

CONTROL ID: 1802617

TITLE: The XRD Amorphous Component in John Klein Drill Fines at Yellowknife Bay, Gale Crater, Mars

AUTHORS (FIRST NAME, LAST NAME): Richard V Morris¹, Douglas W Ming¹, David Blake², David Vaniman³, David L Bish⁵, Steve Chipera⁶, Robert Downs⁴, Shaunna Morrison⁴, Ralf Gellert⁷, Iain Campbell⁷, Allan H Treiman⁸, Cherie Achilles¹, Thomas Bristow², Joy A Crisp⁹, Amy McAdam¹⁰, Paul Douglas Archer¹, Brad Sutter¹, Elizabeth B Rampe¹, MSL Science Team⁹

INSTITUTIONS (ALL): 1. NASA Johnson Space Center, Houston, TX, United States.
2. NASA Ames Research Center, Moffett Field, CA, United States.
3. Planetary Science Institute, Tucson, AZ, United States.
4. University of Arizona, Tucson, AZ, United States.
5. University of Indiana, Bloomington, IN, United States.
6. Chesapeake Energy, Oklahoma City, OK, United States.
7. University of Guelph, Guelph, ON, Canada.
8. Lunar & Planetary Institute, Houston, TX, United States.
9. JPL-Caltech, Pasadena, CA, United States.
10. NASA Goddard Space Flight Center, Greenbelt, MD, United States.

ABSTRACT BODY: Drill fines of mudstone (targets John Klein and Cumberland) from the Sheepbed unit at Yellowknife Bay were analyzed by MSL payload elements including the Chemistry and Mineralogy (CheMin), APXS (Alpha Particle X-Ray Spectrometer), and Sample Analysis at Mars (SAM) instruments. CheMin XRD results show a variety of crystalline phases including feldspar, pyroxene, olivine, oxides, oxyhydroxides, sulfates, sulfides, a tri-octahedral smectite, and XRD amorphous material. The drill fines are distinctly different from corresponding analyses of the global soil (target Rocknest) in that the mudstone samples contained detectable phyllosilicate. Here we focus on John Klein and combine CheMin and APXS data to calculate the chemical composition and concentration of the amorphous component.

The chemical composition of the amorphous plus smectite component for John Klein was calculated by subtracting the abundance-weighted chemical composition of the individual XRD crystalline components from the bulk composition of John Kline as measured by APXS. The chemical composition of individual crystalline components was determined either by stoichiometry (e.g., hematite and magnetite) or from their unit cell parameters (e.g., feldspar, olivine, and pyroxene). The chemical composition of the amorphous + smectite component (~71 wt.% of bulk sample) and bulk chemical composition are similar. In order to calculate the chemical composition of the amorphous component, a chemical composition for the tri-octahedral smectite must be assumed. We selected two tri-octahedral smectites with very different MgO/(FeO + Fe₂O₃) ratios (34 and 1.3 for SapCa1 and Griffithite, respectively). Relative to bulk sample, the concentration of amorphous and smectite components are 40 and 29 wt.% for SapCa1 and 33 and 36 wt.% for Griffithite. The amount of smectite was calculated by requiring the MgO concentration to be ~0 wt.% in the amorphous component. Griffithite is the preferred smectite because the position of its 021 diffraction peak is similar to that reported for John Klein. In both cases, the amorphous component has low SiO₂ and MgO and high FeO + Fe₂O₃, P₂O₅, and SO₃ concentrations relative to bulk sample. The chemical composition of the bulk drill fines and XRD crystalline, smectite, and amorphous components implies alteration of an initially basaltic material under near neutral conditions (not acid sulfate), with the sulfate incorporated later as

veins of CaSO₄ injected into the mudstone.

KEYWORDS: 5470 PLANETARY SCIENCES: SOLID SURFACE PLANETS Surface materials and properties, 5410 PLANETARY SCIENCES: SOLID SURFACE PLANETS Composition, 5415 PLANETARY SCIENCES: SOLID SURFACE PLANETS Erosion and weathering.

(No Image Selected)

(No Table Selected)

Additional Details

Previously Presented Material: 0% previously presented or published

Contact Details

CONTACT (NAME ONLY): Richard Morris

CONTACT (E-MAIL ONLY): richard.v.morris@nasa.gov

TITLE OF TEAM:
