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Analysis of sustained long-period activity at Etna Volcano, Italy

Gilberto Saccorotti ^{a,*}, Ivan Lokmer ^b, Christopher J. Bean ^b,
Giuseppe Di Grazia ^c, Domenico Patanè ^c

^a *Istituto Nazionale di Geofisica e Vulcanologia Sez. di Napoli, Osservatorio Vesuviano. Via Diocleziano, 328- 80124 Napoli, Italy*

^b *School of Geological Sciences, University College Dublin. Belfield, Dublin 4, Ireland, United Kingdom*

^c *Istituto Nazionale di Geofisica e Vulcanologia, Sez. di Catania. Piazza Roma, 2- 95123 Catania, Italy*

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Abstract

Following the installation of a broadband network on Mt. Etna, sustained Long-Period (LP) activity was recorded accompanying a period of total quiescence and the subsequent onset of the 2004–2005 effusive episode. From *c.* about 56000 events detected by an automatic classification procedure, we analyse a subset of about 3000 signals spanning the December 17th, 2003–September 25th, 2004, time interval. LP spectra are characterised by several, unevenly-spaced narrow peaks spanning the 0.5–10 Hz frequency band. These peaks are common to all the recording sites of the network, and different from those associated with tremor signals. Throughout the analysed time interval, LP spectra and waveforms maintain significant similarity, thus indicating the involvement of a non-destructive source process that we interpret in terms of the resonance of a fluid-filled buried cavity. Polarisation analysis indicates radiation from a non-isotropic source involving large amounts of shear. Concurrently with LP signals, recordings from the summit station also depict Very-Long-Period (VLP) pulses whose rectilinear motion points to a region located beneath the summit craters at depths ranging between 800 and 1100 m beneath the surface. Based on a refined repicking of similar waveforms, we obtain robust locations for a selected subset of the most energetic LP events from probabilistic inversion of travel-times calculated for a 3D heterogeneous structure. LP sources cluster in a narrow volume located beneath the summit craters, and extending to a maximum depth of ≈ 800 m beneath the surface. No causal relationships are observed between LP, VLP and tremor activities and the onset of the 2004–2005 lava effusions, thus indicating that magmatic overpressure played a limited role in triggering this eruption. These data represent the very first observation of LP and VLP activity at Etna during non-eruptive periods, and open the way to the quantitative modelling of the geometry and dynamics of the shallow plumbing system.

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1. Introduction

Long-Period (LP) and Very-Long-Period (VLP) events are seismic signals which have been documented at many volcanoes throughout the world. In most cases this seismicity occurs in association with surface eruptive activity, such as at Stromboli, Italy (Neuberg et al., 1994; Chouet

* Corresponding author. Now at Istituto Nazionale di Geofisica e Vulcanologia- Sez. di Pisa. Via U. della Faggiola, 32, 56126 Pisa, Italy. Tel.: +39 050 8311960; fax: +39 050 8311942.

E-mail address: saccorotti@pi.ingv.it (G. Saccorotti).

et al., 1997, 1999), Galeras, Colombia (Gil Cruz and Chouet, 1997), Erebus, Antarctica (Rowe et al., 1998), Soufriere Hills, Montserrat, (Neuberg et al., 1998), Usu, Japan (Matsubara and Yomogida, 2004), Popocatepetl, Mexico (Arciniega-Ceballos et al., 2000). Other observations account instead for LP-VLP signals occurring without any evident correlation with phenomena which are visible at the surface, such as at Lascar, Chile (Hellweg, 1999), Shishaldin, Alaska (Petersen et al., 2006), Tongariro, New Zealand (Hagerty and Benites, 2003), Deception, Antarctica (Ibañez et al., 2000), Kusatsu-Shirane, Japan (Kumagai et al., 2002a). Although different models have been invoked to explain the source mechanism of these events (e.g., Crosson and Bame, 1985; Chouet, 1988; Fujita et al., 1995; Jousset et al., 2003, 2004), all involve the resonance and/or transport of fluids in magmatic and hydrothermal systems. The quantitative assessment of these sources is thus critically important for our ability to model volcanic systems, and to successfully forecast eruptive activity (e.g., Chouet et al., 1994).

With a single exception (Falsaperla et al., 2002), Etna Volcano lacks significant records of long-period seismicity, mostly owing to the reduced bandwidth of the surveillance network, which was traditionally based upon 1 Hz seismometers. Volcanic signals, however, often exhibit a broadband behaviour (see references in the introduction). The undesired effects of the path and recording site are reduced when analysing Long and

Very-Long-Period signals, thus permitting a direct view of the source mechanisms and mass transport phenomena. These considerations suggested the need to extend the frequency bandwidth of Etna's monitoring system, thus leading to the installation of a broadband network as a part of the permanent monitoring systems managed by the Italian National Institute for Geophysics and Volcanology (INGV hereinafter). Following this installation, sustained LP activity was observed throughout the year in 2004, accompanying a period of complete quiescence and the subsequent onset of the 'silent' 2004–2005 eruption (Burton et al., 2005).

In this paper we analyse the wavefield properties and temporal evolution of these signals, in turn retrieving precise locations for a selected subset of the most energetic events. Although our results indicate the absence of any obvious relationship between the 2004 LP swarms and the subsequent resumption of lava effusion, this study makes a significant contribution to the precise mapping of seismo-volcanic sources at Etna, thus setting the stage for the subsequent modelling efforts aimed at investigating the geometry and dynamics of the shallow plumbing system.

2. Instruments and data

Deployed by late November, 2003, INGV's broadband network at Etna consists of 8 Nanometrics TRILIUM seismometers with flat amplitude response within the 40–

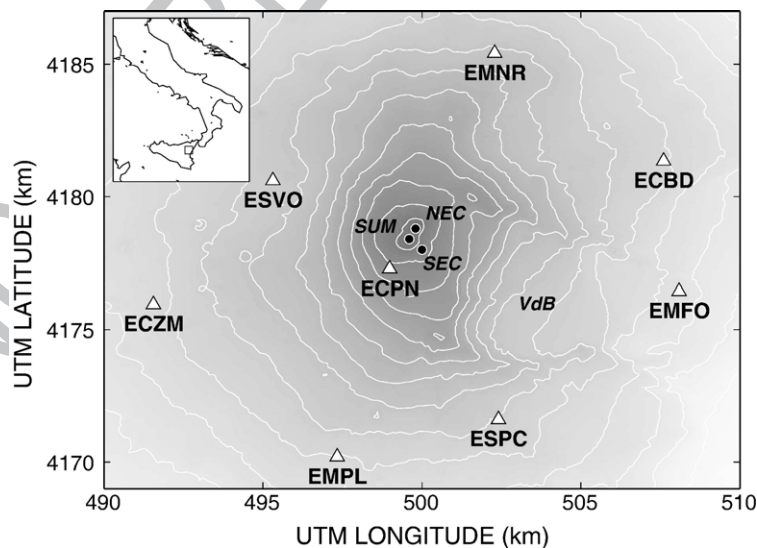


Fig. 1. Map of the summit of Etna Volcano with the location of the broadband stations of the INGV permanent monitoring network (triangles). Solid, black circles mark the position of the main craters: SUM- Summit Craters area; NEC — North-East Crater; SEC — South-East Crater. VdB indicates the Valle del Bove volcano-tectonic depression. The inset shows the location of Etna with respect to Italy.

0.01 s period range, installed at elevations between 1100 and 3000 m a.s.l. and distances from the summit craters between 9 and 1.5 km (Fig. 1). Data from the remote sites are transmitted via mixed radio-satellite links to the ING V monitoring center in Catania, where they are stored with a sampling interval of 0.01 s over consecutive, 2-minute-long digital archives. These new broadband observations resulted in the detection of previously unobserved LP signals, consisting of spindle-shaped, weak pulses rapidly attenuating with distance. Most of their energy is concentrated within a narrow frequency interval spanning 0.5–10 Hz, which we attribute to a source effect based on its persistence to all the widely-spaced stations of the network (Fig. 2, top plot).

