

1 **Topographic and stochastic influences on pāhoehoe lava lobe emplacement**

2

3 Christopher W. Hamilton^{1,2}, Lori S. Glaze¹, Mike R. James³, and Stephen M. Baloga⁴

4 ¹ Planetary Geodynamics Laboratory, NASA Goddard Space Flight Center, Greenbelt,
5 MD, 20771, USA, christopher.hamilton@nasa.gov.

6 ² Department of Astronomy, University of Maryland, College Park, MD, USA.

7 ³ Lancaster Environment Centre, Lancaster University, Lancaster, UK.

8 ⁴ Proxemy Research, Laytonsville, MD, USA

9

10 **Abstract**

11 A detailed understanding of pāhoehoe emplacement is necessary for developing
12 accurate models of flow field development, assessing hazards, and interpreting the
13 significance of lava morphology on Earth and other planetary surfaces. Active pāhoehoe
14 lobes on Kīlauea Volcano, Hawai‘i, were examined on 21–26 February 2006 using
15 oblique time-series stereo-photogrammetry and differential global positioning system
16 (DGPS) measurements. During this time, the local discharge rate for peripheral lava lobes
17 was generally constant at $0.0061 \pm 0.0019 \text{ m}^3/\text{s}$, but the areal coverage rate of the lobes
18 exhibited a periodic increase every 4.13 ± 0.64 minutes. This periodicity is attributed to
19 the time required for the pressure within the liquid lava core to exceed the cooling-
20 induced strength of its margins. The pāhoehoe flow advanced through a series of down-
21 slope and cross-slope breakouts, which began as ~ 0.2 m-thick units (i.e., toes) that
22 coalesced and inflated to become approximately meter-thick lobes. The lobes were
23 thickest above the lowest points of the initial topography and above shallow to reverse-

24 facing slopes, defined relative to the local flow direction. The flow path was typically
25 controlled by high-standing topography, with the zone directly adjacent to the final lobe
26 margin having an average relief that was a few centimeters higher than the lava-
27 inundated region. This suggests that toe-scale topography can, at least temporarily, exert
28 strong controls on pāhoehoe flow paths by impeding the lateral spreading of the lobe.
29 Observed cycles of enhanced areal spreading and inflated lobe morphology are also
30 explored using a model that considers the statistical likelihood of sequential breakouts
31 from active flow margins and the effects of topographic barriers.

32

33 **1. Introduction**

34 Basalt is the most common rock type on the surface of terrestrial bodies
35 throughout the solar system and—by total volume and areal coverage—pāhoehoe flows
36 are the most abundant form of basaltic lava in subaerial and submarine environments on
37 Earth (Self et al., 1994, 1998). Pāhoehoe flow fields are composed of flows, lobes, and
38 toes, which represent a continuum of lava emplacement scales, with toes being the
39 smallest elements. Pāhoehoe flows commonly advance as toes break out along lobe
40 margins (Hon et al., 1994; Crown and Baloga, 1999; Hoblitt et al., 2012). These new toes
41 quickly cool and develop a rheological gradient that includes an inferred three-part
42 structure composed of a brittle outer crust, underlying viscoelastic layer, and inner molten
43 core (Hon et al., 1994). As toes develop sufficient strength to retain incoming lava, they
44 can pressurize, coalesce, and inflate to form lobes that are interconnected with other
45 portions of the flow through internal fluid pathways (Walker, 1991, 2009). The evolution
46 of these pathways can vary considerably (Kauahikaua et al., 1998), but the initial

47 emplacement of pāhoehoe flow fronts can influence the subsequent development of the
48 flow. Consequently, a detailed understanding of processes operating on the scale of toes
49 and lobes can provide important information for modeling aspects of flow field
50 development, assessing hazards, and interpreting the significance of lava flow
51 morphology on Earth and other planetary bodies (Peitersen and Crown, 2000; Byrnes and
52 Crown, 2001).

53 Hon et al. (1994) suggest that “microtopography” (i.e., tens of centimeters relief),
54 herein termed “toe-scale topography”, can strongly affect the emplacement of pāhoehoe
55 lava flows that reach a final inflated thickness of several meters. Previous studies have
56 attempted to determine the relationship between topography and pāhoehoe emplacement
57 (Crown and Balgoa, 1999; Peitersen and Crown, 2000; Byrnes and Crown, 2001), but
58 these studies have been unable to quantify the effects of toe-scale topography at the flow
59 front due to data resolution limitations. Here, oblique digital photogrammetry and
60 Differential Global Positioning System (DGPS) measurements are combined to quantify
61 key physical parameters involved in pāhoehoe flow advance through development of toes
62 and lobes, which constitute the fundamental building blocks of all pāhoehoe flows
63 (Walker, 1991; Self et al., 1996, 1998; Thordarson and Self, 1998). These field
64 observations are also used to inform and constrain a new stochastic model that describes
65 pāhoehoe lobe emplacement in terms of a balance between random and non-random
66 processes.

67

68 **2. Background**

69 Pāhoehoe flow fields exhibit a multitude of branching events at several scales
70 (Crown and Baloga, 1999) and include self-similar morphologies (Bruno et al., 1994) that
71 enable local observations to be used to understand larger aspects of flow fields and their
72 emplacement. However, these patterns are complicated and irregular because of the
73 influence of stochastic processes operating on multiple scales (Kilburn, 1996). This
74 makes it important to identify the scales over which recurring processes and structures
75 combine to form the overall pattern of a pāhoehoe flow field. Understanding how
76 eruption parameters, deterministic factors, and random influences affect the emplacement
77 and modification of pāhoehoe flows, also makes it possible to develop improved
78 probabilistic models to describe likely patterns of flow growth (Glaze and Baloga, 2013).
79 This study examines active pāhoehoe lava lobes on Kīlauea Volcano, Hawai‘i, to
80 characterize the effects of toe-scale topography on flow emplacement and identify key
81 parameters needed to model statistical aspects of the process.

82 Pāhoehoe flows grow through a combination of areal spreading and inflation, with
83 new toes breaking out along lobe margins and quickly cooling by radiation (Hon et al.,
84 1994; Keszthelyi and Denlinger, 1996; Harris et al., 2007a, 2013). Once a thin skin
85 develops on the surface of the new toes, cooling becomes increasingly dominated by
86 conduction (Castruccio et al., 2013). The crust also helps to retain incoming lava, which
87 slows the advance of the new toes as their internal pressure decreases relative to the
88 increasing confining strength of the growing crust (Hon et al., 1994). Given a constant
89 influx of lava, this leads to a reduction in the areal coverage rate of the toes and an
90 increase in the rate of inflation. Inflation will then continue to dominate the emplacement
91 process until the internal pressurization of the lobe reaches a threshold required to rupture

92 the flow margins and initiate a new cycle of enhanced areal spreading (Hoblitt et al.,
93 2012).

94 Within an active pāhoehoe flow, internal connections between numerous lobes
95 can form extensive lava pathways (Kauahikaua et al., 1998). Preferred pathways within
96 these lava networks tend to concentrate in existing topographic lows and may develop
97 into a tube-system as the surrounding flow cools and stagnates. This allows pāhoehoe
98 flows to advance great distances by transporting lava to the flow front through thermally
99 insulated pathways that can reduce cooling rates to only 0.5–1.0°C/km (Helz et al.,
100 1991). As lava is distributed from a master pathway into peripheral lobes and toes, the
101 local discharge rates will progressively decrease with each branch in the system. Local
102 discharge rates may therefore be constant over short periods of time (Hon et al., 1994),
103 but are expected to evolve as the lobes at the former flow front are incorporated into the
104 growing body of the lava transport system. In this study, lobes along the periphery of a
105 low discharge pāhoehoe flow are examined to explore flow front emplacement processes
106 and inflation.

107

108 **3. Study Area**

109 The Prince Kūhiō Kalanianaʻole (PKK) flow was erupted from Puʻu ʻŌʻō on
110 Kīlauea Volcano, Hawaiʻi, from March 2004 to June 2007 (Fig. 1). Active margins along
111 the flow were examined in the field between 21–26 February 2006. During this time, lava
112 from Puʻu ʻŌʻō was primarily transferred through an established tube system that reached
113 the ocean (Koeppen et al., 2013). However, during February 2006, the master tube-
114 system also fed a series of small surface flows located ~7 km south of Puʻu ʻŌʻō. Harris

115 et al. (2007a) referred to this portion of the PKK flow as the “Hook flow”. Field
116 observations and thermal imagery were used to establish that this flow exhibited
117 breakouts of S-type pāhoehoe (Wilmoth and Walker, 1993) from three small flows on 21
118 February. However, by 23 February, the activity had dwindled to the distal margin of just
119 one flow. Waning activity was also evidenced by a decrease in maximum core
120 temperature from 1141°C on 21 February, to 1131°C on 23 February, and a decrease in
121 areal coverage rate from $\sim 0.22 \text{ m}^2/\text{s}$ on February 22 to $\sim 0.03 \text{ m}^2/\text{s}$ on 23 February (Harris
122 et al., 2007a). Additionally, there was a corresponding decrease in discharge rate feeding
123 small surface flows, from $\sim 0.14 \text{ m}^3/\text{s}$ on February 22 to $\sim 0.03 \text{ m}^3/\text{s}$ on 23 February
124 (Harris et al., 2007a). These surface flows stagnated on February 24. Measurements
125 reported in this study were made contemporaneously with the observations of Harris et al.
126 (2007a), who used thermal-image chronometry to establish broad relations among
127 cooling, discharge rate, and areal coverage rates for the whole flow.

128 To place the local discharge rates discussed within this study into a broader
129 context, the Pu‘u ‘Ō‘ō-Kupaianaha eruption began erupting in Kīlauea Volcano’s East
130 Rift Zone (ERZ) in January 3, 1983, and between 1983 and 2002 its time-averaged
131 discharge rate (Harris et al., 2007b) was $\sim 3.81 \text{ m}^3/\text{s}$ dense-rock-equivalent lava (Heliker
132 and Mattox, 2003). However, the Pu‘u ‘Ō‘ō-Kupaianaha eruption has undergone
133 numerous changes throughout its history (Heliker and Mattox, 2003) and in 2003–2007
134 the magma supply rate to the ERZ increased significantly during a mantle-derived surge
135 to Kīlauea Volcano (Poland et al., 2012). During this time, the magma supply to Kīlauea
136 Volcano increased by a factor of two (Poland et al., 2012). By 2005 the supply rate to the
137 ERZ increased to $\sim 5.11 \text{ m}^3/\text{s}$ (Poland et al., 2012) and so at the time of the field campaign

138 in February 2006, effusion rates from Pu‘u ‘Ō‘ō were probably higher than the long-term
139 average. Tilt meter records from the Hawaiian Volcano Observatory’s Pu‘u ‘Ō‘ō cone
140 station also show that there was an oscillatory pattern of inflation and deflation between
141 21–26 February 2006 (Orr, 2011). This pattern began with an inflation event on 21
142 February, which was followed by gradual deflation over the next five days through a
143 series of smaller tilt cycles (Tim Orr, personal communication, 2013). Orr (2011) showed
144 that these tilt cycles were associated with fluctuations in lava discharge through the tube
145 system and may have contributed to the observed breakouts from the master tube.
146 Discharge rates associated with the small surface flows examined within this study were
147 therefore much lower than for the concurrently active PKK tube and the eruption as a
148 whole.

149

150 **4. Methods of Digital Terrain Modeling**

151 To quantify aspects of the pāhoehoe emplacement processes, two complementary
152 methods were employed. First, time-series oblique stereo-imaging and photogrammetry
153 (Robson and James, 2007) were used to determine the changing geometric properties of
154 an active lava lobe, herein referred to as Lobe A. Second, Differential Global Positioning
155 System (DGPS) measurements of topography before and after the emplacement of a
156 second pāhoehoe lobe (Lobe B) were used to assess the relationship between initial
157 topography and final lobe morphology. Lobes A and B were located along the same lava
158 pathway system, with Lobe B located ~10 m directly down-flow of Lobe A.

