Asteroids, Comets, Meleors 1991, pp. 293-290 Lunar and Planetary Institute, Houston, 1992

N932-19183

570-90

293

MODIFICATION OF PRIMORDIAL ICES BY COSMIC RAYS AS SIMULATED BY CYCLOTRON IRRADIATION

## R.I. Kaiser and K. Roessler

Institut für Nuklearchemie, Forschungszentrum Jülich, Postfach 1913, D-5170 Jülich, Germany

### ABSTRACT

Frozen CH<sub>4</sub> and CH<sub>4</sub>/Ar mixtures closed into metal cuvettes and open to the vacuum were irradiated at 15 and 77 K with 10 to 20 MeV p and <sup>3</sup>He<sup>2+</sup> ions in order to simulate the effect of cosmic rays on solid organic matter in space. Ices exposed to vacuum represent surfaces of icy systems whereas closed systems stand for bulk ices. The products were analysed by MS, SEM, RBS, ERDA, <sup>1</sup>H-NMR, HPLC, GC-MS, NEXAFS and FT-IR. Volatile products consisted of a mixture of low molecular species, e.g.  $C_2H_2$ ,  $C_2H_4$ ,  $C_2H_6$ , and long linear aliphatic and olefinic compounds. The formation of polycyclic aromatic hydrocarbons (PAHs) and related species in solid CH<sub>4</sub> is due to a multi center reaction within one collision cascade and is governed by energy density effects with critical linear energy transfer values  $L_T$  between 2 and 10 keV  $\mu$ m<sup>-1</sup>. Open ices exhibit preferential hydrogen release resulting in an increased carbonisation as compared to more hydrogen rich molecules protected inside large icy bodies.

### INTRODUCTION

High energetic particles may modify primordial matter in space to complex compounds and even to precursor molecules for biological evolution. Solid  $CH_4$  has frequently been chosen as a deliberately simple model substance to study the modification of hydrocarbons (Strazzulla and Johnson 1991; Roessler 1991). However, a detailed comparison of closed and open  $CH_4$  targets and systematic studies of dose and energy density effects has not yet been executed. Thin ice condensates in the vacuum may represent surfaces of ices whereas  $CH_4$  targets, closed into metall cuvettes, stand for the bulk.

## EXPERIMENTAL

In the actual experiments, 11.1  $\mu$ m layers of CH<sub>4</sub> and 9.7  $\mu$ m of CH<sub>4</sub>/Ar (1:12) (CH<sub>4</sub>: 99.9995 %, Ar: 99.999 %) were condensed at 10 - 15 K on an (111) Si wafer attached to an aluminium cold finger of a bath cryostat. A second set of experiments was performed with closed targets consisting of 1 mm frozen CH<sub>4</sub> layers between a stainless steel backing and a Ti foil at 77 K. Both irradiation arrangements are described in detail in (Patnaik et al. 1990; Kaiser 1991). The ices were irradiated at 2\*10<sup>-7</sup> mbar with 250 nA cm<sup>-2</sup> beams (fluence some 10<sup>15</sup> cm<sup>-2</sup>) of 17.6 MeV protons and 16.9 MeV <sup>3</sup>He<sup>2+</sup> ions of the CV 28 compact cyclotron of Forschungszentrum Jülich. Irradiation times varied from 5 to 90 min giving rise to doses between 0.07 and 16.8 eV per carbon atom. The linear energy transfer with respect to CH<sub>4</sub>, L<sub>T</sub>(CH<sub>4</sub>), was increased from 161 (p//Ar/CH<sub>4</sub>) via 1800 (p//CH<sub>4</sub>) to 10810 eV  $\mu$ m<sup>-1</sup> (<sup>3</sup>He<sup>2+</sup>/CH<sub>4</sub>).

