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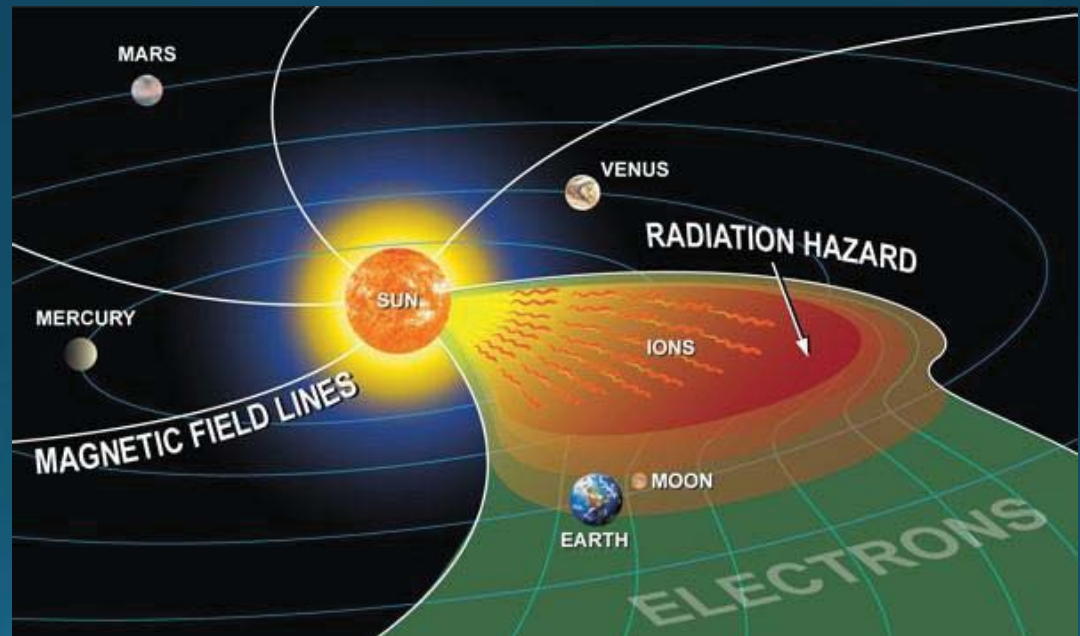
Joseph Minow, NASA/MSFC

AGU, Fall 2014

# Spacecraft Charging in Geostationary Transfer Orbit

# Outline

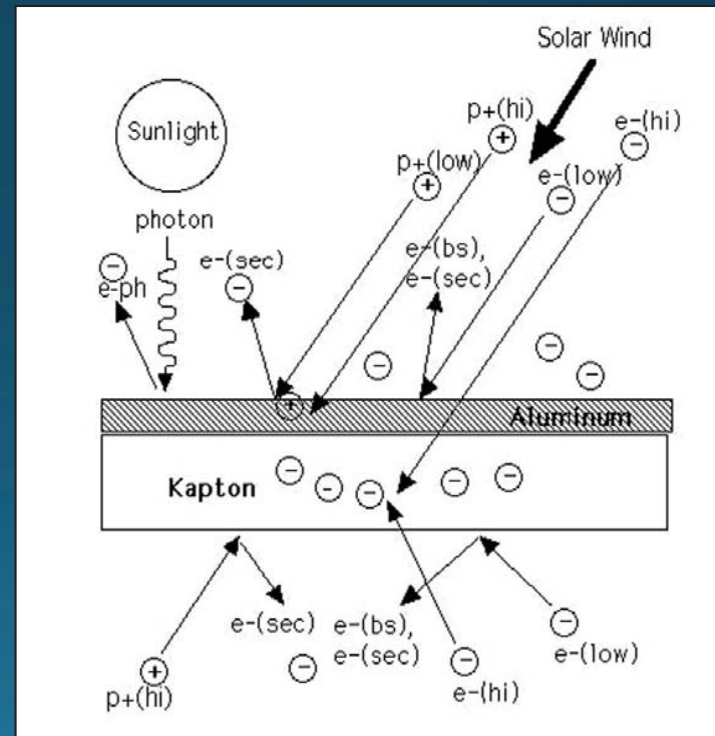
- Background
- Observations
- Model
- Event characteristics
- Future work