159 Time series stereo-image pairs of Lobe A were acquired on 23 February using
160 two synchronized tripod-mounted digital Single Lens Reflex cameras (6 mega-pixel

161 Canon EOS 300D), which collected image pairs every minute over a duration of 35
162 minutes (Robson and James, 2007). The cameras were mounted approximately one meter
163 apart and had pre-calibrated, fixed-focus 28 mm lenses. Photogrammetric control was
164 provided using spherical targets (0.25 mm in diameter) with positions determined by
165 DGPS measurements. The analysis was carried out using Vision Measurement System
166 (Robson and Shortis, <http://www.geomsoft.com>) with image matching by the GOTCHA
167 dense matching algorithm (Gruen, 1985; Day and Muller, 1989; Otto and Chau, 1989). A
168 typical stereopair provided topographic point cloud data with positional precisions of 25,
169 65 and 20 mm in easting (x), northing (y), and elevation (z), respectively. To quantify
170 surface changes, these point cloud data were gridded in Surfer using a Kriging algorithm
171 over a 0.1 m/pixel x - y grid. The data were then imported into ArcGIS.

172 For Lobe B, x , y , and z coordinates for topographic surfaces were acquired before
173 (23 February) and after lava emplacement. Post-emplacement surveys were performed on
174 24–26 February; however, no changes were observed after 24 February and so data from
175 these three surveys were combined together. Data were collected at a sampling rate of 0.5
176 seconds using two Ashtech Z-Xtreme DGPS receivers—one operating in kinematic mode
177 and the other as a static base-station. These data were post-processed using Ashtech
178 Solutions 2.60 with differential, L1/L2 (1575.42 MHz/1227.60 MHz) frequency
179 correction for ionospheric delay error, and International Global Navigation Satellite
180 System Service Standard Product 3 precise orbit file corrections. Resulting data points
181 were filtered to remove measurements with more than 0.05 m residual accuracy error. On
182 volcanic terrains, Hamilton et al. (2010) estimate the vertical precision of single traverse
183 kinematic DGPS surveys to be ± 0.08 m, and double traverses being ± 0.06 m, with

184 horizontal precision being better than the vertical measurements. Most gridded cells in
185 this dataset contain multiple DGPS observations and should have a vertical precision of
186 ± 0.06 m. The resulting datasets for the initial and final topography include 14,618 and
187 20,585 records, respectively. To avoid spatial aliasing and to eliminate redundant data,
188 the BLOCKMODE function in Generic Mapping Tools (GMT; Wessel and Smith, 1991)
189 was used to average z within each pixel. Filtered x , y , and z , data were then interpolated
190 using NEARNEIGHBOR in GMT to form a 0.1 m/pixel x - y grid.

191

192 **5. Time Series Observations of Pāhoehoe Lobe Emplacement**

193 Stereogrammetric point clouds and interpolated Digital Terrain Models (DTMs)
194 were used to identify and digitize the active margins of Lobe A at 1-minute intervals (Fig.
195 2). These areal spreading maps show that the lava lobe grew by a series of breakouts that
196 broadly fit into two categories: narrow (0.2–0.6 m-wide) toes that grew preferentially
197 down-slope and broad (1.4–3.5 m-wide) breakouts that formed along the sides of the
198 lobe, nearly perpendicular to the down flow axis (Figs. 3 and 4). Initial toe thicknesses
199 were typically ~ 0.2 m and, during the 35 minute period of observation, they inflated to a
200 maximum thickness of ~ 1.1 m, with most toes inflating to half their final thickness within
201 10 minutes. Within the first minute of emplacement toes generally extended ~ 0.2 m, with
202 a range of 0.05 to 0.55 m. After 28 minutes, the active flow margin began to exit the field
203 of view of the stereo cameras and toe lengths could not be accurately constrained.

204 The mean local discharge rate measured at Lobe A over the first 19 minutes was
205 0.0061 ± 0.0019 m³/s ($N = 19$, Fig. 5a). Note that all uncertainties in this study are
206 reported at 1 standard deviation, σ , unless otherwise noted. Standard deviation describes

207 how much variation there is from the mean, whereas standard error on the mean, σ_μ ,
208 provides a measure of how well the mean is known based on the number of data points,
209 N , used in the calculation. We therefore provide N as well as σ to enable the calculation
210 of $\sigma_\mu = \sigma/\sqrt{N}$. After 35 minutes, the lobe covered a new area of 11.7 m² and achieved a
211 cumulative volume of 10.4 m³.

212 This local discharge rate was remarkably constant during the observation period,
213 except for the interval between 19 and 24 minutes, when it decreased by nearly a factor
214 of four. During this interval, the lobe experienced little inflation and generally lower than
215 average areal spreading rates (Fig. 5a), which may have been due to breakouts occurring
216 outside the field of view. At 24 minutes, the local discharge rate of 0.0061 m³/s resumed,
217 but after 28 minutes, the measured rate of volume change gradually decreased as the
218 active flow front began to exit the field of view of the stereo-cameras. The mean areal
219 coverage rate was 0.0055 ± 0.0038 m²/s ($N = 35$). The areal coverage exhibits a
220 periodicity with peaks ~3 times greater than the mean (Fig. 5b) and a period of $4.13 \pm$
221 0.64 minutes ($N = 8$). During each cycle of areal growth (measured from trough to trough
222 in areal coverage rates shown in Fig. 5b), the lobe covers a new area of 1.59 ± 0.28 m² (N
223 $= 7$).

224 Excess volume refers to the portion of a lava lobe that is not directly associated
225 with its initial areal expansion. In this case, excess volume is defined as the volume
226 change in the field of view per minute minus the change in area during that minute
227 multiplied by the typical initial thickness of the new breakouts. The typical thickness of
228 new breakouts from Lobe A was ~0.2 m, but may vary for other flows—generally scaling
229 with local discharge rate (Self et al., 1998). Excess volume changes for Lobe A appears

230 to be weakly anti-correlated with its areal spreading rate. However, to quantify the degree
231 of anti-correlation, the cross-correlation coefficient was calculated between the areal
232 growth and excess volume. The correlation analysis was restricted to the first 19 minutes
233 of observation when the volume flow rate feeding the lobe was relatively constant. The
234 areal growth rate during this 19 minute period exhibits a slight overall increase over time
235 (slope = $0.013 \text{ m}^2/\text{minute}$ and intercept = $0.206 \text{ m}^2/\text{minute}$), and therefore the trend was
236 removed prior to analysis. After trend removal, the cross correlation coefficient, at lag =
237 0 (the measure of direct anti-correlation) is -0.22. The negative sign of this statistic
238 indicates anti-correlation; however, the magnitude is well within the 90% confidence
239 interval (± 0.34) indicating that the anti-correlation is very weak and not statistically
240 significant at this level. This suggests that there may be a trade-off between areal
241 spreading and inflation given an approximately constant lava supply, but that the
242 relationship may be complicated by other factors. See Section 7.2 for further discussion.

243

244 **6. Lava Thickness, Inflation, and the Effects of Topographic Relief on Pāhoehoe**

245 **Lobe Emplacement**

246 Topographic maps of the terrain before and after the emplacement of Lobe B are
247 shown in Figure 6. Lava covered 111.83 m^2 within the DGPS survey region, with a total
248 lobe volume of 59.45 m^3 . Forward Looking Infrared Radiometer (FLIR) measurements of
249 the flow at 06:00 on 24 February 2006, revealed no anomalously hot pixels within the
250 study region, which indicates that the flow terminated at least 5 hours before (Harris et
251 al., 2007). Based on the time of the last field observations on 23 February, this constrains
252 the emplacement duration to <8 hours. If the mean areal coverage and discharge rates

253 were similar for Lobes A and B, then the emplacement duration of Lobe B was likely
254 between 2.6 and 5.2 hours.

255 Using the before and after DTMs (Figs. 6a and b), topographic profiles have been
256 extracted down-slope (Fig. 7a) and cross-slope (Fig. 7b). The lava entered the study
257 region from the north and the lobe grew to a length of ~18 m and a maximum width of
258 9.1 m. However, the distal part of the flow focused into a 4.3 m-wide topographic
259 constriction before being bifurcated by a 1.0 m-high topographic obstacle into two
260 smaller lobes that were 2.4 m and 4.0 m-wide. The profiles show that inflation of the lava
261 tended to topographically invert the landscape, with the thickest portions of the lobe
262 developing above topographic depressions and shallow to reverse-facing slopes, defined
263 relative to the local flow direction. Thus, apparent barriers to the spreading of the flow
264 are typically much smaller than the maximum height of the lobe and the thicker portions
265 of the lobe appear confined by the thinner peripheral parts of the flow.

266 The thickness distribution of Lobe B (Fig. 6c) above the existing surface was
267 calculated by taking the simple difference between the final and initial topography. This
268 yields a mean lobe thickness of 0.57 ± 0.24 m, with a maximum of 1.21 m. However,
269 simple topographic difference overestimates inflation because lobes are initially
270 emplaced as toes with an initial thickness. Additionally, lava had to fill small closed
271 depressions before thickening above the surrounding topography.

272 To account for the infilling of small topographic depressions, a new initial surface
273 was created by smoothing the initial topography in a process that is analogous to tightly
274 stretching a rubber sheet through the data. The smoothed surface was created using the
275 GMT function SURFACE, which interpolates a grid using continuous curvature splines

276 in tension (Smith and Wessel, 1990). Specifically, SURFACE applies the Laplace
277 transform (L) to the initial topography (z):

$$278 \quad (1-T) L \times (L(z)) + T \times L(z) = 0. \quad [1]$$

279 Varying the tension factor T between 1 (i.e., the harmonic surface solution) and 0
280 (i.e., the minimum curvature solution) enables this data interpolation method to smooth
281 the initial surface to different degrees. For instance, a grid produced with high tension (T
282 $\rightarrow 1$) will suppress local maxima and minima by fitting an increasingly taut surface
283 between the data constraints, whereas applying a looser tension factor ($T \rightarrow 0$) will
284 produce a smooth polynomial function analogous to a flexed elastic plate that is anchored
285 by the data (Smith and Wessel, 1990).

286 To constrain the inflated thickness of Lobe B, a smoothed local datum was fit to
287 the initial topography using a T -value of 1. This removes high frequency oscillations in
288 the topography, while preserving the overall shape of the terrain. Subtracting this surface
289 from the final topography isolates the thickness of lava above the local datum and
290 provides a better proxy for inflation than the simple difference map because it reduces the
291 thickness contributions due to small basin infilling. Using this method, the mean
292 thickness due to inflation was 0.34 ± 0.21 m ($N = 9614$), with a maximum of 1.00 m (Fig.
293 6d).

294 To isolate local relief within the initial topography, a second interpolated surface
295 was created using Eq. (1) with $T = 0.25$. A loose tension factor of $T = 0.25$ is
296 recommended for low-tension interpolations to suppress extraneous oscillations (i.e.,
297 minima and maxima) that can otherwise be generated within the $T = 0$ (i.e., end-member
298 minimum curvature) solution (Wessel and Smith, <http://gmt.soest.hawaii.edu/>). The new

299 surface preserves more variations in the initial topography than the harmonic surface
300 solution ($T = 1$) and was subtracted from the original initial topography to isolate high
301 frequency changes in elevation (i.e., relief) above and below a new local datum centered
302 on 0 m.

303 In the lava inundated region, the distribution of initial relief is unimodal with a
304 mean of -0.02 ± 0.07 m ($N = 9614$ pixels; Fig. 8). For comparison, a 1-pixel buffer region
305 was also defined around the outer extent of the lobe and, in this buffer region, the
306 frequency distribution of relief had a mean of 0.00 ± 0.07 m ($N = 499$ pixels). However,
307 relative to the lava inundated region, there are more pixels with a relief of 0.03 to 0.09 m
308 and fewer pixels with a relief of -0.09 to -0.03 m (Fig. 8). This implies that there are
309 systematically more high relief pixels bounding the pāhoehoe lava lobe than in the region
310 that was inundated. Consequently, positive relief on the order of only a few centimeters
311 may exert an important influence on bounding the extent of low-effusion-rate pāhoehoe
312 flows.

313 Toe-scale topography (i.e., on the order of centimeters in this case) appears to
314 influence the areal spreading of pāhoehoe lobes by blocking peripheral toes, which may
315 then stagnate and help confine the interior portions of the flow. If the interior of the flow
316 continues to be supplied by lava it may then inflate above the height of the surrounding
317 barriers, thereby topographically inverting the landscape. Barriers to flow may therefore
318 be much less than the final inflated thickness of a lobe. Based on observations of Lobe B,
319 relief that is at least 15% of a toe's initial thickness may be significant in terms of
320 affecting the path of low-discharge pāhoehoe lava lobes, but given sample-size-
321 dependent restrictions on the optimum bin-size for the data (Stuges, 1926; Doane, 1976),

322 influences of even smaller variations in topography cannot be precluded. Obstacle
323 heights affecting the path of higher local discharge rate flows may similarly scale with
324 initial toe height.