These events are superimposed on a background of continuous volcanic tremor, whose energy is concentrated within a frequency range similar to that observed for LP signals (Fig. 2, middle plot). The quantitative comparison between LP and tremor spectra must account for the filtering effects associated with both the propagation paths and the local geology beneath individual recording sites. In order to remove these unwanted contributions, we calculated network-averaged spectra for both LP and tremor signals (Fig. 2, bottom plot). Since the average station spacing is generally larger than the dominant wavelength of the signal, this procedure allows us to highlight the contributions from the primary source while removing those associated with the path and site effects. For the specific example shown in Fig. 2, LP spectra are dominated by 2 main spectral peaks at frequencies of ≈ 0.6 Hz and ≈ 3 Hz, while tremor spectra exhibit a single peak centered at a frequency of ≈ 1 Hz. This observation suggests that the continuous tremor and LP transients result from two distinct source processes.

Exploiting the different spectral signatures of LP signals with respect to those of either tremor or volcano-tectonic (VT) waveforms, we developed an automatic detection procedure based on the similarity between the spectra of triggered signals and a reference LP spectrum. Application of this procedure to data from summit station ECPN (see Fig. 1) resulted in the detection of about 56,000 LP signals spanning the November 2003–December 2004 time interval (Fig. 3a), with a peak in the maximum amplitude and daily rate between early January and late April, 2004 (Fig. 3b). The daily rate further peaked a few days before the beginning of the 2004–2005 eruption, but in association with signals much weaker than those observed during the previous months.

At the beginning of this study (early summer 2004), we had selected the data by thresholding at $155 \mu\text{V}$ the peak-to-peak amplitude of the raw vertical-component

velocity seismograms from station ECPN. This first selection includes 1510 events encompassing the December 17, 2003–May 4, 2004 time span. Following the onset of the 2004–2005 eruption, we used an amplitude threshold of $75 \mu\text{V}$ for extracting additional 1430 events spanning the May 10–September 25, 2004 time interval. This second catalog, however, includes only data from summit station ECPN, as the Signal-to-Noise-Ratio (SNR) at the remaining sites was definitively too low for the application of any reliable analysis. The complete data set analysed in this paper thus amounts to 2940 events. Data were organised in digital archives containing 2-minute-long recordings starting 60 s before the maximum peak-to-peak LP amplitude measured at station ECPN. Before the analysis, data were corrected for instrument response.

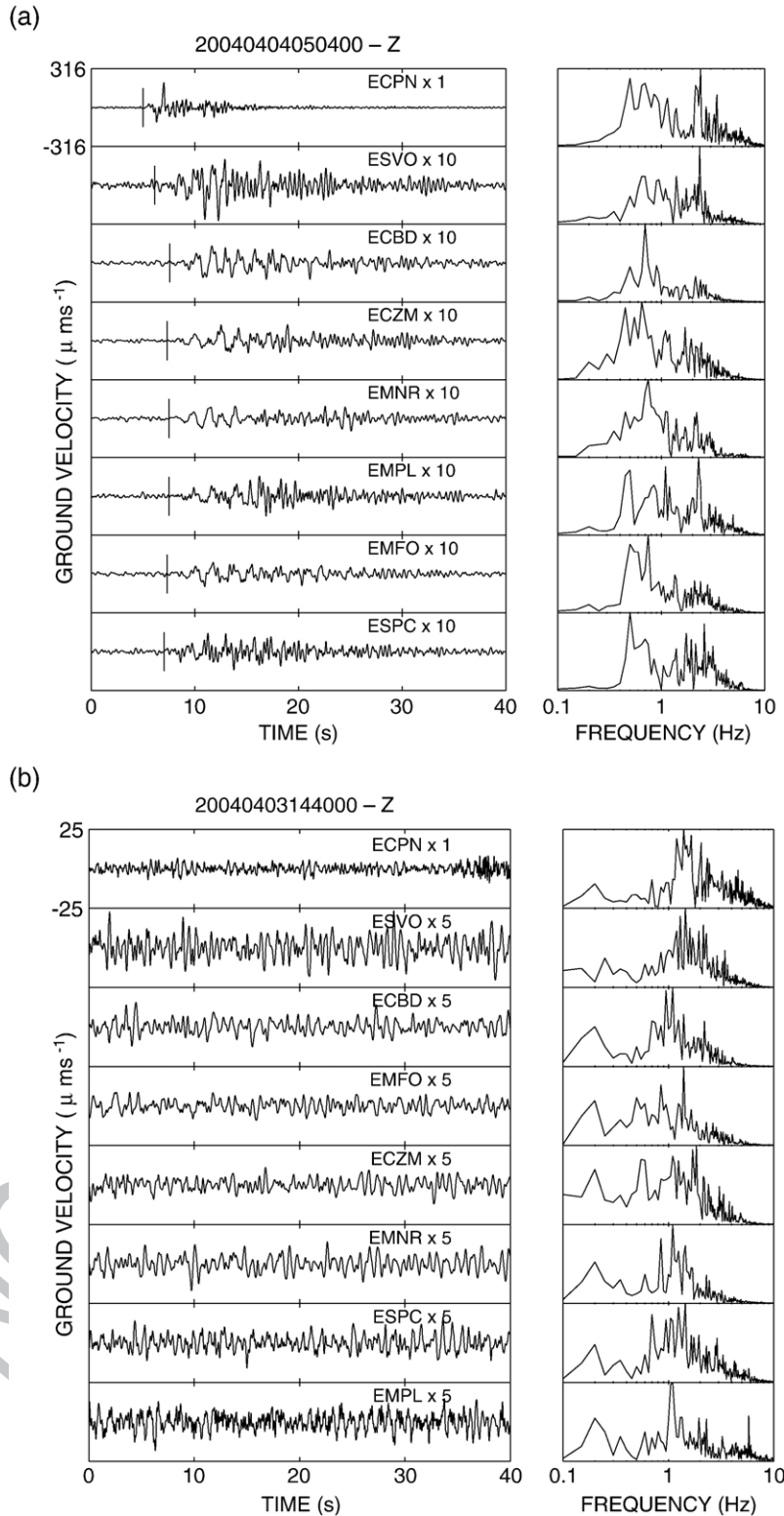
3. Data analysis

3.1. Spectral properties

In order to investigate the temporal evolution of LP signals, we conducted spectral analysis over the whole catalog derived from summit station ECPN, using a 5.12 s long window encompassing the maximum peak-to-peak amplitude of the vertical-component recordings. The frequency resolution df of spectral estimates thus corresponds to 0.19 Hz. As mentioned above, our LP catalog was generated by an automatic classification procedure based on the spectral similarity of triggered signals with respect to a reference LP spectrum. Therefore, we expect that all the different LP spectra will exhibit a high level of mutual similarity. Following this assumption, we averaged the LP spectra on a daily basis, in turn deriving a periodogram estimate (Press et al., 1992) of the power spectrum. These power spectra were then arranged in a time-frequency representation (spectrogram) extending over the whole period spanned by our catalog (Fig. 4a). Integration over frequency of individual power spectra allowed us to obtain an estimate of the daily energy associated with the LP catalog (Fig. 4a, top plot). An average view of LP spectral signature is eventually gained by stacking the different daily power spectra (Fig. 4a, right plot).

LP spectra are dominated by a main peak at a frequency of ≈ 0.6 Hz, and higher-frequency overtones at frequencies of $\approx 2, 3.2, 5, 6.5$ and 9 Hz. Although the relative weight of these peaks varies with time, most of them depict a marked persistency throughout the analysed sequence.

