

Detection of Ridge-like Structures in the Pacific Large Low-Shear-Velocity Province

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Abstract

Waveform tomography is conducted for SH velocity structures of the entire mantle using approximately 3.5 times the data used for obtaining the previous model, SH18CE. The resultant new model, SH18CEX, exhibits a cluster of ridge-like low-velocity anomalies in the western part of the Pacific Large Low-Shear-Velocity Province (LLSVP). The location of the ridge-like anomalies is in good agreement with the location of the abrupt change in the topography of the D'' discontinuity. These results suggest that the LLSVP is associated with a cluster of ridge-like-piles, rather than a single large pile spread over the entire region. The piles probably consist of intrinsically dense material; however, either their volume or density contrast may not be sufficiently large to develop large-scale domes.

Keywords: tomography, Earth's internal structure, lowermost mantle, core-mantle boundary, mantle convection

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1. Introduction

As is well known, there exist two large low-shear-velocity provinces (LLSVPs) in the lowermost mantle beneath the Pacific and Africa. Ridge-like structures (low-velocity anomalies that are horizontally long and narrow) were detected in the African LLSVP via array analyses (e.g., Ni and Helmberger, 2003; Wang and Wen, 2007). Global tomography models also show generally consistent features (e.g., Grand, 2002; Takeuchi, 2007). It is important to verify whether such ridge-like structures are also observed in the Pacific LLSVP; if so, the ridge-like plume can be considered as the fundamental morphology of the upwellings.

Several array analyses were conducted in order to obtain regional structure models for the Pacific LLSVP. For instance, Takeuchi et al. (2008) suggested that the vertical extent of the low-velocity anomalies is approximately 400 km on the western side, whereas He and Wen (2009) suggested that it is approximately 740 and 340 km on the north-western and south-eastern ends, respectively. In these studies, the structure models were obtained by analyzing the data for event-station pairs on, or in the vicinity of, a particular great circle plane. Two-dimensional models were obtained by assuming that the structure is homogeneous in the direction perpendicular to the plane.

However, according to the global tomography model obtained by Schubert et al. (2004), the Pacific LLSVP consists of clusters of small-scale anomalies, and the validity of the aforementioned assumption is not evident. Furthermore, for some regions, such an assumption is clearly invalid; for example, Fig. 4 of Takeuchi and Obara (2010) shows a rapid variation in the ScS - S residuals (~ 8 s variation within 400–500 km) in the direction along which the

26 structures are assumed to be homogeneous by He and Wen (2009). Therefore,
27 further efforts to obtain three-dimensional models are required.

28 Takeuchi (2007) conducted global waveform tomography using three-
29 dimensional Born kernels and obtained the three-dimensional SH velocity
30 model, SH18CE. The tomography method adopted by Takeuchi (2007) uti-
31 lizes all the phases in the waveform data (including $ScSn$ and various major
32 and multi-orbit body phases); thus, the resolution of the LLSVPs is improved
33 significantly (see Fig. 2 of Takeuchi, 2007). This method is also advantageous
34 in that it can recover smaller-scale structures by fully considering the finite-
35 frequency effects (see Figs. 3 and 4 of Panning et al., 2009). In this study, we
36 improve the resolution by using a larger data set than that used by Takeuchi
37 (2007) and obtain a three-dimensional SH velocity model of the entire mantle.
38 The obtained model, SH18CEX, exhibits ridge-like low-velocity anomalies in
39 the western part of the Pacific LLSVP, where the resolution of the model is
40 high. In addition, we discuss the plausibility of the obtained features.

41 **2. Data and Method**

42 We invert the transverse component of the broadband waveform data from
43 IRIS GSN and GEOSCOPE for 679 events (Figure 1a). The data set used
44 in this study is a combination of the data set of Takeuchi (2007) (hereafter,
45 referred to as “Data Set 1”) and the new data set (hereafter, referred to as
46 “Data Set 2”). Data Set 1 includes only data for large events ($M_W \geq 6.5$),
47 whereas Data Set 2 includes data for smaller events (the smallest M_W is
48 6.0). The event distribution for Data Set 2 covers the area that had no or
49 very few events in Data Set 1 (such as Hawaii, the East African rift zone,

50 and mid-ocean ridges). We use a particularly large number of events in the
51 western Pacific region, thereby improving the resolution of the western part
52 of the Pacific LLSVP. The entire data set used in this study consists of 54,790
53 traces (271,798 time windows), which is approximately 3.5 times the number
54 of traces used by Takeuchi (2007). With the exception of the data set used,
55 the methods and parameters employed in this study are exactly the same
56 as those employed by Takeuchi (2007). The basic information required for
57 further discussion is summarized below.

58 The periodic ranges of the data set are exactly identical to those of the
59 data set used by Takeuchi (2007). The data set consists of velocity waveforms
60 with three different periodic ranges (200-400 s, 100-200 s, and 50-100). The
61 methods for data selection are exactly identical to those adopted by Takeuchi
62 (2007). We extracted time windows in which the residuals of the phase
63 and the amplitudes between the observed and synthetic seismograms are
64 reasonably small. These data selections were made to avoid the breakdown
65 of the Born approximations used in the inversion in this study. The resultant
66 data set for 200-400 s primarily consists of surface waveform data, whereas
67 the data set for 50-100 s primarily consists of body waveform data.

68 The model parameters and the damping parameters are also identical
69 to those of Takeuchi (2007). We used the anisotropic PREM (Dziewonski
70 and Anderson, 1981) and the Global CMT solutions as the initial models
71 for the structures and the source parameters, respectively, and we perturbed
72 only the elastic constants (i.e., the other parameters such as density, quality
73 factors, and source parameters were fixed). We expanded the perturbation
74 of the elastic constants N and L (notations follow those of Love, 1927) in

75 terms of 14 radial functions (13 linear spline functions in the mantle and 1
76 box-car function in the crust) for the vertically dependent part, and spherical
77 harmonics with a maximum angular order of 18 for the horizontally depen-
78 dent part. We defined the expansion coefficients as the model parameters.
79 Appropriate scaling relations were assumed between the perturbation of N
80 and L . The damping method and parameters are exactly identical to those
81 of Takeuchi (2007). Therefore, we can directly compare the new model,
82 SH18CEX, with the previous model, SH18CE.

83 **3. Obtained Model**

84 *3.1. Overall Features, Resolution, and Variance Improvements*

85 The resolution of SH18CEX is considerably better than that of SH18CE
86 (Figure 1b). The resolution of SH18CEX for the western Pacific region is
87 sufficient to recover the checkerboard pattern of heterogeneities whose scale
88 mimics the scale of the structures observed in Fig. 5. Note that the checker-
89 board patterns exist in both horizontal and vertical directions. The S waves
90 bottoming at various depths should primarily provide the vertical resolution.

