}
\mathrm{D} & =\eta_{\mathrm{op}} \times \eta_{\mathrm{p}} \times \eta_{\mathrm{m}}=0.857 \\
\mathrm{P}_{\mathrm{sm}} & =\frac{\eta_{T} \times I \times \mathrm{D}}{\mathrm{M}_{\mathrm{se}}}=\frac{(.217)\left(1372 \mathrm{~W} / \mathrm{m}^{2}\right)(0.857)}{.85 \mathrm{~kg} / \mathrm{m}^{2}} \\
& =300.2 \mathrm{~W} / \mathrm{kg} \\
\mathrm{P}_{\mathrm{se}} & =255.2 \mathrm{~W} / \mathrm{m}^{2} \text { at } 25 \cdot \mathrm{C}
\end{aligned}
$$

Therefore, at $\mathrm{T}=115^{\circ} \mathrm{C}$ :

$$
\begin{aligned}
P_{s \mathrm{se}} & =255.2 \mathrm{~W} / \mathrm{m}^{2}+\mathrm{TD} \times 25 \cdot \mathrm{C} \\
& =198.5 \mathrm{~W} / \mathrm{m}^{2}
\end{aligned}
$$

Area of total array -- $\mathrm{A}_{\mathrm{a}}=3.5 \mathrm{MW} / 198.5 \mathrm{~W} / \mathrm{m}^{2}$

$$
=17,632 \mathrm{~m}^{2}
$$

Mass of total array $-\mathrm{M}_{\mathrm{M}}=\mathrm{A}_{2} \times \mathrm{M}_{\mathrm{s}}=17,632 \mathrm{~m}^{2} \times 0.85 \mathrm{~kg} / \mathrm{m}^{2}$ $=14,987 \mathrm{~kg}$

Using the hexagonal design:

$$
\begin{aligned}
& s=d /\left(2 \times \cos 30^{\circ}\right)=0.57735 \times d \\
& b=\left(d \times \tan 30^{\circ}\right) / 2=0.28868 \times d
\end{aligned}
$$

$$
\begin{aligned}
\mathrm{A} \Omega & =\mathrm{sd}+\mathrm{bd} \\
& =0.57735 \mathrm{~d}^{2}+\left(0.28868 \mathrm{~d}^{2}\right. \\
\mathrm{A}_{\mathbf{a}} \quad & =1.73206 \mathrm{~d}^{2}
\end{aligned}
$$

Therefore:

$$
\begin{aligned}
& \mathrm{d}=10,910 / 1.73206=93.0 \mathrm{~m} \\
& \mathrm{~s}=53.7 \mathrm{~m} \\
& \mathrm{~b}=26.85 \mathrm{~m}
\end{aligned}
$$

## A-3.4 Flat Panel Array Wiring Connections

$V_{\text {max }}=200 \mathrm{~V}$
$\mathrm{~V}_{\mathrm{oc}}=945 \mathrm{mV}$
Area of cell -- $\mathrm{A}_{\mathrm{c}}=2 \mathrm{~cm} \times 2 \mathrm{~cm}=0.0004 \mathrm{~m}^{2}$
Number of cells per $\mathrm{m}^{2}-\mathrm{N}=1 / 0.0004=2500$
Max number of cells per series -- $\mathrm{n}_{\mathrm{i}}=\mathrm{V}_{\mathrm{max}} / V_{\mathrm{oc}}=211$
Max number of resistance per series - $n_{R}=210$
Resistance of Al Gage \#7 -- $\mathrm{R}_{\mathbf{1}}=2.68 \Omega / \mathrm{km}$