Starting in April, 2004, the dominant peak at ≈ 0.6 Hz broadens and appears to shift toward slightly



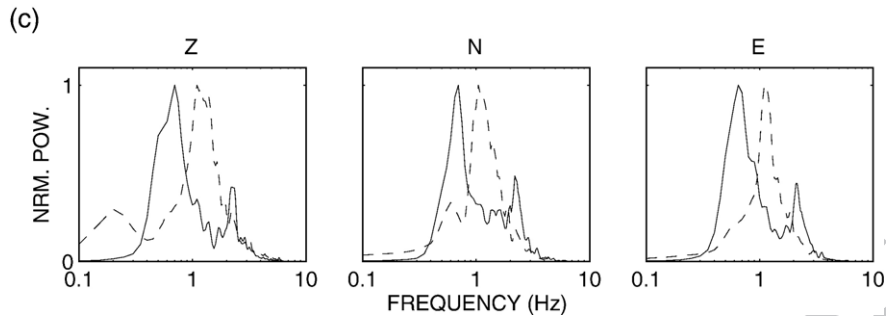


Fig. 2 (continued).

193 higher frequencies; this effect, however, is likely due to
 194 contamination of LP spectra by the tremor peak at
 195 frequency ≈ 1 Hz (see Fig. 4 2). Over the same period, in
 196 fact, LP energy decreases while the energy associated
 197 with the background gets relatively larger (Fig. 4).
 198 Moreover, it is difficult to assess such subtle frequency
 199 changes on the basis of our Fourier-based spectral
 200 measurements; the brief duration of LP signals con-
 201 strains us to perform spectral estimates over short time
 202 windows, thus necessarily leading to poor frequency
 203 resolution. Any future effort aimed at investigating the
 204 fine spectral structure of our LP data should therefore be
 205 based upon techniques, such as instantaneous frequency
 206 determinations or Autoregressive Modelling of the
 207 signal (Lesage et al., 2002), which allow for the efficient
 208 separation of closely-spaced spectral peaks even over
 209 short-duration time series.

210 Similar analyses were conducted over tremor data, by
 211 computing spectral estimates over 5.12-s-long windows
 212 of signal taken from the beginning of the LP recordings
 213 (Fig. 4b). Tremor spectra broadens over the 1–5 Hz
 214 frequency range, and their peaks are less marked than
 215 and different from those observed for LP spectra. No
 216 noticeable variations of either the energy or spectral
 217 content of tremor signal appear to herald the onset of the
 218 2004–2005 eruptive episode.

219 3.2. Waveform similarity

220 Although individual, raw LP signals depict quite
 221 different waveforms, once band-pass filtered around the
 222 main frequency peak they all share a common signature.

223 We quantified this observation by conducting correla-
 224 tion analyses for all the independent event pairs using 5-s-

long windows encompassing the maximum peak-to-peak
 amplitude of the 0.1–1 Hz band-pass-filtered, vertical-
 component seismograms from summit station ECPN. We
 then scanned this large correlation matrix in ascending
 chronological order and extracted the elements for which
 the correlations among all the possible permutations were
 greater than the arbitrary threshold of 0.8. In this
 procedure, we didn't allow any overlap between clusters;
 in other words, once an event was assigned to a given
 group, it was removed from the correlation matrix thus
 preventing it from being associated with any other cluster.
 At the end of the process we then discarded the small
 clusters constituted by less than 5 elements, and found that
 1759 out of the initial 2940 events grouped into four large
 families, whose temporal distributions seem to depict a
 weak evolutionary trend (Fig. 5a). Rather than the
 consecutive evolution of distinct sources, however, such
 a trend is probably an artifact of our sequential, exclusive
 clustering procedure. For each of these clusters, we used
 the inter-event delay times derived from correlation
 analyses to align seismograms, and computed stacked
 waveforms (e.g., Rowe et al., 2004). The stacked signals
 associated with the different clusters display a marked
 similarity (Fig. 5b), thus reinforcing the idea that
 individual groups are most likely representative of the
 same source. We also repeated the correlation and
 clustering procedures over the two horizontal components
 of ground velocity, the result of which support our
 findings for the vertical component case.

254 3.3. Polarisation analysis

255 Fig. 6 illustrates an example of particle motion
 256 trajectories observed at station ESVO in association

Fig. 2. Top: Vertical-component velocity seismograms for a LP event recorded by all the stations of the network, and corresponding normalised amplitude spectra calculated over a 15 s long window starting at the onset of the event (vertical lines on the seismograms). Traces are arranged in order of decreasing amplitude. Middle: Vertical-component velocity seismograms for a 40 s long section of volcanic tremor, and corresponding normalised amplitude spectra calculated over a 15 s long window. Bottom: Network-averaged normalized power spectra estimates for the three components of ground velocity associated with the LP and tremor signals shown in two panels above (continuous and dashed lines, respectively).

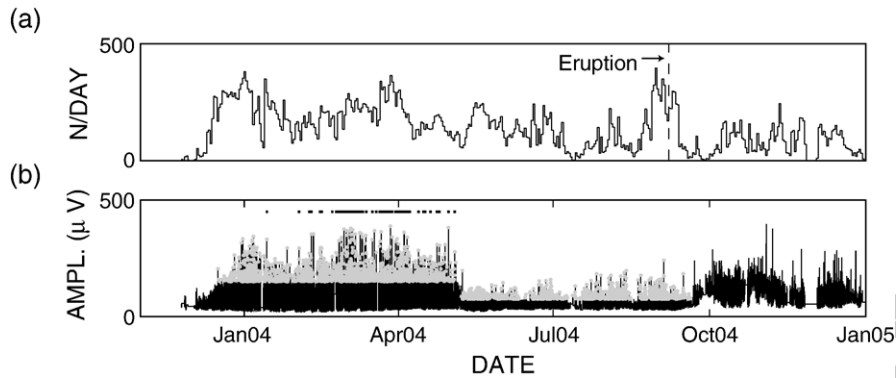


Fig. 3. Time evolution of LP activity from November 1st, 2003, through the end of 2004. (a) Daily count of events at station ECPN resulting from the automatic classification procedure. The dashed vertical line marks the onset of the 2004–2005 eruption. (b) Maximum peak-to-peak amplitude of the raw, vertical-component velocity seismograms from station ECPN. Gray dots indicate the events selected for our analysis after the application of two different amplitude thresholds. Black dots at the top of the plot mark the origin times of the events successfully located.

257 with the event's onset and the subsequent, largest amp-
 258 plitude phase. The signal initiates with a low-amplitude,
 259 quasi-horizontal rectilinear pulse; even accounting for
 260 the expected effect of topography on P-wave incidence
 261 angle, (e.g., Neuberg and Pointer, 2000), this phase is

consistent with a direct P-wave arrival from a source 262
 located at shallow depths beneath the summit craters. 263
 The subsequent, largest-amplitude phase depicts instead 264
 a transverse orientation, which is compatible with a SH- 265
 Love wave arrival. The delay time between this phase 266