#### RESULTS

The volatile species released from the open samples during irradiation and in the warm-up phase to 293 K were detected by a quadrupole mass spectrometer (MS). Immediately after inset of irradiation, species containing one carbon atom such as  $CH_2$  and  $CH_4$  appear and dominate all mass spectra. CH and  $CH_3$  show lesser intensity. In all experiments,  $C_2H_4$  and  $C_2H_2$  were liberated some minutes after onset of beam. The yields of both products increased with time but were observed in lower yields than  $CH_2$ .  $C_2H_4$  emerged later at the expense of the unsaturated species. Higher molecular-weight hydrocarbons were detected only in later stages of irradiation and in the warm-up phase in yields of approx. 1 - 2 X each. In closed cuvettes, up to 94 X of  $CH_4$  were converted into volatiles, i.e.  $C_2H_2$ ,  $C_2H_4$ ,  $C_2H_6$  and  $C_3H_6$  (each up to 20 X) and lower yields of molecules containing up to 9 C atoms (~ 0.5 X each). Furtheron, He-irradiation of closed targets induced a formation of cycloalkanes and cycloalkenes such as cyclopropane, cyclopentane, cyclohexane and cyclohexene besides small amounts of benzene (Patnaik et al. 1990) whereas  $CH_4$  ices

exposed to vacuum yielded mostly unsaturated species, i.e. mono- and disubstituted benzenes, 1,2-dihydronaphthaline, 1,2,3,4-tetrahydronaphthaline and anthracene and/or phenanthrene (Kaiser et al. 1992a). The solid residues were analysed by visual inspection, optical microscopy (OM), Fourier transform infrared spectroscopy (FT-IR) in transmission at 293 K, scanning electron microscopy (SEM), Rutherford backscattering and elastic recoil detection analysis (RBS-ERDA), and near edge X-ray absorption fine structure spectroscopy (NEXAFS). In the targets exposed to vacuum CH, was finally converted into inhomogeneous solid residues, stable at room temperature and with a thickness of approx. 0.5 µm (Kaiser et al. 1992b). Aliphatic features dominate the IR spectra, whereas no unsaturated modes could be detected. With a method of higher sensitivity, H-C- bonds were analysed by NEXAFS, manifestating their concentration between 1 µmole and 1 nmole. FT-IR half widths increasing with irradiation time (factor 3 to 4) can be interpreted as an increase in oligomerization with dose. The seggregation of different phases observed by SEM and OM substantiate these results as shown in Fig. 1a-b. It coincides with an overall H:C ratio decrease from originally 4 (CH4) to 1.4, whereas samples in closed metal cuvettes exhibit a H:C ratio of 2. This indicates an preferential hydrogen release from open targets. The soluble fraction of the residues was dissolved in CDCl, (99.95 %) and 'H-NMR, HPLC and GC-MS spectra ware recorded. Concerning the open targets, up to approx. 90 % of the soluble phase consist of linear alkanes and 5 to 6 % of alkadienes, both containing up to 28 C atoms. <sup>1</sup>H-NMR spectra show 4-5 % of hydrogen located at olefenic carbon atoms. Eventually, about 0.1 % aromatic species were detected. Substituted benzenes were synthesized in all irradiated samples, but higher annulated rings and complex polycyclic aromatic hydrocarbons (approx. 0.1 %) such as naphthaline and pyrene were limited to "He<sup>2+</sup>-irradiations with highest  $L_{T}(CH_{4})$ . They could not be detected in proton irradiated open and closed samples, although proton irradiation doses (1.4 - 2.8 eV per C atom) were similar to those of "He2+-irradiations (0.1 - 15 eV per C atom), Fig.2.

1.00000

III MARANA MANANA I AMARANA

. .

# DISCUSSION

. . . . . . . . .