## At Earth:

Power per cell -- $P_{c}=P_{s} / \mathbb{N}$

$$
=0.079 \mathrm{~W}
$$

Power per series -- $P_{s}=n_{s} \times P_{c}$

$$
=16.7 \mathrm{~W}
$$

Length of series connections $-\mathrm{L}=(0.02 \mathrm{~m})(210)$

$$
=4.2 \mathrm{~m}
$$

Resistance of series -- $\mathrm{R}=(0.0042 \mathrm{~km} \times 2.68 \Omega / \mathrm{km})$

$$
=0.0113 \Omega
$$

Current through each series wire $-\mathrm{I}=\mathrm{P} / \mathrm{V}=16.7 \mathrm{~W} / 200 \mathrm{~V}$

$$
=0.0835 \mathrm{~A}
$$

Power loss per series -- $P_{L S}=I^{2} R=(0.084 A)^{2}(0.0113 \Omega)$

$$
=0.00008 \mathrm{~W}
$$

Number of series in parallel -- $n_{p}=P / P_{s}=209,970$
Power loss in series connections -- $\mathrm{P}_{\mathrm{TS}}=\mathrm{P}_{\mathrm{LS}} \times \mathrm{n}_{\mathrm{p}}$

$$
=16.7 \mathrm{~W}
$$

Length of series wire at $0.084 \mathrm{~A}=654,234 \mathrm{~m}$
Length of parallel wire length at $20 \mathrm{~A}=277,830 \mathrm{~m}$
Using Al Gage \#7 at $28.71 \mathrm{~kg} / \mathrm{km}$ :
Estimated total wiring mass $=11,100 \mathrm{~kg}$
Power loss in series connections -- $\mathrm{P}_{\mathrm{TS}}=\mathrm{I}^{2} \mathrm{R}$

$$
\begin{aligned}
& =(20)^{2} \times 277.83 \times 2.68 \\
& =297,834 \mathrm{~W}
\end{aligned}
$$

Estimated total power loss -- $P_{T \mathrm{~L}}=\mathrm{P}_{\mathrm{TP}}+\mathrm{P}_{\mathrm{TS}}$

$$
\begin{aligned}
& =297,833.3+16.7 \\
& =297,850 \\
& =80,170 \mathrm{~W}
\end{aligned}
$$

## APPENDIX 4

## Control System

The mechanics of a solar-powered electric propulsion vehicle poses important demands on the vehicle's control system, which stem from two main maneuverability requirements: 1) the solar array panels must be directed so as to maximize the amount of sunlight captured by the solar cells; and 2) attitude adjustments are necessary in order to perform orbital maneuvering.

The most effective spacecraft configuration allows one degree of freedom between the solar arrays and main body of the ship. The extra degrees of freedom needed for attitude control are achieved by roll and yaw of the spacecraft. As the craft spirals out, the longer orbit period permits the angular acceleration of the solar panels to decrease. Therefore, the pitch moments are the most significant orientation maneuvers in the initial orbit.

These conditions allow us to predict that the most drastic moment and attitude adjustments will occur during the low earth orbit height of 400 km .

Three main reaction systems were considered for the attitude controllers: 1) Reaction Wheels, 2) Control Moment Gyros, 3) Engine Thrusters.

Comparison of the Reaction Wheel (RW) with the Control Moment Gyro (CMG), indicated that a cluster of CMG's offers significant advantages over a reaction wheel when dealing with a large vehicle to control. CMG efficiency is better since the CMG wheel operates at one speed for which efficiency can be optimized [1]. Larger maximum moments can be achieved at constant momentum rates. Finally, the reaction wheel's maximum rates tend to become non-linear as it approaches saturation; thus, complex feedback is necessary to control
reactions.
While lighter in weight, thrusters appear to be less suitable for this application because a very slow rate of angular adjustments required (as the ship spirals out or in) call for nearcontinuous minute thruster impulses, perhaps increasing complexity and demands on reliable and long-life performance.

Therefore, a gyroscopic control system appears to be a preferred solution to our control needs. This system will be supplemented by a system of small thrusters (RCS). Momentum exchange devices can handle cyclic torques on a continuous basis over a limited period of time. Consequently, the external moments applied to the vehicle will cause the controllers to reach maximum capacity; thus becoming saturated. This condition requires the removal of momentum from the device. This may be accomplished at convenient times during the mission. For our case, the optimum desaturization period will be after the spiral-out from Earth, during the geocentric transfer from Earth to Mars. This interval is over a very long orbit period; there will be little solar array rotation involved as well as smaller disturbances from the tangential thrust vector.

