UNIVERSITÄT BONN Physikalisches Institut

Towards Low Energy Physics from the Heterotic String

von Saúl Noé Ramos Sánchez

We investigate orbifold compactifications of the heterotic string, addressing in detail their construction, classification and phenomenological potential. Based on the insight gained from grand unification theories, we develop a successful strategy to search for models resembling the minimal supersymmetric extension of the standard model (MSSM) in \mathbb{Z}_6 –II orbifold compactifications. We find about 200 MSSM candidates with the gauge group and the exact spectrum of the MSSM, and supersymmetric vacua below the compactification scale. Among them, there are several models with the following realistic features: R-parity, seesaw suppressed neutrino masses, and intermediate scale of supersymmetry breakdown.

Post address: Nussallee 12 53115 Bonn Germany



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Chapter 1

Introduction

One of the current goals of theoretical physics is to formulate a theory which explains all observed forces in a unified way. The Standard Model (SM) of particle physics [1–3] is one of the major advances in physics of the second half of the last century. Including three of the four known fundamental forces of the universe through local $SU(3)_c \times SU(2)_L \times U(1)_Y$ gauge invariance, the SM describes with great precision the interactions between particles at currently probed energies ($\sim 100 \text{ GeV}$). It also predicts the existence of an SU(2) doublet, called Higgs boson, which gives masses to all quarks and leptons once it acquires a vacuum expectation value (VEV). Although the Higgs boson is still to be discovered, there are good reasons to believe that this will occur in the next collider generation.

Despite its predictive power, from a theoretical point of view, there remain still some open questions, such as the stability of the electroweak scale (hierarchy problem), the number of free parameters, the source of the accelerated expansion of the universe (dark energy), the origin of the observed repetition of families and, of course, the absence of a description of gravity in the SM. These issues can indicate that the SM is not a fundamental theory, but rather an effective limit of more general physics at higher energies.

A possible extension of the SM is to introduce a symmetry between bosons and fermions, known as supersymmetry (SUSY) [4–7]. SUSY explains elegantly how a reasonable Higgs mass can be protected from (quadratically divergent) quantum corrections without fine-tuning the parameters of the theory. Therefore, the Higgs mass remains of the order of the electroweak scale, ensuring the stability of this scale even if the theory is valid up to very high energies. Apart from providing a plausible solution to the hierarchy problem, it also predicts unification of all gauge couplings at the unification scale $M_{GUT} \sim 2 \times 10^{16}$ GeV. However, SUSY requires new particles associated to those already known. The so-called superpartners of the SM particles differ only by their spin, so that the superpartner of a fermion is a boson and vice versa. Since no superpartner has been detected so far, SUSY must be broken. Yet one can argue that its breakdown occurs in a fashion such that it lets the main properties of SUSY remain at low energies.

Even without SUSY, the running of the couplings and the symmetries of the particle content of the SM suggest a unified picture of strong and electroweak interactions through grand unified theories (GUTS) [8,9]. The fundamental feature of these theories is that,

at the scale M_{GUT} , all gauge interactions of the (non)supersymmetric SM are gathered together in a single and bigger gauge group, such as SU(5), SO(10) or E₆. This situation seems much more natural in the minimal supersymmetric extension of the SM (MSSM) as the couplings do (almost) meet at M_{GUT} in such a model, stressing the key role that SUSY might play in physics beyond the SM.

One particularly appealing GUT is the SO(10) unified model [10, 11], in which one generation of matter is accommodated in a single spinor representation, according to

$$16 = (3,2)_{1/6} + (\overline{3},1)_{-2/3} + (\overline{3},1)_{1/3} + (1,2)_{-1/2} + (1,1)_1 + (1,1)_0, q \overline{u} \qquad \ell \overline{e} \overline{\nu}$$
(1.1)

where quantum numbers with respect to $SU(3)_c \times SU(2)_L$ are shown in parentheses and the subscript denotes hypercharge. A remarkable prediction of this theory is the existence of right-handed neutrinos, which were not expected in the SM¹ and are very important to explain massive left-handed neutrinos. Moreover, the embedding of the hypercharge in $SU(5) \subset SO(10)$ predicts the weak mixing angle θ_w and provides thereby an explanation of the electric charge quantization.

Beside their attractive properties, GUTs introduce some problems of their own and leave some questions unanswered. A puzzling feature is that, while matter generations are described by complete GUT representations, Higgs and gauge bosons appear only as incomplete or *split* GUT multiplets. In addition, questions like why there are three families in the SM, why their mass eigenstates mix as they do, what the explanation for dark energy is, are still riddles that await their resolution in these scenarios. Some proposals such as incorporating additional discrete (family) symmetries might answer some of these questions. However, we have still to deal with the fact that gravity –whose description by general relativity has been also very well tested– is very different in nature and cannot be directly described along with the other interactions.

Unification of gravity with the rest of the fundamental forces into a single theory led Kaluza and Klein to introduce a fifth spatial dimension compactified on a minute circle [12, 13]. The proposal consisted in extending general relativity to a five-dimensional spacetime. The resulting theory contains surprisingly a set of equations equivalent to those of general relativity, and another set equivalent to Maxwell's equations. That the fifth dimension escapes to our observations was justified by reducing its size. This mechanism introduces an infinite set of heavy particles (modes) termed Kaluza-Klein tower. Despite its beauty, this early attempt revealed soon not to be appropriate for the unified description of gravity and electromagnetism, since the resulting theory cannot be quantized. Therefore, with the growing success of quantum mechanics, the interest in it receded considerably. However, this idea came again to life several years later in a theory that, with its evolution, turned into a good candidate to unify consistently all known forces: string theory.

¹Right-handed neutrinos, however, can be naturally embedded in the sm.

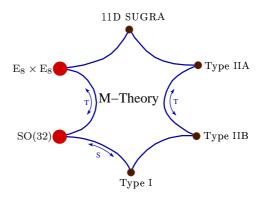


Figure 1.1: (Super)string theories are connected by a web of dualities.

String Theory

String theory arose by attempting to describe the strong interactions, but, once it was noted that string theory includes a spin 2 particle corresponding to the graviton, it became clear that its purpose was very different.

String theory is a perturbative quantum theory in which ordinary point particles are replaced by one-dimensional objects, whose various vibrational modes at the string scale $M_{\rm str}$ can be identified at low energies with different particles. These extended objects, named strings, cover a two dimensional space called *worldsheet*, in which many of its properties acquire a description through conformal field theory. Depending on their boundary conditions, they can be closed or open.

Not only does string theory contain the graviton as one of the vibration modes of the strings, but it indeed reduces to Einstein's theory of gravity at low energies and, due to the extended nature of the string theoretical graviton, this theory also avoids the ultraviolet divergences of graviton scattering amplitudes. Therefore, a quantized description of gravity is possible in string theory [14–16].

Several constraints are inherent to the quantum nature of strings. For instance, requiring that quantum anomalies do not spoil Lorentz invariance of the theory constrains the dimension of the spacetime in which the strings can consistently propagate. This is a striking theoretical achievement because no theory before offered a prediction about the dimensionality that our spacetime must have. At the same time, this poses a major challenge since, as we will see, no consistent string theory describes a spacetime with four dimensions like the one that is so familiar to us.

Historically, the first string theory discovered was the bosonic string, that is consistent in 26 dimensions. This theory was immediately discarded for it contains unphysical particles with imaginary rest mass (negative square mass) called tachyons. Furthermore, this theory clearly cannot yield a description of our universe because the particles composing the observed matter are fermions.

Consistent tachyonic-free string theories require (local worldsheet) SUSY at very high energies [17, 18] and predict a ten-dimensional spacetime. There exist only five consistent (super)string theories: type IIA, type IIB, type I, and the $E_8 \times E_8$ and SO(32) heterotic

theories. These theories are connected by a web of (conformal) dualities and thought of as different limits of an underlying 11-dimensional theory (M-theory) [19, 20], as we depict in fig. 1.1.

The type I and the heterotic string theories are attached to a remarkable discovery, namely that the SO(32) gauge symmetry of the type I and SO(32) heterotic string as well as the $E_8 \times E_8$ gauge group of the other heterotic theory [21, 22] follow from anomaly cancellation [23, 24]. In contrast to these theories, both type II string theories do not consist of nonabelian gauge groups. This is why most of the early works on string theory were mostly focused on the $E_8 \times E_8$ heterotic string which, as we shall confirm, seemed from the beginning the most promising candidate to be a theory capable to describe physics at low energies. It was only after the discovery of extended higher dimensional objects called D-branes [25–27] that type II theories regained interest.

Ever since the discovery of the five consistent string theories, one of the most important issues for string theory has been to make contact with reality. In fact, this is the primary motivation of this thesis. There are, unfortunately, many aspects of string theory that make difficult to believe that it has something to do with the observed universe. Apart from the fact that the SM gauge group does not appear automatically in these theories, the most disturbing feature is that all superstring theories predict a ten-dimensional spacetime. Happily, the old idea of Kaluza and Klein can be adapted effectively in string theories to obtain a consistent reduction from ten to four dimensions.

String Compactifications

Consistent string theories are formulated in ten dimensions. In order to make contact with our four dimensional world, one has to get rid of six of these dimensions. To do it, one usually chooses them to be small and compact enough to escape experimental detection. Notwithstanding their undetectability, they can influence the physics of our observable universe.

Some of the schemes proposed in order to obtain four-dimensional models are Kaluza-Klein (KK) compactifications and D-brane worlds [28]. Let us spend some words on their general properties.

KK compactifications are a natural extension of the five-dimensional KK approach. One considers the ten dimensional spacetime \mathcal{M}^{10} of the string to be a direct product of a four-dimensional flat (Minkowski) spacetime $\mathbb{M}^{3,1}$ and some unknown compact manifold \mathcal{M}^6 , i.e. $\mathcal{M}^{10} = \mathbb{M}^{3,1} \times \mathcal{M}^6$. Further, one assumes that the metric of the space is block-diagonal, such that

$$ds^{2} = g_{\mu\nu}^{(4)}(x) dx^{\mu} dx^{\nu} + g_{mn}^{(6)}(y) dy^{m} dy^{n}, \qquad (1.2)$$

where $g_{\mu\nu}^{(4)}$ is the Minkowski metric and $g_{mn}^{(6)}$ is the metric of the compact internal space.

The expansions of the ten-dimensional fields in the modes of the internal manifold \mathcal{M}^6 yield the theory in four dimensions. As a result of these expansions, an infinite tower of massive modes appears in the lower-dimensional theory. The masses of these four-dimensional states depend inversely on the size of the extra dimensions. If one chooses the

size of the internal manifold to be sufficiently small, the massive KK states become heavy and thus decouple from the spectrum.

Not every six-dimensional manifold is admissible as internal manifold \mathcal{M}^6 . In particular, if one insists on preserving (some) SUSY in the four-dimensional theory, the internal space must have SU(3) holonomy. Furthermore, it is necessary to choose the six-dimensional manifold to be Ricci flat, i.e. such that the Ricci tensor vanishes everywhere. Manifolds with these properties are called Calabi-Yau manifolds [29]. Unfortunately, in compactifications on Calabi-Yau manifolds, the computation of important quantities and properties of the resulting four-dimensional models can be very difficult (if not impossible).

An option to circumvent this problem is provided by compactifying on orbifolds, which will be the main focus of this thesis. Orbifold compactifications [30–34] are very similar to Calabi-Yau manifolds in the sense that both of them can lead to supersymmetric four-dimensional theories. Orbifolds are defined to be the quotient of a six-dimensional torus divided by a discrete set of its isometries. In comparison to Calabi-Yau manifolds, the advantage of orbifolds is that these are Riemann flat, with the exception of a finite set of points, where the curvature of the space concentrates. Therefore, the metric, which for (almost) all Calabi-Yau manifolds is still unknown, can be easily computed in orbifolds. Moreover, in these constructions it is comparatively straightforward to investigate phenomenological properties, such as the low–energy gauge symmetry, the particle spectrum, the Yukawa couplings and the Kähler potential, among others. For these reasons, orbifold compactifications are a rich and natural source of inspiration for phenomenological investigations.

In **D**-brane world scenarios, D-branes play a crucial role. They are subspaces of the ten-dimensional spacetime on which open string can end. This property equip them in general with a nonabelian gauge symmetry. Filling the space with several stacks of D-branes intersecting at angles in the type II string theory can not only reproduce the gauge group of the MSSM, but also its matter content. Matter then lives on a four-dimensional hypersurface while the mediators of gravity propagate in the full ten dimensions. In that sense, brane world constructions do not really compactify the spacetime. Apart from a number of $ad\ hoc$ assumptions, such as the placement of the fields on the branes and the number of branes involved in the stacks, a disadvantage of these constructions compared to orbifold compactifications is that GUTs like SO(10) or E_6 cannot be realized in these setups.

The Road to the MSSM

Although there are only five consistent ten-dimensional string theories, there is a huge number of four-dimensional string compactifications [35, 36], too big to be ever computed precisely. This leads to the picture that string theory has a vast landscape of vacua [37, 38], and that the vacuum we are living in is, from a top-down point of view, as good as many others.

To obtain useful predictions from string theory one might employ the following strategy: first seek vacua that are consistent with observations and then study their properties.

Optimistically, one might then be able to identify certain characteristics common to all realistic vacua, which would lead to predictions. Even if this is not the case, one might still be able to assign probabilities to certain features, allowing to exclude some patches of the landscape on a statistical basis. It turns out, however, that the first step of obtaining realistic models is highly nontrivial. For instance, in the context of orientifolds of Gepner models, the fraction of models with the chiral matter content of the SM is about 10^{-12} [39–41]. Based on a statistical approach [42], the probability of getting something close to the MSSM in the context of intersecting D-branes in an orientifold background is about 10^{-9} [43–45], even if one allows for chiral exotics. However, some patches on the landscape remain still unexplored. This is the case of orbifold compactifications.

In this thesis, we will focus on orbifold compactifications of the heterotic string theory (for recent reviews, see [46–48]). As it was very early noticed, it is preferable to consider the heterotic string with $E_8 \times E_8$ gauge group since, on the one hand, it includes naturally the so-called chain of GUTS

$$G_{SM} \subset SU(5) \subset SO(10) \subset E_6 \subset E_8$$
 sometimes also labeled as "E₄" "E₅" E₆ E₈ (1.3)

and, on the other, the presence of spinors like the 16-plet of SO(10) is more frequent than in the SO(32) theory. This particular fact facilitates enormously the task of getting models with SM generations.

In spite of its relative simplicity, orbifolds have not been systematically studied yet. Furthermore, although it is known that models resembling the MSSM exist [34, 47, 49–51], attempts to accommodate simultaneously all properties of the MSSM in isolated models might fail. One is thus encouraged to turn to a more general strategy that might serve as guiding principle through this search.

In order to get phenomenological viable models from orbifold compactifications, one can draw on the insight gained from grand unified theories by introducing the concept of local GUTs [50, 52–54]. In scenarios with local GUTs, there are special (fixed) points in the internal space where the gauge symmetry is locally that of certain GUTs while the four-dimensional gauge symmetry is that of the SM(up to additional gauge factors that compose the hidden sector). If matter fields are localized at such special points, they form complete GUT representations. This applies, in particular, to a 16-plet of a local SO(10). On the other hand, bulk fields form incomplete GUT multiplets. In the particular case of SO(10), from the four-dimensional viewpoint, the localized states are complete matter generations whereas the bulk fields can adopt the form of, say, Higgs doublets (an incomplete representation of SO(10)). This might offer an intuitive explanation for the observed family structure of the SM and, at the same time, a solution to the so-called doublet-triplet splitting problem.

If there are orbifolds models that get close to the MSSM, addressing their phenomenological viability is of utmost importance. Different questions can be posed in this direction. One challenge is, for instance, to verify whether these models admit supersymmetric vacuum configurations, that is, whether a combination of fields can attain vacuum expectation values (VEVs), such that exotic particles that might be in the spectrum decouple along D-flat (and F-flat) directions. Vacuum configurations are further constrained by the requirement that the Fayet–Iliopoulos (FI) term produced by the anomalous gauge U(1), intrinsic to many orbifold models [55,56], cancels. If such a vacuum configuration is realized, the corresponding model is a "good" MSSM candidate, where SUSY is present at very high energies and has eventually to be broken.

There are different mechanisms to break supersymmetry in string theory, such as composite goldstinos, instantons and gaugino condensation. Perhaps the simplest and most beautiful method is gaugino condensation. The basic notion in this mechanism is that of asymptotic freedom of QCD. At a given (intermediate) scale, usually labeled by Λ , interactions of a particular gauge group become infinitely strong. This can happen in any nonabelian gauge group, including those that appear in the hidden sector of orbifold models. The only requisite is that the gauge theory be of pure Yang-Mills type, that is, that contain (almost) no matter (and that couple to gravity). In this scenario, a couple of gauginos λ and $\bar{\lambda}$ condense together and acquire a VEV $\langle \lambda \bar{\lambda} \rangle$ proportional to Λ^3 .

A remarkable observation is that gaugino condensation can induce breakdown of SUSY [57–59]. The gravitino mass is related to the gaugino condensation scale Λ by

$$m_{3/2} \sim \frac{\Lambda^3}{M_{\rm Pl}^2} \,,$$
 (1.4)

where $M_{\rm Pl}$ denotes the Planck mass scale $\sim 10^{18}$ GeV. Notice that if the condensation scale is $\sim 10^{13}$ GeV, we obtain SUSY breaking in the phenomenological favored interval, around 1 TeV. It might be, of course, interesting to investigate if such scenarios are possible in orbifold constructions.

Another essential property of the MSSM that a realistic model must exhibit is R-parity. It is attached to fundamental issues such as recognizing a suitable candidate for dark-matter and suppressing rapid proton decay. Obtaining a conserved R-parity in string constructions sets frequently an insurmountable hurdle, which must however be overcome in order to reach the MSSM. One possible route is provided by first identifying a $U(1)_{B-L}$ gauge symmetry and then giving VEVs to some SM singlets that break $U(1)_{B-L}$ to an unbroken discrete \mathbb{Z}_2 symmetry, frequently called family reflection symmetry or matter parity [60], which forbids dangerous R-parity violating operators of dimension three and four.

The smallness of neutrino masses may be also a powerful probe of string constructions. The seesaw mechanism [61–63] is perhaps the most attractive way to explain the smallness of the neutrino masses. Its essential ingredients are heavy Majorana neutrinos and their Yukawa couplings to the left–handed neutrinos. Difficulties arise when one attempts to embed this attractive mechanism in string theory [64]. One of the problems is that string GUTs usually do not allow for large GUT representations, which are frequently required for the seesaw mechanism. Other obstacles are posed by string selection rules for the superpotential couplings which forbid some of the necessary couplings in certain models [65]. For some string models one can obtain Dirac neutrino masses [66,67]. Recently, large Majorana masses have been obtained from instanton effects [68–70] as well as from higher order operators [71] and small Majorana masses from large volume compactifications [72]. However,

it turned out to be very difficult to obtain the standard seesaw mechanism in consistent string models which lead to the gauge group and the matter content of the MSSM.

Can string theory with $10^{(1)500}$ vacua make any prediction relevant for the LHC? There are certainly many arguments in favor and many more against this idea. However, like in every new theory, one has to work out all the details before being able to give a definitive answer. It is clear that the beauty and mathematical consistency of string theory is very appealing. And it is also true that in the past theories with such properties, like quantum mechanics, have brought surprising predictions that affect us in our everyday life.

The ultimate goal of physics beyond the SM is still to identify a theory that can reproduce our current knowledge and improve our understanding of physics. Unification provides doubtless a framework where physics beyond the SM takes an appealing form. Adopting this idea into more elaborated theories, such as orbifold compactifications of the heterotic string, can certainly shed light on some of the puzzles of contemporary science. In the present work, we utilize the beauty and mathematical consistency of string theory as a tool in order to build a bridge between confirmed or foreseeable physics and a theory possibly capable to describe all fundamental forces in a unified way.

Overview

To guide the reader through the present work, let us provide some details about the contents of the chapters included.

Chapter 2. After a brief introduction to the heterotic string, we proceed to explain the details of abelian orbifold compactifications of the heterotic string. This chapter intends to be as general as possible. Hence, we do not focus particularly on any of its two variants, $E_8 \times E_8$ and SO(32). We also consider both \mathbb{Z}_N and $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds on the same footing. The effect of the choice of the six-dimensional compactification lattice as well as the constraints on the orbifold parameters are given. We present then in all detail how to compute the massless matter spectrum of orbifold models with and without Wilson lines and illustrate the method with a simple example. Discrete torsion is introduced as an additional degree of freedom which, contrary to previous claims, can also appear in \mathbb{Z}_N orbifolds even in the absence of Wilson lines. We propose an interpretation of models with discrete torsion and a special type of gauge embeddings. To close the chapter, we provide a succinct discussion about Yukawa couplings on heterotic orbifolds. In particular, the selection rules determining the nonvanishing couplings are enumerated.

Chapter 3. Due to the enormous number of redundancies in string constructions, it is always recommendable to have a useful method to classify these constructions, that is, to determine the all different parameters leading to inequivalent orbifold models. We summarize in this chapter two methods. The first one goes by the name of *Dynkin diagram strategy* for it makes extensive use of the properties of the Dynkin diagram of a Lie algebra to determine all admissible gauge embeddings. The second method of classification consists in suggesting a suitable ansatz that describes all shifts and/or Wilson lines of a given order,

minimizing duplicities. We discuss their advantages and drawbacks and give examples of their application. Some of the topics discussed in this chapter were presented in refs. [73,74].

Chapter 4. We describe a general strategy we proposed in ref. [75] to obtain orbifold models that resemble the MSSM, using as guiding principle the concept of local GUTs. Our search is performed by compactifying the $E_8 \times E_8$ heterotic string on the \mathbb{Z}_6 -II orbifold, since it has shown to house some models with realistic properties. After providing the criteria comprising our search strategy, we analyze the results obtained. We find that our approach, as opposed to a random scan, is successful and that a considerable fraction of the models with SO(10) and E_6 local GUT structures passes our criteria. In particular, out of about 3×10^4 inequivalent models with two Wilson lines, 223 can serve as an ultraviolet completion of the SM. In comparison to similar searches in different scenarios, we observe that the probability to find models with realistic features here is much larger. We consider this to be one of the central results of this thesis.

Chapter 5. The study of some aspects of the phenomenology of our MSSM candidates is presented in this chapter, following our previous discussions from refs. [76–78]. To illustrate our results, we describe one characteristic model with the exact spectrum of the MSSM. In that example, we find that the top quark can acquire a heavy mass while the proton cannot decay due to dimension four operators at trilinear level. This means that proton decay is generically suppressed. We also note that all vectorlike exotics can acquire masses in a SUSY-preserving vacuum configuration. This model is used all through the chapter to illustrate other phenomenological properties also treated here. We point out that these qualities are shared by all models and therefore a statistical analysis can yield some string theoretical predictions about low-energy physics. In that spirit, we investigate the scale of SUSY breaking typically realized in orbifold models with realistic features. We find that this kind of models favors hidden sector gauge groups such as SU(4) and SO(8), implying low-energy SUSY breaking in the context of gaugino condensation. We continue then with the discussion of a crucial issue in any MSSM-like model: R-parity. Many models with realistic properties admit the definition of $U(1)_{B-L}$, whose spontaneous breakdown can yield matter parity and thereby proton stability. However, to avoid rapid proton decay due to dimension five operators is still a challenge beyond the scope of this thesis. We conclude this chapter with another interesting question, namely, the viability of the seesaw mechanism in string-derived models. We argue that, even though large representation such as $\overline{126}$ of SO(10) do not appear in this class of models, all the necessary ingredients for the seesaw mechanism are present. These results are very important for future investigations and, therefore, we would like to stress them as our main findings.

Publications

Some results of this work have been published in

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- O. Lebedev, H. P. Nilles, S. Raby, S. Ramos-Sánchez, M. Ratz, P. K. S. Vaudrevange and A. Wingerter, "The Heterotic Road to the MSSM with R parity," Phys. Rev. D 77 (2008) 046013 [arXiv:0708.2691 [hep-th]].
- W. Buchmüller, K. Hamaguchi, O. Lebedev, S. Ramos-Sánchez and M. Ratz, "Seesaw neutrinos from the heterotic string," Phys. Rev. Lett. **99**, 021601 (2007) [arXiv:hep-ph/0703078].
- F. Plöger, S. Ramos-Sánchez, M. Ratz and P. K. S. Vaudrevange, "Mirage Torsion," JHEP **0704**, 063 (2007) [arXiv:hep-th/0702176]
- O. Lebedev, H. P. Nilles, S. Raby, S. Ramos-Sánchez, M. Ratz, P. K. S. Vaudrevange and A. Wingerter, "Low Energy Supersymmetry from the Heterotic Landscape," Phys. Rev. Lett. **98**, 181602 (2007) [arXiv:hep-th/0611203]
- O. Lebedev, H. P. Nilles, S. Raby, S. Ramos-Sánchez, M. Ratz, P. K. S. Vaudrevange and A. Wingerter, "A mini-landscape of exact MSSM spectra in heterotic orbifolds," Phys. Lett. B **645**, 88 (2007) [arXiv:hep-th/0611095]
- H. P. Nilles, S. Ramos-Sánchez, P. K. S. Vaudrevange and A. Wingerter, "Exploring the SO(32) heterotic string," JHEP **0604**, 050 (2006) [arXiv:hep-th/0603086]

Chapter 2

Orbifold Compactifications

In this chapter, we study the heterotic string theory compactified on orbifolds. We start by reviewing briefly the heterotic string. Then we explain one method to get four-dimensional supersymmetric models based on abelian orbifold compactifications. Our discussion is abstract at some level, but it is addressed to people willing to get acquainted with orbifold constructions. We also introduce here the notation to be used along the entire work.

2.1 Heterotic String

It is well known that in closed-string theories left- and right-moving modes are decoupled [21, 22]. This offers the possibility of a new consistent string theory in which left- and right-movers are of different types. The heterotic string arises as the result of combining a ten-dimensional right-moving superstring [79, 80] (ensuring thereby space-time supersymmetry) with a 26-dimensional left-moving bosonic string.

Right-movers

The right-moving bosonic and fermionic degrees of freedom of the superstring are denoted by X_R^i and Ψ_R^i , respectively, where $i=1,\ldots,10$. We can assume that the first four coordinates correspond to the observed minkowskian spacetime. This situation is depicted in fig. 2.1. Since not all these degrees of freedom are independent, we choose the light-cone gauge, in which the coordinates corresponding to i=1,2 are fixed.

The masses of the right-moving states are given by

$$\frac{m_R^2}{4} = N_R - a_R \,, \tag{2.1}$$

where N_R counts the oscillator excitations of the states, and a_R is the zero point energy, which turns out to be 0 for states with (periodic) Ramond boundary conditions, and $\frac{1}{2}$ for states with (antiperiodic) Neveu-Schwarz boundary conditions. We are interested in physical states that are massless at the string scale. From eq. (2.1), we observe that

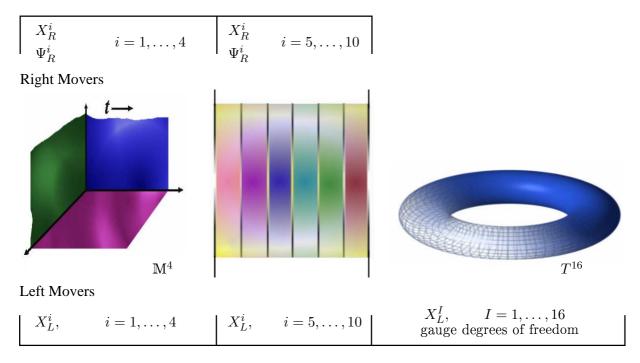


Figure 2.1: The space of the heterotic string. The supersymmetric right-movers live in 10 dimensions whereas the bosonic left-movers, in 26 dimensions. The 16 additional bosonic coordinates compactified on a torus T^{16} give rise to the gauge degrees of freedom. The first four dimensions of the ten dimensional heterotic string correspond to the Minkowski space \mathbb{M}^4 .

 $m_R = 0$ for states $b^i_{-1/2}|0\rangle_R$ in the Neveu-Schwarz sector, and for $d^i_0|0\rangle_R$ in the Ramond sector. The eight transverse excitations $b^i_{-1/2}|0\rangle_R$ form the vectorial representation $\mathbf{8}_v$ of SO(8) and behave as bosons from the spacetime perspective.

On the other hand, since the oscillators d_0^i obey the (Clifford) algebra

$$\left\{ i\sqrt{2}d_0^i, i\sqrt{2}d_0^j \right\} = 2\eta^{ij},$$
 (2.2)

the ground state $d_0^i|0\rangle_R$ forms a spinorial representation with 16 real components. In order to match the number of on-shell fermionic and bosonic degrees of freedom, one has to introduce a GSO projection [17], which does not only reduce the number of massless degrees of freedom by a factor $\frac{1}{2}$, but also ensures an equal number of bosons and fermions at each mass level, as required by supersymmetry. After the GSO projection, we are then left with the $\mathbf{8}_s$ representation of SO(8) in the Ramond sector.

For convenience, $|q\rangle_R$ will denote the right-moving ground state for both sectors, where q stands for the weights of the corresponding SO(8) representation in Cartan-Weyl labels

$$|q\rangle_R = \begin{cases} \left| \underline{\pm 1, 0, 0, 0} \right\rangle_R & \sim \mathbf{8_v}, & \text{NEVEU-SCHWARZ} \\ \left| \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2} \right\rangle_R & \sim \mathbf{8_s}, & \text{RAMOND} \end{cases}$$
(2.3)

where the spinor representation has an even number of plus signs. The underline denotes permutation of the entries. In this notation, eq. (2.1) becomes

$$\frac{m_R^2}{4} = \frac{1}{2}q^2 - \frac{1}{2} \,. \tag{2.4}$$

Left-movers

The coordinates of the bosonic string are denoted by X_L^i , $i = 1, \ldots, 10$, and X_L^I , I = $1, \ldots, 16$. As for the right-movers, the coordinates corresponding to i = 1, 2 are fixed by the light-cone gauge. The evident mismatch in the number of spatial dimensions of leftand right-movers is amended by compactifying the coordinates X_L^I on a 16-dimensional torus T^{16} with radii as small as the string scale, as illustrated in fig. 2.1. As a result of this compactification, the 16-dimensional internal momenta p are nonvanishing and proportional to the winding of states in the compactified space.

The masses of the left-moving states after the compactification are given by

$$\frac{m_L^2}{4} = \frac{1}{2}p^2 + \widetilde{N} - 1 \,\,\,\,(2.5)$$

where \widetilde{N} counts left-moving oscillator excitations and -1 is the zero point energy of the bosonic string. Each compactified dimension contributes with $\frac{1}{2}(p^I)^2$ to the mass.

Gauge and gravitational anomaly cancellation is guaranteed by one-loop modular invariance. It imposes severe constraints on the theory. In particular, the underlying lattice Λ of the 16-dimensional torus T^{16} must be euclidean, even and self-dual. There are only two such lattices in 16 dimensions: the root lattice of $E_8 \times E_8$ and the weight lattice of $\mathrm{Spin}(32)/\mathbb{Z}_2$. Consequently, the (nonabelian) gauge group \mathcal{G} of rank 16 provided by the compactification can be either $E_8 \times E_8$ or SO(32), depending on the choice of Λ . Each lattice yields an independent consistent heterotic string theory.

According to eq. (2.5), at the massless level we have the following left-moving states:

$$\widetilde{\alpha}_{-1}^i|0\rangle_L \qquad i=3,\ldots,10 \,,$$
 (2.6a)

$$\widetilde{\alpha}_{-1}^{i}|0\rangle_{L}$$
 $i=3,\ldots,10$, (2.6a)
 $\widetilde{\alpha}_{-1}^{I}|0\rangle_{L}$ $I=1,\ldots,16$, (2.6b)

$$|p\rangle_L \qquad p^2 = 2. \tag{2.6c}$$

There are 480 internal momenta p fulfilling $p^2 = 2$. As we will see below, it is not a coincidence that the adjoint representation of \mathcal{G} contains 480 charged bosons, too. In fact, the states $|p\rangle_L$ correspond to the left-moving part of the gauge bosons (and gauginos) of this theory. For that reason, we will represent p by the vectors in Cartan-Weyl labels for the corresponding charged bosons

$$\mathcal{G} = E_8 \times E_8: \qquad p \in \begin{cases} \frac{(\pm 1)^2, 0^6}{(0^8)} (0^8), & (0^8) (\pm 1)^2, 0^6, \\ ((\pm \frac{1}{2})^8) (0^8), & (0^8) ((\pm \frac{1}{2})^8), & \text{even } \# \text{ of } +, \end{cases}$$

$$\mathcal{G} = SO(32): \qquad p \in (\pm 1)^2, 0^{14}, \qquad (2.7b)$$

$$\mathcal{G} = SO(32): \qquad p \in \left(\underline{(\pm 1)^2, 0^{14}}\right), \qquad (2.7b)$$

where the exponent of an entry counts the number of times that such an entry appears in the 16-dimensional vector, and the underline stands, as before, for all permutations.

Massless Heterotic Spectrum

Let us analyze the spectrum of the heterotic string. Physical states must fulfill the level matching condition

$$m_R^2 = m_L^2$$
, (2.8)

which follows from the variation of the worldsheet metric. This constraint implies that, in contrast to other string theories, in the heterotic string the GSO projection is not implemented in order to avoid the presence of tachyons. Since the mass of the lowest energy left-moving state ($\tilde{N}=0,\ p=0$) is $m_L^2=-1$ whereas the mass of the right-moving (Neveu-Schwarz) tachyon ($N_R=0$) is $m_R^2=-\frac{1}{2}$, then eq. (2.8) enforces the absence of states with negative mass square in the spectrum of the heterotic string.

At the massless level, combining the right and left-moving states of eqs. (2.3) and (2.6) gives rise to the following states:

• a ten-dimensional $\mathcal{N}=1$ supergravity multiplet

$$|q\rangle_R \otimes \widetilde{\alpha}_{-1}^j |0\rangle_L , \qquad i = 3, \dots, 10 ,$$
 (2.9)

including the graviton g^{ij} , the dilaton ϕ , the antisymmetric tensor B^{ij} and their SUSY partners;¹

• 16 uncharged gauge bosons (and gauginos)

$$|q\rangle_R \otimes \widetilde{\alpha}_{-1}^I |0\rangle_L , \qquad I = 1, \dots, 16$$
 (2.10)

which comprise the set of Cartan generators H_I of the gauge group \mathcal{G} ;

• 480 charged gauge bosons (and gauginos)

$$|q\rangle_R \otimes |p\rangle_L$$
, $p^2 = 2$, (2.11)

with p given by eq. (2.7).

Uncharged and charged gauge bosons together form the 496-dimensional adjoint representation² of the gauge group \mathcal{G} . We notice that the effective theory with the massless content provided before is $\mathcal{N}=1$ supergravity in ten dimensions coupled to Yang-Mills. Such a theory has a gravitational anomaly of 496 units, which can be cancelled by including the 496 gauginos that we have at hand. This cancellation is not surprising because the heterotic string is, by construction, modular invariant and thus anomaly free.

¹Notice that $|q\rangle_R$ represents both bosonic and fermionic degrees of freedom, according to eq. (2.3). Recall also that the (bosonic) q carries a spacetime index i.

²The adjoint representation of $E_8 \times E_8$ reads $(248, 1) \oplus (1, 248)$.

2.2 Compactification on Orbifolds

It is clear that the heterotic string by itself is not a theory which describes the observable universe. A striking difference of this theory with respect to the four-dimensional spacetime of everyday experience is that the heterotic string is ten-dimensional. Further, the gauge symmetry group \mathcal{G} is too big compared to the one of the SM or its supersymmetric version, the MSSM. Therefore, in order for the heterotic string theory to make contact with low–energy physics, one has to introduce a mechanism in the theory ensuring that the additional dimensions are as small as to escape detection.

Compactification of extra dimensions on circles and tori have been discussed since the pioneering works by Kaluza and Klein [12, 13]. However, if the heterotic string is compactified to four dimensions on a six-torus, the resulting theory is far from being phenomenologically acceptable. In torus compactifications no supersymmetry is broken, so one gets $\mathcal{N}=4$ Susy in four dimensions. In supersymmetric theories with $\mathcal{N}\geq 2$, both vector and matter supermultiplets transform according to the same gauge group representation. Hence, these theories have the undesirable feature of being nonchiral.

Orbifold compactifications [30,31] are much more attractive. Orbifolds are very similar to Calabi-Yau manifolds in the sense that both of them can lead to four-dimensional theories with $\mathcal{N}=1$. The advantage of orbifolds is that they are Riemann flat, with the exception of a finite set of points. Therefore, the metric, which for almost all Calabi-Yau manifolds is still unknown, can be easily computed in orbifolds. Consequently, in these constructions it is comparatively straightforward to investigate phenomenological properties, such as the low–energy gauge symmetry, the particle spectrum, the Yukawa couplings and the Kähler potential, among others.

In general, an orbifold is defined to be the quotient of a manifold by a discrete set of its isometries, called the point group P. The simplest example is a one dimensional circle S^1 divided by the point group $P = \mathbb{Z}_2$. As illustrated in fig. 2.2, the points x and -x are identified by \mathbb{Z}_2 . This identification originates the space of points describing the complete orbifold to lie on the interval $x \in [0, \pi]$ (i.e. one half of S^1), which we shall call the fundamental region of the orbifold. A special feature of orbifolds is the appearance of singularities. In the case of S^1/\mathbb{Z}_2 , the points x = 0 and $x = \pi$ are left invariant (or fixed) by the action of \mathbb{Z}_2 . Although not obvious in our one dimensional example, points left invariant under any nontrivial element of P map to singular points of the orbifold.

2.2.1 Toroidal Orbifolds

To construct a heterotic orbifold, one first compactifies six dimensions on a torus T^6 . The six-torus is understood as the quotient \mathbb{R}^6/Γ , where Γ is the lattice of a semisimple Lie algebra of rank 6. Points in \mathbb{R}^6 differing by lattice vectors are identified on the torus, i.e.

$$X^i \sim X^i + n_\alpha e^i_\alpha, \qquad n_\alpha \in \mathbb{Z}, \ i = 5, \dots, 10,$$
 (2.12)

where e_{α} denote the basis vectors of the lattice Γ .

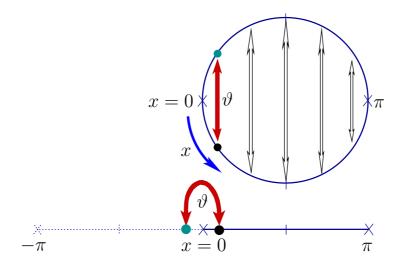


Figure 2.2: Compactification of one dimension on an \mathbb{Z}_2 orbifold. Points in the compactified dimension S^1 are identified by the \mathbb{Z}_2 point group element ϑ . Clearly, the points at x = 0 and $x = \pi$ are left fixed by the \mathbb{Z}_2 action. The fundamental region of S^1/\mathbb{Z}_2 is $x \in [0, \pi]$.

In a second step, one has to conceive a suitable point group P.³ Considering only abelian point groups, we are left with the cyclic groups

$$\mathbb{Z}_N = \left\{ \theta = \vartheta^k \mid 0 \le k < N \right\} \quad \text{and} \quad (2.13a)$$

$$\mathbb{Z}_N \times \mathbb{Z}_M = \left\{ \theta = \vartheta^k \omega^\ell \mid 0 \le k < N; \ 0 \le \ell < M \right\} , \qquad (2.13b)$$

where $\vartheta^N = \omega^M = \mathbb{1}$, M is an integer multiple of N and $k, \ell \in \mathbb{Z}$. The point group generators ϑ and ω , also called *twists*, are discrete rotation generators acting crystallographically on the torus lattice Γ . All point groups of these kinds have been already classified [30,83]. It is common to combine the action of the point group on Γ with the identification of points due to the torus compactification, eq. (2.12). The result is the so-called *space group*, defined by

$$S = P \ltimes \Gamma = \{ g = (\theta, n_{\alpha} e_{\alpha}) \mid \theta \in P, n_{\alpha} \in \mathbb{Z} \}, \qquad (2.14)$$

where the sum over α is understood. Due to the properties of the semidirect product ' \ltimes ', the multiplication of two space group elements is given by

$$g_1g_2 = (\theta_1, n_\alpha e_\alpha)(\theta_2, m_\alpha e_\alpha) = (\theta_1\theta_2, n_\alpha e_\alpha + m_\alpha \theta_1 e_\alpha) \in S$$
. (2.15)

One can also verify that the inverse of a space group element is given by

$$g^{-1} = (\theta, n_{\alpha}e_{\alpha})^{-1} = (\theta^{-1}, -n_{\alpha}\theta^{-1}e_{\alpha}).$$
 (2.16)

³In principle, one could freely choose two different tori, T_L^6 and T_R^6 , for left- and right-movers and, correspondingly, two different point groups, P_L and P_R . Orbifolds containing these elements are called asymmetric [81,82]. In the present work, however, we will focus on the case of symmetric orbifolds, where $T_L^6 = T_R^6$ and $P_L = P_R$.

Moreover, the action of a space group element $g = (\theta, n_{\alpha}e_{\alpha})$ on the six compact dimensions is provided by

$$X^i \xrightarrow{g} (\theta X)^i + n_\alpha e_\alpha^i, \qquad i = 5, \dots, 10.$$
 (2.17)

Modular invariance requires the action of the space group S to be embedded into the 16 gauge degrees of freedom

$$S \hookrightarrow G$$
, (2.18)

where the gauge twisting group G is, in general, a subgroup of the automorphisms of the $E_8 \times E_8$ or SO(32) Lie algebras. Space group elements are mapped to elements of the gauge twisting group, according to

$$\mathbb{Z}_{N} : (\vartheta^{k}, n_{\alpha}e_{\alpha}) \mapsto (kV, n_{\alpha}A_{\alpha})
\mathbb{Z}_{N} \times \mathbb{Z}_{M} : (\vartheta^{k}\omega^{\ell}, n_{\alpha}e_{\alpha}) \mapsto (kV_{1} + \ell V_{2}, n_{\alpha}A_{\alpha})
k, \ell, n_{\alpha} \in \mathbb{Z},$$
(2.19)

where the 16-dimensional shift vectors V, V_i parametrize the embedding automorphisms of the respective twists.⁴ The shifts A_{α} represent Wilson lines [32, 34, 85], i.e. they are gauge transformations associated to the noncontractible loops generated by e_{α} . An element of the gauge twisting group acts on the 16 gauge degrees of freedom of the heterotic string as

$$\mathbb{Z}_{N} : X^{I} \longrightarrow X^{I} + kV^{I} + n_{\alpha}A_{\alpha}^{I}
\mathbb{Z}_{N} \times \mathbb{Z}_{M} : X^{I} \longrightarrow X^{I} + kV_{1}^{I} + \ell V_{2}^{I} + n_{\alpha}A_{\alpha}^{I}$$

$$I = 1, \dots, 16. \qquad (2.20)$$

Finally, we are in position to define a heterotic orbifold. A heterotic orbifold is made up of the product of the quotient spaces of T^6/P and T^{16}/G :

$$\mathcal{O} = T^6/P \otimes T^{16}/G = \mathbb{R}^6/S \otimes T^{16}/G,$$
 (2.21)

where we have made use of the definition of the space group.

