https://ntrs.nasa.gov/search.jsp?R=19920024981 2020-03-17T09:56:04+00:00Z NASA-CR-190739 1EHUS 41 117782 117782 P.11 N92-34225 0117782 Unclas DEVELOPMENT OF A VERSATILE LASER 63/36 LIGHT SCATTERING INSTRUMENT ۷ ۵ (Case Western Reserve Univ.) 17 r by DEVELOPMENT OF LIGHT SCATTERING William V. Meyer VERSATILE LASER LIGHT SCATT INSTRUMENT Final Technical and Rafat R. Ansari NASA Lewis Research Center Case Western Reserve University (NASA-CR-190739) Cleveland, Ohio

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Case Western Reserve University Cleveland, OH 44106

LASER LIGHT SCATTERING CAPABILITIES

A modular approach for assembling a laser light scattering instrument from basic building blocks (Optical Lego) will allow us to implement dynamic light scattering, static light scattering, dynamic depolarized laser light scattering, etc. in a manner ideal for each of NASA's Principal Investigators. Dvnamic light scattering, as used in this report, is a digital technique for measuring and correlating the intensity fluctuations in the scattered light that arise from the Brownian motion of particles in the sample. This gives the diffusion coefficient from which particle sizes in the range of 3 nanometers to above 3 microns are determined. Static light scattering - the measurement of time-averaged intensity scattered by dispersions of particles and macromolecules (e.g. polymers, proteins, micelles, microemulsions, etc.) - is an important tool for the determination of particle structure, weight-average molecular weight and particle interactions. Dynamic depolarized laser light scattering is different from regular dynamic light scattering in that it examines the weak, horizontally polarized scattered light coming from the sample. This depolarized signal contains dynamic and structural information which is not otherwise readily available.

JUSTIFICATION

NASA Lewis Research Center is developing and coordinating the development of a versatile light scattering instrument for use in microgravity. Laser Light Scattering (LLS) will be used in microgravity to measure microscopic particles of 30 angstroms to above 3 microns. Since it is an optical technique, LLS does not affect the sample being studied.

Critical phenomena, nucleation, spinodal decomposition, gelation, aggregation, diffusion, etc. are influenced by gravity and can be better studied with LLS in a microgravity environment. LLS can enhance protein crystal growth experiments which need quantitative information about the growth process and an indication of the onset of nucleation. The above experiments and many others have already been presented to NASA in the form of "pre-proposals", which were solicited by this ATD project from experts in the field desiring an opportunity to use LLS in microgravity. A number of very well qualified investigators have identified a group of basic science experiments which require microgravity for their performance and which will be enabled by LLS instrumentation. These potential Principal Investigators (PIs) will be responding to the Laser Light Scattering section of NASA's forthcoming Announcement of Opportunity (AO) when it is released at the end of 1990.

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PROGRAM OBJECTIVES

The major objectives of this Advanced Technology Development program are to:

- 1. Define and ascertain instrument needs of potential PIs.
- 2. Modify a commonly available instrument to make it compact, rugged, power efficient, and microgravity-ready.
- 3. Make the LLS instrument modular to allow it to be easily reconfigured and optimized for a wide range of experiments which await the forthcoming Microgravity Fluids NASA Research Announcement (NRA or AO).
- 4. Automate and enhance data taking and analysis for the modular LLS instrument.
- 5. Collect data sets on the traditional state-of-the-art, roomsize instrument in our laboratory for evaluation of the hardware and software under development and to further develop in-house expertise in preparing and analyzing different classes of samples.
- Test miniature LLS modules attached to fiber optic probes with protein crystal growth experiments in collaboration with Professor Wm. Wilson of Mississippi State University / MSFC.
- 7. Simulate vibration and low-gravity performance.
- 8. Test back-scatter fiber optic probes.
- 9. Acquire and evaluate a simultaneous multiangle LLS instrument.
- 10. Provide assistance to the STDCE (Surface Tension Driven Convection Experiment) flight hardware project when it is in need of LLS diagnostics. While not a specific objective of the work, since we have the capability we have taken on this other program of interest of MSAD.

