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Earth total water content Anne Peslier

Jacobs, NASA-JSC, Houston TX, USA















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- \diamond (Water in the exosphere)
- \diamond Water in the crust
- \diamond Water in the mantle
- \diamond Water in the core





Image Credit: Japan Times

Glacier in the Miller range, Antarctica

\diamond Type of data

- ♦ Water in the mantle
 - \diamond Definition
 - \diamond Importance
 - \diamond Distribution & controls
 - ♦ H diffusion
 - $\diamond~$ Water in the continental mantle lithosphere
 - ♦ Water in the lithosphere
 - \diamond Water in the oceanic lithosphere
- \diamond The big picture
 - \diamond Water in the Earth layers \diamond Fluxes
- ♦ Comparison with other differentiated planetary bodies



Image Credit: Japan Times

Type of data for water in the deep Earth

Crust

 \diamond Direct samples: crustal rocks

♦ Mantle lithosphere

♦ Direct samples: Peridotites, pyroxenites samples

♦ Asthenosphere

♦ Direct samples: Rare deep diamond inclusions

♦ Indirect samples: Oceanic basalts

(undegassed glasses, melt inclusions)

♦ HP Experiments

 \diamond First principle calculations

∻Core

♦ H solubility in Fe-Ni metal
 ♦ Accretion models



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Image Credit: Japan Times

Disclaimer about "water" in the mantle

Water" in anhydrous minerals = Hydrogen

Ol, Px, Gr, Fd, Apatite, Amph, Mica, Ring, Wads, Perov Phase B, D, H and Egg







 \diamond Calculated in ppm wt H₂O (<1-1000')

Disclaimer about "water" in the mantle

Water" in anhydrous minerals = Hydrogen

Ol, Px, Gr, Fd, Apatite, Amph, Mica, Ring, Wads, Perov Phase B, D, H and Egg









\diamond Calculated in ppm wt H₂O (<1-1000')





Equivalent of several Earth's oceans in the mantle



Techniques

- ♦ FTIR \rightarrow H₂O content
 - \rightarrow speciation
 - \rightarrow location H mineral defects
- ↔ SIMS → H₂O content

 \rightarrow H isotopes

 \rightarrow Cl, F



Detection limits FTIR <0.5 ppm H2O SIMS \leq 2 ppm H2O

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Image Credit: Japan Times

Why is water important?

♦ Water lowers the solidus of mantle lithologies:

- Facilitates partial melting, lower T

♦ Influence on magmatism

- Magma composition
- Magma transfer and eruption style
 - ➔ origin of oceans & atmosphere

♦ Water and rheology:

- Presence of water in olivine makes it weaker
 - \rightarrow crucial for plate tectonics
- Melt circulation and eruption style

♦ Water and remote sensing of the deep Earth:

- Seismic properties: seismic wave attenuation & anisotropy
- Electrical conductivity
- Thermal conductivity

Gaetani & Grove, 1998; Green, 1973; Hirose & Kawamoto, 1995; Chopra & Paterson, 1984; Dixon et al., 2004; Drury, 1991; Hirth & Kohlstedt, 1996; Hirth et al., 2000; Justice et al., 1982; Karato, 1993; 2004; 2006; 2010; Mackwell et al., 1985; Mei & Kohlstedt, 2000; Walker et al., 2007; Jung & Karato, 2001; Demouchy et al 2012; Jones, Evans, Muller, Fullea, Pommier 1990-2015; Lizzaralde et al 1995; Sarafian et al. 2015; Udata et al 2003; Hofmeister 2004

Image Credit: J. Head Brown U., Stromboli

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Image Credit: Japan Times

H loss during xenolith ascent in host magma?

♦ H diffuses quickly through OI & Px

(e.g. Mackwell & Kohlstedt 1990; Ingrin et al 1995-2006, Stalder & Skogby 2003) 🔪

- ♦ Px have homogeneous water contents
- ♦ OI can record H loss

Figure stolen from AGU Chapman conference website (xenoliths added)

Olivine from Mexican mantle xenolith

Absorbance (cm⁻²)

H loss during xenolith ascent in host magma?

