

## **5th International Conference on Tethers in Space**

University of Michigan

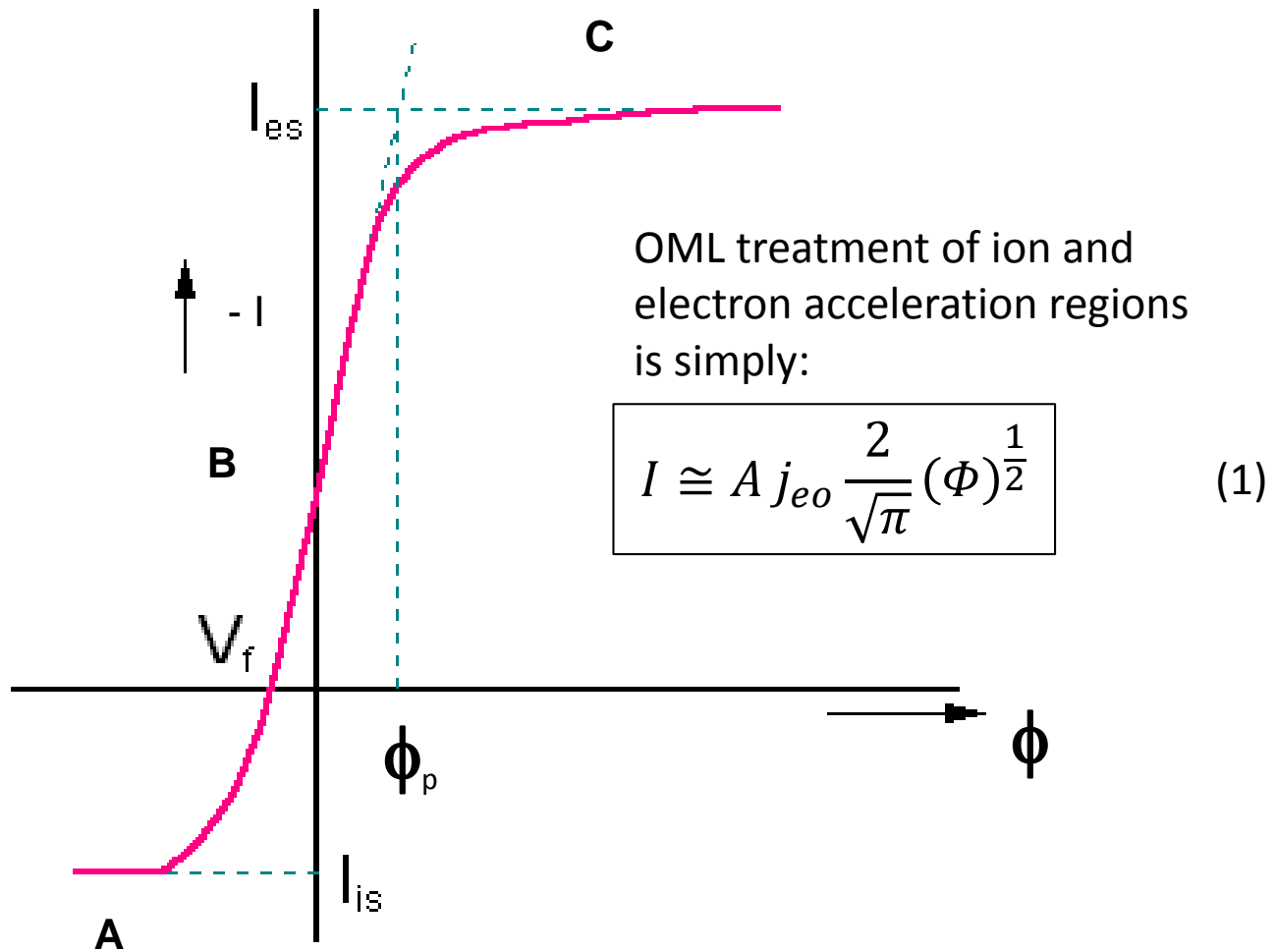
May 24-26, 2016

# **Re-Thinking the Use of the OML Model in Electric-Sail Development**

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# **General Conditions on Application of the OML Model**