325

326 **7. Stochastic Modeling of Pāhoehoe Lobe Emplacement**

327 **7.1. Overview of the Model**

328 Random effects strongly influence processes of pāhoehoe flow growth and pose a
329 fundamental difficulty in developing models for their emplacement (Hon et al. 1994;
330 Thordarson and Self, 1998; Keszthelyi et al., 1999; Crown and Baloga, 1999). Glaze and
331 Baloga (2013) have recently developed a new simulation approach for modeling
332 pāhoehoe emplacement that builds on the ideas presented in Baloga and Glaze (2003).
333 This model is based on (1) conservation of lava volume and (2) prescribed stochastic
334 rules for lava movements within a pāhoehoe lobe. This approach differs significantly
335 from other models that treat lava emplacement as a gravity-driven viscous fluid flow on
336 an inclined plane (e.g., Danes, 1972; Hulme, 1974; Baloga and Pieri, 1986; Baloga, 1987;
337 Crisp and Baloga, 1990; Harris and Rowland, 2001; Rowland et al., 2004; Baloga and
338 Glaze, 2008). The Glaze and Baloga (2013) model simulates the 3-Dimensional shape for
339 pāhoehoe lobes as they evolve in time, subject to a wide variety of ambient and internal
340 conditions and processes. Here, the new model is used to explore the effects of inflation
341 on pāhoehoe lobes analogous to Lobes A and B.

342 The fundamental building block of the Glaze and Baloga (2013) model is the lava
343 “parcel”. A “parcel” of lava is defined here as a volumetric unit sampled from a
344 probability distribution of lava volumes that typically construct a pāhoehoe lobe. Thus

345 the smallest toe observed on the surface or along the margins of a pāhoehoe lobe
346 represents a lava parcel volume from the small end of the distribution. However, larger
347 toes may either represent a larger parcel volume, or may be comprised of multiple
348 parcels.

349 A parcel is observable as a “toe” when it is permanently affixed at the surface or
350 margin of a lobe. In the Glaze and Baloga (2013) model, lava parcels originate from a
351 source region and are transported through fluid internal pathways. For the purposes of
352 discussion here, the model assumes a constant parcel size for all parcels. Crown and
353 Baloga (1999) measured dimensions of hundreds of toes and their mean thickness is very
354 well constrained at ~ 0.20 cm, (0.19 ± 0.08 m, $N = 448$). Based on the geometric mean toe
355 sizes measured by Crown and Baloga (1999), the parcel volume is assumed to be 0.09 m³
356 (Glaze and Baloga, 2013). Assuming a square x - y grid and a typical parcel thickness of
357 0.2 m—comparable to the observed initial thickness of toes in our study area—the
358 corresponding cell spacing is 0.67 m, with a parcel area of 0.45 m².

359 The Glaze and Baloga (2013) random simulation model begins with an initial
360 source region (a single point, linear source, etc.), a constant supply of lava (i.e., one
361 parcel per time step), and probabilistic rules governing where each parcel will be
362 transferred and allocated. For example, in the most basic scenario of complete
363 randomness, two random choices are made for each breakout. First, a parcel is randomly
364 selected to be the source for the next parcel transfer. Second, a random selection is made
365 to determine the orientation of the new parcel transfer relative to its source (i.e., North,
366 South, East, or West). In the purely random case, every face of every parcel has an equal
367 probability of being the source for the next parcel transfer. As various factors that

368 influence emplacement are examined (e.g., barriers, slope, etc.), different probabilistic
369 rules are applied that determine where and when each parcel will be transferred.

370

371 **7.2. Model Simulations of Areal Growth**

372 Figure 9a shows a typical example of a completely random simulation involving
373 500 parcels. In this simulation, each transfer location and direction is chosen at random.
374 Figure 10a (solid line) shows the corresponding cumulative area as a function of time
375 step. However, Glaze and Baloga (2013) showed that topographic profiles of pahoehoe
376 lobes are more consistent with simulations that include sequential breakouts at the
377 margin. In this case, the momentum of new parcels breaking out from a lobe margin
378 results in an increased probability that the next parcel will be transferred from the most
379 recent parcel in the same direction. Figure 9b shows the influence of allowing sequential
380 breakouts at the lobe margin on lobe morphology. Specifically, every time a new parcel is
381 transferred to the margin (increasing the lobe area) probabilistic rules are used to
382 determine whether 0, 1, or 2 extra parcels will be added at that location. For the
383 simulation shown in Figure 9b, the probability of 0 additional parcels is $P(0) = 0.25$, of 1
384 additional parcel being added is $P(1) = 0.25$, and of 2 additional parcels is $P(2) = 0.5$. As
385 expected, the cumulative area of the simulation that includes sequential breakouts at the
386 margin rapidly diverges from the purely random simulation (Fig. 10a), exhibiting a
387 steeper gradient and overall greater cumulative area. Although a general trend in the areal
388 growth (m^2 per time step) is evident in both simulations shown in Figure 10a, the curves
389 are irregular with periods where the area increases rapidly and periods where the area
390 does not change at all. Also as expected, the rate of overall areal growth tapers off over

391 time as the number of available lava transfer locations along the margins decreases
392 relative to the total area of the lobe.

393 The Glaze and Baloga (2013) simulation approach explicitly requires that lava
394 parcels will either expand the area of the flow or contribute to inflation if they are
395 allocated to the flow interior. This naturally results in a trade-off between areal growth
396 and interior lobe inflation similar to the pattern observed during the emplacement of Lobe
397 A. The natural periodicity of the areal growth rate in the simulated lobes is illustrated by
398 the solid line in Figure 10b. Also shown in Figure 10b is the excess volume, which was
399 defined in Section 5 as the volume in addition to what is directly associated with the
400 change in lobe area because all breakouts will be emplaced with some initial thickness.
401 For the simulated data, the definition of excess volume results in a periodic behavior that
402 exactly mirrors the areal growth (i.e., perfect anti-correlation). This suggests that the
403 weak anti-correlation shown in Figure 5b may be relevant, but also indicates that other
404 processes are present in the field that have not yet been accounted for by the model. Also
405 in the field example, there was a period of anomalously low average lava flux between
406 20–23 minutes, which coincided with a period of negative excess volume change. This
407 indicates that areal spreading can also be accommodated by the depletion of lava stored
408 within the parent lobe, rather than always being related to a simple balance between
409 pressurization and confining strength under conditions of constant lava supply. Future
410 studies would therefore benefit from longer time-series observations of multiple active
411 lava lobes to further explore the statistical significance of the relationship between areal
412 spreading and inflation.

413

414 7.3. Comparisons Between Model Simulations and Observations

415 Given the measured volume of Lobe A and typical parcel volume of 0.09 m^3 ,
416 Lobe A would include ~ 117 parcels. The simulation of a lobe composed of 117 parcels
417 begins with a 1×7 cell linear source region trending East-West (depicted as the gray
418 shaded region in Figure 11a crossing through the mid-point of the array). This source
419 region, with one parcel in each of the seven cells, is approximately equivalent to the
420 initial length of Lobe A (i.e., ~ 4.69 m; see Fig. 4a). The model also includes two parallel
421 barriers located at ± 3 cells along the North-South axis. The barriers are considered
422 transitional zones in which the advancing lava enters, but becomes blocked, thereby
423 reflecting the lava parcel back toward the interior of the flow. The distance between these
424 barriers is 4.69 m (including the barrier cells as well), which corresponds to the
425 topographically-bounded maximum width of Lobe A. The probabilities of sequentially
426 adding 0, 1, or 2 extra parcels from an existing flow margin are $P(0) = P(1) = 0.25$, and
427 $P(2) = 0.5$.

428 One cannot expect any specific realization of the stochastic model to exactly
429 reproduce the morphology of an observed lava lobe, but characteristics of the simulated
430 lobes are broadly consistent with the morphology of Lobe A. Figure 11b shows examples
431 of topographic profiles taken perpendicular to the flow direction through the simulated
432 lobes averaged over 10, 20, 30, and 60 realizations. These examples demonstrate that
433 there are stochastic variations between simulations, but also persistent morphological
434 characteristics. For example, the simulated lobes are thickest in the middle (i.e., along the
435 axis of the linear source region) with maximum heights that are consistent with those
436 observed for Lobe A. The models also exhibit the effects of lava confinement along lobe

437 margins due to lava parcels being reflected back toward the interior of the flow. Plan
438 views of the simulated lobes (e.g., Figs. 9a and 9b) also show irregular margins and
439 thickness variations that are typical of inflated pāhoehoe lobes.

440

441 **8. Discussion**

442 **8.1. Relationships between Pāhoehoe Breakouts and Inflation**

443 Pāhoehoe breakouts will be emplaced with some initial thickness and then grow
444 through a process of inflation. This inflation process includes two dominant components:
445 thickening of the molten core and thickening of the crust (Hon et al., 1994), with the crust
446 consisting of both a brittle outer layer and underlying viscoelastic layer. After an initial
447 chilled margin forms around a new toe, it will retain incoming lava and inflate as its
448 molten core thickens to equalize the pressure with other connected portions of the fluid
449 lava pathway. During the early stages of emplacement, the rate of core thickening will
450 greatly exceed the rate of crustal thickening (Hon et al., 1994), but as the fluid interior
451 achieves its equilibrium thickness, inflation will be increasingly dominated by the process
452 of crustal growth (Cashman and Kauahikaua, 1997).

453 Two conditions must be met for the upper crust to continue to accrete new
454 material and contribute to inflation. First, there must be a continuous influx of new lava
455 through the molten core and, second, there must be physical coupling of the molten core
456 to the upper crust (Kauahikaua et al., 1998). If the flow of lava through the core stagnates
457 then the lava will cool *in situ* and will not increase the flow's total thickness.

458 Additionally, if the molten core partially drains and decouples from the upper crust then
459 new material will not be added to the base of the surface layer and the crust may even

460 subside and “deflate” (Kauahikaua et al., 1998). Cooling histories inferred from the
461 thicknesses of the upper brittle and viscoelastic layers (Hon et al., 1994) therefore
462 provide an estimate of the time that a lava pathway flowed at full capacity, but not
463 necessarily the total duration of emplacement if the core of the flow and the upper crust
464 were decoupled. In some cases, the molten core of a flow can also undergo cycles of
465 pressurization related to changes in the local discharge rate (Orr, 2011). This can lead to
466 variations in molten core pressure that induce cycles of surface disruption and
467 subsidence, thereby producing a range of surface textures that include slabbly pāhoehoe,
468 rubbly pāhoehoe, and shatter rings.

469 Initial cooling can strengthen the margins of a lava flow (Castruccio et al., 2013),
470 but the process is complicated by the development of numerous cooling fractures. These
471 fractures will mechanically weaken the outer brittle crust and make the underlying
472 viscoelastic layer more important in terms of governing the overall confining strength of
473 the flow. The frequency of breakouts will therefore depend on the balance between the
474 combined strength of the brittle and viscoelastic layers relative to the pressure within the
475 molten core (Hoblitt et al., 2012). If the internal pressure exceeds the confining strength
476 of the flow’s outer layers, then the periphery of the flow will rupture and generate new
477 breakouts that promote increased areal growth. In contrast, if the exterior of the flow does
478 not rupture, it will pressurize and inflate through gradual deformation of the viscoelastic
479 layer and opening of fractures in the overlying brittle crust.

480 In this study, peripheral lobes along the margins of the PKK flow on the flanks of
481 Kīleaua Volcano, exhibited a generally constant local discharge rate (0.0061 ± 0.0019
482 m^2/s). However, approximately every four minutes there was a significant increase in the

483 areal spreading rate. The observed periodicity in areal coverage may relate to the
484 changing balance between the confining strength of the lava as it cools and strengthens,
485 and the internal pressure within the core of the flow. This balance would give rise to an
486 increased frequency of new breakouts when the core pressure exceeds the confining
487 strength of the exterior. Enhanced areal spreading rates would then persist until the core
488 pressure decreases below a critical threshold required to continuously rupture the active
489 flow margin. As the rate of areal spreading decreases, continued influx of lava into the
490 lobe would begin to re-pressurize its fluid core and perpetuate a new cycle. This process,
491 operating on the scale of an individual lobe, is similar to what Hoblitt et al. (2012) have
492 observed on the scale of an entire pāhoehoe flow, which demonstrates the self-similarity
493 of the inflation mechanism over a range of scales.

