

ified matrix that meets the above criteria is

$$A'_p = \begin{bmatrix} u & \rho & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & u & 0 & 0 & 0 & \frac{B_y}{\rho} & \frac{B_z}{\rho} & \frac{1}{\rho} \\ 0 & 0 & u & 0 & 0 & -\frac{B_x}{\rho} & 0 & 0 \\ 0 & 0 & 0 & u & 0 & 0 & -\frac{B_x}{\rho} & 0 \\ 0 & 0 & 0 & 0 & u & 0 & 0 & 0 \\ 0 & B_y & -B_x & 0 & 0 & u & 0 & 0 \\ 0 & B_z & 0 & -B_x & 0 & 0 & u & 0 \\ 0 & \gamma p & 0 & 0 & 0 & 0 & 0 & u \end{bmatrix}. \quad (9)$$

The eigensystem of this matrix is composed of the following eight waves, with their corresponding eigenvalues λ , left eigenvectors $\tilde{\ell}$ and right eigenvectors \tilde{r} :

One Entropy Wave

$$\begin{aligned} \lambda_e &= u \\ \tilde{\ell}_e &= \left(1, 0, 0, 0, 0, 0, 0, -\frac{1}{a^2}\right) \\ \tilde{r}_e &= (1, 0, 0, 0, 0, 0, 0, 0)^T; \end{aligned} \quad (10)$$

Two Alfvén Waves

$$\begin{aligned} \lambda_a &= u \pm \frac{B_x}{\sqrt{\rho}} \\ \tilde{\ell}_a &= \left(0, 0, -B_z, B_y, 0, \pm \frac{B_z}{\sqrt{\rho}}, \mp \frac{B_y}{\sqrt{\rho}}, 0\right) \\ \tilde{r}_a &= (0, 0, -B_z, B_y, 0, \pm \sqrt{\rho} B_z, \mp \sqrt{\rho} B_y, 0)^T; \end{aligned} \quad (11)$$

Four Magneto-acoustic Waves

$$\begin{aligned} \lambda_{f,s} &= u \pm c_{f,s} \\ \tilde{\ell}_{f,s} &= \left(0, \pm \rho c_{f,s}, \mp \frac{B_x B_y \rho c_{f,s}}{\rho c_{f,s}^2 - B_x^2}, \mp \frac{B_x B_z \rho c_{f,s}}{\rho c_{f,s}^2 - B_x^2}, 0, \frac{B_y \rho c_{f,s}^2}{\rho c_{f,s}^2 - B_x^2}, \frac{B_z \rho c_{f,s}^2}{\rho c_{f,s}^2 - B_x^2}, 1\right) \\ \tilde{r}_{f,s} &= \left(\rho, \pm c_{f,s}, \mp \frac{B_x B_y c_{f,s}}{\rho c_{f,s}^2 - B_x^2}, \mp \frac{B_x B_z c_{f,s}}{\rho c_{f,s}^2 - B_x^2}, 0, \frac{B_y \rho c_{f,s}^2}{\rho c_{f,s}^2 - B_x^2}, \frac{B_z \rho c_{f,s}^2}{\rho c_{f,s}^2 - B_x^2}, \gamma p\right)^T; \end{aligned} \quad (12)$$

One ‘‘Divergence’’ Wave

$$\begin{aligned}
\lambda_d &= u \\
\tilde{\ell}_d &= (0, 0, 0, 0, 1, 0, 0, 0) \\
\tilde{r}_d &= (0, 0, 0, 0, 1, 0, 0, 0)^\top.
\end{aligned} \tag{13}$$

It is important to note that the first seven waves yield the same eigenvectors and eigenvalues as the seven-wave Riemann problem, with the additional information that none of them carries a change in B_x (the fifth entry of each right eigenvector is zero), and none of the wave strengths is proportional to a jump in B_x (the fifth entry of each left eigenvector is zero). The new eighth wave travels with the x -component of the flow speed (its eigenvalue is u), and it carries a jump in B_x (the only non-zero entry in the left eigenvector is the entry corresponding to B_x), and affects only the x -component of the magnetic field (the only non-zero entry in the right eigenvector is the entry corresponding to B_x).

