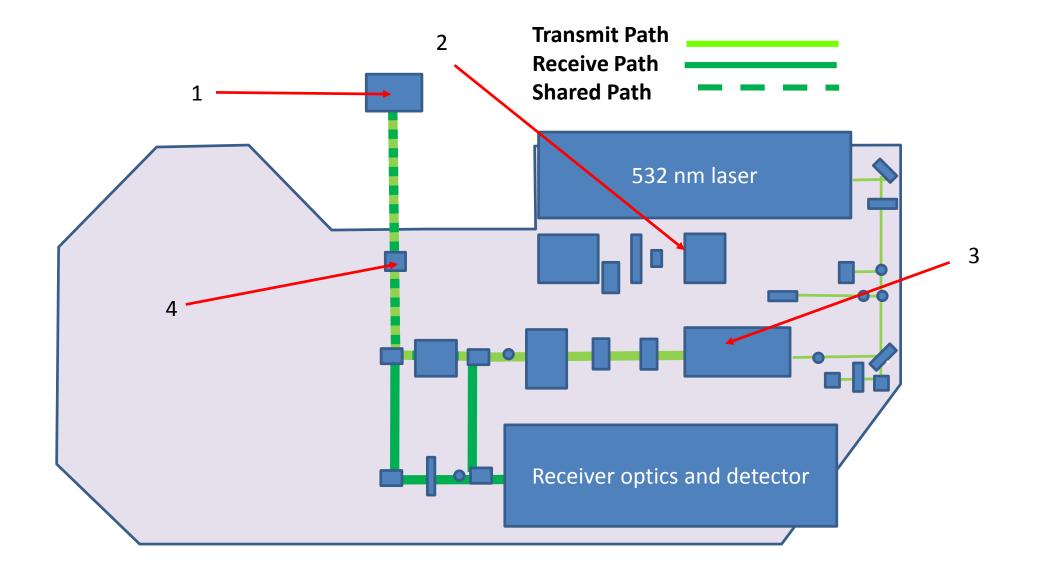
Proposed SLR optical bench required to track debris using ~1550 nm lasers.

M. Shappirio⁽¹⁾, D.B. Coyle⁽¹⁾, J.F. McGarry⁽¹⁾, J. Bufton⁽²⁾, J.W. Cheek⁽³⁾, G. Clarke⁽⁴⁾, S.M. Hull⁽¹⁾, D.R. Skillman⁽¹⁾, P.R. Stysley⁽¹⁾, X. Sun⁽¹⁾, R.P. Young⁽¹⁾, T. Zagwodzki⁽⁵⁾ (1) NASA GSFC, (2) GST, (3) Sigma Space Co., (4) American University, (5) Cybioms Inc.

Abstract: A previous study has indicated that by using ~1550 nm wavelengths a laser ranging system can track debris objects in an "eye safe" manner, while increasing the expected return rate by a factor of ~2/unit area of the telescope[1]. In this presentation we develop the optical bench required to use ~1550nm lasers, and integration with a 532nm system. We will use the optical bench configuration for NGSLR as the baseline, and indicate a possible injection point for the 1550 nm laser. The presentation will include what elements may need to be changed for transmitting the required power on the ~1550nm wavelength, supporting the alignment of the laser to the telescope, and possible concerns for the telescope optics.

Maximum Eye Safe Power

| | 532 nm | 1064 nm | 1550 nm |
|--------------|----------|---------|----------|
| 10 sec exp. | 0.0001 J | 0.001 J | 0.982 J |
| 0 25 sec exp | 0 0001 1 | 0 001 1 | 37 767 1 |



-Based on results using USAF LHAZ6.0 which is calculated using 2014 ANSI standards -Calculated using 1 ns pulse, 50 Hz rep rate, 25 cm beam diameter

-1550 nm remains eye safe at **orders of magnitude** higher power than 1064 or 532 nm