APPROACH

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NASA Lewis Research Center is providing and coordinating the technology for placing a compact Laser Light Scattering instrument in a microgravity environment. This will be accomplished by defining and assessing user requirements for microgravity experiments, coordinating needed technological developments and filling technical gaps. This effort is striving to brassboard and evaluate a miniature multiangle LLS instrument. We expect to see this ATD effort become a flight hardware development project once NASA's Microgravity Fluids AO is released and the many LLS proposals we have solicited are then submitted in response to NASA Headquarters. FY90 funding supports one of the two Lewis people on the project and has forced a freeze in the developments supported by grants and in some of the hardware advances. We do believe that the reader will still be impressed with what follows. £ ...

PROGRESS (Elaborates Program Objectives)

1. NASA Lewis held an Advanced Technology Development (ATD) Laser Light Scattering Workshop in September of 1988. Major LLS contributors, manufacturers, and potential PIs from around the world attended the workshop. Many issues were settled in formal and informal discussions during the two day workshop. A draft version of the Workshop Proceedings was edited and issued soon after the workshop and a final, 308 page version of the Workshop Proceedings (NASA CP-10033) was published August 1989. : EEN

In January of 1990 another meeting of potential LLS PIs took place in Washington DC. At this meeting, more than a dozen researchers discussed and shared their ideas and proposal outlines.

We exhibited the back scatter fiber optic probes discussed below at the Microgravity Fluids workshop in mid 1990. This gave us a chance to discuss the LLS project with additional potential PIs and to excite the LLS community with a live demonstration of technology that is on the cutting edge.

The forthcoming Microgravity Fluids Announcement of Opportunity (December 1990) contains a section for laser light scattering which will solicit not only proposals for microgravity LLS experiments, but additional PI instrument parameters and details of how the PIs plans to store samples and initiate experiments (e.g. - chemically, thermally, through photopolymerization, etc.).

2. Dr. Robert G.W. Brown has served as a hardware consultant to this project. He invented and has patented many of the advances needed for making the miniature modules and improving the detectors used in LLS. He has provided detailed lists of the necessary miniature module specifications and ways of insuring these specifications have been met. The core of these specifications are provided for the reader in Appendix 1. At the January Washington DC meeting, Dr. Brown, associated with Royal Signals and Radar Establishment (RSRE), showed RSRE developments in miniature LLS systems in a presentation to potential PIs. Some of these systems had been designed specifically for NASA needs, and they are about a year ahead of the rest of the world. This meeting gave potential PIs a chance to look ahead at what is technologically possible and it gave NASA and RSRE important feedback on how well the needs of the potential PIs are being met.

3. Single angle detector and laser modules have been breadboarded, but have yet to be finished and delivered to the LLS laboratory at NASA Lewis for evaluation. This delay is the result of a reorganization of Defense Technology Enterprises (DTE), which was the commercial arm of RSRE. The Laser Dynamics group of DTE has since broken off and formed their own company (in alliance with a large parent company) and should be able to deliver the modules before the year is out. We intend to check the miniature instrumentation by repeating measurements which are difficult to make with traditional LLS instrumentation. Details follow in section 5.

4. The maximum likelihood correlator algorithms being developed by Professor Robert Edwards of Case Western Reserve University (CWRU) for this project are in their second year of development. The analysis portion of the prototype software for extracting particle size distributions, errors, and goodness of fit from single angle Quasi Elastic Light Scattering (QELS) data and later automating the LLS instrument is nearly completed and early results are very promising. The prototype code was written in APL, which allowed it to be rapidly modified. However, the APL code now needs to be translated to a language that can be compiled to give rapid results. The code is in the process of being documented and one of our next programming tasks at the LLS laboratory at NASA Lewis will be recoding this program using a fast compiled language. This portion of the project is on a no cost extension while we await funding for its completion. The addition of multiangle algorithms will give more information and promises to decrease the stiffness of the numerical problem and make it easier to fit histograms to the data sets. The development of the multiangle code by Professor Cheung at the University of Akron, in collaboration with Professor Edwards, is

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going well, and is now consistently extracting bimodal and trimodal particle size distributions from the data libraries discussed in the next section. Professor Cheung's work is also on a no cost extension since the ATD project was only funded for one of its two Lewis scientists (the other's salary was provided by the MMSL at NASA Lewis) and hence could not offer continued support for its grant research.