# Surface Charging

- Accumulation of charge on the outer surfaces of a spacecraft
- The net charge is due to the sum of the incident currents

$$\frac{dQ}{dt} = C \frac{dV}{dt} = \frac{d\sigma}{dt} A = \sum_k I_k$$



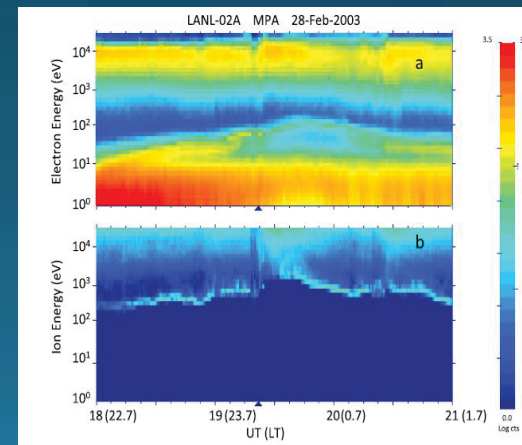
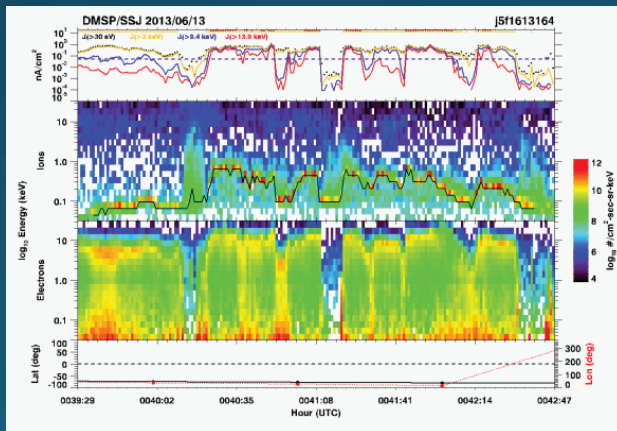
(Garrett and Minow, 2004)

# Ion Line Charging Signature

- “Ion line” is due to the low energy ( $E_0$ ) background ions accelerated to an additional energy ( $q \phi$ ) due to the spacecraft potential

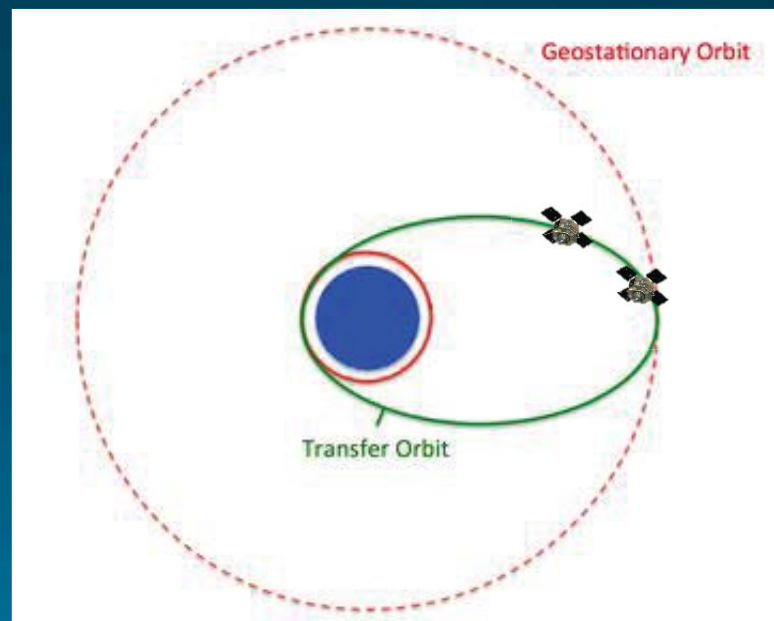
$$E = E_0 + q \phi$$

- Auroral, 1-2 kV, eclipse, low ambient density and high flux for high energy electrons
- GEO, 1-10 kV, midnight through dawn sector