As discussed later, our system has been designed so that saturation will not occur during the spiral-out/in phases. When desaturization occurs, it will be accomplished by small electrical thrusters of the RCS system and countering each CMG gimbal controller to a prescribed gimbal angle on a momentum feedback loop (Figure A-4.1a).

From the mass moment of inertia of the solar arrays, $\mathrm{I}_{\mathrm{xx}}=6.19 \times 10^{6} \mathrm{~kg} \mathrm{~m}^{2}(\mathrm{x}-\mathrm{x}=$ solar arrays axis of rotation) and the initial angular velocity of the ship in LEO, $\varpi_{i} 1.131 \times 10^{-3} \mathrm{sec}^{-1}$, we can estimate that the total spiral-out angular momentum change required will be about 7,000 $\mathrm{kg}-\mathrm{m}^{2} / \mathrm{s}$. A single Sperry M4500 CMG [2] pmuides an angular momentum capacity of $6,101 \mathrm{~kg}-$ $\mathrm{m}^{2} / \mathrm{s}$, thus it would be able to control the solar arrays orientation during most of the spiral out

Figure A-4.1: Control Loop Diagram
without becoming saturated.
To provide maximum flexibility and redundancy we selected a system using a "six pac" configuration which consists of three double gimballed control moment gyros. This arrangement will allow for 100 percent momentum utilization about 3 primary axes. Additionally, this configuration provides 100 percent redundancy of total control; two CMG's can be in an active state while the third is not operating.

Navigation software (beyond scope of this study) will convert the steering commands-Vehicle Control Law, CMG Control Law-issued by the ADCOLE Model 20470 V-SLIT sun sensor assembly into drive commands [7]. Each CMG gimbal pivot will be directly coupled with a D.C. motor by a rate generator which drives the individual momentum vectors of each CMG. Thus changing the angular momentum and causing reaction moments on vehicle. Actual control loop functions are as follows:

1) Position Sensor (Star Sensor)[5]

Measures error between actual and desired vehicle attitude
2) Vehicle Control Law

Formulates vehicle moment command
3) CMG Control Law

Processes command moment and provides input commands to the CMG gimbal servos
4) CMG Servos

Drive individual momentum vectors of each CMG, thus changing angular momentum

## 5) Vehicle Dynamics

Actual change of vehicle attitude
As discussed above, a saturation of the CMG will not occur during the spiral phases. Desaturization, can then be accomplished during the heliocentric transfer phase. At that point, the same D.C. motor that was used to drive the CMG servos can be used to desaturate the
system.

## CMG PROJECTED SPECIFICATIONS:

Double Gimbal<br>Angular momentum $\quad 6,000 \mathrm{~kg}-\mathrm{m}^{2} / \mathrm{sec}$<br>Total system mass<br>700 kg<br>Total power required<br>1 kW

## References

1. B.J. O'Conner, L.A. Morine, "A Description of the CMG and its Application to Space Vehicle Control", AIAA 67-589. 1967.
2. Sperry Flight Systems. Honeywell/Sperry Corporation. DOC. 61-1720-08-00, 1982.
3. H.S. Oh, S.R. Vadali, "Feedback Control and Steering Laws for Spacecraft Using Single Gimbal Control Moment Gyros". AIAA 89-3475-CP. 1989.
4. Zondervan,et al,Guidance, "Navigation, and Control Trades for an Electric Orbit Transfer Vehicle." AIAA 90-3458-CP. 1990.
5. Ninomiya, et al,"High Accuracy Sun Sensor Using CCD's." AIAA 88-4180-CP. 1988.
6. Potts, "The Mars Observer Spacecraft." AIAA 89-0255. Jet Propulsions Laboratory, 1989.
7. ADCOLE Aerospace Products, Excerpts from general catalog.