The core of correlator control program for single detector input is now finished and is being used at Lewis to automatically take data, analyze it, and print out the results. The analysis algorithms are being enhanced at Lewis with additional Mie algorithms. These allow the experimenter to account for scattering asymmetries that arise with samples larger than 1/10 the wavelength of the incoming light. Two out of three Mie routines planned for the enhancements are now coded and running. The automatic LLS instrument parameter setup routines provided by the maximum likelihood algorithms remain to be added to the correlator program.

5. LLS experiments conducted in our laboratory have given us many data sets for analysis and testing of both the software and hardware being developed for this ATD effort. While this library of data is needed for evaluating the modular miniaturized hardware in its different configurations, it also perfects in-house expertise in challenging LLS experiment areas. An eight page listing of ATD LLS experiments prior to March of 1990 was previously submitted to Stuart Glazer at NASA Headquarters. We have attached several tables giving an overview of our more recent dynamic LLS work with single, bimodal, trimodal, and multimodal distributions made in-house from standards. Additional work in static light scattering on polymeric systems is provided Zimm/Berry plots and has been included in Appendix 2. These data libraries have been and will continue to be essential for evaluating both our multiangle algorithm developments and the miniature LLS modules when they arrive.

6. Our collaboration with Professor Wm. W. Wilson (MSFC) has been on hold while we await the arrival of our miniature LLS modules. We have kept in touch and have shared some recent experimental published work from the Journal of Crystal Growth with Professor Wilson discussing the detection of nucleation in protein crystals using LLS.

7. An electronic vibration isolation system with seismic control (a next generation EVIS design from Newport) has been installed and is ready to test the miniature single angle LLS modules when they are delivered.

8. Significant advances in back-scatter fiber optic probes have resulted from a grant initiated by this LLS ATD project with Professor Dhadwal at SUNY-Stony Brook. These fiber optic probes will allow LLS to be used in a solution that does not allow a laser beam to pass through it. Their larger viewing volume may prove to be of great importance when LLS is applied to protein crystal growth experiments. While waiting for the LLS modules to arrive, we have invested significant amounts of research time characterizing the revolutionary fiber probes. Tables summarizing some of the hundreds of data sets taken in this study are included in Appendix 3. With the back scatter probes, we have been able to study milky concentrations ranging up to 10% weight concentration without multiple scattering problems. These probes will be useful for both protein crystals and, in another form, for fractal studies. Multiple angle and low angle probes will be developed when funding allows.

9. Evaluation of the miniature modular multiangle LLS unit will begin when it is received. It can hopefully be purchased with 1991 funding. Note that the currently planned level of funding is sufficient only to cover salaries of the two Lewis investigators. Additional funds are required for the purchase of brassboard items. It will be used to study polydispersity, cross-correlation, and dynamic depolarized light scattering (DDLS). This brassboard of optics and simultaneous detectors will allow testing of our multiangle software with simultaneous multiangle scattering on samples that change too quickly to be evaluated with currently available equipment.

10. We have run hundreds of LLS tests on samples of particles (alumina, glass microspheres, pliolyte) proposed for use by the STDCE project. Two particle size distributions are illustrated in Appendix 4. We have presented them with written reports outlining our findings. These studies showed aggregation problems and a background of fine residual (dust) particles which destroy the background contrast for the cameras. After finding the problem, we identified a possible solution. Professor John Ugelstad in Norway supplied us with 70% porosity particles which are nearly density matched when they fill with the solution they are placed into. They are also all spherically symmetric and monodisperse (ie. - identical in size). Professor Ugelstad is willing to make and sell 90% porosity particles. This information has been passed to the STDCE project.