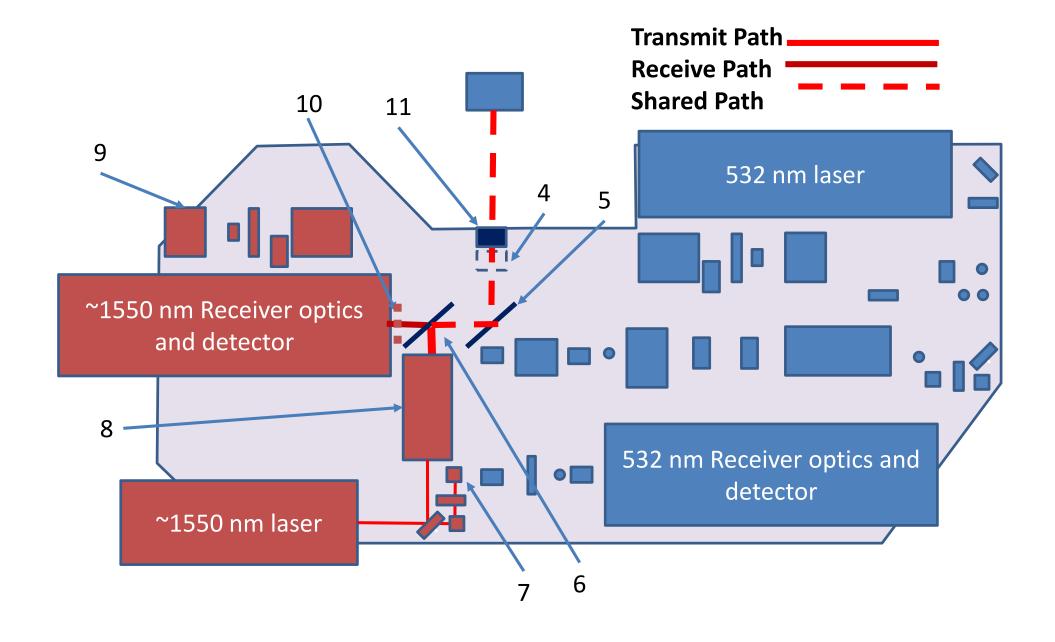


Figure 2: Debris tracking components on NGSLR optical bench. Components in blue 532 nm, red 1550 nm and black are components that need to be added/removed to swap operation from one wavelength to the other. Figure 1: NGSLR optical bench [2]

- 1) Mirror at base of Coude path
- 2) Camera for alignment of laser to telescope
- 3) Removable mirror, installed for alignment, removed for operation
- 4) Variable power Beam Expander

Discussion

- Three optics need to be removable without realignment when replaced,
 - the 1550 nm injection mirror,
 - the optic splitting the 532 nm signal for alignment
 - the optic splitting the 1550 nm signal for alignment
- The 1550 nm beam expander might need to have the ability to be adjusted
- The 1550 nm side is an aperture share setup
- The 1550 nm transmission mirror in this design is mostly a transmission optic with a small mirrored section (aperture sharing)
 Due to shared paths for the transmission and receive the detector should be gated to protect from backscatter light from the transmission
- 4) The one 532 nm optic that needs to be removed for 1550 nm operation the alignment mirror
- 5) 1550 nm injection mirror
- 6) 1550 nm transmission mirror (see discussion section)
- 7) ~1550 nm diode to monitor outgoing laser power
- 8) 1550 nm Beam Expander (x10 fixed)
- 9) Camera for alignment of laser to telescope
- 10) Beam chopper for stray light suppression
- 11) Removable mirror, installed for alignment of the 1550 nm laser, removed for operation.

References:

[1] Tracking orbital debris in a busy airspace environment (3115). M. Shappirio et al., 2014 ILRS conference proceedings

- Detector may also require chopper wheel for additional stray light reduction
- The 1550 nm laser might be large enough that placement on the optical bench is impracticable, could use a fiber to couple the laser to the bench
- Parts or all of the 1550 nm optical bench could be placed above the 532 nm optical bench

Additional considerations

-Telescope optics, particularly the Coude path mirrors, need to be able to handle high power laser pulses in the 1550 nm wavelength

-The optical bench should be designed to be as modular as possible to facilitate installation into different systems