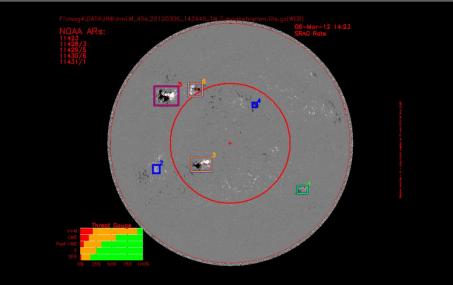
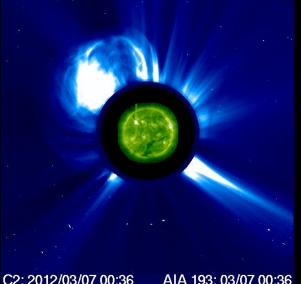


SCIENCE & TECHNOLOGY OFFICE





An All-Clear Space Weather Forecasting System Based on Magnetogram in Near Real Time

David Falconer, Nasser Barghouty, Igor Khazanov, and Ron Moore

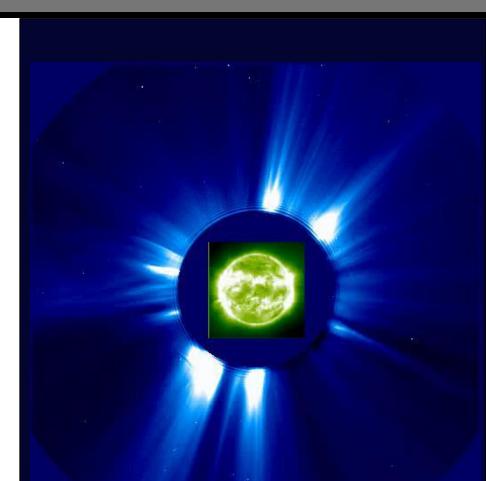
Jan 24, 2016

http://www.uah.edu/cspar/research/mag4-page





- 1. MAG4 background
- 2. Measurement
- 3. Effectiveness
- 4. Transitioning to Vector
- 5. Future improvements

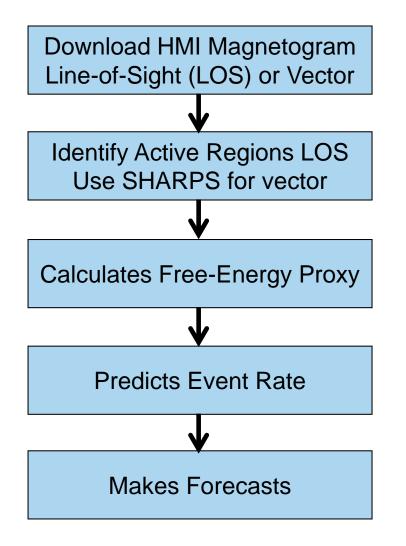


C2: 2000/07/14 09:30:05 EIT: 07/14 09:24:10

Bastille Day (2000 July 14) Flare, Coronal Mass Ejection and Solar Energetic Particle Event







MAG4 is completely automated, from downloading magnetograms to outputting and storing forecast products.





- Flares and CMEs are known to be drivers of the most severe space weather
- Flares and CMEs typically originate in active regions (aka sunspots)
- Flares and CMEs are examples of exceptionally large explosive releases of magnetic energy stored in the corona
- While the amount of free energy cannot be measured directly, freeenergy proxies can be measured
- Event rates have been shown to be correlated with the magnitude of the free-energy proxies





- MAG4 (Magnetogram Forecast), developed originally for NASA/SRAG (Space Radiation Analysis Group), is an automated program that analyzes magnetograms from the HMI (Helioseismic and Magnetic Imager) instrument on NASA SDO (Solar Dynamics Observatory), and automatically predicts the event rate (or probability) of major flares (Mand X-class), Coronal Mass Ejections (CMEs), and Solar Energetic Particle Events
- MAG4 does not forecast that a flare will occur at a particular time in the next 24 or 48 hours; rather the probability of one occurring!
- GONG (<u>Global Oscillations Network Group</u>) magnetograms, can be used instead as a **backup** but at a lower forecast accuracy
- Present cadence of new forecasts: **96 minutes** (60 minutes for CCMC) Vector magnetogram actual cadence: **12 minutes**





