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# **Photon number-phase uncertainty relation in the evolution of the field in a Kerr-like medium**

**Fan An-fu**

*Department of Opto-electronics* Science *and Technology,* Sichuan *University, Chengdu 610064, China* and

*China Center of Advanced* Science *and Technology(World Laboratory), P. O. Box 8730, Beijin9 100080, China*

Sun Nian-Chun *Southwest Institute of Technical Physics Chengdu, 610041, China*

#### **Abstract**

**A model of a single-mode field,** initially **prepared in a coherent state, coupled to a two-level atom surrounded by** a **nonlinear Kerr-like medium contained inside a very good quality cavity is considered. We derive the photon number-phase** uncertainty **relation** in **the evolution of the field for a weak** and **strong nonlinear coupling respectively, within the Hermitian phase operator formalism of Pegg** and **Barnett,** and **discuss the effects of nonlinear coupling of thr Kerr-like medium on photon number-phase uncertainty relation of the field.**

## **1 Introduction**

Recently, **Agarwal** et al[1] have considered the propagation **of** a single-mode resonant field through a nonlinear Kerr-medium. Bužek et al[2] have dealed with a combination of two models: the Jaynes-Cummings model (JCM) describing the interaction of a single-mode cavity field with **a single** two-level **atom,** and a nonlinear Kerr-like medium inside **a** cavity which may be modelled by an anharmonic oscillator $[1,3]$ . Particularly, they have showed that with increasing nonlinear coupling the period between the revivals of the atomic inversion is shortened and its time evolution becomes more regular. Besides, they also described the squeezing of the cavity mode and the time evolution of the photon-number distribution.

As is well-known, the phase properties of light field is very important in quantum optics. Lately, Pegg and Barnett[4-6] have shown that an Hermitian phase operator of radiation field exists. It can be constructed from the phase states. This new phase operator formalism makes it possible to describe the quantum properties of optical phase in a fully quantum mechanics. **Gerry['/']** has studied the phase fluctuations of coherent light interacting with the anharmonic oscillator using the Hermitian phase operator. Gantsog et al[8,9] have studied the phase properties of self-squeezed states generated by the anharmonic oscillator, elliptically polarized light propagating through **a** Kerr medium and a damped anharmonic oscillator using the Hermitian phase operator.

In this paper we consider a generalized JCM with an additional Kerr-like medium, namely, a combined model that comprises the JCM and the anharmonic oscillator model (AOM) used to describe a Kerr medium. We deal with not only the field-Kerr medium interaction, but also the field-atom interaction. We derive the photon number-phase uncertainty relation in the evolution of the field for a weak and strong nonlinear coupling respectively, within the Hermitian phase operator formalism of Pegg and Barnett, and discuss the effects of nonlinear coupling of the Kerr medium on the number-phase uncertainty relation of the field.

## **2** The model

We consider a model which consists **of** a single two-level atom surrounded by a nonlinear Kerr-like medium contained in a high-Q single-mode cavity. The cavity mode is coupled to the Kerr-like medium as well as to the two-level atom. The Kerr-like medium can be modelled as an anharmonic oscillator [1,3]. In the adiabatic limit, the effective Hamiltonian of the system involving only the photon and atomic operators in rotating-wave approximation, is  $[1,2]$ 

$$
H_{eff} = \hbar\omega(a^+a + \frac{1}{2}) + \hbar\omega_0 S^z + \hbar\chi a^{+2}a^2 + \hbar g(S^+a + a^+S^-), \tag{2.1}
$$

where *a* and  $a^+$  are the annihilation and creation opreators of the field mode,  $S^{\pm}$  and  $S^z$  are the spin-flip and inversion operators of the atom respectively, *g* is the field-atom coupling constant and  $\chi$  describes the strength of the quadratic nonlinearity modelling the Kerr medium,  $\omega_0$  is the frequency of the atomic transition, the frequency  $\omega$  is

$$
\omega = \omega_f - \lambda^2 / (\omega_k - \omega_f), \qquad (2.2)
$$

Where  $\omega_f$  and  $\omega_k$  are the frequency of the field mode and the anharmonic oscillator modelling the Kerr medium respectively, and  $\lambda$  is the field-Kerr medium coupling constant.

