

Please ensure that your abstract fits into one column on one page and complies with the *Instructions to Authors* available from the Abstract Submission web page.

An Amoeboid Olivine Aggregate in LEW 85300

KOMATSU M.^{1,2*}, YAMAGUCHI A.³, FAGAN T.J.²,
ZOLENSKY M.E.⁴, SHIRAI N.⁵, MIKOUCHI T.⁶

¹SOKENDAI, The Graduate University for Advanced Studies, Japan. (*correspondence:komatsu_mutsumi@soken.ac.jp)

²National Institute for Polar Research, Japan.

³Dept. of Earth Sciences, Waseda University, Japan.

⁴ARES, NASA Johnson Space Center, USA.

⁵Dept. of Chemistry, Tokyo Metropolitan University, Japan.

⁶Dept. of Earth and Planetary Science, University of Tokyo.

Introduction: Amoeboid Olivine aggregates (AOAs) are irregularly shaped objects commonly observed in carbonaceous chondrites. Because they are composed of fine-grained olivine and Ca-Al-rich minerals, they are sensitive indicators for nebular process and parent body alteration of their parent bodies [e.g., 1].

Recently an AOA was found in a carbonaceous clast in polymict eucrite LEW 85300 [2]. The bulk major element composition of the clast matrix in LEW 85300 suggests a relation to CM, CO and CV chondrites, whereas bulk clast trace and major element compositions do not match any carbonaceous chondrite, suggesting they have a unique origin [3]. Here we characterize the mineralogy of AOA in LEW 85300 and discuss the origin of the carbonaceous clasts.

Results and Discussion: The AOA is located in an impact melt vein. Half of the aggregate shows recrystallization textures (euhedral pyroxene and molten metal/FeS) due to impact melting, but the remaining part preserves the original texture. The AOA is composed of olivine, FeS and Mg,Al-phyllsilicate. Individual olivine grains measure 1-8 μm , with Fe-rich rims, probably due to impact heating.

Olivines in the AOA are highly forsteritic (Fo_{95-99}), indicating that the AOA escaped thermal metamorphism [4]. Although no LIME (Low-Fe, Mn-Enriched) olivine is observed, forsterite composition and the coexistence of Mg,Al-phyllsilicate suggest that the AOA is similar to those in the Bali-type oxidized CV (CV_{oxB}) and CR chondrites. However, it should be noted that fayalitic olivine, which commonly occurs in CV_{oxB} AOA, is not observed in this AOA. Also, the smaller grain size ($<8 \mu\text{m}$) of olivine suggests they may be related to CM or CO chondrites. Therefore, we cannot exclude the possibility that the AOA originated from a unique carbonaceous chondrite as suggested by [3].

REFERENCES: [1] Krot A. N. et al. (2004) *Chem. Erde*, 64, 185-282. [2] Shirai N. et al. (2016) *EPSL*, 437, 57-65. [3]

**This abstract is too long to be accepted for publication.
Please revise it so that it fits into the column on one
page.**

Zolensky M.E. et al. (1992) *MaPS*, 27,596-604. [4] Komatsu
M. et al. (2015) *MaPS*, 50, 1271–1294.