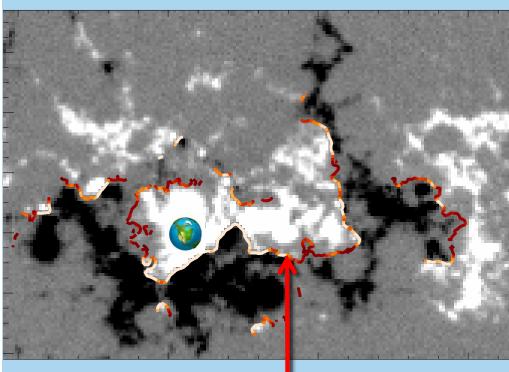
- 1973 The MSFC (Marshall Space Flight Center) vector magnetograph was built to support Skylab
- 2000-present MSFC used vector magnetograms to study CME correlation with freeenergy proxy
- 2007-12 A DOD/Multidisciplinary University Research Initiative/Neutral Atmosphere Density Interdisciplinary Research helped support funding of the basic research
- 2008 Partnered with JSC/SRAG (Space Radiation Analysis Group) and won an R20 NASA/Technical Excellence Initiative grant: Began building a database that grew to ~40,000 magnetograms of ~1,300 active region, covering years 1996-2004 with event catalog from SOHO/MDI (Solar and Heliospheric Observatory/Michelson Doppler Imager) observations
- 2010-present NASA's HEOMD (Human Exploration and Operations Mission Directorate) support
- 2010 SDO is launched began transitioning from MDI to HMI line-of-sight magnetograms.
- 2011 MAG4 installed at SRAG a <u>NRT (Near-Real-Time) forecasting tool</u>, and SRAG began pre-operations testing
- 2012 Provided NOAA web access to MAG4 NRT forecasts
- 2013 Improve MAG4 so that it can use a combination of free-energy proxy and previous flare activity
- 2015 Transition to HMI line-of-sight to vector magnetograms





- When the transverse gradient of the vertical (or line-of-sight) magnetic field is large, there is more freeenergy stored in the magnetic field
- For each Active Region: The integral of the gradient along the neutral line is the free-energy proxy

A magnetogram of an active region

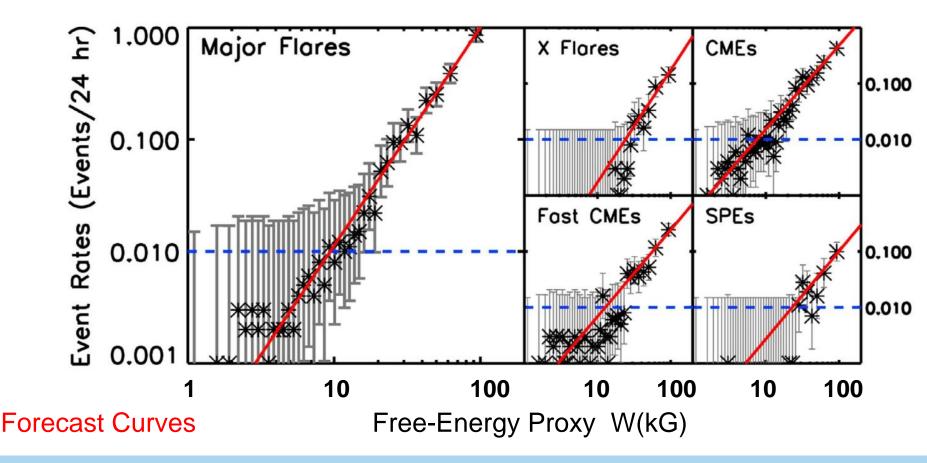


Neutral Line, color coded for gradient









These empirical forecast curves are used to convert our free-energy proxy into predicted event rates. Curves are derived from a sample of 40,000 magnetograms, from 1300 active regions observed between 1996-2004.