# Van Allen Probes

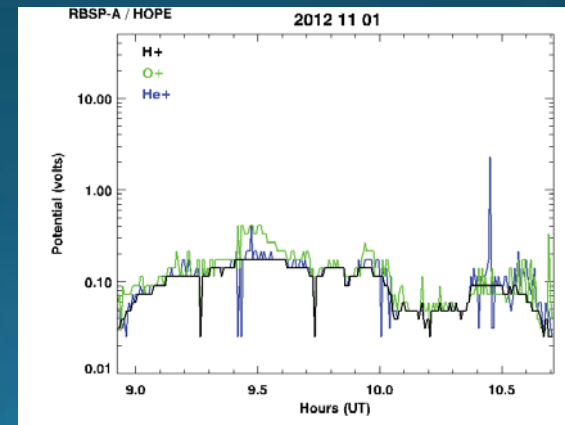
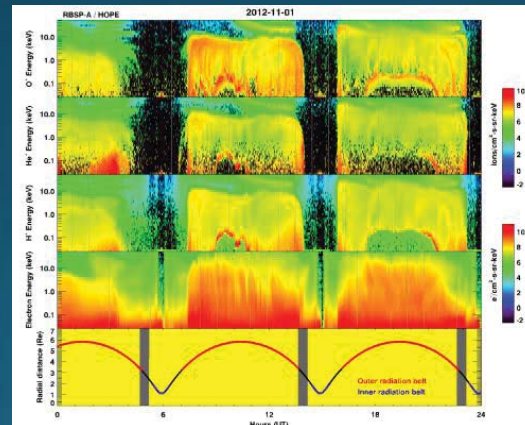
- 700 km x 5.8  $R_e$  orbit, geostationary transit orbit
- Study includes data from the start of mission through December 2013.
- Helium Oxygen Proton Electron (HOPE) plasma spectrometer to identify candidate surface charging events
- Level III moments files



HOPE data courtesy of  
[http://www.rbsp-ect.lanl.gov/rbsp\\_ect.php](http://www.rbsp-ect.lanl.gov/rbsp_ect.php)

# Charging Line Extraction Program

- Read in RBSP HOPE data
- Look for ion charging line in the proton differential energy flux channel
- Program automatically extracts the H<sup>+</sup>, He<sup>+</sup>, O<sup>+</sup> charging lines
  - Plots charging line
  - Prints variables to file including  $t_o$ ,  $t_f$ ,  $\phi$ , ephemeris, moments.