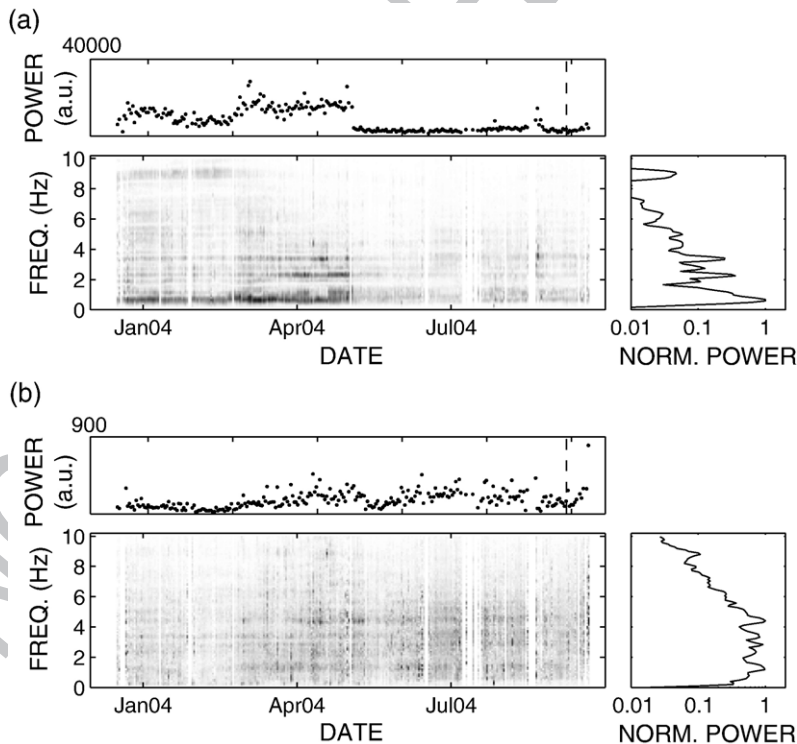


Fig. 4. (a) Spectral amplitude versus time and frequency (spectrogram) for vertical-component LP recordings at station ECPN. Each spectrogram's column represents the daily average of LP spectra calculated over 5.12 s long time windows encompassing the maximum peak-to-peak amplitude of the velocity seismograms. The top plot represents the temporal behaviour of the overall spectral power, obtained after integrating the spectrogram over frequency. The plot at the right depicts the stacked, normalised power spectrum obtained after time-integration of the spectrogram. (b) The same as in (a), but for tremor data preceding the onset of LP signals.

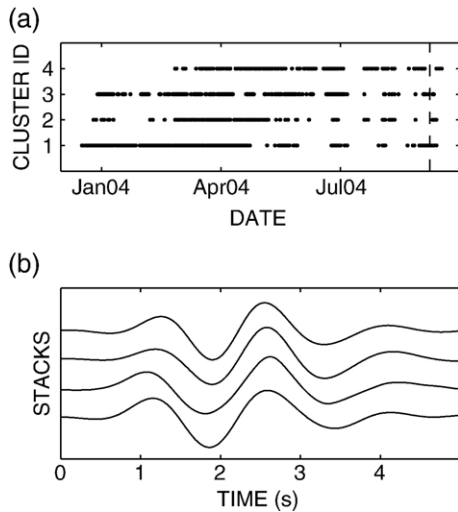


Fig. 5. (a) Temporal location of events pertaining to the four different clusters of similar events derived from correlation analysis. (b) Stacked waveforms obtained after aligning and summing vertical-component velocity seismograms pertaining to the four different clusters shown in (a).

267 and the P-wave onset is on the order of 5 s, thus much
 268 larger than the 1–2 s which would be expected for a
 269 direct S-wave radiated by a shallow source beneath the
 270 summit craters. For the specific example shown in
 271 Fig. 6, the origin of the large-amplitude transverse phase
 272 is thus most likely associated with a path or recording
 273 site effect. For the subset of the most energetic events
 274 used for location (see next section), we derived the
 275 polarisation attributes (azimuth and incidence angles,
 276 coefficient of rectilinearity) using the covariance-matrix

method of Kanasevich (1981). Covariance estimates 277
 were obtained over 2-s-long time windows sliding with 278
 0.1 s increments along the three-component velocity 279
 seismograms band-pass filtered over the 0.1–1 Hz 280
 frequency band. Fig. 7 illustrates the frequency distri- 281
 bution of particle motion azimuths associated with the 282
 onset of the signal and the subsequent largest amplitude 283
 of motion, here taken to correspond with the maximum 284
 of the largest eigenvalue of the 3-component covariance 285
 matrix. The resulting picture is quite complex: at most of 286
 the recording sites, the event's onset is dominated by 287
 radial motion, whose direction points to the summit 288
 craters. A notable exception is station ECPN (see 289
 Fig. 8), whose signals completely lack radial compo- 290
 nents. This observation may be easily interpreted when 291
 we consider that, for a source located at shallow depths 292
 beneath the summit craters, the time difference between 293
 the S- and P-wave arrivals would be much lower than 294
 the dominant period of the signal, thus hindering the 295
 possibility of P-wave discrimination on the basis of 296
 particle motion attributes. For this particular site, 297
 moreover, near-field effects are also expected to be 298
 significant. Station EMFO 22 also lacks clear radial 299
 onsets. A possible interpretation of this observation 300
 involves the combined effects of topography and 301
 structural complexities. That station is in fact located 302
 in close proximity to the NE margin of the Valle del 303
 Bove, as a consequence of which severe wave con- 304
 version and ray bending phenomena are expected (e.g., 305
 O'Brien and Bean, 2004). At several stations (e.g., 306
 ESVO, EMPL, ESPC) the polarisation azimuths asso- 307
 ciated with the largest amplitudes indicate the presence 308

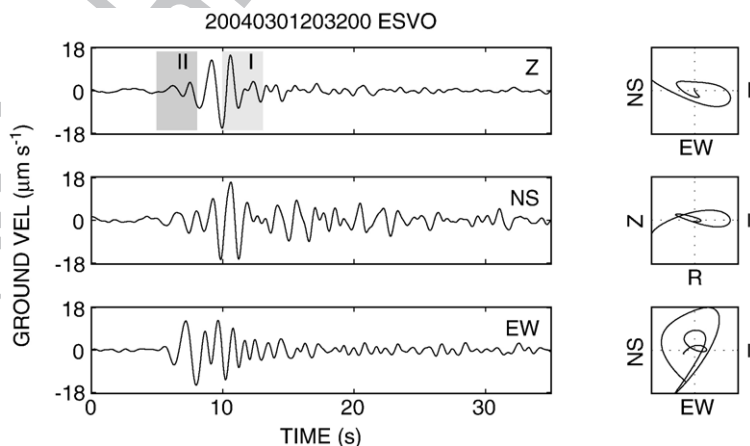


Fig. 6. Left: three-component velocity seismograms at station ESVO filtered over the 0.1–1 Hz frequency range. The two numbered gray strips mark the intervals selected for the particle motion analysis shown on the right. The onset of the LP signal (window I) is marked by a radial pulse, while the subsequent, largest-amplitude of motion (window II) is associated with shear energy.

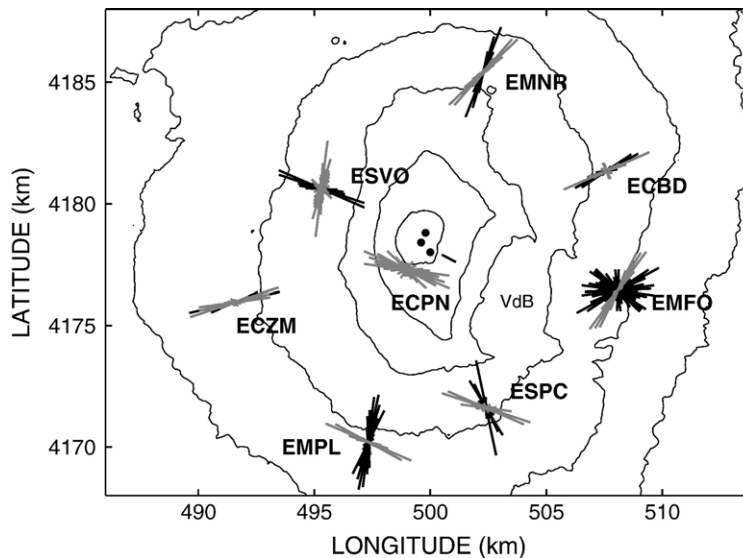


Fig. 7. Histograms of polarisation azimuths (rose diagrams) at all the stations of the network for the located events. Black and gray lines refer to the event's onset and following, largest-amplitude phases, respectively. Black dots and line mark the location of summit craters and 2004–2005 eruptive fissure, respectively.

309 of a large amount of shear energy. This observation may
 310 be interpreted in terms of a topography effect (e.g.,
 311 Ohminato and Chouet, 1997; Ripperger et al., 2003), a
 312 source effect, or a combination of both.