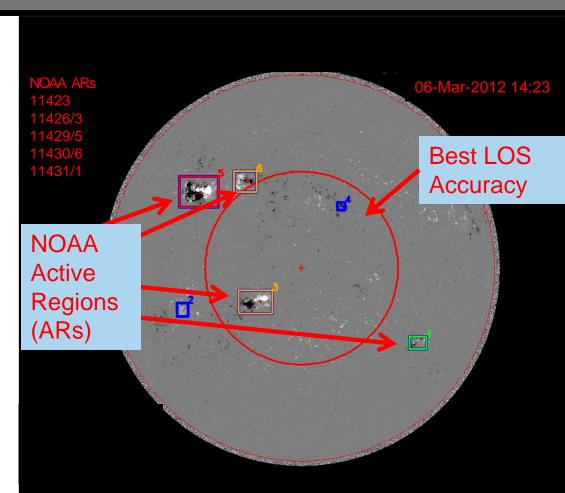


Identifying Active Regions



- Magnetograms are spatial maps of the magnetic field strengths
- They come in two basic types
 - line-of-sight (right)
 - vector magnetograms
- Free-energy proxies can be measured for Active Regions (areas with sunspots) from either type of magnetogram
- Line-of-sight magnetograms suffer reduced accuracy further from disk center

Magnetograms & identify ARs

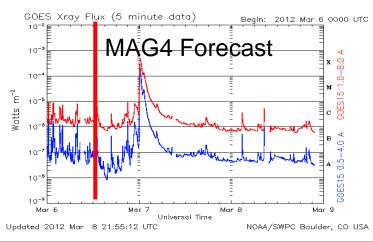


A full-disk line-of-sight magnetogram of the Sun, from SDO/HMI.





Multiplicative uncertainty example Probability Rate 1 sigma Events/day 66% Confidence 0.02 0.01-0.05 0.7-5% 0.7 0.3-1.9 20-80% For a Multiplicative Uncertainty of 2.7x GOES13 Proton Flux (5 minute data) Begin: 2012 Mar 6 0000 UTC Threat Gauae /e/ NOAA/SWPC Boulder, CO



2012/03/06	6 14:23	
# AR#	WL_{SG}	Lng Lat
	(kG)	(deg)
3 11428	9	-21 -17
5 11429	65	-41 17
6 11430	11	-25 20
1 11431	1	36 -27
	Disk Fore	cast Rates
Multi	plicative U	Incertainties
		ast Probabilities
	Uncer	rtainties

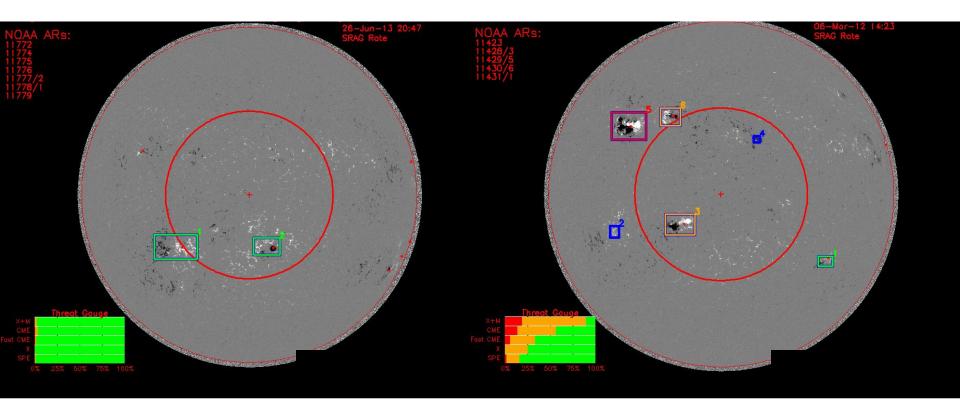
24 Ho	our Eve	ent Rate	9		Dist
M&X	CME	FCME	X	SPE	(deg)
0.020	0.020	0.009	0.002	0.003	27
0.700	0.400	0.200	0.100	0.080	44!
0.020	0.030	0.010	0.004	0.005	32!
0.000	0.001	0.001	0.000	0.000	45!
0.800	0.400	0.200	0.100	0.090	
2.7x 2	2.1x 2.	2x 3.0)x 2.4	х	
50.00%	% 70.00)% 80.0	0% 90	.00% 9	2.00%
40.00	% 20.0	0% 10.	00% 10	0.00%	7.00%