91 The obtained model, SH18CEX, is shown in Figures 2 and 3. First,
92 we compare the lower mantle models of SH18CEX and SH18CE via vi-
93 sual inspection (Figure 2). We see that the overall patterns, i.e., the long-
94 wavelength features, of the two models are nearly invariant, but the signifi-
95 cant differences between the models are the relatively small-scale anomalies
96 observed only in SH18CEX (such as the features indicated by the green
97 arrows in Figure 2). This can probably be attributed to the resolution im-
98 provement in the new model.

99 Next, we compute the correlation coefficients between SH18CEX and
100 SH18CE as a function of depth and degree (Figure 4). The correlation coef-
101 ficients fluctuate among degrees (Fig. 4a) partly because the heterogeneities
102 are very small for some degrees. To clearly observe the overall features of the
103 correlations, we plot the correlation coefficients for each degree bin (degrees
104 1-3, degrees 4-6, \dots , degrees 16-18) (Fig. 4b). The thick black boxes denote
105 the ranges where the correlation coefficients are less than 0.70. Although we
106 have a few exceptions, we can confirm that the primary ranges with lower
107 correlation coefficients are higher degree components (degrees 16-18) in the
108 lower mantle, showing that the small-scale features in the lower mantle are
109 the primary differences between the models SH18CEX and SH18CE. The
110 lowermost mantle is the region with lower coefficients for a larger degree
111 range (degrees 10-18), and we will discuss their small-scale features in the
112 next subsection.

The newly identified small-scale features appear to be constrained primar-
ily by the body waveforms in Data Set 2. Table 1 summarizes the variance
improvements due to SH18CE and SH18CEX. The variance improvement is
defined by

$$\left(1 - \sum_i \frac{\int |u_{\text{obs}}^{(i)}(t) - u_{\text{final}}^{(i)}(t)|^2 dt}{\int |u_{\text{obs}}^{(i)}(t) - u_{\text{init}}^{(i)}(t)|^2 dt} \right) \times 100 (\%), \quad (1)$$

113 where $u_{\text{obs}}^{(i)}$ is the i -th time window of the observed seismograms, and $u_{\text{init}}^{(i)}$
114 and $u_{\text{final}}^{(i)}$ are the i -th time window of the synthetic seismograms for the
115 initial and the final model (either SH18CE or SH18CEX), respectively. The
116 evaluation of the variance improvements for Data Set 2 required extensive

117 computational resources; hence, we used an approximation. We selected
118 the data for 220 out of 488 events of Data Set 2, and we computed the
119 variance improvements for the selected 220 events. We assumed that these
120 improvements are identical to those for all 488 events of Data Set 2. Note that
121 the event selection was based only on the event date (events between 01/2006
122 and 09/2007 were selected), and no other selection rules were applied.

123 SH18CEX exhibits improvements comparable with those of SH18CE for
124 the periodic ranges of 200-400 s and 100-200 s (Table 1, top). For the periodic
125 range of 50-100 s, SH18CEX also exhibits comparable improvements for the
126 existing data (34% for SH18CE and 31% for SH18CEX); on the other hand, it
127 exhibits greater improvements for the incremental data (23% for SH18CE and
128 31% for SH18CEX) (Table 1, top). Improvements for the incremental data
129 themselves are not surprising because they are included only in the inversion
130 for SH18CEX, but note the larger improvements for the periodic range of
131 50-100 s compared with the other ranges. Considering that the data set of
132 50-100 s primarily consists of body waveforms, the results suggest that the
133 incremental constraints on the Earth's structures are primarily attributable
134 to the body waveforms in the incremental data set.

135 For the periodic range of 200-400 s, the improvements for Data set 1
136 are greater than those for Data set 2 (e.g., 44% and 32%, respectively, for
137 SH18CEX) (Table 1, top). This is probably due to the fact that the signal-
138 to-noise ratios of Data Set 2 are not adequate for longer periods because Data
139 Set 2 includes data for smaller events. Indeed, for the periodic range of 200-
140 400 s, variance improvements for larger events ($M_w \geq 6.5$) are significantly
141 larger than those for the data for smaller events ($M_w < 6.5$) (38% and 21%,

142 respectively, for SH18CEX) (Table 1, bottom). However, note that this does
143 not hold for the periodic range of 50-100 s (30% and 32%, respectively, for
144 SH18CEX) (Table 1, bottom), which suggests that such problems are not
145 encountered in this periodic range. Therefore, we can conclude that the
146 small-scale features in the lower mantle would be better constrained by the
147 incremental data set.

148 *3.2. Small-Scale Features in the Western Pacific Region*

149 We investigate the small-scale features observed in SH18CEX. We focus
150 on the western Pacific region, where the resolution of SH18CEX was con-
151 firmed to be high in Fig. 1b. The enlarged figures (Figure 5, top) indicate
152 that the strong low-velocity anomalies are horizontally long and narrow in
153 the vicinity of the core-mantle boundary (CMB). These ridge-like anomalies
154 surround the relatively high-velocity region (represented by the green dot in
155 Fig. 5), suggesting that the observed strong low-velocity anomalies are asso-
156 ciated with the return flow of the downwelling at the center. Such features
157 are not well observed in SH18CE (Figure 5, bottom).

158 Part of the strong anomalies (those intersected by the line $A-A'$ in Figure
159 5) extend to the shallower region. The vertical cross sections (Figure 6, top)
160 show that the extent of the anomalies is wide in the NW-SE direction, narrow
161 in the NE-SW direction, and high upwards. These features are similar to
162 those observed in the African LLSVP (e.g., Ni and Helmberger, 2003; Wang
163 and Wen, 2007).

164 We can confirm some similarities between SH18CEX and several recent
165 models. Figure 6 shows a comparison of SH18CEX, HMSL-S06 (Houser
166 et al., 2008) and S40RTS (Ritsema et al., 2011). In each section of $A-A'$

167 (Fig. 6, left), we can confirm tall and wide low-velocity anomalies; however,
168 the anomalies in HMSL-S06 and S40RTS appear to be slightly less tall and
169 slightly less wide, respectively, as compared to those in SH18CEX. In each
170 section of $B-B'$ (Fig. 6, right), we can confirm two piles of low-velocity anom-
171 lies. The right pile is taller than the left in each model; however, the pile
172 in SH18CEX appear to be tallest. Therefore, we can say that the ridge-like
173 anomalies are more pronounced in the model obtained in this study.

174 The observed ridge-like anomalies are not likely to be caused by resolution
175 smearing. Figure 7 shows the resolution kernels for the input anomalies
176 having a point-wise distribution in the horizontal direction. The extents of
177 the input anomalies in the vertical direction are different between Figs. 7a
178 and 7b. The kernels are more or less isotropic in the horizontal direction,
179 and few elongations are observed. Moreover, the smearing in the vertical
180 direction is small. Therefore, we can conclude that the ridge-like anomalies
181 are not due to the smearing effects.

