THE INVESTIGATION OF CHLORATES AS A POSSIBLE SOURCE OF OXYGEN AND CHLORINE DETECTED BY THE SAMPLE ANALYSIS AT MARS (SAM) INSTRUMENT IN GALE CRATER, MARS. B. Sutter<sup>1,2</sup>, P.D. Archer<sup>1,2</sup>, D.W. Ming<sup>2</sup>, P.B. Niles<sup>2</sup>, J.L. Eigenbrode<sup>3</sup>, H. Franz<sup>3</sup>, D. P. Glavin<sup>3</sup>, A.C. McAdam<sup>3</sup>, P. Mahaffy<sup>3</sup>, J.C. Stern<sup>3</sup>, R. Navarro-Gonzalez<sup>4</sup>, C.P. McKay<sup>5</sup>. <sup>1</sup>Jacobs, Houston,TX 77058, <sup>2</sup>NASA Johnson Space Center, Houston TX 77058, <sup>3</sup>NASA Goddard Space Flight Center, Greenbelt, MD 20771, <sup>4</sup>Universidad Nacional Autónoma de México, México, D.F. 04510, Mexico, <sup>5</sup>NASA Ames Research Center, Moffett Field, CA 94035.

**Introduction:** The Sample Analysis at Mars (SAM) instrument onboard the Curiosity rover detected O<sub>2</sub> and HCl gas releases from the Rocknest (RN) eolian bedform and the John Klein (JK) and Cumberland (CB) drill hole materials in Gale Crater (Fig. 1) [1,2]. Chlorinated hydrocarbons have also been detected by the SAM quadrupole mass spectrometer (QMS) and gas chromatography/mass spectrometer (GCMS) [1,2,3,4]. These detections along with the detection of perchlorate  $(ClO_4)$  by the Mars Phoenix Lander's Wet Chemistry Laboratory (WCL) [5] suggesting perchlorate is a possible candidate for evolved  $O_2$  and chlorine species. Laboratory thermal analysis of perchlorates has yet to provide an unequivocal temperature match to the SAM O<sub>2</sub> and HCl release data [1,2]. Iron mineralogy found in the Rocknest materials when mixed with Ca-perchlorate does cause O2 release temperatures to be closer match to the SAM O<sub>2</sub> release data but more work is required in evaluating the catalytic effects of Fe mineralogy on perchlorate decomposition [6]. Chlorates  $(ClO_3)$  are relevant Mars materials and potential  $O_2$  and Cl sources [7,8]. The objective of this work is to evaluate the thermal decomposition of select chlorate (ClO<sub>3</sub>) salts as possible sources of the  $O_2$  and HCl releases in the Gale Crater materials.

**Materials and Methods:** The Rocknest material examined by SAM consists of unconsolidated sand and dusty material [1]. John Klein and CB are drill hole samples (~6 cm deep) derived from the Sheepbed mudstone and are 3 m and 10cm apart, horizontally and vertically, respectively [2]. The < 150 um size fraction was examined by SAM. Samples were heated (35 °C min<sup>-1</sup>) from 35 to ~840°C in a 30 mb He purge at ~0.8 sccm. Evolved gases were analyzed by the SAM-QMS over the entire temperature range.

A laboratory Setaram Sensys-Evo differential scanning calorimeter (DSC) coupled to a Stanford Research Systems Universal Gas Analyzer at Johnson Space Center (JSC) were configured to operate similarly to the SAM oven/QMS system. Samples are heated from 25 to 730°C under flowing He (3 ml/min) at 30 mb total pressure. Reagent grade (Sigma-Aldrich) NaClO<sub>3</sub>•H<sub>2</sub>O and KClO<sub>3</sub>•H<sub>2</sub>O were used for this work. Mg(ClO<sub>3</sub>)<sub>2</sub>•nH<sub>2</sub>O, and Ca(ClO<sub>3</sub>)<sub>2</sub>•nH<sub>2</sub>O were synthesized by combining dissolved reagent BaClO<sub>3</sub>•H<sub>2</sub>O in water and with equal molar dissolved reagent MgSO<sub>4</sub> and CaSO<sub>4</sub>, respectively. Precipitated BaSO<sub>4</sub> was centrifuged out. The remaining solution was boiled down to ~20 ml and then frozen and subsequently freezedried to allow for Ca- and Mg-chlorate precipitation. Chlorate synthesis was verified by X-ray diffraction.

**Results and Discussion:** The  $O_2$  peak temperatures are ~315 and 385°C for CB and RN, respectively (Fig. 1) suggesting differing  $O_2$  producing species for each material. John Klein has two  $O_2$  peaks at ~225 and 370°C (Fig. 1) which suggests the presence of two  $O_2$ evolving species, consumption of  $O_2$  during organic combustion, or thermal oxidation of a ferrous phase (e.g., magnetite to maghemite transition) [2].

The HCl temperature release characteristics from CB were different than JK and RN. Cumberland was marked by two HCl peaks at ~350 and ~730°C, while RN and JK possessed a gradual HCl release that peaked near ~760°C (Fig. 2). The first HCl phase in CB is coincident with the O<sub>2</sub> release indicating that HCl is sourced from an oxychlorine species like perchlorate or chlorate (Figs. 1,2) [2]. The second CB HCl peak and John Klein and RN peak HCl releases do not coincide with peak O<sub>2</sub> releases.