494

495 **8.2. Static and Dynamic Pressure Considerations**

496 During early stages of pāhoehoe emplacement, the inflation process will be
497 dominated by molten core thickening as the liquid interior of the flow moves toward an
498 equilibrium thickness. However, there are multiple factors that can contribute to this
499 equilibrium. Calculations generally assume a Bingham liquid rheology, with lava
500 descending a slope as an unconfined laminar flow under the influence of gravity (Hulme,
501 1974; Dragoni et al., 1986; Castruccio et al., 2013). Within this scenario, the critical
502 depth h_s required to induce down-slope flow will be

$$503 \quad h_s = \sigma_0 / \rho g \sin(\beta), \quad [2]$$

504 where σ_0 is the flow's yield stress, ρ the flow's density, g is gravitational acceleration,
505 and β is the underlying slope (see Eq. [14] in Hulme, 1974). However, such calculations

506 of a lava flow's equilibrium thickness will be complicated by cooling-induced changes in
507 rheology and other factors that influence its internal pressure once the molten core is
508 confined by a coherent crust. Several studies (e.g., Rossi and Gudmundsson, 1996;
509 Anderson et al., 1999, 2012) have considered the "magma-static" pressure (herein referred
510 to as "hydrostatic" pressure) generated by fluid elevation changes along a hydrostatically
511 connected pathway, as well as the pressure required to drive lava through pipe-like
512 pathways and lift the crust. These models generally assume that the pathways (i.e., lava
513 tubes) are cylindrical pipes with constant cross-sectional area, but the geometry of lava
514 pathways can vary widely within a pāhoehoe flow (e.g., Guest et al., 1984; Kauahikaua et
515 al., 1998; Self et al., 1998; Calvari and Pinkerton, 1999). This raises the possibility that
516 changes in pathway geometry can introduce additional changes in lava core pressure due
517 to dynamic processes.

518 These dynamic effects may be understood within the context of Bernoulli's
519 principle, which describes the inverse relationship between fluid flow velocity and fluid
520 pressure within incompressible laminar flows through a pipe (Resnick and Halliday,
521 1977; Batchelor, 1998). For lava moving within a system of combined pathways, changes
522 in fluid core pressure could affect inflation rates by changing the magnitude of the
523 stresses applied to the overlying crust. This may lead to increased inflation rates where
524 the flow velocity decreases and the core pressure increases. Conversely, inflation rates
525 would be lower where the flow velocity increases and the core pressure decreases. For
526 otherwise equivalent lava flows through a pipe-like pathway, the most important controls
527 on fluid velocity will be the pipe's cross-sectional area and underlying slope. In general,
528 lava velocities will be relatively low where the flow pathway is broad and its cross-

529 sectional area is large and/or where it descends a shallow slope or encounters a reverse-
530 facing slope. Alternatively, flow velocities will be higher when lava enters a narrow
531 topographic constriction and/or descends a steeper slope. These scenarios are broadly
532 consistent with the observed relationship between the initial topography and final
533 morphology for Lobe B. In this case, maximum inflation was observed where the flow
534 was the broadest and above shallow to reverse-facing slopes, whereas the least inflation
535 was observed where the flow passed through narrow constrictions and descended steep
536 slopes (Figs. 6 and 7). These observations support the viability of a conservation of
537 energy approach to modeling the spatial variability of inflation and its relationship to
538 emplacement processes. They also highlight the importance of including both fluid static
539 and dynamic components in understanding core pressure and thickening during the early
540 stages of pāhoehoe inflation. Early stage inflation processes related to core pressurization
541 and thickening may also help to establish preferred pathways through the thickest parts of
542 the flow, which would continue to preferentially inflate through crustal accretion if the
543 pathways remain full. Continued crustal accretion and uplift must also be coupled with
544 overpressurization of the molten core, which Cashman and Kauahikaua (1997) inferred
545 from their observations of pāhoehoe crustal vesicularity profiles that exhibit decreased
546 vesicle number density with depth.

547

548 **9. Conclusions**

549 The confining strength of a lava lobe depends strongly on the age distribution of
550 its surfaces, with the youngest surfaces (i.e., sites of the most recent breakouts) being the
551 weakest. Therefore, when new breakouts occur, growth tends to concentrate at these

552 localities. Otherwise, the next most likely location for a breakout to occur is from an
553 older fractured surface overlying a still molten core. This process is treated with a
554 stochastic model by introducing correlation with a probability distribution describing the
555 preferential growth of the flow in the direction of new breakouts through sequential
556 emplacement of lava parcels (i.e., toes) from that locality. Including correlated growth
557 into the model also recreates observed periods of enhanced areal spreading and lobe
558 morphologies.

559 Time-series data for Lobe A shows that periods of enhanced areal spreading
560 roughly alternate with periods of increased excess volume change. For an approximately
561 constant local discharge rate, this suggests that the volume of lava not utilized during the
562 process of areal spreading can contribute to thickening through inflation. Trade-offs
563 between areal spreading and inflation should be detectable in thermal infrared (e.g.,
564 FLIR) imagery, provided that a single lobe is isolated in the field of view. For instance,
565 when the flow enters a period of above average areal spreading, newly exposed lava
566 surfaces will radiate more thermal energy and will contribute to an increased frequency
567 distribution of high temperature pixels within the field of view. As the flow shifts toward
568 an inflation-dominated mode, older lava surfaces will cool dominantly by conduction
569 rather than radiation (Hon et al., 1994), and this would shift the frequency distribution of
570 temperatures toward lower values. Temperature distributions in FLIR data may therefore
571 be used to identify patterns of pāhoehoe flow emplacement, such as periodicities in areal
572 spreading rate and the trade-off with inflation. Statistical models of flow emplacement
573 may be improved by coupling them with a cooling model to estimate the temperature
574 distribution of the lava based on its emplacement age.

575 New toes are highly sensitive to subtle topographic irregularities and for low
576 discharge rate flows may be deflected or stopped by obstacles measuring only a few
577 centimeters in height. This process was observed during the emplacement of Lobe B and
578 demonstrates that the final morphology of a lobe may inherit boundaries influenced by
579 variations in the initial toe-scale topography that are orders of magnitude smaller than the
580 flow's final thickness. Even though toes may be locally deflected or impeded by toe-scale
581 topography, inflating pāhoehoe lobes may overcome local barriers through a combination
582 of stochastic breakouts in down-flow and cross-flow directions that allow them to find
583 and follow the maximum regional slope. These behaviors are described in the model by
584 considering the sequential emplacement of new lava parcels (i.e., toes) from existing flow
585 margins, with boundaries that reflect parcels back towards the interior of the flow.

586 Combining measurements of initial topography with time-series observations of
587 active pāhoehoe lobes provides new insights into the growth of pāhoehoe flows. These
588 observational constraints also help to improve statistical models of lava emplacement
589 processes and better understand the relationship between areal expansion and inflation.
590 This information is important for understanding how subtle topographic influences can be
591 inherited by inflated lobes to develop preferred lava pathways that influence the overall
592 development of a flow and for interpreting the emplacement history of pāhoehoe lava
593 flows on Earth and other planetary surfaces.

594

595 **Acknowledgments:**

596 *We thank Benjamin Brooks and the Pacific GPS facility for providing access to*
597 *DGPS survey equipment and post-processing resources, Samuel Hulme for his assistance*

598 *with GMT, Richard Herd for kindly providing DGPS data for photogrammetry control,*
599 *Andy Harris for his assistance in the field making contemporaneous FLIR observations,*
600 *Tim Orr for providing historical lava flow data shown in Figure 1, as well as Sarah*
601 *Fagents, Thorvaldur Thordarson, and Jacob Bleacher for many insightful discussions*
602 *relating to lava flow emplacement. Prof. S. Robson and Prof. J. P. Muller are thanked for*
603 *their ongoing support through provision of VMS and GOTCHA, respectively. Christopher*
604 *Kilburn and Jim Kauahikaua are sincerely thanked for their thorough and constructive*
605 *reviews. Field work was conducted in Hawaii Volcanoes National Park under Scientific*
606 *Research and Collecting Permit # HAVO-2006-SCI-0003. CWH was supported by an*
607 *appointment to the NASA Postdoctoral Program at the Goddard Space Flight Center,*
608 *administered by Oak Ridge Associated Universities through a contract with NASA. LSG*
609 *and SMB research was supported by the NASA Planetary Geology and Geophysics, and*
610 *Mars Data Analysis programs (LSG: 811073.02.01.04.44 and 203959.02.03.17.56; SMB:*
611 *NNX08AF16G and NNX10AP63G).*

612 **References**

- 613 Anderson SW, Stofan ER, Smrekar SE, Guest JE, Wood B (1999) Pulsed inflation of
614 pahoehoe lava flows: implications for flood basalt emplacement. *Earth Planet Sci*
615 *Lett* 168:7–18
- 616 Anderson SW, Smrekar SE, Stofan ER (2012) Tumulus development on lava flows:
617 insights from observations of active tumuli and analysis of formation models. *Bull*
618 *Volcanol* 74:931–946. doi:10.1007/s00445-012-0576-2
- 619 Baloga S (1987) Lava flows as kinematic waves. *J Geophys Res* 92 (B9):9271–9279
- 620 Baloga S, Pieri DC (1986) Time-dependent profiles of lava flows. *J Geophys Res*
621 91:9543–9552
- 622 Baloga S, Glaze LS (2003) Pahoehoe transport as a correlated random walk. *J Geophys*
623 *Res* 108:2031. doi:10.1029/2001JB001739
- 624 Baloga S, Glaze LS (2008) Self-replication model for long channelized lava flows on the
625 Mars plains. *J Geophys Res* 113:E05003. doi:10.1029/2007JE002954
- 626 Batchelor GK (1980) *An introduction to fluid dynamics*. Cambridge Univ Press, New
627 York, 615 p
- 628 Bruno BC, Taylor GJ, Rowland SW, Baloga SM (1994) Quantifying the effect of
629 rheology of lava-flow margins using fractal geometry. *Bull Volcanol* 56:193–206
- 630 Byrnes JM, Crown DA (2001) Relationships between pahoehoe surface units,
631 topography, and lava tubes at Mauna Ulu, Kilauea Volcano, Hawaii. *J Geophys*
632 *Res* 106(B2):2139–2151

633 Byrnes JM, Ramsey MS, Crown DA (2004) Surface unit characterization of the Mauna
634 Ulu flow field Kilauea Volcano, Hawai‘i, using integrated field and remote
635 sensing analyses. *J Volcano Geotherm Res* 135:169–193

636 Calvari S, Pinkerton H (1999) Lava tube morphology on Etna and evidence for lava flow
637 emplacement mechanisms. *J Volcanol Geotherm Res* 90:263–280

638 Cashman KV, Kauahikaua JP (1997) Reevaluation of vesicle distributions in basaltic lava
639 flows. *Geology* 25:419–422. doi:10.1130/0091-7613

640 Castruccio A, Rust AC, Sparks RJS (2013) Evolution of crust- and core-dominated lava
641 flows using scaling analysis. *Bull Volcanol* 75:681. doi:10.1007/s004455-012-
642 0681-2

643 Crisp JA, Baloga SM (1990) A model for lava flows with two thermal components. *J*
644 *Geophys Res* 95:1255–1270

645 Crown DA, Baloga SM (1999) Pahoehoe toe dimensions morphology, and branching
646 relationships at Mauna Ulu, Kilauea Volcano, Hawai‘i. *Bull Volcanol* 61:288–305

647 Danes ZF (1972) Dynamics of lava flows *J Geophys Res* 77:1430–1432

648 Doane DP (1976) Aesthetic frequency classifications. *American Statistician* 30:181–1905

649 Dragoni M, Bonafede M, Boschi E (1986) Downslope models of a Bingham liquid:
650 implications for lava flow. *J Volcanol and Geotherm Res* 30:305–325

651 Day T, Muller JP (1989) Digital elevation model production by stereo-matching SPOT
652 image-pairs: A comparison of algorithms. *Image Vision Comp* 7:95–101

653 Gruen AW (1985) Adaptive least squares correlation: A powerful image matching
654 technique. *S Afr J Photogramm Remote Sens Cartogr* 14:175–187

655 Glaze LS, Baloga SM (2013) Simulation of inflated pahoehoe lava flows. *J*

656 Volcanol Geotherm Res, 255:108–123. doi:10.1016/j.jvolgeores.2013.01.018

657 Guest JE, Wood C, Greeley R (1984) Lava tubes, terraces and megatumuli on the 1614–
658 24 pahoehoe lava flow field, Mount Etna, Sicily. Bull Volcanol 47(3):635–648

659 Hamilton CW, Thordarson T, Fagents SA (2010) Explosive lava–water interactions I:
660 architecture and emplacement chronology of volcanic rootless cone groups in the
661 1783–1784 Laki lava flow, Iceland. Bull Volcanol 72(4):449–467.
662 doi:10.1007/s00445-009-0330-6.