KEY PERSONNEL

William Meyer - NASA/LeRC via CWRU contract Project Manager

- * overall management of the ATD effort
- * reporting to NASA Headquarters
- * submitting progress reports to NASA/LeRC
- * Progress Items 1-4, 9

Dr. Rafat Ansari Research Scientist - NASA/Lewis via CWRU contract * Progress Items 5-10

CONCERNING THE APPENDICES WHICH FOLLOW

The following appendices are provided to show individuals knowledgeable in the field of LLS the kind of data analysis presently available for this project. While enhancements are being added, the software available to and developed by this project is quite advanced. The data presented in the following appendices is presented prior to publication and we ask that it be cited accordingly by the reader.

APPENDIX 1

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SUMMARY OF DR. R.G.W. BROWN'S 26 PAGE REPORT CONCERNING THE SPECIFICATION OF AVALANCHE PHOTODIODE DETECTORS FOR USE IN STATIC AND DYNAMIC LIGHT SCATTERING

PARAMETER Pulse height distribution	VALUE Clean, unimodal	PROOF/DOCUMENTATION Graph/Plot
Dark count Rate	<300 cts/sec PCS <1000 cts/sec LDV	Value @ Operating Temperature
Dead Time	<50nsec (<20ns if possible)	Pulse picture
Count rate stability	<+/- 1/2% over experimental time	Graph/plot over over 30 minutes
Temperature stability	<+/- 1/50 °C	Plot of constant source
Voltage stability (bias)	<+/-2 or 3 mV	Plot of constant source
Quantum efficiency	>25% SLIK (if available) >7.5% C30921S	State method of proof
Linearity of count rate	up to 10 MHz at rate	Graph
Afterpulsing	<0.04%	Calculations +9
Factorial moments	up to 3rd or 4th	Calculations for $n>2$
Power consumption	depends on cooling typically less than 6 Watts	Result of measurement
Operational temperature range	Constant ct rate over T	Graph/Plot
Anti-lock up after saturation		
Output pulses	Compatible with chosen correlator	

SUMMARY OF DR. R.G.W. BROWN'S 20 PAGE REPORT CONCERNING THE SPECIFICATION OF SEMICONDUCTOR LASER DIODES FOR USE IN STATIC AND DYNAMIC LIGHT SCATTERING

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PARAMETER	VALUE	PROOF/DOCUMENTATION
Mode structure	TEM ₀₀	Data sheet/far field photograph
Mode hopping	Zero after warm-up	Graph/plot over 30 minutes
Wavelength	780 or 850 nm typically	Monochromator output
Output power	>20mW over 30 minutes	Graph of plot over 30 minutes
Factorial Moments	Up to 3rd or 4th +/- error	Calculations for n>2
Excess correlations	<0.04%	Results of Auto- measurements at 20K and 100K cts/sec
Coherence length/ linewidth	>3m / <80MHz	Manufactures data sheet or Spectrum- analyzer plot
Laser beam diameter/ collimation	as chosen	Photo, Ronchi test or knife edge convolution
Polarization	500:1 if possible +/- 1/10° alignment	Alignment accuracy fading plot
Power consumption	depends on cooling	Result of measurement
Operational temperature range	Constancy over T	Graph/Plot
Temperature stability	+/-0.5 °C	Graph/Plot
Voltage stabilities	+/-2 or 3mV	Plot of constancy

APPENDIX 2

(GLOSSARY)

This glossary is provided with the assumption that the reader has some knowledge of LLS. Many volumes have been written on the terms listed below. We are providing this glossary to insure that the abbreviations in the following tables are meaningful.

1. Conc.: Weight/volume solids concentration in percent.

2. Diff.Coeff.: Diffusion coefficient. The rate at which a particle diffuses is strongly dependent upon this parameter. In simpler terms, the driving force results from a gradient in concentration which causes particles to move from regions of high concentration to regions of low concentration.

3. 2dCum: Second cumulant of a fourth order cumulant analysis. This analysis describes the size distribution of the particles by its moments. Information on the shape of the distribution function is not provided by this method.

4. PDP: Polydispersity parameter of the system being studied.

5. **DblExp:** Double exponential. A double exponential often fits data quite well, but not much can be inferred unless additional evidence of a bimodal model exists.