The  $O_2$  release peak temperatures for Ca-chlorate and Mg-chlorate are the closest match to the RN and second JK  $O_2$  peaks (Fig. 1). The Mg-chlorate  $O_2$  peak nearly coincides with RN and second JK  $O_2$  peaks while the Ca-chlorate  $O_2$  peak are ~25 to 40°C higher than the RN and JK peaks, respectively. The Nachlorate and K-chlorate  $O_2$  peaks are higher than all the  $O_2$  peak temperatures for RN, JK, and CB (Fig. 1).

None of the chlorates evaluated here coincide with the CB and first JK  $O_2$  peaks (Fig. 1). Iron-chlorate has yet to be evaluated and may yield  $O_2$  peak temperatures consistent with CB and the first JK  $O_2$  peaks.

The Mg-chlorate HCl peak is closest to the CB HCl peak 1 but does not yield any HCl that could explain the high temperature (>700°C) HCl peaks in RN, JK, or CB (Fig. 2). The HCl evolution is caused by Mg-chlorate decomposition to MgO which liberates Cl<sub>2</sub> that reacts with water to form HCl as indicated by heated stage X-ray diffraction (data not shown). The Mg-chlorate peak HCl release temperature is near ( within 25°C) to the CB5 peak 1 HCl release. This suggests that Mg-chlorate could be a reasonable candidate for explaining the CB HCl peak 1. The Mg-chlorate O<sub>2</sub> release; however, does not provide an ideal match



Fig. 1. Oxygen release versus temperature from Rocknest-4(RN4), John Klein-4(JK4) and Cumberland-5 (CB5) materials as measured by the SAM-QMS and select chlorate salts analyzed by the JSC-SAM testbed. JK4 is multiplied  $\times$ 5 relative to RN4 and CB5. Arrowed dotted lines indicate Mg- and Ca-chlorate O<sub>2</sub> peak positions relative to O<sub>2</sub> releases in the Gale Crater materials.

to the CB  $O_2$  peak (Fig. 1) suggesting that Mg-chlorate may not be a candidate for CB.

There are two minor HCl releases just above baseline from Ca-chlorate at 70°C and 400°C (Fig. 2). The 400°C release is related to Ca-chlorate decomposition but the low temperature release may be attributed to a temporary HCl baseline rise caused by significant water release associated Ca-chlorate dehydration (data not shown). Much less HCl is produced from Ca-chlorate decomposition because most of the Cl combines with Ca to form CaCl<sub>2</sub>. No HCl releases were detected for Na- and K-chlorate because all Cl is used to form NaCl and KCl [9]. Above 450°C, the Ca-chlorate HCl baseline levels appear to rise slightly which is similar to the observed HCl baseline rise in RN and JK (Fig. 2). The DSC maximum temperature is 730°C and it is unknown if the baseline would have risen and peaked, similar to what was observed for RN and JK. Nevertheless, the Ca-chlorate HCl baseline rise at high temperature coupled with the close O<sub>2</sub> release temperature match to RN and JK suggests that Ca-chlorate may be



Fig. 2. Hydrochloric acid (mass 36) release versus temperature from Rocknest-4 (RN4), John Klein-4 (JK4) and Cumberland-5 (CB5) materials as measured by the SAM-QMS and select chlorate salts analyzed by the JSC-SAM testbed.

a candidate oxychlorine species for RN and JK.

No chlorates matched O2 releases for JK peak 1 and CB. Mg-chlorate O<sub>2</sub> release temperature is consistent with RN and JK but does not have a matching HCl release. Ca-chlorate O2 and HCl release characteristics are both close with RN and JK (O<sub>2</sub> peak 2) which suggests that Ca-chlorate may be a better candidate oxvchlorine species than Mg-chlorate. The thermal decomposition of Fe(II)- and Fe(III)-chlorate will be evaluated next to determine if these chlorate species can better explain the O2 and HCl release characteristics of RN, JK and/or CB. Mixing iron phases (e.g., hematite) with chlorates are known to have catalytic effects that cause chlorate O2 and HCl release temperatures to differ from pure chlorate phases [9]. Future work will mix iron phases (e.g., hematite, magnetite, pyrite, fayalite, amorphous Fe material) known to occur in RN, JK, and CB with chlorates to determine if improved O<sub>2</sub> and HCl release temperature matches with the Gale materials can be obtained.

## **References:**

[1]Leshin et al. (2013) Science, 341 DOI: 10.1126/science.1238937 [2]Ming et al. (2013) Science Express 10.1126/science. 1245267. [3]Glavin et al. (2013) JGR, 118,1955. [4]Glavin et al. (2013) LPS XLV. [5]Hecht et al. (2009) Science, 325, 64. [6] Bruck et al. (2013) LPS XLV. [7] Hanely et al. (2012),GRL, 39, L08201. [8] Quinn et al. (2011) LPS XLIII. [9] Rudlof & Freeman (1970) J.Phys.Chem.74,3317.