182 **4. Consistency with the Travel Time Data**

183 We confirm the plausibility of the obtained model by checking its con-
184 sistency with the observed travel time data. We plot the distribution of the
185 $ScS-S$ travel time residuals observed by using Japanese broadband seismic
186 arrays (Figure 8a). 3,469 residuals were measured between 45.3° and 80.7°
187 using bandpass-filtered velocity seismograms with corner periods of 3.3 and
188 100 s. It should be noted that these residuals are independent of the data
189 set used in the waveform tomography in that: (i) the former is data from
190 the regional array, whereas the latter is data from global networks, (ii) the

191 former is relatively short-period data (around 3.3 s), whereas the latter is
192 longer-period data (around 50 s), and (iii) the former is relative travel time
193 data, whereas the latter only contains information regarding absolute travel
194 times.

195 As in Fig. 4 of Takeuchi and Obara (2010), the measurements in this study
196 indicate an 8 s variation in the residuals within a region of around 400–500
197 km (at the green line labeled *P* in Figure 8a of this paper). Although the
198 fluctuations are large, we also observe a variation of around 5 s in the residuals
199 (at the green line labeled *Q* in Figure 8a); the residuals of north-eastern part
200 are approximately 5 s larger than those of the south-western part. These
201 results intuitively suggest the existence of large velocity gradients in the NE-
202 SW direction. Other regions with relatively abrupt changes in the residuals
203 (*R* and *S* in Figure 8a) suggest the existence of velocity gradients in the NW-
204 SE direction. These features are generally consistent with those reported
205 previously (e.g., Fig. 1 of Schubert et al., 2004); however, the features in
206 Figure 8a appear to be clearer. This is probably because the results in Fig.
207 8a are obtained from a single regional array.

208 The low-velocity anomalies in SH18CEX effectively explain the observed
209 distribution of the *ScS-S* residuals (Figure 8b, left; see also Figure 8d, left).
210 By introducing a ridge-like structure, we can explain the abrupt change in
211 the residuals at *P*, *Q*, and *R* in Figures 8a and 8c. We can also explain the
212 abrupt change at *S* by other low-velocity anomalies in the lowermost mantle.
213 In contrast, SH18CE does not explain the observations (Figures 8b, right
214 and 8d, right). Larger residuals are observed in the region surrounded by the
215 lines *P*, *Q*, and *R* (Figs. 8a and 8c), whereas the model SH18CE predicts

216 smaller residuals (Figs. 8b, right and 8d, right). The results strongly suggest
217 that the low-velocity anomalies in SH18CEX are more plausible than those
218 in SH18CE.

219 The obtained low-velocity structures have good correlations with the D''
220 topography observed by Takeuchi and Obara (2010), who analyzed *ScS-SdS*
221 times for the Fiji-Tonga events. The sampling region extends across the
222 ridge-like structure (Figure 9a). The *ScS-S* residuals observed by Takeuchi
223 and Obara (2010) were indeed large at the center of the ridge-like structure,
224 and they linearly decreased with increasing distance from the center (Figure
225 9b, left). The D'' discontinuity was deep at the center, became slightly shall-
226 lower at the side, and abruptly became very shallow beyond the side of the
227 ridge-like structure (Figure 9b, right). The abrupt jump in the discontinuity
228 suggests that the ridge-like structure is probably associated with a chemically
229 distinct pile (Figure 9c), as discussed by Takeuchi and Obara (2010).

230 5. Discussion and Implications

231 In several previous studies, the LLSVPs have been interpreted as isolated
232 piles of intrinsically dense materials (e.g., Ni et al., 2002; Ni and Helmberger,
233 2003; Wang and Wen, 2007). However, such piles are often expected to
234 have larger-scale structures (e.g., Tackley, 1998, 2002; McNamara and Zhong,
235 2005), which seems to contradict the cluster of small plumes observed in
236 this study. In contrast, Schubert et al. (2004) proposed that LLSVPs are
237 clusters of isochemical thermal plumes, which seems to contradict the abrupt
238 change in the topography of the D'' discontinuity observed in this study.
239 One solution for these contradictions may be as follows: the piles consist

240 of intrinsically dense materials; however, either the volume or the density
241 contrast of the dense materials is small. Under these circumstances, the
242 thermo-chemical plumes are expected to be similar to the isochemical thermal
243 plumes (see, for example, the discussions by Bull et al., 2009).

244 The morphology of the plumes has long been debated, even for simple
245 Rayleigh-Bénard convections. Bercovici et al. (1989), for example, suggested
246 that the upwellings in the earth-like spherical shells are conduit-like, whereas
247 Houseman (1990) and Yanagisawa and Yamagishi (2005) suggested that the
248 upwellings are sheet-like. The most fundamental difference between these
249 studies is the Rayleigh numbers that were considered. The existence of ridge-
250 like structures suggests that the convection in the lower mantle is as vigorous
251 as that for large Rayleigh numbers (more than, say, 1000 times the critical
252 Rayleigh number).

253 In the new model, SH18CEX, we see ridge-like structures in both the
254 African LLSVP and the Pacific LLSVP (Figure 2, top). The structures in
255 the African LLSVP are similar to those obtained by Schubert et al. (2004).
256 The results suggest that both the Pacific and the African LLSVPs consist of
257 clusters of chemically distinct piles. It is notable that piles are not spread over
258 the entire region of the LLSVPs, but confined only to the ridge regions. The
259 recent high-P,T elasticity simulation of deep mantle minerals suggests that
260 small volume fractions of mid-ocean ridge basalt (MORB) in the lowermost
261 mantle are sufficient for explaining the amplitude of V_s and V_ϕ anomalies
262 observed in tomographic studies (Tsuchiya, 2011). The ridge-like pile clusters
263 seem to be compatible with this mineralogical interpretation.

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Table 1: Comparison of variance improvements.

	Data Set 1		Data Set 2	
	SH18CE	SH18CEX	SH18CE	SH18CEX
200-400 s	42 %	44 %	30 %	32 %
100-200 s	41 %	42 %	33 %	37 %
50-100 s	34 %	31 %	23 %	31 %

	Data Set 2 ($M_w \geq 6.5$)		Data Set 2 ($M_w < 6.5$)	
	SH18CE	SH18CEX	SH18CE	SH18CEX
200-400 s	36 %	38 %	20 %	21 %
100-200 s	34 %	38 %	32 %	36 %
50-100 s	21 %	30 %	24 %	32 %