663 Harris AJL, Rowland SK (2001) FLOWGO: A kinematic thermo-rheological model for
664 lava flowing in a channel. Bull Volcanol 63:20–44

665 Harris AL, Dehn J, James MR, Hamilton C, Herd R, Lodato L, Steffke A(2007a)
666 Pahoehoe flow cooling, discharge, and coverage rates from thermal image
667 chronometry. Geophys Res Lett 34:L19303. doi:10.1029/2007GL030791

668 Harris AJL, Dehn J, Calvari S(2007b) Lava effusion rate definition and measurement: a
669 review. Bull Volcanol 70:1–22

670 Harris AJL (2013) Lava flows. In: Modeling volcanic processes: the physics and
671 mathematics of volcanism (eds. S.A. Fagents, T.K.P. Gregg, and R.M.C. Lopes).
672 Cambridge Univ. Press, New York, pp 85–106

673 Hon K, Kauahikaua J, Denlinger R, MacKay R (1994) Emplacement and inflation of
674 pahoehoe sheet flows: Observations and measurements of active lava flows on
675 Kilauea Volcano, Hawaii. Geol Soc Am Bull 106:351–370

676 Hoblitt RP, Orr TR, Heliker C, Denlinger RP, Hon K, Cervelli PF (2012) Inflation rates,
677 rifts, and bands in a pāhoehoe sheet flow. Geosphere 8(5):179–195.
678 doi:10.1130/GES00656.1

679 Hulme G (1974) The interpretation of lava flow morphology, *Geophys J R astr Soc*
680 39:361–383

681 Kauahikaua J, Cashman KV, Mattox TN, Heliker CC, Hon LA, Mangan MT, Thornber
682 CR (1998) Observations on basaltic lava streams in tubes from Kilauea Volcano,
683 island of Hawai'i. *J Geophys Res* 103(B11):27,303–27,323

684 Keszthelyi L, Denlinger R (1996) The initial cooling of pahoehoe flow lobes. *Bull*
685 *Volcanol* 58:5–18

686 Keszthelyi L, Self S, Thordarson T (1999) Application of recent studies on the
687 emplacement of basaltic lava flows to the Deccan Traps, *Memoirs – Geolog Soc*
688 *India* 43:485–520

689 Keszthelyi L, Self S, Thordarson T (2006) Flood lavas on Earth, Io and Mars. *J Geol Soc*
690 163(2):253–264

691 Keszthelyi L, McEwen AS, Thordarson T (2000) Terrestrial analogs and thermal models
692 for Martian flood lavas. *J Geophys Res* 105(E6):15,027–15,049

693 Kilburn CRJ (1996) Patterns and predictability in the emplacement of subaerial lava
694 flows and flow fields. In: Scarpa R., R.I. Tilling (eds.) *Monitoring and mitigation*
695 *of volcanic hazards*. Springer, New York: pp 491–537

696 Koeppen WC, Patrick M, Orr T, Sutton AJ, Dow D, Wright R (2013) Constraints on the
697 portioning of Kīlauea's lavas between surface and tube flows, estimated from
698 infrared satellite data, sulfur dioxide emission rates, and field observations. *Bull*
699 *Volcanol* 75:716. doi:10.1007/s00445-013-0716-3

700 Orr TR (2011) Lava tube shatter rings and their correlation with lava flux increases at
701 Kīlauea Volcano, Hawai‘i. *Bull Volcanol* 73:355–346. doi:10.1007/s00445-010-
702 0414-3

703 Otto GP, Chau TKW (1989) Region-growing algorithm for matching of terrain images,
704 *Image Vision Comp* 7:83–94

705 Peitersen MN, Crown DA (2000) Correlations between topography and intraflow width
706 behavior in Martian and terrestrial lava flows. *J Geophys Res* 105(E2):4123–4123

707 Poland MP, Miklius A, Jeff Sutton A, Thornber CR (2012) A mantle-driven surge in
708 magma supply to Kilauea Volcano during 2003–2007. *Nature Geosci* 5(4):295–
709 300

710 Ramsey MS, Harris AJL (2012) Volcanology 2010: How will thermal remote sensing of
711 volcanic surface activity evolve over the next decade? *J Volcanol Geotherm Res.*
712 249:217–233

713 Resnick R, Halliday D (1977) *Physics: Part One* (3rd ed.). John Wiley & Sons, New York,
714 608 p

715 Robson S, James MR (2007) Photogrammetric image sequence processing to determine
716 change in active lava flows. *Proc Remote Sensing and Photogrammetry Society*
717 *Ann Conf, 2007 (RSPSoc 2007)*, 11–14th September, Newcastle upon Tyne, U.K.

718 Rossi MJ, Gudmundsson A (1996) The morphology and formation of flow-lobe tumuli
719 and Icelandic shield volcanoes. *J Volc Geotherm Res* 72:291–308

720 Rowland SK, Harris AJL, Garbeil H (2004) Effects of martian conditions on numerically
721 modeled, cooling-limited channelized lava flows. *J Geophys Res* 109(E10010).
722 doi:10.1029/2004JE002288

723 Self S, Thordarson T, Keszthelyi L, Walker GPL, Hon K, Murphy MT, Long P,
724 Finnemore D (1996) A new model for the emplacement of Columbia River
725 basalts as large, inflated pahoehoe lava flow fields. *Geophys Res Lett* 23:2689–
726 2692

727 Self S, Keszthelyi L, Thordarson T (1998) The importance of pāhoehoe. *Ann Rev Earth*
728 *Planet Sci* 26(1):81–110

729 Smith WHF, Wessel P (1990) Gridding with continuous curvature splines in tension.
730 *Geophysics* 55(3):293–305

731 Sturges H (1926) The choice of a class-interval. *J Amer Statist Assoc* 21:65–66

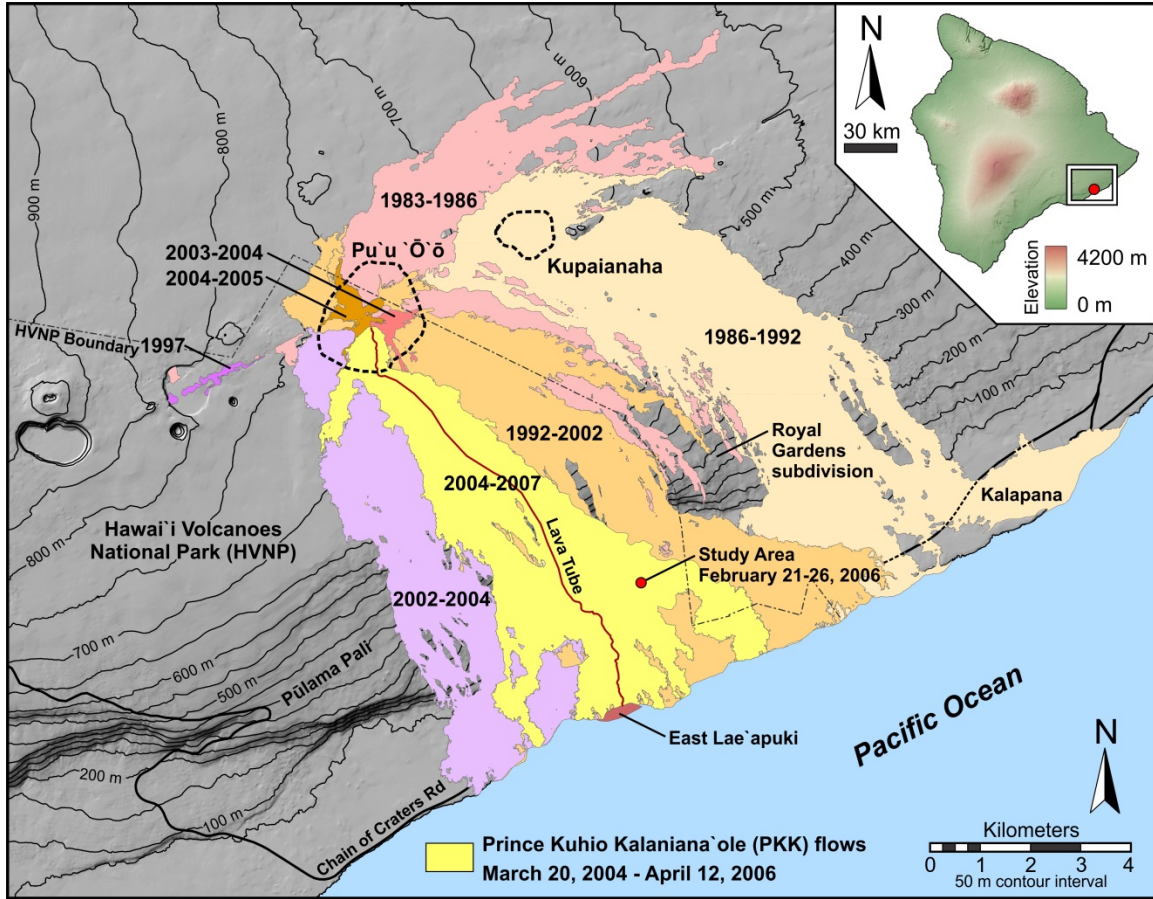
732 Thordarson T, Self S (1998) The Roza Member, Columbia River Basalt Group: A
733 gigantic pahoehoe lava flow field formed by endogenous processes? *J Geophys*
734 *Res* 103(B11):27411–27445

735 Walker GPL (1991) Structure, and origin by injection of lava under surface crust, of
736 tumuli, 'lava rises', 'lava-rise pits', and 'lava-inflation clefts' in Hawaii. *Bull*
737 *Volcanol* 53:546–558

738 Walker GPL (2009) The endogenous growth of pahoehoe lava lobes and morphology of
739 lava-rise edges. In: Thordarson T, Self S, Larsen G, Rowland SK, Hoskuldsson A
740 (eds) *Studies in volcanology—the legacy of George Walker (Special Publications*
741 *of IAVCEI No. 2)*. The Geol Soc: pp 17–32

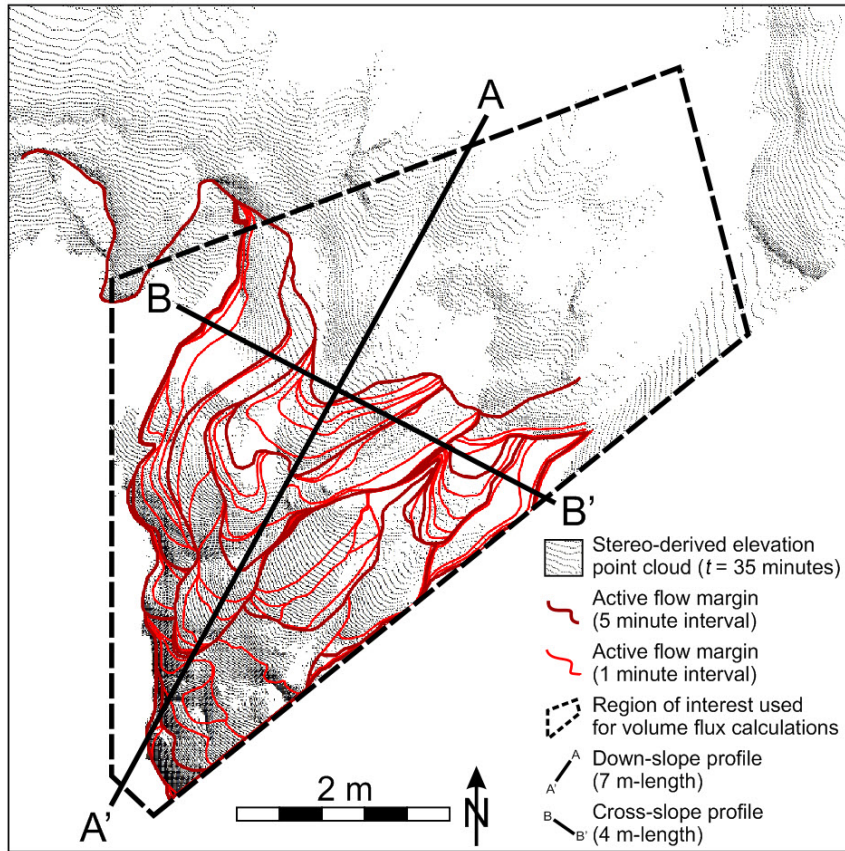
742 Wilmoth RA, Walker GPL (1993) P-type and S-type pahoehoe: a study of vesicle
743 distribution and patterns in Hawaiian lava flows. *J Volcanol Geotherm Res*
744 55:129–142