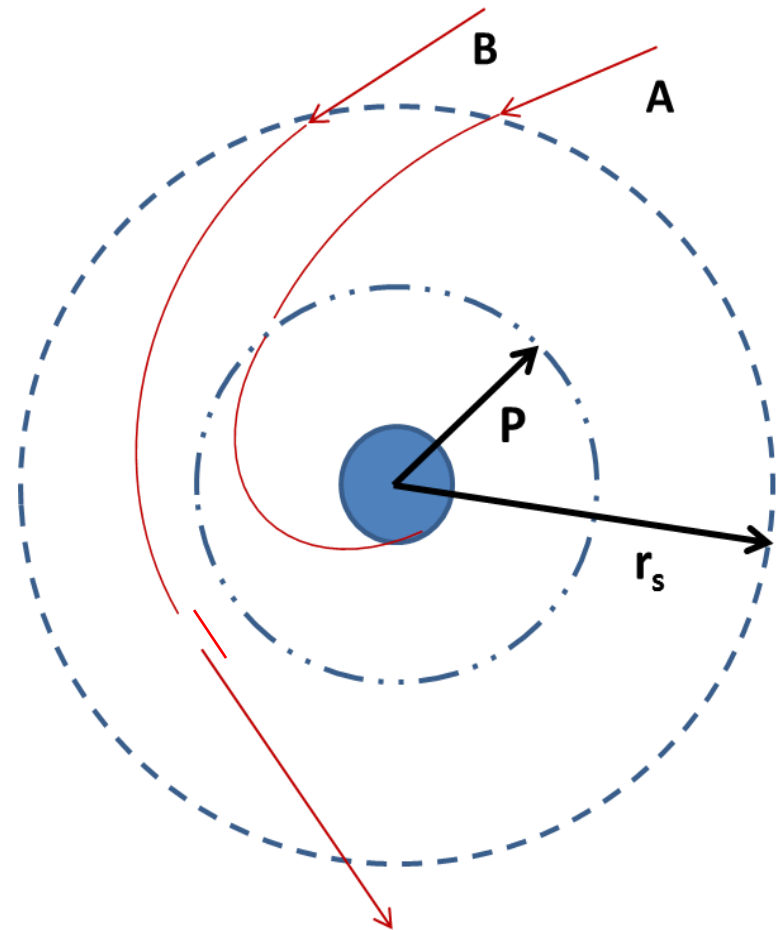


**Langmuir Probe i-v Characteristic**

# OML Assumed Electron Trajectories and Sheath E-Field Distribution

## Assumptions:

- (1) There exists an impact parameter,  $P$ , such that:
  - All particles with trajectories for which  $r_{\min} > P$  are lost.
  - All particles for which  $r_{\min} \leq P$  are collected.
- (2) Electric field within the sheath is symmetrical and undistorted.



# OML Derivation

$$f(u, v) = \left( \frac{m}{2\pi kT} \right) \exp \left( - \frac{m(u^2 + v^2)}{2kT} \right)$$

$$j_{eo} = \frac{en}{4} \left( \frac{8kT}{\pi m} \right)^{\frac{1}{2}}.$$

$$P = \left( \frac{r_s}{r_w} \right) \operatorname{erf} \left( Z \frac{1}{2} \right) e^{\Phi} \left[ 1 - \operatorname{erf} \left( \Phi + Z \right) \frac{1}{2} \right]$$

where  $\Phi = \left( \frac{e\phi_w}{kT} \right)$ , and  $Z = \left( \frac{r_s^2}{r_s^2 - r_w^2} \right) \Phi$ ,