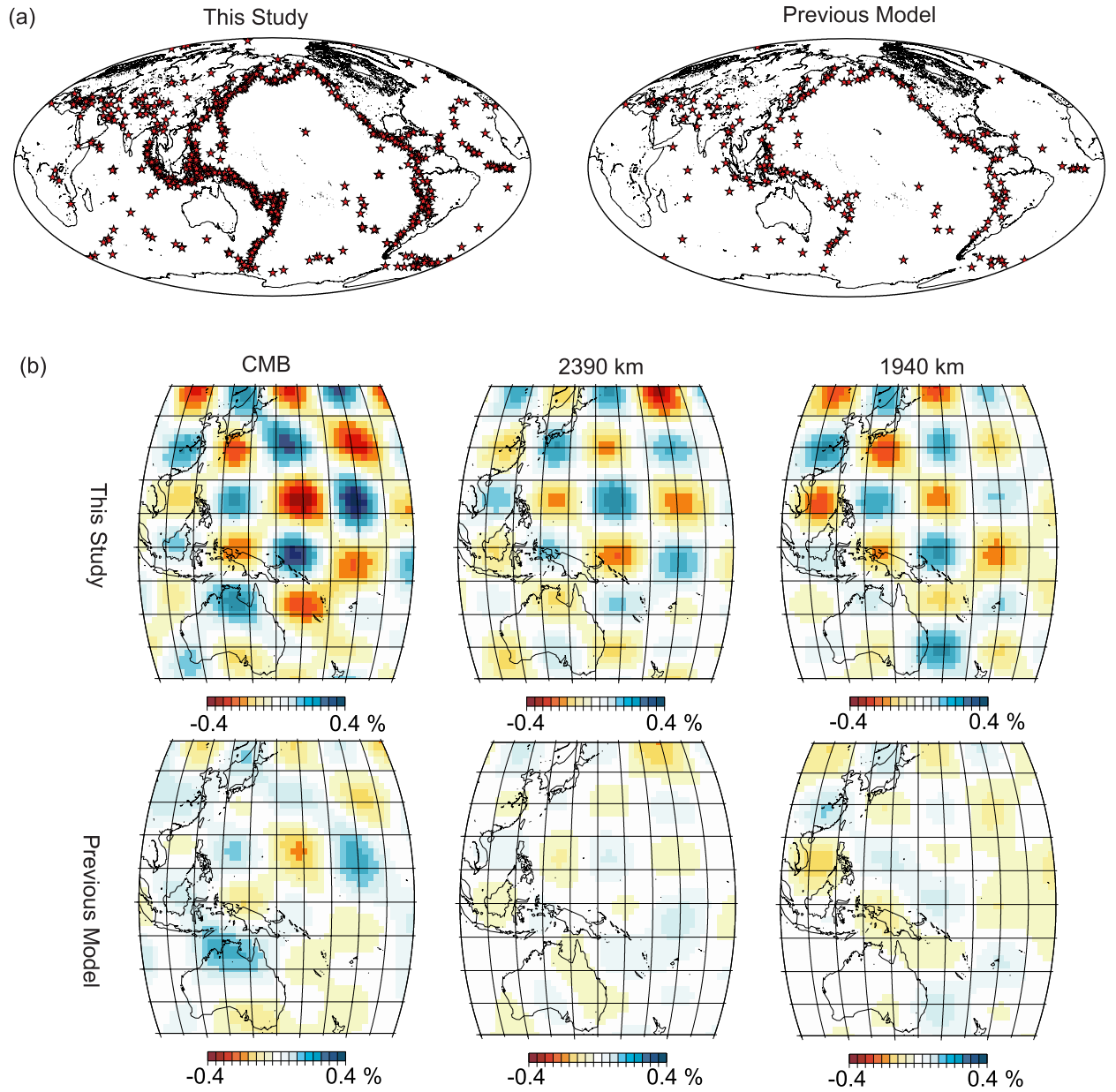


Figure 1: (a) Events used for obtaining SH18CEX in this study (left) and SH18CE (Takeuchi, 2007) (right). (b) Recovered models for checkerboard patterns of the heterogeneities when we use the data sets for SH18CEX (upper figures) and SH18CE (bottom figures).

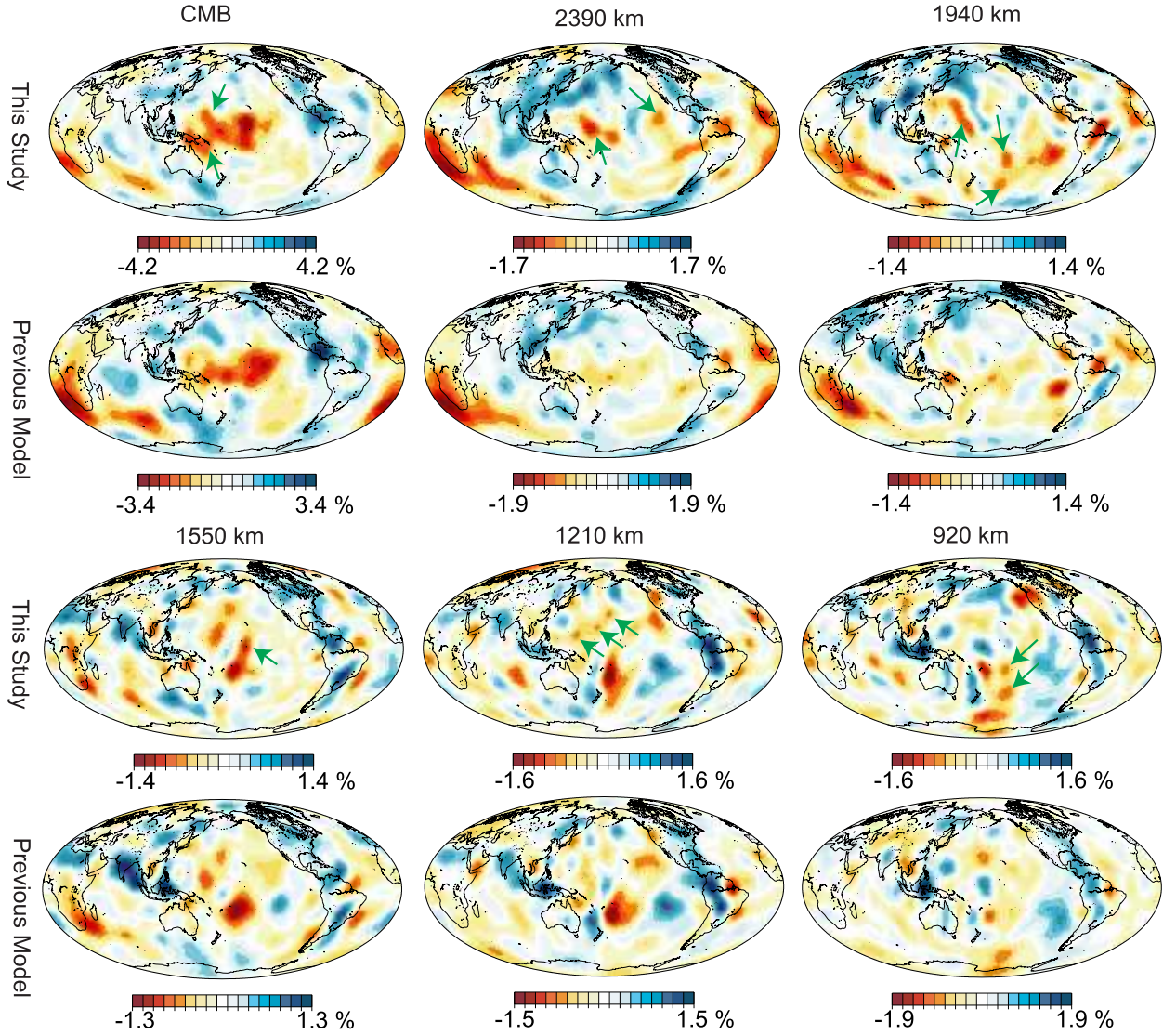


Figure 2: Comparison between SH18CEX (upper figures) and SH18CE (lower figures) at various depths in the lower mantle. The SH velocity perturbations with respect to the initial model, PREM (Dziewonski and Anderson, 1981), are shown. The green arrows indicate examples of the prominent features observed only in SH18CEX.