313 At present, we are unable to evaluate the relative
 314 contributions of these two different causes in shaping
 315 the observed signal. It is worth observing, however, that
 316 the largest-amplitude polarisations depict some spatial
 317 symmetry which could be suggestive of the radiation
 318 pattern from a non-isotropic source. At sites ECZM and
 319 ECBD, which are located at opposite directions with
 320 respect to the summit craters, the largest amplitudes are

321 associated with radial motion, while energetic shear
 322 components are observed at the ESVO-EMFO sites pair.
 323 Thus, the proximity to a P-wave nodal plane might offer
 324 an additional interpretation for explaining the lack of
 325 initial radial motion observed at station EMFO.

326 At present, it is impossible to speculate any further
 327 about the nature of the 12 LP source. Quantitative
 328 constraints about the geometry and dynamics of this
 329 source will hopefully be gained through future efforts
 330 aimed at retrieving the time-histories of the Moment
 331 Tensor components via full-waveform modelling (e.g.,
 332 Kumagai et al., 2002b, 2005).

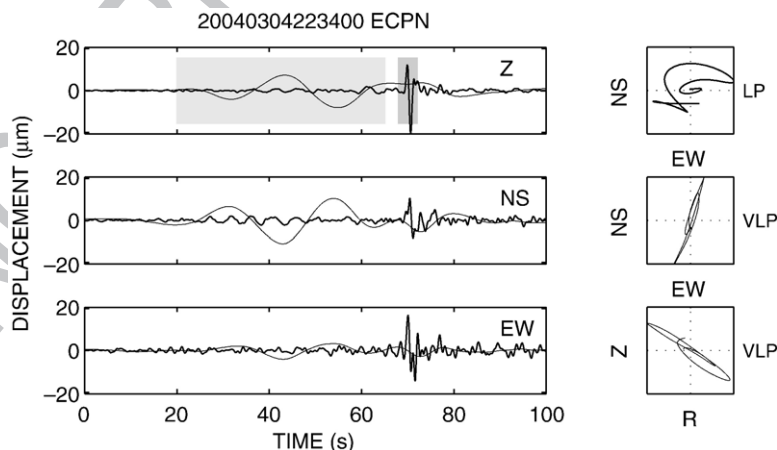


Fig. 8. Left: three-component displacement seismograms at summit station ECPN filtered over the 10–1 s (LP) and 40–10 s (VLP) period intervals (bold and thin lines, respectively). VLP data are magnified 2 times. The light and dark gray strips respectively mark the intervals selected for the VLP and LP particle motion analysis shown on the right. The onset of the LP signal is marked by a transverse pulse, while the VLP particle motion orbits point to the summit craters with a rather shallow incidence angle.

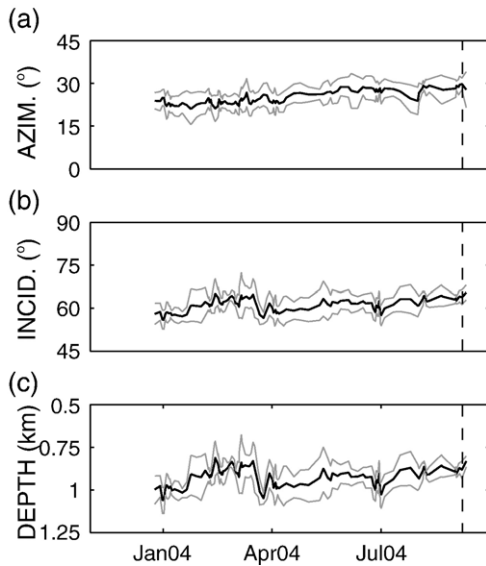


Fig. 9. Temporal evolution of azimuth (a) and incidence (b) angles for VLP pulses observed at summit station ECPN. The bold, black line is the fit to the original data, obtained by averaging subsequent groups of 40 measurements overlapping by 20. The thin, gray lines are the error bounds derived from the standard deviation of each set of observations. The plot in (c) indicates the depth of the VLP source obtained by projecting the polarisation vector onto the EW vertical plane passing through the craters. The depth is calculated with respect to the summit craters.

At station ECPN we also observed weak VLP oscillations at a period of about 20 s (Fig. 8), occurring either concurrently with or independently from the LP signals. These pulses exhibit rectilinear particle motion oriented toward the summit craters, with dominant incidence angles between 50° and 70° .

Using band-pass filtering over the 50–10 s period interval, we extended measurements of the VLP polarisation azimuth and incidence angles to all data segments from station ECPN. The low energy of these pulses means that they are not visible or cannot be analysed at the remaining recording sites.

Fig. 9 illustrates the temporal variation of the direction of VLP particle motion. By constraining the VLP sources to the EW vertical plane passing through the crater, and neglecting the particle motion distortion as a consequence of free-surface interaction (Neuberg and Pointer, 2000), we then convert VLP polarisation angles to source depths. Under such simplifying assumptions, and accounting for measurement uncertainties, the depth of the VLP source would vary between ≈ 750 m and ≈ 1100 m beneath the summit craters. The shallowest locations are associated with the January–April, 2004, interval, and with the later part of the catalog preceding the onset of the eruption. Finally, we compare the daily average of the peak-to-peak amplitudes of LP and VLP signals (Fig. 10). Within the

inherent limitations due to the lack of completeness of both catalogs, significant correlation exists between the two sets of observations, thus suggesting an inter-relationship between LP and VLP seismicity.

4. Locations

Determining the location of our LP data is challenging due to the emergent onsets and poor SNR at most stations. However, the similarity of waveforms may be exploited to get consistent estimates of inter-event differential times and high-SNR stacked waveforms to be used for reliable phase-picking procedures. From the LP catalog, we used a threshold of $50 \mu\text{s}^{-1}$ on the peak-to-peak amplitude of the vertical-component velocity seismograms at station ECPN to extract the largest 123 events for which recordings by at least 6 stations were available. For these events, we used the preliminary time pickings at station ECPN to guess P-wave onsets at the remaining sites of the network, based on the travel times predicted for a source located 500 m beneath the craters. Using this set of guessed time pickings, we repeated the correlation analyses using short (2.5 s long) time windows encompassing the expected signal onset. We then applied the Equivalence Class (EC) clustering algorithm (Press et al., 1992) to the resulting correlation matrices, using a correlation threshold of 0.9. In this manner, we found that all the recordings from individual stations grouped into a single, large family. In contrast to the clustering procedure presented above, the group formed by the EC algorithm is an ‘open’ tree, in the sense that not every event pair in the tree needs to be correlated. Events A, B and C form a tree if (A,B) and (B,C) are correlated,

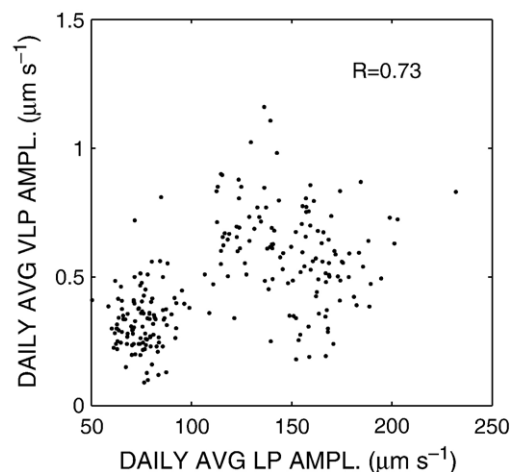


Fig. 10. Relationships between the daily-average of LP and VLP signal amplitudes.

390 regardless the correlation between A and C. This
391 approach is particularly advantageous for analysing
392 large catalogues where, for a given correlation threshold,
393 there are sets of events connected by at least one
394 differential time path (Shearer, 1997).