Science for Space Weather





June 26, 2013 C1, C1.5 flares March 7, 2012 X5.4, X1.3, C1.6 CME 2684, 1825 km/sec, Solar Energetic Proton Event reaches 6530 particle flux unit >10MeV







2. Skill Metrics Significance of Upgrade

Forecast Method	YY	YN	NY	NN	PC(%)	POD	FAR	HSS	TSS
McIntosh/ NOAA	259	638	631	18476	93.7	0.29	0.71	0.26	0.26
Free-Energy Proxy Present MAG4	273	284	618	18830	95.5	0.31	0.50	0.35	0.47
Free-energy proxy and previous flare activity Upgraded MAG4	340	317	551	18797	95.7	0.38	0.48	0.42	0.49
Best	890	0	0	19114	100	1	0	1	1

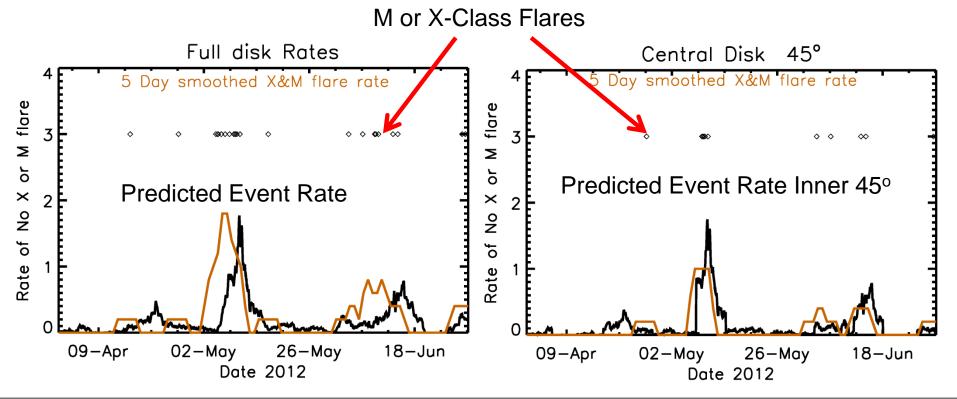
Improvement in Metric	PC(%)	POD	FAR	HSS	TSS
McIntosh/NOAA	1.8±0.5	0.03±0.05	0.21±0.07	0.10±0.04	0.21±0.07
Present MAG4	(4σ)	(0.3σ)	(3σ)	(2σ)	(3σ)
Present MAG4	0.2±0.2	0.08±0.03	0.02±0.05	0.06±0.03	0.03±0.05
Upgraded MAG4	(0.7σ)	(2σ)	(0.5σ)	(2σ)	(0.5σ)





1. Situational Awareness

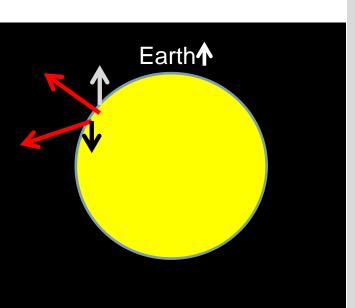
- During periods when flare-productive active regions cross the disk, the predicted rate and actual rate both increase, providing situational awareness
- The results are best when flares and predicted rates are limited to inner 45 degree circle (Right)