For  $r_s \gg r_w$  and  $\Phi \gg 1$ ,

$$P \cong \frac{2}{\sqrt{\pi}} (\Phi + 1)^{\frac{1}{2}} \approx \frac{2}{\sqrt{\pi}} \Phi^{1/2}$$

(2)

## OML Derivation (Cont'd)

$$I = A_w j_{eo} P$$

Inserting Eqn. (2) for P,

$$I \cong A_w j_{eo} \frac{2}{\sqrt{\pi}} (\Phi)^{\frac{1}{2}}$$

$$\boxed{I = \frac{en}{\pi} A_w \left( \frac{2e\phi_w}{m} \right)^{\frac{1}{2}}} \quad (3)$$

where  $A_w = 2\pi r_w L$

Eqn. (3) is identical to the standard OML representation given in Eqn. (1). Therefore, all assumptions and approximations used in arriving at Eqn. (3) apply to the OML model.

## Assumptions and Approximations Made in the LMS Treatment

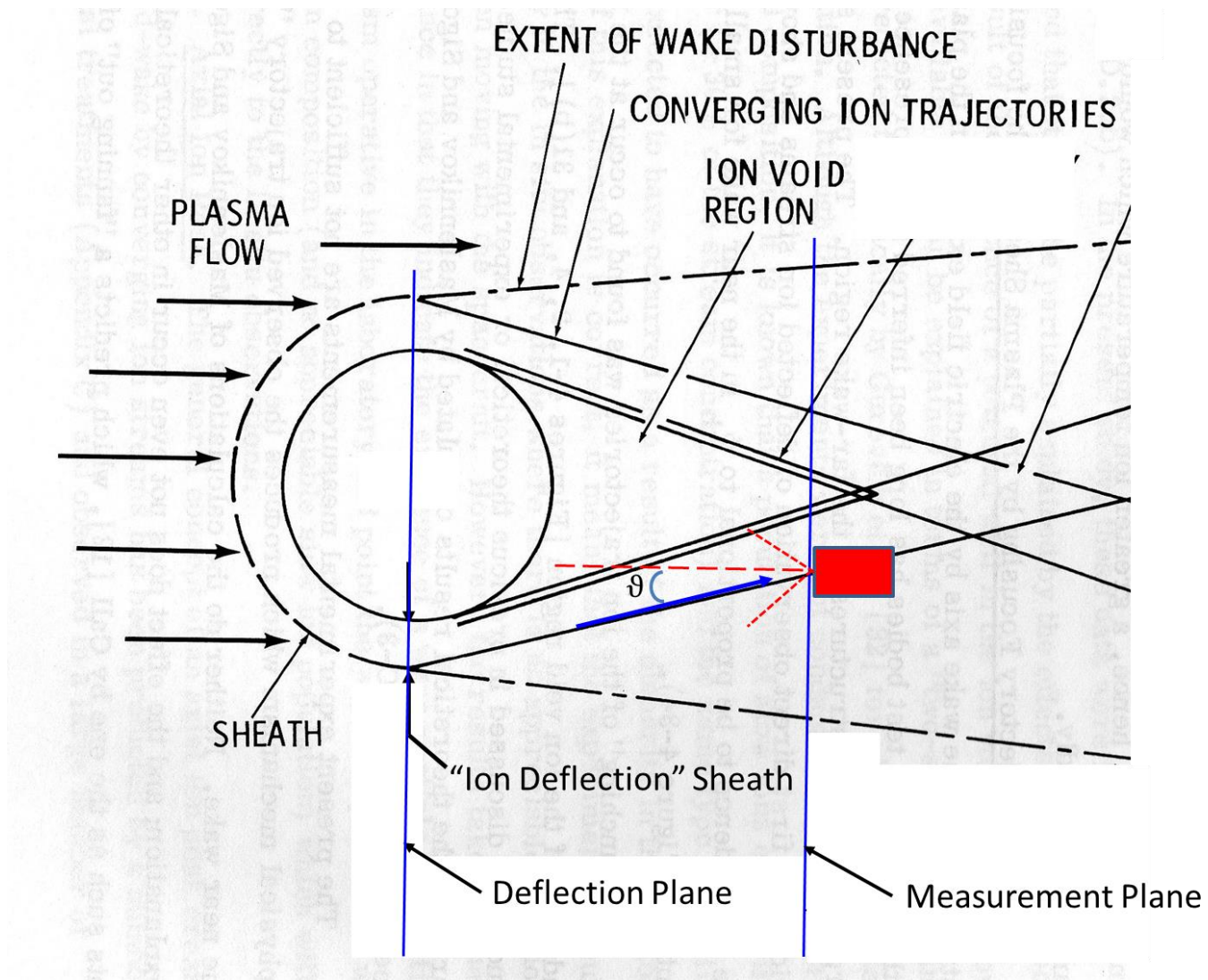
- (1) Small body ( $r_w \ll \lambda_D$ ).
- (2) Quasi-static conditions ( $V_{\text{drift}} \ll v_{\text{th}}$ ).
- (3) Maxwellian distribution of the collected plasma component at the sheath boundary.
- (4) Total absorption of the collected species that contact the electrode—particles are either collected or lost.
- (5) Cylindrically symmetric E-field.
- (6) No collisional effects on particle trajectories. Electron trajectories through the sheath region are determined totally by their initial velocity at the sheath boundary and the sheath electric field.