6. ExpSam: Exponential sampling is also known as Pike-Ostrowski analysis. This technique does not assume a fixed number of modes.

7. NNLS: A non-negatively constrained least squares algorithm described by C.L. Lawsen and R.J. Hansen in "Solving Least Squares Problems", Prentice-Hall, 1974.

8. Contin: Contin is capable of analyzing a variety of related numerical problems for the regularized solution of linear algebraic and linear Fredholm integral equations of the first kind. It contains options for peak constraints and linear equality and inequality constraints. It was developed by S.W. Provencher in the Federal Republic of Germany.

9. **RMS ERROR**: Root-mean-square error from residual plots as a function of correlator channel.

10. CNTRT: Total number of photons counted in one second.

11. GAMMA: Half-width at half-maximum of the intensity fluctuation spectra due to Brownian motion of particles. The diffusion coefficient is extracted from Gamma.

12. **% ERROR** indicates % error in Gamma, and hence in diffusion coefficient measurement.

13. Radius of gyration is defined as the square root of the weight average of all the mass elements.

14. 2nd Virial Coeff. provides a measure of particles interactions in the system being studied.

TABLE 1

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LASER-LIGHT SCATTERING STUDIES OF MONODISPERSE SYSTEMS (SPHERICALLY SYMMETRICAL IDENTICAL PARTICLES)

			T DMETER	TN NA	JOME TEL			REST	RMS	CNTRT	GAMMA	 0%
degrees	Cm ² Sec ⁻¹	2dCum 1	PDP DD1E	txp Exp	Sam NN	ILS COI	ntin	FIT	ERROR	KHZ	rad/s	ERROR
(DOW 91	nm)		 		- 	 	 	• 1 1 1 1 1 1				
15	4.83E-8	92.9	0.15	N/A	159	82	101	SINN	2.25E-3	485	95.5	1.00
20	4.58E-8	90.5	0.24	88	117	90	75	2dCum	7.23E-4	356	173.3	0.32
30	5.19E-8	91.0	0.07	89	79	85	N/A	NNLS	4.26E-4	288	384.1	0.29
45	5.60E-8	85.1	0.05	N/A	68	83	101	Contin	5.68E-4	164	896	0.33
60	5.73E-8	85.0	0.01	N/A	85	84	101	ExpSam	4.15E-4	218	1530	0.27
06	5.62E-8	87.2	0.002	N/A	88	86	N/A	2dCum	6.22E-4	134	2983	0.42
120	5.45E-8	89.6	0.008	N/A	06	87	N/A	2dCum	5.70E-4	171	4356	0.34
135	5.43E-8	88.7	0.03	N/A	88	90	101	2dCum	З.69Е-4	198	5009	0.22
150	5.32E-8	87.1	0.09	N/A	144	83	N/A	SINN	1.92E-4	262	5571	0.85
160	5.26E-8	89.4	0,07	87	90	84	92	ExpSam	2.25E-4	344	5645	0.17
(DOW 48	2 nm)											
30	1.06E-8	466.4	0.005	N/A	472	461	N/A	2dCum	1.31E-4	2155	466.4	0.19
60	7.34E-9	627.7	0.11	690	404	571	N/A	SJINN	5.94E-4	88.7	627.8	0.86
06	9.68E-9	486.3	0.07	484	769	499	486	2dCum	5.01E-4	57.5	535	0.31
135	3.10E-9	1087	0.45	952	1583	786	1938	NNLS	7.45E-4	26.9	408	1.34
(DOW 1.	082 micron)											
30	3.77E-9	1364	0.005	N/A	1122	1277	N/A	2dCum	9.07E-4	649	25.5	0.84
45	З.69Е-9	1244	0.14	1188	1132	1026	1328	Contin	5.03E-4	136	61.3	0.53
60	3.22E-9	1387	0.20	1101	762	1097	1620	DblExp	8.23E-4	48.6	93.8	1.17
06	3.20E-9	1194	0.28	1241	1269	872	1008	SLIN	7.88E-4	30	218	0.71

Parameters: Temperature: 25°C; Laser Wavelength: 514.5 nm; Particles: Polystyrene polymer.