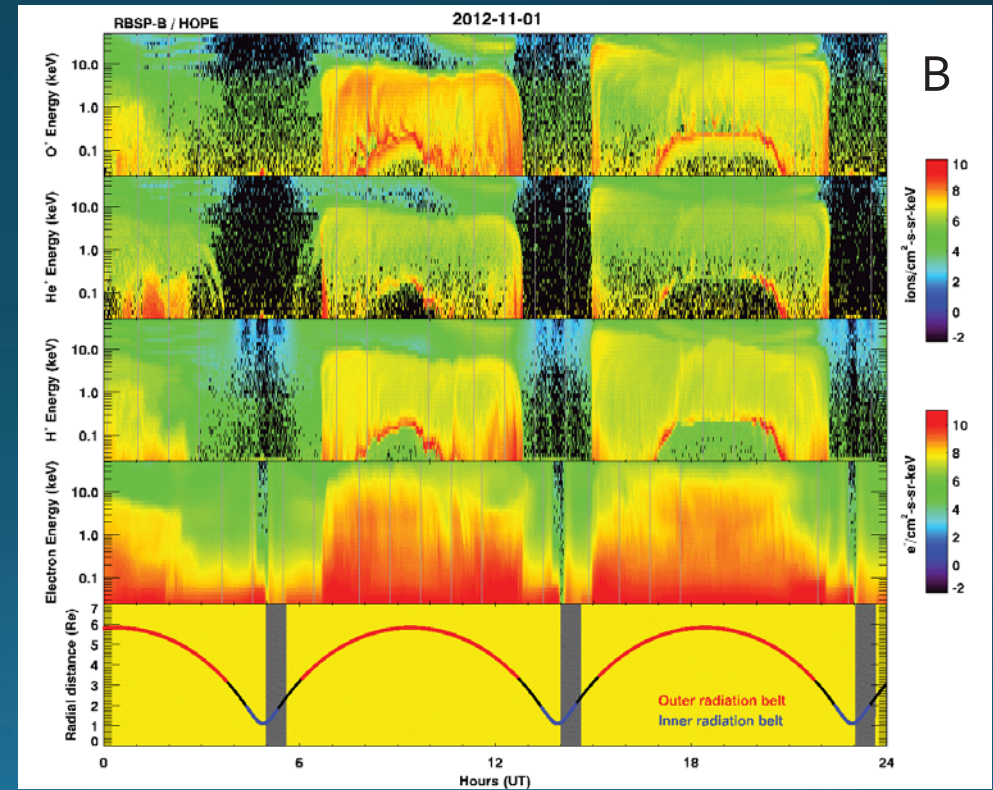
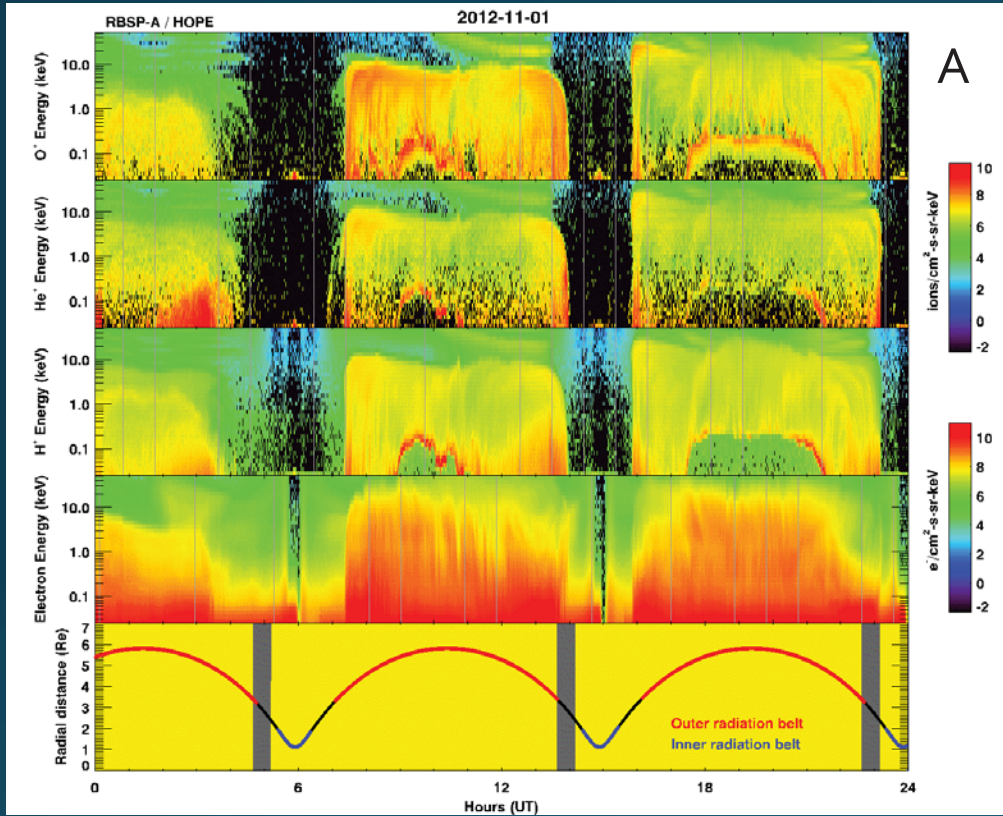
Extracted ion lines.

# Dual Satellite Observations

- By looking at RBSP-A and -B, we can explore the temporal and spatial information of the charging events
- Satellites exhibited charging on the same day in 12 out of 30 days
- When both satellites exhibit charging on the same day, the charging was of similar magnitude