# Potential Effects of the Assumptions and Approximations

- (1) Quasi-static conditions. (The Solar Wind has a very large drift velocity and on TSS-1R, Parker-Murphy erred by 200-300%—most probably because of orbital velocity effects. )
- (2) Particles are either collected or. (Trapped particles occur in lab plasma sheaths and in the TSS-1R Satellite HV sheath.)
- (3) Symmetric E-field. (Proton deflection and/or particle collisions or trapping can distort the field.)
- (4) No collisional effects on collected particle trajectories.  
This implies:
  - No photo-emission from surfaces.
  - No recombination of charged particles.
  - No charge exchange collisions.
  - No trapped particles.

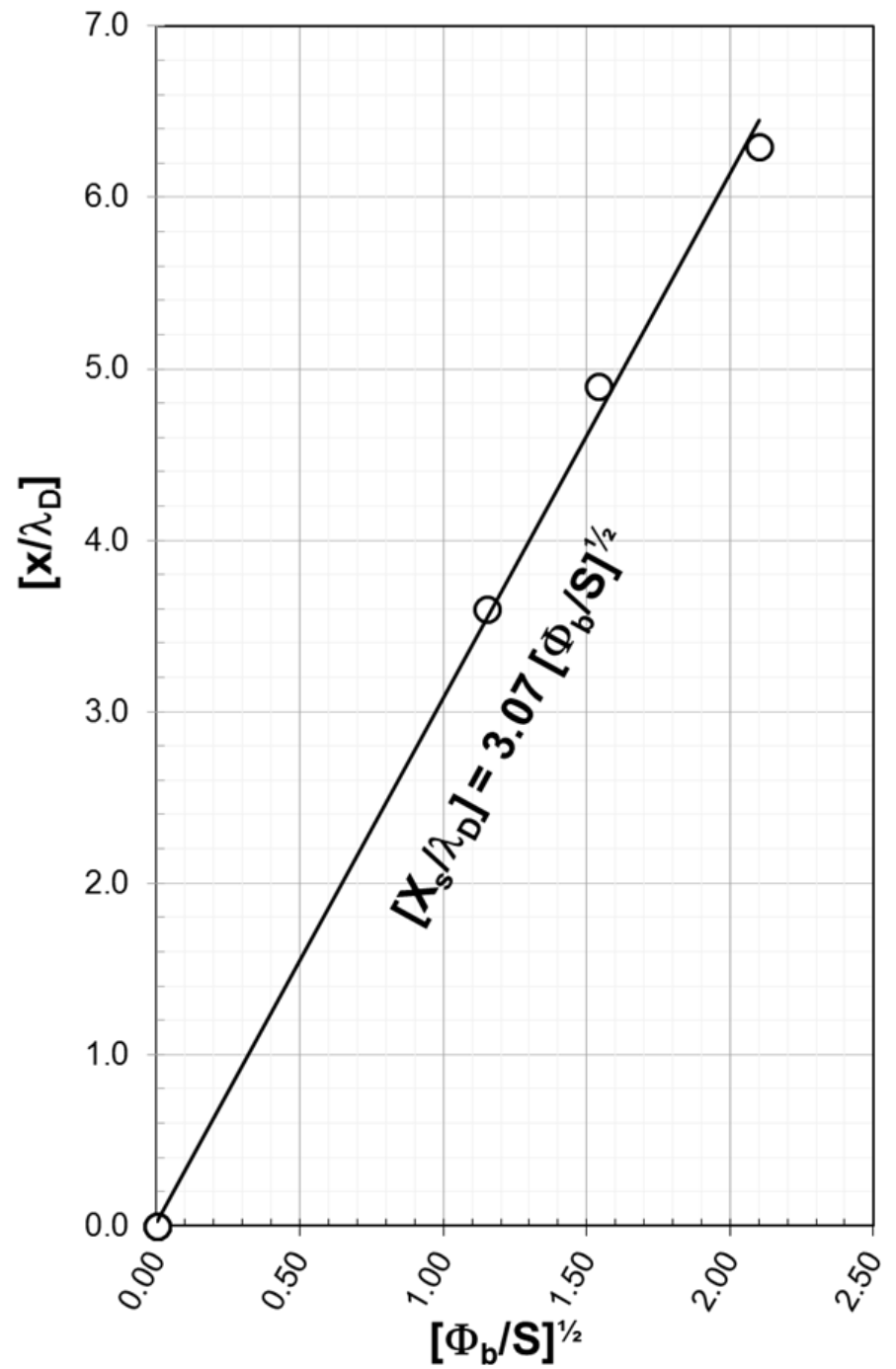


# **Laboratory Observations of Body-Plasma Interactions**

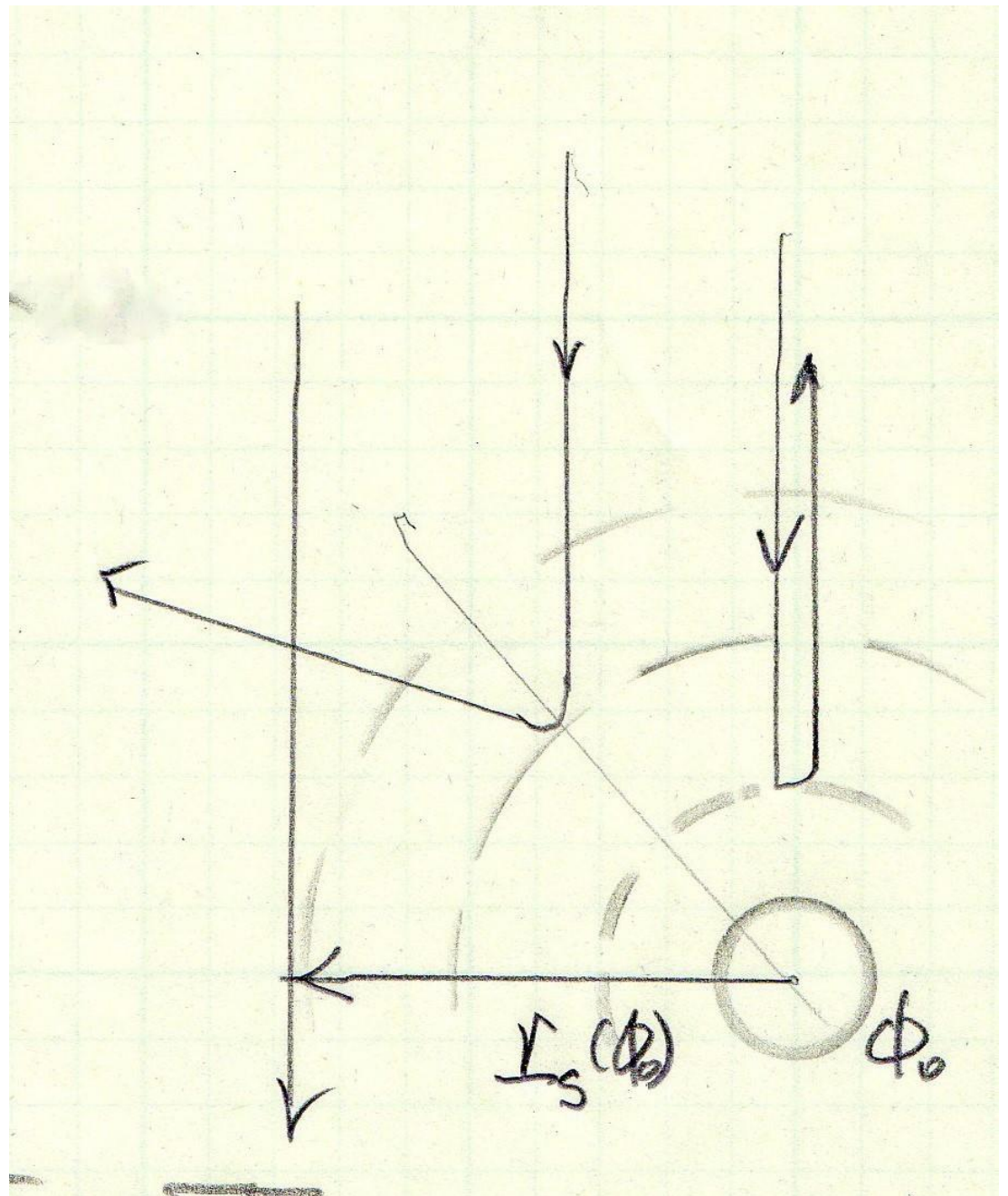