It is clear that the eigensystem of this modified matrix has all of the desired properties. In the case $B_x = \text{constant}$, the strength of the eighth wave is zero, and the model reverts to that of the seven-wave problem. The new wave simply gives a rational procedure for dealing with non-zero jumps in B_x across the cell faces, which will in general occur when problems in two or three dimensions are being solved. The question remains, however, of what the modification of the matrix A_p (and the corresponding changes to B_p and C_p) has done to the system of conservation laws.

This can be seen by collecting the source terms due to the modifications to A_p , B_p and C_p and transforming to conserved variables. The new equation set, which has the eight-wave eigensystem described above, is

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho \mathbf{u} \\ \mathbf{B} \\ E \end{pmatrix} + \nabla \cdot \begin{pmatrix} \rho \mathbf{u} \\ \rho \mathbf{u} \mathbf{u} + \mathbf{I} \left(p + \frac{\mathbf{B} \cdot \mathbf{B}}{2} \right) - \mathbf{B} n \mathbf{B} \\ \mathbf{u} \mathbf{B} - \mathbf{B} \mathbf{u} \\ \left(E + p + \frac{\mathbf{B} \cdot \mathbf{B}}{2} \right) \mathbf{u} - \mathbf{B} (\mathbf{u} \cdot \mathbf{B}) \end{pmatrix} = - \begin{pmatrix} 0 \\ \mathbf{B} \\ \mathbf{u} \\ \mathbf{u} \cdot \mathbf{B} \end{pmatrix} \nabla \cdot \mathbf{B}. \tag{14}$$

This is a noteworthy result: *the source term that must be added to Equation 2 is proportional to $\nabla \cdot \mathbf{B}$* . At the partial differential equation level, only terms that are equal to zero have been added to the conservative form of the governing equations. So, while technically the equations are no longer in conservative form, the deviations from conservation will be very small. It is only by writing the equations in this slightly non-conservative form that the singularity related to $\nabla \cdot \mathbf{B}$ can be removed. It has been previously noted that solving the momentum equation in non-conservative form can remove instabilities related to non-zero

$\nabla \cdot \mathbf{B}$ [1]; the current work hopefully reinforces this earlier result, and sheds further light on the mechanism for stabilizing the equations, as well as applying the idea in a novel way to develop a Riemann solver for multi-dimensional MHD.

It is interesting to note another justification of this particular choice of source term. Rewriting Equation 2 slightly by expanding some of the terms, the following form of the equations

$$\begin{aligned} \frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) &= 0 \\ \frac{\partial}{\partial t} (\rho \mathbf{u}) + \nabla \cdot (\rho \mathbf{u} \mathbf{u}) + \nabla \cdot \left(p + \frac{\mathbf{B} \cdot \mathbf{B}}{2} \right) - \mathbf{B} \cdot \nabla \mathbf{B} - \underline{\mathbf{B} \nabla \cdot \mathbf{B}} &= 0 \\ \frac{\partial \mathbf{B}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{B} + \mathbf{B} \nabla \cdot \mathbf{u} - \mathbf{B} \cdot \nabla \mathbf{u} - \underline{\mathbf{u} \nabla \cdot \mathbf{B}} &= 0 \\ \frac{\partial E}{\partial t} + \mathbf{u} \cdot \nabla \left(E + p + \frac{\mathbf{B} \cdot \mathbf{B}}{2} \right) + \left(E + p + \frac{\mathbf{B} \cdot \mathbf{B}}{2} \right) \nabla \cdot \mathbf{u} - & \\ \mathbf{B} \cdot \nabla (\mathbf{u} \cdot \mathbf{B}) - \underline{(\mathbf{u} \cdot \mathbf{B}) \nabla \cdot \mathbf{B}} &= 0 \end{aligned} \quad (15)$$

is obtained. The terms that are proportional to $\nabla \cdot \mathbf{B}$ have been underlined; they are exactly the same as the source term defined above. Thus it can be seen that the addition of the source term in Equation 14 simply acts to remove the terms proportional to $\nabla \cdot \mathbf{B}$ that appear in Equation 2.