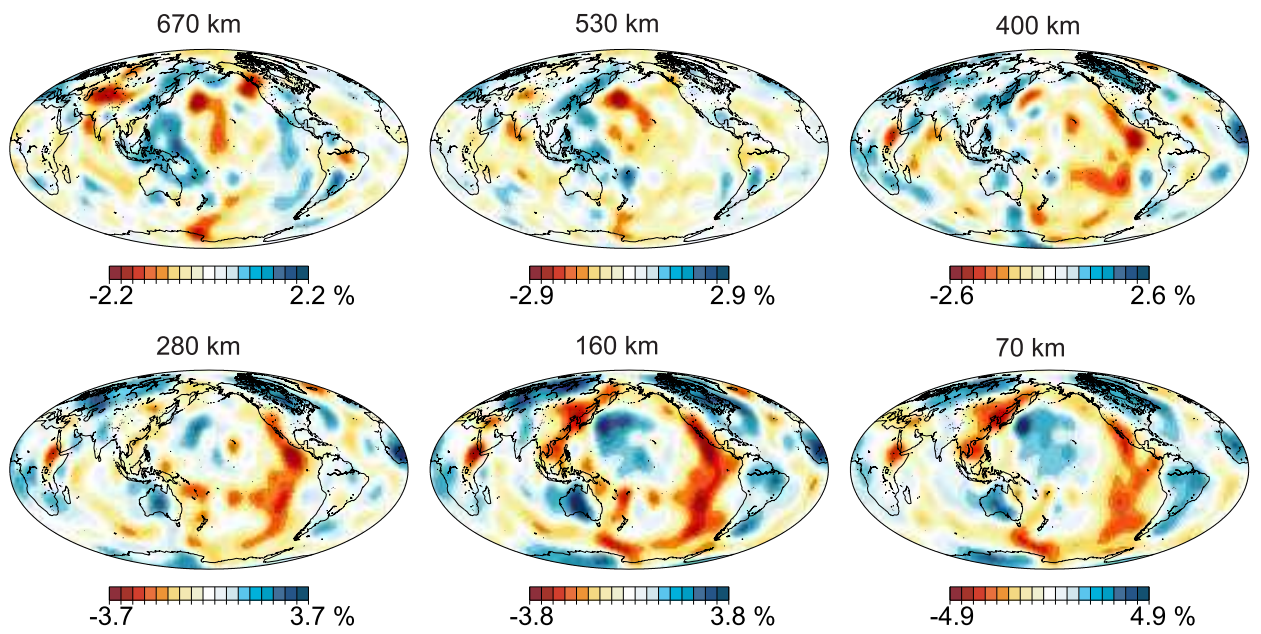


Figure 3: SH18CEX at various depths in the upper mantle.

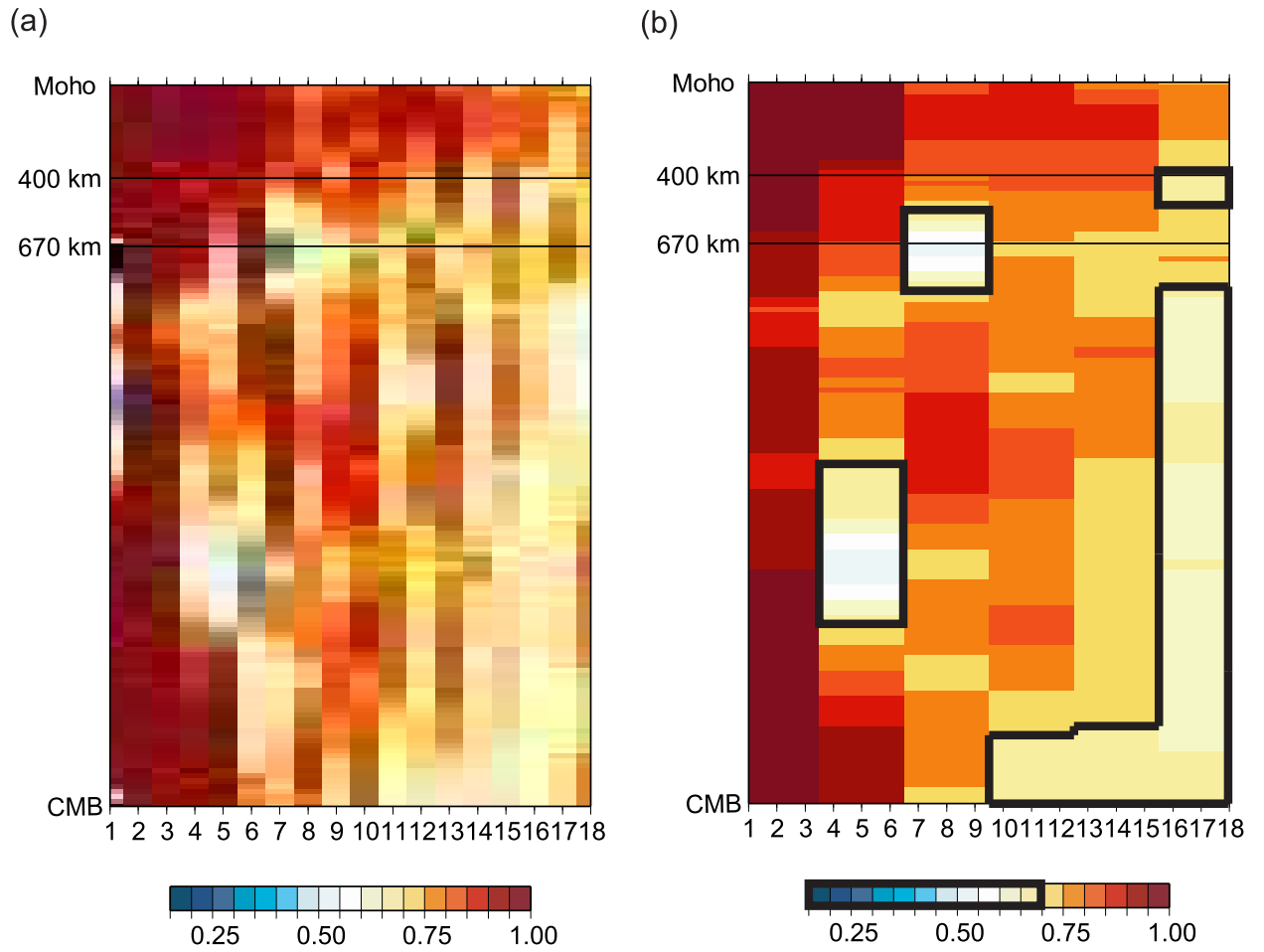


Figure 4: (a) Correlation coefficients between SH18CEX and SH18CE as a function of degree (horizontal axis) and depth (vertical axis). The discontinuities at 400 and 670 km are indicated by solid lines. (b) The same as (a) but showing the correlation coefficients for degree bins (degrees 1-3, 4-6, \dots , 15-18). The thick black boxes denote the regions with correlation coefficients less than 0.70.