## APPENDIX 5 <br> Communications and Navigation Systems

## Nomenclature

| b/s or bps | bit per second |
| :--- | :--- |
| CD\&H | Command, Data, and Handling |
| dB | Decibel |
| DSN | Deep Space Network |
| $\mathrm{E}_{6} / \mathrm{N}_{0}$ | Energy per bit to spectral-noise-density ratio |
| Hz | Herz |
| m | meter |
| SSA | Solid State Amplifier |
| Sr. | steridian (solid angle) |
| W | Watt |

Specific performance and physical parameters of the communications and navigation systems are tabulated in Tables A-5.1 through A-5.4. The information in these tables incorporate projections of technological accomplishments by 2015 such as fabrication of high temperature superconducting surfaces for antenna dishes and improvements in radiation hardness of the components. All calculations were made assuming that modulation and encoding techniques would be improved such that $\mathrm{E}_{\mathrm{o}} / \mathrm{N}_{\mathrm{o}}$ could take the following values: 1) 3.0 dB for X -band telemetry, 2) 3.5 dB for Ka-band telemetry, 3) 5.0 dB for X -band commands, and 4) 6.0 dB for Ka-band commands. All calculations also provide a safety margin of at least 2.0. Table A-5.1 summarizes the chosen communications and navigation systems.

## A-5.1 The RF Subsystem

The RF subsystem employs six antennas, two 1.5 m high gain antennas and four 0.5 m omnidirectional antennas. Two high gain antennas were chosen for redundancy. They will be placed along the symmetric axis of the ship, between the solar arrays to minimize signal blockage by the arrays. The omnidirectional antennas will be placed as perpendicular pairs on opposite ends of the ship. During normal operations only two of these antennas will be functioning, one
for reception and one for transmission. In the case of an emergency, all four will be functioning, two for transmission and two for reception. The omnidirectional antenna operates at X-band ( 8400 MHz ) frequency with a transmitting power output of $2.5,5.0$, or 40.0 Watts. It delivers telemetry and receives and sends commands during the Earth spiral phase of the mission. Telemetry will be transmitted to the 26 m DSN antenna at 5.0 W and $2 \mathrm{Mb} / \mathrm{s}$. Commands will be transmitted at $10 \mathrm{~kb} / \mathrm{s}$ and 2.5 W to the 26 m DSN antenna. The omnidirectional antenna will also be used to receive command data during emergency situations. The Command Data and Handling (CD\&H) system alerts the spacecraft of an emergency if any of the following criteria are met: 1) Unplanned failure of solar array to properly generate power. This usually occurs when either the attitude control system or the motors on the solar arrays fail, or when some phenomena has caused the ship to tumble, 2) the temperature of the communications system drops below $0 \cdot \mathrm{C}$, or 3 ) the spacecraft has not received and Earth or Mars transmission in 3 days. (When the spacecraft orbits Mars near Mars/Earth opposition, the spacecraft will be commanded to increase this requirement to 19 days since communications are not possible for 17 days during opposition.) In emergency cases, the omnidirectional antenna is used because its wide beam angle $\left(90^{\bullet}\right)$ should allow it a greater chance of contacting an Earth or Mars station than the high gain antenna, with a beam angle of $0.875^{\circ}$. The omnidirectional may also transmit during emergencies. However, the low gain of 0 dB allows the antenna to transmit only 10 bps of commands at 40 W to the 64 m antenna from opposition if one wishes to maintain the standard bit error rate of $1 \times 10^{-5}$ for commands. If one uses and array of DSN antennas (a $64 \mathrm{~m}-34 \mathrm{~m}$ $34 \mathrm{~m}-34 \mathrm{~m}-34 \mathrm{~m}$ array, for example) to track the ship in case of an emergency, then the 100 bps transmission can be maintained from as far as 1.8 AU .

One high gain antenna will receive and transmit commands in X-band ( 8.4 GHz ) during the interplanetary phase of the mission. The other will deliver telemetry at Ka-band ( 32 GHz ).