Space Group Conjugacy Classes

Not all elements of the space group describe a distinct action on the orbifold. A useful concept to gather those elements producing the same effect on the orbifold is that of conjugacy class. Two space group elements $g_1, g_2 \in S$ are conjugate if there exists another space group element $h \in S$, such that

$$hg_1h^{-1} = g_2 \quad \Leftrightarrow \quad g_1 \simeq g_2. \tag{2.22}$$

One says then that both g_1 and g_2 belong to the same conjugacy class [g], defined by

$$[g] = \{ h g h^{-1} \mid h \in S \} , \qquad (2.23)$$

and are therefore equivalent.

⁴It is possible to embed the action of a space group element (θ, e_{α}) as (Θ, A_{α}) , where Θ denotes a rotation in the gauge degrees of freedom [33,84].

With help of eqs. (2.15) and (2.16), one can easily verify that conjugation of $g = (\theta, n_{\alpha}e_{\alpha})$ under an arbitrary group element $h = (\theta_h, m_{\alpha}e_{\alpha})$ yields

$$h g h^{-1} = (\theta_h, m_\alpha e_\alpha) (\theta, n_\alpha e_\alpha) (\theta_h^{-1}, -m_\alpha \theta_h^{-1} e_\alpha)$$

= $(\theta, n_\alpha \theta_h e_\alpha + (1 - \theta) m_\alpha e_\alpha).$ (2.24)

Thus, the conjugacy class (2.23) of a general space group element g becomes

$$[q] = \{ (\theta, n_{\alpha}\theta_{h}e_{\alpha} + (\mathbb{1} - \theta)m_{\alpha}e_{\alpha}) \mid \theta, \theta_{h} \in P \}. \tag{2.25}$$

It is evident that there are several different conjugacy classes. It is convenient to organize them into two categories: a) $\theta = 1$ and b) $\theta \neq 1$. The conjugacy classes of the former case compose the so-called *untwisted sector*, denoted U. Those of the second case constitute one or more *twisted sectors*, denoted $T_{k,\ell}$, depending on the number of nontrivial θ s available. The origin and meaning of the tags *untwisted* and *twisted* will be clarified in section 2.3.

Notice that the elements of the untwisted sector are just lattice translations, $(1, n_{\alpha}e_{\alpha})$. The corresponding conjugacy classes, according to eq. (2.25), acquire then the form

$$[(1, n_{\alpha}e_{\alpha})] = \{(1, n_{\alpha}\theta_{h}e_{\alpha}) \mid \theta_{h} \in P\}. \tag{2.26}$$

2.2.2 Consistency Conditions

We have seen in section 2.1 that the heterotic string has intrinsic theoretical constraints on its geometry and spectrum. It is then natural to expect some requirements for the needed parameters in orbifold compactifications. These constraints fall into three classes:

- $\mathcal{N} = 1$ SUSY;
- $S \hookrightarrow G$ embedding conditions; and
- modular invariance.

$\mathcal{N}=1$ SUSY

Compactification of the heterotic string is supposed to provide a theory capable of describing, among other things, physics at low energies. As we have mentioned, theories with $\mathcal{N} > 1$ are nonchiral and thus unrealistic. On the other hand, models with no supersymmetry are phenomenologically disfavored as they cannot alleviate some of the fundamental puzzles of the standard model. Thus, a natural phenomenological requirement for orbifold models is to have $\mathcal{N} = 1$ SUSY. This can be guaranteed by an appropriate choice of the point group P.

In abelian orbifolds, the point group is, in general, a subset of the full holonomy group (i.e. the group of rotations in six dimensions), $P \subset O(6)$. Insisting on $\mathcal{N} = 1$ reduces the holonomy group to SU(3), yielding therefore

$$P \subset SU(3) \subset O(6)$$
. (2.27)

Point group P	6D Lattice Γ	Twist vector v
\mathbb{Z}_3	$SU(3)^3$	$\frac{1}{3}(0, 1, 1, -2)$
\mathbb{Z}_4	$SU(4)^2$	
	$SO(5) \times SU(4) \times SU(2)$	$\frac{1}{4}(0, 1, 1, -2)$
	$SO(5)^2 \times SU(2)^2$	
\mathbb{Z}_6 -I	$G_2^2 \times SU(3)$	$\frac{1}{6}(0, 1, 1, -2)$
\mathbb{Z}_6 -II	$G_2 \times SU(3) \times SU(2)^2$	$\frac{1}{6}(0, 1, 2, -3)$
	$SU(6) \times SU(2)$	$\frac{1}{6}(0, 2, 1, -3)$
	$SU(3) \times SO(8)$	
	$SU(3) \times SO(7) \times SU(2)$	
\mathbb{Z}_7	SU(7)	$\frac{1}{7}(0, 1, 2, -3)$
\mathbb{Z}_8 -I	$SO(9) \times SO(5)$	$\frac{1}{8}(0, 1, -3, 2)$
$\mathbb{Z}_8 ext{-II}$	$SO(10) \times SU(2)$	$\frac{1}{8}(0, 1, 3, -4)$
	$SO(9) \times SU(2)^2$	
\mathbb{Z}_{12} -I	E_{6}	$\frac{1}{12}(0, 1, -5, 4)$
	$F_4 \times \mathrm{SU}(3)$	
$\mathbb{Z}_{12} ext{-II}$	$F_4 \times \mathrm{SU}(2)^2$	$\frac{1}{12}(0, 1, 5, -6)$

Table 2.1: Admissible six-dimensional crystallographic lattices [86] and twist vectors for \mathbb{Z}_N orbifolds. The order of the sublattices in each Γ coincides with the order of the components of the corresponding twist vector.

Allowed point groups are the cyclic groups \mathbb{Z}_N , with N=3,4,6,7,8,12, and $\mathbb{Z}_N \times \mathbb{Z}_M$, with N, M=2,3,4,6 and M an integer multiple of N. As we are about to see, it turns out that the specific form of the twist ϑ is also constrained by $\mathcal{N}=1$.

To simplify the notation, the six compact coordinates of the torus X^i , i = 5, ..., 10, are conveniently combined into three complex coordinates

$$Z^{1} = \frac{1}{\sqrt{2}} \left(X^{5} + iX^{6} \right) , \quad Z^{2} = \frac{1}{\sqrt{2}} \left(X^{7} + iX^{8} \right) , \quad Z^{3} = \frac{1}{\sqrt{2}} \left(X^{9} + iX^{10} \right) .$$
 (2.28)

In the light-cone gauge, the observable spacetime can be represented by $Z^0 = \frac{1}{\sqrt{2}} (X^3 + iX^4)$.

Further, we will assume the torus T^6 to be factorizable,⁵ i.e. it can be written as $T^6 = T^2 \otimes T^2 \otimes T^2$. In the basis (2.28), the twist of a \mathbb{Z}_N orbifold is then a diagonal 3×3 -matrix of the form

$$\vartheta = \operatorname{diag}\left(e^{2\pi i v^{1}}, e^{2\pi i v^{2}}, e^{2\pi i v^{3}}\right).$$
 (2.29)

⁵Orbifolds in nonfactorizable tori have been also extensively studied. For details, see e.g. [87, 88]. A possible relation between orbifolds on factorizable and nonfactorizable lattices was first conjectured in [74].

Point group P	Twist vector v_1	Twist vector v_2
$\mathbb{Z}_2 imes \mathbb{Z}_2$	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{2}(0, 0, 1, -1)$
$\mathbb{Z}_2 imes \mathbb{Z}_4$	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{4}(0, 0, 1, -1)$
$\mathbb{Z}_3 \times \mathbb{Z}_3$	$\frac{1}{3}(0, 1, 0, -1)$	$\frac{1}{3}(0, 0, 1, -1)$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -I	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -II	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{6}(0, 1, 1, -2)$
$\mathbb{Z}_4 imes \mathbb{Z}_4$	$\frac{1}{4}(0, 1, 0, -1)$	$\frac{1}{4}(0, 0, 1, -1)$
$\mathbb{Z}_3 \times \mathbb{Z}_6$	$\frac{1}{3}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$
$\mathbb{Z}_6 imes \mathbb{Z}_6$	$\frac{1}{6}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$

Table 2.2: Twist vectors for $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds leading to $\mathcal{N} = 1$. Six-dimensional compactification lattices can be found in e.g. [87,88] (see also appendix D).

where $v = (0, v^1, v^2, v^3)$ is called the *twist vector* and carries the full information of the point group action.⁶ The action of the point group on the complex compact coordinates is then given by

$$Z^a \xrightarrow{\vartheta} \exp\{2\pi i v^a\} Z^a, \qquad a = 1, 2, 3,$$
 (2.30)

where we can include the trivial action of the twist on the observable spacetime plane Z^0 . In $\mathcal{N}=1$ SUSY there is only one gravitino. So, we need to determine what the form of the twist vector must be so as to ensure that only one gravitino survives after compactification. The ten-dimensional gravitino contained in the supergravity multiplet (2.9) of the heterotic string splits into four gravitini in four dimensions:

$$\left| \pm \frac{1}{2}; \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2} \right\rangle_R \otimes \widetilde{\alpha}_{-1}^{\nu} \left| 0 \right\rangle_L, \quad \text{even } \# \text{ of } + \text{ signs }.$$
 (2.31)

The first component of the right-mover provides both chiralities for the gravitini. Due to the compactification, the last three components are internal indices that account for a multiplicity factor of four for each chirality. Since the left-mover carries lorentzian index, only the right-mover transforms under the orbifold action

$$\left| \pm \frac{1}{2}; \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2} \right\rangle_{R} \stackrel{\vartheta}{\longrightarrow} e^{-2\pi i \left(\pm \frac{1}{2}v^{1} \pm \frac{1}{2}v^{2} \pm \frac{1}{2}v^{3} \right)} \left| \pm \frac{1}{2}; \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2} \right\rangle_{R}. \tag{2.32}$$

The spectrum of an orbifold contains only states that are invariant under the orbifold action. Thus, in order to get $\mathcal{N}=1$, the phase $\exp\{-\pi i (\pm v^1 \pm v^2 \pm v^3)\}$ must be trivial for one gravitino. One notices that the condition $\pm v_1 \pm v_2 \pm v_3 = 0$ for one combination of signs assures the presence of solely one gravitino in the orbifold spectrum. Therefore, one can choose the components of the twist vector to satisfy

$$v^1 + v^2 + v^3 = 0, (2.33)$$

⁶In $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds, there is one twist vector for each of the two point group generators, ϑ and ω .

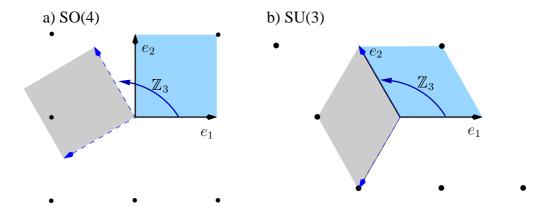


Figure 2.3: A two-dimensional \mathbb{Z}_3 transformation a) does not act crystallographically on the root lattice of SO(4). In contrast, b) the root lattice of SU(3) is mapped to itself under \mathbb{Z}_3 .

so that $|\pm(\frac{1}{2}; \frac{1}{2}, \frac{1}{2})\rangle_R \otimes \widetilde{\alpha}_{-1}^{\nu}|0\rangle_L$ be the surviving gravitino. Moreover, as the twist vector v of \mathbb{Z}_N orbifolds corresponds to a twist $\vartheta \in P$ of order N ($\vartheta^N = \mathbb{1}$), then its components fulfill in general $Nv^a \in \mathbb{Z}$, a = 1, 2, 3. Note that if one v^a is zero (or integer), we obtain $\mathcal{N} = 2$ in four dimensions, implying that all components v^a of a \mathbb{Z}_N orbifold twist vector must be nontrivial. In the third column of table 2.1, we present our choice of twist vectors for all admissible \mathbb{Z}_N orbifolds.

In $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds, there are two twist vectors, v_1 and v_2 , that also fulfill eq. (2.33). Only their combined action must lead to $\mathcal{N} = 1$. In this case, one can choose the twist vectors as shown in table 2.2. Notice that each of the twist vectors leads to $\mathcal{N} = 2$, but their conjoint action provides a theory with $\mathcal{N} = 1$, as required.

A secondary effect of the restriction $\mathcal{N}=1$ occurs in the six-dimensional compactified space. As already mentioned, the point group P has to act crystallographically on the root lattice Γ of a six-dimensional Lie algebra. In other words, the lattice of the torus T^6 must be mapped to itself under the action of P. Provided a point group P, not any root lattice is admissible. As an example, consider fig. 2.3: in two dimensions, the action of \mathbb{Z}_3 can be regarded as a rotation by $2\pi/3$. Clearly, whereas the root lattice of SU(3) is left invariant under \mathbb{Z}_3 , the root lattice of SO(4) is not. Suitable root lattices for \mathbb{Z}_N orbifolds are given in table 2.1. We have also compiled a list of some allowed lattices for $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds in table D.2.

$S \hookrightarrow G$ Embedding Conditions

The embedding of the space group into the gauge degrees of freedom imposes some conditions on the shift vector(s) V and the Wilson lines A_{α} . First of all, embedding the twist satisfying $\vartheta^{N} = 1$ implies that the associated shift vector has to be also of order N, i.e.

$$NV \in \Lambda$$
, (2.34)

where Λ is the 16-dimensional weight lattice of the corresponding heterotic string.

On the other hand, Wilson lines are subject to certain conditions depending on the compactification lattice Γ . Wilson lines are the embeddings of the six lattice generators e_{α} , therefore one might be enticed to think that there are six distinct Wilson lines A_{α} . However, not all directions e_{α} are independent in the orbifold. Consider for example the two-dimensional SU(3) lattice of fig. 2.3. One sees that a \mathbb{Z}_3 transformation maps e_1 to e_2 , implying that the corresponding Wilson lines A_1 and A_2 have to be equivalent in the orbifold. Consequently, in \mathbb{Z}_3 orbifolds with compactification lattice $\Gamma = \mathrm{SU}(3)^3$ (cf. table 2.1), one finds

$$e_{\alpha} \xrightarrow{\vartheta} e_{\alpha+1} \implies A_{\alpha} \approx A_{\alpha+1}, \quad \alpha = 1, 3, 5; \ \vartheta \in \mathbb{Z}_3,$$
 (2.35)

where ' \approx ' indicates that the Wilson lines are identical up to lattice translations in Λ .

In general, relations between the Wilson lines are effect of equivalences between space group elements. As we have seen before, elements of a given conjugacy class are equivalent. According to eq. (2.26), the conjugacy class of a lattice translation $g = (1, n_{\alpha}e_{\alpha})$ is given by

$$[(1, n_{\alpha}e_{\alpha})] = \{(1, n_{\alpha}\theta e_{\alpha}) \mid \theta \in P\}$$

$$(2.36)$$

and contains elements describing the same orbifold action. For example, in the case of \mathbb{Z}_3 orbifolds, the relation

$$(1, e_1) \simeq (\vartheta, 0)(1, e_1)(\vartheta, 0)^{-1} = (1, e_2),$$
 (2.37)

implies that $(1, e_1)$ and $(1, e_2)$ are indistinguishable from the orbifold perspective. Then, their embedding into the gauge degrees of freedom should also be identified. This restricts A_1 and A_2 to be equal up to lattice vectors, as we had already shown.

It is not hard to realize that also the order of the Wilson lines gets restricted by the choice of Γ . Consider in the \mathbb{Z}_3 orbifold the following elements of the same conjugacy class:

$$(1, e_2) \simeq (\vartheta, 0)(1, e_2)(\vartheta, 0)^{-1} = (1, -e_1 - e_2).$$
 (2.38)

Embedding this relation into the gauge degrees of freedom and using $A_1 \approx A_2$ yields

$$A_2 \approx -A_1 - A_2 \approx -2A_2 \quad \Leftrightarrow \quad 3A_2 \approx 0.$$
 (2.39)

In other words, A_2 (as well as A_1) has to be a Wilson line of order 3, $3A_2 \in \Lambda$. Similar relations apply also for the other two SU(3) factors of the lattice Γ . Hence, all in all, we obtain

$$3A_1 \approx 3A_3 \approx 3A_5 \approx 0$$
; $A_1 \approx A_2$, $A_3 \approx A_4$, $A_5 \approx A_6$. (2.40)

The number of independent Wilson lines A_{α} and their order N_{α} for admissible choices of Γ in \mathbb{Z}_N orbifolds are provided in table D.1.⁷ See table D.2 for constraints on Wilson lines of $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds.

⁷Notice that several typos of the literature have been corrected there.

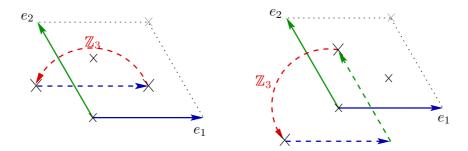


Figure 2.4: Fixed points arise naturally by the action of a discrete symmetry on the torus. In an SU(3) lattice, the action of \mathbb{Z}_3 is counteracted by lattice translations, leaving three points fixed.

Modular Invariance

Terms of the one-loop partition function of abelian orbifolds acquire in general a nontrivial phase under modular transformations [89, 90]. Demanding the partition function to be modular invariant safeguards the resulting theory from anomalies. Therefore, we are committed to requiring the phase that arises from modular transformations to vanish. This imposes constraints on the orbifold parameters which, for \mathbb{Z}_N orbifolds without Wilson lines, are usually expressed as [31, 90, 91]

$$N(V^2 - v^2) = 0 \mod 2. (2.41)$$

In including Wilson lines (and a second twist of the point group in the case of $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds), eq. (2.41) has to be replaced by [74]

$$N(V_1^2 - v_1^2) = 0 \mod 2, \qquad (2.42a)$$

$$M(V_2^2 - v_2^2) = 0 \mod 2,$$
 (2.42b)

$$M(V_1 \cdot V_2 - v_1 \cdot v_2) = 0 \mod 2, \qquad (2.42c)$$

$$N_{\alpha} (A_{\alpha} \cdot V_i) = 0 \mod 2, \qquad (2.42d)$$

$$N_{\alpha} \left(A_{\alpha}^{2} \right) = 0 \mod 2 \,, \tag{2.42e}$$

$$Q_{\alpha\beta} (A_{\alpha} \cdot A_{\beta}) = 0 \mod 2 \quad (\alpha \neq \beta),$$
 (2.42f)

where N_{α} corresponds to the order of the Wilson line A_{α} ($N_{\alpha}A_{\alpha} \in \Lambda$), and $Q_{\alpha\beta} \equiv \gcd(N_{\alpha}, N_{\beta})$ denotes the greatest common divisor of N_{α} and N_{β} . For \mathbb{Z}_N orbifolds, one has $V_1 = V$, $v_1 = v$ and eqs. (2.42c) and (2.42d) are clearly unnecessary.

2.2.3 Orbifold Geometry

Fixed points appear naturally in orbifold compactifications due to the action of the twist on the compact space. As an example, consider a one-complex-dimensional \mathbb{Z}_3 orbifold on

⁸In the case of two different \mathbb{Z}_2 Wilson lines we find that (2.42f) can be relaxed, i.e. $\gcd(N_{\alpha}, N_{\beta})$ can be replaced by $N_{\alpha} N_{\beta} = 4$, provided there exists no $g \in P$ with the property $g e_{\alpha} \neq e_{\alpha}$ but $g e_{\beta} = e_{\beta}$. Imposing the weaker condition leads, as we find, to anomaly-free spectra.

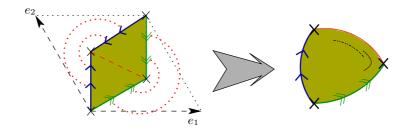


Figure 2.5: The fundamental region of a one-complex-dimensional \mathbb{Z}_3 orbifold is reduced to one third of the fundamental cell of the SU(3) root lattice. Folding appropriately the fundamental region, one can see that the \mathbb{Z}_3 orbifold is a flat manifold everywhere except at the three singularities corresponding to the fixed points.

an SU(3) torus lattice. \mathbb{Z}_3 acts as a rotation by $2\pi/3$ on the complex plane. Evidently, the point at the origin is not affected by the \mathbb{Z}_3 action and is therefore fixed. Furthermore, considering torus translations, one finds that there are two additional fixed points inside the fundamental cell of the torus. Observe the situation depicted in figure 2.4. The discrete rotation of those points is counteracted by translations in the torus lattice, so that the points remain unaffected in the orbifold.

Let us make two remarks. First, note that the three fixed points described above can be related neither by further lattice translations nor by the repeated action of \mathbb{Z}_3 . This independence characterizes all fixed points on the orbifold. Secondly, if one considers not only the fundamental cell of the torus illustrated in figure 2.4, but the entire (infinite) root lattice of SU(3), clearly, one finds an infinite set of fixed points. Yet all of them are identified to the three fixed points on the orbifold by the conjoint action of \mathbb{Z}_3 rotations and lattice translations, that is, by the action of the space group. Hence, the fixed points belong to only three different classes and, for describing all of them, it suffices to take one representative out of each of these classes.

Orbifolds are generalizations of manifolds in the sense that they are smooth (almost) everywhere, with exception of a constrained set of points, where the curvature concentrates; i.e. orbifolds admit singularities. That can be realized in the last example, by observing the space resulting after moding out the discrete group \mathbb{Z}_3 . In that case, the entire space can be described by one third of the area of the torus, the so-called fundamental region, as sketched in figure 2.5. Then, one notices that points along the boundaries of the fundamental region of the orbifold are identified under the action of the space group. To visualize the orbifold space, one has to fold the fundamental region and paste the edges together. The outcome is a triangular pillow-like object with sharp corners located at the fixed points. One can proof that such corners are conical singularities that concentrate the curvature of the orbifold. This means that only at the fixed points the holonomy group is nontrivial (generically, it is a subgroup of \mathbb{Z}_N).

⁹As we will see, there are situations in which some points are fixed under the action of one point group element, but are connected to other fixed points by the action of another one. In that case, not all fixed points in the torus are independent fixed points on the orbifold.

In any six-dimensional orbifold, fixed points are determined by considering the underlying lattice and the space group action. To be more precise, consider an arbitrary space group element $g = (\theta, n_{\alpha}e_{\alpha})$. Following from eq. (2.17), its action on the complex coordinates Z^a is given by

$$Z \stackrel{g}{\longrightarrow} \theta Z + n_{\alpha} e_{\alpha} \,, \tag{2.43}$$

where the basis vectors e_{α} of the six-torus are now expressed in the complexified basis eq. (2.28), and θ denotes an arbitrary point group element taking the form $\theta = \vartheta^k$ for \mathbb{Z}_N orbifolds or $\theta = \vartheta^k \omega^\ell$ for $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds. A point Z_f in the compact space is said to be fixed in the orbifold if it is invariant under the action of a particular space group element g_f . This means that fixed points satisfy

$$Z_f = g_f Z_f = \theta Z_f + n_\alpha e_\alpha \quad \Leftrightarrow \quad (\mathbb{1} - \theta) Z_f \in \Gamma \tag{2.44}$$

for a given $g_f \in S$.

It is convenient to label a fixed point by the corresponding space group element g_f instead of by its spatial coordinates Z_f^a . The space group elements g_f will be called constructing elements. This notation is rather convenient for several reasons. We know that, even though the number of solutions of eq. (2.44) is infinite, only a reduced finite number of points are inequivalent in the orbifold. In fact, inequivalent points in the compact space are related to space group elements from different conjugation classes. Secondly, we can say that points expressed by $g_f = (1, n_\alpha e_\alpha)$ belong to the untwisted sector. Further, fixed points represented by $g_f = (\theta, n_\alpha e_\alpha)$ with $\theta \neq 1$ 'live' in one of the twisted sectors. For example, fixed points left invariant under the action of $\theta = \vartheta^k \neq 1$ of a \mathbb{Z}_N orbifold are said to belong to the k-th twisted sector $(k = 1, \ldots, N - 1)$. Analogously, invariant points under $\theta = \vartheta^k \omega^\ell \neq 1$ in $\mathbb{Z}_N \times \mathbb{Z}_M$ are called fixed points of the (k, ℓ) -th twisted sector.

The number of distinct (conjugacy classes of) fixed points varies for different sectors of an orbifold. One first notices that in the untwisted sector, every point of the space is evidently invariant and $g_f = (1, 0)$. Thus, we end up with a six-dimensional fixed torus without singularities in the untwisted sector. A less boring situation appears in the twisted sectors. There, the solutions of eq. (2.44) are either isolated fixed points or one-complex-dimensional invariant surfaces, commonly called fixed tori. The former case applies to points fixed under θ such that $1 - \theta$ is nonsingular (i.e. $\det(1 - \theta) \neq 0$). The latter appears when $1 - \theta$ is singular. This is easy to understand because $1 - \theta$ is singular only if one of the eigenvalues of θ is one or, stated differently, only if one complex plane is left invariant under θ .

In case that $det(1 - \theta) \neq 0$ and the lattice of the compact space is factorizable, ¹⁰ the number of isolated fixed points in the twisted sector corresponding to θ is given by an (over)simplified version of the Lefschetz fixed point theorem [92]

$$\#Z_f = \det(\mathbb{1} - \theta) = \prod_{a=1}^3 4 \sin^2(\pi v^a),$$
 (2.45)

 $^{^{10}}$ i.e. the six-dimensional lattice can be written as the product of three two-dimensional sublattices, each of which corresponds to a complex plane with coordinate Z^a and a fixed.

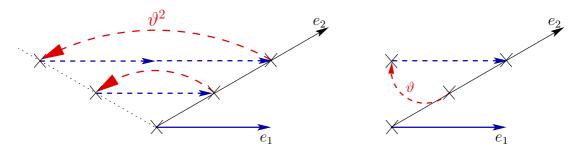


Figure 2.6: Fixed points of the ϑ^2 twisted sector on the torus spanned by the root lattice of G_2 . The point group generator ϑ acts as a rotation by $2\pi/6$. Some points invariant under ϑ^2 on the torus are identified on the orbifold by the action of ϑ .

where v^a are the entries of the twist vector. Formula (2.45) is, at first sight, very appealing, since it does not depend on the particular geometry of the underlying lattice Γ . Nonetheless, there are too few cases for which eq. (2.45) applies. For example, in most of the \mathbb{Z}_N orbifolds, only the number of fixed points in the first twisted sector ($\theta = \vartheta^1$) are determined by that formula. Few other twisted sectors of both \mathbb{Z}_N and $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds can also be addressed in this way.

In presence of fixed tori, that is, when $\det(1-\theta)=0$, one might conjecture that it suffices to extract the nontrivial two-complex-dimensional part of θ and then to apply formula (2.45). Unfortunately, the result obtained in that way is, in general, wrong. The reason can be traced back to the origin of formula (2.45). In a more complex version, the Lefschetz fixed point theorem¹¹ states that the number of fixed points (or fixed tori) is given by the index of the space of elements g_f (associated to the fixed points) divided by the largest set of (sub)symmetries of the space group relating fixed points among each other. Only if the set of modded out symmetries comprises exclusively lattice translations on the lattice Γ (and Γ is factorizable), then formula (2.45) (or a lower dimensional version of it) yields a correct result.

Let us examine more closely this situation in an example. In \mathbb{Z}_N orbifolds, $\mathbb{1} - \theta$ is singular only for higher twisted sectors corresponding to $\theta = \vartheta^k$ with 1 < k < N. Suppose ϑ to be a \mathbb{Z}_6 generator and the compact space to be spanned by the root lattice of $G_2 \times SU(3) \times SO(4)$. As it will be detailed shortly, the action of ϑ^2 is trivial in the sublattice spanned by SO(4), so that we get fixed tori in the second twisted sector. Now, focus on the sublattice spanned by G_2 . The point group generator ϑ acts as a rotation by $2\pi/6$ on the G_2 plane; therefore, fixed points (tori) of the ϑ^2 twisted sector are those points left invariant under a rotation by 120 degrees. In figure 2.6, we present the three points of the G_2 sublattice left invariant by this rotation. Notice that a one-complex-dimensional version of formula (2.45) also leads to three invariant points:

$$\#Z_f^{G_2} = \left| 1 - e^{2 \times 2\pi i/6} \right| = \left| \frac{3}{2} - i \frac{\sqrt{3}}{2} \right| = \frac{9}{4} + \frac{3}{4} = 3.$$
 (2.46)

¹¹See e.g. appendix A of ref. [93].

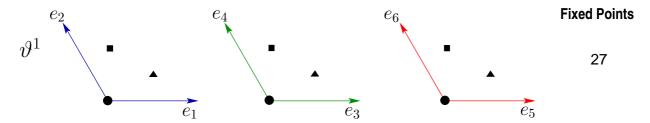


Figure 2.7: The geometry of \mathbb{Z}_3 orbifolds compactified on an SU(3)³ lattice. The 27 fixed points of the ϑ^1 and ϑ^2 twisted sectors are identical.

The points invariant under ϑ^2 in the G_2 sublattice are represented by the space group elements $g_1 = (\vartheta^2, 0)$, $g_2 = (\vartheta^2, e_1)$ and $g_3 = (\vartheta^2, 2e_1)$. Nevertheless, to conclude that there are three fixed points in the G_2 plane of the second twisted sector is wrong. It is not hard to verify that

$$g_3 = (\vartheta, e_1) g_2 (\vartheta, e_1)^{-1},$$
 (2.47)

situation that is also depicted in figure 2.6. Then g_2 and g_3 belong to the same conjugacy class, implying that only the conjugacy classes of g_1 and g_2 are independent. Therefore, only two points (tori) are truly fixed under the action of the orbifold in the G_2 sublattice.

Before proceeding to the details of two important examples, let us add a remark. The compactification lattice plays a very important role in the number of fixed points. For example, the usual $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifold compactification on the factorizable lattice of $SU(2)^6$ admits twice the amount of fixed tori of the same orbifold on the nonfactorizable root lattice of SO(12) [94]. In fact, it is possible to state that the number of fixed points (or tori) of orbifolds with factorizable compact space is, in general, bigger than that of the same orbifold with a nonfactorizable lattice. Our discussion here will restrict to the factorizable case.

Standard Example: The \mathbb{Z}_3 Orbifold

The \mathbb{Z}_3 orbifold has been long studied since the mid-eighties [30, 32, 95, 96] mainly because it is the simplest orbifold and because, even in that scope, there are chances to get semirealistic models [34, 49].

The compact space of the \mathbb{Z}_3 orbifold is spanned by the root lattice of $SU(3)^3$. The twist vector preserving $\mathcal{N}=1$ SUSY is given by

$$v = (0, 1/3, 1/3, -2/3),$$
 (2.48)

implying that the \mathbb{Z}_3 point group generator ϑ acts as a simultaneous rotation by $2\pi/3$ on all three SU(3) sublattices. The \mathbb{Z}_3 orbifold has three sectors: the untwisted sector ($\vartheta^0 = 1$) and two twisted sectors (ϑ^1 and ϑ^2). In the untwisted sector the action of the point group is trivial and, therefore, all points are invariant.

To find the points fixed under ϑ , we use the fact that the underlying lattice is factorizable. We have already studied the \mathbb{Z}_3 action on a single SU(3) lattice. We have seen

that three independent points are left invariant by \mathbb{Z}_3 . As an extension of that case, we find that each of the three sublattices has three independent fixed points, so that the six-dimensional space of the first twisted sector contains a total of 27 fixed points, displayed in figure 2.7. This result can be verified by using formula (2.45).

The fixed points of the ϑ^2 twisted sector are identical to those of the first twisted sector. Therefore, there is no need to consider these two sectors separately. We will see in section 2.3 that this structure will be reflected in the matter spectrum of the orbifold. In any orbifold, one can show that the fixed point structure of the sector θ is equal to that of the sector θ^{-1} . Note that in the \mathbb{Z}_3 case $\vartheta^2 = \vartheta^{-1}$.

The Z₆-II Orbifold

The \mathbb{Z}_6 -II orbifold was first studied in detail in ref. [47,97]. In those works, the structure of the fixed points was suggested as a tool to get models with phenomenologically acceptable features. Since the present thesis is based on the \mathbb{Z}_6 -II orbifold, we discuss in detail the structure of its fixed points. We illustrate our results in figure 2.8.

We will consider the compact space of the \mathbb{Z}_6 -II orbifold to be spanned by the factor-izable lattice¹²

$$\Gamma_{\mathbb{Z}_6-II} = G_2 \times SU(3) \times SO(4). \tag{2.49}$$

An advantage of the lattice $\Gamma_{\mathbb{Z}_6-II}$ being factorizable, is that we can find the fixed points independently for each sublattice without loss of information, and then put them all together in order to obtain the entire fixed point structure.

The point group \mathbb{Z}_6 -II is generated by ϑ which acts simultaneously as a rotation by $2\pi/6$ on the G_2 plane, a rotation by $2\pi/3$ on the SU(3) plane, and a reflection on the origin of the SO(4) sublattice. The action of ϑ is then described by the twist vector

$$v = (0, 1/6, 1/3, -1/2).$$
 (2.50)

There are five twisted sectors corresponding to the different powers of ϑ .

Let us consider the first twisted sector. Formula (2.45) tells us that there are 12 fixed points. Their precise location in the compact space can however only be found by means of eq. (2.44). The constructing elements corresponding to the fixed points are given by

$$g_f^{\vartheta} \in \{(\vartheta, n_5e_5 + n_6e_6), (\vartheta, e_3 + n_5e_5 + n_6e_6), (\vartheta, e_3 + e_4 + n_5e_5 + n_6e_6)\}$$
 (2.51)

with $n_5, n_6 = 0, 1$.

The action of ϑ^2 is encoded in 2v = (0, 1/3, 2/3, -1). This means particularly that all points of the SO(4) sublattice are left invariant. Therefore, the action of ϑ^2 on the compact space introduces fixed tori. In this case, formula (2.45) does not count correctly the number of fixed points. However, from previous discussions we know that, under a rotation of $2 \times 2\pi/3$ on the SU(3) sublattice, three points are left fixed. Further, we have

 $^{^{12}\}mathbb{Z}_6$ -II orbifolds on nonfactorizable lattices can also lead to interesting results. Some of their properties are briefly discussed in appendix C.

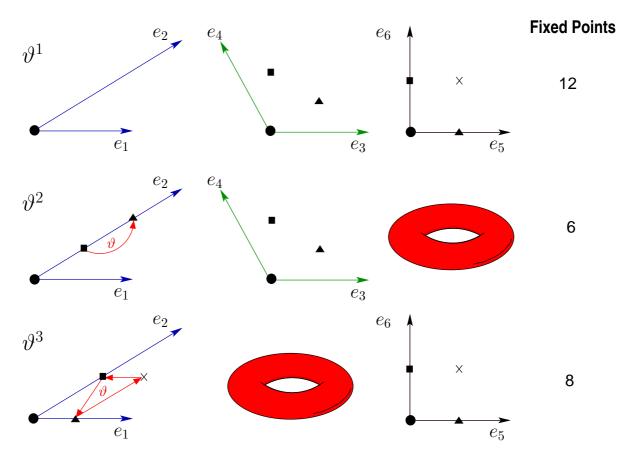


Figure 2.8: The geometry of \mathbb{Z}_6 -II orbifolds compactified on a $G_2 \times SU(3) \times SO(4)$ lattice. Fixed points of the ϑ^1 , ϑ^2 and ϑ^3 twisted sectors are presented. The fixed points of the ϑ^4 and ϑ^5 twisted sectors are identical to those of the ϑ^1 and ϑ^2 twisted sectors, respectively.

also seen that there are only two inequivalent fixed points in the G_2 sublattice. All in all, we find 2×3 fixed tori in the second twisted sector. The constructing elements are

$$g_f^{\vartheta^2} \in \{(\vartheta^2, 0), (\vartheta^2, e_1), (\vartheta^2, e_4), (\vartheta^2, e_1 + e_4), (\vartheta^2, e_3 + e_4), (\vartheta^2, e_1 + e_3 + e_4)\}.$$
 (2.52)

In the third twisted sector, the point group action, described by 3v = (0, 1/2, 1, -3/2), acts trivially in the SU(3) sublattice, hence, we obtain fixed tori in this case too. It is easy to see that eq. (2.44) leads to four invariant points under ϑ^3 in each of the other two sublattices. However, the points out of the origin of the G_2 sublattice are equivalent. Let us denote the associated space group elements by $g_1 = (\vartheta^3, e_1)$, $g_2 = (\vartheta^3, e_1 + e_2)$ and $g_3 = (\vartheta^3, e_2)$. It is easy to verify that

$$g_1 \simeq (\vartheta, e_1) \ g_1 \ (\vartheta, e_1)^{-1} = g_2$$
 and $g_2 \simeq (\vartheta, 2e_1 - e_2) \ g_2 \ (\vartheta, 2e_1 + e_2)^{-1} = g_3$, (2.53)

whence it follows that all three elements belong to the same conjugacy class. We are then left with a total of 2×4 fixed tori, described by the following constructing elements:

$$g_f^{\vartheta^3} \in \{(\vartheta^3, n_5e_5 + n_6e_6), (\vartheta^3, e_2 + n_5e_5 + n_6e_6)\}$$
 (2.54)

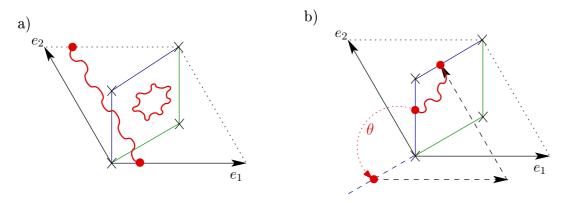


Figure 2.9: Closed strings in orbifold compactifications. a) Untwisted strings are closed on the torus, whereas b) twisted strings only close through the action of the twist θ .

with $n_5, n_6 = 0, 1$.

The fourth and fifth twisted sectors posses the structure of the second and first twisted sectors, respectively. Therefore, it is enough to study the three twisted sectors depicted in figure 2.8.

2.3 Strings on Heterotic Orbifolds

The final ingredient of orbifold compactifications is their spectrum of matter. Matter in heterotic orbifolds is described by closed strings. A special feature of orbifold compactifications is that they admit two types of closed strings: *untwisted* and *twisted* strings (see figure 2.9). Untwisted strings fulfill the following boundary conditions:

$$Z(\tau, \sigma + \pi) = Z(\tau, \sigma) + n_{\alpha}e_{\alpha}. \tag{2.55}$$

This indicates that they are closed already on the torus and free to propagate in the compact space. In contrast, twisted strings are closed only after identifying points by the twist θ , according to their boundary conditions

$$Z(\tau, \sigma + \pi) = \theta Z(\tau, \sigma) + n_{\alpha} e_{\alpha}, \qquad (2.56)$$

whence it follows that they are bound to the fixed points. The relation between the types of strings in an orbifold and the action of the point group is the reason why we have called untwisted to the space group elements with a trivial point group action, and twisted to those other which include a nontrivial element of the point group.

Since massive strings have masses of the order of M_{str} , they are too heavy to contribute directly to low–energy physics. Therefore, we will consider the spectrum of heterotic orbifolds to be composed only of massless string states, satisfying additionally the level matching condition for right- and left-movers, eq. (2.8).

Massless states differ for the several constructing elements g. On the one hand, untwisted states (with constructing element g = (1, 0)) are written in general as

$$|q\rangle_R \otimes \widetilde{\alpha}_{-1}^x |p\rangle_L \,,$$
 (2.57)

where q is a weight of SO(8) and $p \in \Lambda$ as given in eq. (2.7). $\widetilde{\alpha}_{-1}^x$ represents an oscillator excitation in one direction of either the Minkowski space $(\widetilde{\alpha}_{-1}^{\mu})$ or the compact space $(\widetilde{\alpha}_{-1}^{a}, \widetilde{\alpha}_{-1}^{\bar{a}})$ or the gauge degrees of freedom $(\widetilde{\alpha}_{-1}^{I})$.

On the other hand, massless states associated to a twisted constructing element $g = (\vartheta^k, n_{\alpha}e_{\alpha})$ (with k = 1, ..., N - 1) are in general expressed as

$$|q_{\rm sh}\rangle \otimes \widetilde{\alpha} |p_{\rm sh}\rangle \equiv |q + v_q\rangle_R \otimes \widetilde{\alpha} |p + V_q\rangle_L,$$
 (2.58)

where $\tilde{\alpha}$ denotes in this case a product of oscillators of the form $\tilde{\alpha}_{-\eta^a}^a$ (or $\tilde{\alpha}_{-1+\eta^a}^a$) in the complex directions a=1,2,3 (or their conjugates $\bar{a}=\bar{1},\bar{2},\bar{3}$), with $\eta^a=k\,v^a\mod 1$, such that $0<\eta^a\leq 1$. Here we also defined the *local twist* and the *local shift* vectors associated to the constructing element g as

$$v_g = k v$$
 and $V_g = k V + n_\alpha A_\alpha$, (2.59)

respectively. One can trivially extend these results to $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds by including a second twist vector v_2 and a second shift vector V_2 in eq. (2.59).

Orbifold Projections

As we have already mentioned, the matter spectrum of orbifolds will be composed only of those massless states which are invariant under the action of the space group S and its counterpart, the gauge twisting group G, that is, under the orbifold group $O \subset S \otimes G$. The problem here is to find a prescription to figure out whether a massless state is projected out by the orbifold action. Let us make some general observations.