To isolate the effects of the nonlinear coupling of the Kerr medium from that of the finite detuning, we restrict in the case of the resonance  $(i.e.,\omega_0 = \omega)$ . Let us assume that the atom is initially in the excited state  $|e \rangle$  and the field mode is prepared in a coherent state  $|\alpha \rangle$ . The initial state vector  $|\psi(0)\rangle$  of the system is

$$
|\psi(0)\rangle = |\alpha\rangle \otimes |e\rangle = \sum_{n=0}^{\infty} b_n e^{in\beta} |n, e\rangle, \qquad (2.3)
$$

where

$$
b_n = \exp(-\bar{n}/2)(\bar{n}^n/n!)^{1/2}.
$$
 (2.4)

In the interaction picture, the state vector of the system at a later time *t* is found from the Hamiltonian (2.1) to be

$$
|\psi(t)\rangle = \sum_{n=0}^{\infty} b_n e^{in\beta} e^{-i\chi n^2 t} \cdot \{ [\cos(\frac{1}{2}\Omega_{n\chi}t) + i\frac{2\chi n}{\Omega_{n\chi}} \sin(\frac{1}{2}\Omega_{n\chi}t)] |n, e\rangle
$$

$$
-i\frac{2g\sqrt{n+1}}{\Omega_{n\chi}} \sin(\frac{1}{2}\Omega_{n\chi}t) |n+1, g\rangle \}, \tag{2.5}
$$

where  $|g| > 0$  is the ground state of the atom,  $\Omega_{n\chi}$  is the generalized Rabi frequency defined by

$$
\Omega_{n\chi} = [4g^2(n+1) + 4\chi^2 n^2]^{1/2},\tag{2.6}
$$

It is obvious that, for  $\chi = 0$ , the state vector  $|\psi(t) >$  given by Eq.(2.5) describes the dynamics of the ordinary JCM.

## 3 The phase variance of the cavity field

Based on the Hermitian phase formalism of Pegg and Barnett[4-6], The complete set of  $s + 1$ orthonormal phase state is defined by

$$
|\theta_m\rangle = \frac{1}{\sqrt{s+1}} \sum_{n=0}^{s} \exp(in\theta_m)|n\rangle, \tag{3.1}
$$

where  $\theta_m = \theta_0 + 2\pi m/(s+1)$ ,  $m = 0, 1, 2, \dots$ , *s*, and  $\theta_m$  is an arbitrary real number. The Hermitian phase operator is given by

$$
\hat{\Phi}_{\theta} = \sum_{m=0}^{s} \theta_m |\theta_m\rangle \langle \theta_m|, \tag{3.2}
$$

Clearly, phase state  $|\theta_m\rangle$  are eigenstates of  $\hat{\Phi}_{\theta}$  with the eigenvalues  $\theta_m$ . The eigenvalues  $\theta_m$  are restricted to lie within a phase window between  $\theta_0$  and  $(\theta_0 + 2\pi)$ . It has to be noted That, after all expectation values of the phase variables associated with the phase properties of the field have been calculated in the finite  $(s + 1)$ -dimensional space, *s* is allowed to tend to infinity. The phase distribution of the state given by Eq.(2.5) is

$$
p(\theta_m, t) = \vert < \theta_m \vert \psi(t) > \vert^2,\tag{3.3}
$$

with the expectation value and the variance

$$
\langle \hat{\Phi}_{\theta} \rangle = \sum_{m} \theta_{m} P(\theta_{m}, t), \tag{3.4}
$$

$$
\langle \Delta \hat{\Phi}_{\theta}^{2} \rangle = \sum_{m} (\theta_{m} - \langle \hat{\Phi}_{\theta} \rangle)^{2} P(\theta_{m}, t). \tag{3.5}
$$

We choose the reference phase  $\theta_0 = \beta - \pi s/(s + 1)$ , and introduce a new phase label  $\mu = m - s/2$ , which goes in integer steps from  $(-s/2)$  to  $(s/2)$ . Then the phase distribution becomes symmetric in  $\mu$ . In the limit as *s* tends to infinity, the continuous phase variable can be introduced replacing  $\mu 2\pi/(s+1)$  by  $\theta$  and  $2\pi/(s+1)$  by  $d\theta$ . Then we can find a continuous phase distribution

$$
P(\theta, t) = \frac{1}{2\pi} \{ 1 + 2 \sum_{n > n'} b_n b_{n'} [A_{nn'} \cos[(n - n')\theta + (n^2 - n'^2)\chi t] + B_{nn'} \sin[(n - n')\theta + (n^2 - n^2)\chi t]] \},
$$
\n(3.6)