**Regions of Disturbed Plasma Flow**

# Effective Sheath Width for Detectable Ion Deflection



# Max Extent of Proton Deflection.



## Figure 4. Specular Proton Deflection Calculation

At any deflection point,  
 $r^*(\vartheta, \phi)$ ,

$$e\phi(r) = \frac{1}{2}m_p V_0^2 \\ = \frac{1}{2}m_p (V_0 \cos\vartheta)^2$$

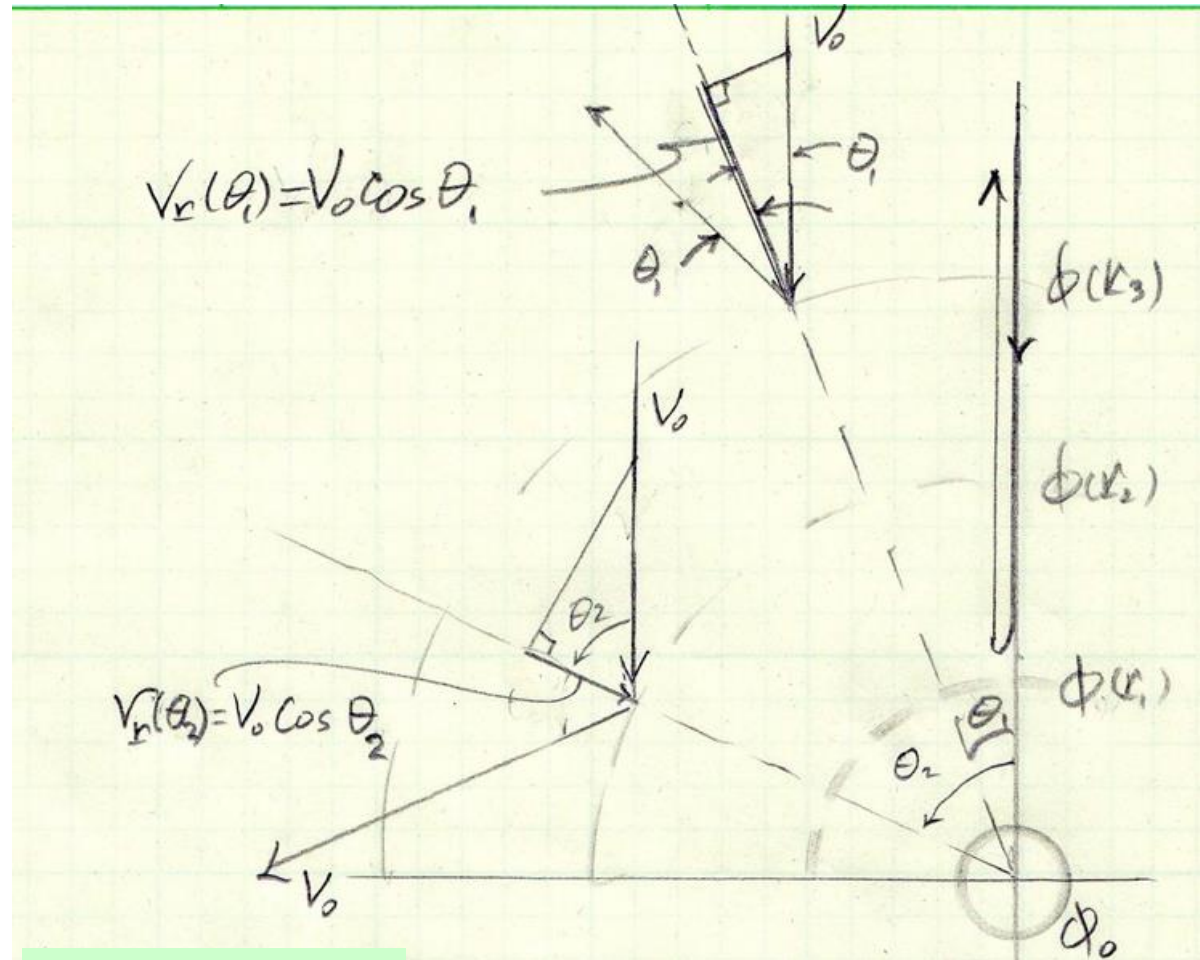
and

$$\phi(r) = \phi_0 \ln(r_s/r) / \ln(r_s/r_w), \quad \text{where } \phi(r=0) = 0$$