The boundary conditions for closed strings on orbifolds, eqs. (2.55) and (2.56), are summarized by

$$Z(\tau, \sigma + \pi) = gZ(\tau, \sigma) \tag{2.60}$$

for a constructing element g. The set of all states with compact coordinates fulfilling eq. (2.60) define a *Hilbert space* \mathcal{H}_g .

To ensure compatibility of the states from \mathcal{H}_g with the orbifold, we let an arbitrary space group element $h \in S$ act on the coordinates describing the strings. The boundary condition becomes

$$(hZ)(\tau, \sigma + \pi) = hgZ(\tau, \sigma) = hgh^{-1}(hZ)(\tau, \sigma),$$
(2.61)

where we have made use of eq. (2.60). To interpret this, we have to distinguish between two cases: a) h commutes with g, and b) h does not commute with g.

a) Let us consider first that h commutes with q, i.e.

$$[g, h] = 0. (2.62)$$

In this case, eq. (2.61) translates to

$$(hZ)(\tau, \sigma + \pi) = q(hZ)(\tau, \sigma) \tag{2.63}$$

This boundary condition indicates that states described by the coordinates h Z belong to the same Hilbert space \mathcal{H}_q . Consequently, h must act trivially on the states from \mathcal{H}_q :

$$|q_{\rm sh}\rangle_R \otimes \widetilde{\alpha} |p_{\rm sh}\rangle_L \stackrel{h}{\longmapsto} \Phi |q_{\rm sh}\rangle_R \otimes \widetilde{\alpha} |p_{\rm sh}\rangle_L \stackrel{!}{=} |q_{\rm sh}\rangle_R \otimes \widetilde{\alpha} |p_{\rm sh}\rangle_L.$$
 (2.64)

States from \mathcal{H}_g that do not fulfill eq. (2.64) have to be projected out.

Let us define the *centralizer* \mathcal{Z}_g of a constructing element g as the set of all space group elements h commuting with g:

$$\mathcal{Z}_q = \{ h \in S \mid [g, h] = 0 \} . \tag{2.65}$$

The massless matter spectrum of orbifold compactifications is then formed by the massless states which are invariant under all elements of the centralizer.

b) Consider now a noncommuting space group element h,

$$[g, h] \neq 0.$$
 (2.66)

In this case, eq. (2.61) indicates that h maps states from a given Hilbert space \mathcal{H}_g onto a different Hilbert space $\mathcal{H}_{h\,g\,h^{-1}}$. Subsequent application of h then leads to the sequence

$$\mathcal{H}_{g} \xrightarrow{h} \mathcal{H}_{h q h^{-1}} \xrightarrow{h} \mathcal{H}_{h^{2} q h^{-2}} \xrightarrow{h} \mathcal{H}_{h^{3} q h^{-3}} \xrightarrow{h} \dots$$
 (2.67)

The crucial point is now that, since g and hgh^{-1} belong to the same conjugacy class, hZ and Z are identified on the orbifold. This means that, on the orbifold, the different Hilbert spaces $\mathcal{H}_{h^n g h^{-n}}$ are to be combined into a single orbifold Hilbert space. Invariant states are then linear combinations of states from all $\mathcal{H}_{h^n g h^{-n}}$. Such linear combinations involve, in general, relative phase factors (often called γ -phase) [78].

Let us emphasize here that the action of noncommuting space group elements on physical states of \mathcal{H}_g does not project out any state from the spectrum.

2.3.1 Untwisted Sector

In the untwisted sector U, the level matching condition for massless states of orbifolds coincides with that of the uncompactified heterotic string, that is,

$$\frac{m_L^2}{4} = \frac{1}{2}p^2 + \widetilde{N} - 1 = 0 = \frac{1}{2}q^2 - \frac{1}{2} = \frac{m_R^2}{4}, \qquad (2.68)$$

where p is a root of either $E_8 \times E_8$ or SO(32) (see eq. (2.7)), q denotes the SO(8) weight vector of the right-mover and \widetilde{N} counts the number of oscillator excitations. As in the uncompactified heterotic string, eq. (2.68) has solution only if $q^2 = 1$ for the right-movers and either $p^2 = 0$, $\widetilde{N} = 1$ or $p^2 = 2$, $\widetilde{N} = 0$ for the left-movers. The spectrum of the untwisted sector, nevertheless, is different from that of the heterotic string discussed in section 2.1. The reason being that some states are projected out by the action of elements of the centralizer.

Transformation Phase

In the projection condition, eq. (2.64), we have considered that massless states of \mathcal{H}_g acquire a phase Φ under the action of an arbitrary element h of the centralizer \mathcal{Z}_g . In fact, every element of the massless states transform differently under the action of an element h of the centralizer. Embedding h into the gauge degrees of freedom shifts the bosonic coordinates X^I by V_h^I , where V_h is the local shift vector of h (see eq. (2.59)). In the momentum space, this accounts for a phase that depends on the momentum p of the state and the local shift vector V_h :

$$|p\rangle_L \stackrel{h}{\longmapsto} e^{2\pi i p \cdot V_h} |p\rangle_L .$$
 (2.69)

The action of h provides the right-moving states with a similar phase

$$|q\rangle_R \stackrel{h}{\longmapsto} e^{2\pi i q \cdot v_h} |q\rangle_R,$$
 (2.70)

where v_h is the local twist vector associated to h. Finally, oscillators are transformed as

$$\widetilde{\alpha}_{-1}^{I} \stackrel{h}{\longmapsto} \widetilde{\alpha}_{-1}^{I}, \qquad I = 1, \dots, 16,
\widetilde{\alpha}_{-1}^{\mu} \stackrel{h}{\longmapsto} \widetilde{\alpha}_{-1}^{\mu}, \qquad \mu = 2, 3,
\widetilde{\alpha}_{-1}^{a} \stackrel{h}{\longmapsto} e^{2\pi i v_{h}^{a}} \widetilde{\alpha}_{-1}^{a},
\widetilde{\alpha}_{-1}^{\bar{a}} \stackrel{h}{\longmapsto} e^{-2\pi i v_{h}^{a}} \widetilde{\alpha}_{-1}^{\bar{a}}.$$
(2.71)

Left- and right-moving momenta lie on an even, self-dual lattice of lorentzian signature (22.6) [98], implying a relative sign between the phases of the left- and right movers. Therefore, the phase acquired by untwisted massless states under the action of h is given by

$$|q\rangle_{R} \otimes |p\rangle_{L}: \qquad \Phi = e^{2\pi i [p \cdot V_{h} - q \cdot v_{h}]} |q\rangle_{R} \otimes \widetilde{\alpha}_{-1}^{x} |0\rangle_{L}: \qquad \Phi = e^{2\pi i \left[-q \cdot v_{h} + (\delta_{x,a} - \delta_{x,\bar{a}})v_{h}^{a}\right]},$$

$$(2.72)$$

where x stands for the direction in which the oscillator excitations acts.

Untwisted Spectrum

The constructing element of the untwisted sector is g = (1, 0) and, therefore, the associated centralizer contains all elements of the space group. The spectrum of the untwisted sector is composed by those massless states of the uncompactified heterotic string which are invariant under all elements of the space group. In the following, we evaluate the effect of the orbifold projection on the massless spectrum of the heterotic string, discussed in section 2.1.

The ten-dimensional supergravity multiplet of the heterotic string splits into:

• a four-dimensional graviton $g^{\mu\nu}$, dilaton φ , antisymmetric tensor $B^{\mu\nu}$ (whose dual is the model-independent axion a_{MI}), and their superpartners. They are given by the space-group-invariant components of

$$|q\rangle_R \otimes \alpha_{-1}^{\nu}|0\rangle_L. \tag{2.73}$$

Since the left-movers $\alpha_{-1}^{\nu}|0\rangle_L$ do not transform under any h, then the right-movers $|q\rangle_R$ must transform trivially too. From eq. (2.70), we see that this occurs only if the six-dimensional momenta q in eq. (2.73) correspond to¹³

$$q = \begin{cases} \pm \left(\frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}\right) \\ \pm (1, 0, 0, 0) \end{cases}$$
 (2.74)

This is a trivial consequence of requiring to preserve $\mathcal{N} = 1$ SUSY in orbifold compactifications (cf. section 2.2.2);

• some **geometrical moduli** given by states of the type

$$|q\rangle_R \otimes \widetilde{\alpha}_{-1}^{a/\bar{a}} |0\rangle_L$$
 (2.75)

satisfying the invariance condition

$$q \cdot v_h \pm v_h^a = 0 \mod 1, \qquad (2.76)$$

where the relative sign - (+) is associated with an oscillator carrying holomorphic index a (antiholomorphic index \bar{a}). These states are gauge singlet fields, arising from the compact components of the ten-dimensional graviton and antisymmetric tensor (and their superpartners). In particular, the symmetric combinations are the moduli for the flat metric of the compact space, which can be written as

$$G_{\alpha\beta} = e_{\alpha} \cdot e_{\beta} \,, \tag{2.77}$$

where e_{α} correspond to the basis vectors of the compact space. The surviving components of the antisymmetric tensor give rise to the so-called model-dependent axions a_{MD} . Note that, additionally to those states that do not satisfy the invariance condition eq. (2.76), also states such as $|q\rangle_R \otimes \tilde{\alpha}_{-1}^{\nu/I}$ are projected out by the orbifold action.

The action of h on the 16 **ten-dimensional uncharged gauge bosons** leaves invariant only their four-dimensional components specified by

$$|q\rangle_R \otimes \widetilde{\alpha}_{-1}^I |0\rangle_L \tag{2.78}$$

with the right-mover momenta q also given by eq. (2.74) due to the invariance of $\widetilde{\alpha}_{-1}^{I}$. These states are the **16 Cartan generators of** the four-dimensional gauge group \mathcal{G}_{4D} . Therefore, the rank of the gauge group cannot be reduced by compactifying on this kind of orbifolds.¹⁴

The 480 charged gauge bosons are of the form $|q\rangle_R \otimes |p\rangle_L$. Those states left invariant under the action of the space group acquire different natures depending on their transformation properties:

¹³The SO(8) weights q in eq. (2.74) carry implicitly minkowskian index μ .

¹⁴It is possible to reduce the rank of the algebra by embedding the point group generator θ into the gauge degrees of freedom as a rotation Θ instead of a shift vector V.

• the **charged gauge bosons** (and gauginos) of the four-dimensional gauge group \mathcal{G}_{4D} are those states where both left- and right-movers transform trivially under the action of any element h of the space group, that is, where $q \cdot v_h = 0 \mod 1$ and $p \cdot V_h = 0 \mod 1$, independently. The only right-moving momenta satisfying the former constraint are those provided in eq. (2.74), which are also the right-movers of the Cartan generators of the unbroken gauge group. The condition for the left-moving momenta must be fulfilled for any h, then it can be restated neatly as

$$p \cdot V = 0 \mod 1$$
, (for V_1 and V_2 in $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds) (2.79a)

$$p \cdot A_{\alpha} = 0 \mod 1, \qquad \alpha = 1, \dots, 6, \tag{2.79b}$$

where A_{α} are Wilson lines. These states transform in the adjoint representation of \mathcal{G}_{4D} . Provided that not all 480 left-moving momenta p of the ten-dimensional charged bosons satisfy eqs. (2.79), even though the rank is not reduced, the gauge symmetry can be broken:

• those states whose left- and right-moving components transform nontrivially and satisfy the invariance condition

$$p \cdot V_h - q \cdot v_h = 0 \mod 1 \tag{2.80}$$

constitute the so-called **untwisted charged matter** of orbifold models. The various matter states form several SUSY chiral-multiplets. Their gauge transformation properties with respect to \mathcal{G}_{4D} depend on their momenta p.

2.3.2 Twisted Sectors

Zero modes of twisted sectors $T_{k(\ell)}$ are associated to the constructing elements g of the fixed points. Requiring the states to be massless accounts for the following conditions on the left- and right-moving momenta:

$$\frac{1}{4}m_L^2 = \frac{1}{2}p_{\rm sh}^2 + \widetilde{N} - 1 + \delta c \stackrel{!}{=} 0,$$

$$\frac{1}{4}m_R^2 = \frac{1}{2}q_{\rm sh}^2 - \frac{1}{2} + \delta c \stackrel{!}{=} 0,$$
(2.81)

where δc corresponds to a change in the zero point energy related to the appearance of twisted oscillators $\widetilde{\alpha}_{-\eta^a}^{a/\bar{a}}$. It is expressed by

$$\delta c = \sum_{a} \eta^{a} (1 - \eta^{a}) \tag{2.82}$$

with $\eta^a = v_g^a \mod 1$, such that $0 < \eta^a \le 1$. For massless states, one can write the twisted (fractional) oscillator number \widetilde{N} as

$$\widetilde{N} = \sum_{a=1}^{3} \eta^a \left(\widetilde{N}_g^a + \widetilde{N}_g^{*a} \right) . \tag{2.83}$$

Here, \widetilde{N}_g^a and \widetilde{N}_g^{*a} are integer oscillator number, counting respectively the number of excitations in the holomorphic a and antiholomorphic \bar{a} directions.

Transformation Phase

The transformation of left- and right-moving states $|p_{\rm sh}\rangle_L$, $|q_{\rm sh}\rangle_R$ under the action of an arbitrary centralizer element h can also be read off from eqs. (2.69) and (2.70), where we have only to substitute p for $p_{\rm sh}$ and q for $q_{\rm sh}$.

Further, just as in the untwisted sector, only the oscillator excitations on the compact directions a and \bar{a} transform nontrivially. Their transformations are given by

$$\widetilde{\alpha}_{-\eta^{a}}^{a} \stackrel{h}{\longmapsto} e^{2\pi i v_{h}^{a}} \widetilde{\alpha}_{-\eta^{a}}^{a}
\widetilde{\alpha}_{-\eta^{a}}^{\bar{a}} \stackrel{h}{\longmapsto} e^{-2\pi i v_{h}^{a}} \widetilde{\alpha}_{-\eta^{a}}^{\bar{a}}$$
(2.84)

Putting everything together, the complete transformation phase of a massless twisted states reads

$$\Phi = e^{2\pi i \left[p_{\rm sh} \cdot V_h - q_{\rm sh} \cdot v_h + (\tilde{N}_g - \tilde{N}_g^*) \cdot v_h\right]} \Phi_{\rm vac} , \qquad (2.85)$$

where the vacuum phase

$$\Phi_{\text{vac}} = e^{2\pi i \left[-\frac{1}{2} (V_g \cdot V_h - v_g \cdot v_h) \right]} \tag{2.86}$$

arises as consequence of the geometrical properties of twisted strings (cf. appendix of ref. [74]).

2.3.3 A \mathbb{Z}_3 Example

The model studied here was presented in refs. [34,49]. In those works, the $E_8 \times E_8$ heterotic string was compactified on a \mathbb{Z}_3 orbifold with abelian embedding and two Wilson lines. Here, let us first study the spectrum of the model in the absence of Wilson lines and then consider the effect of these background fields.

The embedding of the point group into the gauge degrees of freedom is chosen to be given by the shift vector

$$V = \left(\frac{1}{3}^4, \frac{2}{3}, 0^3\right) \left(\frac{2}{3}, 0^7\right), \tag{2.87}$$

which, together with the twist vector of a \mathbb{Z}_3 orbifold, v = (0, 1/3, 1/3, -2/3), satisfies the modular invariance condition, eq. (2.41).

Untwisted Sector

Aside from the fields $G^{\mu\nu}$, $B^{\mu\nu}$ and φ (and their superpartners), we have nine geometrical moduli of the type

$$|q\rangle_R \otimes \widetilde{\alpha}^{\bar{a}}|0\rangle_L$$
 (2.88)

They arise as follows. In the absence of Wilson lines, the transformation properties of the states in the orbifold spectrum are determined by the action of the point group generated

by θ . Therefore, denoting $\varrho = e^{2\pi i/3}$, we can classify the right-movers according to their eigenvalues with respect to the space group element $h = (\vartheta, 0)$:

$$\varrho^{0}: |q\rangle_{R} \text{ with } q = \pm (\frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}), \pm (1, 0, 0, 0),
\varrho^{1}: |q\rangle_{R} \text{ with } q = (\frac{1}{2}, \frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}), (0, \underline{1, 0, 0}),
\varrho^{2}: |q\rangle_{R} \text{ with } q = (-\frac{1}{2}, \underline{-\frac{1}{2}, \frac{1}{2}}, \frac{1}{2}), (0, \underline{-1, 0, 0}),$$
(2.89)

where, as usual, the underscore denotes all permutations. On the other hand, left-moving oscillators with internal indices also transform:

$$\begin{array}{ll}
\varrho^1: & \widetilde{\alpha}_{-1}^a, \\
\varrho^2: & \widetilde{\alpha}_{-1}^{\overline{a}}.
\end{array}$$
(2.90)

Therefore, for each complex direction a we obtain $\operatorname{six}(\varrho^1)_R(\varrho^2)_L$ invariant states and other six with $(\varrho^2)_R(\varrho^1)_L$. However, note that the former states are conjugate to the latter ones. Together, they enter into a single physical state, so that it is enough to count left-chiral states, i.e. states where the Ramond (half-integer) SO(8) weight has $q^0 = -\frac{1}{2}$. In the present example, we count only the $\operatorname{six}(\varrho^2)_R(\varrho^1)_L$ invariant states. Furthermore, since SUSY is preserved, a physical state of the spectrum must contain a Ramond weight along with an Neveu-Schwarz (integer) weight. In this way, we come up to three invariant states for each of the three complex directions, as given in eq. (2.88). Combinations of these nine moduli correspond to the nine independent deformation parameters of the compact space of a \mathbb{Z}_3 orbifold [99].

The unbroken gauge group has rank sixteen because all Cartan generators given in eq. (2.78) are invariant. The explicit breaking of $E_8 \times E_8$ is obtained through the left-moving momenta p (i.e. the roots of $E_8 \times E_8$ in eq. (2.7a)) satisfying $p \cdot V = 0 \mod 1$. There are 72 invariant momenta p from the first E_8 and 84 from the second E_8 factor. Adding to this number the eight Cartan generators of each E_8 provides the dimensionality of the adjoint of the unbroken gauge group. In this case $72 + 8 \rightarrow 80$ corresponds to the adjoint of SU(9) and $84 + 8 \rightarrow 91 \oplus 1$ correspond to the adjoint of SO(14) × U(1). The corresponding momenta p constitute the roots of the unbroken gauge group. The left-moving states with these momenta $|p\rangle_L$ (transforming as ϱ^0) tensor together with the invariant (ϱ^0) right-movers of eq. (2.89). Since the number of left-chiral states gives the multiplicity of states, we find that these gauge bosons appear, as expected, only once. Therefore, the orbifold action with the shift vector eq. (2.87) breaks the gauge group as

$$\mathcal{G} = \mathcal{E}_8 \times \mathcal{E}_8 \longrightarrow \mathcal{G}_{4D} = \mathrm{SU}(9) \times \mathrm{SO}(14) \times \mathrm{U}(1),$$
 (2.91)

where the U(1) generator is given by the vector

$$t = (0^8) (-18, 0^7).$$
 (2.92)

¹⁵Note that the designation of chirality is arbitrary.

More precisely, in order to determine the unbroken gauge group, one has to take the 72+84 invariant roots p and chose a basis that fixes a semiordering in the weight space [100]. From the semiordering, one can find the simple roots α_i , defined as those roots which are positive and cannot be written as the sum of two positive roots. Provided that the roots have squared length 2, one can then compute the Cartan matrix $A_{ij} = \alpha_i \cdot \alpha_j$ for each gauge group factor G_a (e.g. for SU(9)), which describes uniquely the algebra. A more detailed discussion on how to obtain the unbroken gauge group in orbifold compactifications can be found in section 3.2 of ref. [101].

The untwisted charged matter is formed by tensoring together left-moving $|p\rangle_L$ and right-moving $|q\rangle_L$ states which, separately, acquire nontrivial transformation phases under the action of $h=(\vartheta,0)$. There are 2[84+(14+64)] momenta p from eq. (2.7a) which lead to nontrivial phases for the left-moving states: 84+(14+64) of them transform with a phase ϱ^1 and the remaining 84+(14+64) with ϱ^2 . Moreover, 2×84 momenta come from the first E_8 whereas $2\times(14+64)$ come from the second one. The gauge properties of the states are encoded in the momenta p. The nonabelian gauge quantum numbers with respect to \mathcal{G}_{4D} , eq. (2.91), are obtained by rewriting the momenta in Dynkin labels:

$$p \to p_{DL(G_a)} = (\alpha_1 \cdot p, \dots, \alpha_n \cdot p)$$
 for each G_a , (2.93)

where α_i are the roots of the gauge factor $G_a \subset \mathcal{G}_{4D}$ of rank n. Comparing the momenta in Dynkin labels p_{DL} with the results in the tables of ref. [102], one identifies the representations. In our example, we summarize the properties of the left-moving states $|p\rangle_L$ by

$$|p\rangle_L \text{ with } \varrho^1 \sim (\overline{\bf 84}, \, {\bf 1})_0 \oplus ({\bf 1}, \, {\bf 14})_{+18} \oplus ({\bf 1}, \, \overline{\bf 64})_{-9},$$

 $|p\rangle_L \text{ with } \varrho^2 \sim ({\bf 84}, \, {\bf 1})_0 \oplus ({\bf 1}, \, {\bf 14})_{-18} \oplus ({\bf 1}, \, {\bf 64})_{+9}.$ (2.94)

The U(1) charge, denoted by the subscript, is given by $t \cdot p$, with t representing the U(1) generator given in eq. (2.92). Finally, the physical states are the result of combining the left-movers described in eq. (2.94) with the right movers given in eq. (2.89), which provide the states with a multiplicity factor of three. Again, one finds that $(\varrho^1)_R(\varrho^2)_L$ left-chiral whereas $(\varrho^2)_R(\varrho^1)_L$ are right-chiral invariant states, thus they enter together into physical states. The untwisted charged matter spectrum is

$$3(84, 1)_0 \oplus 3(1, 14)_{-18} \oplus 3(1, 64)_{+9}$$

Twisted Sectors

As discussed in section 2.2.3, the \mathbb{Z}_3 orbifold has 27 fixed points on each of its two twisted sectors, labeled T_1 and T_2 for the action of ϑ^1 and ϑ^2 , respectively. Twisted states are attached to the fixed points. They are described by the momenta satisfying the masslessness conditions eq. (2.81), $q_{\rm sh} = q + v_g$ and $p_{\rm sh} = p + V_g$, where $p \in \Lambda$, and v_g and V_g depend on the constructing element g (see eq. (2.59)).

Let us compute the spectrum of the first twisted sector. Consider the fixed point at the origin. Its constructing element is $g = (\vartheta, 0)$; therefore, the local twist vector is $v_q = v$

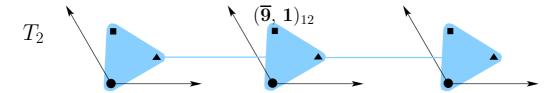


Figure 2.10: In the absence of Wilson lines, all fixed points are degenerate. The twisted matter content correspond then to 27 copies of $(\overline{9}, 1)_{12}$.

and the local shift vector is $V_g = V$. Since the Wilson lines are trivial in this example, notice that the local twist and shift vectors are the same for all constructing elements of this sector. Then it follows that the matter spectrum at each of the fixed points is the same, so that the spectrum of the sector is given by 27 copies of the matter at the origin.

The conditions on the left- and right-moving momenta for massless states read

$$p_{\rm sh}^2 = (p+V)^2 \stackrel{!}{=} 2 - 2\widetilde{N} - 2\delta c = \frac{4}{3} - 2\widetilde{N},$$

$$q_{\rm sh}^2 = (q+v)^2 \stackrel{!}{=} 1 - 2\delta c = \frac{1}{3},$$
(2.95)

where we have used eq. (2.82) to compute $\delta c=1/3$. Further, since $V^2=4/3$ and \widetilde{N} is nonnegative, eq. (2.95) has solution only for $\widetilde{N}=0$. There are only nine $p_{\rm sh}$ and two $q_{\rm sh}$ that solve eq. (2.95):

$$p_{\rm sh}: \quad \left(\frac{-\frac{2}{3}, \frac{1}{3}^{3}}{\frac{1}{3}}, -\frac{1}{3}, 0^{3}\right) \left(\frac{2}{3}, 0^{7}\right), \qquad q_{\rm sh}: \quad \left(0, \frac{1}{3}, \frac{1}{3}, \frac{1}{3}\right), \\ \left(\frac{1}{3}^{4}, \frac{2}{3}, 0^{3}\right) \left(\frac{2}{3}, 0^{7}\right), \qquad \left(\frac{1}{2}, -\frac{1}{6}, -\frac{1}{6}, -\frac{1}{6}\right). \\ \left(-\frac{1}{6}^{4}, \frac{1}{6}, \frac{1}{2}^{2}, -\frac{1}{2}\right) \left(\frac{2}{3}, 0^{7}\right), \\ \left(-\frac{1}{6}^{4}, \frac{1}{6}, -\frac{1}{2}^{3}\right) \left(\frac{2}{3}, 0^{7}\right).$$

Notice that the states $|q_{\rm sh}\rangle_R \otimes |p_{\rm sh}\rangle_L$ of the first twisted sector are right-chiral ($q_{\rm sh}^0 = +1/2$ for the Ramond momentum).

It turns out that in the T_1 (and T_{N-1}) sector of \mathbb{Z}_N orbifolds all solutions to the mass equation enter into an invariant state. In other words, all states from the first twisted sector are invariant under the orbifold action. The reason is as follows. The elements h of the centralizer \mathbb{Z}_g can also be expressed as g^i with $i = 0, \ldots, N-1$. In this case, one can verify that momenta $p_{\rm sh}$ and $q_{\rm sh}$ solving the masslessness conditions always satisfy

$$p_{\rm sh} \cdot V_h - q_{\rm sh} \cdot v_h + (N_g - N_g^*) \cdot v_h - \frac{1}{2} (V_g \cdot V_h - v_g \cdot v_h) = 0 \mod 1$$
 (2.96)

for any centralizer element h.

Therefore, all the momenta $p_{\rm sh}$ and $q_{\rm sh}$ given above survive the orbifold projection, providing, in considering all the fixed points of the T_1 sector, 27 right-chiral supermultiplets which transform as

$$27(9, 1)_{-12}$$
 (2.97)

under gauge transformations.

In the T_2 sector, the situation is similar. As in the T_1 sector, it suffices to consider the constructing element of the fixed point at the origin $g = (\theta^2, 0)$ and count 27 copies of the associated matter content, as illustrated in fig. 2.10. The associated massless momenta are given by

$$p_{\rm sh}: \quad \left(\frac{2}{3}, -\frac{1}{3}^{3}, \frac{1}{3}, 0^{3}\right) \left(-\frac{2}{3}, 0^{7}\right), \qquad q_{\rm sh}: \quad \left(0, -\frac{1}{3}, -\frac{1}{3}, -\frac{1}{3}\right),$$

$$\left(-\frac{1}{3}^{4}, -\frac{2}{3}, 0^{3}\right) \left(-\frac{2}{3}, 0^{7}\right), \qquad \left(-\frac{1}{2}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}\right).$$

$$\left(\frac{1}{6}^{4}, -\frac{1}{6}, -\frac{1}{2}^{2}, \frac{1}{2}\right) \left(-\frac{2}{3}, 0^{7}\right),$$

$$\left(\frac{1}{6}^{4}, -\frac{1}{6}, \frac{1}{2}^{3}\right) \left(-\frac{2}{3}, 0^{7}\right).$$

Since these momenta differ from those of the T_1 by a sign, the states $|q_{\rm sh}\rangle_R \otimes |p_{\rm sh}\rangle_L$ of the T_2 sector correspond to the conjugate of the states in the T_1 sector. One can verify that these states have the following gauge transformations

$$27(\overline{9}, 1)_{+12}$$
. (2.98)

Notice that these states are left-chiral and, therefore, combine with those states from the T_1 sector to form complete SUSY multiplets.

In summary, the matter content of the present \mathbb{Z}_3 orbifold model is given in terms of the gauge representations of the states by

Sector	Matter content
U	$3(84, 1)_0$
	$3(1,14)_{-18}$
	$3(1,64)_9$
T_2	$27(\overline{9}, 1)_{12}$

We omit here the gravity multiplet, the moduli and gauge bosons.

Including Wilson Lines

We introduce now the two Wilson lines given by the vectors [34, 49]

$$A_{1} = \begin{pmatrix} 0^{7}, \frac{2}{3} \end{pmatrix} \begin{pmatrix} 0, \frac{1}{3}^{2}, 0^{5} \end{pmatrix}, A_{3} = \begin{pmatrix} \frac{1}{3}^{3}, \frac{2}{3}, \frac{1}{3}, 0, \frac{1}{3}^{2} \end{pmatrix} \begin{pmatrix} \frac{1}{3}^{2}, 0^{6} \end{pmatrix}.$$
(2.99)

Wilson lines A_{α} are gauge transformations associated to the noncontractible cycles in the directions e_{α} of the compact space. We have seen before that, due to the structure of the compact space, the Wilson lines of \mathbb{Z}_3 orbifolds satisfy $A_{\alpha} = A_{\alpha+1}$ with $\alpha = 1, 3, 5$ (cf. eq. (2.35)).

Geometrical moduli, the graviton, the dilaton and the antisymmetric tensor fields do not feel the presence of the Wilson lines because their transformation properties only depend on the geometry of the compact space (in other words, on the twist vector v and the right-moving momenta q). Gauge bosons with left-moving momenta p, on the other hand, have to satisfy additionally

$$p \cdot A_{\alpha} = 0 \mod 1 \tag{2.100}$$

in order to be invariant (cf. eq. (2.79b)). In this way, the gauge symmetry is further broken.

In the present example, the action of Wilson lines leave invariant only eight momenta p from the first E_8 and 40 from the second E_8 . We find that, together with the 16 invariant Cartan generators, they constitute the adjoint representations of the unbroken gauge group in four dimensions

$$\mathcal{G}_{4D} \xrightarrow{A_{\alpha}} \mathcal{G}'_{4D} = \mathrm{SU}(3) \times \mathrm{SU}(2) \times [\mathrm{U}(1)^5 \times \mathrm{SO}(10) \times \mathrm{U}(1)^3].$$
 (2.101)

Here, we have separated symbolically what we will call the observable sector from the hidden sector, choice justified on the appearance of the $SU(3) \times SU(2)$ gauge factors. Optimistically, one could at this point say that a model with such a gauge group is a good candidate for describing the standard model of particle physics. The eight U(1) generators are labeled t_i with i = 1, ..., 8.

Another effect of the presence of nontrivial Wilson lines is the change of the matter spectrum. Many of the formerly invariant momenta p and $p_{\rm sh}$ are not invariant any more with respect to all elements of the centralizer(s). Let us focus first on the untwisted sector. The centralizer of the constructing element g = (1, 0) is, as mentioned before, the complete space group. A valid basis of \mathcal{Z}_q is given by

$$\mathcal{Z}_g = \{ (\vartheta, 0), (\mathbb{1}, e_1), (\mathbb{1}, e_2), (\mathbb{1}, e_3), (\mathbb{1}, e_4), (\mathbb{1}, e_5), (\mathbb{1}, e_6) \} . \tag{2.102}$$

In particular, the action of the space group elements $h = (1, e_{\alpha})$ imposes new constraints on the momenta p of the untwisted charged matter. From eq. (2.80) (with $V_h = A_{\alpha}$), it follows that these constraints are also given by eq. (2.100). We are then left with $2[(3+3\times 2+2)+16]$ momenta p. After tensoring left- and right-moving states together and computing their gauge quantum numbers, we find the following matter representations of $SU(3) \times SU(2) \times SO(10)$ (we omit the U(1) charges):

$$3(\mathbf{3},\,\mathbf{1},\,\mathbf{1}) \oplus 3(\overline{\mathbf{3}},\,\mathbf{1},\,\mathbf{1}) \oplus 3(\mathbf{1},\,\mathbf{2},\,\mathbf{1}) \oplus 3(\mathbf{1},\,\mathbf{1},\,\overline{\mathbf{16}})$$

The computation of the matter states in the twisted sectors gets more involved than before. In the previous discussion, we have used the fact that, in absence of Wilson lines, the solutions to mass equations and the projection condition are identical at any fixed point of a given twisted sector. Therefore, in that case, all fixed points are degenerate. This situation changes in presence of Wilson lines. Consider, for example, two different

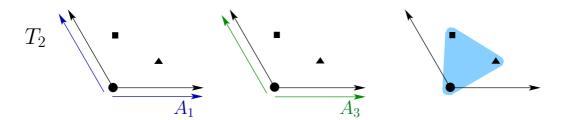


Figure 2.11: The action of the Wilson lines A_1 and A_3 lifts the degeneracy of the fixed points of the first and second complex planes. Note that the Wilson line A_i is related to the directions e_{α} and $e_{\alpha+1}$ of the compact space.

constructing elements of the T_2 sector: $g_1 = (\vartheta^2, 0)$ and $g_2 = (\vartheta^2, e_2)$. The massless momenta $p_{\rm sh} = p + V_{g_1}$ for g_1 correspond to those found in the case without Wilson lines because $V_{g_1} = 2V$. The right-moving momenta is not altered by the introduction of Wilson lines. The states $|q_{\rm sh}\rangle_R \otimes |p_{\rm sh}\rangle_L$ at g_1 form the gauge representations

$$1(1, 2, 1) \oplus 1(\overline{3}, 1, 1) \oplus 4(\overline{1}, 4, 1)$$
 (2.103)

under the gauge group \mathcal{G}'_{4D} given in eq. (2.101) (excluding the U(1) charges).

For g_2 the local shift vector changes to $V_{g_2} = 2V + A_2 = 2V + A_1$. Hence, the solutions to the mass equation for the left-movers take a different form:

$$p_{\rm sh}: \left(-\frac{1}{3}^4 - \frac{2}{3}, 0^3\right) \left(\frac{1}{3}^2, -\frac{2}{3}, 0^5\right), \left(\frac{1}{6}^4, -\frac{1}{6}, -\frac{1}{2}, \frac{1}{2}, \frac{1}{6}\right) \left(-\frac{2}{3}, \frac{1}{3}^2, 0^5\right).$$

These nine momenta along with the right-moving momenta $q_{\rm sh}$ of the orbifold without Wilson lines comprise the matter spectrum of the fixed point with constructing element g_2 , which is written down in terms of the corresponding gauge quantum numbers (excluding U(1) charges):

$$9(1, 1, 1). (2.104)$$

The difference between the matter content for g_1 and g_2 , eqs. (2.103) and (2.104), makes manifest that Wilson lines lift the degeneracy of the fixed points. Only the points with coordinates in the last torus conserve the degeneracy as shown in fig. 2.11. The degeneracy of the last torus gives then a multiplicity factor of three for the twisted states. In principle, we could expect a different matter content at each of the nine fixed points with different coordinates in the first two complex planes. We list in table 2.3 the matter content associated to the different fixed points of the T_2 sector. Notice that there are only five classes of matter content. As we shall discuss shortly, the class of matter content of a constructing element g is related to its local shift V_g .

The matter states in the T_1 sector are, as before, the right-chiral partners of the states in the T_2 sector. Hence, in terms of left-chiral states, the matter spectrum of this models is summarized as follows:

Constructing element(s) g	Matter content
$(\vartheta^2, n_5 e_5 + n_6 e_6)$	$1(1,2,1) \oplus 1(\overline{3},1,1) \oplus 4(1,1,1)$
$(\vartheta^2, e_3 + n_5 e_5 + n_6 e_6)$	
$(\vartheta^2, e_1 + e_2 + e_3 + n_5 e_5 + n_6 e_6)$	$2(1, 2, 1) \oplus 1(\overline{3}, 1, 1) \oplus 1(3, 1, 1) \oplus 8(1, 1, 1)$
$(\vartheta^2, e_1 + e_2 + e_3 + e_4 + n_5 e_5 + n_6 e_6)$	
$(\vartheta^2, e_3 + e_4 + n_5 e_5 + n_6 e_6)$	$1(1,2,1) \oplus 1(3,1,1) \oplus 4(1,1,1)$
$(\vartheta^2, e_2 + e_3 + e_4 + n_5 e_5 + n_6 e_6)$	
$(\vartheta^2, e_2 + n_5 e_5 + n_6 e_6)$	9(1, 1, 1)
$(\vartheta^2, e_1 + e_2 + n_5 e_5 + n_6 e_6)$	
$(\vartheta^2, e_2 + e_3 + n_5 e_5 + n_6 e_6)$	$3(1,2,1) \oplus 3(1,1,1)$

Table 2.3: The matter content attached to the fixed points of the T_2 sector of a \mathbb{Z}_3 orbifold model with two Wilson lines.

Sector	Matter content	Sector	Matter content
U	3(3, 1, 1)	T_2	15(3 , 1 , 1)
	$3(\overline{\bf 3}, {\bf 2}, {\bf 1})$		$12(\overline{\bf 3},{\bf 1},{\bf 1})$
	3(1, 2, 1)		36(1 , 2 , 1)
	$3(1, 1, \overline{16})$		171(1 , 1 , 1)

Let us spend few words on some phenomenological properties of this model. If a string compactification is to reproduce the standard model of particle physics, it must have the gauge group $G_{\rm SM}={\rm SU}(3)\times {\rm SU}(2)\times {\rm U}(1)_Y$. It has been shown that in the present model it is possible to find a combination of U(1)'s producing the correct spectrum of the SM plus additional particles, which turn out to be vectorlike with respect to $G_{\rm SM}$ (see table 1 of ref. [49]). The vectorlike character of the exotic particles insinuates that such particles can acquire large masses provided that adequate couplings exist in the theory. As we will see in section 2.5.1, ensuring that such couplings do not vanish is quite nontrivial, for allowed nonvanishing couplings in string theory have to satisfy strong constraints.

Other qualities of this model include preservation of $\mathcal{N}=1$ SUSY and the spontaneous breaking of the hidden sector. Yet issues such as correct Yukawa-mass textures, proton decay, neutrino masses and compatibility with GUT theories require further investigation.

2.3.4 Local Shift Vectors and Local Spectra

We have seen that there are five classes of local spectra at the different fixed points of a \mathbb{Z}_3 orbifold model with Wilson lines. It turns out that this is a model-independent statement for all \mathbb{Z}_3 orbifold models, as we will explain in the following.

Consider first an arbitrary \mathbb{Z}_N or $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifold model. Locally, at a fixed point

with constructing element g, the 480 gauge bosons of the ten dimensional gauge group \mathcal{G} are affected only by the action of the local shift V_g . Therefore, the local gauge group \mathcal{G}_g and the local matter spectrum coincide with the four-dimensional gauge group and the spectrum at one arbitrary fixed point of an orbifold model with global shift vector $V = V_g$ (and no Wilson lines). The roots of the local gauge group will be those roots p of the ten-dimensional gauge group \mathcal{G} satisfying

$$p \cdot V_g = 0 \mod 1$$
 $p^2 = 2$. (2.105)

Clearly, matter states transform locally under \mathcal{G}_g rather than under the four-dimensional gauge group. This restricts the number of local gauge groups (and local spectra) to be the number of (inequivalent) admissible shift vectors of the studied orbifold.

The intersection of the local gauge groups of the different fixed points in a particular model yields the (global) four-dimensional gauge group \mathcal{G}_{4D} :

$$\mathcal{G}_{4D} = \mathcal{G}_{q_1} \cap \mathcal{G}_{q_2} \cap \mathcal{G}_{q_3} \cap \dots, \tag{2.106}$$

whence it follows that the local symmetry group is, in general, bigger than the \mathcal{G}_{4D} , i.e.

$$\mathcal{G}_{4D} \subset \mathcal{G}_{q_i}, \quad i = 1, 2, 3, \dots$$
 (2.107)

From a global perspective in four dimensions, local matter states form representations of \mathcal{G}_{4D} . These can be derived by the branching rule of the local representations under the breaking $\mathcal{G}_g \to \mathcal{G}_{4D}$. In higher twisted sectors, some of the local states can be projected out from a global perspective, so that only incomplete local representation survive in the full orbifold.

We can now return to our \mathbb{Z}_3 example. Since, in the case of \mathbb{Z}_3 orbifolds, there are only five admissible shift vectors V (see table D.4), any local shift vector V_g must be equivalent to one of them. It follows then that there are also five classes of different local spectra, in agreement with table 2.3. For concreteness, consider the constructing element $g_2 = (\vartheta^2, e_2)$. The corresponding local shift vector is

$$V_{g_2} = 2V + A_1 = \left(\frac{2^4}{3}, \frac{4}{3}, 0^2, \frac{2}{3}\right) \left(\frac{4}{3}, \frac{1^2}{3}, 0^5\right) \xrightarrow{\mathcal{W} + \Lambda} \left(\frac{2}{3}, \frac{1^2}{3}, 0^6\right) \left(\frac{2}{3}, \frac{1^2}{3}, 0^6\right), \quad (2.108)$$

where $W + \Lambda$ denotes Weyl rotations accompanied by $E_8 \times E_8$ lattice translations. One can verify that V_{g_2} corresponds to the shift vector $V^{(2)}$ listed in table 2.3. Therefore, the local gauge group is $\mathcal{G}_{g_2} = E_6 \times SU(3) \times E_6 \times SU(3)$ and the local matter content is given the bifundamental representation $(\mathbf{1}, \mathbf{3}, \mathbf{1}, \mathbf{3})$. It is not difficult to confirm that the nine singlets of eq. (2.104) under \mathcal{G}'_{4D} arise from the breaking $\mathcal{G}_{g_2} \to \mathcal{G}'_{4D}$.

We will see in chapter 4 that the local picture proves to be a useful tool in the search after orbifold models with realistic properties, such as grand unification.