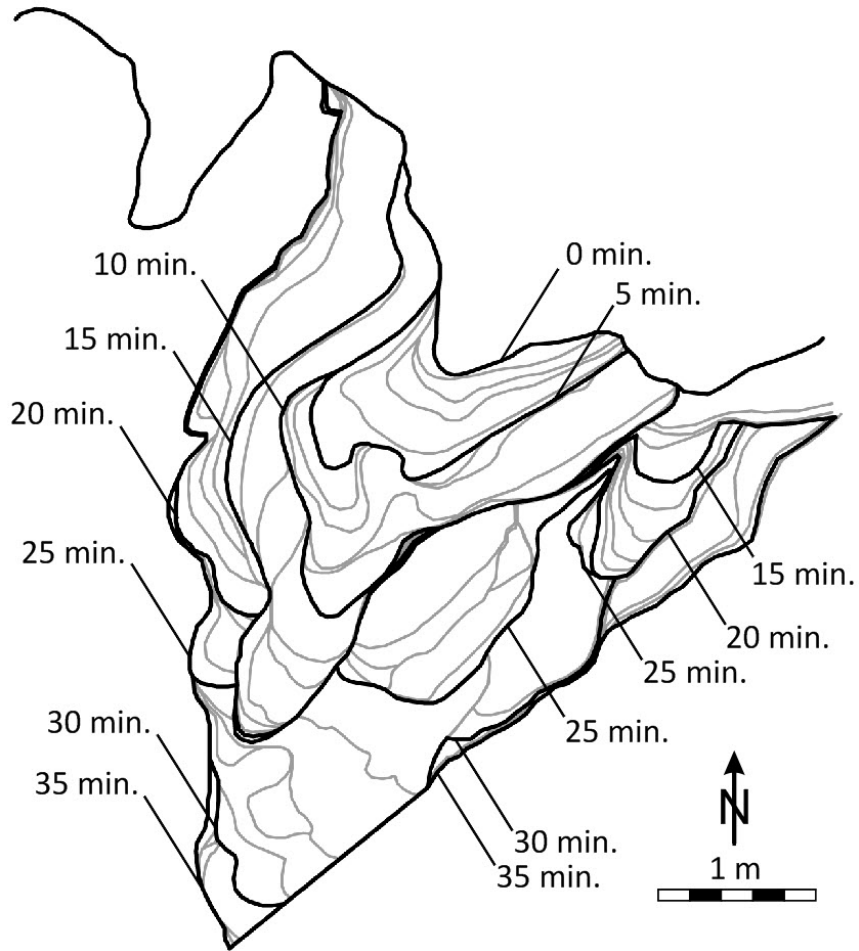
745 **Figures**



746

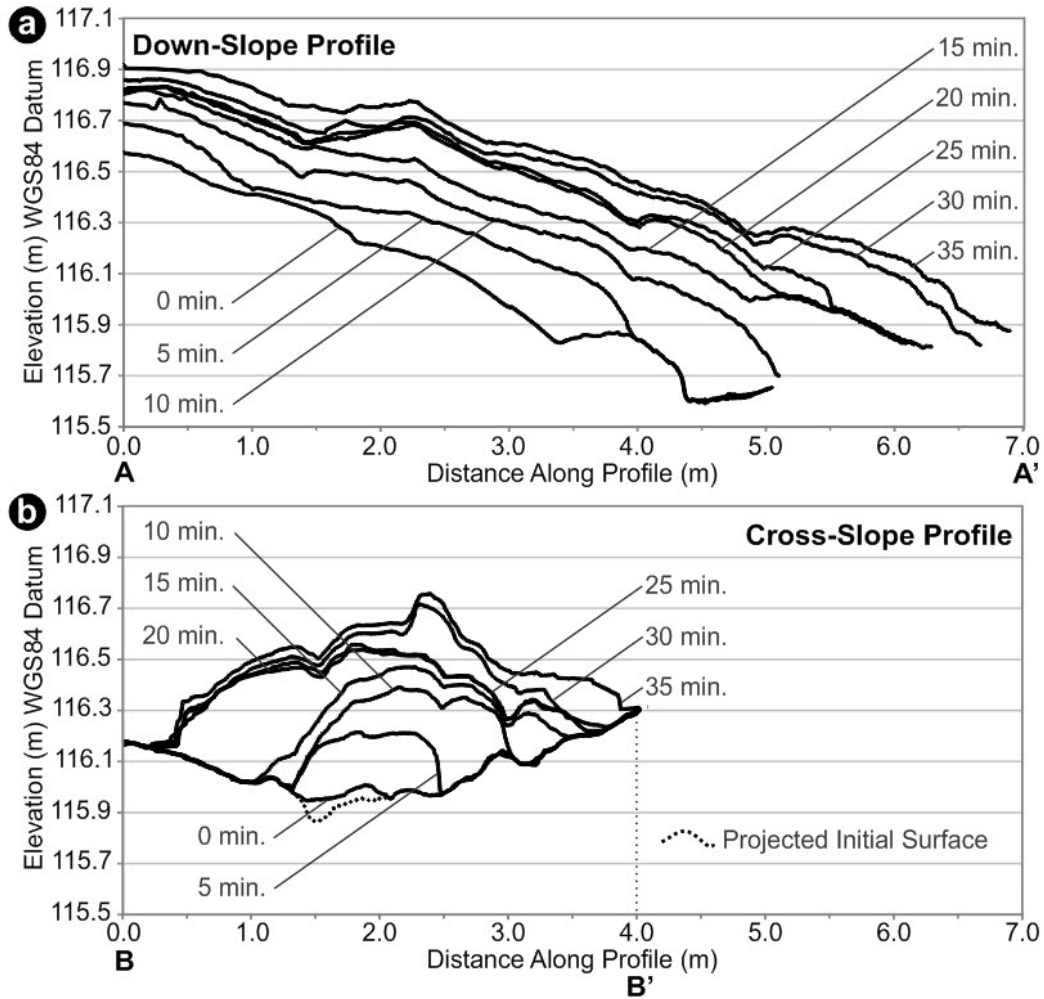
747 Figure 1. Lava flow map showing the study area location (19.336430°N, -
748 155.055235°W) within the Prince Kūhiō Kalanianaʻole (PKK) flow (shown in yellow) on
749 Kīlauea Volcano, Hawaiʻi. Pāhoehoe lava lobes at this locality were fed lava from Puʻu
750 ʻŌʻō. Field measurements were acquired on 21–26 February 2006 for two lobes (A and
751 B) located along the same lava pathway system. Lobe B was located directly down-slope
752 of Lobe A, with a separation distance of ~10 m. The dashed lines approximately bound
753 the edifices of Puʻu ʻŌʻō and Kupaianaha. Lava flow data provided by Tim Orr, United
754 States Geologic Survey (USGS).





759

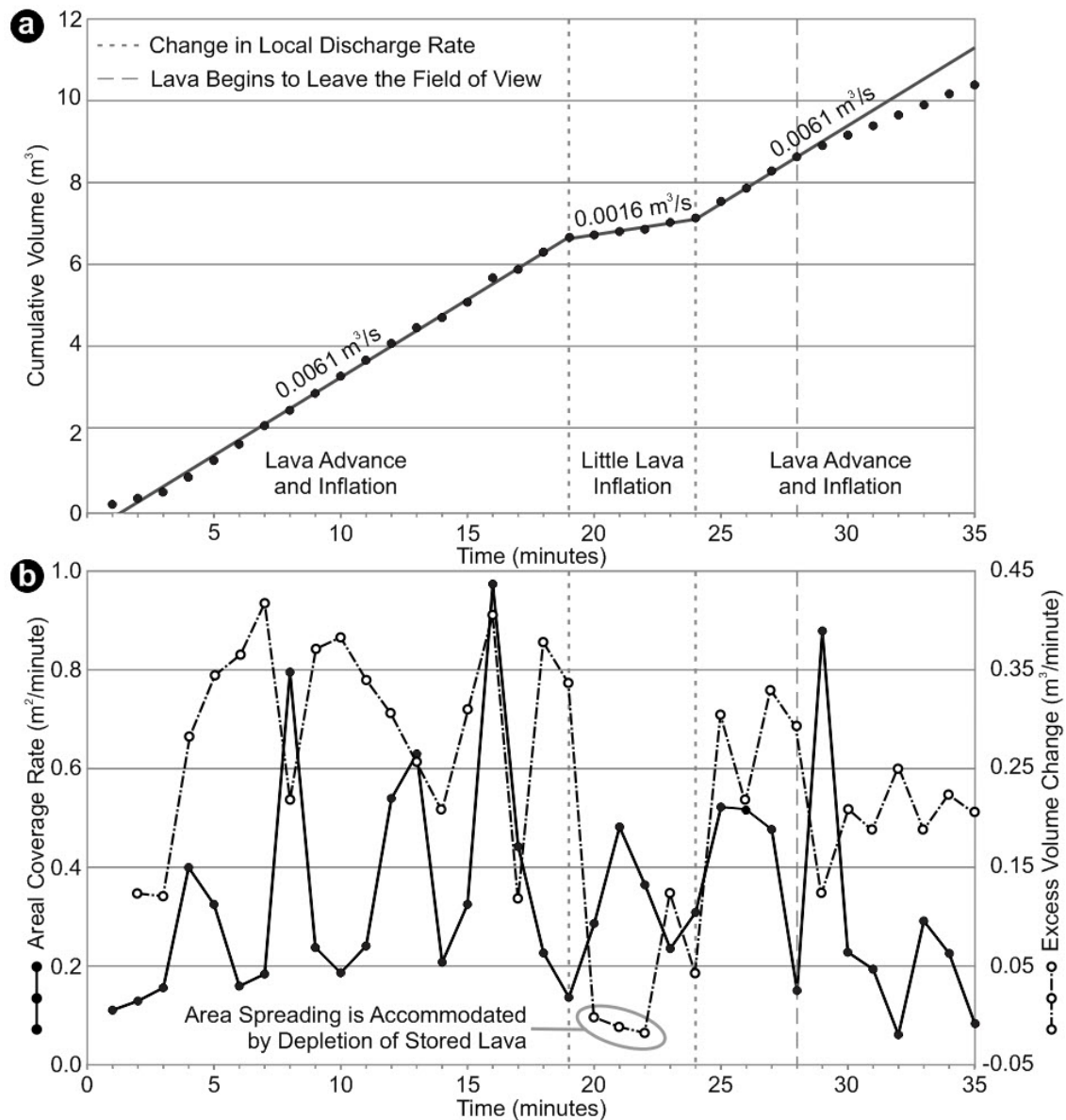
760 Figure 3. Time series evolution of Lobe A. Dark lines show the extent of the flow at 5
 761 minute intervals with grey lines showing the flow margins at 1 minute intervals. The flow
 762 begins to move outside the field of view at 28 minutes.



763

764 Figure 4. (a) Down-slope ($\overline{AA'}$) and (b) cross-slope ($\overline{BB'}$) profiles for Lobe A showing

765 the changing topography of the lava lobe shown in Figures 1 and 2.