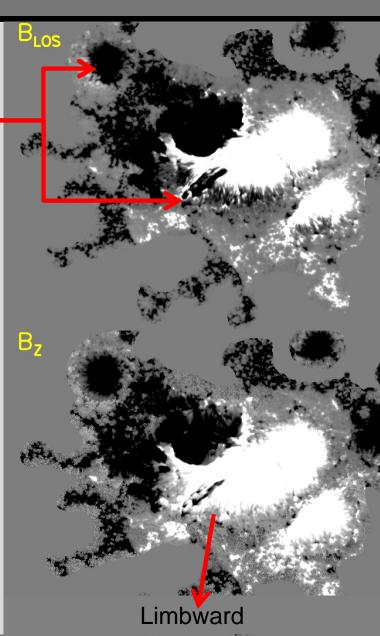
 Both vectors shown in red have positive B_z (magnetic field out of the sun), but have opposite sign B_{LOS} and thus a false (unphysical) neutral line in the line-of-sight (LOS) field.



Actual Examples -

False Neutral Lines occur on limbward sides of sunspots.

Problem fixed by converting from B_{LOS} and $B_{Transverse}$ to B_{Z} and $B_{Horizontal}$

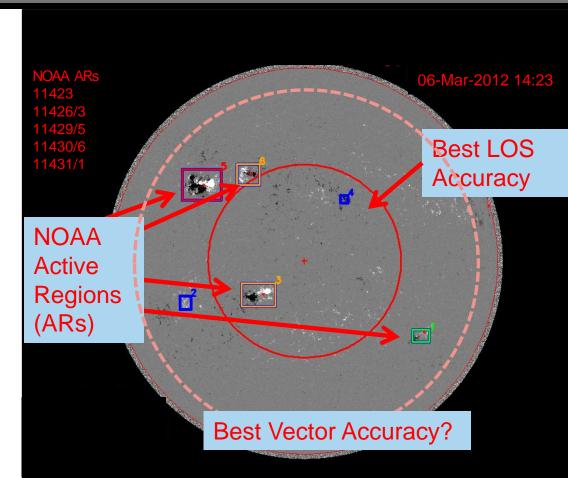






- For LOS most accurate to 30 heliocentric degrees. (RED)
- For Vector similar accuracies should exist to 60 heliocentric degrees.



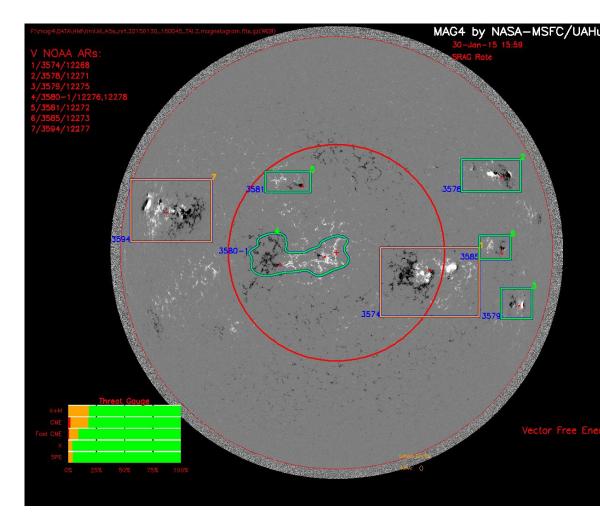


A full-disk line-of-sight magnetogram of the Sun, from SDO/HMI.