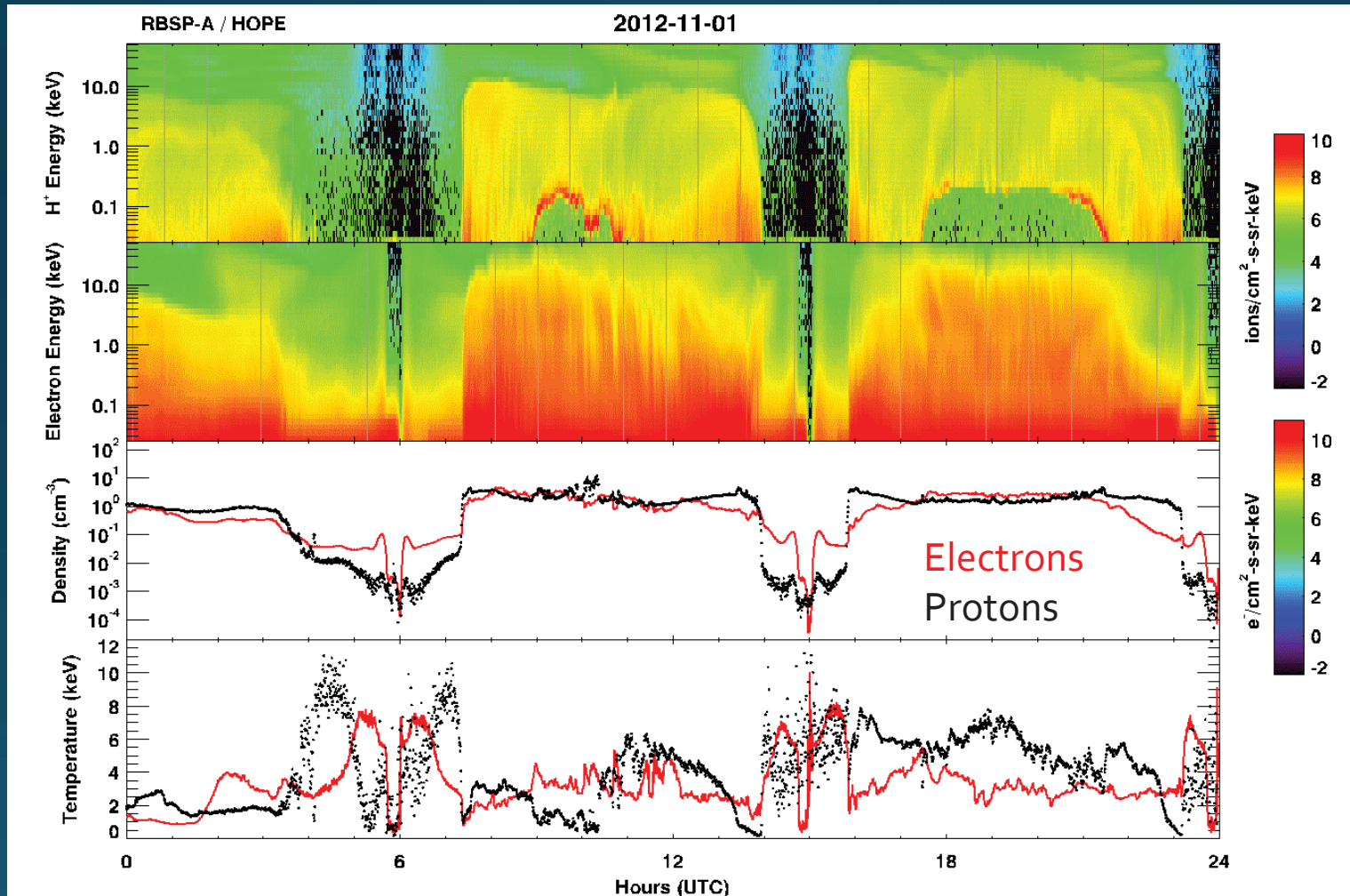


# Example: Nov 1, 2012



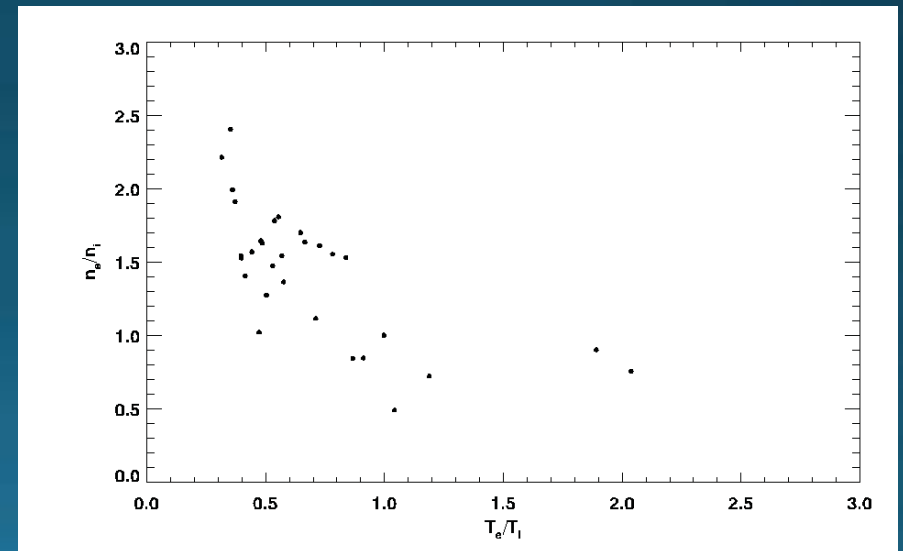
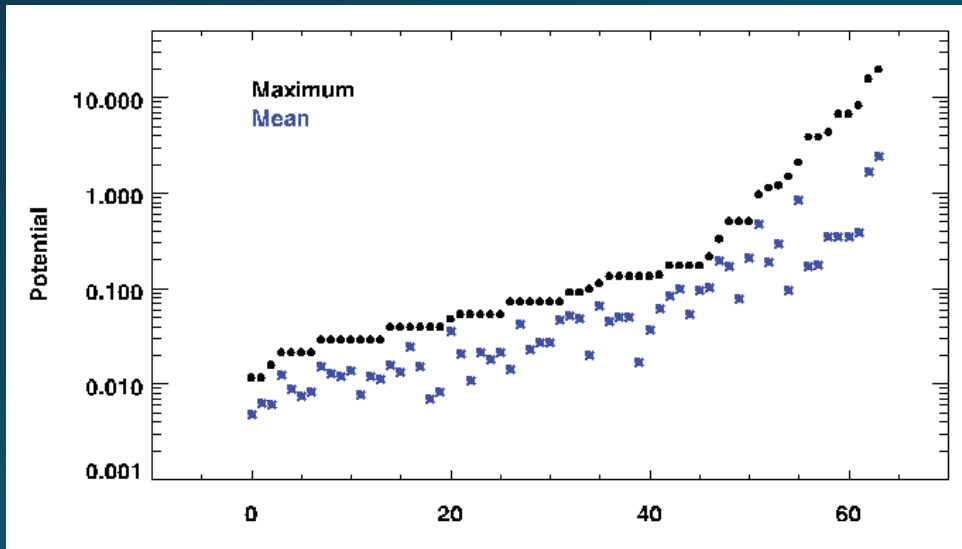