Commands will be carried at $10 \mathrm{~kb} / \mathrm{s}$ and 5 W of power, telemetry will be sent at $2 \mathrm{Mb} / \mathrm{s}$ and 12 W of power. All will be transmitted to a 34 m DSN antenna. Should failure occur in the Kaband transponder, the X-band transponder will send the telemetry at $10 \mathrm{~kb} / \mathrm{s}$ and 40 W of power to the 64 m DSN antenna. Both transponders will have connections to both antennas in case an antenna fails, and both antennas are capable of handling four beams of radio signals, two transmitting and two being received. During the near Mars phase of the mission, it will be assumed that the spacecraft will communicated with antenna stations on or circling Mars. It is also assumed that these stations will have an antenna of at least 8 m in diameter and are capable of transmitting only in X -band. These transmissions will begin at around 0.05 AU from Mars using the high gain antennas. Telemetry will be sent at $2 \mathrm{Mb} / \mathrm{s}$ and 40 W of output power. Commands will be delivered at $1 \mathrm{~kb} / \mathrm{s}$ and 2.5 W of power. When the ship begins to spiral in or executes a spiral out from Mars, the high gain antennas will become inactive and the omnidirectional antennas will be used for communication.

In both near Earth and near Mars transmissions, command and telemetry phases of communication are separated. Usually, commands are transmitted and received on a subcarrier along with the telemetry data. However, the use of the subcarrier demands an extra transponder. This adds weight, power drain, and complexity. If separation of the two data streams is not feasible, the Ka-band transponder will transmit command data as a subcarrier through one omnidirectional antenna. The X-band transponder will transmit telemetry through one of the other antennas. The weight will increase by no more than 1 kg , but the power required during transmission will increase by 14.0 Watts to pay for the decrease in transmission time. This system can transmit $2 \mathrm{Mb} / \mathrm{s}$ telemetry and $10 \mathrm{~kb} / \mathrm{s}$ command.

Maximum bit rates were designed for the possible addition of scientific instruments as secondary functions for a mission. The bit rate of $2 \mathrm{Mb} / \mathrm{s}$ can handle color video transmission.

The $10 \mathrm{~kb} / \mathrm{s}$ telemetry rate is meant to be used only in the emergency caused by the failure of the Ka-band transponder, thus only data from ship sensors will be transmitted at this rate. The gimbal pointing assembly has a pointing error of $0.1^{\cdot}$. It has the ability to change the direction of the antenna by nearly pi Sr. with respect to the axis of the spacecraft. All gimbal systems are run by 5 W motors. Optionally, the omnidirectional antennas may also have a gimbal system that allows pi Sr. of movement. This will add eight kilograms of weight and the possibility of 10 to 20 Watts of additional power during transmission. All transponders use solid state amplifiers (SSA) for boosting transmission signal power. Both transponders also are fitted with bandpass filters to improve performance.

## A-5.2 The CD\&H Subsystem

The heart of the command data and handling subsystem is a low power 32 bit processor. This makes multitasking possible and will increasing the speed and efficiency of the CD\&H subsystem. In turn, and indirectly, this may increase the efficiency and lifetime of many ship components such as the engines and the arrays. It also decreases the chances of a catastrophic failure because it decreases reaction times to unusual sensor readings. No more than 4 MB of RAM should be required to activate commands and manipulate incoming data. The $\mathrm{CD} \& \mathrm{H}$ system places the data from the ship's sensors onto a read/write laser disk for storage until the ship is in the right position to transmit it. The magnetic disk can hold 1 TB of data which, assuming daily transmission, provides more that adequate storage. All software required by the communications system alone could probably fit in 4 k ROM; however, because of the uncertainty of the complexity of programming required to run other ship systems (particularly the engine array), 128 k ROM has been allocated to contain a copy of all software. The system includes and event sequencer to hold series of delayed commands transmitted from Earth until
activated by a "macro command" transmitted from Earth. The CD\&H subsystem uses analog switches for temperature control and requires analog/digital modulators to convert the analog output of certain ship instruments (eg. DRIRU) to digital form for transmission.