395 For the clustered data, we used the preliminary time
396 pickings and the precise, inter-event time delays derived
397 from correlation analyses to derive consistent alignment
398 of waveforms (Shearer, 1997). This procedure is expres-
399 sed in matrix notation as

$$\mathbf{d} = \mathbf{G}\mathbf{m} \quad (1)$$

400 where the vector \mathbf{d} contains the preliminary time
402 pickings and the correlation- derived inter-event time
403 differences; \mathbf{G} is an auxiliary matrix, and \mathbf{m} is a vector
404 containing the adjusted first arrivals. For the specific
405 case of a cluster formed by 4 events, the above equation
406 is written explicitly as:

$$\begin{bmatrix} T_1 \\ T_2 \\ T_3 \\ T_4 \\ \Delta T_{12} \\ \Delta T_{13} \\ \Delta T_{14} \\ \Delta T_{23} \\ \Delta T_{24} \\ \Delta T_{34} \end{bmatrix} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 1 & -1 & 0 & 0 \\ 1 & 0 & -1 & 0 \\ 1 & 0 & 0 & -1 \\ 0 & 1 & -1 & 0 \\ 0 & 1 & 0 & -1 \\ 0 & 0 & 1 & -1 \end{bmatrix} \cdot \begin{bmatrix} t_1 \\ t_2 \\ t_3 \\ t_4 \end{bmatrix} \quad (2)$$

408 where T_i is the preliminary P-wave arrival for event i ,
409 and Δ_{ij} is the time difference between events i and j
410 derived from correlation analysis. We seek a model
411 vector \mathbf{m} , containing the adjusted estimates of arrival
412 times t , that minimizes $\mathbf{d} - \mathbf{G}\mathbf{m}$. This is solved using a
413 weighted generalised inverse:

$$\mathbf{m} = (\mathbf{G}^T \cdot \mathbf{W} \cdot \mathbf{G})^{-1} \mathbf{G}^T \cdot \mathbf{W} \cdot \mathbf{d} \quad (3)$$

414 where the matrix \mathbf{W} contains weights associated with
416 the quality of both the preliminary picks and time
417 difference estimates. We assigned an unitary weight to
418 all the preliminary picks, while weighting the inter-
419 event delay times by the respective correlation coeffi-
420 cient c_{ij} using the relationship (Got et al., 1994):

$$W_{ij} = \frac{c_{ij}^2}{1 - c_{ij}^2} \quad (4)$$

423 Since the number of differential time estimates is
424 generally much larger than that of the preliminary time
425 pickings, the arrival times derived from Eq. (3) provide
426 a consistent alignment of waveforms, but the signal's

427 onsets are not necessarily correct. As a last step,
428 therefore, we derived absolute time pickings by
429 correcting the adjusted arrival times for the visually-
430 estimated onset of the stacked waveforms (Fig. 11;
431 Rowe et al., 2004).

432 For the location procedure, we used a non-linear,
433 probabilistic inversion (Tarantola and Vallette, 1982)
434 acting on reciprocal travel-times calculated using finite-
435 difference for the 3D heterogenous P-wave velocity
436 structure of Patanè et al. (2002). In our approach, we
437 first located the stacked event, and then used the
438 residuals from this location as station terms for the
439 subsequent location of the entire selected catalogue. In
440 this manner, the systematic errors due to our incomplete
441 knowledge of the earth structure are eliminated, while
442 high precision in the relative positions of individual
443 hypocenters is ensured by the previously-obtained least-
444 square adjustment of first arrivals.

445 Fig. 12 displays the final locations, compared with
446 the particle motion attributes for the VLP oscillations
447 measured at summit station ECPN. In agreement with
448 the similarity of waveforms, all the hypocenters are
449 tightly clustered in a small volume extending between
450 the surface and ≈ 800 m beneath the summit craters.
451 Once accounting for the uncertainties in the hypocenter
452 determinations, the source region of LP events thus
453 appear to be slightly shallower than the VLP source, if
454 one assumes the latter one to be located along the
455 crater's axis. It is worth noting that events initially
456 associated with the different clusters do not display any
457 systematic location pattern. This supports the initial

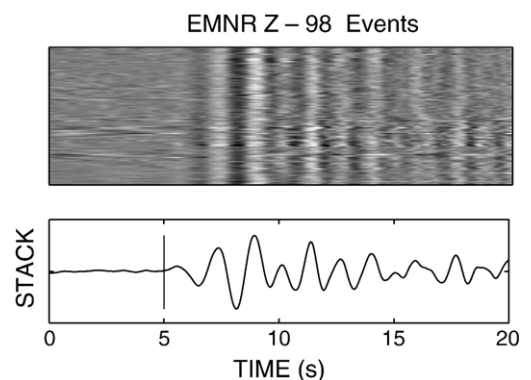


Fig. 11. Top: Vertical-component velocity seismograms for station EMNR after the least-squares adjustment procedure of arrival times. Dark and light gray tones correspond to troughs and peaks in the ground velocity time-history, respectively. Bottom: Stacked trace resulting from summation of the above seismograms. The vertical line marks the manually-picked onset. Note how the constructive summing of the aligned traces allows us to successfully recover the weak, positive first arrival, which is barely visible on the individual seismograms.

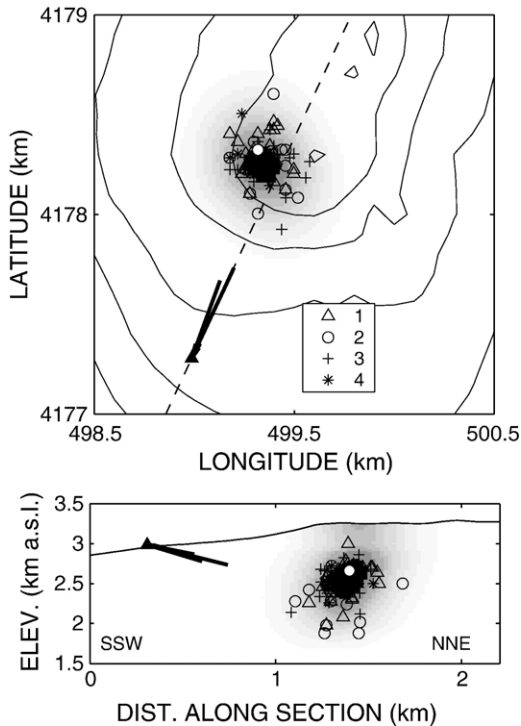


Fig. 12. Maximum-Likelihood hypocenter locations in map view (a) and projected along the vertical section passing through station ECPN (triangle) and the summit craters (b). The shaded regions bound the 90% confidence regions for hypocenter location, obtained by stacking the marginal probability distributions associated with individual events. White circles mark the location obtained from the stacked waveforms. The thick gray lines are the frequency distribution of azimuth and incidence angles for VLP pulses observed at the summit station ECPN during the time interval encompassed by the located LP events. Events pertaining to the different clusters are marked by different symbols, according to the legend shown in (a).

458 hypothesis that these clusters are most likely associated
 459 with slightly different excitation mechanisms of the
 460 same source, rather than being the result of spatially-
 461 distinct sources.

462 5. Discussion and conclusions

463 In this paper we described sustained LP activity
 464 occurring at Etna Volcano from November 2003 through
 465 to the end of 2004. LP amplitudes fluctuate with time,
 466 and climax during the January–April 2004 time span.
 467 The daily rate of occurrence peaked in late December,
 468 2003, late March, 2004, and late August, 2004, a few
 469 days before the onset of the 2004–2005 eruption. This
 470 latter peak is however associated with signals of very
 471 small amplitude, and is not particularly relevant once
 472 compared to the previous ones.