# Moments



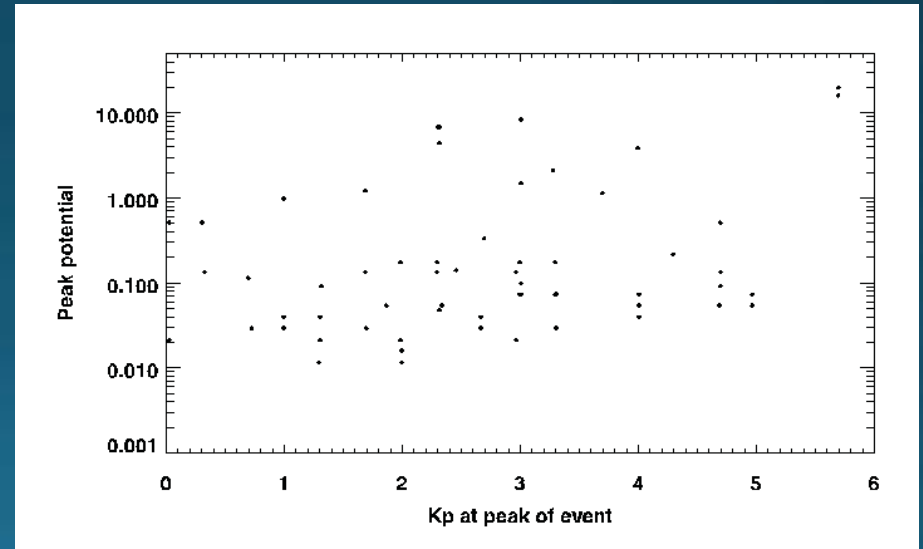
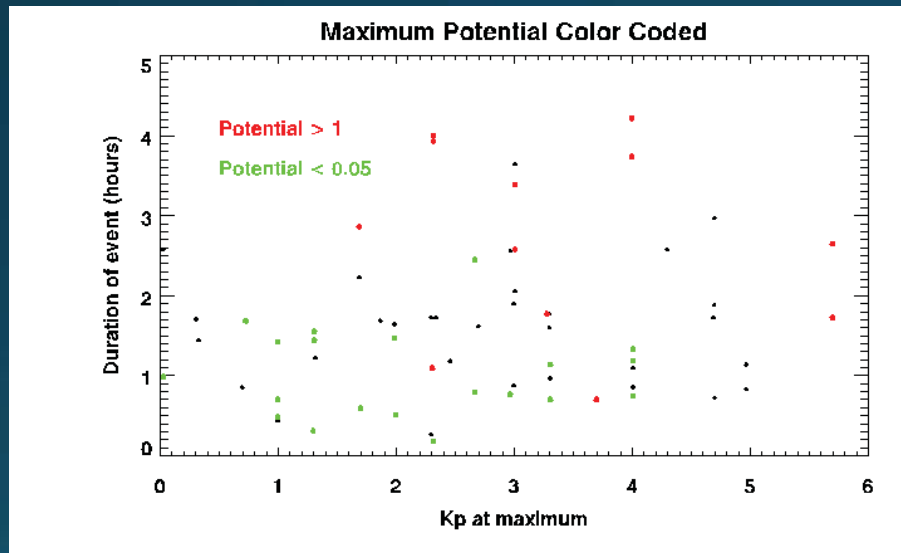
# Charging Levels