- MAG4 presently
 - Predicts what MDI will measure if AR was at disk center.
 - Then uses the forecast curves from MDI to forecast events.
- MAG4 needs to upgrade to take full advantage of vector magnetograms.
- <u>http://www.uah.edu/csp</u> <u>ar/research/mag4-page</u>











2. Skill Metrics Equations

	Actual Yes	Actual No
Predict Yes	YY	YN
Predict No	NY	NN

Metric Equations

Percent Correct Probability of Detection False Alarm Rate Heidke Skill Score

True Skill Score

PC=(YY+NN)/(YY+YN+NY+YY) POD=YY/(YY+NY) FAR=YN/(YY+YN) HSS=2*(YY*NN-YN*NY)/[(YY+NY)* (NY+NN)+(YY+YN)*(YN+NN)] TSS=(YY*NN-NY*YN)/((YY+NY)*(YN+NN))





2. Skill Metrics

Truth Table	Actual Yes	Actual No	PC POD	Percent Correct (0 to 100%) Probability of Detection (0 to 1)
Predict Yes	YY	YN	FAR	False Alarm Rate (1 to 0)
Predict No	NY	NN	HSS	Heidke Skill Score (-1 to 1)
			TSS	True Skill Score (-1 to 1)

Forecast Method	YY	YN	NY	NN	PC (%)	POD	FAR	HSS	TSS
McIntosh/ NOAA	259	638	631	18476	93.7	0.29	0.71	0.26	0.26
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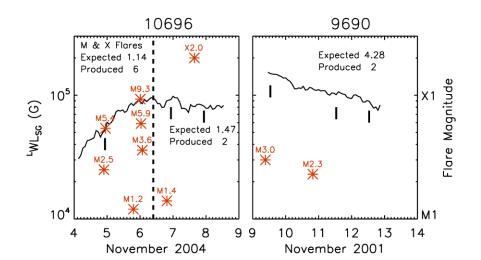


- Correct for systemic errors in measurements due to radial dependence of transverse field noise and foreshortening.
- Access skill scores relative to present LOS
- Investigate if a vector measure is better than LOS measure
- Investigate if change of measure is important
- Investigate if combination produces better measure
- Correct SPE forecasts based on Longitude of source region
- Partner with heliosphere modeling to better forecast where the particles would go.
- Twin CME`

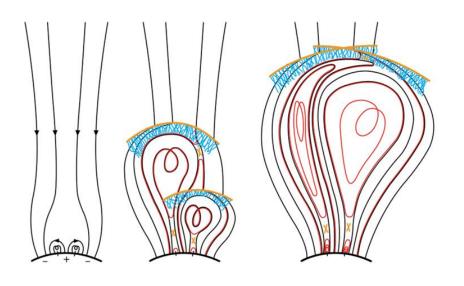




 Determine if growth/decay phase of an active region has additional prediction ability



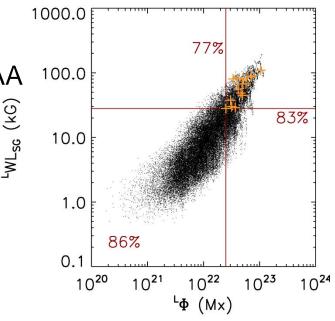
 Determine if looking at twin-CME or single-CME can produce improved SEP forecasts







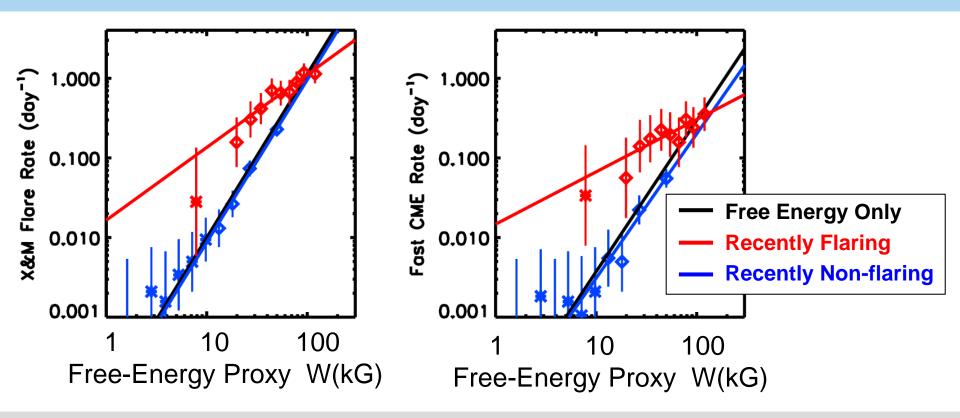
- Continue to supply support, and upgrades to NASA/SRAG
- Finish and evaluate using vector magnetograms
- Fully implement in consultation with SRAG using combined free-energy proxy and previous flare activity
- Improve All Clear Forecast for SEP
- Partner with Air Force, and through them NOAA