where

$$
A_{nn'} = \cos(\frac{1}{2}\Omega_{n\chi}t)\cos(\frac{1}{2}\Omega_{n'\chi}t) + \frac{4g^2\sqrt{(n+1)(n'+1) + 4\chi^2nn'}}{\Omega_{n\chi}\Omega_{n'\chi}}\sin(\frac{1}{2}\Omega_{n\chi}t)\sin(\frac{1}{2}\Omega_{n'\chi}t)
$$

$$
B_{nn'} = (2\chi n/\Omega_{n\chi})\sin(\frac{1}{2}\Omega_{n\chi}t)\cos(\frac{1}{2}\Omega_{n'\chi}t) - (2\chi n'/\Omega_{n'\chi})\cos(\frac{1}{2}\Omega_{n\chi}t)\sin(\frac{1}{2}\Omega_{n'\chi}t). \tag{3.7}
$$

The function  $P(\theta, t)$  is normalized so that

$$
\int_{-\pi}^{\pi} P(\theta, t) d\theta = 1. \tag{3.8}
$$

If the mean photon number in the field is large,  $\bar{n}$  >> 1, the coefficient  $b_n$  in Eq.(2.4) can be well approximated by a continuous Gaussian distribution

$$
b_n = (2\pi \bar{n})^{-1/4} \exp[-(\bar{n} - n)^2/4\bar{n}]. \tag{3.9}
$$

If Eq. (3,9) **is substituted into** Eq.(3.6) and the **summation** in Eq. (3.6) is replaced by an appropriate integral over the variable *n,* one can appraximately work the phase distribution.

We will consider two limit cases:

(1) The weak nonlinear coupling which is defined by the condition  $g^2\bar{n} >> \chi^2\bar{n}^2$ , with  $\bar{n} >> 1$ . In this case the generalized Rabi frequency can be approximated as

$$
\Omega_{n\chi} \approx g[(\bar{n})^{1/2} + n(\bar{n})^{-1/2}](1 + \frac{1}{2}\varepsilon_{\chi}^2), \tag{3.10}
$$

**where**

$$
\varepsilon_{\chi} = (\chi^2 \bar{n}^2 / g^2 \bar{n})^{1/2}.
$$
\n(3.11)

Then using Eqs.(3,9) and (3.10) **and** replacing **the summation** in Eq.(3.6) **by** an **appropriate** integral over *n,* we obtain

$$
P(\theta, t) = \frac{1}{2\pi} \left( \frac{2\pi \bar{n}}{1 + 16\bar{n}^2 \chi^2 t^2} \right)^2 \cdot \left\{ (1 + \varepsilon_\chi) \exp\left[-\frac{2\bar{n}}{1 + 16\bar{n}^2 \chi^2 t^2} [2\bar{n}\chi t + (\theta - \frac{gt}{2\sqrt{\bar{n}}})^2] \right] + (1 - \varepsilon_\chi) \exp\left[-\frac{2\bar{n}}{1 + 16\bar{n}^2 \chi^2 t^2} [2\bar{n}\chi t + (\theta + \frac{gt}{2\sqrt{\bar{n}}})^2] \right] \right\}.
$$
 (3.12)

According to Eqs.(3.4) and (3.5), using Eq.(3.12), and taking into account  $\theta_m = \theta + \beta$ , we can directly find an expectation value of the phase operator and its variance

$$
\langle \hat{\Phi}_{\theta} \rangle = \beta + \varepsilon_{\chi} \frac{gt}{2\sqrt{\bar{n}}}(1 - 4\bar{n}\varepsilon_{\chi}), \tag{3.13}
$$

$$
<\Delta \hat{\Phi}_{\theta}^2> = \frac{1}{4\bar{n}} + \frac{(gt)^2}{4\bar{n}} (1 + 16\bar{n}\varepsilon_{\chi}^2). \tag{3.14}
$$

(2) The strong nonlinear coupling which is defined by the condition  $g^2\bar{n} << \chi^2\bar{n}^2$ , with  $\bar{n} >> 1$ . In this case the generalized Rabi frequency can be approximated as

$$
\Omega_{nx} \approx 2\chi \bar{n} (1 + \frac{1}{2} \varepsilon_g^2), \tag{3.15}
$$

where

$$
\varepsilon_g = (g^2 \bar{n}/\chi^2 \bar{n}^2)^{1/2}.
$$
\n(3.16)

