

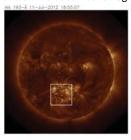
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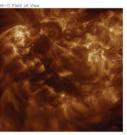


Abstract: The High-Resolution Coronal Imager (Hi-C) flew aboard a NASA sounding rocket on 2012 July 11 and captured roughly 345 s of high spatial and temporal resolution images of the solar corona in a narrowband 193 Angstrom channel. We have analyzed the fluctuations in intensity of Active Region 11520. We selected events based on intensities greater than a threshold determined from the photon and readout noise. We compare the Hi-C events with those determined from AIA. We find that Hi-C detects shorter and smaller events than AIA. We also find that the intensity increase in the Hi-C events is ~3 times greater than the intensity increase in the AIA events. We conclude the events are related to linear sub-structure that is unresolved by AIA.

Background:

- Hi-C flew on July 11, 2012 and collected 37 full frame images. The first 7 were blurred due to spacecraft motion, we omit these images in this analysis.
- Hi-C observed the Sun in the 193 A passband, similar to the 193 passband on the Atmospheric Imaging Array (AIA).
- Hi-C has 0.1"/pixel, $\sim 0.3"$ resolution, 2 s exposure time, 5.5 s cadence.
- Hi-C field of view included Active Region 11520.

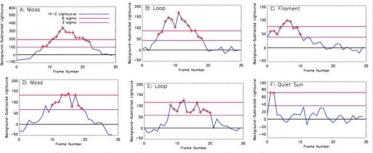




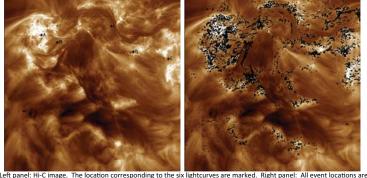
Analysis:

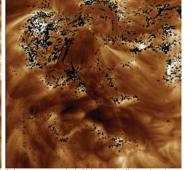
- The Hi-C data set was co-aligned and converted to photons. The lightcurve of each pixel was analyzed individually. The 8th brightest frame in the lightcurve was used to approximate background, $I_{background}$. The noise in the signal was calculated as: $\sigma^2 = I_{background} + \sigma^2_{readout}$ The readout noise of the Hi-C detector varied by quadrant. We used the value
- from the noisiest quadrant ($\sigma = 100$ electrons = 5.6 photons).
- We identified pixels where the intensity went above 6σ as event pixels.
- We measured the lifetime of the event as the time the intensity stayed above 3σ. We only consider events that started and finished during the observation
- For each 6σ pixel, we found all surrounding pixels that are at least 3σ above the background at the same time to calculate the event area.

Example Lightcurves:



Events were identified as having at least one frame at the 6 sigma level. All frames with intensity > 3 sigma around the 6 sigma frame (shown with red +) were included in the lifetime

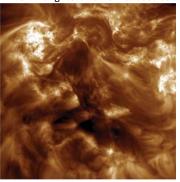


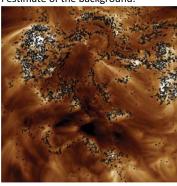


with the black contours. Events tend to occur in moss regions (e.g., A and E) or in locations of quickly evolving loops (e.g., B and D). The long extended loops do not show dynamic behavior

AIA Analysis:

- Identical analysis was applied to the AIA data. The AIA data was analyzed in its native resolution and orientation (i.e., it was not re-binned to Hi-C resolution or rotated to be aligned to Hi-C.)
- The 4th brightest frame was used as an estimate of the background.

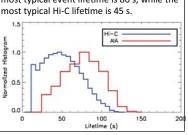


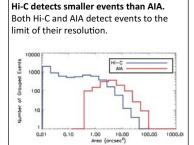


Left panel: AIA image, Right panel: All event locations are shown with the black contours. The location of events are

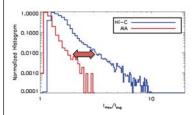
Results:

Hi-C detects shorter events than AIA. The most typical event lifetime is 80 s, while the

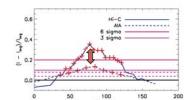




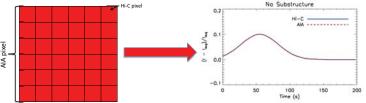
Hi-C events are brighter than events in AIA. The intensity enhancement in Hi-C is ~3 times that in AIA.



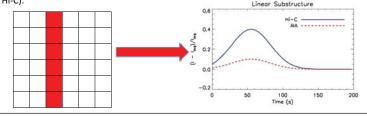
Below is the lightcurve corresponding to point A in the moss. AIA (shown as dashed lines) has a weaker intensity enhancement than Hi-C.



Why is this important? If a the brightening occurs in a structure that is resolved by AIA, then the intensity enhancement in the AIA pixel should be the same as any Hi-C pixel.



If a the brightening occurs in a structure that is not resolved by AIA, then the Hi-C intensity enhancement should be larger than the AIA intensity enhancement. If the substructure is linear, the difference should be the ratio of the resolutions (~4-6 for AIA to Hi-C).



ARW was supported by NASA Living with a Star and TS was supported by the NSF REU program. The authors thank Drs. Leon Golub, Ed DeLuca, Jonathan Cirtain and the Hi-C