766

767 Figure 5. (a) Cumulative volume in the region of interest identified in Figure 2 for Lobe

768 A. The mean local discharge rate ($0.0061 \pm 0.0019 m^3 s^{-1}$, calculated for the first 19

769 minutes) is generally constant during the observation period, except for the interval

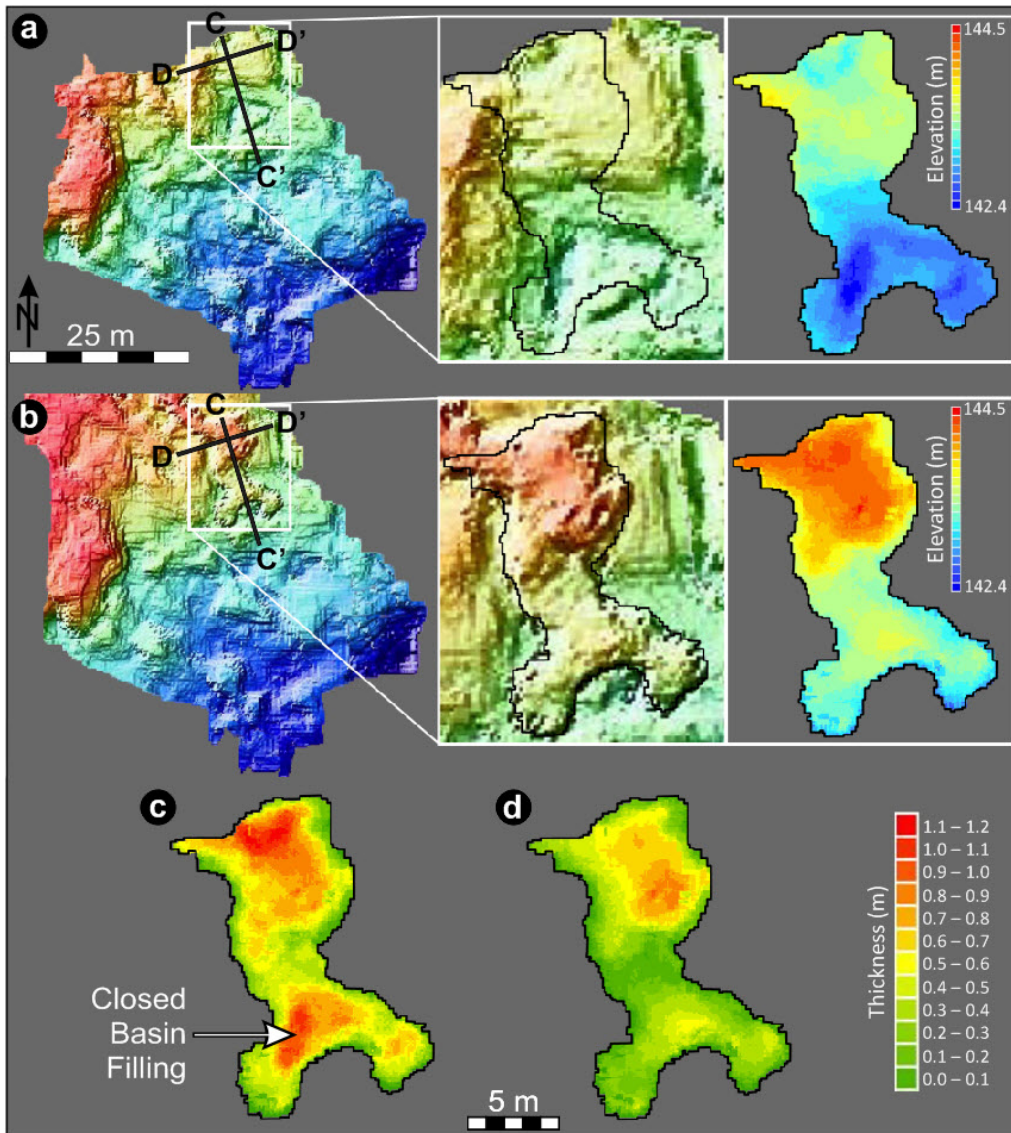
770 between 19 and 24 minutes where the mean local discharge rate decreases by nearly a

771 factor of four and the lobe experiences little inflation and generally lower than average

772 areal spreading rates. At 28 minutes, the rate of volume change gradually decreases as the

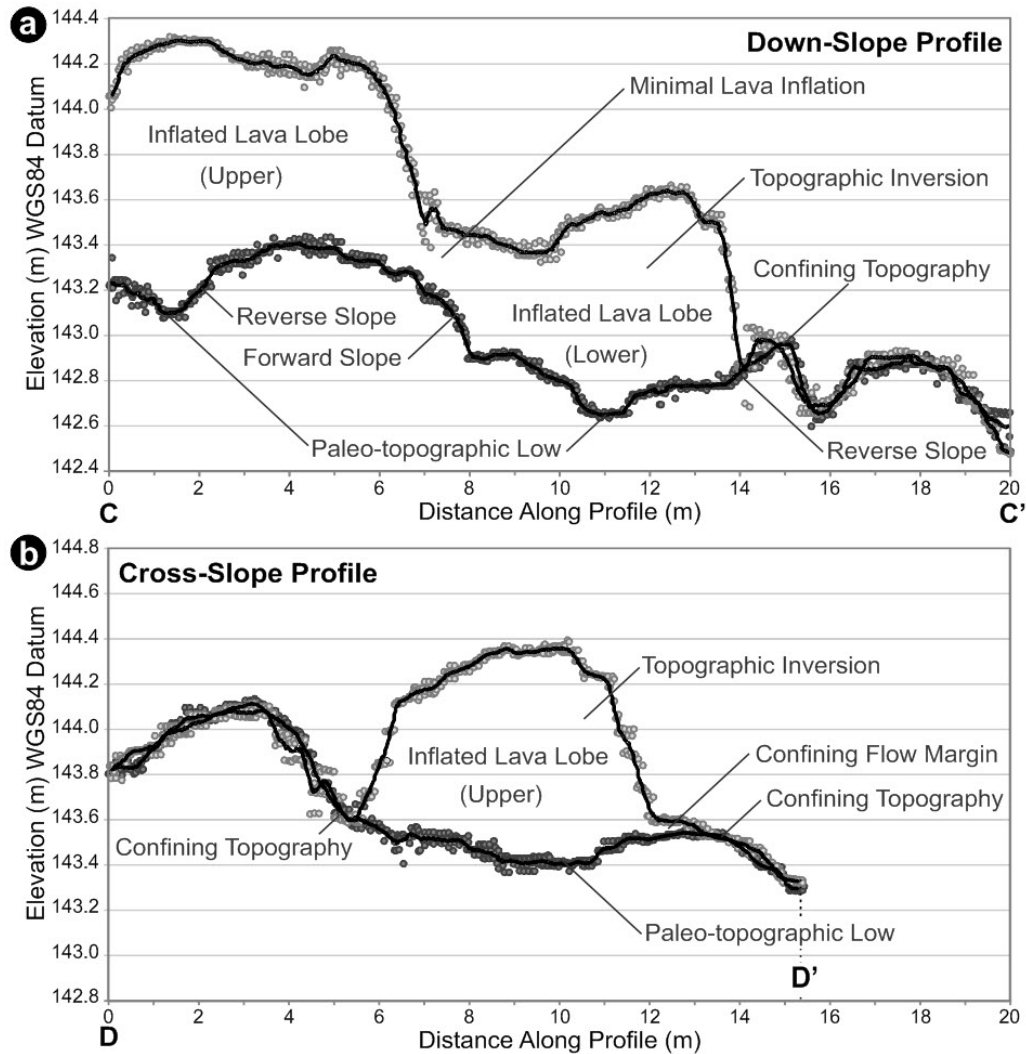
773 active flow front begins to exit the field of view of the stereo-cameras. (b) Areal coverage

774 rate (mean $0.33 \pm 0.23 \text{ m}^2/\text{minute}$, filled black circles with solid black line) and mean
775 excess volume change ($0.23 \pm 0.12 \text{ m}^3/\text{minute}$, open circles with broken line), where
776 excess volume change is defined in the text. Excess volume changes less than zero,
777 between 20–23 minutes, implies that areal growth was accommodated by depletion of
778 stored lava during this time interval. The areal coverage rate shows a periodicity with
779 peaks at intervals of 4.13 ± 0.64 minutes. All uncertainties are reported at 1σ .



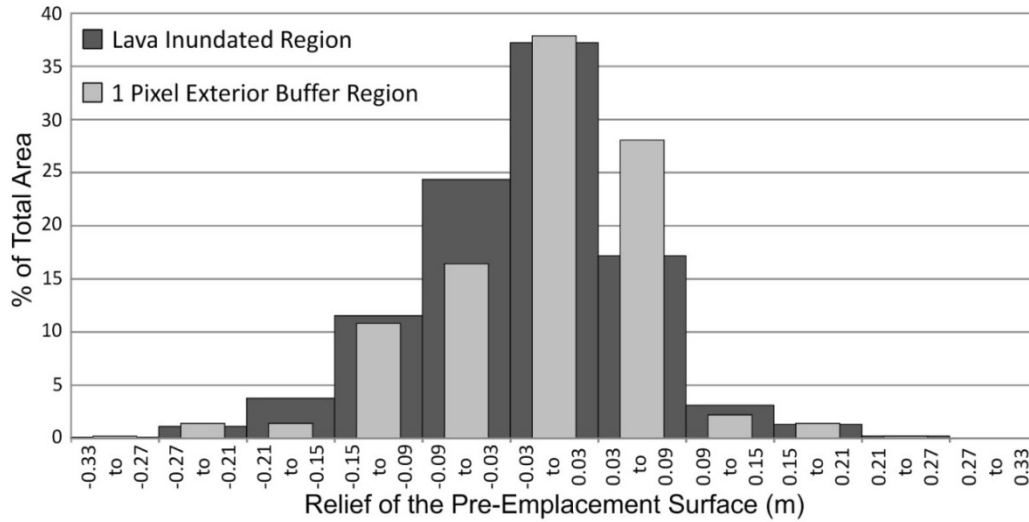
780

781 Figure 6. (a) Initial and (b) post-lava lobe emplacement digital terrain models (DTMs)
 782 generated from Differential Global Positioning System (DGPS) measurements for Lobe
 783 B, with cool colors showing low elevation and warm colors showing relatively higher
 784 elevation. (c) Simple difference between the final and initial topography. (d) Difference
 785 between the final topography and a Laplacian harmonic surface solution, which removes
 786 the effects of closed basin filling. Profiles $\overline{CC'}$ and $\overline{DD'}$ are shown in Figures 7a and 7b,
 787 respectively. Elevations in (a) and (b) are reported relative to the WGS84 datum, and lava
 788 thicknesses in (c) and (d) are shown to the same scale.