Then

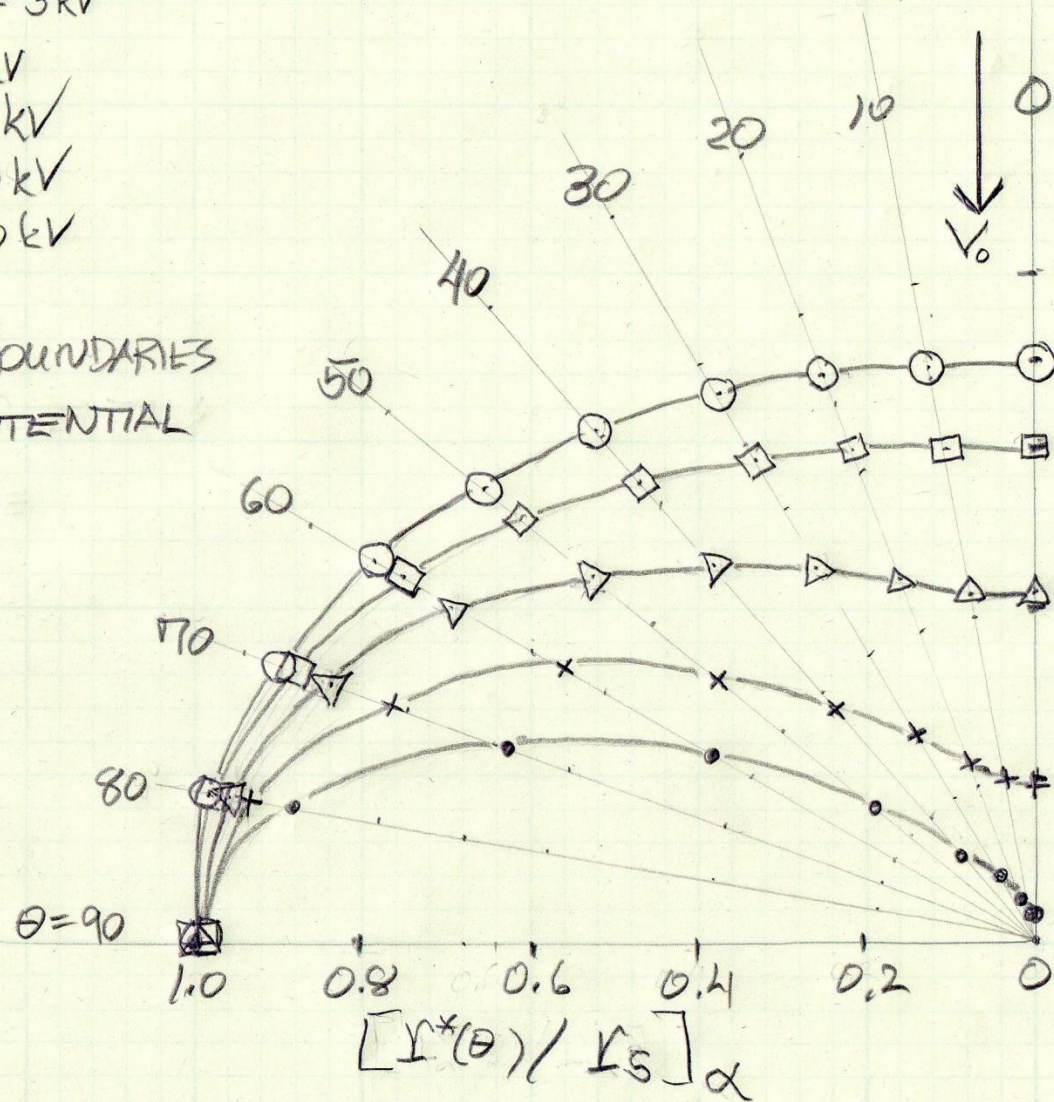
$$r^*(\vartheta) = r_s (r_s/r_w)^{\beta \cos^2 \vartheta}$$

(4)



- -  $\phi_0 = 3 \text{ kV}$
- + -  $6 \text{ kV}$
- $\Delta$  -  $12 \text{ kV}$
- $\square$  -  $20 \text{ kV}$
- $\circ$  -  $30 \text{ kV}$

REFLECTION BOUNDARIES  
VS BIAS POTENTIAL  
AT  $1 R_E$



**Reflection Boundaries Within Sheath—at  $r^*(\vartheta, \phi)$  given by Eqn. (4)—for Various Proton Drift Energies**

# Thrust Calculations Based On Laboratory Experimental Results

Calculation of Momentum Exchange

$$f = n_o v_o (M_{in} - M_{out})$$

$$M_{in} = (m_p v_o)$$

$$M_{out} = m_p v_o \int_0^{\pi/2} \cos(2\vartheta) d\vartheta = 0$$

$$F = 2r_s(\phi_b)f = 2r_s n_o m_p v_o^2$$

$$F = 0.87 \mu\text{N/m}$$

F is thrust generated per m of wire for nominal solar wind ( $V_o = 400$  km/s,  $n_o = 7 \times 10^6$  m<sup>-3</sup>, and  $T_e = 1.5 \times 10^5$  °k).

Schematic of the complex array of physical effects observed in the near plasma environment of the TSS satellite.