2.3.5 Anomaly Cancellation

Orbifold models present generically one anomalous U(1) symmetry, U(1)_A [55]. However, it is reasonable to expect that the theory in four dimensions should be anomaly free as it arises from the heterotic string, where anomaly cancellation is achieved by the Green-Schwarz mechanism [23]. It has been shown [103] that modular invariance of the orbifold guarantees that the anomaly polynomial in four dimensions factorizes as $\operatorname{tr} F^2 - \operatorname{tr} R^2$. From there it follows that the anomaly can be cancelled by the generalized Green-Schwarz mechanism [104–107]. Besides, the so-called U(1)_A universality condition holds automatically [56, 108]

$$\frac{1}{2}\operatorname{tr}(\ell Q_A) = \frac{1}{2|\mathsf{t}_j|^2}\operatorname{tr} Q_j^2 Q_A = \frac{1}{6|\mathsf{t}_A|^2}\operatorname{tr} Q_A^3 = \frac{1}{24}\operatorname{tr} Q_A \left(\equiv 8\pi^2 \delta_{GS} \right) \quad j \neq A, \quad (2.109)$$

where ℓ denotes the index of a given representation under a nonabelian gauge group. Furthermore, t_j are the generators of the U(1) factors (t_A corresponds to U(1)_A) that define the charge Q_j as:

$$Q_j |p_{\rm sh}\rangle_L = (\mathsf{t}_j \cdot p_{\rm sh}) |p_{\rm sh}\rangle_L$$

The constant δ_{GS} enters in the transformation of the dilaton in the Green-Schwarz mechanism. Orbifold models contain at most one anomalous U(1). Hence the remaining U(1)'s satisfy

$$\frac{1}{2}\operatorname{tr}(\ell Q_i) = \frac{1}{2|\mathbf{t}_i|^2}\operatorname{tr} Q_j^2 Q_i = \frac{1}{6|\mathbf{t}_i|^2}\operatorname{tr} Q_i^3 = \frac{1}{24}\operatorname{tr} Q_i = 0 \quad i, j \neq A.$$
 (2.110)

The universality condition, eq. (2.109), states also that pure abelian, mixed abelian–nonabelian, and mixed abelian–gravitational anomalies are not independent of each other. Therefore, all of them cancel simultaneously. In ref. [109] it is shown that this holds also for discrete symmetries present in orbifold constructions. Other anomalies, such as pure nonabelian anomalies or Witten's anomaly [110], vanish automatically in orbifolds. The conditions (2.109) and (2.110) have been confirmed for all models appearing in the present work.

2.4 Discrete Torsion in Orbifold Models

It is well known that discrete torsion [91] introduces an additional degree of freedom in $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifold constructions [83]. Nonetheless, there are some features of discrete torsion on orbifolds which have received little attention. Only recently we have noted [74] that \mathbb{Z}_N orbifolds also admit discrete torsion and, moreover, that discrete torsion can add more than one degree of freedom to orbifolds with Wilson lines. In this section, we study briefly $\mathbb{Z}_N \times \mathbb{Z}_M$ and \mathbb{Z}_N orbifold models with discrete torsion and introduce the concept of generalized discrete torsion. Then we analyze an interesting equivalence between models with generalized discrete torsion and torsionless brother models. For a more detailed discussion, ref. [111] is recommended.

The one-loop partition function Z for orbifold compactifications has the overall structure

$$Z = \sum_{\substack{g,h\\[g,h]=0}} \varepsilon(g,h) Z(g,h) , \qquad g,h \in S.$$
 (2.111)

The relative phases $\varepsilon(g,h)$ are called discrete torsion phases. Their values can vary between the different terms in the partition function and thus between the different sectors. Different assignments of phases lead, in general, to different orbifold models. The arbitrariness of $\varepsilon(g,h)$ corresponds to the freedom of turning on a background antisymmetric field on the torus [91].

Although the discrete torsion phases appear to be arbitrary, modular invariance at one loop and factorizability of the partition function at two loops impose certain constraints on the torsion phases. They are given by

$$\varepsilon(g_1g_2, g_3) = \varepsilon(g_1, g_3) \varepsilon(g_2, g_3) , \qquad (2.112a)$$

$$\varepsilon(g_1, g_2) = \varepsilon(g_2, g_1)^{-1}. \tag{2.112b}$$

$$\varepsilon(g,g) = 1. \tag{2.112c}$$

The last equation is a convention, rather than a constraint, and can be seen as a sort of normalization of the phases.

2.4.1 Discrete Torsion without Wilson Lines

In orbifolds without Wilson lines, g and h are chosen to be elements of the point group P. It follows then that for \mathbb{Z}_N orbifolds the solution of eqs. (2.112) is trivial:

$$\varepsilon(g,h) = 1 \quad \forall g, h \in P .$$
 (2.113)

Therefore, in the case of \mathbb{Z}_N orbifolds without Wilson lines, non-trivial discrete torsion cannot be introduced.

In $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds, still without Wilson lines, the situation is different because there are independent pairs of elements which commute with each other. If we take two point group elements $g = \vartheta^{k_1} \omega^{\ell_1}$ and $h = \vartheta^{k_2} \omega^{\ell_2}$, then the discrete torsion phase is determined by eqs. (2.112) to be [83]

$$\varepsilon(g,h) = \varepsilon(\theta^{k_1}\omega^{\ell_1}, \theta^{k_2}\omega^{\ell_2}) = \exp\left\{\frac{2\pi i a}{N}(k_1\ell_2 - k_2\ell_1)\right\}, \quad a = 0, 1, \dots, N,$$
 (2.114)

where N is the order of the twist ϑ .¹⁶ Note that there are only N inequivalent assignments of ε .

¹⁶Our convention for $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds is that M = n N with $n \in \mathbb{Z}$.

2.4.2 Generalized Discrete Torsion

More recently, the concept of discrete torsion has been extended by introducing a generalized discrete torsion phase in the context of type IIA/B string theory [112]. This generalized torsion phase depends on the fixed points rather than on the sectors of the orbifold. Therefore, one has to consider g and h to be elements of the space group S.

Considering the space group elements $g = (\vartheta^{k_1} \omega^{\ell_1}, n_{\alpha} e_{\alpha})$ and $h = (\vartheta^{k_2} \omega^{\ell_2}, m_{\alpha} e_{\alpha})$, the general solution to eqs. (2.112) for the discrete torsion phase is written down as

$$\varepsilon(g,h) = e^{2\pi i \left[a (k_1 \ell_2 - k_2 \ell_1) + b_\alpha (k_1 m_\alpha - k_2 n_\alpha) + c_\alpha (\ell_1 m_\alpha - \ell_2 n_\alpha) + d_{\alpha\beta} n_\alpha m_\beta\right]}. \tag{2.115}$$

where the sum over α , β is understood. The values of $a, b_{\alpha}, c_{\alpha}, d_{\alpha\beta}$ are required by modular invariance and the geometry of the lattice Γ to satisfy

$$N a, N_{\alpha} b_{\alpha}, N_{\alpha} c_{\alpha}, N_{\alpha\beta} d_{\alpha\beta} = 0 \mod 1,$$

 $d_{\alpha\beta} = -d_{\beta\alpha},$ (2.116)

for each $\alpha, \beta = 0, \ldots, 6$. Here N is the order of the twist ϑ , N_{α} the order of the Wilson line A_{α} , and $N_{\alpha\beta}$ is the greatest common divisor of N_{α} and N_{β} (compare with the conditions for modular invariance, eqs. (2.42)). Additional constraints on the parameters b_{α} , c_{α} , $d_{\alpha\beta}$ due to the choice of the lattice Γ appear in a similar fashion as those on the order of Wilson lines, explained in section 2.2.2. It is not hard to see that if $e_{\alpha} \simeq e_{\beta}$ on the orbifold, then $b_{\alpha} = b_{\beta}$, $c_{\alpha} = c_{\beta}$ and $d_{\alpha\beta} = 0$ must hold.

The generalized discrete torsion is not restricted only to $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds, as it was commonly believed, but will likewise appear in the \mathbb{Z}_N case. Clearly, since in \mathbb{Z}_N orbifolds there is only one shift, the parameters a and c_{α} vanish.

Role of Discrete Torsion on Orbifolds

The most important consequence of nontrivial ε -phases for our discussion is that they modify the boundary conditions for twisted states and thus change the twisted spectrum. This can be seen from the transformation phase Φ of eq. (2.64), which is modified in the presence of discrete torsion according to

$$\Phi \longmapsto \Phi' = \varepsilon(q, h) \Phi . \tag{2.117}$$

Clearly, the phases Φ' depend on the constructing element g and an element h of its centralizer \mathcal{Z}_g , i.e. $\Phi' = \Phi'(g,h)$. The projection phases Φ' and Φ project out different twisted matter states. Very frequently all the states located at some fixed point are projected out by the effect of the modified phase Φ' . One could say that these 'empty' fixed points disappear from the spectrum or, in other words, that they are nonphysical (at massless level). This feature is interesting because it allows to interpret the effect of discrete torsion in terms of a change in the metric of the compact space, as is discussed in section 2.4.4.

Examples

 \mathbb{Z}_3 Orbifolds. Let us consider the \mathbb{Z}_3 orbifold compactified on an SU(3)³ lattice. As we have seen section 2.2.2, the lattice vectors of Γ are related by the action of the point group generator. In particular, we have that $e_{\alpha} \simeq e_{\alpha+1}$, for $\alpha = 1, 3, 5$, on the orbifold. This implies that there are only three independent b_{α} , namely b_1 , b_3 , b_5 , while the other b-parameters satisfy $b_2 = b_1$, $b_4 = b_3$, $b_6 = b_5$. Further, the antisymmetric matrix $d_{\alpha\beta}$ takes the form

$$d_{\alpha\beta} = \begin{pmatrix} 0 & 0 & d_1 & d_1 & d_2 & d_2 \\ 0 & 0 & d_1 & d_1 & d_2 & d_2 \\ -d_1 & -d_1 & 0 & 0 & d_3 & d_3 \\ -d_1 & -d_1 & 0 & 0 & d_3 & d_3 \\ -d_2 & -d_2 & -d_3 & -d_3 & 0 & 0 \\ -d_2 & -d_2 & -d_3 & -d_3 & 0 & 0 \end{pmatrix}.$$

$$(2.118)$$

Therefore, there are six independent discrete torsion parameter, which can take the values $0, \frac{1}{3}$ or $\frac{2}{3}$.

 $\mathbb{Z}_3 \times \mathbb{Z}_3$ Orbifolds. The $\mathbb{Z}_3 \times \mathbb{Z}_3$ orbifold is very similar to \mathbb{Z}_3 . Since the compactification lattice is that of the \mathbb{Z}_3 orbifold, the six discrete torsion parameters b_{α} and $d_{\alpha\beta}$ (eq. (2.118)) are also admissible in this case. Additionally, the parameters c_{α} appear in the theory. In the same way as b_{α} , they are restricted by the geometry, so that only c_{α} with $\alpha = 1, 3, 5$ are independent. Including a, we obtain ten independent discrete torsion parameters with values $0, \frac{1}{3}$ or $\frac{2}{3}$.

 \mathbb{Z}_6 -II Orbifolds. For the \mathbb{Z}_6 -II orbifold on a $G_2 \times SU(3) \times SO(4)$ lattice an analogous consideration shows that there are only few nontrivial discrete torsion parameters. We find that the only discrete torsion parameters that accept nonzero values are $b_3 = b_4 = 0, \frac{1}{3}, \frac{2}{3}, b_5, b_6 = 0, \frac{1}{2}$ and $d_{56} = -d_{65} = 0, \frac{1}{2}$; that is, only four independent parameters. However, very frequently the factors accompanying these parameters in eq. (2.115) (such as $(k_1 m_{\alpha} - k_2 n_{\alpha})$ for the parameters b_{α}) vanish, implying that the corresponding parameters are nonphysical. This is explained by taking into account the space group elements g, h entering the ε -phases, that is, all constructing and centralizer elements. It turns out that this happens for all discrete torsion parameters of \mathbb{Z}_6 -II orbifolds. Hence, discrete torsion is irrelevant in \mathbb{Z}_6 -II orbifolds. This will be important in chapter 4 for the classification of orbifolds of this type.

2.4.3 Brother Models

In most of the studies of orbifold models, it is claimed that two models whose parameters (V_1, V_2, A_{α}) differ only by lattice translations are equivalent. This is, in general, not true. The reason being that lattice translations influence the projection condition of twisted states, eq. (2.64).

A (torsionless) model M is defined by (V_1, V_2, A_α) . A brother model M' appears by

adding lattice vectors to the shifts and Wilson lines, i.e. M' is defined by

$$(V_1', V_2', A_\alpha') = (V_1 + \Delta V_1, V_2 + \Delta V_2, A_\alpha + \Delta A_\alpha), \qquad (2.119)$$

with $\Delta V_i, \Delta A_\alpha \in \Lambda$. From the conditions (2.42), the choice of lattice vectors $(\Delta V_i, \Delta A_\alpha)$ is constrained by

$$M(V_1 \cdot \Delta V_2 + V_2 \cdot \Delta V_1 + \Delta V_1 \cdot \Delta V_2) = 0 \mod 2 \equiv 2x,$$
 (2.120a)

$$N_{\alpha} \left(V_i \cdot \Delta A_{\alpha} + A_{\alpha} \cdot \Delta V_i + \Delta V_i \cdot \Delta A_{\alpha} \right) = 0 \mod 2 \equiv 2 y_{i\alpha} , \qquad (2.120b)$$

$$Q_{\alpha\beta} \left(A_{\alpha} \cdot \Delta A_{\beta} + A_{\beta} \cdot \Delta A_{\alpha} + \Delta A_{\alpha} \cdot \Delta A_{\beta} \right) = 0 \mod 2 \equiv 2 z_{\alpha\beta} , \qquad (2.120c)$$

where $x, y_{i\alpha}, z_{\alpha\beta} \in \mathbb{Z}$.

One can verify that the inclusion of the lattice vectors $(\Delta V_1, \Delta V_2, \Delta A_{\alpha})$ alters the projection phase of brother models as

$$\Phi \longmapsto \Phi' = \widetilde{\varepsilon}(g, h) \Phi , \qquad (2.121)$$

where the 'brother phase' $\widetilde{\varepsilon}$ is given by

$$\widetilde{\varepsilon}(g,h) = \exp\left\{-2\pi i \left[\left(k_1 \ell_2 - k_2 \ell_1\right) \left(V_2 \cdot \Delta V_1 - \frac{x}{M}\right) + \left(k_1 m_\alpha - k_2 n_\alpha\right) \left(A_\alpha \cdot \Delta V_1 - \frac{y_{1\alpha}}{N_\alpha}\right) + \left(\ell_1 m_\alpha - \ell_2 n_\alpha\right) \left(A_\alpha \cdot \Delta V_2 - \frac{y_{2\alpha}}{N_\alpha}\right) + n_\alpha m_\beta \left(A_\beta \cdot \Delta A_\alpha - \frac{z_{\alpha\beta}}{Q_{\alpha\beta}}\right) \right] \right\}, \quad (2.122)$$

corresponding to the constructing element $g = (\theta^{k_1} \omega^{\ell_1}, n_{\alpha} e_{\alpha})$ and the centralizer element $h = (\theta^{k_2} \omega^{\ell_2}, m_{\alpha} e_{\alpha})$. One can see that $D_{\alpha\beta} \equiv A_{\beta} \cdot \Delta A_{\alpha} - z_{\alpha\beta}/Q_{\alpha\beta}$ is (almost) antisymmetric in α , β ,

$$D_{\alpha\beta} = -D_{\beta\alpha} \bmod 1. \tag{2.123}$$

Like the discrete torsion phase, the brother phase is not the same for all fixed points; hence the local spectrum is changed.

The brother phase $\tilde{\varepsilon}$ and the generalized discrete torsion phase eq. (2.115) are not only very similar, but can, in fact, be made coincide. This implies an unexpected connection between lattice translations of the parameters $(\Delta V_1, \Delta V_2, \Delta A_{\alpha})$ and discrete torsion. In other words, we find that models with discrete torsion can be mimicked by torsionless models with modified shift vector(s) and background fields. We find however that the noninteger values of the parameter $d_{\alpha\beta}$ do not allow an interpretation in terms of lattice translations in models with trivial Wilson lines. Therefore, discrete torsion is, in this sense, more general than the concept of brother models.

As an illustration of the relation between brother models and orbifolds with discrete torsion, we have investigated the distinct $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifold models with standard embedding that one can find for different nonzero values of the discrete torsion parameter a. We have found that one can trade the parameter a for a pair of lattice vectors $(\Delta V_1, \Delta V_2)$ which, added the standard embedding shift vectors, lead to the spectra obtained in the models with discrete torsion. Our findings are listed in table D.6.

2.4.4 Discrete Torsion and Nonfactorizable Lattices

Let us comment on one last interesting observation. We have found that in many cases orbifold models M with certain geometry, i.e. compactification lattice Γ , and generalized discrete torsion switched on are equivalent to torsionless models M' based on a different lattice Γ' . Model M' has less fixed points than M, and the mismatch turns out to constitute precisely the 'empty' fixed points of model M due to the discrete torsion phase.

The simplest examples are based on $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifolds with standard embedding and without Wilson lines. By varying the allowed discrete-torsion parameters of this orbifold (especially, $d_{\alpha\beta}$), we have found eight different models with nonzero net number of **27**-plets of E_6 (see ref. [74]). Interestingly, these models have already been discussed in the literature, but in a somewhat different context. They appeared first in ref. [113] in the context of free fermionic string models related to the $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifold with an additional freely acting shift. More recently, new $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifold constructions have been found in studying orbifolds of non-factorizable six-tori [87, 94]. For each of the models found by adding nonvanishing discrete torsion phases, there is a corresponding 'non-factorizable' model M' with the following properties:

- 1. Each 'non-empty' fixed point, i.e. each fixed point with local zero-modes, in the model M can be mapped to a fixed point with the same spectrum in model M'.
- 2. The number of 'non-empty' fixed points in M coincides with the total number of fixed points in M'.

These relations are not limited to $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifolds, rather we find an analogous connection also in other $\mathbb{Z}_N \times \mathbb{Z}_M$ cases $(\mathbb{Z}_N \times \mathbb{Z}_M \text{ orbifolds based on non-factorizable compactification lattices have recently been discussed in [88]). This result hints towards an intriguing impact of generalized discrete torsion on the interpretation of orbifold geometry. What the (zero-mode) spectra concerns, introducing generalized discrete torsion (or considering generalized brother models) is equivalent to changing the geometry of the underlying compact space, <math>\Gamma \to \Gamma'$. To establish complete equivalence between these models would require to prove that the couplings of the corresponding states are the same, which is beyond the scope of the present study. It is, however, tempting to speculate that non-resolvable singularities (fixed points with no states attached) do not 'really' exist as one can always choose (for a given spectrum) the compactification lattice Γ in such a way that there are no 'empty' fixed points.

2.5 String Interactions: Yukawa Couplings

To close this chapter, let us examine one crucial element necessary in order to study the low-energy field theory limit of orbifold compactifications: field interactions. In contrast to pure field theory, where couplings between matter fields are chosen *ad hoc*, in orbifold compactifications they are determined by strict rules derived from string theory.

2.5.1 String Selection Rules

Consider the *n*-point correlation function of two fermions and n-2 bosons. The corresponding physical states shall be denoted by Ψ_i , $i=1,\ldots,n$. Then, in the field theory limit, a non-vanishing correlation function induces the following term in the superpotential

$$W \supset \Psi_1 \Psi_2 \Psi_3 \dots \Psi_n . \tag{2.124}$$

A complete evaluation of the correlation function has only been performed for 3-point couplings and yields a moduli dependent coupling strength [114, 115, 99, 116]. Recently, the correlation function of n-point couplings has been discussed at some extent [117].

On the other hand, symmetries of the correlation function give rise to the so-called string selection rules. These rules determine whether a given coupling vanishes or not. We use the following notation: the constructing elements of Ψ_i are denoted by $g_i = (\theta_i, n_{\alpha}^i e_{\alpha}) \in S$ and their left- and right-moving shifted momenta, by $p_{\text{sh},i}$ and $q_{\text{sh},i}$, respectively. Then, the string selection rules read:

• Gauge invariance.

Since the 16-dimensional left-moving momenta describe the gauge quantum numbers, the sum over all left-moving shifted momenta $p_{\text{sh},i}$ must vanish:

$$\sum_{i} p_{\text{sh},i} = 0 \tag{2.125}$$

This translates to the field theoretic requirement of gauge invariance for allowed terms in the superpotential. However, note that summing all momenta is, in practice, very cumbersome. Instead, one can verify gauge invariance directly by computing all abelian and nonabelian representations corresponding to the particles in the spectrum and then use well-known rules to form gauge invariant combinations of (super)fields. We would rather apply this second approach.

• Conservation of R-charge.

In orbifold constructions, R-symmetries are discrete symmetries in the six-dimensional space inherited from Lorentz invariance of the ten-dimensional theory. Basically, they arise from demanding invariance of the compact space under the twist. Generically, there are three such symmetries—one for each complex plane—whose quantum numbers can be identified with the last three components of the SO(8) weight momenta $q_{\rm sh}$ of the right movers. Since $q_{\rm sh}$ is not invariant under the ghost picture changing [47], the R-charges have to be amended by some oscillator contributions, resulting in

$$R_i = q_{\text{sh},i} - \widetilde{N}_{q,i} + \widetilde{N}_{q,i}^* \tag{2.126}$$

which lie in the SO(8) weight lattice. Here $\widetilde{N}_{g,i}$ and $\widetilde{N}_{g,i}^*$ are (vectors of) integer oscillator numbers, counting the number of holomorphic and antiholomorphic oscillator excitations, respectively. From the definition of the charges (2.126), one notices that the corresponding R-symmetries do distinguish between bosons and fermions.¹⁷

¹⁷Recall that fermions and bosons have different right-moving momenta.

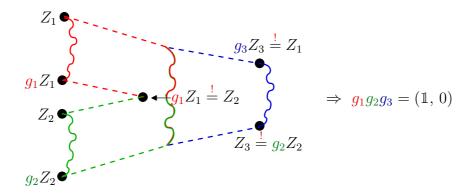


Figure 2.12: The space group selection rule can be visualized as the ability of twisted strings to join. g_i denote three constructing elements.

In the untwisted sector, the bosonic R-charges, eq. (2.126), have only three different values. This allows to split the untwisted sector U in further untwisted sectors U_1 , U_2 and U_3 comprised by (super)fields with R-charges (0,1,0,0), (0,0,1,0) and (0,0,0,1), respectively.

Invariance of the theory under R-symmetries constrains the superpotential of the theory. Since these symmetries are discrete (as they arise from discrete rotations), the invariance conditions can be stated as

$$\sum_{i} R_i^a = 0 \mod N^a \quad \text{for } a = 1, 2, 3 , \qquad (2.127)$$

where N^a denotes the order of the twist action on the a^{th} complex plane, i.e. it is the smallest integers such that $N^a v^a \in \mathbb{Z}$ (no summation). Here, two of the R_i come from fermions and the rest from bosons in order to be allowed in the superpotential. For computational purposes, it is more convenient to use the purely bosonic notation, where eq. (2.127) becomes

$$\sum_{i} R_i^a = -1 \mod N^a \ . \tag{2.128}$$

A caveat is in order here: in all moment we have assumed that the lattice Γ of the compact space is factorizable, that is, that Γ can be written as the product of three one-complex-dimensional sublattices, each of which corresponds to a complex plane Z^a of the internal dimensions. If Γ is nonfactorizable, eq. (2.127) has to be modified. A brief discussion on this issue is provided in appendix Γ .

• Space group selection rule.

Since the states entering the superpotential carry also some information about their space group properties, the product of constructing elements g_i must lie in the same conjugacy class as the identity, i.e.

$$\prod_{i} g_i \simeq (\mathbb{1}, 0) \ . \tag{2.129}$$

In terms of conjugate elements $h_i g_i h_i^{-1}$ of g_i , this condition can be reformulated as [118]

$$\prod_{i} h_i g_i h_i^{-1} = (\mathbb{1}, v) \quad \text{with} \quad v \in \sum_{i} (\mathbb{1} - \theta_i) \Lambda.$$
 (2.130)

Notice that this rule implies in particular that

$$\prod_{i} \theta_i = 1, \qquad (2.131)$$

condition which is known in the literature as the point group selection rule. Stated in this way, the point group selection rule can be reinterpreted as a discrete symmetry [109]. In fact, also the translational part on the space group rule $(v \in \sum_i (1 - \theta_i)\Lambda)$ can be seen as a discrete (flavor) symmetry.

This selection rule can be visualized as the geometrical ability of twisted strings to join (see fig. 2.12).

On the γ -Rule(s)

In the literature, there exists an additional selection rule, usually called γ -rule. In our notation, it reads [47,99]

$$\sum_{i} \gamma_i = 0 \mod 1 \,\,, \tag{2.132}$$

where γ_i denotes the so-called γ -phase of Ψ_i . If we suppose that the states Ψ_i are associated to the conjugacy classes of $g_i \in S$, a phase $e^{2\pi i \gamma_i(h)}$ arises from the action of a noncommuting element h ($[g_i, h] \neq 0$) on the geometrical part of Ψ_i . However, since the physical states Ψ_i are, by definition, space group invariant, this phase comes always along with a phase $\Phi(p_{\rm sh}, q_{\rm sh}, h)$, such that

$$\Psi_i \xrightarrow{h} e^{2\pi i \gamma_i(h)} \Phi_i(p_{\rm sh}, q_{\rm sh}, h) \Psi_i = \Psi_i. \tag{2.133}$$

It follows then that the gamma phases $\gamma_i(h)$ can also be written in terms of $p_{\rm sh}, q_{\rm sh}$ and the embedding of h in the gauge degrees of freedom. Considering a coupling $\Psi_1\Psi_2\Psi_3\ldots$, it turns out that this relation implies that

$$e^{-2\pi i \sum_{i} \gamma_i(h)} = \prod_{i} \Phi_i(p_{\rm sh}, q_{\rm sh}, h)$$
 (2.134)

Further, we have shown [78] that $\prod_i \Phi_i(p_{\rm sh}, q_{\rm sh}, h) = 1$ follows from the selection rules listed above. Therefore, the expression

$$\sum_{i} \gamma_i(h) = 0 \mod 1 \tag{2.135}$$

with an arbitrary space group element $h \in S$, is always true for those couplings allowed by the previous rules. Notice that if we consider that the elements $h_1 = (\vartheta, 0)$, $h_2 = (\omega, 0)$ and $h_{\alpha+2} = (\mathbb{1}, e_{\alpha})$ form a basis of S, we can have at most eight γ -"rules". The one traditionally considered corresponds to that associated to the space group element $h = (\vartheta, 0)$ of \mathbb{Z}_N orbifolds. For a very detailed discussion on the derivation and triviality of the γ -rule(s), we refer to [111].

Chapter 3

Classification of Orbifolds

This chapter is devoted to the techniques used to arrive systematically to inequivalent orbifold compactifications of the heterotic string. We review first the frequently called $Dynkin\ diagram\ method$, based on a theorem by Kač. This method is used mainly to classify models without background fields (Wilson lines). Models with Wilson lines can be classified more effectively by using a proper ansatz that characterizes Wilson lines of a given order, as discussed in section 3.2. Finally, we introduce the C++ Orbifolder, a computer program developed to classify orbifold models and compute their properties.

3.1 \mathbb{Z}_N Orbifolds without Wilson Lines

In the absence of Wilson lines, the *Dynkin diagram method* [93, 119, 120] is the standard strategy to obtain the gauge embeddings of the point group P. It consists in identifying the inner automorphisms of the Lie algebra of the gauge group \mathcal{G} (either $E_8 \times E_8$ or SO(32)), as we will now describe.

Consider the extended Dynkin diagram of E_8 and SO(32) given in fig. 3.1. The numbers attached to the nodes are the Coxeter or Kač labels k_i , which are by definition the expansion coefficients of the highest root α_H in terms of the simple roots, that is

$$\alpha_H = k_1 \alpha_1 + \ldots + k_r \alpha_r \,, \tag{3.1}$$

where r is the rank¹ of the algebra. For convenience, the Kač label of the most negative root $\alpha_0 \equiv -\alpha_H$ is set to $k_0 = 1$. Then, by a theorem due to Kač [121], all order-N inner automorphisms of an algebra up to conjugation are given by

$$\sigma_s(E_{\alpha_i}) = \exp(2\pi i s_i/N) E_{\alpha_i}, \quad i = 0, \dots, r,$$
(3.2)

where E_{α} are the step operators of the Lie algebra of \mathcal{G} and the sequence $s = (s_0, \ldots, s_r)$ may be chosen arbitrarily subject to the conditions that the coefficients s_i be nonnegative,

¹The rank of E_8 is r = 8 whereas the rank of SO(32) is r = 16.

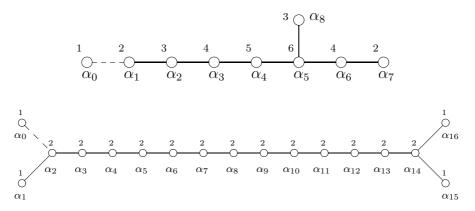


Figure 3.1: Extended Dynkin diagram of E_8 and SO(32) and the associated Coxeter or Kač labels.

relatively prime integers and

$$\sum_{i=0}^{r} k_i s_i = N. \tag{3.3}$$

The Shift Vector

The embedding of the point group in the gauge degrees of freedom, described by $X^I \mapsto X^I + V^I$ (see eq. (2.19)), induces the transformations

$$\sigma_V(H_i) = H_i, \qquad \sigma_V(E_\alpha) = \exp(2\pi i \alpha \cdot V) E_\alpha$$
 (3.4)

on the Cartan generators H_i and step operators E_{α} of the Lie algebra of \mathcal{G} , with α being a root of \mathcal{G} , and V the shift vector. These transformations clearly describe an automorphism of the algebra.

To derive the shift vector corresponding to a given automorphism is now particularly easy. Comparing eq. (3.2) to eq. (3.4), it follows that

$$\alpha_i \cdot V = \frac{s_i}{N}, \quad i = 1, \dots, r, \tag{3.5}$$

for the r linearly independent simple roots α_i . Using that the simple roots α_i and the fundamental weights (their duals) satisfy $\alpha_i \cdot \alpha_j^* = \delta_{ij}$, one can expand the shift vector V in terms of the fundamental weights as

$$V = \sum_{i=1}^{r} \frac{s_i}{N} \alpha_i^* \,, \tag{3.6}$$

i.e. the integers s_i divided by the order N are the Dynkin labels of V. A direct calculation reveals that this shift vector also gives the correct transformation for the step operator corresponding to the most negative root α_0 , $\sigma_V(E_{\alpha_0}) = \exp(2\pi i \alpha_0 \cdot V) E_{\alpha_0}$.

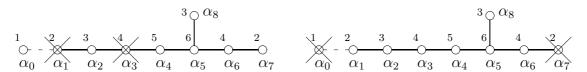


Figure 3.2: Symmetry breaking induced by an $E_8 \times E_8$ inner automorphism of order N = 6 described by $s^a = (0, 1, 0, 1, 0, 0, 0, 0, 0, 0)$ and $s^b = (4, 0, 0, 0, 0, 0, 0, 1, 0)$.

Constructing shift vectors of $E_8 \times E_8$ requires to find two sequences $s^a = (s_0^a, \ldots, s_8^a)$ and $s^b = (s_0^b, \ldots, s_8^b)$, each of which leads to independent automorphism, V^a and V^b , acting differently on each of the two E_8 gauge factors. One can show that the combination

$$V = (V^a)(V^b) (3.7)$$

form an inner automorphism of $E_8 \times E_8$.

To compute the explicit form of the shift vectors, we use tables D.7 and D.8, where we provide our choice of the basis for the simple roots and fundamental weights of SO(32) and E_8 , respectively.

The Unbroken Gauge Group \mathcal{G}_{4D}

To determine the unbroken gauge group \mathcal{G}_{4D} , one has to verify the action of an inner automorphism on the Cartan generators H_i and the step operators E_{α} for the simple roots α_i of the algebra of \mathcal{G} . Their transformations due to the shift vector V are given by eq. (3.4). We see first that the Cartan generators do not transform, confirming that the rank is not reduced by the action of the shift vector, as we learnt in section 2.3.1. Besides, from the transformation properties of the step operators, eq. (3.4), we note that the only step operators that are invariant under the automorphism are those that satisfy

$$\alpha_i \cdot V = 0 \mod 1. \tag{3.8}$$

Using that $\alpha_i \cdot \alpha_j^* = \delta_{ij}$, we conclude then that in the extended Dynkin diagram(s) depicted in fig. 3.1, the simple root α_i (i = 0, ..., r) is projected out, if and only if the coefficient s_i in eq. (3.3) is nonzero.

This allows an interpretation in terms of the Dynkin diagram of the corresponding Lie algebra. Given the parameters (s_1, \ldots, s_r) , the Dynkin diagram of the unbroken gauge group \mathcal{G}_{4D} is obtained after deleting the nodes for which $s_i \neq 0$ from the Dynkin diagram of the original gauge group \mathcal{G} .

To illustrate our discussion, let us consider the $E_8 \times E_8$ Lie algebra. The parameters

$$s^{a} = (0, 1, 0, 1, 0, 0, 0, 0, 0)$$
 and $s^{b} = (4, 0, 0, 0, 0, 0, 0, 1, 0)$ (3.9)

acting on the first and second E_8 factors, respectively, describe an automorphism of order N=6 because

$$\sum_{i=0}^{8} k_i s_i^a = \sum_{i=0}^{8} k_i s_i^b = 6 , \qquad (3.10)$$

where k_i are the mentioned Kač labels. By using the basis given in table D.8, we find that the corresponding eight-dimensional shift vectors are given by

$$V^{a} = \frac{1}{6} \left(\alpha_{1}^{*} + \alpha_{3}^{*} \right) = \frac{1}{6} \left(4, 2, 1^{2}, 0^{4} \right) \quad \text{and} \quad V^{b} = \frac{1}{6} \alpha_{7}^{*} = \frac{1}{6} \left(2, 0^{7} \right). \tag{3.11}$$

These vectors are the two components of a shift vector acting on the 16-dimensional degrees of freedom of the heterotic string. Therefore, we obtain

$$V^{SO(10),1} = \frac{1}{6} (4, 2, 1^2, 0^4) (2, 0^7) . \tag{3.12}$$

The unbroken gauge group is found from fig. 3.2 to be

$$\mathcal{G} = E_8 \times E_8 \xrightarrow{V^{SO(10),1}} \mathcal{G}_{4D} = [SO(10) \times SU(2) \times SU(2) \times U(1)] \times [SO(14) \times U(1)].$$
 (3.13)

The U(1)'s appear because the rank is not reduced by $V^{SO(10),1}$.

As a side remark, we would like to point out that the shift vector $V^{SO(10),1}$ turns out to be of particular interest. We will see in chapter 4 that the phenomenology of models with this shift vector is very promising. This feature will be associated to the existence of an SO(10) gauge group in the resulting \mathcal{G}_{4D} and certain properties of the spectrum produced by $V^{SO(10),1}$.

Restrictions on the Shift Vector

Not every shift vector V which describes an automorphism of the algebra is an admissible choice for model construction. We have already seen in section 2.2.2 that for a twist $\vartheta \in P$ of order N, $\vartheta^N = 1$ implies that NV should act as the identity on the gauge degrees of freedom. Hence, a consistency condition on the shift vector V is

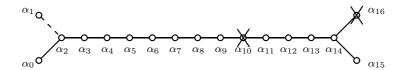
$$NV \in \Lambda. \tag{3.14}$$

Further, modular invariance of the partition function requires that

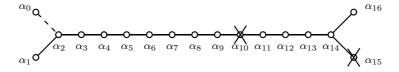
$$N(V^2 - v^2) = 0 \mod 2$$
 (3.15)

has to be satisfied, where v is the twist vector, acting in the complexified coordinates of the compact space.

From eq. (3.14) it is clear that for a given order N of the twist ϑ , all shifts V of order N' are also admissible, as long as N' divides N. In principle, we could determine the admissible shifts for each N' separately, but a more practical approach is to run through the outlined procedure for N, dropping the condition on the relative-primeness of the sequence $s = (s_0, \ldots, s_r)$. It is not hard to verify that, by dropping that condition, the order of the shift can be some N' which is smaller than N.



(a) Breaking due to $V^{(3)}$ of table D.5.



(b) Breaking due to $V^{(4)}$ from table D.5.

Figure 3.3: Extended Dynkin diagram of SO(32) corresponding to the breaking due to the shifts $V^{(3)}$ and $V^{(4)}$ of the \mathbb{Z}_4 orbifold

In the previous example, the shift vector $V^{\mathrm{SO}(10),1}$ of order 6 given in eq. (3.12) fulfills trivially eq. (3.14) because, by construction, $6V \in \Lambda^*$ and the lattice Λ is self-dual. In general, all shift vectors generated by the Dynkin diagram method satisfy $NV \in \Lambda^* = \Lambda$. Further, we see that this shift vector combined with the \mathbb{Z}_6 -II twist vector v = (1/6, 1/3, -1/2) comply with the modular invariance condition:

$$6\left(\left(V^{\text{SO}(10),1}\right)^2 - v^2\right) = 6\left(\frac{13}{18} - \frac{7}{18}\right) = 2 = 0 \mod 2 \checkmark. \tag{3.16}$$

Hence, $V^{\mathrm{SO}(10),1}$ is an allowed \mathbb{Z}_6 -II shift vector.

Discriminating Equivalent Models

Notice that the form of the shift V is not unique because the choice of the basis for the simple roots α_i and their duals is not unique. This implies the possibility that two apparently different shift vectors can lead to identical orbifold models. The reason is that the automorphisms σ_V of a Lie algebra respect certain symmetries. Let us first enumerate them.

Weyl reflections. The action of an element w of the Weyl group \mathcal{W} on a shift vector V is given by

$$V \xrightarrow{w} w(V) = V - 2\frac{\alpha \cdot V}{\alpha^2} \alpha,$$
 (3.17)

where α is an arbitrary root. The Weyl reflections corresponding to $\alpha_1, \ldots, \alpha_r$ form a basis w_1, \ldots, w_r , so that any other element of the Weyl group is described by products of these.

In SO(32), it is not hard to verify that the group of Weyl reflections alters the shift vector just by permuting its entries or changing simultaneously the sign of two entries. This follows from the fact that all simple roots have integer entries 0, ± 1 . Weyl reflections

$V^{(\#)}$	U Sector	T_1 Sector	T_2 Sector
3	$2(20, \overline{6})_{-\frac{1}{2}},$	$16(1,15)_{\frac{1}{4}},$	$10(1,15)_{\frac{1}{2}}, 6(1,\overline{15})_{\frac{1}{2}},$
	$1(1,15)_1, 1(1,\overline{15})_{-1}$	$80(1,1)_{-\frac{3}{4}}$	$10(1,1)_{\frac{3}{2}}^{2}, 6(1,1)_{\frac{3}{2}}^{2}$
4	$2(20,6)_{-\frac{1}{2}},$	$16(20,1)_{-\frac{3}{4}},$	$10(1,\overline{15})_{\frac{1}{2}}, 6(1,15)_{\frac{1}{2}},$
	$1(1,15)_{-1}, 1(\overline{1},\overline{15})_{1}$	$32(1, \overline{6})_{-\frac{1}{4}}$	$10(1,1)_{\frac{3}{2}}, \ 6(1,1)_{\frac{3}{2}}$

Table 3.1: Spectra of two \mathbb{Z}_4 orbifold models of the SO(32) heterotic string with similar symmetry breakdown patterns. The U(1) charges are written as subindices.

in $E_8 \times E_8$ are more complicated. Those related to roots with integer entries act on the shift permuting or pairwise sign-flipping its entries. However, if α is a half-integer weight of $E_8 \times E_8$ (also called spinorial), the shift vector can change nontrivially. Consider, for instance, the Weyl element w associated to the E_8 root $\alpha = \left(\frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}\right)$ applied to both eight-dimensional components of the shift vector $V^{\text{SO}(10),1}$ given in eq. (3.12). The result reads

$$w\left(V^{\text{SO}(10),1}\right) = \frac{1}{6} \left(2, 0, -1, 2, 1, -2^{3}\right) \cdot \left(\frac{3}{2}, -\frac{1}{4}^{7}\right) \tag{3.18}$$

Despite the difference between the original shift V and its rotated counterpart w(V), both have the same effect on the algebra. This can be seen as follows. A Weyl rotation does not affect only the shift, but the complete space, including the simple roots of the algebra. In fact, the Weyl group is an isomorphism of the algebra. Hence, scalar products are not modified. This implies particularly that $V \cdot \alpha_i = w(V) \cdot w(\alpha_i)$, i.e. the unbroken gauge group is not modified by Weyl reflections. A less obvious consequence is that the orbifold spectrum is not altered either.

Lattice translations. We have seen in section 2.4.3 that lattice translations can have an influence in orbifold models. Namely, models whose gauge embedding is modified by lattice vectors, i.e. brother models, can mimic the effect of discrete torsion. We also pointed out, nevertheless, that in the case of \mathbb{Z}_N orbifold models without Wilson lines, this cannot occur. It follows then that adding lattice vectors to \mathbb{Z}_N shift vectors is an allowed symmetry of the automorphisms σ_V .

Symmetries of the Dynkin diagram. It is usually argued that symmetries of the (extended) Dynkin diagram can help to avoid redundancies. Whereas $E_8 \times E_8$ does not have any symmetry of this type, SO(32) has at least the symmetry defined by the operation $\alpha_i \leftrightarrow \alpha_{16-i}$, i.e. a reflection on a vertical axis crossing α_8 . The models described by automorphisms obeying this symmetry lead to identical spectra and gauge group, confirming thereby this symmetry.

However, not all symmetries of the Dynkin diagram are symmetries of the automorphisms, as we now explain. Another symmetry of the SO(32) Dynkin diagram is spinor conjugation, that is, the redefinition of the simple roots $\alpha_{15} \leftrightarrow \alpha_{16}$. Even though this operation is a symmetry of the Dynkin diagram, it is not a symmetry of orbifolds. As an example, consider the \mathbb{Z}_4 shifts $V^{(3)}$ and $V^{(4)}$ of table D.5. They both induce the unbroken

\mathbb{Z}_N	$E_8 \times E_8$	SO(32)
\mathbb{Z}_3	5	6
\mathbb{Z}_4	12	16
\mathbb{Z}_6 -I	58	80
\mathbb{Z}_6 -II	61	75
\mathbb{Z}_7	40	56
\mathbb{Z}_8 -I	145	196
\mathbb{Z}_8 -II	146	194
\mathbb{Z}_{12} -I	1669	2295
\mathbb{Z}_{12} -II	1663	2223

Table 3.2: Comparison between the number of inequivalent \mathbb{Z}_N shifts in the $\mathbb{E}_8 \times \mathbb{E}_8$ heterotic string and in the SO(32) heterotic string [73].

gauge group $\mathcal{G}_{4D} = SO(20) \times SU(6) \times U(1)$ (see fig. 3.3), but their twisted matter spectra differ, as evident from table 3.1.