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TABLE 2

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LASER-LIGHT SCATTERING STUDIES OF POLYDISPERSE SYSTEMS (MIXTURES OF DIFFERENT SIZE PARTICLES)

Angle	Diff.Coef1	f. D	IAMETE	N NI Y	ANOMETI	ERS		BEST	RMS	CNTRT	GAMMA	%
degrees	Cm ² Sec ⁻¹	2dCum 1	Idd 909	Exp E	¢pSam №	SULS	Contin	FIT	ERROR	KHz	rad/s	ERROR
BIMODAL	(CWRU 39 I	nm and D	1 16 MO	(mī								
30	7.62E-8	56.1	0.20	55	50	49	71	Contin	2.03E-4	335	622	0.40
60	7.86E-8	55.6	0.17	53	51	52	70	ExpSam	3.92E-4	186	2341	0.39
06	8.07E-8	53.7	0.17	55	99	47	54	Contin	4.64E-4	157	4849	0.24
135	7.98E-8	54.4	0.18	54	53	50	54	Contin	4.39E-4	212	8161	0.32
160	7.89E-8	54.2	0.20	54	85	53	78	DblExp	2.68E-4	408	9314	0.17
TRIMODAL	L (CWRU 39	nm, DOW	91 nm	and 4	82 nm)			,				
20	_ 1.06E-8	394	0.36	313	215	290	523	DblExp	3.52E-4	2665	39.9	1.32
30	1.00E-8	396	0.36	325	461	314	499	NNLS	1.02E-4	1563	87.9	1.29
30	1.01E-8	391	0.29	328	215	245	518	Contin	5.42E-4	1604	89.0	1.40
60	4.36E-8	82.9	0.27	78	78	69	171	SINN	3.33E-4	138	1568	1.14
06	4.00E-8	88.8	0.28	84	LL	70	167	Contin	3.66E-4	114	2929	1.08
135	5.72E-8	70.0	0.22	70	11	59	136	SINN	6.58E-4	130	6345	0.55
150	3.16E-8	94.0	0.35	85	85	52	62	NNLS	4.22E-4	216	5165	1.49
MULTIMO	DAL (CWRU	39 nm, D	1 16 MO	DD , mc	W 482	nm and	I DOW 1	(mn 080				
30	0.70E-8	551	0.36	439	557	431	882	NNLS	2.31E-4	1492	63.2	1.52
60	1.90E-8	141	0.48	116	72	59	172	SJUN	3.10E-4	129	925	2.10
06	2.42E-8	125	0.28	96	91	46	261	SINN	6.37E-4	06	2083	3.15
135	2.55E-8	106	0.27	68	81	62	134	STIN	4.96E-4	102	4187	2.85
150	1.23E-8	173	0.40	111	148	61	410	Contin	3.32E-4	170	2802	4.59

Parameters: Temperature: 25°C; Laser Wavelength: 514.5 nm; Particles: Polystyrene polymer.

APPENDIX 3

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FIBER OPTIC LIGHT SCATTERING RESULTS

TABLE 3: 35 NM (Lot # 1900223A)

PARAMETERS:

Particle Standard: Bangs Polystyrene 35 nm. Wavelength=632.8 nm, T=20°C, Angle=157° Run time: 5 minutes

Conc.	Diff.Coef	f.	DIAME	TER IN I	NANOMETEI	RS		BEST	RMS	CNTRT	GAMMA	
ф	Cm ² Sec ⁻¹	2dCum	PDP	DblExp	ExpSam	SINN	Contin	FIT	ERROR	KHZ	rad/s	ERROR
0.01	9.99E-8	42.1	0.03	N/A	42	42	N/A	2dCum	9.91E-4	16	6858	0.72
0.05	10.6E-8	39.2	0.06	39	49	38	40	2dCum	2.84E-4	72	7362	0.17
0.10	10.7E-8	39.3	0.04	N/A	39	37	41	2dCum	3.51E-4	53	7342	0.22
0.50	11.1E-8	37.8	0.04	N/A	37	35	40	ExpSam	1.40E-4	343	7633	0.05
1.00	11.4E-8	36.6	0.05	37	37	37	38	SINN	6.73E-5	300	7888	0.03
2.50	11.8E-8	35.4	0.05	35	35	35	36	SINN	8.18E-5	665	8146	0.07
5.00	12.2E-8	33.8	0.07	33	34	19	N/A	DblExp	6.81E-5	615	8544	0.26
10.0	12.0E-8	33.6	0.09	34	33	33	35	ExpSam	2.62E-4	177	8575	0.14