The dualism of radiation dose and energy density and the definite differences in open and closed ice targets constitute ideal tools to discuss the experimental results. The correlation of the C2 species in the mass spectra shows unequivocally that  $C_2H_2$ ,  $C_2H_4$ , and  $C_2H_6$  descend from the same precursor, i.e. excited methylcarbene [CH-CH<sub>3</sub>]\* formed by insertion of hot carbon into a C-H bond of CH<sub>4</sub> (Stöcklin 1969). Hydrogen elimination of methylcarbane leads to ethine. H-rearrangement yields ethene. Conspicously, C2Ha is only formed at higher doses with a critical H concentration necessary to stimulate hydrogen pick-up of methylcarbene followed by deexcitation. The higher molecular species like propane were synthesized by additional radical attack and combination. The formation of a few % polycyclic aromatic hydrocarbons and partially hydrogenated analogs is governed by a critical linear energy transfer  $L_{T}$  (CH<sub>4</sub>) of about 2 to 10 keV µm<sup>-1</sup> which induces a multi center reaction (Roessler et al. 1990), i.e. aggregation of individual insertion products of hot carbon secondaries in CH, and radicals such as CH, CH<sub>2</sub> and CH<sub>3</sub> in a single collision cascade. If  $L_{\tau}(CH_{\star})$  is below this critical value, the concentration of aggregating species is too low in the activated zone to lead to the formation of complex two- or three-dimensional molecules in the relaxation phase. Thus, the accumulated radiation dose seems to be of minor importance in the synthesis of PAHs as compared to the  $L_{T}(CH_{4})$  value. Classical hot atom reactions such as insertion, abstraction and addition (Stöcklin 1969, Roessler 1991) cannot explain the synthesis of PAHs at lowest doses of 0.1 eV per C atom. Step-by-step production in overlapping cascades should contribute to a radiolytical formation only at doses exceeding some 100 eV per C atom . Thus, He-ions and heavier components of cosmic energetic particles are very effective in processing organic material. The synthesized species point to different chemical conversion of the original CH, in the bulk and near surfaces. Ices exposed to vacuum exhibit H2 formation in aligned cascades and elimination from excited intermediates which can be considered as a kind of oxidation. B-outdiffusion as well as the loss of reactive intermediates is restricted in closed targets. Consequently, molecules are protected from being oxidized: saturated cyclic compounds were solely formed in closed samples. The hydrogen loss of icy surfaces results in an increasing carbonization, such as shown by the low H:C ratio of 1.4 and the formation of amorphous carbon and amorphous hydrogenated carbon in specific areas.

The linear energy transfer is the limiting factor in formation of complex molecules in space (Fig. 3). At the relatively long irradiation times at low fluxes in space, the annealing of defects and reactive species may decrease the necessary concentration for build-up of larger units.

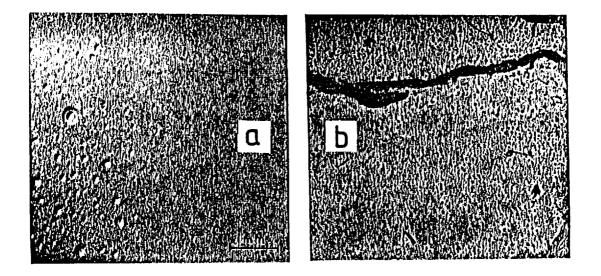
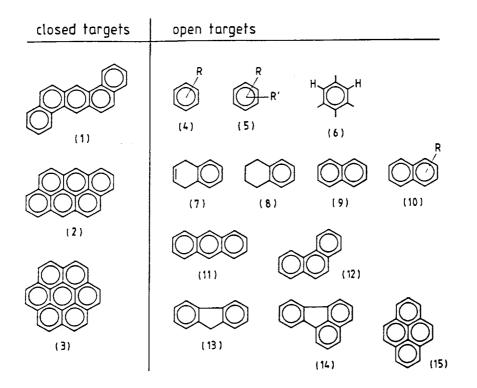


Fig.la-b : Microscopy of residues on Si wafer. a)  $p/(Ar/CH_4)$ ,  $D^* = 0.07 \text{ eV/C}$ ,  $L_T(CH_4) = 161 \text{ eV } \mu \text{m}^{-1}$ ; b)  $^3\text{He}^{2*}//CH_4$ ,  $D^* = 16.87 \text{ eV/C}$ ,  $L_T(CH_4) = 10810 \text{ eV } \mu \text{m}^{-1}$ .