Another interesting note is what the evolution equation for $\nabla \cdot \mathbf{B}$ is for the two forms of the governing equations. This may be seen by taking the divergence of the evolution equation for the magnetic field in Equations 2 and 14. For the original form of the equations, the evolution equation is

$$\begin{aligned} \nabla \cdot \left(\frac{\partial \mathbf{B}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{B} + \mathbf{B} \nabla \cdot \mathbf{u} - \mathbf{B} \cdot \nabla \mathbf{u} - \mathbf{u} \nabla \cdot \mathbf{B} \right) &= 0 \\ \frac{\partial}{\partial t} (\nabla \cdot \mathbf{B}) &= 0. \end{aligned} \quad (16)$$

From the partial differential equation point of view, this might well seem the correct result; $\nabla \cdot \mathbf{B} = 0$ is an initial condition, and this equation guarantees that $\nabla \cdot \mathbf{B} = 0$ is maintained throughout the evolution. For the modified form of the equations, the evolution equation for the magnetic field is

$$\begin{aligned} \nabla \cdot \left(\frac{\partial \mathbf{B}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{B} + \mathbf{B} \nabla \cdot \mathbf{u} - \mathbf{B} \cdot \nabla \mathbf{u} \right) &= 0 \\ \frac{\partial}{\partial t} (\nabla \cdot \mathbf{B}) + \nabla \cdot (\mathbf{u} \nabla \cdot \mathbf{B}) &= 0. \end{aligned} \quad (17)$$

Thus the addition of the source term has modified the evolution equation for $\nabla \cdot \mathbf{B}$ so that the quantity $\nabla \cdot \mathbf{B}/\rho$ is treated as a passive scalar. This is clearly the more numerically stable of the two evolution equations; any local $\nabla \cdot \mathbf{B}$ that is created is convected away.

The above derivation gives all the pieces for building an ideal MHD solver that works for two-dimensional problems, without having to resort to non-located variables or a projection algorithm. Specifically, a Roe-type approximate Riemann solver has been implemented, where the wave strengths and speeds are derived from the above left eigenvectors and eigenvalues. The eigenvectors are properly normalized to avoid difficulties associated with coinciding wave speeds [5]. The average state needed at cell interfaces is computed by a simple average of left and right states (although a Roe average does exist for the ideal MHD equations [4]). The source term, though small, is calculated in each cell, and added to the residual. The resulting code is first order in space and time.

3 A Test of the Eight-Wave Riemann Solver

Brio and Wu [2] developed a test problem for one-dimensional MHD solvers based on the shock-tube problem of Sod [6]. Two stationary plasmas are separated by a membrane which is removed at $t = 0$, allowing the plasmas to interact. The test problem used here for the two-dimensional MHD solver is a rotated version of the Brio-Wu problem. The left and right input states, and the orientation of propagation of disturbances to the grid, is shown in Figure 2. In the Brio-Wu problem (the top figure), the boundary conditions are that the problem is periodic along a line $y = \text{constant}$; in the current test problem (the bottom figure), the boundary conditions are that the problem is periodic along a line $x + y = \text{constant}$.