By using effectively these symmetries, one can discriminate the number of inequivalent \mathbb{Z}_N models without Wilson lines.

Advantages and Disadvantages

The classification strategy described in this section possesses certain features which make it useful to obtain all shift embeddings of \mathbb{Z}_N orbifolds. By construction, the only requirement on the shift vectors is imposed by modular invariance, eq. (3.15). One can directly determine the unbroken gauge group without need of further information about the massless spectrum. It is also straightforward to eliminate redundancies in the classification. However, this method also has some drawbacks. In considering $\mathbb{Z}_N \times \mathbb{Z}_M$ models or including Wilson lines, a complete classification requires to invoke additional mechanisms, such as adding linear combinations of U(1) directions to the shift vector(s) and Wilson lines (see e.g. [93]). In that case, the beautiful simplicity of this method is lost.

3.1.1 The Classification

We have used the method described in section 3.1 to classify all shift embeddings of $E_8 \times E_8$ and SO(32) heterotic orbifolds. In table 3.2 we list the number of models without Wilson lines found for each \mathbb{Z}_N orbifold. With exception of the number of models for \mathbb{Z}_8 -II and \mathbb{Z}_{12} orbifolds, our results for the $E_8 \times E_8$ agree with those presented in ref. [122]. We have verified that the results listed here are correct. Since not all SO(32) orbifold models were known, we have thought it would be useful to display them in a web page [123], where the following details of each of these models are provided:

• the twist vector v and the root lattice Γ , which specifies the geometry,

- the shift vector V and the corresponding gauge group,
- the matter content, listed by sectors, including all U(1) charges, where we have denoted the anomalous one by U_{1A} .

Comparing the numbers of inequivalent SO(32) models to those of $E_8 \times E_8$, we find that the SO(32) heterotic string leads to a larger amount of models. This difference can become important if Wilson lines are present. We have learnt in section 2.3.4 that the local spectra at the fixed points of orbifold models with Wilson lines correspond always to the twisted spectra of the models without Wilson lines. In other words, if one defines a local shift vector V_g for each of the constructing elements denoting the fixed points, then each V_g (which includes Wilson lines) has to be equivalent to one of the shift vectors V we have before nontrivial Wilson lines are switched on. In this sense, SO(32) orbifolds lead to a richer variety of models.

A remark is in order. In classifying orbifold models based on the SO(32) heterotic string, we were somehow surprised by the common presence of spinors in the spectra. We give some details about this feature in appendix B. This situation contrasts with the popular notion that only $E_8 \times E_8$ orbifolds admit, say, SO(10) spinors and therefore SM families of quarks and leptons. Yet we have to admit, that spinors are more frequently encountered in $E_8 \times E_8$ than in SO(32) orbifolds, what renders $E_8 \times E_8$ more attractive.

Our classification of \mathbb{Z}_N orbifolds closes one of the unfinished tasks started already in the late eighties. Nonetheless, a classification of all \mathbb{Z}_N orbifolds is far from being useful if one does not include Wilson lines, for they trigger further symmetry breakdown and, hence, the appearance of models resembling the SM, which is, by the way, the true aim of any study of this kind. Therefore, one is forced to find a useful way to study systematically models with Wilson lines.

3.2 Including Wilson Lines and $\mathbb{Z}_N \times \mathbb{Z}_M$ Orbifolds

A convincing approach to classify orbifold models with Wilson lines was first proposed in the context of \mathbb{Z}_3 orbifolds in ref. [96]. The proposal consists in a generic ansatz that describes any (shift vector or) Wilson line of order three. As explained in more detail in appendix A, we find that this approach can easily be extended to any order. Let us summarize here our findings.

By identifying those models whose shifts and Wilson lines match after the action of lattice translations and Weyl reflections, one finds the shifts of order N, M and Wilson lines of order N_{α} take a block-form:

$$V_j = (V_{j,\text{block }1}, V_{j,\text{block }2}, V_{j,\text{block }3}, \dots), \qquad (3.19a)$$

$$A_{\alpha} = (A_{\alpha, \text{block 1}}, A_{\alpha, \text{block 2}}, A_{\alpha, \text{block 3}}, \dots), \qquad (3.19b)$$

where the blocks $V_{j,\text{block i}}$ and $A_{\alpha,\text{block i}}$ are m_i -dimensional vectors. The specific form of these blocks depends on the order of the respective shift vector V_j and Wilson line A_{α} . Let

us denote by $X_{\text{block i}}$ an arbitrary block of a shift or Wilson line of order \widetilde{N} . The form of this block is then expressed by (see appendix A)

$$X_{\text{block i}} = \begin{cases} \frac{1}{\tilde{N}} \left((\pm j)^{\alpha}, -(\tilde{N} - j)^{\beta}, 0^{n_0}, 1^{n_1}, \dots, (\tilde{N} - j)^{n_{(\widetilde{N} - j)} - \alpha - \beta}, \dots, \left(\frac{\tilde{N}}{2}\right)^{n_{\frac{\widetilde{N}}{2}}} \right) & \text{case a} \\ \frac{1}{2\tilde{N}} \left((\pm j)^{\alpha}, -(2\tilde{N} - j)^{\beta}, 1^{n_1}, 3^{n_3}, \dots, (2\tilde{N} - j)^{n_{(2\widetilde{N} - j)} - \alpha - \beta}, \dots, (\tilde{N} - 1)^{n_{\widetilde{N} - 1}} \right) & \text{case b} \\ \frac{1}{\tilde{N}} \left((\pm j)^{\alpha}, -(\tilde{N} - j)^{\beta}, 0^{n_0}, 1^{n_1}, \dots, (\tilde{N} - j)^{n_{(\widetilde{N} - j)} - \alpha - \beta}, \dots, \left(\frac{\tilde{N} - 1}{2}\right)^{n_{(\frac{\widetilde{N} - 1}{2})}} \right) & \text{case c} \end{cases}$$

where α and β are either 0 or 1 such that $\alpha + \beta = 0, 1$ and $\sum_k n_k = m_i$. Furthermore, j depends on the value of \widetilde{N} and the kind of building block, according to the following cases:

- a) even order \widetilde{N} , 'vectorial' block. 'Vectorial' means here that the entries have a maximal denominator of \widetilde{N} . In this case j takes the values $\{\frac{\widetilde{N}}{2}+1,\ldots,\widetilde{N}-1\}$;
- b) even order \widetilde{N} , 'spinorial' block. 'Spinorial' means that the entries have a maximal denominator of $2\widetilde{N}$ and odd numerators. We have $j \in \{\widetilde{N}+1,\widetilde{N}+3,\ldots,2\widetilde{N}-1\}$;
- c) odd order \widetilde{N} . In this case 'vectorial' and 'spinorial' blocks are equivalent; so it suffices to give one ansatz, where $j \in \{\frac{\widetilde{N}+1}{2}, \frac{\widetilde{N}+3}{2}, \dots, \widetilde{N}\}$.

Shift vectors and Wilson lines obtained by the repeated use of the ansatz (3.20) have to satisfy the modular invariance conditions, eq. (2.42), and the embedding constraint $NV_1, MV_2, N_{\alpha}A_{\alpha} \in \Lambda$. In contrast to the Dynkin diagram strategy, the latter requirement is not automatically fulfilled. Indeed, that requisite implies that all blocks of a shift or Wilson line must be of the same type (either a, b or c). Once both consistency conditions are met by a set of parameters, one has encountered an admissible orbifold model. In this way, one arrives at a more general classification of orbifold models.

There is however a subtlety. As we have explained in section 2.4.3, brother models can appear by considering lattice translations of the shift(s) and/or Wilson lines. The matter spectrum of a brother to an orbifold model differs in the twisted sectors. Therefore, a model and its brother are not equivalent. Since in the derivation of the ansätze (3.20) we have taken lattice translations as a symmetry of the theory, brother models have been disregarded and thus a classification based on this method is not comprehensive. On the other hand, it is worth to mention that models with the same matter spectrum of the brother models can be found by varying the discrete torsion parameters. In a sense, then, a classification based on the proposed ansätze provides us with the full set of inequivalent torsionless models.

Discriminating Equivalent Models

In this method, we have tried to avoid redundancies by disregarding those gauge embeddings differing by lattice translations, and pairwise sign-flips and permutations of their entries. However, in $E_8 \times E_8$ models, as mentioned in section 3.1, there are additional Weyl

reflections (those based on the spinorial simple root), which modify in general shifts and Wilson lines in a very unpredictable way. In other words, our present strategy has not projected out all duplicities. Inspecting the symmetries due to those other Weyl symmetries is hopeless due to the huge amount of elements of the Weyl group \mathcal{W} . Hence, we have to look for an alternative to eliminate equivalent models.

It turns out more practical to use shifts and Wilson lines to compute the massless spectrum of each model and, then, consider two models as inequivalent if and only if their massless spectra are different. Of course, this job gets too ambitious if one has at hand a great many models and one lacks computer support. That might be one of the reasons why a comprehensive classification of orbifolds has been postponed for several years.

To perform the comparison of the spectra in a reasonable time, we compare the non-Abelian massless spectra and the number of singlets, that is, we avoid a comparison of U(1) charges. This underestimates the true number of models somewhat. Ideally, we should not only compare the full massless spectra (including U(1) charges), but also the localization of the particles and the couplings among them. Furthermore, it might be also necessary to consider the Kaluza-Klein tower of nonzero modes. This is evidently unfeasible.

We would like to point out that the problem of identifying inequivalent models does not appear only in this method. Also Dynkin diagram techniques suffer of the same problem when considering Wilson lines. An interesting question would be to find out whether there is another way to discriminate duplicities in a classification.

Advantages and Disadvantages

The classification method presented in this section is an alternative to classify orbifold models with Wilson lines. By means of computer programs, one can generate all models of a given orbifold in a direct manner. However, this method neglects brother models, or analogously, models with discrete torsion. Therefore, one is forced to vary the discrete torsion parameters in order to arrive to a complete set of models. Further, one cannot avoid to overcount some models, which can only be identified by directly comparing the states of their matter spectra.

3.2.1 Sample Classification of $\mathbb{Z}_3 \times \mathbb{Z}_3$ without Wilson Lines

As a concrete application, let us describe the classification of $\mathbb{Z}_3 \times \mathbb{Z}_3$ orbifolds without Wilson lines. Here we make use of both classification methods introduced in this chapter. A $\mathbb{Z}_3 \times \mathbb{Z}_3$ orbifold model is described by the twist vectors

$$v_1 = (0, 1/3, 0, -1/3)$$
 and $v_2 = (0, 0, 1/3, -1/3)$, (3.21)

two shift vectors V_1 and V_2 of order three and three Wilson lines, which are not important for our current discussion. We aim here basically at a classification of all admissible shift embeddings.

$V^{(\#)}$	Shift vector V_1	6D gauge group \mathcal{G}_{6D}
1	$\frac{1}{3} \left(0^6, 1^2\right) \left(0^8\right)$	$E_7 \times E_8$
2	$\frac{1}{3}(0^6, 1^2)(0^5, 1^2, 2)$	$E_7 \times E_6 \times SU(3)$
3	$\frac{1}{3}(0^3, 1^4, 2)(0^8)$	$SU(9)\times E_8$
4	$\frac{1}{3}(0^6, 1^4, 2)(0^5, 1^2, 2)$	$SU(9)\times E_6\times SU(3)$
5	$\frac{1}{3}(0^6, 1^6, 2)(0^6, 1^6, 2)$	$SO(14) \times SO(14)$

Table 3.3: Five inequivalent $E_8 \times E_8$ shift vectors V_1 of the $\mathbb{Z}_3 \times \mathbb{Z}_3$ orbifold and their induced gauge symmetry breakdown.

First, by using the Dynkin diagram strategy depicted in section 3.1, one finds that there are only five consistent shift vectors V_1 , which can be written in the generic form

$$V_1 = \frac{1}{3}(0^{n_0}, 1^{n_1}, 2^{\alpha})(0^{n'_0}, 1^{n'_1}, 2^{\beta}), \tag{3.22}$$

where α and β can be either 0 or 1, and $n_i, n_i' \in \mathbb{Z}$, such that $n_0 + n_1 + \alpha = n_0' + n_1' + \beta = 8$. The admissible shifts (listed in table 3.3) can be obtained by verifying the consistency conditions

$$3(V_1^2 - v_1^2) = 0 \mod 2 \qquad \text{and} \qquad 3V_1 \in \Lambda.$$

Note that the modular invariance condition forbids the trivial shift $V_1 = (0^8)(0^8)$. This is a difference with respect to the \mathbb{Z}_3 orbifold.

From the ansatz (3.20), it follows that the second shift V_2 of order three has the generic form

$$V_{2} = \frac{1}{3} \left(\begin{pmatrix} 3 \\ \vdots \\ -2 \end{pmatrix}, \begin{pmatrix} 1 \\ 0 \end{pmatrix}^{n_{0}-1}, \begin{pmatrix} 1 \\ 0 \\ -1 \end{pmatrix}^{n_{1}+\alpha} \right) \left(\begin{pmatrix} 3 \\ \vdots \\ -2 \end{pmatrix}, \begin{pmatrix} 1 \\ 0 \end{pmatrix}^{n'_{0}-1}, \begin{pmatrix} 1 \\ 0 \\ -1 \end{pmatrix}^{n'_{1}+\beta} \right) . (3.23)$$

Together with one of the shift vectors V_1 from table 3.3, the shift vectors generated by this ansatz are subject to the usual lattice conditions and modular invariance, eq. (2.42). Disregarding those models that are equivalent at massless level, we get 109 shift embeddings.

Generalized Discrete Torsion

We use now the set of shift embeddings to generate all admissible models. As discussed in the examples of section 2.4.2, there are ten independent generalized discrete torsion parameters, whose values are $0, \frac{1}{3}$ or $\frac{2}{3}$. One might be tempted to deduce that the total number of models is 109×3^{10} , but out of them only 1082 models (or more precisely, massless spectra) are inequivalent. These models comprise the complete set of admissible models without Wilson lines. The model definitions and the resulting spectra are detailed in our web page [123].

3.3 The C++ Orbifolder

The extensive work that an orbifold classification requires, would not be feasible without computer programs. In collaboration with P.K.S. Vaudrevange and A. Wingerter, we have written the C++ Orbifolder. This program, which will be available to the scientific community, has been thought to cover the needs of a researcher interested in the phenomenological consequences of orbifold compactifications without dealing with the time-consuming and cumbersome computation.

Provided the shift vector(s) and Wilson lines, the C++ Orbifolder can compute the massless spectrum of the $E_8 \times E_8$ or SO(32) heterotic string compactified on any symmetric \mathbb{Z}_N and $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifold in much less than a second on a 2.66 GHz computer. Its results are given in a human readable format, which can be saved as TXT- or TEX-file. The underlying lattice Γ as well as the inclusion of generalized discrete torsion can be chosen at will. In computing orbifold spectra, the program verifies all consistency constraints, such as the conditions for cancellation of anomalies and modular invariance.

One can profit of the speed of the C++ Orbifolder in a classification of orbifold models. Some routines of this program use the classification methods outlined in this chapter. Thus, one can generate a large amount of models with and without discrete torsion in a reasonable time. In particular, the comparison of two different spectra takes less than a second, independently of how complicated they are.

A very useful feature of this program is that it is capable to determine all nonvanishing couplings entering in the superpotential of a given model, according to the string selection rules enumerated in section 2.5.1. This process has to be performed order by order in the superpotential and therefore is very demanding. As yet the program can determine the superpotential for \mathbb{Z}_6 -II orbifolds up to order eight in the fields within one week, in average. For other orbifolds, like e.g. \mathbb{Z}_3 or \mathbb{Z}_{12} , this computation can take significantly less time because there are less particles in the spectra and/or less allowed couplings.

A phenomenological viable orbifold model must have at least the SM gauge group and all possible exotic particles have to be decoupled from the low-energy spectrum. To pick up a nonanomalous linear combination of U(1)'s that play the role of the hypercharge is a challenge. We have developed routines to perform that task in the case when the hypercharge can be embedded in an SU(5) GUT theory. Moreover, the Orbifolder can identify the SM particles and determine whether additional exotics acquire large masses when some SM singlets get VEVs of the order of the string mass scale.

The C++ Orbifolder is, however, still under development. Some of the aspects that are currently in an advanced stage include: discrete flavor symmetries of the superpotential, continuous global accidental symmetries of the superpotential, and resolution of orbifold singularities. One might be interested also in considering asymmetric orbifold constructions and a systematic study of discrete accidental symmetries of the superpotential, which can shed light, for instance, in the solution to the problem of proton stability.

Chapter 4

A Mini-Landscape of Z₆-II Orbifolds

Inspired by the works by Kobayashi et al. [97] and Buchmüller et al. [48], we aim at promising models of the \mathbb{Z}_6 -II orbifold. We perform a systematic search of models with two Wilson lines guided by the concept of local GUTs, which is introduced in section 4.1. It turns out that out of 3×10^4 \mathbb{Z}_6 -II orbifold models with local GUTs, about 200 models have the exact matter spectrum of the MSSM and some other realistic properties. This chapter focuses mainly on the methodology of our search. We list the main results here and the discussion of the phenomenology of the models is left for chapter 5.

4.1 Orbifold Local GUTS

Let us focus on \mathbb{Z}_N orbifold compactifications. Consider a fixed point or fixed torus with constructing element $g = (\vartheta^k, n_\alpha e_\alpha)$. Thus, the associated local gauge shift and local twist vector are given by $V_g = kV + n_\alpha A_\alpha$ and $v_g = kv$, respectively. We have seen in section 2.3.4 that at the fixed point g, the local gauge group \mathcal{G}_g is larger than the four-dimensional unbroken gauge group \mathcal{G}_{4D} . Furthermore, the matter states living at the fixed point g furnish complete representations under \mathcal{G}_g , which can nonetheless be partly projected out when considering the complete orbifold by the action of Wilson lines. On the other hand, massless modes of untwisted (bulk) fields also build representations under \mathcal{G}_g , but they are always partly projected out (even in the absence of background fields). Therefore, bulk states are, in general, incomplete (also called 'split') multiplets from the four-dimensional perspective.

A useful concept in orbifold compactifications is that of local grand unification, or *local* GUTs, which appear when the local symmetry \mathcal{G}_g at some fixed point(s) is a GUT symmetry, such as SO(10) or E₆. In presence of local GUTs, the SM gauge symmetry can arise as the intersection of different local GUT groups.

If we suppose, for example, that the local gauge symmetry is SO(10) whereas the unbroken gauge group in four dimensions is $\mathcal{G}_{4D} = G_{SM} = SU(3)_c \times SU(2)_L \times U(1)_Y$, then the states attached to the fixed points can transform locally as **16**-plets, which under the

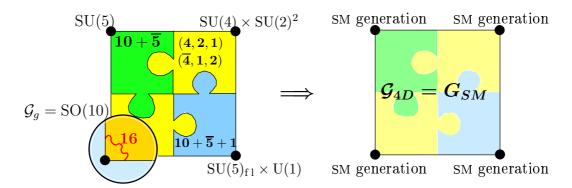


Figure 4.1: Local GUTs at different fixed points: local vs. global picture. The intersection of all local groups is $\mathcal{G}_{4D} = G_{SM}$.

SM gauge group, form the representations

$$\mathbf{16} = (\mathbf{3}, \mathbf{2})_{1/6} + (\overline{\mathbf{3}}, \mathbf{1})_{-2/3} + (\overline{\mathbf{3}}, \mathbf{1})_{1/3} + (\mathbf{1}, \mathbf{2})_{-1/2} + (\mathbf{1}, \mathbf{1})_1 + (\mathbf{1}, \mathbf{1})_0. \tag{4.1}$$

Similarly, $\mathbf{10}$ -plets of SO(10) can appear in the same (or in a different) local spectrum. They correspond to the SM representations

$$\mathbf{10} = (\mathbf{3}, \mathbf{1})_{-1/3} + (\mathbf{1}, \mathbf{2})_{1/2} + (\overline{\mathbf{3}}, \mathbf{1})_{1/3} + (\mathbf{1}, \mathbf{2})_{-1/2}, \tag{4.2}$$

which are those of the Higgs doublets and Higgs triplets. As we already mentioned, once we abandon the local picture and consider the orbifold globally, some parts of the **16**– or **10**–plets might not be present in the global spectrum due to the orbifold projections. In the best of the cases, the parts of a local **10**–plet which are projected out will be those corresponding to Higgs triplets, whereas the complete **16**–plets of the matter generation will survive. This situation would then endow the theory with a natural explanation of the structure of the SM families [48, 50, 53, 54]. It would additionally provide a solution to the doublet-triplet splitting problem of 4D GUTS [124].

It is noteworthy that such an ideal case as the example just presented, including other GUT symmetries, can be realized in orbifold constructions. We have to demand that matter generations appear from local GUTs localized at specific points, where the action of the orbifold group on the associated states is trivial. Furthermore, it is necessary that the Higgs representations come from the bulk or from fixed points where the orbifold projections can partly alter their structure. Since solely higher twisted-sector fields are affected by orbifold projections, the first requirement is met when local GUTs appear at the fixed points of the first twisted sector and the local matter spectrum includes GUT generations (such as 27-plets of E_6 or 16-plets of SO(10)).

In section 2.3.4, we have concluded that a local shift vector V_g has to be equivalent to one of the shift vectors V that define \mathbb{Z}_N orbifold models without Wilson lines. Therefore,

That the states in the first twisted sector are not affected by the orbifold projection follows from the fact that the projection phase, eq. (2.85), is trivial for T_1 (and T_5) twisted states satisfying the masslessness condition, eq. (2.81).

finding local GUTs with SM matter generations at fixed points of the T_1 sector amounts to first identifying those shifts leading to the desired GUT symmetry as the gauge group of the theory in four dimensions and GUT generations in the T_1 sector, and then selecting a suitable set of Wilson lines inducing the symmetry breakdown $\mathcal{G}_{GUT} \to G_{SM}$ (see fig. 4.1).

One of the qualities of this method is that it ensures the existence of standard GUT hypercharge, which is consistent with gauge coupling unification, although no GUT appears in 4D. Another advantage is that, even though matter fields form locally complete GUT representations, interactions generally break GUT relations since different local GUTs are supported at different fixed points. In particular, mixing of the localized generations with vectorlike bulk states breaks the unwanted GUT relations for the fermion masses [48, 125]. As already mentioned, this approach can explain why the SM gauge and Higgs bosons do not form complete GUT multiplets, while the matter fields do.

In the following sections, the concept of local GUTs is used as a guiding principle in the search for orbifold models with realistic features.

4.2 MSSM Search Strategy

It is well known that with a suitable choice of Wilson lines it is not difficult to obtain the SM gauge group up to U(1) factors. The real challenge is to get the exact matter spectrum of the SM or, more precisely, of its minimal supersymmetric extension, the MSSM. Demanding additionally gauge coupling unification constrains possible models.² In order to find models which combine both properties, we base our strategy on the concept of local GUTs described in the previous section.

An SO(10) GUT has very compelling features which make it appropriate for model building [52]. Among other reasons, it incorporates the success of Georgi-Glashow and Pati-Salam GUTs, provides a single gauge coupling, predicts the existence of right-handed neutrinos and gathers a complete SM generation within a single representation. Hence, we shall be mostly interested in the gauge shifts V which allow for a local SO(10). Also E₆ local GUTs can be reasonable promising as they embed the SO(10) structure. That is, the shift vectors V we will consider are such that the left-moving momenta p satisfying

$$p \cdot V = 0 \mod 1, \quad p^2 = 2 \tag{4.3}$$

are roots of SO(10) or E_6 (up to extra group factors). Furthermore, the massless states of the T_1 sector are required to contain **16**-plets of SO(10) at the fixed points with SO(10) symmetry or **27**-plets of E_6 at the fixed points with E_6 symmetry.

Since these massless states from T_1 are automatically invariant under the orbifold action, they all survive in 4D and appear as complete GUT multiplets. In the case of SO(10), that gives one complete SM generation, while in the case of E_6 we have 27 = 16 + 10 + 1 under SO(10). It is thus necessary to decouple all (or part) of the 10-plets from the low-energy theory.

²In refs. [34,49], only models without standard hypercharge are investigated. We have verified that, in fact, \mathbb{Z}_3 and \mathbb{Z}_4 orbifolds do not allow for GUT hypercharge normalization.

Shift	4D gauge group \mathcal{G}_{4D}
$V^{\mathrm{SO}(10),1}$	$SO(10) \times SU(2)^2 \times SO(14) \times U(1)^2$
$V^{SO(10),2}$	$SO(10) \times SU(3) \times E_7 \times U(1)^2$
$V^{\mathrm{E}_6,1}$	$E_6 \times E_8 \times U(1)^2$
$V^{\mathrm{E}_6,2}$	$E_6 \times SU(3) \times E_7 \times U(1)$

Table 4.1: Gauge group associated to shift vectors with local GUT structure. Their matter spectrum in the T_1 (or equivalently T_5) sector includes 12 GUT families.

The Wilson lines will be chosen such that the standard model gauge group is embedded into the local GUT as

$$G_{SM} \subset SU(5) \subset SO(10) \text{ or } E_6$$
. (4.4)

It follows then that the hypercharge is that of standard GUTs and thus consistent with gauge coupling unification. Hypercharge will appear as a linear combination of several U(1)s and, thus, the generator of hypercharge (in a standard basis),

$$\mathsf{t}_Y = \left(-\frac{1}{2}, -\frac{1}{2}, \frac{1}{3}, \frac{1}{3}, \frac{1}{3}\right),\tag{4.5}$$

will be embedded in a 16-dimensional vector acting on the gauge degrees of freedom. Note that gauge coupling unification in orbifold models is due to the fact that the 10D (not 4D) theory is described by a single coupling.

Our model search is carried out in the \mathbb{Z}_6 -II orbifold compactification of the $E_8 \times E_8$ heterotic string. Some properties of this orbifold were described in section 2.2.3 (see the discussion around eq. (2.49)). In this construction, there are two gauge shifts leading to a local SO(10) GUT [93, 126],

$$V^{\text{SO}(10),1} = \left(\frac{1}{3}, \frac{1}{2}, \frac{1}{2}, 0, 0, 0, 0, 0\right) \quad \left(\frac{1}{3}, 0, 0, 0, 0, 0, 0, 0, 0\right) ,$$

$$V^{\text{SO}(10),2} = \left(\frac{1}{3}, \frac{1}{3}, \frac{1}{3}, 0, 0, 0, 0, 0\right) \quad \left(\frac{1}{6}, \frac{1}{6}, 0, 0, 0, 0, 0, 0, 0\right) ,$$

$$(4.6)$$

and 2 shifts leading to a local E_6 GUT,

$$V^{E_6,1} = \left(\frac{1}{2}, \frac{1}{3}, \frac{1}{6}, 0, 0, 0, 0, 0\right) \quad (0, 0, 0, 0, 0, 0, 0, 0) ,$$

$$V^{E_6,2} = \left(\frac{2}{3}, \frac{1}{3}, \frac{1}{3}, 0, 0, 0, 0, 0\right) \quad \left(\frac{1}{6}, \frac{1}{6}, 0, 0, 0, 0, 0, 0\right) . \tag{4.7}$$

The symmetry breakdown of $E_8 \times E_8$ due to these shift vectors is given in table 4.1. The shift $V^{SO(10),1}$ of eq. (3.12) and the one presented here are equivalent. They differ because we have used Weyl transformations to reduce the number of nonzero entries in order to facilitate later analysis. We will focus on these shifts and scan over possible Wilson lines, employing the classification method of section 3.2, to get the SM gauge group. Due to its geometry, the \mathbb{Z}_6 -II orbifold with lattice $G_2 \times SU(3) \times SO(4)$ allows for up to two Wilson lines of order 2, A_5 and A_6 , and for one Wilson line of order 3, A_3 , as illustrated in fig. 4.2.

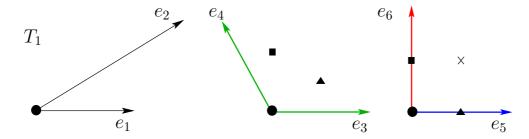


Figure 4.2: T_1 sector of the \mathbb{Z}_6 -II orbifold on a $G_2 \times SU(3) \times SO(4)$ lattice. The SU(3) torus admits one Wilson line A_3 of order 3 (green) associated to the directions e_3 and e_4 while the SO(4) torus can allocate up to two Wilson lines, A_5 (blue) and A_6 (red), of order 2.

The next question is how to get three matter generations. A geometric explanation of the origin of the three generations is somehow attractive and can be realized in orbifold compactifications. As depicted in fig. 4.2, there are 3×4 fixed points in the T_1 sector, which, in the absence of Wilson lines, are degenerate, that is, are furnished with identical matter spectrum. Wilson lines lift this degeneracy in the direction they act. For example, if the Wilson line A_5 related to the torus direction e_5 is nontrivial, then in the SO(4) torus the points \blacksquare and \bullet as well as \times and \blacktriangle are equivalent, but the horizontal symmetry is lost. The points that keep their local GUT structure are in this case the six ones related to \blacksquare and \bullet in the SO(4) plane. Hence, the simplest possibility to get three matter generations is to use three equivalent fixed points with 16-plets [53] which appear in models with two Wilson lines of order 2. If the extra states are vectorlike and can be given large masses, the low-energy spectrum will contain three matter families. However, if one starts from the shifts given above, this strategy fails since all such models contain chiral exotic states [48]. In the case of E_6 , it does not work either since a simple analysis of the Dynkin diagram of E_6 reveals that one cannot obtain $G_{\rm SM} \subset {\rm SU}(5) \subset E_6$ with two Wilson lines of order 2.

The next-to-simplest possibility is to use two equivalent fixed points which give rise to two matter generations. This situation can be obtained in models with one Wilson line of order 3 and another of order 2. The third generation would then have to come from other twisted or untwisted sectors. The appearance of the third complete family can be linked to the SM anomaly cancellation. Indeed, since the shift vectors of eqs. (4.6) and (4.7) lead to GUT families in the untwisted sector, that sector contains part of a 16-plet after the inclusion of Wilson lines. Then the simplest options consistent with the SM anomaly cancellation are that the remaining matter either completes the 16-plet or provides vectorlike partners of the untwisted sector. In more complicated cases, additional 16- or $\overline{16}$ -plets can appear. The localized 16- and 27-plets are true GUT multiplets, whereas the third or "bulk" generation only has the SM quantum numbers of an additional 16-plet.

We find that the above strategy is successful. First, we notice that the model presented

³The remaining vertical symmetry is usually identified with D_4 or S_2 , which in adequate models can be interpreted as a family symmetry.

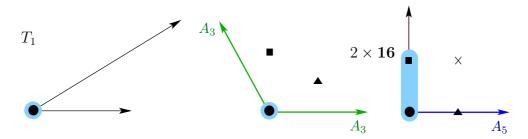


Figure 4.3: A successful approach to get models with three matter generations. Two GUT generations are located at fixed points of the T_1 sector while the third one comes from other untwisted or twisted sectors.

in refs. [48, 50] adjusts precisely to this scheme. Secondly, also in the case of E_6 GUTs we observe that it is not difficult to obtain the SM gauge group. As we will detail in section 4.4, in fact, other massless states present in the spectra of models based on this strategy are often vectorlike with respect to the SM gauge group and can be given large masses consistent with string selection rules.

Once established the basis of our study, let us specify now all other details of our search strategy. We enumerate the steps of our search as follows:

- ① Generate all models with two Wilson lines.
- ② Identify "inequivalent" models.
- ③ Select models with $G_{SM} \subset SU(5) \subset SO(10)$ or E_6 .
- 4 Select models with three net (3,2).
- ⑤ Select models with nonanomalous $U(1)_Y \subset SU(5)$.
- © Select models with net three SM families + Higgses + vectorlike.

Notice that in step \oplus we have decided not to specialize to the case of one Wilson line of order 2 and one of order 3. Even though we already know that only in that case we can obtain realistic models, we would like also to know how frequently those models appear in the landscape patch that we have chosen. In other words, it might be useful to count with the total number of models in order to draw some reliable statistical results out of our search. To do so, we apply the classification method described in section 3.2, that is, we propose an ansatz describing generically Wilson lines of order 2 and 3, based on eqs. (3.19b) and (3.20).

Not all models obtained in step 1 are inequivalent. In step 2, we consider two models to be "equivalent" if they have identical spectra with respect to nonabelian gauge groups and have the same number of nonabelian singlets. Thus, models differing only in U(1) charges are treated as equivalent. In addition, some models differ only by the localization of states on the different fixed points. We know that these ambiguities occur and it is possible that in some cases Yukawa couplings are affected. Hence our criterion may underestimate somewhat the number of truly inequivalent models.

The criterion @ is crucial for obtaining viable models. We require particularly that $SU(3)\times SU(2)$ appears embedded in an SU(5) subgroup of the local GUTs. This constraint

is imposed in order for the hypercharge to be of Georgi-Glashow type and, hence, consistent with gauge coupling unification.

Of course, we could try immediately to identify an appropriate hypercharge. It proves however to be more practical to first reduce further the number of models available. Criterion \oplus amounts to finding those models where three copies of quark-doublet-like particles exist. Let us just remark that, in general, both representations (3, 2) and $(\overline{3}, 2)$ under $SU(3) \times SU(2)$ appear in the spectra. They form some vectorlike pairs that will eventually decouple from the low-energy spectrum. Thus, only models satisfying $|\#(3, 2) - \#(\overline{3}, 2)| = 3$ will lead to realistic phenomenology.

After the above procedure, it is then straightforward to identify the hypercharge. In an ideal case, the hypercharge generator would be given by (4.5) embedded in a 16-dimensional vector. However, this form is basis-dependent. In general, what one does is to compare

the simple roots of the SU(5) embedded in the GUT group to those of $SU(3)\times SU(2)$. Since the roots of SU(5) span a four-dimensional space whereas those of $SU(3)\times SU(2)$ only span a three-dimensional space, the hypercharge generator t_Y

is uniquely determined by the sole direction orthogonal to the simple roots of $SU(3)\times SU(2)$ which lies in the space of SU(5). In other words, the hypercharge results directly from the symmetry breakdown $SU(5) \longrightarrow G_{SM}$. Ambiguities arise in certain seldom cases when $U(1)_Y$ can be defined in different ways. In this case, we have to count the model twice. On the other hand, the hypercharge generator can unfortunately mix with the anomalous U(1) of the orbifold, what leads to the undesirable consequence that it has to be broken at very high energies. Therefore, at step s, one is forced to disregard those models where the hypercharge appears anomalous.

We are in position to identify the MSSM candidates. In step ©, we have to verify that the resulting spectrum contains exactly the spectrum of the MSSM, i.e. three net matter generations and, at least, one pair of Higss-doublets with the correct quantum numbers under $G_{SM} = SU(3)_c \times SU(2)_L \times U(1)_Y$. Other states in the spectra of admissible models, generically called exotics must form vectorlike pairs.

4.3 A Fertile Patch in the Landscape

The results of our search strategy are presented in table 4.2. Interestingly enough, this strategy has allowed us to find 223 explicit examples with the MSSM matter content plus additional vectorlike particles. These models represent a major result in the context of string compactifications and are considered one of the central results of this work. We see that most of the MSSM candidates arise from a local SO(10) GUT. Naïvely, we can consider this result to be another argument why SO(10) GUTs might be preferred to other GUTs. On the other hand, it might just reflect the fact that two Wilson lines cannot break E_6 enough to get a comparable amount of models with G_{SM} gauge symmetry. One might then expect that adding a third Wilson line, many more models with promising properties arise

Criterion	$V^{\mathrm{SO}(10),1}$	$V^{\mathrm{SO}(10),2}$	$V^{{ m E}_6,1}$	$V^{{ m E}_6,2}$
② inequivalent models with 2 Wilson lines	22,000	7,800	680	1,700
③ SM gauge group $\subset SU(5) \subset SO(10)$ (or E_6)	3,563	1,163	27	63
④ 3 net (3,2)	1,170	492	3	32
© nonanomalous $U(1)_Y \subset SU(5)$	528	234	3	22
© spectrum = 3 generations + vectorlike	128	90	3	2

Table 4.2: Statistics of \mathbb{Z}_{6} -II orbifold models based on the shifts $V^{SO(10),1}, V^{SO(10),2}, V^{E_{6},1}, V^{E_{6},2}$ with two Wilson lines.

from the E_6 shifts.

We find that the properties of the models with the chiral MSSM matter content, such as the number and type of vectorlike exotics and SM singlets, are so similar that there is no model which can be preferred a priori. This implies that, even though one is also interested in studying the phenomenological features of each model, some interesting conclusions can be drawn in a generic form. The details of the models are listed in [123].

It is instructive to compare our model scan to others. In certain types of intersecting D-brane models, it was found that the probability of obtaining the SM gauge group and three generations of quarks and leptons, while allowing for chiral exotics, is 10^{-9} [43, 44]. More recently, in ref. [45] intersecting D-branes on \mathbb{Z}_6 -II orbifolds have lead to the conclusion that the probability of finding such matter configurations can be enhanced to 10^{-8} . The criterion which comes closest to the requirements imposed in refs. [43–45] is 4. We find that within our sample the corresponding probability is 5%.

In refs. [40, 41], orientifolds of Gepner models were scanned for chiral MSSM matter spectra, and it was found that the fraction of such models is 10^{-14} . In our set of models, the corresponding probability, i.e. the fraction of models passing criterion ©, is of order 1%. Note also that, in all of our models, hypercharge is normalized as in standard GUTs and thus consistent with gauge coupling unification.

This comparison shows that our sample of heterotic orbifolds is unusually "fertile" compared to other constructions. The probability of finding something close to the MSSM is much higher than that in other patches of the landscape analyzed so far. It would be interesting to extend these results to other regions of the landscape where promising models exist [51,127–129].

We would like to make an additional remark. Our strategy to determine the hypercharge is, of course, not the only choice. One could instead express $U(1)_Y$ as an arbitrary linear combination of all U(1)'s (not only of those embedded in the local GUT symmetry) which gives the correct values of hypercharge to the MSSM particles. This approach was followed in ref. [130]. The authors of ref. [130] find that the majority of the models at step \oplus allow for a definition of a nonanomalous $U(1)_Y$. However, only in 12% of those models, hypercharge is in harmony with coupling unification. That means, in particular, that even

in a more general scheme, relaxing the demand $U(1)_Y \subset SU(5)$, (almost) only those 223 models at step © of our search meet all the phenomenological properties we require.

4.4 Towards Realistic String Models

Taking as a base the MSSM candidates passing criteria 3-6, there are now many ways to address the question of their phenomenological viability. One could, for example, opt for a model-by-model approach, in an attempt to identify one model with many features that match the known low-energy physics and, simultaneously, to find the solution to as many as possible of the currently open questions in theoretical high energy physics. By that approach, certainly one might find one model with some beautiful properties that also gives answer to some puzzles. However, a different model will have some other nice features and will solve some other problems.

We would rather follow an alternative approach. Provided that the MSSM candidates are very similar, one could ask general questions, such as whether the vectorlike exotics decouple, or whether SUSY preserving vacuum configurations are realizable from these models, among other matters. The answers to those questions are, of course, model dependent, but, on a statistical footing, they will probably yield some predictions or, at least, exclude some regions of the landscape.

4.4.1 Coupling Selection Rules in \mathbb{Z}_6 -II

Let us start by stating the string selection rules which will be repeatedly used in our study. Assume a coupling between a number of states Ψ_i . Based on our discussion of section 2.5.1, couplings entering the superpotential must satisfy the following rules.

Gauge invariance. The usual rule on the left-moving momenta $p_{\text{sh},i}$ of the states Ψ_i , $\sum_i p_{\text{sh},i} = 0$, applies.

R-charge conservation.⁴ Since the twist vector reads v = (0, 1/6, 1/3, -1/2), the order of the corresponding (discrete) R-symmetries is respectively 6, 3 and 2. Therefore, a coupling in the superpotential is invariant if the R-charges of the states involved fulfill the conditions

$$\sum_{i} R_i^1 = -1 \mod 6, \quad \sum_{i} R_i^2 = -1 \mod 3, \quad \sum_{i} R_i^3 = -1 \mod 2. \tag{4.8}$$

Space group selection rule. In general, this rule reads $\prod_i g_i \simeq (1, 0)$ with $g_i = (\vartheta^i, n^i_{\alpha} e_{\alpha}) \in S$ denoting the constructing elements of the states entering the coupling. It can be restated in \mathbb{Z}_6 -II orbifolds as the following set of constraints:

$$\frac{1}{6}\sum_{i}k_{i} = 0 \mod 1, \tag{4.9a}$$

 $^{^4}$ R-symmetries do distinguish between fermions and bosons, therefore, in some approaches [131], the remaining discrete subgroup after breaking these symmetries can well be identified with an R-parity. We do not follow this approach.

$$\frac{1}{3}\sum_{i}(n_3^i + n_4^i) = 0 \mod 1, \tag{4.9b}$$

$$\frac{1}{2} \sum_{i} n_5^i = 0 \mod 1, \tag{4.9c}$$

$$\frac{1}{2} \sum_{i} n_6^i = 0 \mod 1. \tag{4.9d}$$

Notice that, in this notation, these rules can be identified with conservation of certain discrete \mathbb{Z}_n symmetries with (integer) charges k, $n_3 + n_4$, n_5 and n_6 . Some consequences of the appearance of these symmetries have been studied in ref. [109].