♦ H diffuses quickly through OI & Px at magmatic temperatures

(e.g. Mackwell & Kohlstedt 1990; Ingrin et al 1995, 1999, Stalder & Skogby 2003)

- But H not always lost because H diffusion is coupled with that of slower elements (e.g. Berry et al 2007, Novella et al 2015, Skogby, Stalder, Sundval 1989-2009)
- \diamond Check:

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Image Credit: Japan Times

♦ Cratons

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014, Novella et al 2015, Taylor et al 2016

Bell et al., 1992; Peslier et al 2012; Doucet et al 2014; Peslier et al unpublished

Xenoliths NOT representative of whole mantle lithosphere

- Over-representation of metasomatized water-rich peridotite near melt/fluid channels
- Overall mantle lithos dry (preserved in diamond mineral inclusions)
- > Water has a role in cratonic root long term longevity

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014, Novella et al 2015, Taylor et al 2016

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Image Credit: Japan Times

Image Credit: Demouchy & Bolfan-Casanova 2016

Off-craton olivine: most have lost H during xeno ascent

Peslier et al 2010- 2015-Unpub; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Doucet et al 2014, Novella et al 2015, Taylor et al 2016; Warren & Hauri 2014

Image Credit: Demouchy & Bolfan-Casanova 2016

- Off-craton olivine: most have lost H during xeno ascent
- Max [H2O] in olivine limited by solubility?
- Max [H2O] in olivine depends on local melt/fluid water activity
- > Mantle lithosphere is unsaturated in water

Peslier et al 2010- 2015-Unpub; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Doucet et al 2014, Novella et al 2015, Taylor et al 2016; Warren & Hauri 2014; Férot & Bolfan-Casanova 2012, Tenner, Ardia et al 2012; Yang et al 2014-15; Demouchy & Bolfan-Casanoca 2016

Image Credit: Demouchy & Bolfan-Casanova 2016

- Similar [H₂O] for all tectonic settings
- Subduction peridotite not more wet (except CO Plateau)

Skogby & Co 2006-2015; Peslier & Co 2002- 2015-Unpub; Demouchy & Co 2006-2016; Xia & Co 2004-2015 ; Stalder et al 2005; Schmädicke-Göse 2009-2015; Warren & Hauri 2014

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Image Credit: Japan Times

Water in the oceanic mantle

- Water in melt inclusions & undegassed glasses
- Oceanic island basalts (OIB) have more water than MORB
- Oceanic basalts: water & Ce similarly incompatible

McDonough & Sun 1995; Palme & O'Neill 2007; Dixon et al 2001-08; Salters & Stracke 2004; Cabral et al. 2014; Workman et al. 2004-06; Kendrick et al. 2014-15; Jackson et al. 2007, 2015; Métrich et al. 2014; Seaman et al. 2004

Water in the oceanic mantle

> Oceanic basalts: water & Ce similarly incompatible

- Oceanic peridotites: water & Ce decoupled
- Large scale re-equilibration of water?
 Partition coefficient problem?

McDonough & Sun 1995; Palme & O'Neill 2007; Dixon et al 2001-08; Salters & Stracke 2004; Cabral et al. 2014; Workman et al. 2004-06; Kendrick et al. 2014-15; Wallace 20024; Jackson et al. 2007, 2015; Métrich et al. 2014; Seaman et al. 2004; Peslier & Bizimis, 2015; Bizimis & Peslier 2015; Warren & Hauri, 2014; Peslier et al unpublished

Image Credit: Demouchy & Bolfan-Casanova 2016

Plume-lithosphere interaction

Image Credit: Demouchy & Bolfan-Casanova 2016

- Similar [H₂O] for all tectonic settings
- Subduction peridotite not more wet (except CO Plateau)
- Plume interaction (Hawaii, Ontong-Java, Tanz craton) lithosphere not more wet

Skogby & Co 2006-2015; Peslier & Co 2002- 2015-Unpub; Demouchy & Co 2006-2016; Xia & Co 2004-2015 ; Stalder et al 2005; Schmädicke-Göse 2009-2015; Warren & Hauri 2014; Peslier & Bizimis 2015; Bizimis & Peslier 2015; Hui et al 2015