Both the rotated and non-rotated problems were run on coarse (600 cells in x) and fine (1200 cells in x) grids. The time step was taken as $\Delta t/\Delta x = 0.2$, which corresponds to a CFL number of approximately 0.8 on the non-rotated problem. The ratio of specific heats, γ , was 2.0. The number of time steps taken on the coarse and fine grids were 100 and 200, respectively. The x -axis in the plotted results from the rotated problem was scaled by a factor of $\sqrt{2}$, to account for the fact that the CFL number is lower for the rotated problem than for the non-rotated problem.

Figures 3–7 show comparisons of the results on the fine grid of the non-rotated shock-tube problem with the (scaled) results of the rotated shock-tube problem for

3. density (ρ);
4. pressure (p);
5. velocity component normal to the original discontinuity (u_n);

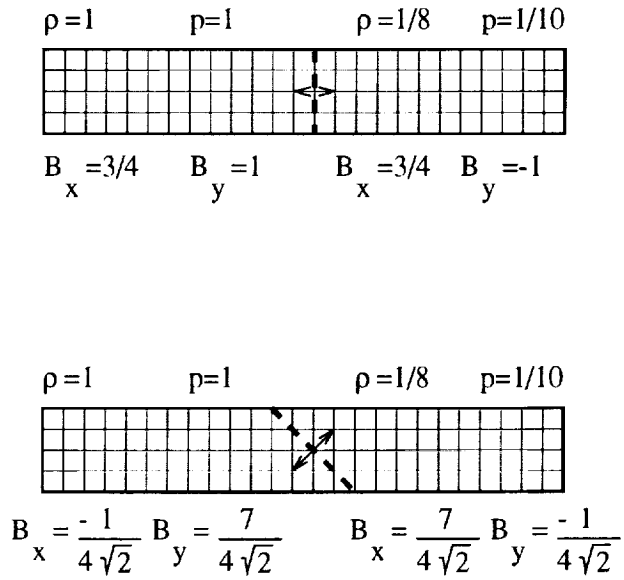


Figure 2: A Test Problem for Two-Dimensional MHD

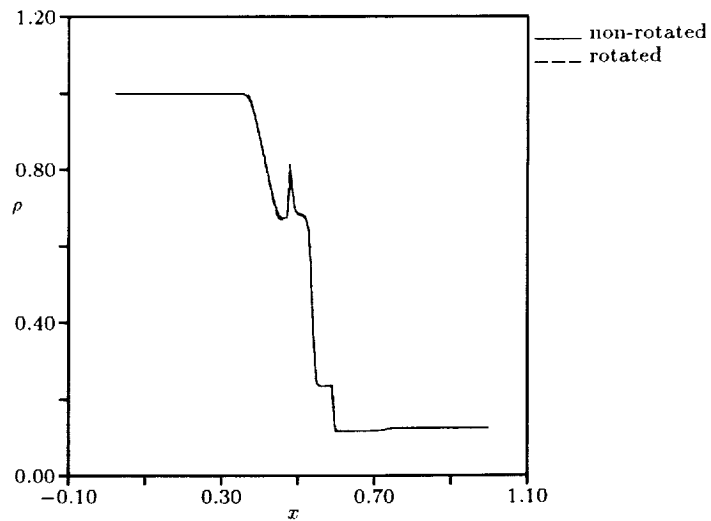


Figure 3: Density in the Rotated and Non-Rotated Shock Tubes

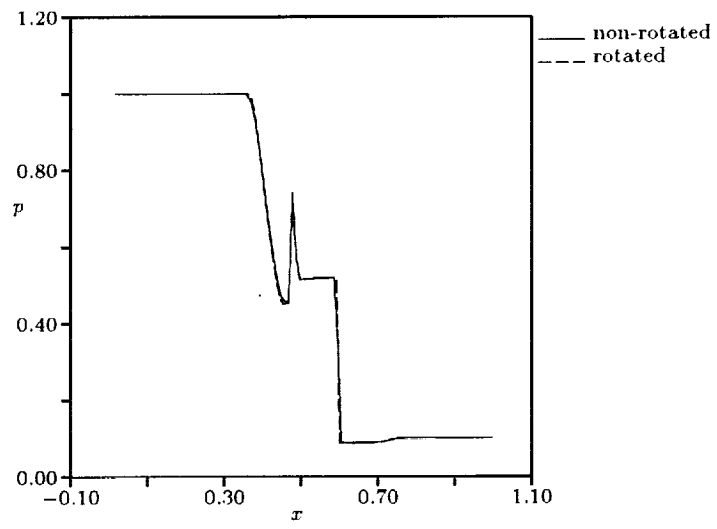


Figure 4: Pressure in the Rotated and Non-Rotated Shock Tubes

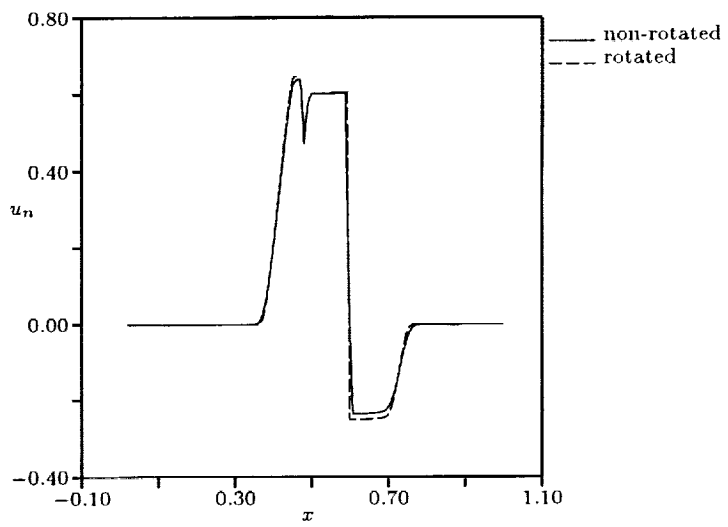


Figure 5: Normal Velocity in the Rotated and Non-Rotated Shock Tubes