4.4.2 Decoupling Exotic Particles

Now we are in position to investigate the decoupling of vectorlike extra matter $\{x_i\}$. One could argue that the existence of these exotics in the spectra of our MSSM candidates constitutes a problem by itself since vectorlikeness does not guarantee immediately that such exotics get large masses. Therefore, we have to corroborate whether the extra matter can be given a large mass by explicitly computing the couplings that endow exotics with masses. The mass terms for such states are provided by the superpotential

$$W = x_i \,\bar{x}_j \,\mathcal{M}_{x\bar{x}}^{ij} \equiv x_i \,\bar{x}_j \,\langle s_1 \, s_2 \dots \rangle \,, \tag{4.10}$$

where s_1, s_2, \ldots are SM singlets. Some singlets are required to get large (close to $M_{\rm str}$) vevs in order to cancel the Fayet–Iliopoulos (FI) term of the anomalous U(1) intrinsic to the majority of orbifold compactifications. Then, if the relevant Yukawa couplings are allowed by the string selection rules discussed above, mass terms as that of eq. (4.10) appear, making the vectorlike matter heavy. Thus the exotic particles decouple from the low–energy theory.

Clearly, one cannot switch on the singlet VEVs at will. Instead, one has to ensure that they are consistent with SUSY. Supersymmetry requires vanishing F- and D-terms. The number of the F-term equations equals the number of complex singlet fields s_i , therefore there are in general nontrivial singlet configurations with vanishing F-potential. The D-terms can be made zero by complexified gauge transformations [132] if each field enters a gauge invariant monomial [133]. Thus, to ensure that the decoupling of exotics is consistent with SUSY, one has to show that all SM singlets appearing in the mass matrices for the exotics enter gauge invariant monomials involving only SM singlets and carrying anomalous charge. However, for simplicity, we will assume momentarily that all SM singlets entering the mass matrices develop large supersymmetric VEVs.⁵

In many cases, vectorlike exotics happen to have the same quantum numbers as the SM particles. Hence, in the process of decoupling, the vectorlike states can mix with the

 $^{^{5}}$ This assumption will be confirmed in the next chapter by examining carefully the decoupling of exotics in SUSY preserving vacua.

localized 16– and 27–plets (as long as it is allowed by the SM quantum numbers) such that the physical states at low energies are *neither* localized *nor* "true" GUT multiplets. Even though part of the beauty of a geometrical explanation of the family structure dissolves thereby, it is clear that whatever the mixing, in the end exactly three SM families will be left if the mass matrices $\mathcal{M}_{x\bar{x}}$ have maximal rank.

Decoupling of exotics in SUSY preserving vacua has additional effects on the low-energy theory. Generally, there are supersymmetric vacua in which all or most of the SM singlets get large VEVs. As a consequence, many of the (abelian and nonabelian) gauge group factors get spontaneously broken, such that the low-energy gauge group can be G_{SM} up to a hidden sector:

$$G_{SM} \times G_{\text{hidden}}$$
, (4.11)

where the SM matter is neutral under G_{hidden} . This (true) hidden sector is important to deal with the problem of low-energy SUSY breaking, as we shall outline in section 5.3.

In practice, to show that the decoupling of exotics is consistent with string selection rules is a technically involved and time consuming issue. In order to simplify the task and to reduce the number of models, we first impose an additional condition. We require that the models possess a renormalizable top—Yukawa coupling as motivated by phenomenology (namely, by the large mass of the top—quark). We point out that this requirement, nevertheless, is not imperative since a large top—Yukawa coupling can also be obtained as a result of a conspiracy of the VEVs of the fields or by other means.

As a second technical simplification, we consider only superpotential couplings up to order eight. Thus, the next two steps in our selection procedure are:

- ② Select models with a "heavy top".
- ® Select models in which the exotics decouple at order 8.

In step \mathcal{O} , we require a renormalizable $\mathcal{O}(1)$ Yukawa coupling of the type⁶

$$q \bar{u} h_u \sim (\mathbf{3}, \mathbf{2})_{1/6} (\overline{\mathbf{3}}, \mathbf{1})_{-2/3} (\mathbf{1}, \mathbf{2})_{1/2}.$$
 (4.12)

To accomplish this condition, we have first to verify whether such a coupling is allowed by string theory. In the \mathbb{Z}_6 -II orbifold, combining the point group selection rule, eq. (4.9a), with R-charge conservation, one finds that only renormalizable couplings of fields from the sectors

$$U_1 U_2 U_3$$
, $U_3 T_2 T_4$, $U_2 T_3 T_3$, $T_1 T_2 T_3$, $T_1 T_1 T_4$ (4.13)

are nonvanishing. On the other hand, the coupling strength of string interactions is given by their correlation function, which for renormalizable couplings has been computed analytically [114, 115]. It turns out that the coupling strength has a dependence on the localization of the fields involved. In particular, the coupling for twisted fields is proportional to e^{-A} , where A denotes the area of the triangle formed by the fixed points where the strings are located; hence, it can be highly suppressed if the strings are located far

⁶Notice that at this level it is impossible to distinguish between lepton doublets $\bar{\ell}$ and up-type Higgses h_u .

	V^{SO}	(10),1	V^{SO}	(10),2	$V^{{ m E}_6,1}$	$V^{\mathrm{E}_6,2}$
Criterion	7	Ø	7	X	7	7
© spectrum = 3 generations + vectorlike	72	56	37	53	3	2
® exotics decouple at order 8	56	50	32	53	3	2

Table 4.3: A subset of the MSSM candidates. The number of models with "heavy top" are listed under \mathcal{O} . \boxtimes denotes no "heavy top".

apart from each other. Therefore, couplings including an untwisted field are unsuppressed whereas TTT-couplings are significant only when the twisted fields are localized at the same fixed point. We discard models in which the above couplings vanish or are suppressed.

In the next step @, we select models in which the mass matrices $\mathcal{M}_{x\bar{x}}$ for the exotics (cf. eq. (4.10)) have a maximal rank such that no exotic states appear at low energies. Here, we consider only superpotential couplings up to order eight and for this analysis we assume that all SM singlets can obtain supersymmetric VEVs.

In table 5.2 we summarize our results (see ref. [123] for further details). We make a distinction between those having a "heavy top", as in step \mathcal{O} , and those without it, \boxtimes . We identify 93 models that pass requirements \mathcal{O} and \otimes . This means that a significant fraction of our models can serve as an ultraviolet completion of the MSSM in string theory.

There are also many models which do not fulfill criterion \mathcal{D} . As already mentioned, although our naïve approach indicates that no heavy top can be found directly in these models, it can be that further analysis reveals that such conclusion is incorrect. For that reason, we have analyzed also the decoupling of the exotics in these MSSM candidates. In this set, exotic particles can get large masses in 103 models. All in all, we see that if the criterion \mathcal{D} is skipped, out of 223 MSSM candidates, 196 do not present exotics at low energies.

A word on possible light exotics is in order. In ref. [134] it has been considered the possible presence of a special kind of vectorlike exotics (baptized exotica) in the low–energy spectrum. These exotic particles, which appear in some of our models, would not affect gauge coupling unification and could have a fractional electric charge. Although unlikely, one wonders whether such particles might appear in future colliders. If it happens, we would have to reconsider not to decouple all exotics.

We would like to point out an observation. We find that there are in our models several pairs of Higgs doublets with a matrix of μ -like mass terms. Generally, all Higgs doublets can acquire large masses just as the (other) vectorlike exotics do. The challenge is then to decouple all Higgs-doublet pairs minus one, which will trigger the Higgs mechanism of the SM model. To get only one pair of massless Higgs doublets usually requires fine-tuning in the VEVs of the SM singlets such that the mass matrix for the $(1,2)_{-1/2}$, $(1,2)_{1/2}$ states gets a zero eigenvalue. This is the notorious supersymmetric μ -problem. The fine-tuning can be ameliorated if the vacuum respects certain (approximate) symmetries [135,136]. We will find in section 5.4 that such symmetries might appear regularly in promising orbifold

constructions.

In conclusion, we have found a very interesting set of semi-realistic models. In the most optimistic case, we find that any of 196 models could very well house the physics of the SM. However, we are still far of making any phenomenological conjecture from string theory about physics at low energies. To verify whether our MSSM candidates are consistent with phenomenology requires addressing several questions. Some important issues we must still clarify include

- SUSY preserving vacua at high energies,
- hierarchically small SUSY breaking,
- realistic flavor structures, and
- absence of fast proton decay.

These and some other phenomenological questions will be discussed in the next chapter.

Chapter 5

Low Energy Physics from Heterotic Orbifolds

In this chapter we examine the phenomenological properties of the models found in the previous chapter. First, we study the details of the matter spectra of a characteristic model with the exact spectrum of the MSSM, which serves as example all through this chapter. Then we proceed to evaluate the implications of the hidden sectors of the MSSM-like models. We investigate a correlation between the realistic properties of these models and SUSY breaking via gaugino condensation. A surprising result is that the breakdown of SUSY occurs generically at an intermediate energy scale. Another interesting problem is whether it is possible to impose R-parity in order to guarantee proton stability. We find a chance, defining a matter parity from the spontaneous breaking of $U(1)_{B-L}$ which however can only be defined in some models. Other plausible possibilities are not discussed here. Finally, we address the question of the viability of the seesaw mechanism in orbifold models. Contrary to previous statements, we find that the presence of $\mathcal{O}(100)$ right-handed neutrinos in MSSM-like models supports the success of the seesaw mechanism in these models.

5.1 An Orbifold-MSSM: Part I

In the last chapter we found more than 200 models with the MSSM matter spectrum and vectorlike exotics. By demanding that the vectorlike exotics decouple from the low–energy spectrum, we have seen that the number of models is barely reduced. We then opted to take a more aggressive approach: we disregarded those models where a trilinear Yukawa coupling $q \bar{u} h_u$ is vanishing or suppressed according to string selection rules. Note that this requirement is not imperative and was implemented only as a technical simplification. It turns out that this criterion is rather strict and reduces the amount of realistic models by a factor 1/2. The remaining $\mathcal{O}(100)$ models have many common properties and will lead us to interesting phenomenological conclusions along the lines of this chapter.

We study now one generic model out of the set of promising MSSM candidates with "heavy top", which will be referred to as *orbifold*-MSSM. The model is based on the gauge

#	Representation	Label	#	(Anti-)Repr.	Label
3	$(3,2;1,1)_{1/6}$	q_i			
8	$({f 1},{f 2};{f 1},{f 1})_{-1/2}$	ℓ_i	5	$({f 1},{f 2};{f 1},{f 1})_{1/2}$	$ar{\ell}_i$
3	$({f 1},{f 1};{f 1},{f 1})_1$	$ar{e}_i$			
	$({f \overline{3}},{f 1};{f 1},{f 1})_{-2/3}$	\bar{u}_i			
7	$({f \overline{3}},{f 1};{f 1},{f 1})_{1/3}$	$ar{d}_i$	4	$(3,1;1,1)_{-1/3}$	d_i
4	$(3,1;1,1)_{1/6}$	v_i	4	$({f \overline{3}},{f 1};{f 1},{f 1})_{-1/6}$	\bar{v}_i
20	$({f 1},{f 1};{f 1},{f 1})_{1/2}$	s_i^+	20	$({f 1},{f 1};{f 1},{f 1})_{-1/2}$	s_i^-
2	$({f 1},{f 1};{f 1},{f 2})_{1/2}$	x_i^+	2	$({f 1},{f 1};{f 1},{f 2})_{-1/2}$	x_i^-

#	Representation	Label
4	$({f 1},{f 2};{f 1},{f 1})_0$	m_i
2	$({f 1},{f 2};{f 1},{f 2})_0$	y_i
47	$({f 1},{f 1};{f 1},{f 1})_0$	s_i^0
26	$({f 1},{f 1};{f 1},{f 2})_0$	h_i
9	$({f 1},{f 1};{f 8},{f 1})_0$	w_i

Table 5.1: Massless spectrum. The quantum numbers are shown with respect to $SU(3)_c \times SU(2)_L \times SO(8) \times SU(2)$, the hypercharge is given by the subscript.

shift

$$V^{\text{SO}(10),1} = \left(\frac{1}{3}, -\frac{1}{2}, -\frac{1}{2}, 0, 0, 0, 0, 0\right) \left(\frac{1}{2}, -\frac{1}{6}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, \frac{1}{2}\right) . \tag{5.1}$$

where we have added an $E_8 \times E_8$ lattice vector to simplify computations. The Wilson lines are chosen as

$$A_{3} = \left(-\frac{1}{2}, -\frac{1}{2}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}\right) \left(\frac{10}{3}, 0, -6, -\frac{7}{3}, -\frac{4}{3}, -5, -3, 3\right),$$

$$A_{5} = \left(\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, \frac{1}{4}, \frac{1}{4}, \frac{1}{4}, \frac{1}{4}\right) \left(1, -1, -\frac{5}{2}, -\frac{3}{2}, -\frac{1}{2}, -\frac{5}{2}, -\frac{3}{2}, \frac{3}{2}\right). \tag{5.2}$$

The hypercharge generator is identified with the SU(5) standard form

$$\mathbf{t}_{Y} = \left(0, 0, 0, -\frac{1}{2}, -\frac{1}{2}, \frac{1}{3}, \frac{1}{3}, \frac{1}{3}\right) \left(0, 0, 0, 0, 0, 0, 0, 0\right). \tag{5.3}$$

The gauge group after compactification is

$$G_{SM} \times SO(8) \times SU(2) \times U(1)^7$$
, (5.4)

while the massless spectrum is given in table 5.1. From there, we see that the three SM particles ℓ_i and \bar{d}_i come accompanied with additional vectorlike pairs $\ell - \bar{\ell}$ and $\bar{d} - d$, respectively. Exotics with nonzero hypercharge, such as x_i^+ and x_i^- , are also vectorlike. Since the exotics with no hypercharge m_i and y_i are doublets under $\mathrm{SU}(2)_L$, they can allow couplings such as $m_i m_j$ and decouple from the low–energy spectrum. The remaining particles are either SM singlets or the three SM matter generations which need not be decoupled.

Note that Higgs doublets and lepton doublets cannot be distinguished in a supersymmetric theory, therefore both classes of particles are denoted by ℓ . That implies particularly that the down-type Higgs h_d is given by a linear combination of some ℓ_i , whereas the uptype Higgs h_u (required in supersymmetric theories) is a superposition of $\bar{\ell}_i$.

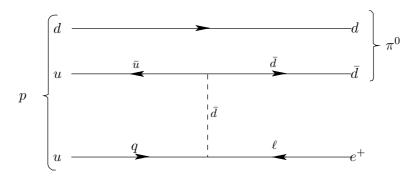


Figure 5.1: Dangerous contribution to proton decay through exchange of the scalar component of \bar{d} due to dimension four operators of the type $q_i \ell_j \bar{d}_k$ and $\bar{u}_i \bar{d}_j \bar{d}_k$. The proton decays rapidly as $p \to \pi^0 e^+$.

5.1.1 Renormalizable Couplings, "Heavy Top" and Proton Decay

At trilinear level, string selection rules (see section 4.4.1) allow only few couplings of SM particles. There is one coupling of the type $q_i \bar{\ell}_j \bar{u}_k$ with only untwisted fields, which we consider to be the Yukawa coupling that gives mass to the top-quark. Notice that this coupling allows us to identify the right-handed top, the up-type Higgs doublet h_u and the quark doublet of the third generation.

There are also four couplings of the type $q_i \ell_j \bar{d}_k$ and four of the type $\bar{e}_i \ell_j \ell_k$. They can produce the down-type quark and lepton masses as well as lepton number violating interactions. Clearly, depending on the specifics of the vacuum configuration and, particularly, on the choice of the down-type Higgs h_d , it can be that lepton violating operators do not appear in this model.

Further, we note that lepton number violating interactions $q_i \, \ell_j \, d_k$ are harmless to the stability of the proton as long as they do not come accompanied by quark interactions, $\bar{u}_i \bar{d}_j \bar{d}_k$ (see fig. 5.1). Thus, in this model, due to the absence of $\bar{u}_i \bar{d}_j \bar{d}_k$ operators, the proton is stable at trilinear level, so that dangerous effective dimension four operators can appear only suppressed by different powers of $M_{\rm str}$.

5.1.2 Spontaneous Symmetry Breaking and Decoupling of Exotics

Let us suppose that all SM singlets s_i^0 develop nonzero VEVs. Since they are only charged under some hidden U(1)'s, the gauge group gets spontaneously broken to

$$G_{SM} \times G_{\text{hidden}}$$
, (5.5)

where $G_{\text{hidden}} = \text{SO}(8) \times \text{SU}(2)$. Note that if more than one fields h_i were allowed to develop VEVs, the unbroken hidden gauge group would certainly be $G_{\text{hidden}} = \text{SO}(8)$.

Additionally, the vectorlike states get large masses. In order to verify this, one has to obtain the mass matrices $\mathcal{M}_{x\bar{x}}$ of the exotics and assume that the VEVs of the singlets s_i^0 are of the same order $\langle s_i^0 \rangle \approx s \approx M_{\rm str}$. Replacing the singlets by their VEVs, one can then compute the rank of the effective mass matrices. A mass matrix with maximal rank ensures

that all exotics of the corresponding class obtain masses proportional to the VEVs s. We make at this point a strong assumption, namely that all singlets acquire VEVs consistent with supersymmetry, that is, along D- and F-flat directions. This assumption will be confirmed later.

We have checked that the rank of all the mass matrices is maximal, such that the exotics do decouple from the effective low-energy theory. Below we present most of the matrices. There, s^n indicates the dominant mass term for each entry, i.e. s^n denotes that the corresponding coupling appears first when n singlets are involved. Each entry usually contains many terms and involves different singlets as well as coupling strengths, which are presumed to be of order 1 in string units.

$$\mathcal{M}_{mm} = \begin{pmatrix} 0 & s^{5} & s^{6} & s^{6} \\ s^{5} & 0 & s^{6} & s^{6} \\ s^{6} & s^{6} & 0 & s^{5} \\ s^{6} & s^{6} & s^{5} & 0 \end{pmatrix}, \qquad \mathcal{M}_{d\bar{d}} = \begin{pmatrix} s^{6} & s^{6} & s^{3} & s^{6} & s^{6} & s^{1} & s^{1} \\ s^{6} & s^{6} & s^{3} & s^{6} & s^{6} & s^{1} & s^{1} \\ s^{6} & s^{6} & s^{3} & s^{6} & s^{6} & s^{1} & s^{1} \\ s^{3} & 0 & 0 & s^{3} & 0 & s^{6} & s^{6} \\ s^{6} & s^{3} & 0 & s^{6} & s^{3} & s^{6} & s^{6} \end{pmatrix},$$

$$\mathcal{M}_{v\bar{v}} = \begin{pmatrix} s^{5} & s^{5} & s^{5} & s^{5} \\ s^{5} & s^{5} & s^{5} & s^{5} \\ s^{6} & s^{6} & s^{1} & s^{5} \\ s^{6} & s^{6} & s^{5} & s^{1} \end{pmatrix}, \qquad \mathcal{M}_{\ell\bar{\ell}} = \begin{pmatrix} s^{3} & s^{1} & s^{1} & s^{1} & s^{1} \\ s^{1} & s^{3} & s^{3} & s^{3} & s^{3} \\ s^{1} & s^{3} & s^{3} & s^{3} & s^{3} \\ s^{1} & s^{3} & s^{3} & s^{3} & s^{3} \\ s^{1} & s^{3} & s^{3} & s^{3} & s^{3} \\ s^{1} & s^{3} & s^{3} & s^{3} & s^{3} \\ s^{1} & s^{3} & s^{6} & s^{6} & s^{2} \\ s^{4} & s^{2} & s^{6} & s^{2} & s^{2} \end{pmatrix}.$$

Similarly, the 20×20 mass matrix $\mathcal{M}_{s^+s^-}$ has also maximal rank. The $d\bar{d}$ mass matrix is 4×7 such that there are three massless \bar{d} states. The $\ell\bar{\ell}$ mass matrix is 8×5 , so there are effectively three massless lepton doublets.

As we already mentioned, without additional information, it is not possible to distinguish between lepton doublets and Higgs doublets. Therefore, since the mass matrix $\mathcal{M}_{\ell\bar\ell}$ has maximal rank, all possible Higgs pairs are massive. If one wishes to recover the Higgs mechanism, it is necessary to choose a special vacuum configuration where the rank of the matrix $\mathcal{M}_{\ell\bar\ell}$ is reduced. This corresponds precisely to the supersymmetric μ -problem, that in a first approach is not automatically solved. Note that, since the matrix $\mathcal{M}_{\ell\bar\ell}$ is not diagonal per se, the resulting Higgs doublets will be indeed linear combinations of the fields ℓ_i and $\bar\ell_i$. Unfortunately, this spoils our argument about the identification of the Higgs fields through the trilinear coupling of the top quark.

Summing up all the ingredients just described, we end up with the exact MSSM spectrum.

Let us stress that these properties, although particular to this orbifold-MSSM, are quite similar to those of other models. That is, all 93 MSSM candidates with "heavy top" accept this description. Perhaps one of the most relevant differences will be the number of lepton

doublets ℓ and $\bar{\ell}$ that can be interpreted as Higgs doublets. In some cases, there are only four SUSY fields ℓ_i and one $\bar{\ell}_i$, which may be seen as three lepton doublets and one pair of Higgses. This can lead to considering a particular model more predictive than another.

Let us stop our discussion on this model here and proceed to explain how to guarantee SUSY at high energies in these models. This model will appear regularly in the following sections in order to exemplify some of the new elements introduced.

5.2 Supersymmetric Vacua

In the previous section, we have assumed that all SM singlets can acquire nonvanishing VEVs without destroying some of the properties of our orbifold-MSSM. However, this is in general not true. In this section we discuss the constraints to get supersymmetric vacuum configurations, their consequences and a method to obtain such vacua.

Momentarily, let us assume only global supersymmetry. We would like to verify whether the MSSM candidates admit supersymmetric vacuum configurations. This amounts to inspecting whether a combination of fields can attain VEVs, such that the D- and F-terms vanish. This is explained as follows. The scalar potential in a supersymmetric gauge theory is given, in terms of the auxiliary fields D_a and F_i of the gauge and chiral multiplets by

$$V(\phi_i, \phi_i^{\dagger}) = \frac{1}{2} \sum_a D_a(\phi_i, \phi_i^{\dagger})^2 + \sum_i |F_i|^2,$$
 (5.6)

where the auxiliary fields are expressed in terms of the fields ϕ_i, ϕ_i^* and the gauge group generators t_a in the representation of ϕ_i . The auxiliary fields are given by

$$D_a = \sum_i \phi_i^{\dagger} \mathsf{t}_a \phi_i \,, \tag{5.7}$$

$$F_i = \frac{\partial W}{\partial \phi_i} \tag{5.8}$$

with $W = W(\phi_i)$ being the superpotential. From a field theoretical perspective, solutions of eqs. (5.7) and (5.8) represent supersymmetric vacua. However, string theory introduces a new element, namely an anomalous U(1) symmetry, U(1)_A (such that $\operatorname{tr} \mathbf{t}_A \neq 0$). In presence of U(1)_A, the corresponding D-term analogous to eq. (5.7) gets an additional contribution¹ ξ [137], giving rise to the so-called FI D-term

$$D_A = \sum_i \phi_i^{\dagger} t_A \phi_i + \xi \equiv \sum_i q_i^A |\phi_i|^2 + \frac{g M_{\text{Pl}}^2}{192\pi^2} \operatorname{tr} t_A, \qquad (5.9)$$

where t_A is the generator of $U(1)_A$ and g is the four-dimensional coupling constant, given by the VEV of the four-dimensional dilaton φ as $g = e^{\langle \varphi \rangle}$. In our conventions, $\operatorname{tr} t_A > 0$

¹To maximize confusion, normally ξ is called FI term. Note that D_A is called here FI D-term.

and thus $\xi > 0$. Note that the 'traditional' vacuum $\langle \phi_i \rangle = 0$ is not supersymmetric as (for $g \neq 0$) the extra term ξ induces SUSY breakdown at a scale too large to be realistic.

Supersymmetry remains unbroken if and only if the scalar potential admits a minimum $\phi_i = \langle \phi_i \rangle$ such that $V(\langle \phi_i \rangle, \langle \phi_i^{\dagger} \rangle) = 0$. Therefore, any supersymmetric ground state has to satisfy the D- and F-flatness conditions

$$\langle D_a \rangle = 0 \quad \forall a,$$
 (5.10a)
 $\langle F_i \rangle = 0 \quad \forall i.$ (5.10b)

$$\langle F_i \rangle = 0 \qquad \forall i. \tag{5.10b}$$

Note that there are as many F_i -constraints as fields ϕ_i .

Provided a set of fields ϕ_i , supersymmetric field configurations are given by the sets of VEVs $\langle \phi_i \rangle$ satisfying eqs. (5.10). Naïvely, it appears that the number of constraints is larger than the number of variables, so that the system seems to be overconstrained. However, this is not the case. As well known, complexified gauge transformations along (F-) flat directions allow us to eliminate the D-term constraints², such that the number of variables equals the number of equations.

5.2.1D–Flatness

Let us review first the issue of D-flatness and cancellation of the FI term [104,133,139–141]. Since, in particular, D_A must vanish in a supersymmetric vacuum, at least some of the scalars are forced to attain large VEVs, typically not far below the string scale. This sets the scale of the breaking of $U(1)_A$ and other gauge symmetries, under which the scalars might be charged. In other words, vanishing of the FI D-term triggers spontaneous symmetry breaking at a very high scale. On the other hand, the condition $D_A = 0$ also implies that there must be at least one field whose anomalous charge q^A is (in our convention) negative (i.e. opposite in sign to trt_A).

D-flatness can be achieved by the noticeable observation that, in general, every holomorphic gauge invariant monomial (HIM) represents a D-flat direction [133]. Particularly, in theories without an anomalous U(1), the condition (5.10a) is satisfied if there exists a monomial $I(\phi_i)$. Then, the VEVs of the scalar fields in a D-flat vacuum configuration are read off from

$$\left\langle \frac{\partial I}{\partial \phi_i} \right\rangle = c \langle \phi_i^{\dagger} \rangle, \tag{5.11}$$

where c is a complex dimensional constant, $c \neq 0$. This remarkable result follows simply from the gauge invariance of $I(\phi_i)$ [133]. In the case that $U(1)_A$ is present, the monomial $I(\phi_i)$ must be invariant with respect to all gauge symmetries except $U(1)_A$. The anomalous charge of this monomial must be negative [141]. Let us explain this in more detail and propose a method to find such HIMs.

²For a field-theoretic discussion, see e.g. ref. [138, p.57-58]. Details on heterotic orbifolds can be found in ref. [48]

In supersymmetric theories with a single U(1) gauge theory, the so-called D-term potential is given by

$$V_D \propto \left[\sum_i q_i |\phi_i|^2\right]^2 . \tag{5.12}$$

Consider as a first example a U(1) gauge theory with two fields ϕ_{\pm} carrying the charges ± 1 . Since $V_D^{1/2} \propto |\phi_+|^2 - |\phi_-|^2$, V_D vanishes at the ground state $|\langle \phi_+ \rangle| = |\langle \phi_- \rangle|$. That is, one has a D-flat direction, parametrized by $c = |\langle \phi_+ \rangle| = |\langle \phi_- \rangle|$. Note that, if one writes the HIM $I(\phi_{\pm}) = \phi_+ \phi_-$, which is the only possible holomorphic and invariant monomial, eq. (5.11) yields the same result.

Consider now a theory with one field (ϕ_1) with charge 2 and two fields (ϕ_2, ϕ_3) with charges -1. Then we have many flat directions, described by the roots of the equation $2|\langle \phi_1 \rangle|^2 - |\langle \phi_2 \rangle|^2 - |\langle \phi_3 \rangle|^2 = 0$. The solutions are the three directions

$$|\langle \phi_1 \rangle| = |\langle \phi_2 \rangle| = |\langle \phi_3 \rangle| = c; \tag{5.13a}$$

$$|\langle \phi_2 \rangle| = 0, \ |\langle \phi_3 \rangle| = \sqrt{2} |\langle \phi_1 \rangle| = c'; \tag{5.13b}$$

$$|\langle \phi_3 \rangle| = 0, \ |\langle \phi_2 \rangle| = \sqrt{2} |\langle \phi_1 \rangle| = c'' \tag{5.13c}$$

with complex parameters c, c' and c''. It is convenient to associate these directions to the HIMs

$$\phi_1 \, \phi_2 \, \phi_3 \, , \quad \phi_1 \, \phi_3^2 \, , \quad \phi_1 \, \phi_2^2 \, , \tag{5.14}$$

respectively.

From this example, we notice that a monomial $I(\phi_i) = \phi_1^{n_1} \phi_2^{n_2} \cdots \phi_k^{n_k}$ represents a flat direction defined by the relation

$$\frac{|\langle \phi_1 \rangle|}{\sqrt{n_1}} = \frac{|\langle \phi_2 \rangle|}{\sqrt{n_2}} = \dots = \frac{|\langle \phi_k \rangle|}{\sqrt{n_k}} \quad \text{and} \quad |\langle \phi_j \rangle| = 0 \text{ for } n_j = 0.$$
 (5.15)

It is, however, clear that there is only a finite number of linearly independent D-flat directions. In the previous example, the third direction is not independent of the other two. In other words, the requirement $V_D=0$ poses only one constraint on the three real variables $(|\langle \phi_i \rangle|^2)$ entering eq. (5.12). The space of absolute values $|\langle \phi_i \rangle|$ is two-dimensional. The power of using the monomials is that checking whether certain monomials are linearly independent or not is fairly simple. One identifies with each monomial the vector of exponents, $n=(n_1,n_2\ldots)$. The flat directions are independent if and only if the vectors are linearly independent. In the previous example one would get the vectors (1,1,1), (1,0,2), and (1,2,0), out of which only two are linearly independent.

It is also clear how to obtain these vectors: all of them are orthogonal to the vector of charges $q = (q_1, q_2, ...)$. That is, the problem of finding the above monomials (and thus the D-flat directions) is reduced to the problem of finding vectors n with the following properties:

1.
$$q \cdot n = 0$$

$$2. n_i \in \mathbb{N}_0.$$

The property that the n_i be integer-valued does not pose a constraint in our models: since the charges are rational, one can rescale any n having the first property such as to have integer entries. However, the requirement that the entries be nonnegative, which reflects that the monomials ought to be holomorphic, is a constraint.

The discussion so far can easily be extended to $U(1)^m$ theories. Here the D-term potential is

$$V_D \propto \sum_{j=1}^m \left[\sum_i q_i^{(j)} |\phi_i|^2 \right]^2 ,$$
 (5.16)

where $q_i^{(j)}$ is the charge of the field ϕ_i under the j^{th} U(1) factor. Now a D-flat direction has to satisfy the above constraints for each U(1) factor separately. Again, it is advantageous to represent D-flat directions by holomorphic gauge invariant monomials. Then the vector n of exponents has to be orthogonal to every charge vector $q^{(j)} = (q_1^{(j)}, q_2^{(j)}, \dots)$. In other words, n has to be in the kernel of the charge matrix Q,

$$Q \cdot n = 0, \quad \text{with} \quad Q = \begin{pmatrix} q_1^{(1)} & q_2^{(1)} & \dots \\ q_1^{(2)} & q_2^{(2)} & \dots \\ \vdots & \vdots & \vdots \\ q_1^{(n)} & q_2^{(n)} & \dots \end{pmatrix} . \tag{5.17}$$

Hence, the problem of finding the D-flat directions of a $U(1)^m$ gauge theory is reduced to the task of calculating the kernel of the charge matrix Q, and to forming linear combinations of elements of this kernel in such a way that the entries are nonnegative integers. The maximal linear independent set of such linear combinations is in one-to-one correspondence with the independent D-flat directions. It is straightforward to see that the results obtained so far generalize to the nonabelian case [133].

We can now review the issue of cancelling the FI term. For an anomalous U(1), the D-term potential V_D follows from the FI D-term, eq. (5.9),

$$V_D^A \propto \left[\sum q_i^A |\phi_i|^2 + \xi\right]^2 . \tag{5.18}$$

Recalling that $\xi > 0$ in our conventions, to cancel the FI term one thus has to find a holomorphic monomial,

$$I = \phi_1^{n_1} s_2^{n_2} \dots {5.19}$$

with net negative charge under $U(1)_A$, i.e.

$$\sum_{i} n_i \, q_i^A \ < \ 0 \ . \tag{5.20}$$

To summarize, the D-flat directions are in one-to-one correspondence with holomorphic gauge invariant monomials. In the abelian case, such monomials can be identified with elements of the kernel of the charge matrix Q with nonnegative integer entries. Cancellation of the FI term requires the existence of a holomorphic monomial with net negative charge under $U(1)_A$, which is gauge invariant with respect to all other group factors.

5.2.2 F-Flatness

Let us now turn to the discussion of F-flatness. Since the superpotential W is nonrenormalizable, studying this question in detail is somewhat less general than D-flatness because most of the statements one can obtain are order-dependent.

However, ensuring that a vacuum configuration is F-flat (ignoring momentarily D-flatness) is mostly trivial because the number of equations

$$\langle F_i \rangle = \left\langle \frac{\partial W}{\partial \phi_i} \right\rangle = 0 \qquad \forall i$$
 (5.21)

coincides with the number of variables and, in general, when the superpotential is a non-trivial polynomial, some of their solutions are nontrivial.

In particular, nontrivial solutions can always be found in orbifold compactifications. Let us illustrate it in the case of \mathbb{Z}_N orbifolds. If the superpotential W_0 at order x in the fields ϕ_i is allowed by string selection rules, then an 'extended' superpotential $W \sim W_0 + W_0^{N+1} + W_0^{2N+1} + \dots$ is also admissible.

As an example, assume a \mathbb{Z}_3 orbifold toy-model with two particles, ϕ_1 and ϕ_2 . Suppose further that the superpotential $W_0 = \phi_1 \phi_2^2$ is allowed. Clearly, the solution to eq. (5.21) is given by $\langle \phi_2 \rangle = 0$ and $\langle \phi_1 \rangle$ arbitrary. Suppose now that $\langle \phi_2 \rangle = 0$ is unwanted for some reason. An allowed extension of W_0 would be given by $W \sim W_0 + W_0^4 = \phi_1 \phi_2^2 + \phi_1^4 \phi_2^8$. In this scenario, the previous solution is still a valid option, but there is also an additional nontrivial F-flat vacuum configuration, parametrized by $\langle \phi_2^6 \rangle = -1/4 \langle \phi_1^3 \rangle$ and arbitrary $\langle \phi_1^3 \rangle \neq 0$. Notice that $\langle \phi_2 \rangle$ in this vacuum configuration is nontrivial. This discussion makes manifest the order-dependence of F-flatness. In this sense, verifying F-flatness at a given order in the superpotential is physically not so relevant in orbifold models.

After arriving to an F-flat vacuum configuration, the natural question is whether recovering D-flatness imposes additional conditions and, therefore, overconstrains the choice of the vacuum parameters. It has been shown [48] that in orbifold models, given a solution to the F-term equations (5.21), complexified gauge transformations scale this solution to give a family of solutions. Remarkably, particular rescalings succeed in rendering all the D-terms (including the FI D-term) zero.

$$\langle W \rangle = 0$$

Up to now, our considerations are valid only for globally supersymmetric models at perturbative level. An admissible vacuum configuration arising from string theory, however, must also be consistent with local SUSY. In fact, supergravity as well as nonperturbative effects can modify the properties of a particular vacuum configuration.

In an attempt to deal with those additional effects, the condition $\langle W \rangle = 0$ is commonly imposed in the same footing as F-flatness (see e.g. refs. [49,141]). Strictly speaking, this condition has nothing to do with preservation of SUSY. Since in supergravity the gravitino

³Note that we omit all coefficients.

mass is given by

$$m_{3/2} \propto |\langle W \rangle|^2$$
, (5.22)

 $\langle W \rangle = 0$ implies $m_{3/2} = 0$ and therefore, through the relation $\Lambda = -3 \, M_{\rm Pl}^2 \, m_{3/2}^2$, a vanishing cosmological constant Λ . Recall that in the context of supergravity, a nonvanishing gravitino mass does not necessarily imply breakdown of supergravity. That is, imposing the condition $\langle W \rangle = 0$ amounts to requiring the existence of a Minkowski vacuum (similar to the vacuum we are living in) rather than unbroken supersymmetry.

A caveat is in order. In order to compute e.g. the gravitino mass, eq. (5.22), and thus to discriminate a vacuum according to its nature (Minkowski, De Sitter or anti-De Sitter), it is fundamental to consider the complete superpotential W, including, in particular, contributions due to nonperturbative effects. However, the superpotential considered in previous sections does not include those contributions because, at this level, they are hardly controllable in orbifold constructions. Therefore, requiring $\langle W \rangle$ to vanish does not affect directly the VEVs of the fields ϕ_i , contrary to the D- and F-flatness conditions that we discussed above.

5.2.3 A Susy Vacuum of the Orbifold-Mssm

We are now in position to verify whether the orbifold-MSSM introduced in section 5.1 possesses a supersymmetric vacuum. By following the method described above, we find the HIM

$$I = s_1^0 s_2^0 (s_3^0)^3 (s_5^0)^3 (s_8^0) (s_{22}^0) (s_{35}^0)^2 (s_{41}^0)^3 (s_{43}^0)^4 (s_{46}^0)^3 h_2^4 h_3 h_5^5 h_9^2 h_{13}^2 h_{14}^2 h_{20} h_{21}^3 h_{22}^6$$
 (5.23)

with net anomalous charge $\sum_i q_i^A = -52/3$. We further identify that some fields share the same gauge quantum numbers:

$$s_5^0 \leftrightarrow s_{12}^0 \leftrightarrow s_9^0 \leftrightarrow s_{16}^0$$
, $s_8^0 \leftrightarrow s_{15}^0$, $s_{22}^0 \leftrightarrow s_{24}^0$. (5.24)

Therefore we can consider a vacuum configuration where the SM singlets⁴

$$\{\widetilde{s}_i\} = \{s_1^0, s_2^0, s_3^0, s_5^0, s_8^0, s_9^0, s_{12}^0, s_{15}^0, s_{16}^0, s_{22}^0, s_{24}^0, s_{35}^0, s_{41}^0, s_{43}^0, s_{46}^0, h_2, h_3, h_5, h_9, h_{13}, h_{14}, h_{20}, h_{21}, h_{22}\}$$

$$(5.25)$$

develop nonzero VEVs while the expectation values of all other fields vanish. Fields s_i^0 are singlets under all nonabelian gauge factors and hypercharge, but carry (hidden) U(1) charges; fields h_i are doublets under the hidden SU(2) and charged under the U(1)'s excepting hypercharge (cf. tables 5.1 and F.2).

Let us assume that all particles attain (almost) the same VEV $\langle \tilde{s}_i \rangle \approx s$. From eq. (5.9) with tr $t_A = 170/3$ for the orbifold MSSM, we see that the FI *D*-term is given by

$$\langle D_A \rangle \approx -\frac{52}{3}s^2 + \frac{g\,M_{\rm Pl}^2}{192\pi^2} \frac{170}{3} \,.$$
 (5.26)

⁴In table F.2 the two SM singlets $s_{1,2}^0$ have been denoted by $\chi_{1,2}$. The reason will become transparent when we introduce B-L in section 5.4.

Therefore, in order for the chosen vacuum to be D-flat, the expectation values of the SM singlets should be $s \approx \sqrt{g} M_{\rm Pl} \times 10^{-2} \approx ({\rm few}) \times 10^{16} {\rm GeV}$.

In the vacuum configuration (5.25), the gauge symmetry group \mathcal{G}_{4D} given in eq. (5.4) breaks spontaneously, as expected, to

$$G_{SM} \times G_{\text{hidden}}$$
, (5.27)

where $G_{\text{hidden}} = \text{SO}(8)$. Notice that the rank of G_{hidden} is four. This will turn out to be very general in models with realistic features and will lead in the next section to the conjecture that SUSY breaking through gaugino condensation occurs generally at an intermediate energy scale for realistic heterotic orbifold models.

For completeness, let us address F-flatness in the orbifold-MSSM. To verify whether this vacuum is also F-flat, we would need the superpotential W up to any order in SM singlets. Realistically, we have to stop computing the superpotential at a given order. We consider here, for simplicity, the order-six superpotential

$$W = s_{32}^{0}h_{5}(s_{5}^{0}h_{1} + s_{12}^{0}h_{2}) + (s_{15}^{0}h_{15} + s_{8}^{0}h_{13})(s_{42}^{0} + s_{43}^{0})(h_{23} + h_{25}) + (s_{22}^{0}h_{14} + s_{24}^{0}h_{16})(s_{42}^{0} + s_{43}^{0})(h_{18} + h_{20}) + h_{22}(s_{5}^{0}h_{3} + s_{12}^{0}h_{4})\left(s_{41}^{0}(s_{26}^{0} + s_{28}^{0}) + s_{32}^{0}(s_{42}^{0} + s_{43}^{0}) + s_{35}^{0}(s_{45}^{0} + s_{46}^{0})\right).$$

$$(5.28)$$

The resulting F-terms are provided in eqs. (F.26). We observe that with nonvanishing VEVs for the fields (5.25), there are still some nonzero F-terms, implying that some of those fields should have trivial VEVs. However, as we have mentioned before, one can argue that, if we go to higher orders in W, nontrivial solutions to all $F_i = 0$ equations can be found. As a matter of fact, we find that if we go to order eight in the superpotential, such nontrivial solutions exist.⁵

5.3 Supersymmetry Breakdown

In the previous section, we have found that supersymmetric vacua can be achieved in MSSM candidates (at least at perturbative-level). However, as SUSY is broken in nature, realistic models should admit spontaneous SUSY breaking at an intermediate scale. Remarkably, it is known that in most $\mathcal{N}=1$ vacua SUSY is broken spontaneously by nonperturbative effects. Our understanding of nonperturbative breaking of string theory is as yet very limited, but below the string scale we can work in the effective quantum field theory. Indeed, there is a reasonably coherent understanding of nonperturbative breaking of SUSY in field theory, and the low–energy theories emerging, in particular, from heterotic orbifolds are typically of the type in which this breaking happens. This topic is quite involved as there are several symmetry-breaking mechanisms, such as gaugino condensation, instantons or composite goldstinos.

⁵The details are naturally available from the author upon request.

In orbifold models with realistic features, there are frequently additional (nonabelian) gauge symmetries that remain unbroken even after all exotics have acquired large masses. This hidden sector⁶ G_{hidden} usually contains little or no matter at all. These are precisely the ingredients that can trigger spontaneous SUSY breaking via hidden sector gaugino condensation [57–59, 142].