Throughout the analysed time interval, LP spectra
 display a few, narrow peaks spanning the 0.5–10 Hz
 frequency band. Due to their persistence at all the
 widely-spaced stations of the network, we attribute the
 origin of these peaks to a source effect. Once band-pass
 filtered around the main spectral component at frequen-
 cy ≈ 0.6 Hz, waveforms maintain a marked similarity
 throughout the analysed time interval, thus indicating
 the involvement of a non-destructive source process.
 Taken all together, these observations suggest that our
 LP events most likely represent the harmonic oscillation
 of a magmatic/hydrothermal reservoir repeatedly trig-
 gered by time-localized pressure steps (e.g., [Crosson
 and Bame, 1985](#); [Chouet, 1988,1996](#); [Fujita et al., 1995](#);
[Neuberg and Pointer, 2000](#); [Jousset et al., 2003, 2004](#)).
 Within this framework, the low frequency content of

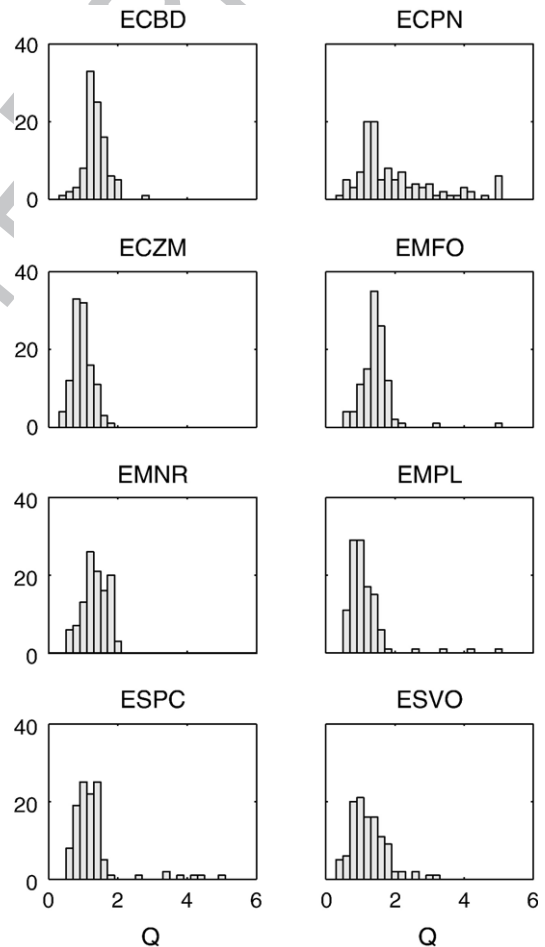


Fig. 13. Frequency distribution of Q measurements associated with the spectral peak at a frequency ≈ 0.6 Hz measured on the vertical components of all the stations of the network, for the set of located events.

489 these signals would be indicative of the existence of
490 interface waves generated at the fluid-rock boundary.

491 Previous studies of tremor and LP suggested that
492 both events represented the response of the same source
493 to different excitation mechanisms. For instance, [Chouet](#)
494 [et al. \(1997\)](#) analysed tremor and LP explosion-quakes
495 at Stromboli Volcano, Italy, and found that both signals
496 depicted several common spectral peaks. [Neuberg and](#)
497 [Pointer \(2000\)](#) analysed LP and tremor from La
498 Soufriere Volcano, Montserrat, and observed that LP
499 signals occur in swarms, occasionally merging in tremor
500 episodes characterised by harmonic spectral lines
501 shifting with time. [Neuberg and Pointer \(2000\)](#) modeled
502 this observation by simulating tremor signals as the
503 regular, repeated activation of the LP source at
504 consecutive time intervals. Thus, at both Stromboli
505 and La Soufriere volcanoes, LPs and tremor appear to
506 share the same source process, and would represent the
507 response of the same resonator to transient or repeated
508 excitation mechanisms, respectively.

509 In contrast, the spectral analyses presented for our
510 Etna data do not depict any common peak between LP
511 and tremor spectra. Moreover, detailed time-frequency
512 analyses (not shown here) indicate that the tremor-LP
513 transition is marked by an abrupt change of the spectral
514 signature, without any evidence of shifting spectral lines.
515 These observations thus suggest that, during the period
516 of our observations, LPs and tremor at Etna originated
517 from two distinct source processes and/or locations.
518 Corroborating this hypothesis are the results from a
519 recent study by [Di Grazia et al. \(2006\)](#), who retrieved
520 tremor locations at Etna throughout the year 2004. For
521 the period relevant to the present work (early April,
522 2004) the tremor sources imaged by [Di Grazia et al.](#)
523 [\(2006\)](#) are clustered beneath the summit craters, at
524 elevations around 1600 m (see their [Fig. 2](#)), and thus
525 considerably deeper than the source centroid of our LP
526 signals. These observations thus point to a complex
527 plumbing system, where different portions of the summit
528 conduits' network are affected by different dynamical
529 processes, thus generating distinct classes of signals.

530 The quantitative investigation about the geometry
531 and dynamics of the LP source requires a definition of
532 the time-history of the Moment-Tensor and Single-
533 Force components via full-waveform modelling of the
534 signals (e.g., [Kumagai et al., 2002b](#); [Nakano et al.,](#)
535 [2003](#); [Kumagai et al., 2005](#)). Some of the observations
536 presented above, however, already give hints about the
537 geometry of the LP source. LP models have been
538 proposed accounting for spherical (e.g., [Crosson and](#)
539 [Bame, 1985](#); [Fujita et al., 1995](#)) or cylindrical ([Chouet,](#)
540 [1985](#)) geometries. The large amount of shear energy,

and the non-isotropic radiation pattern observed during
our polarisation analysis can be explained, however,
only by invoking a crack-like geometry ([Chouet, 1988](#)),
as it is the only source geometry which is able to
generate a significant amount of shear waves. [Chouet](#)
(1988) showed that the fluid-filled crack generates a
very slow wave, that he called the crack wave, which
leads to more realistic estimates of the size of the
resonator than those associated with the spherical
geometry. Moreover, the crack-like geometry is also
appropriate for mass transport conditions beneath a
volcano. Under this perspective, the unevenly-spaced
peaks observed in our LP spectra would be representa-
tive of the different longitudinal and transverse modes
of vibration of the crack.

Following the resonating-fracture hypothesis, hints
about the composition and physical properties of the
fluid contained in the crack may be gained through
examination of the quality factor Q of the resonator,
that we measure from LP spectral estimates using the
relationship:

$$Q = \frac{f}{\Delta f} \quad (5)$$

where f is the frequency corresponding to a given
spectral peak and Δf is the width of that peak at half the
peak's magnitude. We applied Eq. (5) to the dominant
peak at a frequency ≈ 0.6 Hz separately to the three
components of motion for the set of located events,
obtaining similar results. Therefore, in [Fig. 13](#) we only
show results for the vertical component. All the stations
of the network yield consistent results, with distributions
of Q peaked at values around. This value is significantly
smaller than those observed at other volcanoes. Q
estimates from the literature span in fact the 10–500
range, the lower and upper bounds being associated with
LPs observed at Kilauea volcano, Hawaii ([Kumagai et](#)
[al., 2005](#)) and Galeras volcano, Colombia ([Kumagai and](#)
[Chouet, 1999](#)). The quality factor is made up by two
terms: $Q^{-1} = Q_i^{-1} + Q_r^{-1}$, where Q_i^{-1} expresses the
intrinsic attenuation in the fluid, and Q_r^{-1} refers to
energy losses at the fluid-rock interface.

Using [Chouet's \(1988\)](#) crack model, [Kumagai and](#)
[Chouet \(1999, 2000\)](#) performed a detailed investigation
of the acoustic properties of a fluid-filled fracture under
a variety of fluid and rock-matrix properties, in turn
examining ([Kumagai and Chouet, 2001](#)) the dependence
of such properties on the crack geometry and vibration
modes. These studies demonstrated that the wide range
spanned by Q measurements may be explained in terms
of the different physical properties of the multiphase

590 fluid mixtures and the surrounding rock matrix.
591 Although Kumagai and Chouet's results demonstrated
592 that the association between a given Q value and the
593 composition and physical properties of the fluid may be
594 manifold, their conclusions indicate that the high decay
595 rate observed for our LP events can only be explained in
596 terms of a low impedance contrast at the fluid-rock
597 interface which prevents the trapping of elastic energy
598 inside the fissure.