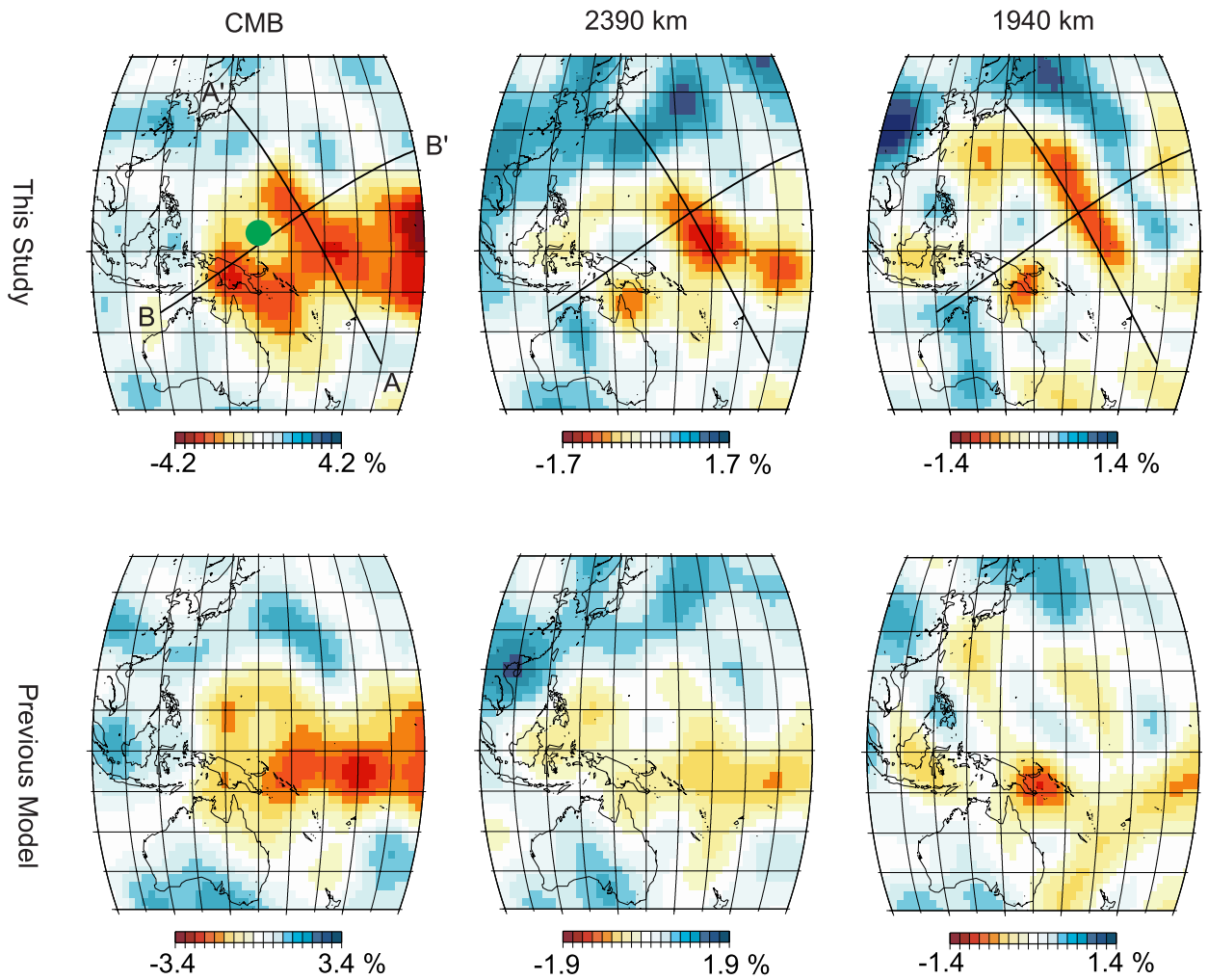


Figure 5: Comparison between SH18CEX (upper figures) and SH18CE (lower figures) in the western Pacific region. The lines A-A' and B-B' denote the locations of the vertical sections shown in Fig. 6. The green dot denotes the relatively high-velocity region discussed in the text.