REMARKS: All measurements were taken using a Neutral Density Filter of OD.40

FIBER OPTIC LIGHT SCATTERING RESULTS

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TABLE 4: 85 NM (Lot # 1900214A)

PARAMETERS:

Particle Standard: Bangs Polystyrene 85 nm. Wavelength=632.8 nm, T=20°C, Angle=157° Run time: 5 minutes

		רבמ										-
Conc.	Diff.Coef1	1.1	DIAMET	ER IN	NANOMET	ERS		BEST	RMS	CNTRT	GAMMA	æ
940 	Cm ² Sec ⁻¹	2dCum	PDP Dł	olExp	ExpSam	NNLS	Contin	FIT	ERROR	KHz	rad/s	ERROR
0.01	4.77E-8	87.3	0.052	N/A	81	86	101	2dCum	3.73E-4	38.01	3304.5	0.23
0.05	4.87E-8	86.4	0.035	N/A	86	82	N/A	2dCum	6.66E-4	17.66	3338.9	0.44
0.05	4.95E-8	86.2	0.008	N/A	87	84	N/A	2dCum	3.58E-4	132.3	3343.8	0.25
0.10	5.12E-8	82.3	0.031	N/A	82	81	N/A	2dCum	2.84E-4	155.4	3502.9	0.22
0.10	5.10E-8	84.4	0.005	N/A	N/A	81	N/A	2dCum	1.49E-4	253.0	3415.9	0.10
0.20	5.16E-8	83.5	0.005	N/A	84	80	N/A	ExpSam	2.87E-4	436.3	3451.8	0.25
0.25	5.30E-8	80.0	0.020	N/A	80	78	N/A	ExpSam	4.04E-4	624.2	3604.0	0.38
0.30	5.20E-8	81.3	0.025	N/A	81	81	N/A	ExpSam	2.24E-4	537.2	3547.0	0.09
0.40	5.31E-8	79.9	0.018	N/A	80	78	N/A	ExpSam	5.76E-4	694.8	3609.4	0.23
0.50	5.39E-8	77.1	0.054	N/A	71	74	76	2dCum	2.64E-4	974.8	3741.3	0.26
0.60	5.43E-8	76.9	0.048	N/A	71	75	N/A	2dCum	1.96E-4	792.4	3751.9	0.18
0.70	5.38E-8	77.0	0.059	N/A	<i>LL</i>	75	N/A	2dCum	1.80E-4	912.0	3745.6	0.18
0.80	5.51E-8	75.0	0.062	N/A	106	72	N/A	2dCum	2.71E-4	949.2	3845.0	0.27
06.0	5.48E-8	75.4	0.061	76	95	74	76	Contin	4.45E-4	990.7	3822.5	0.18
1.00	5.71E-8	73.2	0.045	N/A	71	73	76	ExpSam	5.39E-4	1360	3939.9	0.34
5.00	6.25E-8	64.6	0.099	61	41	51	73	SJUN	9.72E-5	983.2	4464.0	0.51

All measurements were taken using a Neutral Density Filter of OD.40 REMARKS:

FIBER OPTIC LIGHT SCATTERING RESULTS

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TABLE 5: 165 NM (Lot # 1890620C)

PARAMETERS:

Particle Standard: Bangs Polystyrene 165 nm. Wavelength=632.8 nm, T=20°C, Angle=157° Run time: 5 minutes