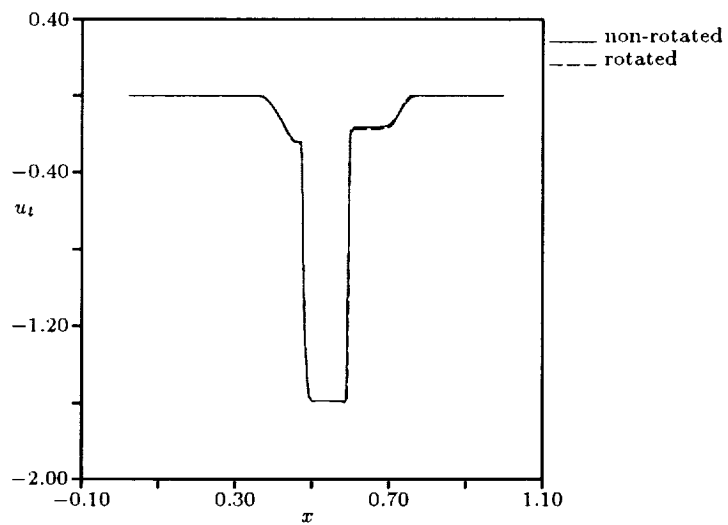


Figure 6: Tangential Velocity in the Rotated and Non-Rotated Shock Tubes

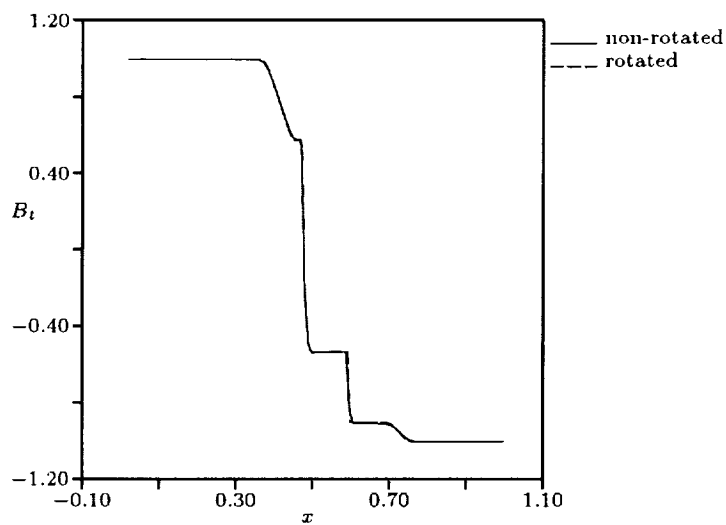


Figure 7: Tangential Magnetic Field in the Rotated and Non-Rotated Shock Tubes

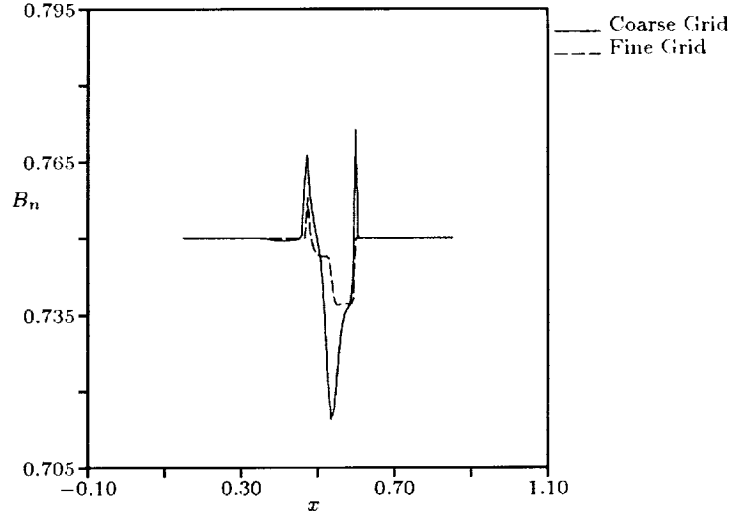


Figure 8: Normal Magnetic Field in the Rotated Shock Tube (Coarse and Fine)

6. velocity component tangential to the original discontinuity (u_t);
7. magnetic-field component tangential to the original discontinuity (B_t).

As can be seen, the agreement is quite good, with the results of the two cases nearly indistinguishable for all but the normal component of velocity. The errors in u_n are balanced by errors in the magnetic-field component normal to the original discontinuity (B_n). Figure 8 shows B_n for the rotated shock-tube problem on the coarse and fine grids. In the non-rotated problem, $B_n = 0.75$ throughout the tube. As can be seen, there are errors on the order of a few percent in B_n on the coarse grid, but the errors are reduced as the grid is refined.

4 Concluding Remarks

In some respects, this paper presents the development of only one-eighth of a Riemann solver. Seven of the eight waves of the Riemann solver are the same as those used in previous work on upwind methods for MHD. The deceptively simple eighth wave that arises from the analysis, however, is of a different character than the other seven — it arises only in multi-dimensional problems, and it is crucial for understanding and solving those problems. It plays the very important role of stabilizing the numerical method with respect to the small amounts of $\nabla \cdot \mathbf{B}$ generated in solving the discrete MHD equations.

Given the meteoric rise of Riemann solvers in the computation of compressible gas dynamics, it is not very risky to predict that schemes based on Riemann solvers will play an

increasingly important role in the computation of compressible conducting flows. The ability of Riemann solvers to capture strong discontinuities robustly and with minimal dissipation, the framework that Riemann solvers provide for implementing stable boundary procedures, and the aesthetically appealing physical basis of Riemann solvers are all strong arguments for their use. The aim of this paper is to remove what is hopefully one of the last major obstacles to the use of Riemann solvers in large-scale codes for computing multi-dimensional conducting flows.

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