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Image Credit: Japan Times

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REVIEW ARTICLE

Diamonds and water in the deep Earth: a new scenario

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ABSTRACT

Earth is a water planet, but how much water exists on and in the Earth? Is the water limited to the Earth's surface and limited depths of our planet (molecular water of the hydrosphere), or do deep reservoirs of hydrogen and oxygen really exist as proposed in recent works but not yet proven? Due to the importance of H₂O for life and geological processes on the Earth, these questions are among the most significant in all of the Earth sciences. Water must be present in the deep Earth as plate tectonics could not work without water as a major driving force that lowers both viscosity and density of the solid mineral phases of the interior and controls the onset of melting. On subduction, water is returned to the hydrosphere first by dewatering of hydrous phases and second by melting and arc magmatism in and above the subducting slab. The mantle is composed of oxygen minerals, and the extent to which hydrogen is dissolved in them constitutes the true reservoir of the planet's water. Are 'deep water and diamonds' intimately related as indicated in the title of the present article? What is the connection between these two important terrestrial materials? The necessity to review this issue arises from the recent discovery of a strongly hydrous ringwoodite in a Brazilian diamond. As ringwoodite constitutes 60% or more of the lower part of the transition zone, between 525 and 660 km depth, this could correspond to a huge amount of water in this region, comparable or greater in mass to all of Earth's hydrosphere. If the water found in this ringwoodite is representative of the water concentrations of the transition zone, then estimates of Earth's total water reservoir are in need of major revision. This work is an attempt at such a revision.

Whole Earth geohydrologic cycle, from the clouds to the core: The distribution of water in the dynamic Earth system

The Geological Society of America Special Paper 500

2013

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ARTICLE HISTORY

KEYWORDS diamond; water; Earth; reservoir; mantle

Transition zone: probably wet

- High [H2O] in ringwoodite, wadsleyite and majorite from exp (Inoue, Smyth, Ohtani, Demouchy, Bolfan-Casanova)
- Phase Egg (AlSiO₃(OH)) & hydrous ringwoodite as diamond inclusions (*Wirth et al 2007; Pearson et al 2014*)
- Consistent with electrical conductivity & seismic tomography (Karato 2001, Yoshino 2010, Koyama et al 2006)

Image Credit: Nestola & Smyth 2015, Kathy Maher

- ♦ TZ: uncertainty on [H2O] of ringwoodite, wadsleyite and majorite
 - Phase Egg & hydrous ringwoodite as diamond inclusions (*Wirth et al 07, Pearson et al 14*)
- ♦ Lower mantle: uncertainty on [H2O] of bridgmanite (Meade, Bolfan-Casanova, Litasov, Murakami...)
- \diamond Back to problem of storage capacity (exp) vs actual water content
- \diamond Uncertainty on amount of hydrous phases (B, Egg, D, H, δ-AlOOH) (Ohtani, Ghosh, Walter)

Rubie et al 2015 From accretion models

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Image Credit: Japan Times

Water deep cycle in the Earth

Fluxes in 10¹¹ kg/year

Oceans: 1.4x10²¹ kg 10^{11} kg = 7 ppb oceans

Image Credit: Hirschmann 2006

Image Credit: Hirschmann 2006

Water deep cycle in the Earth

Image Credit: Hirschmann 2006

Water deep cycle in the Earth

Image Credit: Hirschmann 2006

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Image Credit: Japan Times

Water in the inner solar system

Albarède 2013; Marty 2012; Hauri et al. 2011; McCubbin et al. 2010; Usui et al. 2012; Sarafian et al. 2013-2014; Hui et al 2013; Peslier et al. 2015 ; Rubie 2015

CC

CC

CC

CC

Conclusions

♦ Mantle lithosphere is undersaturated in water

- ♦ Too high [H2O] for degree of melting
- ♦ Decoupled from other incompatible trace elements (Ce)
- \diamond Mostly controlled by metasomatism in continental lithosphere
- ♦ All tectonic settings have similar [H2O]
- \diamond The transition zone is rich in water
- Water content of lower mantle and core?
- ♦ Water deep cycle
 - \diamond ~ 1 mass ocean re-added to mantle over 3 Ga
 - \diamond Residence time of water in Earth: ~ 1 Ga

Tartèse et al. 2013; Greenwood et al. 2011, Saal et al., 2013; Albarède et al., 2009, 2013; Sarafian et al., 2013-2014; Hui et al., 2015

Water deep cycle in the Earth

Image Credit: Hirschmann 2006

Peacock 1990, Wallmann 2001; Hilton et al 2002; Jarrad 2003; Bercovici & Karato 2003; Wallace 2005, Rüpke et al 2006, van Keken et al 2011, Bodnar et al 2013; Parai & Mukhopadhyay 2012

Noachian/Hesperian Mars

Amazonian Mars

McSween 2015

Parai & Mukhopadhyay 2012