## A-5.3 Navigation and Batteries

The ship contains three star trackers for precise determination of position over three axes and for purposes of redundancy. These components will lie along the truss structure near the high gain antennas to minimize blockage of their view of surrounding space. The ship also holds three dry inertial measurement units (DRIRU). These systems are self-redundant. One is used for keeping track of the orientation of the spacecraft; the other two are placed one on each solar array structure to monitor the orientation of the arrays. Solar sensors determine the orientation of the solar array to toward the sun. There are eight sensors placed on the array. The other three are placed along the ship's truss structure to aid in calculating the attitude of the spacecraft. The batteries are employed for use during times of peak power drain and during an emergency caused by solar array failure. They are capable of delivering up to 2 kW of power for approximately 1.3 hours. When not in use, the batteries will be charging. Near Earth, where the period in shadow will be longest during the mission and where solar array power is greatest, the batteries will be under full charge. During the interplanetary trajectory, when no shadow should be encountered, the batteries will charge in trickle mode. Around Mars, the high inclination of the orbit keeps the ship within Mars' shadow for a very short period of time, therefore the batteries need only charge at the intermediate level.

## A-5.4 Power Requirements

Power required by specific components and the total power for navigation and battery systems is given in Table A-5.2. The various power requirements for typical communications
operations are listed in Table A-5.3.

## A-5.5 Mass of System

The masses of the subsystems for communications and navigation are summarized in Table A-5.4.

Table A-5.1 Parameters of the Communications and Navigation Systems

## Telemetry System:

X-band transmission frequency: 8.4 GHz
X-band reception frequency: 7.1 GHz
Ka-band transmission frequency: 32 GHz
Ka-band reception frequency: 28 GHz
1.5 m antenna gain (at 8.4 GHz ): 39.1 dB
1.5 m antenna gain (at 32.0 GHz ): 50.68 dB
1.5 m antenna efficiency: 0.6 (with high temperature superconducting surface.)

Omnidirectional antenna gain: 0 dB
High gain antenna transmission : 2.5/5.0/12.0/15.0/40.0
Omnidirectional antenna output: $2.5 \mathrm{~W} / 5.0 \mathrm{~W} / 40.0 \mathrm{~W}$
Bit error rate for telemetry: $5 \times 10^{-3}$
Bit error rate for command: $1 \times 10^{-5}$
Bit rates for telemetry transmission: $2 \mathrm{Mb} / \mathrm{s}$ (Video/Ship Systems), $4.6 \mathrm{~kb} / \mathrm{s}$ (Ship systems, Ka-band transponder failure emergency)
Bit rates for command transmission: $10 \mathrm{~kb} / \mathrm{s}, 1000 \mathrm{bps}, 100 \mathrm{bps}, 10 \mathrm{bps}$

## Command Data and Handling System

16 bit radiation hardened processor
4 MB RAM
1 Terabyte laser disk storage
Software encoded on 128 k ROM
Event Sequencer
Analog Switches
Analog/Digital Modulators

## Navigation Systems

3 Star Trackers
3 Dry Inertial Reference Units
11 Sun Sensors

## Batteries

Capacity: 39.0 Ampere-Hours
Storage: 1316.0 Watt-Hours
Voltage: 27 V
Current: 1 Ampere
Voltage Required for Communications System: 287 VDC
Operating Temperature Range for Communications System: 0 to $40 \cdot \mathrm{C}$
Operating Temperature Range for Batteries: 0 to $20^{\circ} \mathrm{C}$
2. The different power numbers shown for the DRIRU are the powers required for operation of $1 / 2 / 3$ gyros respectively.
3. The three normal operation powers for the batteries correspond to trickle/intermediate/full charge rates respectively.
4. The navigation and battery power requirements includes only those values listed under navigation subsystem and batteries. It is assumed that all three gyros in the DRIRU will be working during normal operation. The three different numbers under normal operation refer to different battery charge states (see number three).

Table A-5.3 Communications Power Requirement Totals for Possible Operations

Max power required for near-Earth operation: 360.0 W
Components in use: X -band transponder (transmitting at 40 W ), CD\&H subsystem.