- Peak potentials exceed the mean of the charging events by an approximate order of magnitude or less
- In general, more charging events occurred at times when  $n_e > n_i$  and when  $T_i > T_e$



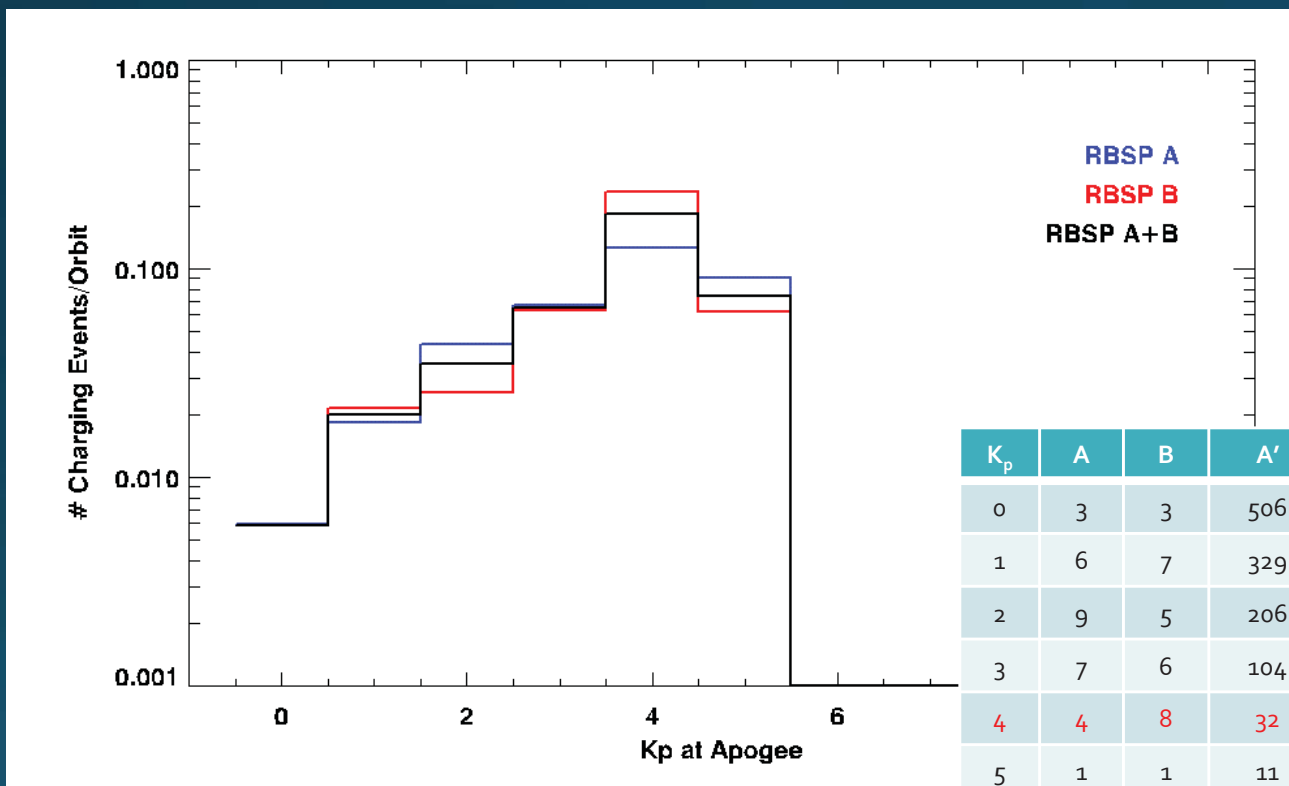
# Relationship to $K_p$

- Weak correlation between the longer the event, higher  $K_p$ , higher charging levels