Fig.2 : Polycyclic aromatic hydrocarbons formed in open targets and in metall cuvettes (1:dibenzo[a,b]anthracene, 2:dibenzo[def,mno]chrysene, 3:coronene, 4-6:mono-, di and tetrasubstituted benzenes, 7:1,2-dihydronaphthaline, 8:1,2,3,4-tetrahydronaphthaline, 9:naphthaline, 10:alkylnaphthaline, 11:anthracene, 12:phenanthrene, 13:fluorene, 14:fluoranthene, 15:pyrene).



÷-

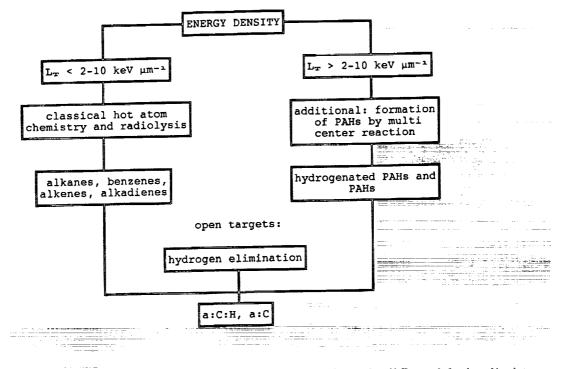


Fig.3 : Mechanisms of modification of pure CH<sub>4</sub> and synthesized species under MeV particle irradiation.

## ACKNOWLEDGEMENTS

The authors thank Prof. G. Götz and Mr. P. Müller, Jena for recording and helpful discussions of RBS-ERDA spectra. They are indebted to the CV 28 cyclotron crew of Forschungszentrum Jülich for irradiations, Prof. J. Lauterwein and Mrs. K. Busse, Münster for recording the NMR spectra, Dr. E. Wallura, Jülich for SEM measurements, and Prof. G. Grunze, Heidelberg for NEXAFS spectra.

## REFERENCES

Kaiser R.I. (1991) <u>Chemische Prozesse durch Zyklotronionen in festem Methan - Simulation der</u> Wechselwirkung kosmischer Strahlung mit einfachen extraterrestrischen Eiskörpern. Report Jül-2492, 130 pp. Kaiser R.I., Mahfouz R.M., and Roessler K.(1992a) Emission of organic products from the surface of frozen methane under MeV ion irradiation. <u>Nucl. Instr. Meth. Phys. Res. B</u>, in press.

Kaiser R.I., Lauterwein J., Müller P., and Roessler K. (1992b) Energy density effects in the formation of organic residues in frozen methane by MeV ions. Nucl. Instr. Meth. Phys. Res. B, in press.

Patnaik A., Roessler K., and Zádor E. (1990) Reactions of suprathermal carbon atoms in solid methane. Radiochim. Acta, <u>50</u>, 75-85.

Ē

INTERNET

Y HANNISH ISH SULULI I

A LA MARKED LA

Roessler K., Eich G., Patnaik A, and Zádor E. (1990) Polycyclic aromatic hydrocarbons via multicenter reactions induced by solar radiation. Lunar Planet. Sci. Conf. XXI, 1035-1036.

Roessler K. (1991) Suprathermal chemistry in space. In <u>Solid State Astrophysics</u> (E. Bussoletti and G. Strazzulla, eds.), pp. 197-266, North Holland, Amsterdam.

Stöcklin G. (1969) <u>Chemie heißer Atome</u>. Verlag Chemie, Weinheim. 283 pp. (1972) Chimie des atomes chauds, Masson et Cie, Paris.

Strazzulla G. and Johnson R.E. (1991). In <u>Comets in the Post-Halley Era</u> (R.L. Newburn Jr., et al., eds), 1, pp. 243-275 Kluwer, Dordrecht.

1-5