1. Recent Flare History



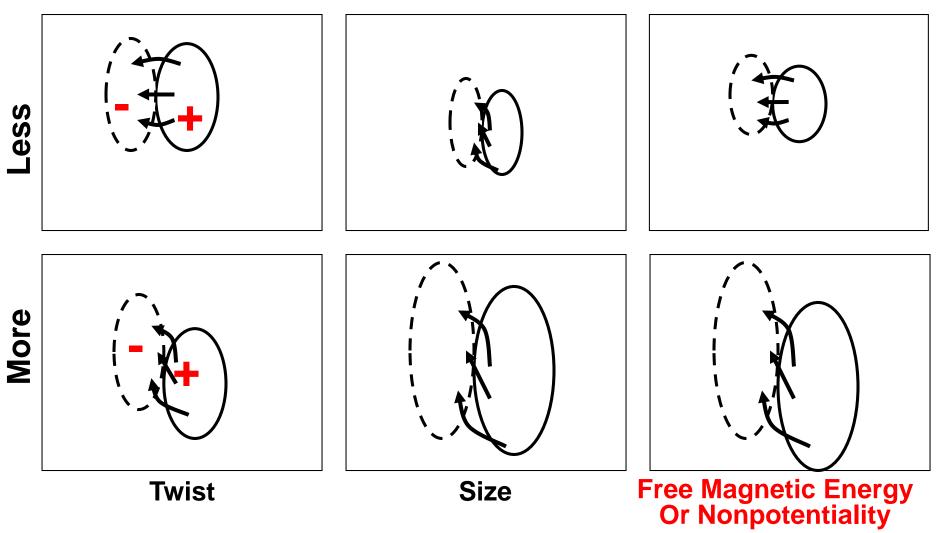
Active regions that have recently produced an X- or M-Class flare are more likely to produce flares in the near future





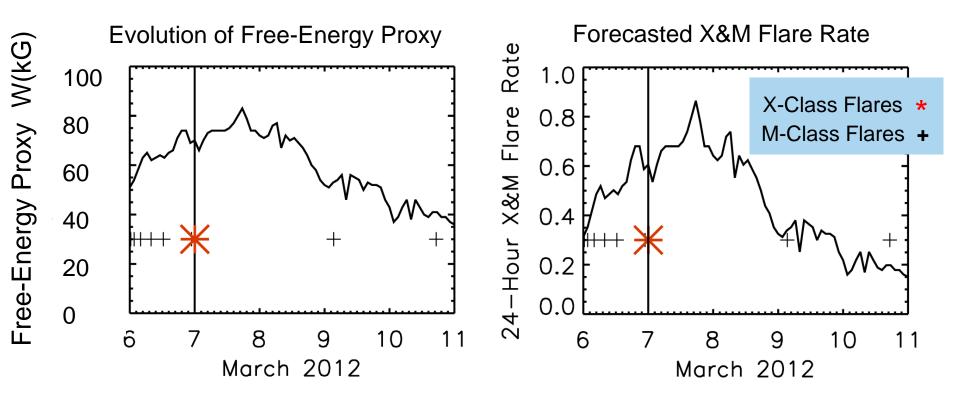


Currents ~10¹² Amps









The free-energy proxy evolves on time periods of days => the forecast is on those time scales





Flares occur when high free-energy proxy active regions crossing disk.