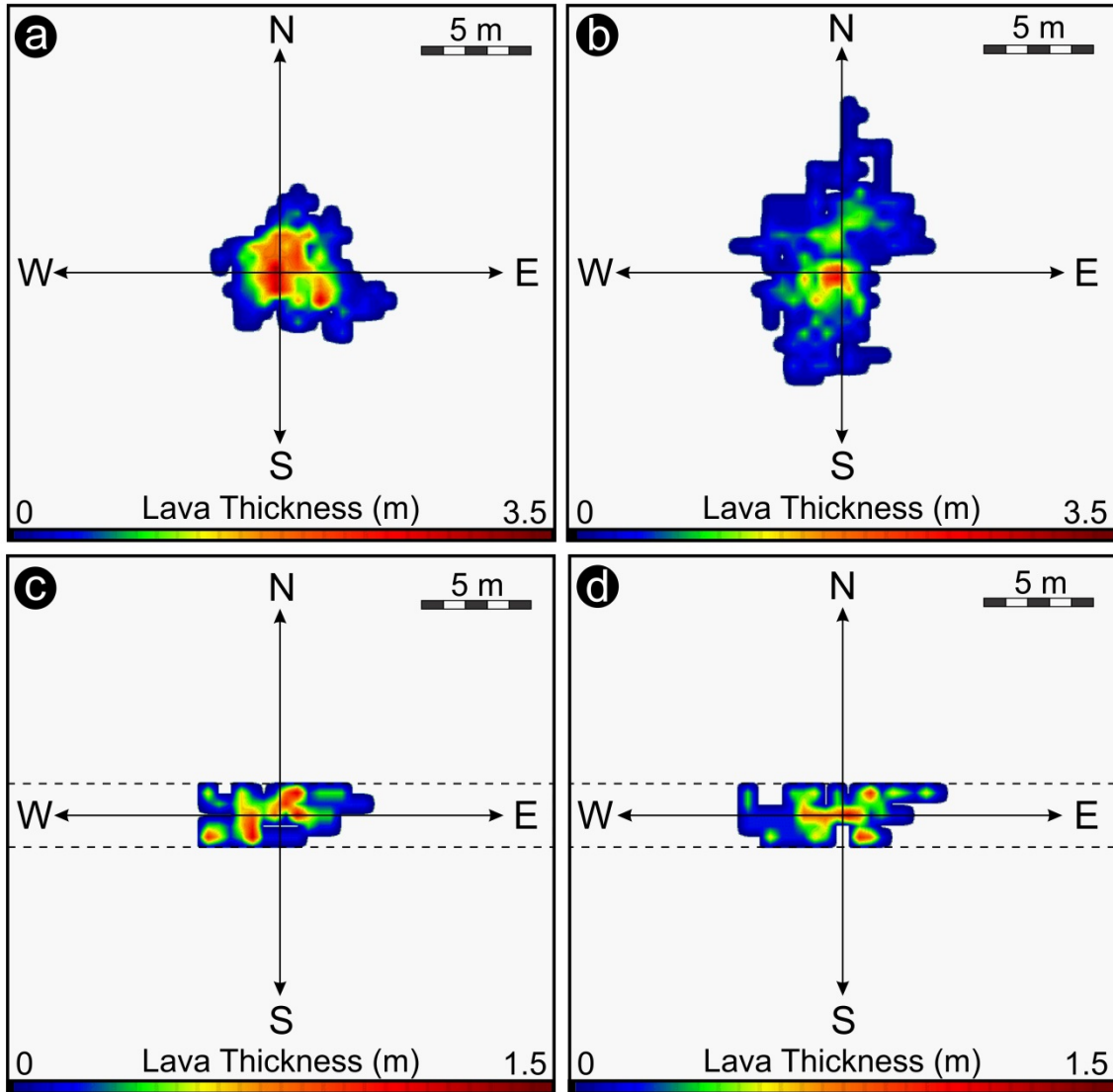
789

790 Figure 7. (a) Down-slope and (b) cross-slope profiles showing the initial and final
 791 topography for Lobe B (see $\overline{CC'}$ and $\overline{DD'}$ in Figure 6). The flow has topographically
 792 inverted the landscape, with inflation being greatest above paleo-topographic lows and
 793 anti-facing slopes, and least above forward-facing slopes. Note the pre-eruption
 794 topography and thin confining flow margins that may help to confine the thicker interior
 795 portions of the lobe.



796

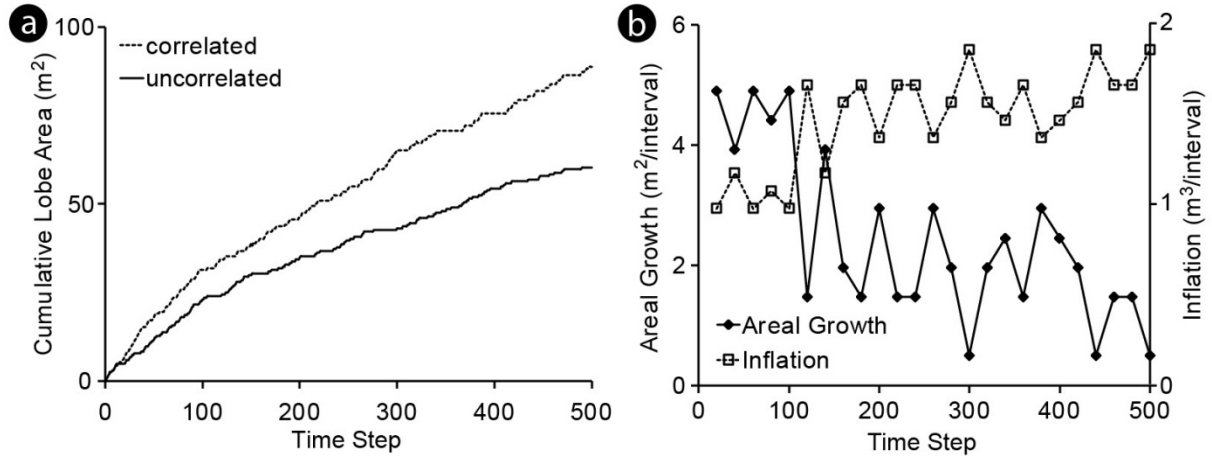
797 Figure 8. Frequency distribution of relief on initial surface within the lava inundated
 798 region (dark grey) and within a 1 pixel buffer region around Lobe B (light grey). The
 799 mean relief within both regions is close to zero (i.e., 0.02 ± 0.07 m in the inundated
 800 region and 0.00 ± 0.07 m in the buffer region), but the buffer region exhibits more +0.03
 801 to +0.09 m relief surfaces and -0.09 to -0.03 m relief surface, which implies that positive
 802 relief on the order of only a few centimeters may act to confine the areal spreading of the
 803 flow. All uncertainties are reported at 1σ .



804

805 Figure 9. (a) Simulation of 500 random emplacement lava parcels (without
 806 sequential breakouts from lobe margins). The maximum flow thickness is 3.2 m. (b)
 807 Simulation of 500 lava parcels using the same probability distribution for correlated toe
 808 growth used to model Lobe A (i.e., $P(0) = P(1) = 0.25$ and $P(2) = 0.5$). The maximum
 809 flow thickness is 2.6 m. Note that including correlated breakouts in (b) increases the
 810 perimeter:area ratio relative to the random example shown in (a). (c and d) Plan view
 811 examples of two simulations of Lobe A using 117 lava parcels, boundaries (dash lines) at
 812 ± 3 cells on the North-South axis, and a probability distribution of $P(0) = P(1) = 0.25$ and

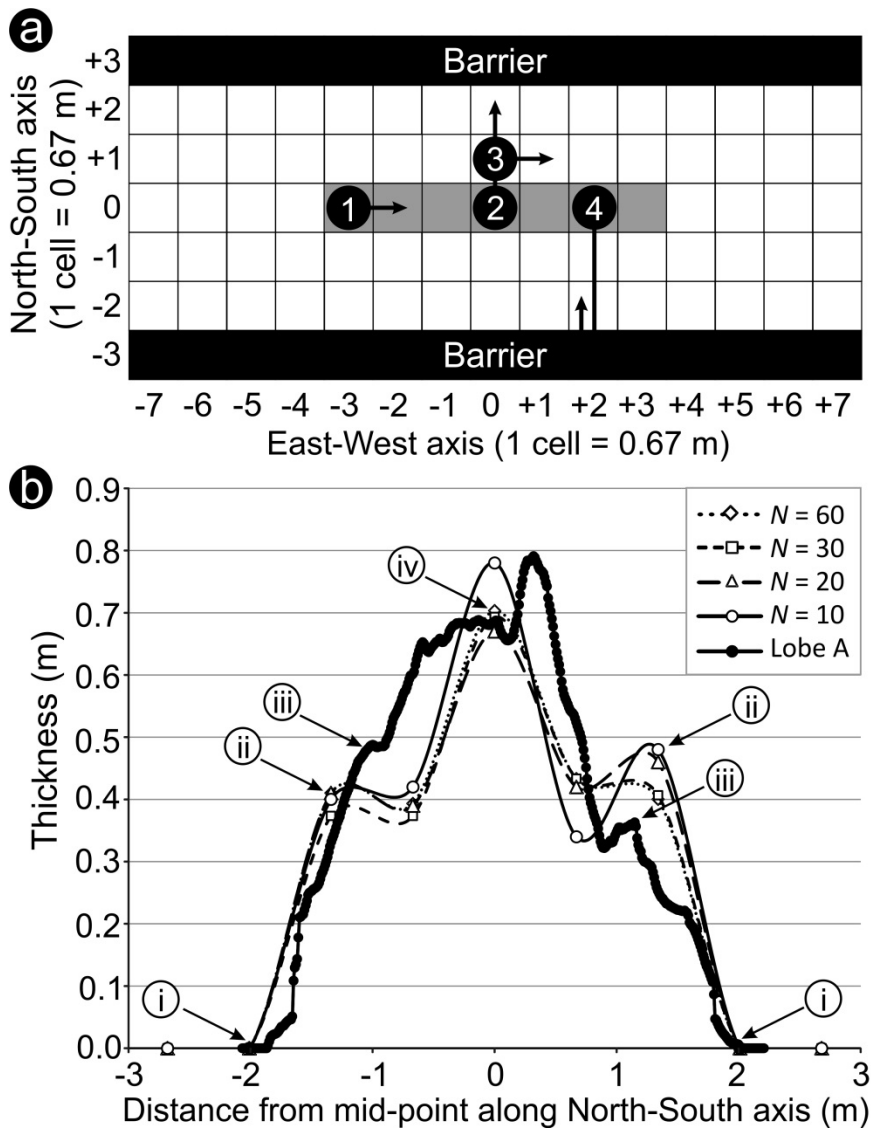
813 $P(2) = 0.5$ for adding 0, 1, and 2 extra parcels in the same direction when a parcel breaks
814 out from the flow margin. The maximum height of the flows shown in (c) and (d) are 1.2
815 m. For visualization purposes, the REBIN function in Interactive Data Language (IDL)
816 was used to magnify the cells by a factor of 10 and perform a default bilinear
817 interpolation. This process smoothes the model output and provides the exterior toes with
818 a rounded appearance. This process does not alter the raw model output shown in Figures
819 10 or 11b.



820

821 Figure 10. (a) Cumulative lobe area versus time for two stochastic simulation realizations
 822 of pāhoehoe lobe emplacement. The uncorrelated realization (lower solid line) was
 823 generated by considering purely random selection of both the location and direction of
 824 each parcel transfer. The correlated scenario (upper dashed line) modifies this model by
 825 introducing an increased probability of sequential parcel emplacement when new parcels
 826 break out at the periphery of the flow. In this case, the probabilities of adding 0, 1, and 2,
 827 correlated parcels are $P(0) = 0.25$, $P(1) = 0.25$, and $P(2) = 0.5$, respectively. (b) Shows
 828 the periodicities and anti-correlation in areal spreading and excess volume change for a
 829 typical purely random simulation and an interval of 20 time steps.

830



831

832 Figure 9. (a) A schematic example showing how the stochastic model works. A linear
 833 source region (shown in gray) passes through the mid-point of the array and trends East-
 834 West. Two parallel barriers are shown ± 3 cells from the linear source along the North-
 835 South axis. In the simulation, a cell is first chosen at random from within the confines of
 836 the flow (initially corresponding to the extent of the linear source region) and then a
 837 random cardinal direction is chosen. A lava parcel is then transferred to the specified cell
 838 (marked with a filled black circle, e.g., "1") and to adjacent cell (identified by the
 839 associated arrow). If the parcel would be allocated beyond the confines of the flow, there

840 is a 0.25, 0.25, and 0.50 probability that 0, 1, or 2 extra parcels, respectively, will be
841 allocated to sequential cells in the same direction (e.g., “2” shows 1 extra, 2 total, cells
842 being added). As the areal extent of the flow increases, new parcel transfer locations
843 become available (e.g., “3”). If a cell allocation would result in a parcel being extended
844 beyond the barrier, then the parcel is reflected back towards the cell interior (e.g., “4”).
845 Each parcel allocated to a given cell increases its thickness by 0.2 m. This process is
846 repeated until all lava parcels are allocated, which in the case of Lobe A, includes a total
847 of 117 parcels. (b) Simulated North-South profiles for Lobe A, averaged over 10, 20, 30,
848 and 40 realizations of the model. The results show that despite the stochastic nature of the
849 model, the average simulations converge to a stable equilibrium shape over a few tens of
850 realizations. Also shown is the flow-perpendicular cross-section through Lobe A (Fig. 4b)
851 with the underlying topographic trend removed. “i”: The maximum width simulated lobes
852 and Lobe A are determined by the influences of topographic barriers. “ii”: Reflecting of
853 lava parcels by the barriers produces an excess thickening in the adjacent interior cells.
854 “iii”: High-standing topography and stagnated peripheral toes helps to confine the interior
855 of the lobe and promote inflation. “iv”: The thickest portion of the simulated lobe is
856 located along the medial axis and is generally consistent with the thickness of Lobe A
857 along the flow axis.