Gaugino condensation occurs when one or more gauge groups in the hidden sector become strongly coupled at an intermediate scale Λ . To determine the exact scale in which it happens, we need to know the running of the coupling of the hidden sector. It is given by

$$g_{G_{\text{hidden}}}^2(\mu) \approx \frac{1}{g^{-2}(M_{\text{GUT}}) - \beta_{\text{hidden}} \ln(M_{\text{GUT}}^2/\mu^2)}, \qquad \beta_{\text{hidden}} = \frac{b_0}{16\pi^2},$$
 (5.29)

where g is the four-dimensional (string) coupling constant given by $g^{-2} = e^{-2\langle\phi\rangle} = \mathrm{Re}\langle S\rangle$, and b_0 is the well known beta-coefficient which depends on the gauge group G_{hidden} and the (hidden) matter content. The Landau pole of the hidden sector is determined by $g_{G_{\mathrm{hidden}}}^2 \to \infty$. Hence, we obtain

$$\Lambda \approx M_{\text{GUT}} \exp \left\{ -\frac{1}{2\beta} \frac{1}{g^2(M_{\text{GUT}})} \right\} , \qquad (5.30)$$

where we have omitted the label $_{\rm hidden}$ to keep the results short. Note that the scale Λ is below the string (and GUT) scale, but above the scale where any of the observable gauge groups become strong. Just as with quarks in QCD, the strong attraction causes the gauginos to condense [57],

$$\langle \lambda \bar{\lambda} \rangle \approx \Lambda^3 \,.$$
 (5.31)

As in QCD, this condensate breaks a chiral symmetry, but in the pure supersymmetric gauge theory (containing only gauge boson and gauginos) it does not break SUSY.

In string theory, the fields of the hidden sector couple to one special moduli, namely the dilaton S. In particular, the auxiliary field of S couples to the (hidden) gauginos inducing, at scales below Λ , a nonperturbative effective interaction (compare with eq. (5.30))

$$F_S \langle \lambda \bar{\lambda} \rangle \approx F_S \Lambda^3 \approx F_S M_{\text{GUT}}^3 e^{-\frac{3}{2\beta}S}$$
 (5.32)

It is usual to define the parameter $a \equiv 3/2\beta$ that, in a way, can be used to determine whether SUSY breaking occurs at a realistic energy scale (see e.g. ref. [143]).

The existence of such coupling implies an effective nonperturbative superpotential

$$W \approx M_{\text{GUT}}^3 e^{-aS}, \qquad a = \frac{3}{2\beta}. \tag{5.33}$$

This superpotential breaks supersymmetry. Despite the fact that a given model could be supersymmetric at all orders of perturbation theory, nonperturbatively, the term

$$F_S = \frac{\partial W}{\partial S} \approx -aM_{\text{GUT}}^3 e^{-aS} \tag{5.34}$$

⁶Notice that a requirement to call it "hidden sector" is that the observable matter (e.g. the MSSM particle spectrum) be uncharged under $G_{\rm hidden}$.

is nonzero and, thus, breaks SUSY. Furthermore, the F-term in eq. (5.34) leads to $S \to \infty$ at the minimum of the resulting scalar potential. This is the notorious dilaton run-away problem that appears in models with a single gaugino condensate and a classical (universal) Kähler potential

$$K = -\ln(S + \overline{S}). \tag{5.35}$$

To solve this problem, one can either employ multiple gaugino condensates or nonperturbative corrections to the Kähler potential. In orbifold models, nevertheless, the first option is generically ruled out since there is mostly just one unbroken gauge group factor in the hidden sector or the condensation scale of a possible second hidden gauge factor is too low (as is the case for e.g. SU(2) or SU(3)). Thus, we are left with the second option only. It consists in amending the classical Kähler potential for the dilaton by a nonperturbative functional form, such that

$$K = -\ln(S + \overline{S}) + \Delta K_{\rm np} . \tag{5.36}$$

The functional form of $\Delta K_{\rm np}$ has been studied in the literature [143–148]. For a favorable choice of the parameters, this correction allows one to stabilize the dilaton at a realistic value, Re $S \simeq 2$, while breaking supersymmetry [143, 147–150]. The T-moduli can be stabilized at the same time by including T-dependence in the superpotential required by T-duality [151, 152]. In simple examples, the overall T-modulus is stabilized at the self-dual point such that $F_T = 0$. This leads to dilaton dominated SUSY breaking. There are many problems attached to moduli stabilization that, in principle, can have some influence on our results. However, they are beyond the scope of this thesis.

If the dilaton is stabilized at a realistic value $\text{Re}\langle S\rangle \approx 2$, gaugino condensation translates into SUSY breaking with the gravitino mass determined by the VEV of the (hidden) gaugino condensate [58]

$$m_{3/2} \approx \frac{\langle \lambda \lambda \rangle}{M_{\rm Pl}^2} \approx \frac{\Lambda^3}{M_{\rm Pl}^2} \,.$$
 (5.37)

In particular, notice that

$$\Lambda \sim 10^{13} \,\text{GeV} \tag{5.38}$$

leads to the gravitino mass in the TeV range. It is clear that for certain (hidden) gauge groups and matter content, Λ can be in the right range.

Let us point out that, even if the scale of gaugino condensation is adequate ($\sim 10^{13}$ GeV), there are many factors that can affect it. In particular, there are string threshold corrections [153–157] which lead to different gauge couplings in the visible and hidden sectors. Whereas in the visible sector, due to gauge coupling unification, one requires $g^{-2}(M_{\rm GUT}) \approx 2$, in the hidden sector string threshold corrections can alter its value as

$$g_{\text{hidden}}^{-2}(M_{\text{GUT}}) \approx 2(1-\Delta)$$
, (5.39)

where Δ parametrizes such corrections. In this case, the running of the hidden gauge coupling changes and, hence, so do the corresponding condensation scale:

$$\Lambda \approx M_{\rm GUT} \exp\left\{-\frac{1}{2\beta}2(1-\Delta)\right\}.$$
 (5.40)

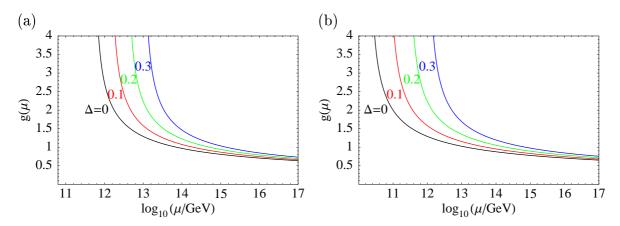


Figure 5.2: Running of the coupling constants of (a) $G_{\text{hidden}} = \text{SO}(8)$ (orbifold-MSSM) and (b) $G_{\text{hidden}} = \text{SU}(4)$ [48] for different threshold corrections Δ .

In the next section, we will have the opportunity to see how these corrections affect the breaking of supersymmetry.

5.3.1 Susy Breakdown in the Orbifold-Mssm

In section 5.2.3, we have seen that in an admissible vacuum configuration that preserves SUSY at high energies (as high as the string scale) the unbroken hidden sector is reduced to $G_{\text{hidden}} = SO(8)$. Furthermore, we observe that there are some 8-plets in this sector (cf. table 5.1). These states can be split in two sets, which form mass terms independently (see table F.1). In the vacuum configuration (5.25), their mass matrices are given by

$$\mathcal{M}_{ww} = \begin{pmatrix} \widetilde{s} & \widetilde{s}^5 & 0 & \widetilde{s}^5 & \widetilde{s}^5 \\ \widetilde{s}^5 & \widetilde{s} & 0 & \widetilde{s}^5 & \widetilde{s}^5 \\ 0 & 0 & 0 & \widetilde{s}^3 & \widetilde{s}^3 \\ \widetilde{s}^5 & \widetilde{s}^5 & \widetilde{s}^3 & \widetilde{s}^6 & \widetilde{s}^6 \\ \widetilde{s}^5 & \widetilde{s}^5 & \widetilde{s}^3 & \widetilde{s}^6 & \widetilde{s}^6 \end{pmatrix}, \qquad \mathcal{M}_{f\bar{f}} = \begin{pmatrix} 0 & \widetilde{s}^3 \\ 0 & \widetilde{s}^3 \end{pmatrix} . \tag{5.41}$$

The first of these matrices has maximal rank, but from the second matrix there is a pair f, \bar{f} that stays massless up to energies lower than M_{GUT} .

Let us proceed to compute the gravitino mass $m_{3/2}$. We will require the beta-coefficient that can be computed by

$$b_0^{SO(2N)} = 6 \cdot (N-1) - \#(2N-\text{plets});$$
 (5.42)

in our case it is $b_0 = 16$. Therefore, with the realistic value for the stabilization of the dilaton, $\text{Re}\langle S\rangle \approx 2$, we obtain from eq. (5.30)

$$\Lambda \approx M_{\rm GUT} e^{-\pi^2} \approx 10^{12} \text{ GeV} \,,$$
 (5.43)

what in turn yields the gravitino mass $m_{3/2} \approx 1$ GeV. This scale is phenomenologically unacceptable because that scale has been already ruled out by experiment. However, as we mentioned before, threshold corrections can modify this scale. In fact, by using eqs. (5.29) and (5.40), we find that threshold corrections enhance the scale of gaugino condensation and, therefore, also the gravitino mass. In fig. (5.2) we present the running of the gauge coupling $g(\mu)$ for two different cases. Fig. (5.2)(a) illustrates the influence of different values of Δ on the behavior of the gauge coupling. We notice that for $\Delta = 0$ we recover the result given in eq. (5.43) whereas for $\Delta = 0.3$ the condensation scale becomes almost 10^{14} GeV. In fig. (5.2)(b), we compare our result with a case where the hidden sector has an SU(4) gauge group and no massless matter in that sector. We might say that the situation in SO(8) is somewhat better than in SU(4). However, we must recall that there are of course other factors that can affect our estimates. For example, we have not described precisely the mechanism to stabilize the dilaton. Further, we have used the symbol '\approx' in many of our equations because those values are not precise. Therefore, in this study the most important result is that the scale of SUSY lies around the phenomenologically interesting interval.

5.3.2 Susy Breakdown in the (Mini-)Landscape

After having examined some of the most important aspects of gaugino condensation, we would like to continue our study on the set of MSSM-candidates obtained in the last chapter through a search guided by grand unification. We have found 196 orbifold models (cf. table 4.3) with the following properties:

- SM gauge group times additional gauge factors,
- nonanomalous hypercharge of the Georgi-Glashow type, i.e. consistent with gauge coupling unification,
 - three MSSM matter generations plus vectorlike exotics,
 - all vectorlike exotics are decoupled from the massless spectrum.

In order to save some computation time, we have imposed an additional constraint on the models. Namely, we have demanded one trilinear coupling of the type $q \ell \bar{u}$ which might be responsible for the heavy mass of the top quark. Although this condition is not arbitrary, it is also not imperative since a heavy top quark might also appear through alternative methods. However, for consistency, we will stick to that constraint in this section.

The strategy we have followed consists in first finding a set of MSSM candidates and then studying common features that could lead to some sort of low-energy predictions. In this section, we concentrate on the question of SUSY breaking. Particularly, we would like to figure out whether our MSSM candidates yield a reasonable scale of SUSY breaking via hidden sector gaugino condensation.

With that purpose, we will impose an additional criterion in our search

Select models where exotics decouple + gaugino condensation

At this step, we select models in which the decoupling of the SM exotic states is possible without breaking the largest gauge group in the hidden sector. We find that all or almost

Criterion	$V^{\mathrm{SO}(10),1}$	$V^{\mathrm{SO}(10),2}$	$V^{{ m E}_6,1}$	$V^{{ m E}_6,2}$
⑦ heavy top	72	37	3	2
® exotics decouple at order 8	56	32	3	2
exotics decouple + gaugino condensation	47	25	3	2

Table 5.2: A subset of the MSSM candidates.

all of the matter states charged under this group also attain large masses which allow for spontaneous SUSY breaking via gaugino condensation.

The models satisfying all our criteria, \mathbb{Q} - \mathbb{Q} , we consider the most promising MSSM candidates. Our results are presented in table 4.2. More details can be found in [158]. We find it remarkable that out of $\mathcal{O}(10^4)$ inequivalent models, $\mathcal{O}(10^2)$ pass all of our requirements, including a hidden sector that allows SUSY breaking. In this sense, the region of the heterotic landscape endowed with local SO(10) and E₆ GUTs is particularly attractive.

A comment is in order. We impose by hand the requirement that gaugino condensation be allowed. By assigning VEVs to all SM singlets, i.e. without verifying explicitely D- and F-flatness, the hidden sector gauge group is broken by matter VEVs charged under this group. Similarly, the SM gauge group is broken in more general "vacuum" configurations where many other fields can also acquire VEVs. Clearly, most of the string landscape is not relevant to our physical world. It is only possible to obtain useful predictions from the landscape once certain criteria are imposed. Here we require that gaugino condensation be allowed so that SUSY can be broken. Since the largest hidden sector group factor would dominate SUSY breaking, we focus on vacua in which this factor is preserved by matter VEVs. Within the set of our promising models, we can now study predictions for the scale of SUSY breaking.

Our MSSM candidates have the necessary ingredients for supersymmetry breaking via gaugino condensation in the hidden sector. In particular, they contain nonabelian gauge groups with no or little matter. The corresponding gauge interactions become strong at some intermediate scale which can lead to spontaneous supersymmetry breakdown. The specifics depend on the moduli stabilization mechanism, but the main features such as the scale of supersymmetry breaking hold more generally. In particular, the gravitino mass is given by

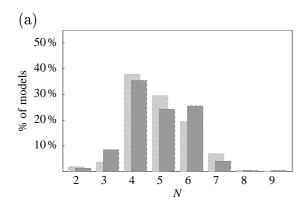
$$m_{3/2} \approx \frac{\Lambda^3}{M_{\rm Pl}^2} \,, \tag{5.44}$$

while the proportionality constant is model-dependent.

The gaugino condensation scale Λ is given by the renormalization group invariant scale of the condensing gauge group,

$$\Lambda \sim M_{\text{GUT}} \exp \left\{ -\frac{1}{2\beta} \frac{1}{g^2(M_{\text{GUT}})} \right\}. \tag{5.45}$$

With an appropriate mechanism, the dilaton can be stabilized at a realistic value $\text{Re}\langle S\rangle\approx 2$



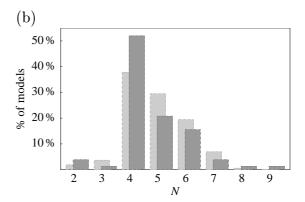


Figure 5.3: (a) Number of models vs. the size of largest gauge group in the hidden sector. N labels SU(N), SO(2N), E_N groups. The background corresponds to step @, while the foreground corresponds to step @ (see table 4.2). (b) As before, but with models of step @ at the foreground.

while breaking SUSY. As already mentioned, we see that for $\Lambda \sim 10^{13}$ GeV, the gravitino mass lies in the TeV range which is favored by phenomenology. SUSY breaking is communicated to the observable sector by gravity [57].

In fig. 5.3(a), we display the frequency of occurrence of various gauge groups in the hidden sector. The preferred size (N) of the gauge groups depends on the conditions imposed on the spectrum. When all inequivalent models with two Wilson lines are considered, N=4,5,6 appear with similar likelihood and N=4 is somewhat preferred. If we require the massless spectrum to be the MSSM + vectorlike matter, the fractions of models with N=4,5,6 become even closer. However, if we further require a heavy top quark and the decoupling of exotics at order eight, N=4 is clearly preferred (figure 5.3(b)). In this case, SU(4) and SO(8) provide the dominant contribution. Since all or almost all matter charged under these groups is decoupled, this leads to gaugino condensation at an intermediate scale. Possible scales of gaugino condensation are shown in fig. 5.4.

The correlation between the observable and hidden sectors comes about for a few reasons. First, it is due to modular invariance which ties the gauge shifts and Wilson lines in the two sectors. Second, the gauge shifts and Wilson lines in the hidden sector affect properties of the massless spectrum, for instance, through the masslessness equations.

We see that among the promising models, just as in the orbifold-MSSM presented in section 5.3.1, intermediate scale supersymmetry breaking is preferred. The underlying reason is that realistic spectra require complicated Wilson lines, which break the hidden sector gauge group. The surviving gauge factors are neither too big (unlike in Calabi–Yau compactifications with the standard embedding) nor too small.

There are significant uncertainties in the estimation of the supersymmetry-breaking scale. First, the identification of Λ with the renormalization group invariant scale is not precise. A factor of a few uncertainty in this relation leads to two orders of magnitude

⁷We note that before step ®, gaugino condensation does not occur in many cases due to the presence of hidden sector matter.

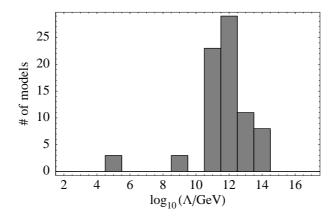


Figure 5.4: Number of models vs. scale of gaugino condensation.

uncertainty in $m_{3/2}$. Also, there could be significant string threshold corrections which can affect the estimate although we see from fig. (5.2) that threshold corrections might enhance the scale of SUSY breaking. Thus, the resulting "prediction" for the superpartner masses should be understood within 2-3 orders of magnitude.

5.4 *R*-Parity and Proton Decay

An essential property of the MSSM that a realistic model must exhibit is R-parity. This has the advantage of greatly reducing the number of arbitrary parameters in the superpotential, forbidding dimension three and four baryon or lepton number violating operators, and preserving a viable dark matter candidate, i.e. the LSP. However, obtaining a conserved R-parity in string constructions sets frequently an insurmountable hurdle, which must however be overcome in order to reach the MSSM.

Our strategy for accomplishing this is, in principle, quite simple. We identify first a $U(1)_{B-L}$ gauge symmetry and then give VEVs to some SM singlets that break $U(1)_{B-L}$ to the discrete subgroup

$$\mathbb{Z}_2^{\mathcal{M}} : (-1)^{3(B-L)}.$$
 (5.46)

This unbroken discrete symmetry is the so-called family reflection symmetry or matter parity [60], which, due to its properties, can be considered an R-parity. This is a global \mathbb{Z}_2 symmetry which is even on the Higgs doublets and odd on all SM quark and lepton fields. Further, it forbids dangerous baryon or lepton number violating operators of dimension three and four:

$$\overline{u}\,\overline{d}\,\overline{d}\,, \quad q\,\overline{d}\,\ell\,, \quad \ell\,\ell\,\overline{e} \quad \text{and} \quad \ell\,h_u\,.$$
 (5.47)

On the other hand, it allows quark and lepton Yukawa couplings as well as the Majorana neutrino mass operator $\bar{\nu}\bar{\nu}$. However, there are certain dimension five operators that are allowed by this symmetry:

$$\kappa_{ijkl}^{(1)} q_i \, q_j \, q_k \, \ell_l \,, \qquad \kappa_{ijkl}^{(2)} \bar{u}_i \, \bar{u}_j \, \bar{d}_k \, \bar{e}_l \,,$$
(5.48)

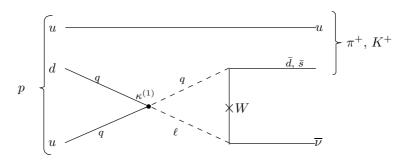


Figure 5.5: Possible contribution to proton decay, involving the dimension five operator $q q q \ell$. The proton decays as $p \to K^+ \bar{\nu}$ or $\pi^+ \bar{\nu}$.

which can induce proton decay (see fig. (5.5)). $\kappa^{(1)}$ and $\kappa^{(2)}$ are coupling constants. In the MSSM $\kappa^{(1)}$ is constrained by [159]

$$\frac{\kappa_{1121}^{(1)}}{\Lambda^{cut}} \approx \frac{\kappa_{1122}^{(1)}}{\Lambda^{cut}} \lesssim 16\pi^2 \frac{M_{\text{SUSY}}}{M_{\text{GUT}}^2}$$
(5.49)

and $\kappa^{(2)}$ can take arbitrary values as long as they are consistent with perturbation theory. Here, Λ^{cut} denotes an intermediate cut-off scale and $M_{\rm SUSY} \approx 10^4$ GeV is the scale at which SUSY is supposed to break.

We build the $U(1)_{B\!-\!L}$ symmetry as a superposition of all gauge U(1)'s of the MSSM candidates excepting the anomalous one.⁸ In our MSSM candidates, a natural choice for the generator of $U(1)_{B\!-\!L}$ follows from the standard breaking

$$SO(10) \rightarrow SU(5) \times U(1)_X \rightarrow G_{SM} \times U(1)_X,$$
 (5.50)

where the generators of hypercharge and B-L are given by

$$\mathsf{t}_Y = \left(-\frac{1}{2}, -\frac{1}{2}, \frac{1}{3}, \frac{1}{3}, \frac{1}{3}\right),\tag{5.51}$$

$$t_{B-L} = \frac{2}{5} (2 t_Y + t_X) = \left(0, 0, \frac{2}{3}, \frac{2}{3}, \frac{2}{3}\right)$$
 (5.52)

with $t_X = (1^5)$ generating the additional $U(1)_X$. Under these symmetries, including $SU(3)_c \times SU(2)_L$, an MSSM matter generation has the quantum numbers

where the subscripts denote hypercharge and B-L charge, respectively. Further, the Higgs doublets ϕ and $\bar{\phi}$ have U(1) charges (-1/2,0) and (1/2,0), respectively.

⁸In fact, this constraint might be relaxed. Yet in that case, one has to deal with the anomalies more carefully.

Unfortunately, the symmetry generated by $\mathbf{t}_{B\!-\!L}$ mixes with $\mathrm{U}(1)_A$ in our MSSM candidates and is, hence, ruled out. We have to search for a different "suitable" definition of $\mathrm{U}(1)_{B\!-\!L}$ which has to satisfy two conditions:

- (i) it must assign the standard B-L charges, eq. (5.53) to all SM particles (including a pair of Higgs fields), and
 - (ii) it must give most SM singlets a value that satisfies $3(B-L) = 0 \mod 2$.

Condition (i) is clear, but the second condition requires some argumentation. Firstly, SM singlets with charges $3(B-L)=0 \mod 2$ can obtain VEVs for decoupling exotics, as well as giving effective quark and lepton Yukawa couplings. Secondly, if there are some singlets (let us call them χ_i) satisfying $3(B-L)=0 \mod 2$ and with nontrivial B-L charge, they will break $\mathrm{U}(1)_{B-L}$, but will leave the R-parity, $\mathbb{Z}_2^{\mathcal{M}}$, unbroken. Moreover, note that if singlets with $3(B-L)=1 \mod 2$ develop nonvanishing VEVs, R-parity is broken and dimension four baryon/lepton number violating operators are typically generated. We would like to remark that this approach to get R-parity is not unique. Alternatives arise from considering other internal symmetries, such as R-charge conservation (inherited from Lorentz invariance in ten dimensions) [131] or T-duality [160].

Searching for generators of $U(1)_{B-L}$ is not difficult, but quite cumbersome. One has to express t_{B-L} as a superposition of all U(1)'s, excepting the anomalous U(1),

$$\mathsf{t}_{B\!-\!L} = \sum_{i \neq A} x_i \mathsf{t}_i \,, \tag{5.54}$$

and solve (non)linear equations on the real parameters x_i guaranteeing that the conditions (i) and (ii) explained above be fulfilled. The precise description of the method is provided in appendix D of ref. [78].

5.4.1 Supersymmetric MSSM Candidates

Before going any further, there is a question that we have to clarify. In section 4.4, it was assumed that all SM singlets can acquire VEVs. Nevertheless, we have learnt in section 5.2 a strategy to identify those singlets which admit a nonzero VEV in a consistent supersymmetric vacuum. In the present case, since B-L must be broken spontaneously, it results impossible to postpone the task of verifying D-flatness any longer.

Let us focus on the models based on the shifts $V^{SO(10),1}$ and $V^{SO(10),2}$, that is, in those that concentrate most promising models. According to our previous discussion, all we have to do is to verify whether there exist holomorphic invariant monomials (HIMs) with net negative anomalous charge. All fields entering the monomials are allowed to attain nonvanishing VEVs and, therefore, contribute to the masses of the exotics.

We recall our results from chapter 4 and complement them with our new findings in table 5.3. In particular, observe that we have incorporated the nontrivial criterion @'. We have verified that all vectorlike exotics acquire large masses. These results are crucial for the discussion on breaking of $U(1)_{B-L}$.

⁹We refer here to one of the selections rules. See section 2.5.

	V^{SO}	(10),1	V^{SO}	(10),2
Criterion	7	X	7	X
© spectrum = 3 generations + vectorlike	72	56	37	53
® exotics decouple at order 8	56	50	32	53
®' exotics decouple in SUSY configurations	55	50	32	53

Table 5.3: A subset of the MSSM candidates. The number of models with "heavy top" are listed under ②. 🕱 denotes no "heavy top". Only in one model, it was not possible to find a supersymmetric vacuum configuration where all exotics decouple.

Notice that we have not discussed F-flatness or any other additional constraint (e.g. $\langle W \rangle = 0$). As we mentioned in section 5.2, the reason is that other conditions are not relevant since they depend on technical particularities of the models, such as the order of the perturbative potential at hand. On the other hand, we do not have any prescription to write the full superpotential including all possible nonperturbative contributions. Additionally, we have shown that once one has found a D-flat configuration, it is always possible to find a nontrivial solution to the F-term equations.

5.4.2 R-Parity Invariant MSSM Candidates

For simplicity, we focus here only on those models from table 5.3 with a cubic top Yukawa coupling, that is, on those 55+32 MSSM candidates satisfying simultaneously criteria ② and ③'.

We notice that the shift $V^{\mathrm{SO}(10),1}$ seems still to be more promising in comparison to $V^{\mathrm{SO}(10),2}$. With our method, we find that 34 of the 55 (5 of the 32) MSSM candidates from the first (second) shift vector admit nonanomalous definitions of B-L. Each model, nevertheless, correspond to a big family of configurations with distinct acceptable $\mathrm{U}(1)_{B-L}$ symmetries. There are two reason for this to happen:

- In many cases there are vectorlike exotics with SM gauge charges identical to those of quarks and leptons. Further, without B-L, each lepton doublet ℓ could very well be a Higgs boson ϕ . Thus there are different ways to choose which of these states have standard B-L charges. Each choice can lead to a different (and nonanomalous) definition of B-L
- For each choice of SM particles above, there may be more than one B-L definition. The system of equations that determine the values of the variables x_i in eq. (5.54) might be underconstrained. Therefore, in some cases, there are (continuous) families of B-L generators (i.e. some of the x_i 's are free or fulfill equations with an undetermined number of solutions).

These ambiguities, although disturbing at first sight, might also be considered a rich source of new physics. However, it is clear that we need to implement new criteria in order to reduce the number of B-L configurations. We require the presence of SM singlets

_

with charges $B-L=0,\pm 2,\pm 2/3,\pm 2/5,\ldots,\pm 6/7$ in the spectrum¹⁰. Configurations with singlets having B-L charges consistent with $3(B-L)=0 \mod 2$, but different from the values just mentioned are automatically disregarded.

Including only these possibilities, we find 3447 (144) consistent B-L generators from the first (second) SO(10) shifts. The question is now whether all these configurations are inequivalent or not. It can be argued that, if one considers one single model (shift vector + Wilson lines + hypercharge generator) with several B-L configurations, we could distinguish the inequivalent configurations by comparing the sets of SM singlets with charges fulfilling $3(B-L)=0 \mod 2$. If these sets (and not the charges) coincide in two models, one might state that they are equivalent and, thus, that one of them can be ignored. This is indeed the approach we follow.¹¹

In this approach, we find that there are "only" 85 (8) inequivalent B-L configurations from $V^{\mathrm{SO}(10),1}$ ($V^{\mathrm{SO}(10),2}$). Further, as we are interested in models containing solely the SM gauge group times a hidden sector for gaugino condensation, one can verify whether all extra unbroken U(1)'s break when fields with proper B-L charges attain VEVs. Requiring the absence of extra unbroken U(1)'s reduces this set to 42 (0) acceptable configurations. Finally, demanding that all exotics decouple along D-flat directions leads to 15 (0) acceptable solutions with an exact low energy R-parity. We list the corresponding shifts, Wilson lines and generators of hypercharge and B-L in table E.1. These input parameters can be used to determine all other properties of the models.

This result is specific to our (B-L)-based strategy and we expect, in general, more acceptable models to exist. We point out that there is a big uncertainty in this number and, therefore, it should be taken as a lower bound.

5.4.3 R-Parity in the Orbifold-MSSM

The orbifold-MSSM allows us to define a suitable B-L generator which leads to the standard charges for the SM matter, eq. (5.53). The generator is given by

$$\mathbf{t}_{B-L} = \left(-1, -1, 0, 0, 0, \frac{2}{3}, \frac{2}{3}, \frac{2}{3}\right) \left(-2x + \frac{1}{2}, -\frac{1}{2}, 0, -x, -x, 0, 0, 0\right) , \tag{5.55}$$

with arbitrary x. Let us take $x=\frac{1}{2}$. Then the B-L generator we shall use is

$$\mathsf{t}_{B-L} = \left(-1, -1, 0, 0, 0, \frac{2}{3}, \frac{2}{3}, \frac{2}{3}\right), \left(-\frac{1}{2}, -\frac{1}{2}, 0, -\frac{1}{2}, -\frac{1}{2}, 0, 0, 0\right) . \tag{5.56}$$

An interesting feature is that the spectrum contains a pair of fields with B-L charges ± 2 , so that, if the B-L gauge symmetry is broken by VEVs of these fields, the matter parity $(-1)^{3(B-L)}$ is conserved and proton decay due to dimension three and four operators

¹⁰Note that we have included also charges that violate the condition $3(B-L) = 0 \mod 2$. This is also admissible because other discrete symmetries resulting from the breaking of $U(1)_{B-L}$ might be as good as $\mathbb{Z}_2^{\mathcal{M}}$. Further details can be found in appendix C of ref. [78]

 $^{^{11}\}mathrm{Nevertheless},$ soon we found that some configurations with identical sets of singlets with correct B-L charges led to different charges in the MSSM matter fields and therefore different phenomenological properties. It seems necessary to try a different approach.

is suppressed. In table F.1, we have summarized the matter spectrum of the model. Since now we include B-L, we can now distinguish between fields with identical SM quantum numbers; in particular, we can distinguish now between Higgs and lepton doublets.

The vacuum configuration (5.25) is still admissible in presence of $U(1)_{B-L}$. In this configuration, SM singlets with B-L charges ± 2 , labeled $\chi_1 \equiv s_1^0$ and $\chi_2 \equiv s_2^0$, attain nonzero VEVs and thus break B-L by two units, such that, e.g. particles with B-L charge 1 are equivalent to particles with B-L charge equal to -1. This example applies to the right-handed neutrinos to be studied in the next section.

Let us now proceed to enumerate some of the properties that the orbifold-MSSM acquires after breaking B-L.

Absence of proton decay operators of dimension three and four. This is, of course, not surprising since we have imposed B-L just to avoid these operators.

Cancellation of the μ -term. The Higgs mass terms are

$$\bar{\phi}_{i} (\mathcal{M}_{\bar{\phi}\phi})_{ij} \phi_{j}, \text{ where } \mathcal{M}_{\bar{\phi}\phi} = \begin{pmatrix} \widetilde{s}^{4} & 0 & 0 & \widetilde{s} \\ \widetilde{s} & \widetilde{s}^{3} & \widetilde{s}^{3} & \widetilde{s}^{6} \\ \widetilde{s}^{5} & 0 & 0 & \widetilde{s}^{3} \\ \widetilde{s} & 0 & 0 & \widetilde{s}^{3} \end{pmatrix}.$$

$$(5.57)$$

The up-type Higgs h_u is a linear combination of $\bar{\phi}_1$, $\bar{\phi}_3$ and $\bar{\phi}_4$,

$$h_u \sim \tilde{s}^2 \bar{\phi}_1 + \bar{\phi}_3 + \tilde{s}^4 \bar{\phi}_4 \,, \tag{5.58}$$

while the down-type Higgs is composed out of ϕ_2 and ϕ_3 ,

$$h_d \sim \phi_2 + \phi_3$$
 (5.59)

Let us remark that this spoils our argument about the existence of "heavy top".

Notice that in this vacuum configuration the μ -term, being defined as the smallest eigenvalue of $\mathcal{M}_{\bar{\phi}\phi}$,

$$\mu = \frac{\partial^2 W}{\partial h_d \, \partial h_u} \bigg|_{h_u = h_d = 0} \,, \tag{5.60}$$

vanishes up to order \tilde{s}^6 , at which we work. It is a way to deal with the supersymmetric μ -problem.

Charged fermion Yukawa matrices. The up-Higgs Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_u)_{ij}^{(k)} q_i \, \bar{u}_j \, \bar{\phi}_k \,,$$
 (5.61)

where the matrices $Y_u^{(i)}$ are given in appendix F. Thus, the physical 3×3 up-Higgs Yukawa matrix is

$$Y_u \sim \tilde{s}^2 Y_u^{(1)} + Y_u^{(3)} + \tilde{s}^4 Y_u^{(4)} = \begin{pmatrix} 0 & 0 & \tilde{s}^8 \\ 0 & 0 & \tilde{s}^8 \\ \tilde{s}^5 & \tilde{s}^5 & \tilde{s}^2 \end{pmatrix}.$$
 (5.62)

Note that due to the Higgs mixing the top quark Yukawa coupling for this vacuum configuration is given by \tilde{s}^2 . Therefore, the corresponding \tilde{s} VEVs are required to be quite large.

The down-Higgs Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_d)_{ij}^{(k)} q_i \, \bar{d}_j \, \phi_k \,,$$
 (5.63)

where again the Yukawa matrices are provided in appendix F.

The physical 3×3 down-Higgs Yukawa matrix emerges by integrating out a pair of vectorlike d- and $\bar{d}-$ quarks,

$$Y_d = \begin{pmatrix} 1 & \widetilde{s}^3 & 0 \\ 1 & \widetilde{s}^3 & 0 \\ \widetilde{s} & \widetilde{s}^4 & 0 \end{pmatrix}. \tag{5.64}$$

We note that both the up and down quarks are massless at order six in SM singlets. However, we have checked that the up quark becomes massive at order seven and the down quark gets a mass at order eight.

Analogously, the physical 3×3 matrix emerges by integrating out a pair of vector-like ℓ - and $\bar{\ell}$ -leptons,

$$Y_e = \begin{pmatrix} 1 & 1 & \widetilde{s} \\ \widetilde{s} & \widetilde{s} & \widetilde{s}^2 \\ 0 & 0 & \widetilde{s}^6 \end{pmatrix}. \tag{5.65}$$

Dimension five baryon and lepton number violating operators. We have looked for effective dimension five baryon and lepton number violating operators in this model. We find that to order \tilde{s}^6 no such operators exist. However, these operators can be generated once the exotics δ_i , $\bar{\delta}_i$ are integrated out. Fortunately, a clever choice of VEVs for the fields $\{\tilde{s}_i\}$ can guarantee sufficient suppression of all induced $q q q \ell$ operators, consistent with current bounds on proton decay, eq. (5.49).

 μ -term and Minkowski space. In many MSSM candidates, it happens that the μ -term is identified with the SM singlet superpotential, $W(\widetilde{s}_i)$. Unfortunately, this relation does not hold the orbifold-MSSM. This is intimately related to the number of Higgs pairs found. In the orbifold-MSSM there are four fields of this type and they mix. In models where the relation $\mu = \langle W(\widetilde{s}_j) \rangle = 0$ holds, there is only one pair of Higgses and generically they are bulk fields.

In the orbifold-MSSM we could not find any argument against the eventual instauration of a deep anti-De Sitter vacuum.

Further details of the orbifold-MSSM regarding the effect of our B-L choice on the neutrino masses will be discussed in the next section. The detailed mass matrices are displayed in appendix F.

5.5 Neutrino Masses

The seesaw mechanism [61–63] is perhaps the most attractive way to explain the smallness of the neutrino masses. Its essential ingredients are heavy Majorana neutrinos and their Yukawa couplings to the left–handed neutrinos. The supersymmetric seesaw mechanism is described by the superpotential

$$W = Y_{\nu}^{ij} \,\bar{\phi} \,\ell_i \,N_j + \frac{1}{2} M_{jk} N_j \,N_k \,, \qquad (5.66)$$

where $\bar{\phi}$ and ℓ_i (i=1,2,3) are the up-type Higgs and lepton doublets, and N_j ($1 \leq j \leq n$) are some heavy standard model singlets. At low energies, this leads to three light neutrinos with masses of order $(Y_{\nu} \langle \bar{\phi} \rangle)^2/M$, where Y_{ν} and M represent typical values of Y_{ν}^{ij} and M_{jk} , respectively. For $Y_{\nu} \sim 1$ and $M \sim 10^{16} \, \text{GeV}$, one has $m_{\nu} \sim 10^{-3} \, \text{eV}$. The scales of the atmospheric and solar neutrino oscillations, [161]

$$\sqrt{\Delta m_{\rm atm}^2} \simeq 0.04 \,\text{eV} \,, \qquad \sqrt{\Delta m_{\rm sol}^2} \simeq 0.008 \,\text{eV} \,, \tag{5.67}$$

are suspiciously close to this scale. This hints at GUT structures behind the seesaw.

In conventional GUTS, N_j are members of GUT matter multiplets, e.g. a **16**-plet of SO(10), and M_{jk} are related to a VEV of a large GUT representation, e.g. a $\overline{126}$ -plet of SO(10). In this case the Majorana mass terms originate from the coupling $16 \ \overline{126}$ (cf. e.g. [162]).

5.5.1 Seesaw Mechanism with Several Neutrinos

Even though in our scheme we have local GUTs, the Yukawa couplings do not necessarily preserve the symmetry of these GUTs. The symmetry of the nonlocal coupling in ten dimensions is an intersection of the local gauge groups at the vertices. This implies, for example, that the Majorana mass terms for the neutrino components of the 16-plets can originate from the coupling

$$\nu_{16} \times \nu_{16} \times (\text{SM singlets})$$
, (5.68)

where the singlets belong neither to $\overline{\bf 16}$ nor to $\overline{\bf 126}$ of SO(10).

Furthermore, any SM singlet can play the role of the right-handed neutrino as long as it has a Yukawa coupling to the lepton doublets and a large Majorana mass. These are abundant in orbifold models and typical models contain $\mathcal{O}(100)$ such singlets.

MSSM candidates contain an anomalous U(1) which induces the FI D-term,

$$D_{\text{anom}} = \frac{g M_{\text{Pl}}^2}{192\pi^2} \operatorname{tr} t_{\text{anom}} + \sum_{i} q_i |\phi_i|^2, \qquad (5.69)$$

where t_A is the anomalous U(1) generator, q_i are the anomalous charges of fields ϕ_i and g is the gauge coupling. This triggers spontaneous gauge symmetry breaking while preserving

supersymmetry [104]. Some of the fields charged under the anomalous U(1) (and, in addition, under other gauge groups) develop nonzero VEVs thereby reducing gauge symmetry. The scale of these VEVs is set by the FI term which is somewhat below the string scale. This eventually determines the seesaw scale. In general, any SM singlets can get large VEVs as long as it is consistent with supersymmetry, and one can obtain the standard model gauge symmetry times that of the hidden sector,

$$\mathcal{G}_{4D} \longrightarrow G_{SM} \times G_{\text{hidden}}$$
 (5.70)

The singlet VEVs are not necessarily associated with flat directions in the field space and generally correspond to isolated solutions to supersymmetry equations [48]. The hidden matter gauge group $G_{\rm hidden}$ can be responsible for spontaneous supersymmetry breaking. In fact, within the class of models with the MSSM spectrum, gaugino condensation in the hidden sector favors TeV-scale soft masses for the observable fields, as we have seen in section 5.3.2.

The nonzero singlet VEVs lead to the mass terms for the vectorlike states,

$$W = x_i \bar{x}_j \langle s_a s_b \dots \rangle, \qquad (5.71)$$

where x_i, \bar{x}_j are the vectorlike exotics and $\langle s_k \rangle$ are the SM singlet VEVs in string units. Such a coupling must be consistent with string selection rules. In section 5.4, it has been shown that many \mathbb{Z}_6 -II models satisfy this requirement and all of the vectorlike exotics can be decoupled. This results in the MSSM spectrum at low energies.

Similarly, the singlet VEVs induce "Majorana" mass terms for the SM singlets as well as the neutrino Yukawa couplings of eq. (5.66),

$$M_{ij} \sim \langle s_a s_b \dots \rangle \ , \ Y_{\nu}^{ij} \sim \langle s_{\alpha} s_{\beta} \dots \rangle \ ,$$
 (5.72)

as long as it is consistent with string selection rules.

Identification of right—handed neutrinos is intimately related to the issue of baryon/lepton number violation. In generic vacua, any SM singlet can play the role of the right—handed neutrino. However, such vacua also suffer from excessive R-parity violating interactions. The simplest way to suppress these interactions is to identify a B-L gauge symmetry and enforce either its approximate conservation or conservation of its discrete (matter parity) subgroup. In local GUTs, the B-L generator resembles the standard GUT B-L, but also requires extra U(1) components beyond SO(10). It is nonanomalous and produces the standard B-L charges for the SM matter. If B-L is broken by VEVs of fields carrying even charges under B-L, the matter parity $(-1)^{3(B-L)}$ is conserved. This forbids dangerous R-parity violating interactions and requires the right-handed neutrino to carry the charge $q_{B-L}=\pm 1$. Another possibility is that $\mathrm{U}(1)_{B-L}$ is broken at an intermediate scale M_{B-L} such that all R-parity violating couplings are suppressed by M_{B-L}/M_{Pl} . In this case, Majorana mass terms for the right-handed neutrinos are allowed only upon B-L breaking, which lowers the seesaw scale to intermediate energies. In what follows, we consider these possibilities in specific heterotic orbifold models.

5.5.2 Seesaw Mechanism on the Orbifold-MSSM

The B-L generator is identified with

$$\mathbf{t}_{B-L} = \left(-1, -1, 0, 0, 0, \frac{2}{3}, \frac{2}{3}, \frac{2}{3}\right) \left(-\frac{1}{2}, -\frac{1}{2}, 0, -\frac{1}{2}, -\frac{1}{2}, 0, 0, 0\right) . \tag{5.73}$$

In general supersymmetric configurations, many SM singlets get nonzero VEVs. Choosing a subset of such singlets with 0 or $\pm 2~B-L$ charges, the unbroken gauge symmetry is

$$G_{\rm SM} \times G_{\rm hidden}$$
, (5.74)

where $G_{\text{hidden}} = \text{SO}(8)$, while all of the exotic states get large masses and decouple. This vacuum preserves the matter parity $(-1)^{3(B-L)}$.