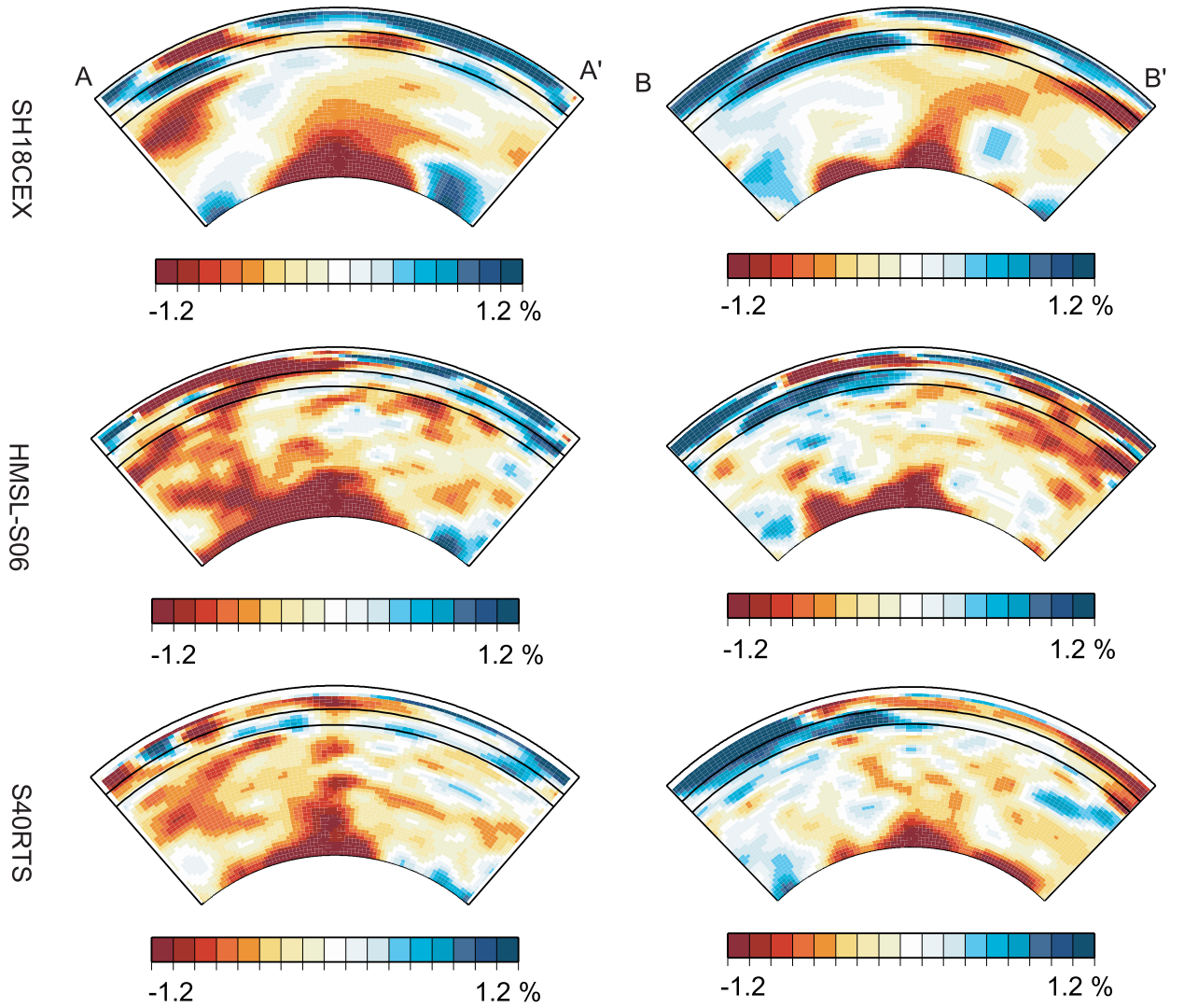


Figure 6: Vertical cross sections of SH18CEX obtained herein (upper figures), of HMSL-S06 obtained by Houser et al. (2008) (middle figures), and of S40RTS obtained by Ritsema et al. (2011) (bottom figures) at the locations indicated by the lines $A-A'$ (left) and $B-B'$ (right) in Fig. 5.

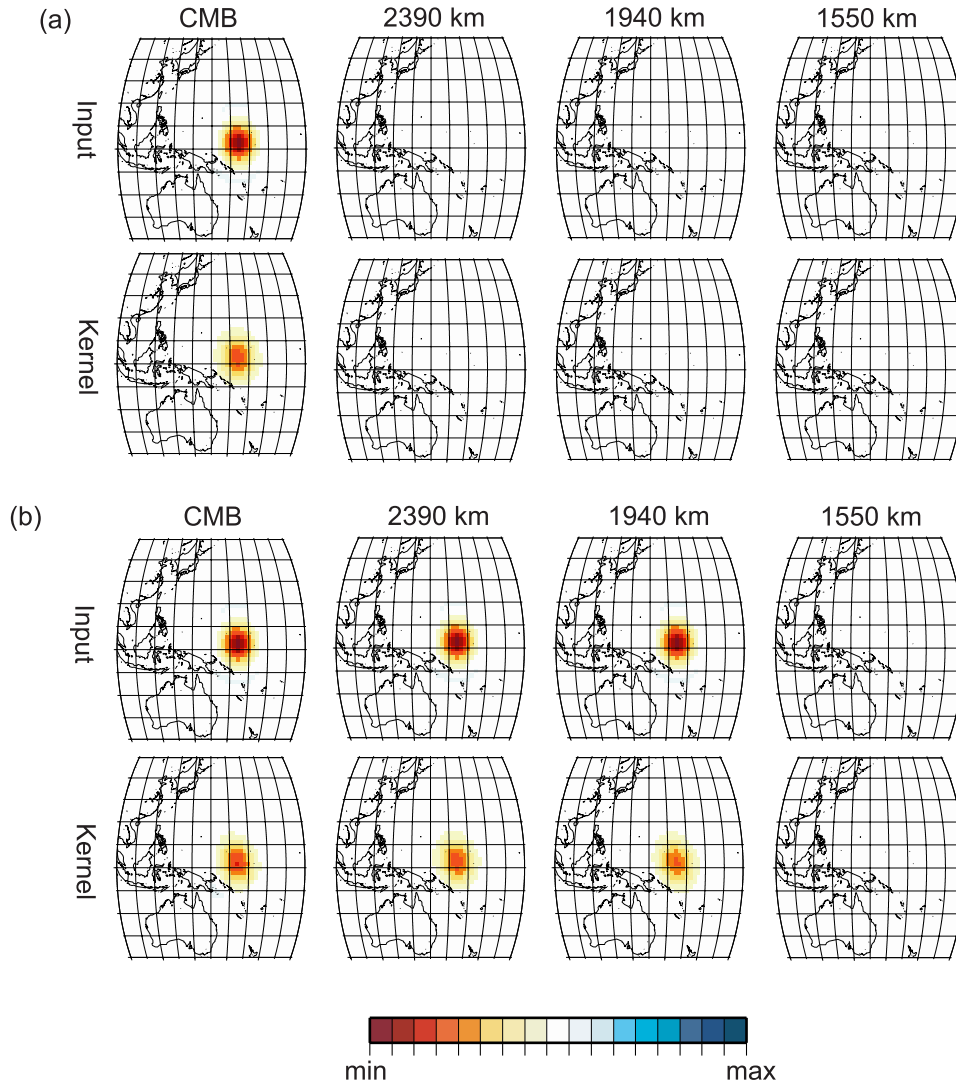


Figure 7: (a) Resolution kernel for low-velocity anomalies having a point-wise distribution in the horizontal direction. The input model (upper figures) and the resolution kernel of SH18CEX (bottom figures) are compared. The vertical dependent part of the input anomalies is given by the perturbations of the model parameters for the linear spline function whose node is at the CMB. (b) The same as (a), except that the input anomalies are given by the perturbations of the model parameters for the linear spline functions whose node is at the CMB, 2390 km depth, and 1940 km depth.