**Then, we find**

$$
P(\theta, t) = \frac{(8\pi\bar{n})^{1/2}}{2\pi} (1 + 16\bar{n}^2 \chi^2 t^2)^{-1/2} \exp[-\frac{2\bar{n}}{1 + 16\bar{n}^2 \chi^2 t^2} (\theta + 2\bar{n} \chi t)^2]
$$
(3.17)

$$
\langle \hat{\Phi}_{\theta} \rangle = \beta - 2\bar{n}\chi t, \tag{3.18}
$$

$$
\langle \Delta \hat{\Phi}_{\theta}^2 \rangle = \frac{1}{4\bar{n}} + 4\bar{n}(\chi t)^2. \tag{3.19}
$$

We see that the average value of the phase is not equal to the initial quantity  $\beta$ , and the **phase variance is** always **enhanced. For the weak nonliner coupling, the enhancement of the phase variance** is proportional to  $(gt)^2/4\bar{n}$ , this is similar to that of the resonance field in coherent state **JCM[10]. For the strong nonlinear** coupling, **the enhancement of the phase variance is proportional**  $\tan \frac{1}{2}$ . Obviously, the enhancement of the phase variance in the strong nonlinear coupling **case is larger than that in the weak nonlinear coupling case.**

### **4 The number-phase uncertainty relation**

**It is not difficult to calculate the variance of the photon-number for the state** given **by Eq.(2.5).** Using  $\hat{n} = a^{\dagger}a$ , for the weak nonlinear coupling we obtain

$$
<(\Delta \hat{n})^2>=<\hat{n}^2>-<\hat{n}>^2\approx \bar{n}+\sqrt{\bar{n}}gt\exp[-(gt)^2/2]\sin(2\sqrt{\bar{n}}gt).
$$
 (4.1)

 $\epsilon$ <sub>i</sub>. From Eqs. (3.14) and (4.1), we find that the number-phase uncertainty relation is

$$
<(\Delta \hat{n})^2><\Delta \hat{\Phi}_{\theta}^2> \approx \frac{1}{4} + \frac{1}{4} (gt)^2 \{1 + 16 \bar{n} \varepsilon_{\chi}^2 + 16 \sqrt{\bar{n}} g t \varepsilon_{\chi}^2 \exp[-(gt)^2/2] \sin(2\sqrt{\bar{n}} gt)\}.
$$
 (4.2)

For the strong nonlinear coupling, **we** find

$$
\langle (\Delta \hat{n})^2 \rangle \approx \bar{n} + 2\bar{n}\chi t \varepsilon_g^2 \exp[-2\bar{n}(\chi t)^2] \sin(2\bar{n}\chi t), \tag{4.3}
$$

$$
<(\Delta\hat{n})^2><\Delta\hat{\Phi}_{\theta}^2>\approx\frac{1}{4}+4\bar{n}^2(\chi t)^2\{1+2\chi t\epsilon_{\theta}^2\exp[-2\bar{n}(\chi t)^2]\sin(2\bar{n}\chi t)\}.
$$
 (4.4)

We **see that, the** uncertainty **product** during **the evolution is expanded. The expansion of the** uncertainty **product is fast in strong nonlinear** coupling case. **This is similar to that of the self-squeezed** state **generated** by **the** anharmonic **oscillator[8]. This occurs because** both **the field**atom **interaction[11]** and **the field-Kerr** medium **interaction** are **nonlinear,** moreover **the nonlinear interaction strength** of **the latter is** larger **than that of the former.**

#### **5 Summary**

**In the present paper we** consider a generalized **JCM in the presence** an additional **Kerr-like** medium, **we have** derived the **photon** munber-phase uncertainty **relation in the evolution of** a **resonant field for** a **weak** mid **strong** nonlinear **coupling, within the Hermitian phase operator** formalism of Pegg and Barnett. We have shown that the nonlinear coupling of the cavity mode to Kerr-like medium leads to the enhancement of the phase variance of the field and the expansion of the uncertainty product. Particularly, the expansion of the uncertainty product is fast in strong nonlinear coupling case. We have indicated that this is similar to that of the self-squeezed state generated by the anharmonic oscillator. We have also indicated that the nonlinear interaction strength of the field-Kerr medium is larger than that of the field-atom.

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