599 Several multiphase mixtures may fulfill such a
600 condition; however, the absence of any surface-visible
601 phenomena indicates that the most likely candidate for
602 the LP source is a crack filled with a basalt-gas or water-
603 vapour mixtures at very low gas-volume fractions. For
604 instance, a bubbly basalt at gas-volume fraction close to
605 0 and pressures between 5 and 20 MPa (thus encom-
606 passing the 13–17 MPa range which is expected at LP's
607 hypocentral depths) yields Q_r^{-1} over the 0.1–1 range
608 (see Fig. 13 in Kumagai and Chouet, 2000), and Q_i^{-1}
609 between 0.06 and 0.001. For this particular fluid, there-
610 fore, the total Q would be in the 1–10 range. Similarly, a
611 fracture filled by bubbly water at 1–1.5% gas-volume
612 fraction would depict Q_r^{-1} and Q_i^{-1} spanning the 0.2–1
613 and 0.05–0.001 ranges, respectively, thus resulting in a
614 Q between 1 and 5.

615 The temporal trends of the acoustical properties
616 (attenuation and dominant frequency) have important
617 implications for monitoring the volcano. Temporal
618 changes of such parameters may in fact be interpreted
619 in terms of variations in the physical properties of the
620 multiphase mixtures filling the resonating cavity (e.g.,
621 Kumagai et al., 2002a; Jousset et al., 2003; De Angelis
622 and McNutt, 2005). Within the measurement limitations
623 inherent the low amplitude and short duration of the
624 signals, the spectral signatures of our LP data appear to
625 be invariant with time (Fig. 4). Moreover, the marked
626 similarity of waveforms throughout the analysed time
627 interval suggests that the attenuation at the source
628 remained roughly constant as well. Taken all together,
629 these observations indicate that the physical conditions
630 of the upper conduit system feeding the summit craters
631 didn't suffer any remarkable variation during the months
632 preceding the eruption, in turn pointing to a general lack
633 of correlation between the LP-generating process and the
634 renewal of effusive activity. These arguments thus
635 support the hypothesis that both the feeding system
636 and triggering mechanism of the 2004–2005 eruption
637 were poorly related to the dynamics of the main
638 plumbing system feeding the summit craters. Rather,
639 the lack of any significant seismological precursor (Di
640 Grazia et al., 2006) and the petrological/geochemical
641 properties of the extruded lavas (Burton et al., 2005)

642 suggest that this eruption was mostly controlled by
643 geodynamic stresses acting on a lateral reservoir filled by
644 degassed, resident magma. Additional work is required
645 to put further constraints on these hypotheses: future
646 studies will therefore concern detailed analysis of the
647 dominant frequencies and decaying rate of LP wave-
648 forms, using methods with greater resolving power than
649 the Fourier-based spectral estimates adopted in this
650 work. Of particular promise is the *Sompi* method which,
651 based on an auto-regressive 24 equation, addresses the
652 problem of resolving the decaying harmonic compo-
653 nents of a time series corrupted by noise (Kumazawa
654 et al., 1990; Nakano et al., 1998). Moreover, the correct
655 interpretation of the acoustic properties in terms of fluid
656 characteristics requires the assessment of the LP source
657 geometry and dynamics. This latter task may be
658 approached using moment-tensor inversion from full-
659 waveform modelling of the most energetic signals. A
660 major challenge in this respect is the calculation of
661 Green's function which, as suggested by our polarisation
662 analysis, must necessarily account for the effects of
663 topography on wave propagation.

664 During our analyses, we also found a significant
665 correlation between the amplitudes of LP and VLP
666 signals. Taken together with the proximity of the respec-
667 tive source regions, these observations suggest that a
668 close link must exist between the two phenomena.
669 Recent studies (e.g., Chouet, 2003, and references
670 therein) indicate that VLP signals most likely represent
671 the dynamic response of the host rock to mass transport
672 processes in volcanic conduits, such as the sudden
673 decompression of large gas bubbles as they approach the
674 terminal part of a magmatic column.

675 Quantitative estimates about the volumetric changes
676 associated with this kind of process would require
677 determination of the Moment-Tensor components, a task
678 which is made impossible by the extremely poor SNR
679 exhibited by these events at all the peripheral stations of
680 the network. With the data presently available, a rough
681 estimate of the volume changes at the VLP source can
682 only be attempted by fixing the source location and
683 assuming a source of geometry. Under these simplifying
684 conditions, we considered an explosive (isotropic) point
685 source located at the barycenter of the LP sources, i.e. at a
686 distance and depth of 1000 m and 500 m, respectively,
687 from summit station ECPN. Using the classic Stokes'
688 solution (Aki and Richards, 2002), we then calculated
689 the Green's functions for an isotropic point source, using
690 $V_p = 3000 \text{ m s}^{-1}$, V_p/V_s ratio of $\sqrt{3}$, and rock density of
691 2500 kg m^{-3} . In this manner, we found that the typical
692 radial displacement of $\approx 10 \text{ }\mu\text{m}$ observed at station
693 ECPN would correspond to an isotropic moment M_i on

694 the order of $\approx 4 \times 10^{12}$ Nm. The isotropic moment is
 695 related to the volume change ΔV at the source by the
 696 relationship (Aki and Richards, 2002):

$$M_i = \Delta V(\lambda + 2\mu) \quad (6)$$

698 where λ and μ are Lamé's elastic coefficients. Assuming
 699 $\lambda = 2\mu$, and $\mu = 7.5$ GPa, from the above relationship we
 700 obtain a ΔV on the order of 1.5×10^2 m³. This value is
 701 smaller than, but comparable to, the ≈ 200 m³ volumetric
 702 changes associated with Strombolian explosions at
 703 Stromboli volcano (Chouet, 2003). Our estimates thus
 704 reinforce the hypothesis that the weak VLP events
 705 observed at Etna could actually be representative of a
 706 mass-transport process involving the movement and
 707 associated decompression of gas slugs as they approach
 708 the terminal part of the magmatic column. Under this
 709 perspective, the injection of these fluids into an
 710 overlying cavity filled by either magmatic or hydrother-
 711 mal fluids at poor gas volume fraction would drive the
 712 periodic pressurisation steps of such a shallow reservoir,
 713 thus triggering its resonant, LP oscillations.

714 In a previous work, Falsaperla et al. (2002) performed
 715 a detailed study of LP swarms occurred during the 1991–
 716 1993 eruption of Mt. Etna. These swarms were
 717 temporally correlated with episodic collapses of the NE
 718 crater floor, and originated in a region located slightly
 719 east of the NE crater, extending from the surface to a
 720 depth of 2000 m. These events depicted narrow spectral
 721 peaks spanning the 1–5 Hz frequency range, and fairly
 722 low quality factors ($Q = 13$ – 18). Falsaperla et al. (2002)
 723 used these features to develop a fluid-filled crack model
 724 of the source, in turn postulating that the locus of LP
 725 events was most likely associated with a dyke connect-
 726 ing the NE crater to a depressurizing magma body.
 727 Falsaperla et al. (2002) postulated that the repeated
 728 collapse of the overlying crater floor could have
 729 induced the transient pressurisations of that dyke, thus
 730 triggering its resonant, LP oscillations.

731 These past results, and the ones presented in this
 732 work, provide two interesting case studies demonstrat-
 733 ing the variety of dynamical processes acting as source
 734 of LP seismic waves. The quantitative assessment of
 735 these processes thus represent a fundamental step
 736 toward a better understanding of volcanic systems, in
 737 turn constituting a critical improvement in our ability to
 738 successfully forecast eruptive activity.

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