# Normalized to $K_p$

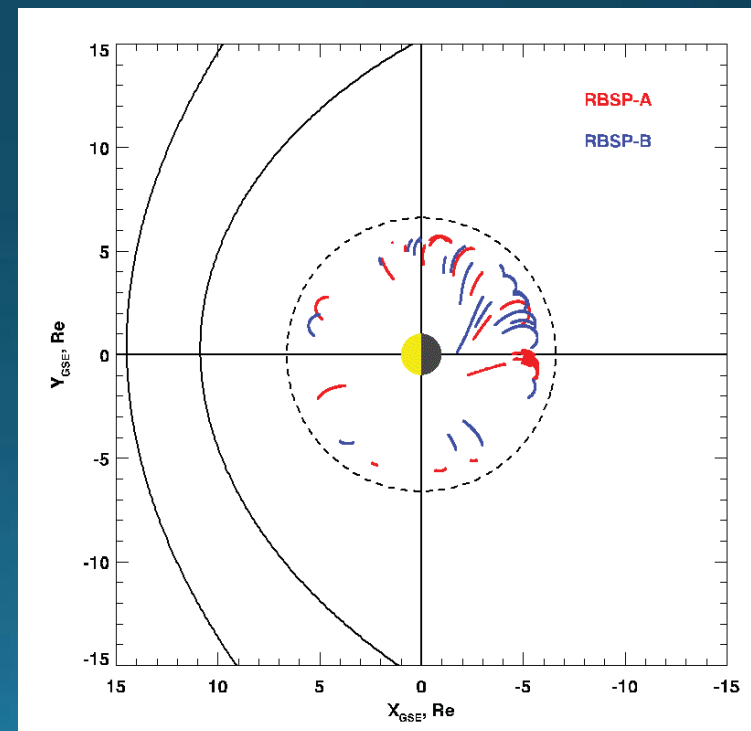
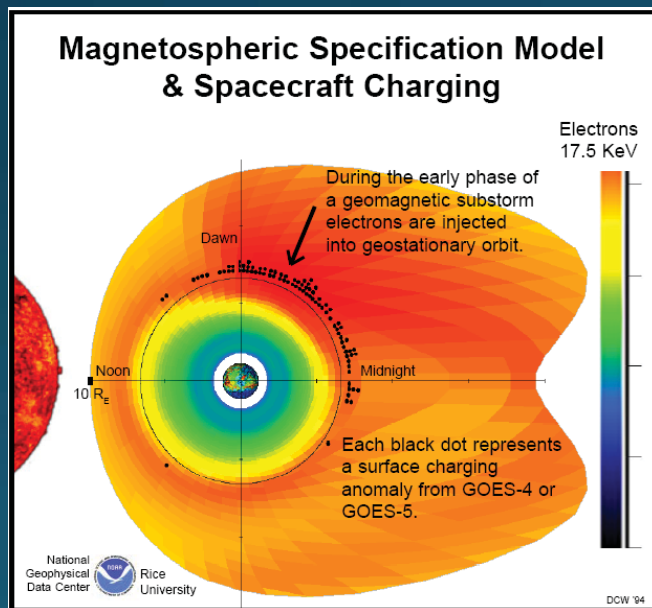
- RBSP-A saw charging 12.5% of the orbits when  $K_p=4$
- 2.5% of the total time (orbits) saw charging



$K_p$	A	B	A'	B'	A/A'	B/B'	A+B/A'+B'
0	3	3	506	512	0.006	0.006	0.006
1	6	7	329	327	0.018	0.021	0.02
2	9	5	206	196	0.044	0.026	0.035
3	7	6	104	95	0.067	0.063	0.065
4	4	8	32	34	0.125	0.235	0.185
5	1	1	11	16	0.091	0.063	0.074
6	0	0	2	2	0	0	0
7	0	0	1	2	0	0	0

# Surface Charging Locations

- GEO charging is more prevalent in the midnight to dawn sector
- GTO, larger number in midnight-dawn sector, but sizable number at other local times



# Summary

- 63 candidate surface charging events in both RBSP-A and B
- All events are in the outer radiation belt
- Charging rates increase with  $K_p$
- Most (55) are in sunlight, however 8 are in eclipse or partial eclipse condition
- Minimum duration charging event ~20 minutes
- Maximum duration charging event ~4 hours
- Maximum potential ~ few kilovolts