Run tı	me: 5 minut	es										
Conc.	Diff.Coeff	•	DIAMET	ER IN	NANOMET	ERS		BEST	RMS (CNTRT	GAMMA	ф
ж	Cm ² Sec ⁻¹	2dCum	I dog	blExp	ExpSam	SINN	Contin	FIT	ERROR	KHZ	rad/s	ERROR
0.005	2.70E-8	161.3	0.005	N/A	163	151	N/A	2dCum	3.95E-4	128	1788	0.12
0.01	2.67E-8	161.3	0.005	N/A	163	153	N/A	2dCum	3.45E-4	200	1788	0.13
0.05	2.77E-8	157.7	0.005	N/A	134	144	N/A	2dCum	8.22E-4	578	1828	0.22
0.10	2.75E-8	151.8	0.048	N/A	140	145	179	2dCum	5.20E-4	811	1900	0.26
0.20	2.78E-8	147.9	0.070	144	148	N/A	106	ExpSam	1.81E-4	1098	1950	0.11
0.30	2.84E-8	147.3	0.043	N/A	143	145	N/A	ExpSam	5.86E-4	984	1958	0.31
0.40	2.82E-8	145.4	0.075	144	140	138	130	ExpSam	2.02E-4	984	1984	0.13
0.50	2.84E-8	143.8	0.083	144	199	142	N/A	2dCum	4.97E-4	883	2006	0.28
0.60	2.85E-8	143.0	0.083	143	141	95	N/A	SINN	3.05E-4	749	2017	0.30
0.70	2.88E-8	140.7	0.095	153	129	106	N/A	ExpSam	7.65E-4	666	2049	0.58
0.80	2.98E-8	133.6	0.11	133	115	89	150	SINN	2.01E-4	651	2158	0.49
0.90	3.15E-8	121.7	0.16	131	113	49	136	Contin	4.48E-4	516	2370	0.82
1.00	3.21E-8	117.3	0.18	102	85	98	158	ExpSam	1.05E-3	536	2458	1.00
2.50	10.6E-8	32.80	0.15	36	35	21	32	SINN	1.63E-4	300	8779	2.54
2.50	14.1E-8	22.30	0.30	19	17	12	44	NNLS	1.98E-4	339	12958	2.10

REMARKS: All measurements were taken using a Neutral Density Filter of OD.40

FIBER OPTIC LIGHT SCATTERING RESULTS

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TABLE 6: 261 NM (Lot # 1B73 SERAGEN)

PARAMETERS:

Particle Standard: Seragen Polystyrene 261 nm. Wavelength=632.8 nm, T=20°C, Angle=157° Run time: 5 minutes

Run ti	me: 5 minut	tes										-
Conc.	Diff.Coef1	ŕ.	DIAMET	ER IN	NANOMETI	ERS		BEST	RMS C	NTRT	GAMMA	ж
с¥о	Cm ² Sec ⁻¹	2dCum	PDP	DblExp	ExpSam	NNLS	Contin	FIT	ERROR	KHz	rad/s	ERROR
0.005	1.63E-8	260.1	0.017	N/A	260	258	267*	2dCum	6.76E-4	N/A	1108	0.23
0.01	1.65E-8	260.8	0.005	N/A	264	257	259*	2dCum	2.54E-4	N/A	1105	0.19
0.05	1.68E-8	255.2	0.005	N/A	258	251	256*	2dCum	1.68E-4	N/A	1130	0.13
0.10	1.67E-8	257.4	0.005	N/A	260	250	250*	2dCum	3.35E-4	N/A	1120	0.28
0.20	1.78E-8	233.0	0.06	224	234	217	245	Contin	3.63E-4	195	1238	0.23
0.30	1.71E-8	230.0	0.13	239	236	201	240	STIN	2.48E-4	229	1255	0.30
0.40	1.72E-8	223.0	0.16	225	200	197	206	SINN	3.04E-4	295	1293	0.39
0.50	1.83E-8	203.4	0.182	210	188	156	341	SINN	4.94E-4	228	1418	0.59
1.00	1.03E-8	226.0	0.179	205	195	186	303	SINN	3.60E-3	309	1457	0.78

* represents Contin calculations without Mie corrections in the limit 200,400 nm. **REMARKS**: