



# **RESOLVE:** an International Mission to Search for Volatiles at the Lunar Poles

(Regolith & Environmental Science and Oxygen & Lunar Volatiles Extraction)

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# Our Evolving Understanding of the Moon and its Resources



**RESOLVE:** Regolith & Environment Science and Oxygen & Lunar Volatile Extraction

Integrated data sets from instruments on LRO support the existence of large quantities of water ice in the PSRs and in partially sunlit regions



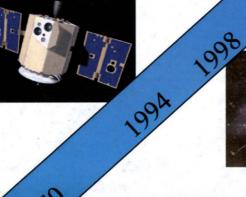
2000

Synthetic Aperture Radar on Chandrayaan 1 returns data that is consistent with water ice in the PSR's



Clementine's Bi-Static Radar suggest Water Ice in permanently shadowed regions near the poles

Watson, Murray and Brown theorize that cold traps at the moon's poles may contain water ice



Apollo samples point to a dry Moon LCROSS impacts Cabeus A and clearly detects significant quantities of water in the ejecta

Neutron Spectrometer aboard Lunar Prospector detects elevated levels of hydrogen that correlates with permanent shadow



## LCROSS & LRO Definitively Prove Existence of Volatiles at the Lunar Poles



**RESOLVE:** Regolith & Environment Science and Oxygen & Lunar Volatile Extraction

					Instrument			
	Column Density (# m <sup>-2</sup> )	Relative to H2O(g) (NIR spec only)	Concentration (%)	Long-term Vacuum Stability Temp (K)	UV/Vis	NIR	LAMP	M3
СО	1.7e13±1.5e11		5.7	15			x	
H <sub>2</sub> O(g)	5.1(1.4)E19	1	5.50	106		x		
H <sub>2</sub>	5.8e13±1.0e11		1.39	10			x	
H <sub>2</sub> S	8.5(0.9)E18	0.1675	0.92	47	x	x		
Ca	3.3e12±1.3e10		0.79				x	
Hg	5.0e11±2.9e8		0.48	135			x	
NH <sub>3</sub>	3.1(1.5)E18	0.0603	0.33	63		x		
Mg	1.3e12±5.3e9		0.19				x	
SO <sub>2</sub>	1.6(0.4)E18	0.0319	0.18	58		x		
C <sub>2</sub> H <sub>4</sub>	1.6(1.7)E18	0.0312	0.17	~50		x		
CO <sub>2</sub>	1.1(1.0)E18	0.0217	0.12	50	x	x		
CH <sub>3</sub> OH	7.8(42)E17	0.0155	0.09	86		x		
CH <sub>4</sub>	3.3(3.0)E17	0.0065	0.04	19		x		
ОН	1.7(0.4)E16	0.0003	0.002	>300 K if adsorbed	x	x		x
H <sub>2</sub> O (adsorb)			0.001-0.002					x
Na		1-2 kg		197	x			
CS					x			
CN					x			
NHCN	×				x			
NH					x			
NH <sub>2</sub>					x			

Volatiles comprise possibly 15% (or more) of LCROSS impact site regolith

# What's the Next Step?

- We now know with certainty that there are volatiles at one spot on the moon.
- Comparison's of orbital instrument data with the LCROSS plume seem to suggest that the water is not evenly distributed.
- Until we know the distribution and accessibility of the volatiles don't really know if we have a usable resource.
- A "Ground Truth" surface mission is the next logical step.
- RESOLVE is the payload that NASA and the CSA are designing to answer these questions



## **RESOLVE Mission Requirements**



**RESOLVE:** Regolith & Environment Science and Oxygen & Lunar Volatile Extraction

## **Primary Mission:**

- ✓ Verify the existence of and characterize the constituents and <u>distribution</u> of water and other volatiles in lunar polar surface materials
  - Map the surface distribution of hydrogen rich materials (Neutron Spectrometer, Near-IR Spectrometer)
  - Extract 1m core sample with minimal loss of volatiles from selected sites (Drill /Auger Subsystem)
    - to a depth of 1m
  - Heat multiple samples from each core to drive off volatiles for analysis (OVEN Subsystem)
    - from 100°K to 473°K
    - from 0 up to 100 psia (reliably seal in aggressively abrasive lunar environment)
  - Determine the constituents and quantities of the volatiles extracted (LAVA Subsystem)
    - Hope to find and quantify H2, He, CO, CO2, CH4, H2O, N2, NH3, H2S, SO2
    - Survive limited exposure to HF, HCI, and Hg

## **Secondary Mission:**

- ✓ Demonstrate the ISRU Hydrogen Reduction Process to extract oxygen from lunar regolith
  - Heat sample to reaction temperature (OVEN Subsystem)
    - from 473°K to 1173°K
  - Flow H2 through regolith to extract oxygen in the form of water (OVEN Subsystem)
  - Capture, quantify, and display the water generated (LAVA Subsystem)



## RESOLVE (Regolith & Environment Science and Oxygen and Lunar Volatile Extraction)



**RESOLVE: Regolith & Environment Science and Oxygen & Lunar Volatile Extraction** 

## Sample Acquisition -

## Auger/Core Drill [CSA provided]

- Complete core down to 1 m; Auger to 0.5 m
- Minimal/no volatile loss
- Low mass/power (<25 kg)</li>
- Wide variation in regolith/rock/ice characteristics for penetration and sample collection
- Wide temperature variation from surface to depth (300K to <100K)</li>

## Sample Evaluation -

## Near Infrared Spectrometer (NIR)

- · Low mass/low power for flight
- Mineral characterization and ice/water detection before volatile processing
- Controlled illumination source

## **Resource Localization –**

### Neutron Spectrometer (NS)

- · Low mass/low power for flight
- Water-equivalent hydrogen ≥ 0.5 wt% down to 1 meter depth at 0.1 m/s roving speed

## **RESOLVE Instrument Suite Specifications**

- Nom. Mission Life = 10+ Cores, 12+ days
- Mass = 60-70 kg
- Dimensions = w/o rover: 68.5 x 112 x 1200 cm
- Ave. Power; 200 W

## Volatile Content/Oxygen Extraction -

## Oxygen & Volatile Extraction Node (OVEN)

- Temperature range of <100K to 900K</li>
- 50 operations nominal
- · Fast operations for short duration missions
- Process 30 to 60 gm of sample per operation (Order of magnitude greater than TEGA & SAM)

## Volatile Content Evaluation -

Lunar Advanced Volatile Analysis (LAVA)

- Fast analysis, complete GC-MS analysis in under 2 minutes
- Measure water content of regolith at 0.5% (weight) or greater
- Characterize volatiles of interest below 70 AMU

## Operation Control – Flight Avionics [CSA/NASA]

Space-rated microprocessor

# Surface Mobility/Operation [CSA mobility platform]

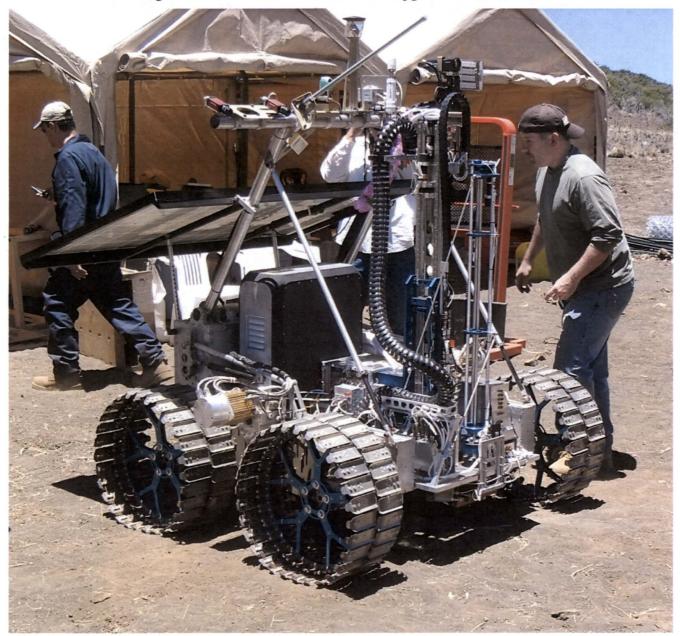
- Low mass/large payload capability
- · Driving and situation awareness, stereo-cameras
- Autonomous navigation using stereo-cameras and sensors
- NASA contributions likely for communications and thermal management



## RESOLVE 3<sup>rd</sup> Generation Prototype Near Flight Mass, Volume and Power



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## Permanent Shadowed Craters?

- LRO radar data suggests large, thick deposits of water ice in some of the Permanently Shadowed Craters
- However, temperatures are extremely low (<40K), and a mission of any significant duration would probably require a nuclear energy source.
  - Mission would be prohibitively expensive for our current budget environment.

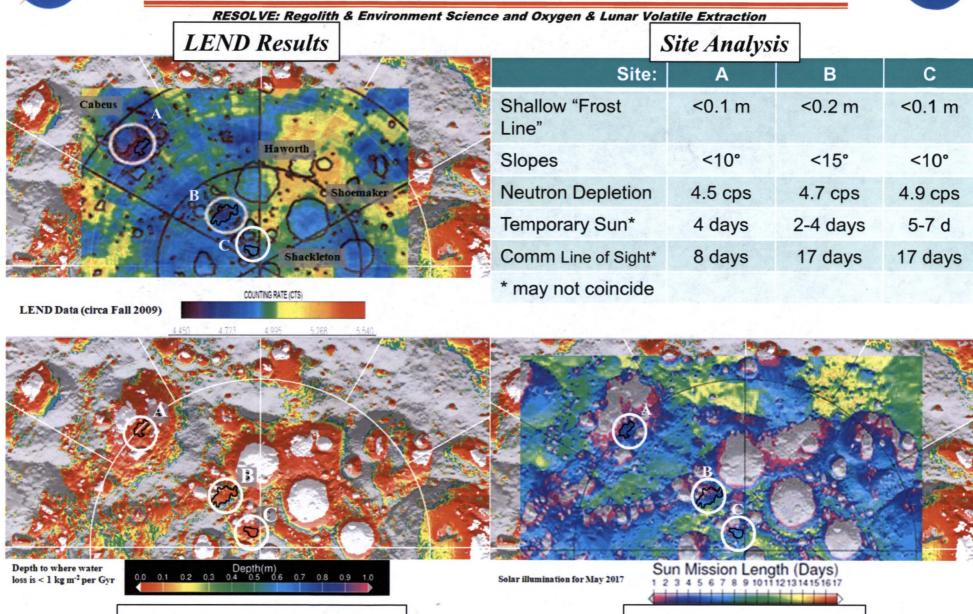
## Partially sunlit regions?

- Lunar Exploration Neutron Detector (LEND) suggests that there are areas of neutron suppression (indicator of hydrogen) outside of the permanently shadowed regions.
- David Paige and the DIVINER radiometer team published results indicating that there are many areas in the polar regions that have subsurface temperatures (<100K) that would support the existence of water ice.</li>
- Solar powered missions are more affordable and the operating environment for hardware is much less harsh.
- Perhaps a location like this would make it easier to set up a future mining operation on the Moon if the resources were plentiful enough.

# NASA

## **RESOLVE Mission Options – Potential South Pole Landing Sites**





**Predicted Volatile Stability** 

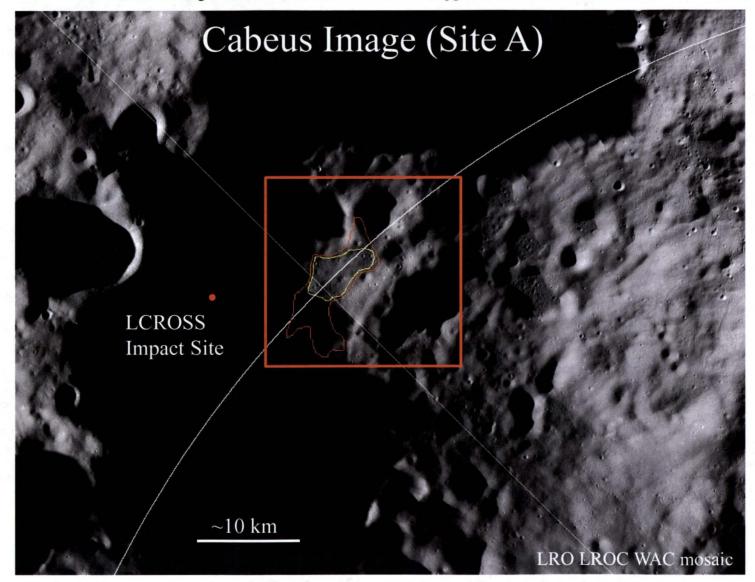
Solar Power Potential



## RESOLVE Mission Options – Potential South Pole Landing Sites



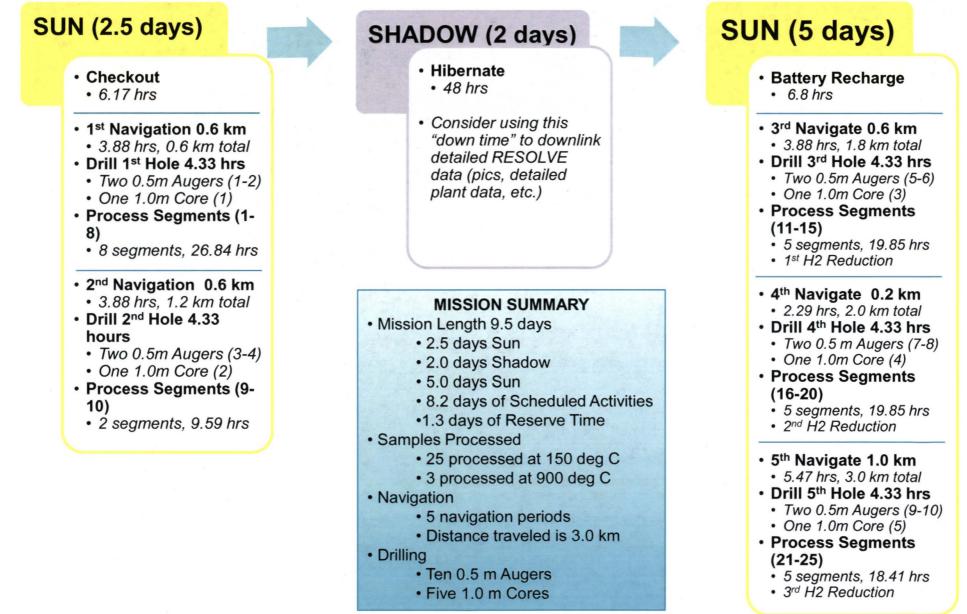
**RESOLVE: Regolith & Environment Science and Oxygen & Lunar Volatile Extraction** 



# Sun and Shadow Ops

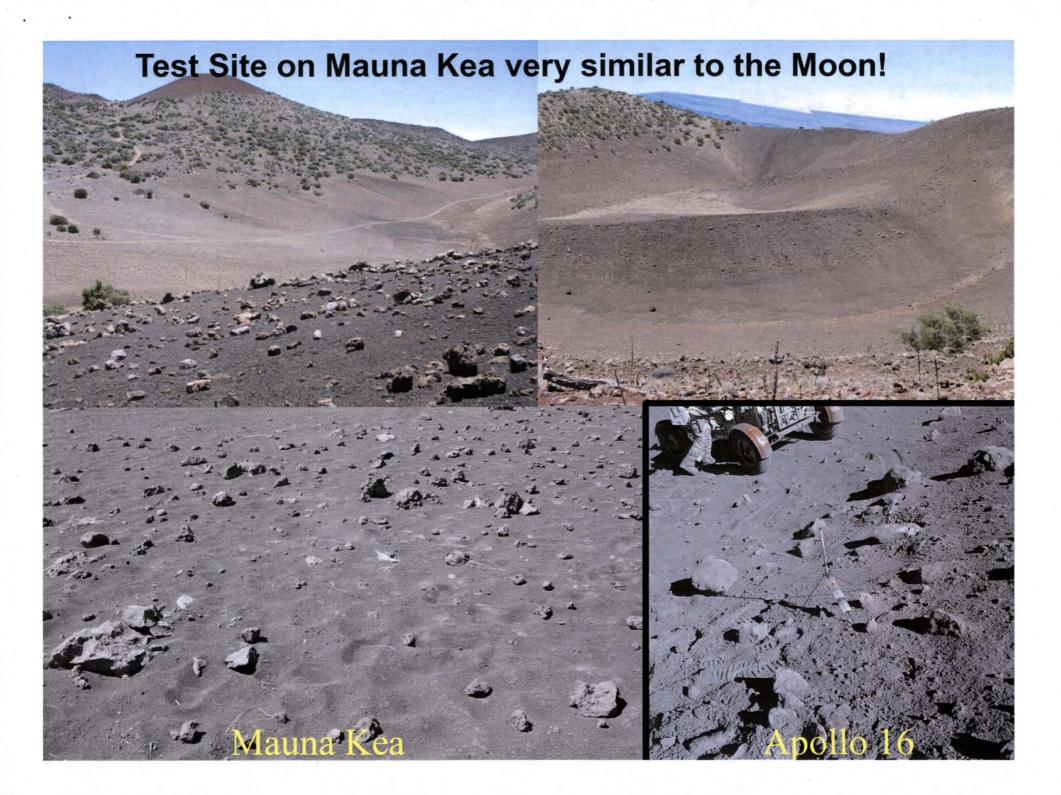


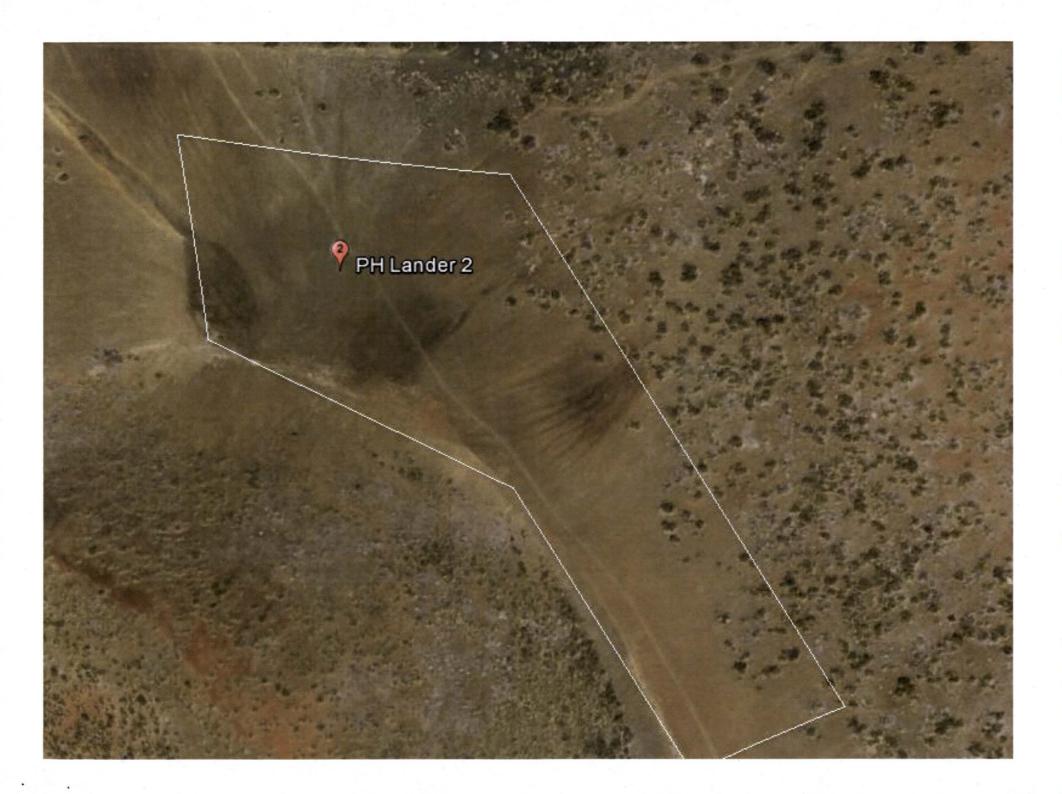
#### **RESOLVE: Regolith & Environment Science and Oxygen & Lunar Volatile Extraction**



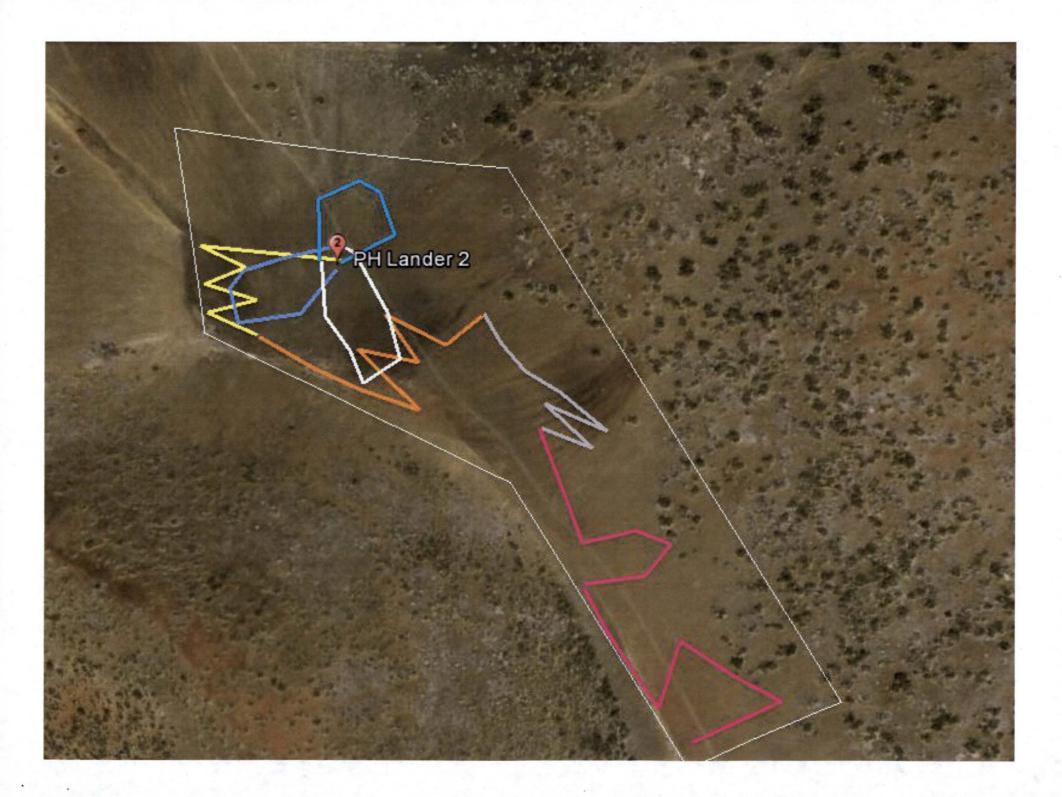
# **Resource Prospector Mission Simulation**

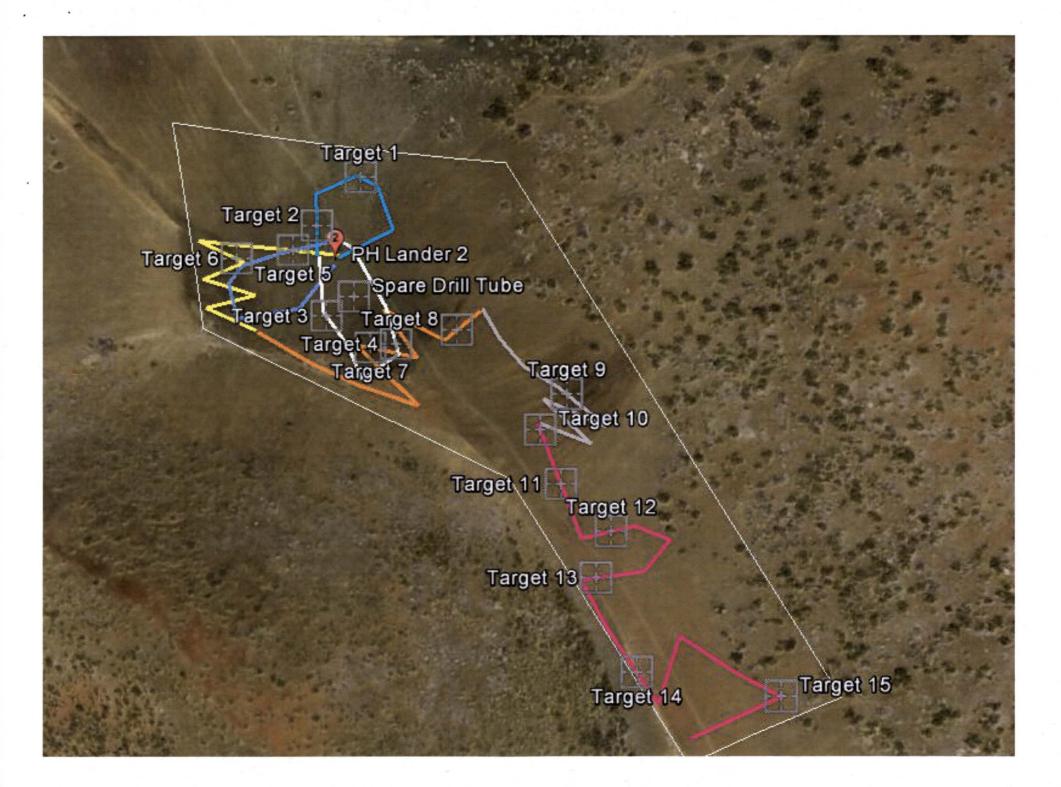
- This mission has the most challenging surface mobility timeline ever considered
  - To insure that the mission objectives can be met a mission simulation using the Field Prototype was executed
- Goal was to test the hardware in a harsh environment, and test our ability to meet mission objectives in a restricted timeline.
- Full Mission Control with Shift Operations planned.
  Test was conducted on the slopes of the volcano Mauna Kea (background image)
  - Some accommodations had to be made due to natural water table, solar power availability, safety, etc.











# **Test Site Preparation**



Clear area of grass

NASA



Loading/Compacting Tephra

Prepare polyethylene targets





Blend target areas in with surrounding

Sample tube in hole

Rover Communications (CSA)

Navigation & Situational Awareness Cameras & Lights (CSA)

Solar Array (CSA)

> Artemis Jr. Rover (CSA)

LAVA Gas Chromatograph/ Mass Spectrometer (NASA)

> OVEN Sample Heating Unit (NASA)

Total Station-Relative Navigation (CSA)

> Situational Awareness Camera & Lights (CSA)

DESTIN Drill System (CSA)

> Avionics & Software (CSA & MASA)

Spectrometer (MASA)

Neutron Spectrometer (NASA)

Communications (NASA)

> Situational Awareness Camera (NASA)

Mission Control: Timeline, Trave & Data Display Software (NASA



# **On-Site & Remote Operations Centers**

5 Centers; 4 different time zones



**On-Site Control Center** 

## ExDoc at CSA HQ

## Johnson Space Center

CSA ASC



Science Backroom at ARC

## Lunar Polar Resource Mission Simulation 'Flight' like hardware and operations





NASA

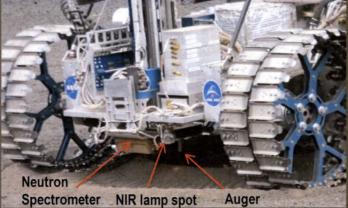
Rover Egress from Lander



Rover Searching Exploration Site



Data from Neutron Spectrometer and Rover Navigation displayed on xGDS showing 'hot spot' found by RESOLVE





Darker cuttings appear at a depth of 15-120 cm

Auger and Examine Cutting Pile for Ice with Near Infrared Spectrometer



Drilling, Sample Collection, Sample Transfer, & Processing to Measure Water and Other Volatiles



# **RESOLVE Mission Objectives/Results**



## CAT 1 Objectives (Mandatory)

✓ Travel at least 100m on the lunar surface to map the horizontal distribution of volatiles (FD1)

## CAT 2 Objectives (Highly Desirable)

✓ Perform at least 1 coring operation. Process all regolith in the drill stem acquired during the coring operation. (FD2)

✓ Perform at least 1 water droplet demo during volatile analysis. (FD2)



## CAT 3 Objectives (Desirable)

- ✓ Map the horizontal distribution of voiatiles over a point to point distance of 500m (FD4) (lunar objective is 1km.)
- Perform coring operations and process regolith at a minimum of 3 locations. (ED4)
- Volatile analysis will be performed on at least 4 segments from each core to achieve a vertical resolution of 25cm or better.
- Perform a minimum of 3 Augering operations (FD4). (Note that the Lunar objective is 6)
- Perform at least 2 total water droplet demos.
   Perform 1 in conjunction with hydrogen reduction and perform 1 during low temperature volatile analysis.

## CAT 4 Objectives (Goals)

- Perform 2 coring operations be separated by at least 500 m straight line distance. (FD4) (lunar objective is 1km.)
- 2. Travel 3 km total regardless of direction
- 3. Travel directly to local areas of interest
- associated with possible retention of hydrogen
- Process regolith from 5 cores
- further develop innar exploration technologies

# **Mission Simulation Conclusions**

As expected, the mission timeline is extremely challenging
 Problems, which are always going to occur, delay ability to achieve objectives

 Essential to develop contingency plans for as many problems that can be envisioned

Lost a lot of time during simulation troubleshooting when we should have just been executing alternate plans

 In spite of the hardware and operational problems experienced we achieved most of the mission objectives

 Vaporizing, transferring and sampling the water vapor took longer than expected (need to increase transfer tank size)

OVEN design may be sensitive to slope rover is on

Bottom line.... Mission appears to be achievable.

Development will continue

# 2013 Design & Test Activities

 In FY13 the major subsystems have been designing and fabricating engineering test prototype

- Risk reduction activity will test prototypes in flight conditions
- Some integration of subsystems that are tightly coupled is likely to occur (e.g. OVEN and LAVA)
- More detailed assessment of potential landing sites will occur
  - We'd like to find an area of partial sunshine close to an area of permanent shadow
  - North Pole sites will be analyzed in more detail
  - Critical Trade Studies will be conducted
    - Thermal management, Communications architecture, Power management, etc.

# The Path to Flight

Some important decisions remain open

- Canada has two open design study contracts on the flight rover and sample acquisition system (decision late 2013)
- NASA must decide on a Lander (decision this summer)
  - Internal design based on Morpheus and/or Mighty Eagle?
  - Commercial Partner?
  - International partner?
- Mission Concept Review scheduled for mid-September
   Flight Program Office anticipates Authority to Proceed (ATP) in 2014
- Assuming we receive ATP
  - Mission design activities would take place in 2014-16
  - Integration and final test in 2017
  - Launch/Landing April/May of 2018
    - Date dependent on final landing site selection

# Questions? ENCE & OXYGEN &