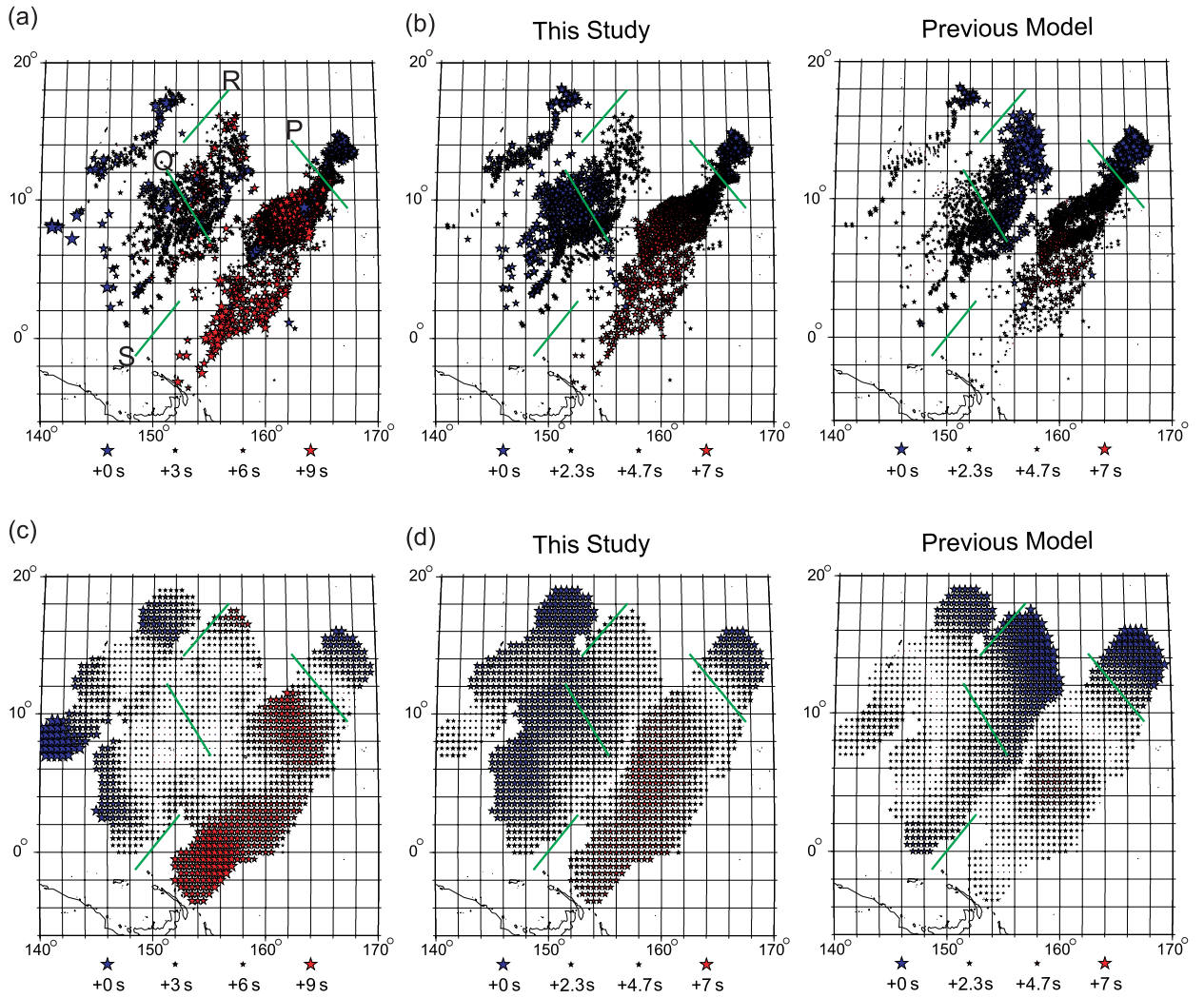


Figure 8: (a) Distribution of the observed ScS - S residuals measured in this study. The residuals are projected at the bouncing point of ScS . The green lines denote the rough locations of the abrupt jump of the residuals discussed in the text. (b) Same as (a), except for the predictions plotted using SH18CEX (left) and SH18CE (right). (c),(d) The same as (a) and (b), respectively, other than plotting the cap averaged residuals with 1.5° radius.

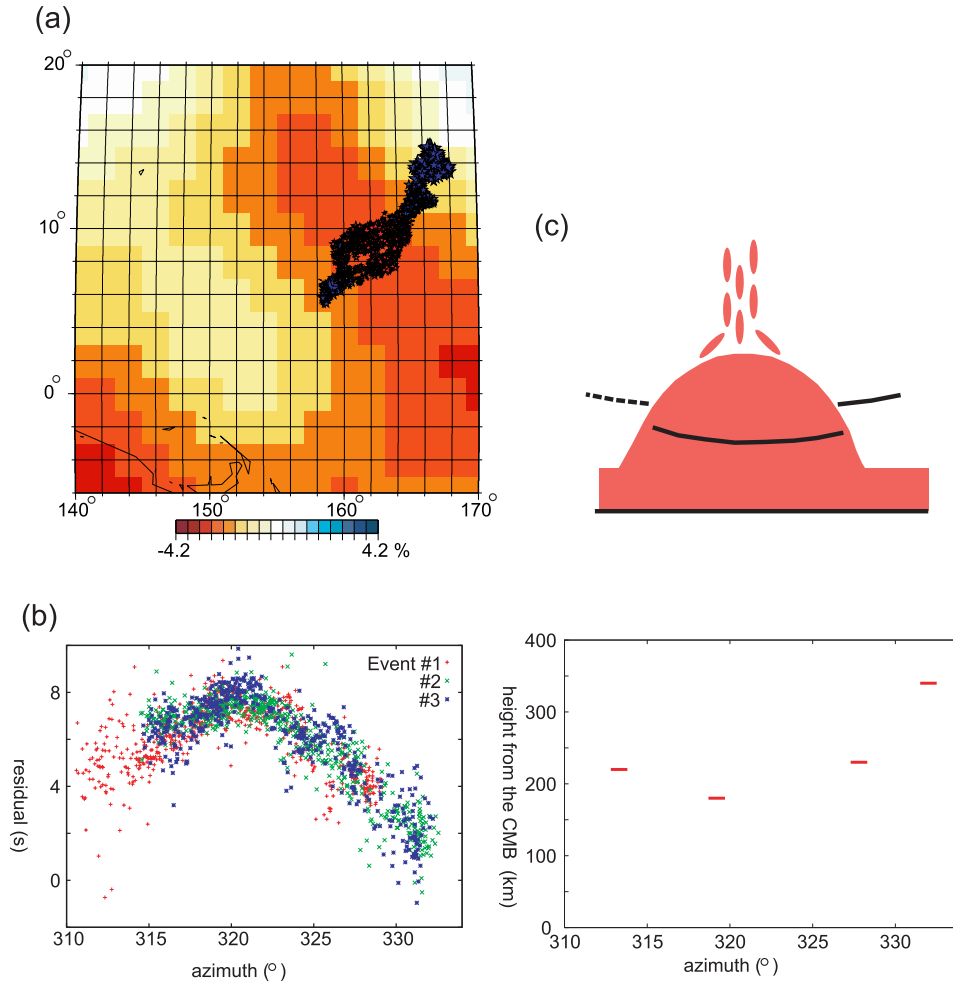


Figure 9: (a) SH18CEX at the CMB overplotted by the $ScS-S$ residuals previously reported by Takeuchi and Obara (2010). Note that the scale for the residuals is not identical to that in Fig. 8, but it is appropriately chosen for the plot. (b) The $ScS-S$ residuals shown in (a), plotted as a function of the azimuth. The azimuth is measured from the centroid of the events analyzed by Takeuchi and Obara (2010) (left). The height of the D'' discontinuity as a function of the azimuth reported by Takeuchi and Obara (2010) (right). (c) Schematic diagram of the structures of the region studied by Takeuchi and Obara (2010). The red part denotes the chemically distinct region and the solid black lines denote the D'' discontinuity.