Nominal power for near-Earth transmission operation: $340.0 \mathrm{~W} / 134.0 \mathrm{~W}$
Components in use: As above. Second value denotes X-band command transmission.

Normal operations power for near-Earth operation: 120.0 W
Components in use: As above, except X-band transponder in receive mode.
Max power for interplanetary operation: 225.0 W
Components in use: X-band transponder transmitting at $5.0 \mathrm{~W}, \mathrm{Ka}$-band transponder transmitting at 12.0 W , Gimble pointing system motors running, and CD\&H systems running.

Nominal power for interplanetary transmission operations: 195.0 W Components in use: As above, except gimble system not running.

Normal operations power for interplanetary transmission operation: 145.0 W Components in use: As above, except X-band and Ka-band transponders in receive mode only.

Max power for Mars orbit operations: 225.0 W
Components in use: Same as for interplanetary operation. Used only if Mars communication systems break down.

Nominal power for Mars orbit transmission: $340.0 \mathrm{~W} / 134.0 \mathrm{~W}$
Components in use: As for "nominal power for near-Earth transmission."
Normal operations power for Mars orbit operations: 120.0 W
Components in use: As above, except the X -band transponder is in receive mode only.

Emergency (solar array failure) operations: 580.0 W Components in use: X-band transmission at 40 W on two omnidirectional antennas, CD\&H subsystem.

NOTE: Max power calculations use the maximum power values for components given in the table. Normal operations and nominal power calculations use numbers listed under "Normal operations power."

Table A-5.4 Communications \& Navigation Systems Weight
Weight (kg.)
Antenna Subsystem
1.5 m High Gain Antenna (2) ..... 6.5
Gimbal Pointing System (2) ..... 4.0
Antenna Feed Array (2) ..... 4.0
0.5 m Omnidirectional Antenna (4) ..... 2.0
ANTENNA SUBSYSTEM WEIGHT ..... 16.5
Telemetry Subsystem
X-band Transponder ..... 15.5
Ka-band Transponder ..... 18.0
TELEMETRY SUBSYSTEM WEIGHT ..... 33.7
CD\&H Subsystem
CD\&H Subsystem Weight ..... 15.0
TOTAL WEIGHT OF COMMUNICATION SUBSYSTEM: ..... 65.0
Navigation Subsystem
Star Tracker (3) ..... 24.0
Dry Inertial Reference Unit (3) ..... 48.0
Sun Sensors (11) ..... 1.0
TOTAL WEIGHT OF NAVIGATION SUBSYSTEM: ..... 73.0
Batteries (4)
Battery Weight ..... 80.0
TOTAL WEIGHT OF NAVIGATION SUBSYSTEM AND BATTERIES: ..... 153.0

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[^0]:    A-1.3 Program ORBITB.BAS
    Prcgram orbitb.bas
     T:PRINT" " OF ENGINES ":ENG:" W/":TG:" NEWTONS THPLST EACH"
    T" MASS FLOW RATE (NET) KG/S: ":MDOT IMLEO (AFTER SPIRAL-OUT BURN): ":MASS:" KG".PRI:T Loop to calculate progressive orbits are acceleration increases
    due to the loss of propellant mass

    AT $=T \emptyset^{\prime} E N G / M A S S \quad: \ln m / s^{2} 2$
    $A T=A T \cdot T U V^{2} 2 / A U$ $R=R \emptyset /\left(1 *-2 w^{\prime} A T \cdot D S / V Q^{-} 2\right)$
    $V 2=2 w^{\prime} D S \cdot A T+(2 * / R-1 * / R \theta)$
    $V A V E=(V G+V) / 2$
    $D R=R-R g$
    $D V=V-V G$
    DVTOT $=$ DVTOT $+A B S(D V)$
    Calculate flight path angle. ang momentum. energy
    eccentricity. aphelion. perihelion. and time for increment
    
    ENERGY=V $/ 2 m-1 / 2 / R$
    $A=-1 \# /(2 \| \cdot E N H: R G Y)$