We find that there are 39 right-handed neutrinos defined by $q_{B-L}=\pm 1$, two of which are members of the localized 16-plets. They have Yukawa couplings to the lepton doublets and large Majorana mass terms. We have calculated the 3×39 Yukawa matrix Y_{ν} and 39×39 Majorana mass matrix M of eq. (5.66) up to order six in the singlet VEVs. That is, for each matrix element, we have determined at which order in the superpotential a nonzero coupling is allowed by string selection rules. Each entry depends on the quantum numbers and the localization of the Majorana neutrinos, and involves products of different singlets and moduli-dependent Yukawa couplings. We then assume that the main hierarchy in these entries comes from products of singlet VEVs so that these matrices can be treated as textures.

The effective mass matrix for the left-handed neutrinos.

$$M_{\text{eff}} = -v_u^2 Y_\nu M^{-1} Y_\nu^T \tag{5.75}$$

with v_u being the up-type Higgs VEV, can be represented by the texture

$$M_{\text{eff}} \sim -\frac{v_u^2}{M_*} \begin{pmatrix} 1 & s & s \\ s & s^2 & s^2 \\ s & s^2 & s^2 \end{pmatrix} .$$
 (5.76)

Here s < 1 represents a generic singlet VEV in string units and M_* is the effective seesaw scale. Y_{ν} contains entries with powers of s between 1 and 5, while the dependence of the eigenvalues of M ranges from s to s^8 (with no massless eigenstates at generic points in moduli space). This results in a strong s-dependence of the effective seesaw scale M_* . This scale is further suppressed by the large multiplicity of heavy singlets N, $M_* \propto N^{-x}$ with 0 < x < 2. The value of x depends on the texture. For example, when all the singlets contribute equally, x = 2, whereas x = 0 if only a fixed number of neutrinos have nonnegligible couplings. For the present model, we find

$$M_* \sim 0.1 \, s^5 \, M_{\rm str} \sim 10^{14} \, {\rm GeV} \,,$$
 (5.77)

for the string scale $M_{\rm str}=2\cdot 10^{17}\,{\rm GeV}$ and $s\sim 0.3$. The obtained texture (5.76) is of course model dependent.

The corresponding charged lepton Yukawa matrix is of the form

$$Y_e \sim \begin{pmatrix} 1 & 1 & c \\ c & c & c^2 \\ 0 & 0 & 0 \end{pmatrix} ,$$
 (5.78)

where "0" denotes absence of the coupling up to order six in the singlet VEVs c. Such zeros are expected to be filled in at higher orders. Here we are again using a single expansion parameter although in practice there are many variables.

These crude estimates show that reasonable fermion masses can in principle be obtained. Inserting order one coefficients in the textures, one finds that the eigenvalues scale as

$$m_{\nu_i} \sim (1, s^2, s^2) \frac{v_u^2}{M_*}, \quad m_{e_i} \sim (1, c, 0) v_d,$$
 (5.79)

where v_d is the down-type Higgs VEV. For $s \sim 0.3$ and $c \sim 0.1$ the textures reproduce roughly the observed lepton mass hierarchy. The above texture favors the normal neutrino mass hierarchy and can accommodate small and large mixing angles. Further details of the model are available in appendix F.

Chapter 6

Conclusions and Outlook

This thesis was based on orbifold compactifications of the heterotic string. We have emphasized its relation to low–energy phenomenology and, therefore, its viability as a unified theory of all observed forces. In particular, we have focused on the phenomenology of abelian orbifold compactifications of the $E_8 \times E_8$ heterotic string theory.

We have described in detail the heterotic string and provided some recipes for the computation of the direct properties of orbifold models, such as their fixed points and their matter spectra. Along with this analysis, we have commented about the inclusion of some additional degrees of freedom denominated discrete torsion. In concrete, we have shown that, contrary to previous statements, \mathbb{Z}_N orbifolds with Wilson lines as well as $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifold models admit discrete torsion. Further, we have found out that the assignments of discrete torsion can be traded for translations of the gauge embedding parameters (shifts and Wilson lines) in the lattice Λ .

We have explained and extended some conventional methods employed to classify orbifold models, paying special attention to the construction of ansätze that allow us to obtain systematically all inequivalent input parameters of orbifold compactifications, that is, shift vectors and Wilson lines. The classification method obtained is very flexible and can be adapted to both $E_8 \times E_8$ and SO(32) heterotic string theories. Our methods have already been tested in all \mathbb{Z}_N and $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds. We have made an exhaustive classification of \mathbb{Z}_N orbifolds without background fields and also a complete classification of $\mathbb{Z}_3 \times \mathbb{Z}_3$ orbifold models, including discrete torsion. Whenever possible, we have compared our results to those of the literature.

It is clear, however, that the task of getting all orbifold models or computing their properties requires some automatization. We have built the C++ Orbifolder, a program written in c++ language capable to compute most of the features of orbifold constructions in a fraction of a second. We consider that such a tool might be of some interest for the scientific community and, therefore, we plan to publish it in the future.

We have analyzed particularly the heterotic $E_8 \times E_8$ string compactified on a \mathbb{Z}_6 -II orbifold, allowing for up to two discrete Wilson lines. Employing a search strategy based on the concept of local GUTs, we have obtained about 3×10^4 inequivalent models. Out of them, 218 models exhibit the MSSM gauge group structure, three light families and vec-

torlike exotics. We show that all the vectorlike exotics can decouple without breaking supersymmetry for 190 of these models. This means that almost 1% out of 3×10^4 models have the gauge group and the chiral matter content of the MSSM. This result shows that orbifold compactifications of the heterotic string correspond to a particularly fertile region in the landscape and that the probability of getting something close to the MSSM is significantly higher than that in other constructions. It would be interesting to extend these results to other regions of the landscape where some promising models also exist.

Furthermore, we have found that requiring realistic features in the set of MSSM candidates is correlated with the supersymmetry breaking scale such that, in the context of gaugino condensation, intermediate scale ($\sim 1 \text{ TeV}$) of SUSY breaking is favored. This occurs because most of the models with realistic features have a hidden sector with SU(4) or SO(8) nonabelian symmetry.

In order to get closer to the MSSM, we defined a successful strategy for obtaining models with an exact R-parity, based on spontaneous breaking of a gauge $\mathrm{U}(1)_{B-L}$ symmetry. We find 87 models which have a renormalizable top Yukawa coupling. Out of these, we identify 15 models with an exact R-parity, no light exotics or $\mathrm{U}(1)$ gauge bosons and an order one top quark Yukawa coupling. We would like to remark that the number 15 is a lower bound, mainly since our search is based on a specific strategy related to B-L symmetry. Additionally, many of our constraints, such as demanding a renormalizable coupling for the top quark are just artifacts invoked in order to simplify our analysis. Therefore, an interesting question would be to obtain similar results in cases where: a) no trilinear coupling for the top quark is demanded, b) no vectorlikeness of the SM matter generations with respect to B-L is imposed, c) also models with anomalous B-L are admited, and d) R-parity results directly from the so-called R-charge conservation (string) selection rule.

On the other hand, we notice that dimension five baryon and lepton number violating operators, like $q q q \ell$ appear in orbifold models quite regularly even if B-L is imposed. They are either generated in the superpotential to some order in SM singlets, or they may also be generated when integrating out heavy exotics. In some cases, only fine-tuning the VEVs of some SM singlets can alleviate this problem. One is thus encouraged to investigate other possible solutions, such as the identification of some discrete symmetries.

Another aspect of phenomenology that captured our attention is the seesaw mechanism. Since there are no upper bounds on the amount of right-handed neutrinos that might appear, we find that in a scenario with $\mathcal{O}(100)$ right-handed neutrinos, the seesaw mechanism is realized from orbifold constructions. Moreover, (left-handed) neutrino masses are then enhanced (in comparison to the naïve estimate of the neutrino masses in GUTs) to more realistic values. We consider this setup to be plausible and perhaps even desirable, since some consequences of the existence of many right-handed neutrinos could shed some light in some open issues, such as leptogenesis [163,164]. We might even consider the abundance of right-handed neutrinos to be a prediction of string theory.

We have also studied in detail one model that we have called orbifold-MSSM. In fact, this model is not special in the sense that other MSSM candidates possess very similar qualities: only the exact matter spectrum of the MSSM (up to many SM singlet fields); admissible B-L symmetry which is spontaneously broken to matter parity in a supersymmetric vacuum

configuration; gaugino condensation scale leading to SUSY breaking at an intermediate scale; vanishing μ -term; and proton decay through dimension five operators can be avoided by some fine-tuning. The top Yukawa coupling is order \tilde{s}^2 due to Higgs doublet mixing in this model. Further, both the up and down quarks are massless at order six in SM singlets. However, the up quark becomes massive at order seven and the down quark gets a mass at order eight. These properties are very interesting, but we believe that the best model awaits still in some corner of the landscape.

There are some phenomenological issues that we have not addressed in this thesis. In particular, one issue concerns proton stability. The examples we studied are challenged by the presence of dimension five proton decay operators. Their suppression may require additional (discrete) symmetries or even a change in the geometry of the orbifold, as suggested in appendix C. There are also dimension six operators, generated by GUT gauge boson exchange, which we have not discussed.

Further, we have not studied precision gauge coupling unification. Although hypercharge is normalized as in four-dimensional GUTs thus allowing gauge coupling unification in the first approximation, there are various corrections that can be important. First, a detailed analysis would require the calculation of string threshold corrections in the presence of discrete Wilson lines. However in specific cases these corrections are known to be small [165]. Second, there are corrections from the vectorlike exotic states. It is possible that precision gauge coupling unification may require anisotropic compactifications, leading to an effective orbifold GUT. [47, 166–168].

Finally, there are the usual questions of moduli stabilization and supersymmetry break-down in a Minkowski vacuum. Some of them we discussed previously in [76]. We have not addressed all of these issues here. On the other hand, it is clear that if given the freedom of arbitrarily tuning moduli VEVs we are not able to find the MSSM, the whole approach would be futile. However, with a number of MSSM candidates in this fertile patch of the landscape, it is now imperative to tackle the hard problems just mentioned.

Appendix A

Form of Shift Vectors and Wilson lines

Even though there is a straighforward prescription to obtain all \mathbb{Z}_N shift vectors (see section 3.1), it is necessary to implement a method which can be easily extended to $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds and orbifolds with Wilson lines. An option is an ansatz that describes any shift or Wilson line. To start with, let us specify an ansatz for \mathbb{Z}_N shifts of SO(32) orbifolds. We will see in section A.2 that the resulting ansatz can be then trivially generalized.

A.1 \mathbb{Z}_N Shift Vectors of SO(32) Orbifolds

We will call vectorial shifts to those shift vectors of order N ($NV \in \Lambda$) whose entries have a maximal denominator of N. Spinorial shifts will be those other which have a denominator of 2N and an odd numerator. As an example, consider the \mathbb{Z}_4 shift vectors of table D.5. The first 12 of them are vectorial, whereas the last four are spinorial. It is convenient to describe separately these two cases and further distinguish between \mathbb{Z}_N shifts with N even and N odd.

To obtain the general form of a shift, we take for granted that two shift vectors are equivalent if they are related by lattice vectors or by Weyl reflections, i.e. by any permutation of the entries and pairwise sign flips.¹

a) Vectorial Shifts and N Even

Invariance under lattice translations implies that the entries V_i of a shift of order N are constrained by $-(N-1)/N \le V_i \le 1$. Therefore, we can start with the ansatz

$$V = \frac{1}{N} \left(-(N-1)^{n_{-(N-1)}}, -(N-2)^{n_{-(N-2)}}, \dots, 0^{n_0}, 1^{n_1}, \dots, N^{n_N} \right)$$
 (A.1)

for a vectorial shift. As usual, the exponent n_i of an entry i counts how many times that entry is repeated; for instance, the entry -(N-1) appears $n_{-1(N-1)}$ times. Therefore, the

 $^{^{1}\}mathbb{Z}_{N}$ orbifold models without Wilson lines do not change under lattice translations, unless one introduces lattice-valued Wilson lines. In such case, the new model can be a *brother model*. Further details are provided in section 2.4.3.

exponents n_i are integers satisfying

$$\sum_{i} n_i = 16. \tag{A.2}$$

Since pairwise sign-flips are allowed, we see that all negative entries can be made nonnegative as long as at least one n_{-i} is even (i = 1, ..., N). However, if there is an odd number of negative entries and $n_0 = 0$, one negative entry will remain. Combining both results, we get

$$V = \frac{1}{N} ((-j)^{\alpha}, 0^{n_0}, 1^{n_1}, \dots, j^{n_j - \alpha}, \dots, (N - 1)^{n_{N-1}}), \qquad 1 \le j \le N - 1$$
 (A.3)

with $\alpha=0,1$. This ansatz contains still some redundancies. Lattice translations of the shifts amounts to adding ± 1 to an even number of entries of V. One can thus apply $V_i \to V_i - 1$ for an even number of entries fulfilling $V_i > \frac{N}{2}$ and then flip their signs. By this operation, a \mathbb{Z}_N vectorial shift can be finally expressed by

$$V = \frac{1}{N} \left((\pm j)^{\alpha}, -(N-j)^{\beta}, 0^{n_0}, 1^{n_1}, \dots, (N-j)^{n_{(N-j)}-\alpha-\beta}, \dots, \left(\frac{N}{2}\right)^{n_{\frac{N}{2}}} \right), \tag{A.4}$$

where $\alpha, \beta = 0, 1$ such that $\alpha + \beta = 0, 1$ and j can take the values $\{\frac{N}{2} + 1, \dots, N - 1\}$.

As a side remark, using the shift V of eq. (A.4) and the SO(32) simple roots α_i provided in table D.7, one can compute $\alpha_i \cdot V$. Since those roots with $\alpha_i \cdot V \neq 0$ are projected out while the others form the Dynkin diagram of the unbroken gauge group \mathcal{G}_{4D} , one finds that the symmetry breakdown induced by a generic vectorial shift of SO(32) orbifolds is given by

$$SO(32) \longrightarrow SO(2n_0) \times U(n_1) \times ... \times U(n_{\left(\frac{N}{2}-1\right)}) \times SO(2n_{N/2}),$$
 (A.5)

recalling that $U(n) = SU(n) \times U(1)$.

b) Spinorial Shifts and N Even

An analogous analysis to that presented in the previous case leads to the conclusion that a generic spinorial shift can be written as

$$V = \frac{1}{2N} \left((\pm j)^{\alpha}, -(2N-j)^{\beta}, 1^{n_1}, 3^{n_3}, \dots, (2N-j)^{n_{(2N-j)}-\alpha-\beta}, \dots, (N-1)^{n_{N-1}} \right), \quad (A.6)$$

with $j \in \{N+1, N+3, \dots, 2N-1\}$. It is again not hard to confirm that a shift vector of this kind leads to the breaking

$$SO(32) \longrightarrow U(n_1) \times U(n_3) \times ... \times U(n_{N-3}) \times U(n_{N-1}).$$
 (A.7)

c) N Odd

The general form of the shift vectors changes slightly. First, as explained in ref. [169], in this case it is enough to determine either the vectorial or the spinorial shifts, since one spinorial shift can always be transformed into a vectorial one by the action of Weyl reflections and lattice vectors. Choosing the vectorial form, we obtain the following shift vector of odd order:

$$V = \frac{1}{N} \left((\pm j)^{\alpha}, -(N-j)^{\beta}, 0^{n_0}, 1^{n_1}, \dots, (N-j)^{n_{(N-j)}-\alpha-\beta}, \dots, \left(\frac{N-1}{2} \right)^{n_{(\frac{N-1}{2})}} \right), \quad (A.8)$$

where $j \in \{\frac{N+1}{2}, \frac{N+3}{2}, \dots, N\}$. The resulting four dimensional gauge group is

$$SO(32) \longrightarrow SO(2n_0) \times U(n_1) \times ... \times U(n_{\left(\frac{N-3}{2}\right)}) \times U(n_{\left(\frac{N-1}{2}\right)}).$$
 (A.9)

The shifts obtained by the ansätze given in eqs. (A.4)–(A.8) are further restrained by $NV \in \Lambda$ and the requirement that follows from invariance of the orbifold partition function:

$$N\left(V^2 - v^2\right) = 0 \bmod 2 \tag{A.10}$$

for a given twist vector v of order N. Only those shifts satisfying these conditions are consistent with orbifold compactifications. The ansätze (A.4)–(A.8) constitute, as we will see in the next section, the basic constructing blocks of general $\mathbb{Z}_N \times \mathbb{Z}_M$ shifts and Wilson lines.

A.2 A General Ansatz

Now we are in position to propose an ansatz which covers all orbifolds, including Wilson lines. Consider a model with a given shift vector V. Independently of whether V belongs to an $E_8 \times E_8$ or an SO(32) orbifold, Weyl reflexions and lattice translations allow to make its entries as small as possible and to gather them in blocks $V_{\text{block i}}$, just as in the ansätze (A.4)–(A.8):

$$V = \frac{1}{N} \left(V_{\text{block 1}}, V_{\text{block 2}}, V_{\text{block 3}}, \dots \right), \tag{A.11}$$

where each subvector $V_{\text{block i}}$ contains m_i identical integer or half-integer entries, such that $\sum_i m_i = 16$. Our task is then to find a universal second shift V_2 of order M, in the case of $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds, or/and the general form of a Wilson line A_{α} of order N_{α} , trying to avoid redundancies at maximum. From the block-structure of the shift vector, we can expect V_2 and A_{α} also to be splitted in blocks of length m_i as

$$V_2 = (V_{2,\text{block }1}, V_{2,\text{block }2}, V_{2,\text{block }3}, \dots),$$
 (A.12a)

$$A_{\alpha} = (A_{\alpha, \text{block } 1}, A_{\alpha, \text{block } 2}, A_{\alpha, \text{block } 3}, \dots) . \tag{A.12b}$$

This would mean that each block $V_{2,\text{block i}}$ and $A_{\alpha,\text{block i}}$ is independent and can, therefore, be described separately. In other words, we have just to find an ansatz for a subvector

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 $V_{2,\mathrm{block}\ i}$ or $A_{\alpha,\mathrm{block}\ i}$ of the corresponding order $(M\ \mathrm{or}\ N_{\alpha})$ and then to apply it consecutively over all blocks.

In the previous section, we have seen that the ansätze (A.4)–(A.8) describe the minimal form of a shift of order N. They can be used in order to build the blocks $V_{2,\text{block i}}$ and $A_{\alpha,\text{block i}}$, by replacing N by M or N_{α} and demanding that

$$\sum_{i} n_{j} = m_{i} \tag{A.13}$$

for each block. We have additionally to guarantee that the vectors obtained in this way lie on the lattice Λ , i.e. one has to impose MV_2 , $N_{\alpha}A_{\alpha} \in \Lambda$. This condition implies that, for even N, all blocks must be of the same type (either vectorial or spinorial). Clearly, admissible orbifold models are only those with shift(s) and Wilson lines fulfilling the modular invariance conditions provided in eq. (2.42).

Appendix B

Spinors in SO(32) Orbifolds

Early works on orbifold compactifications considered the SO(32) heterotic string to be phenomenologically disfavored. One of the reasons being that the theory by itself does not include spinors in its massless spectrum. This situation was shown to be preserved in the simplest orbifold compactifications, leading to the somewhat naïve conclusion that, if spinors can be found from the SO(32) heterotic string at all, they must appear in a quite unnatural way.

Only recently, the interest in four-dimensional heterotic SO(32) orbifold constructions has been revived [170,171]. In the case of \mathbb{Z}_N orbifolds, a complete classification of gauge embeddings in absence of background fields has been achieved [73,169]. Interestingly, that classification has shown that spinors of SO(2n) gauge groups appear rather frequently in the twisted sectors of SO(32) orbifolds. Spinors of SO(10), in particular, are found locally at fixed points of the first twisted sector of many orbifold models. Moreover, that classification reveals that the amount of available SO(32) orbifold models is comparable, at the same level, to that of its more famous brother: the $E_8 \times E_8$ string. Thus, it is not adventurous rather to conclude that model building based on the heterotic SO(32) string theory might be as interesting as that based on the $E_8 \times E_8$ theory.

On the other hand, the appearance of spinors on models derived from the SO(32) heterotic string might also be important for a possible understanding of the SO(32) heterotic type I duality in four spacetime dimensions. We know that spinors do not appear in the perturbative type I theory. Thus, the mentioned duality requires the implementation of nonperturbative effects.

B.1 SO(10) Spinors and Shift Vectors

Let us investigate here the possibility of having the 16-dimensional spinor representation of SO(10) in \mathbb{Z}_N orbifolds of the SO(32) heterotic string theory. In the standard basis of the

roots of SO(10), the highest weight of the 16-plet is given by the five-dimensional vector ¹

$$\mathbf{16}_{\mathrm{HW}} = \left(\frac{1}{2}^{5}\right). \tag{B.1}$$

To be included in the spectrum of an orbifold compactification, this vector must be part of a 16-dimensional solution $p_{\rm sh}$ to the mass formula for massless states, eq. (2.68) or eq. (2.81). Therefore, $p_{\rm sh}$ must be written as

$$p_{\rm sh} = p + V_g = \left(\frac{1}{2}, a_1, \dots, a_{11}\right),$$
 (B.2)

where $p \in \Lambda_{\text{Spin}(32)/\mathbb{Z}_2}$, V_g is the local shift of a constructing element g and $a_i \in \mathbb{R}$ are selected so that p_{sh} fulfills $Np_{\text{sh}} \in \Lambda_{\text{Spin}(32)/\mathbb{Z}_2}$ and the mass formula. There can be more than one combination of different a_i 's for which the resulting p_{sh} satisfies all the conditions.

A first consequence of eq. (B.2) is that one cannot get **16**-plets of SO(10) in the untwisted sector of any orbifold compactification. The reason is that the only $p_{\rm sh}$ available for untwisted massless states correspond to the roots of SO(32), which are expressed by the 480 vectors (± 1 , ± 1 , 0^{14}).

As a second consequence, one finds that it is not possible to get 16-plets of SO(10) in the \mathbb{Z}_3 orbifold. This can be understood as follows. Since in \mathbb{Z}_3 orbifolds $3p_{\rm sh}$ must lie on the lattice $\Lambda_{\rm Spin(32)/\mathbb{Z}_2}$ and the first five entries of $3p_{\rm sh}$ are half-integer, all a_i 's must also be half-integer, $3a_i \in \mathbb{Z} + 1/2$. Assuming in the best of the cases that $3a_i = 1/2$, it follows that $p_{\rm sh}^2 \geq 14/9$. However, masslessness of the states in \mathbb{Z}_3 orbifolds constrains $p_{\rm sh}$ to have at most squared length 4/3.

It is possible to find the form of the local shift vectors producing SO(10) spinors. From eq. (B.2), for all \mathbb{Z}_N orbifolds with N > 3 the local shift(s) giving rise to the **16** representation of SO(10) can be written as

$$V_q = p_{\rm sh} - p \tag{B.3}$$

with $p_{\rm sh}$ given as in eq. (B.2).

Hence, we can determine the shift vector V of an orbifold model containing SO(10) spinors in the first twisted sector, relevant for phenomenology, as we have discussed in section 4.1. In absence of Wilson lines, the local shift vector V_g at any fixed point of the first twisted sector coincides with V, therefore

$$V = p_{\rm sh} - p \xrightarrow{p=0} p_{\rm sh}, \tag{B.4}$$

where we have chosen p = 0 because two shifts are equivalent if they differ by an arbitrary lattice vector. This shift vector is automatically modular invariant. For each allowed $p_{\rm sh}$ from eq. (B.2), one can verify that the shift vector V obtained through eq. (B.4) leads to a four-dimensional SO(10) gauge group by applying the patterns provided in appendix A.1.

¹ As usual, powers of an entry represent how often that entry is repeated in the vector.

As an example, we consider the \mathbb{Z}_4 orbifold without Wilson lines. The only possible shift vector consistent with $4V \in \Lambda_{\mathrm{Spin}(32)/\mathbb{Z}_2}$, the masslessness condition eq. (2.81) and eq. (B.4) is

$$V = p_{\rm sh} = \left(\frac{1}{2}^5, \frac{1}{4}^2, 0^9\right),$$
 (B.5)

which is one of the admissible shift vectors for the \mathbb{Z}_4 orbifold, listed in table D.5.

Appendix C

Z₆-II on Nonfactorizable Lattices

 \mathbb{Z}_6 -II orbifold compactifications on the lattice $G_2 \times SU(3) \times SO(4)$ have proven to lead to realistic models. The question here is how the properties of such models are influenced by the use of one of the other allowed lattices of \mathbb{Z}_6 -II. In the best of the cases, the possible changes shall be of relevance when dealing with unsolved phenomenological issues such as proton decay.

C.1 Lattices and Spectrum

The allowed lattices of \mathbb{Z}_6 -II orbifolds are presented in table C.1. Note that, with exception of the lattice (A), $G_2 \times SU(3) \times SO(4)$, all lattices are nonfactorizable. These lattices are independent in the sense that they cannot be deformed continuously into one another. Therefore, one knows that orbifolds on those lattices will have certainly distinct properties.

The geometry of the compact space will affect some other properties of orbifold models. Firstly, one finds that the number of fixed points of the orbifold varies according to

	Lattice	T_1	T_2	T_3
(A)	$G_2 \times SU(3) \times SO(4)$	12	6	8
(B)	$SU(3) \times SO(8)$	12	3	8
(C)	$SU(3) \times SO(7) \times SU(2)$	12	3	8
(D)	$SU(6) \times SU(2)$	12	3	4

where, as usual, T_i denotes the i^{th} twisted sector. Secondly, the number and order of the allowed Wilson lines changes, leading e.g. to the presence of a (previously disregarded) Wilson line of order six in the case of $SU(6) \times SU(2)$.

This two features have crucial effects on the matter spectrum of orbifold models. Evidently, the number of states of orbifolds on nonfactorizable lattices will be reduced compared to the number of states on the usual lattice. What is somehow more interesting is that the missing states in nonfactorizable orbifolds form vectorlike pairs.

Consider, for instance, the spectrum of the orbifold-MSSM listed in table C.2. We see that the number of exotics can be considerably reduced compactifying on nonfactorizable

6D L	attice Γ	Twist v	Conditions on Wilson lines
(A) $G_2 \times SU$	$(3) \times SO(4)$ $\frac{1}{6}$	$\frac{1}{5}(0, 1, 2, -3)$	$3A_3 \approx 2A_5 \approx 2A_6 \approx 0; A_1 \approx A_2 \approx 0, A_3 \approx A_4$
(B) SU(3)	\times SO(8) $\frac{1}{6}$	$\frac{1}{5}(0, 2, 1, -3)$	$3A_1 \approx 2A_5 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4 \approx 0, \ A_5 \approx A_6$
(C) $SU(3) \times S(3)$	$O(7) \times SU(2)$ $\frac{1}{6}$	$\frac{1}{5}(0, 2, 1, -3)$	$3A_1 \approx 2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4 \approx 0$
(D) SU(6)	$\times SU(2)$ $\frac{1}{6}$	$\frac{1}{5}(0, 2, 1, -3)$	$6A_1 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5$

Table C.1: Allowed lattices for \mathbb{Z}_6 -II orbifolds and constraints on the Wilson lines. With exception of (A), all lattices are nonfactorizable.

lattices. All nonfactorizable lattices (B,C,D) yield in the observable sector only three MSSM generations plus two pairs of Higgses. That means that unwanted quark triplets and two Higgs pairs have disappeared. However, we lose simultaneously some necessary particles, such as some SM singlets s^0 , h, used to decouple exotics, and some right-handed neutrinos n, \bar{n} , η , $\bar{\eta}$. This is rather a curiosity than a problem.

This situation is similar to what is observed between brother models of $\mathbb{Z}_2 \times \mathbb{Z}_2$ orbifolds. In that case, nonvanishing discrete torsion on orbifolds with a given torus lattice projects out some of the states living at particular fixed points, reducing thereby the effective number of fixed points. Hence, the massless spectra of models with discrete torsion on a given lattice turn out to be identical to the spectra of models without discrete torsion on a different lattice [74]. A natural question is then whether discrete torsion also relates the spectra of \mathbb{Z}_6 -II orbifolds compactified on different lattices. Unfortunately, a close inspection of the parameters of discrete torsion in \mathbb{Z}_6 -II orbifolds convinces us that discrete torsion (as introduced in ref. [74]) does not have any effect on the spectrum of the models. However, it was during long time, without clear justification, believed that discrete symmetry only affected $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds. Therefore, a very interesting question would be to try to find out whether we are still missing something.

C.2 String Selection Rules

Compactifying on nonfactorizable lattices affects more than the spectrum. The computation of Yukawa couplings in factorizable orbifold models is based on certain selection rules stated in a form that makes use of the factorizability of the lattice. Particularly, conservation of R-charge in \mathbb{Z}_6 -II factorizable orbifolds is expressed as

$$\sum_{i} R_i^1 = -1 \mod 6, \quad \sum_{i} R_i^2 = -1 \mod 3, \quad \sum_{i} R_i^3 = -1 \mod 2, \tag{C.1}$$

where the sum runs over the states of a given coupling. R_i^a corresponds to the R-charge of the i^{th} particle of the coupling on the a^{th} complex plane of the compact space.

In the literature [140, 172], it is usually stated that R-charge conservation, eq. (C.1), follows from the symmetries of the three complex planes of a factorizable orbifold. In the case of factorizable \mathbb{Z}_6 -II orbifolds, it is clear that the three independent (discrete) rotations generated by the Cartan generators H^a of the underlying SO(6) (\simeq SU(4)) Lorentz

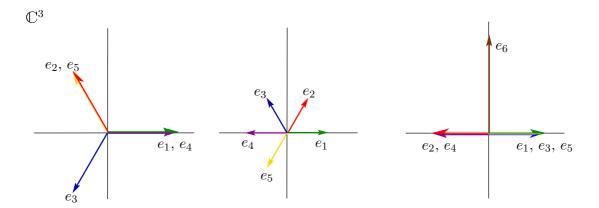


Figure C.1: The root lattice of $SU(6) \times SU(2)$ for \mathbb{Z}_6 -II orbifolds in complex coordinates.

symmetry leave the compactification lattice (A) invariant. Therefore, the original SO(6) Poincaré symmetry of the compact space is broken to $\mathbb{Z}_6 \times \mathbb{Z}_3 \times \mathbb{Z}_2$ in the factorizable orbifold.

We might assume that this is also true for compactifications on other lattices and then we could apply the same R-charge conservation given in eq. (C.1). However, as we will see below, independent rotations on each of the complex planes are not symmetries of the lattice.

C.3 (D) \mathbb{Z}_{6} -II on $SU(6) \times SU(2)$

Let us consider, as an example, the $SU(6) \times SU(2)$ lattice. The basis vectors of the root lattice of $SU(6) \times SU(2)$ are expressed in the six-dimensional orthogonal coordinates by

$$e_{1} = (1, 0, 1/\sqrt{3}, 0, \sqrt{2/3}, 0),$$

$$e_{2} = (-1/2, \sqrt{3}/2, 1/2\sqrt{3}, 1/2, -\sqrt{2/3}, 0),$$

$$e_{3} = (-1/2, -\sqrt{3}/2, -1/2\sqrt{3}, 1/2, \sqrt{2/3}, 0),$$

$$e_{4} = (1, 0, -1/\sqrt{3}, 0, -\sqrt{2/3}, 0),$$

$$e_{5} = (-1/2, \sqrt{3}/2, -1/2\sqrt{3}, -1/2, \sqrt{2/3}, 0),$$

$$e_{6} = (0, 0, 0, 0, 0, \sqrt{2}).$$
(C.2)

Notice that the twist vector in this basis is given by v = (0, 1/3, 1/6, -1/2). The lattice is illustrated in fig. C.1.

One sees that the action of $\mathbf{R}^1 \equiv e^{2\pi i \frac{1}{3}H_1} \in \mathbb{Z}_3 \subset SO(6)$ on e.g. the basis vector e_1 is

$$\mathbf{R}^{1} e_{1} = \frac{1}{2} (e_{1} + e_{2} - e_{4} + e_{5}) .$$
 (C.3)

Since the coefficients on the right-hand side are not integer, then \mathbb{R}^1 is not a symmetry of the lattice $SU(6)\times SU(2)$. One can show that the only element of SO(6) containing only

Cartan generators that leaves the lattice invariant is

$$\mathbf{R}^{t} \equiv e^{2\pi i \frac{1}{6}(2H_1 + H_2 - 3H_3)} \tag{C.4}$$

and powers of it.

We can then conjecture that the corresponding "R-charge conservation rule", in analogy to the factorizable case, is

$$\sum_{i} R_i^t \equiv \sum_{i} (2R^1 + R^2 - 3R^3)_i = 0 \mod 6.$$
 (C.5)

Nevertheless, one notices immediately one problem: charges R_i^t of bosonic and fermionic superpartners coincide, i.e. the \mathbb{Z}_6 symmetry generated by $2H_1 + H_2 - 3H_3$ is not an R-symmetry. This issue is not only interesting *per se*, but its consequences might be very important for phenomenological considerations. In particular, we have verified the effect of the "rule" eq. (C.5) on the couplings of the MSSM orbifold. We find that, given that we have less constraints in the nonfactorizable case, many more couplings are allowed.

Since the number of states in the massles spectrum is reduced, also the couplings will be restricted. This might have important consequences on both the decoupling of exotics and the appearance of dangerous proton decay operators, which are still to be studied.

			Latt	ice Γ	1
sm Repr.	Label	(A)	(B)	(C)	(D)
$(1, 2)_{(-1/2,0)}$	ϕ	4	2	2	2
$(1,2)_{(1/2,0)}$	$ar{\phi}$	4	2	2	2
$(\overline{\bf 3},{\bf 1})_{(1/3,-1/3)}$	$ar{d}$	4	3	3	3
$(3,1)_{(-1/3,1/3)}$	d	1	0	0	0
$(1, 2)_{(-1/2, -1)}$	ℓ	4	3	3	3
$(1, 2)_{(1/2,1)}$	$ar{\ell}$	1	0	0	. 0
$(3,2)_{(1/6,1/3)}$	q	3	3	3	3
$(1,1)_{(1,1)}$	$ar{e}$	3	3	3	. 3
$(\overline{\bf 3},{\bf 1})_{(-2/3,-1/3)}$	\bar{u}	3	3	3	3
$(1,1)_{(0,1)}$	$ar{n},ar{\eta}$	21	14	14	12
$(1,1)_{(0,-1)}$	n, η	18	11	11	9
$(1,1)_{(0,0)}$	h,s^0,w	98	60	60	52
$(1,1)_{(0,-2)}$	χ_1	1	1	1	1
$(1,1)_{(0,2)}$	χ_2	. 1	. 1	. 1	. 1
$(1,1)_{(0,1/2)}$	$ar{f}$	16	8	8	8
$(1,1)_{(0,-1/2)}$	f	16	8	8	8

			Latt	ice I	1
SM Repr.	Label	(A)	(B)	(C)	(D)
$(1,1)_{(1/2,1)}$	s^+, x^+	14	14	14	10
$(1,1)_{(-1/2,-1)}$	s^{-}, x^{-}	14	14	14	10
$(1, 1)_{(1/2,0)}$	s^+	8	8	8	6
$(1,1)_{(-1/2,0)}$	s^-	8	8	8	6
$(1, 1)_{(1/2,2)}$	s^+	2	2	2	2
$(1,1)_{(-1/2,-2)}$	s^-	2	2	2	2
$(3,1)_{(-1/3,-2/3)}$	δ	3	1	1	1
$(3,1)_{(1/3,2/3)}$	$ar{\delta}$	3	1	1	1
$(\overline{\bf 3},{\bf 1})_{(-1/6,2/3)}$	$ar{v}$	2	2	2	0
$(3, 1)_{(1/6, -2/3)}$	v	2	2	2	0
$(3,1)_{(-1/6,-1/3)}$	$ar{v}$	2	2	2	2
$(3,1)_{(1/6,1/3)}$	v	2	2	2	2
$(1, 2)_{(0,-1)}$	m	2	2	2	2
$(1, 2)_{(0,1)}$	m	. 2	2	2	2
$(1, 2)_{(0,0)}$	y	4	4	4	4

Table C.2: Multiplicity of the SM fields of the orbifold-MSSM in the different admissible compactification lattices of \mathbb{Z}_6 -II. The hidden-sector quantum numbers have been omitted. The labels coincide with those of table F.2. (A) denotes the factorizable lattice $G_2 \times SU(3) \times SO(4)$. Note that some pairs of exotics disappear in nonfactorizable lattices (B,C,D).

Appendix D

Orbifold Tables

P	6D Lattice Γ	Conditions on the Wilson lines
\mathbb{Z}_3	$SU(3)^3$	$3A_1 \approx 3A_3 \approx 3A_5 \approx 0$; $A_1 \approx A_2$, $A_3 \approx A_4$, $A_5 \approx A_6$
\mathbb{Z}_4	$SU(4)^2$	$4A_1 \approx 4A_4 \approx 0; \ A_1 \approx A_2 \approx A_3, \ A_4 \approx A_5 \approx A_6$
	$SU(4) \times SO(5) \times SU(2)$	$4A_1 \approx 2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_3 \approx A_4, \ A_4 \approx 0$
	$SO(5)^2 \times SU(2)^2$	$2A_2 \approx 2A_4 \approx 2A_5 \approx 2A_6 \approx 0; A_1 \approx A_3 \approx 0$
\mathbb{Z}_6 -I	$G_2^2 \times SU(3)$	$3A_5 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx 0, \ A_5 \approx A_6$
\mathbb{Z}_6 -II	$SU(6) \times SU(2)$	$6A_1 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5$
	$G_2 \times SU(3) \times SU(2)^2$	$3A_3 \approx 2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx 0, \ A_3 \approx A_4$
	$SU(3) \times SO(8)$	$3A_1 \approx 2A_5 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4 \approx 0, \ A_5 \approx A_6$
	$SU(3) \times SO(7) \times SU(2)$	$3A_1 \approx 2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4 \approx 0$
\mathbb{Z}_7	SU(7)	$7A_1 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5 \approx A_6$
\mathbb{Z}_8 -I	$SO(9) \times SO(5)$	$2A_4 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_5 \approx 0$
\mathbb{Z}_8 -II	$SO(10) \times SU(2)$	$2A_4 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx 0, \ A_4 \approx A_5$
	$SO(9) \times SU(2)^2$	$2A_4 \approx 2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx 0$
\mathbb{Z}_{12} -I	E_{6}	$3A_1 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5, \ A_6 \approx 0$
	$F_4 \times \mathrm{SU}(3)$	$3A_5 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx 0, \ A_5 \approx A_6$
\mathbb{Z}_{12} -II	$F_4 \times \mathrm{SU}(2)^2$	$2A_5 \approx 2A_6 \approx 0; \ A_1 \approx A_2 \approx A_3 \approx A_4 \approx 0$

Table D.1: Wilson lines are constrained in orbifold compactifications due to the compactification lattice Γ . The symbol ' \approx ' means equivalence up to translations in the gauge lattice Λ , thus $N_{\alpha}A_{\alpha}\approx 0$ means that $N_{\alpha}A_{\alpha}\in \Lambda$. Some frequent typos in the current literature have been corrected here.

P	6D Lattice Γ	Conditions on the Wilson lines
$\mathbb{Z}_2 \times \mathbb{Z}_2$	$SU(2)^6$	$2A_1 \approx 2A_2 \approx 2A_3 \approx 2A_4 \approx 2A_5 \approx 2A_6 \approx 0$
	$SU(3) \times SU(2)^4$	"
	$SU(3)^2SU(2)^2$	"
	$SU(4) \times SU(2)^3$	"
	$SU(4) \times SU(3) \times SU(2)$	"
	$SU(4)^2$	"
	$SU(3)^3$	"
$\mathbb{Z}_2 \times \mathbb{Z}_4$	$SU(2)^2 \times SO(5)^2$	$2A_1 \approx 2A_2 \approx 2A_4 \approx 2A_6 \approx 0; \ A_3 \approx A_5 \approx 0$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -I	$SU(2)^2 \times SU(3) \times G_2$	$2A_1 \approx 2A_2 \approx 0; \ A_3 \approx A_4 \approx A_5 \approx A_6 \approx 0$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -II	$G_2^2 \times SU(3)$	$A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5 \approx A_6 \approx 0$
$\mathbb{Z}_3 \times \mathbb{Z}_3$	$SU(3)^3$	$3A_1 \approx 3A_3 \approx 3A_5 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4, \ A_5 \approx A_6$
	E_{6}	$3A_1 \approx 3A_3 \approx 0; \ A_1 \approx A_2 \approx A_4 \approx A_5 \approx A_6$
$\mathbb{Z}_3 \times \mathbb{Z}_6$	$SU(3) \times G_2^2$	$3A_1 \approx 0; \ A_1 \approx A_2, \ A_3 \approx A_4 \approx A_5 \approx A_6 \approx 0$
$\mathbb{Z}_4 \times \mathbb{Z}_4$	$SO(5)^3$	$2A_2 \approx 2A_4 \approx 2A_6 \approx 0; \ A_1 \approx A_3 \approx A_5 \approx 0$
$\mathbb{Z}_6 \times \mathbb{Z}_6$	G_2^3	$A_1 \approx A_2 \approx A_3 \approx A_4 \approx A_5 \approx A_6 \approx 0$

Table D.2: Constraints on the Wilson lines of $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds. The notation is the same as in table D.1.

\mathbb{Z}_N or	\mathbb{Z}_N orbifolds		$\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds		
Point group P	Twist vector v	Point group P	Twist vector v_1	Twist vector v_2	
\mathbb{Z}_3	$\frac{1}{3}(0, 1, 1, -2)$	$\mathbb{Z}_2 imes \mathbb{Z}_2$	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{2}(0, 0, 1, -1)$	
\mathbb{Z}_4	$\frac{1}{4}(0, 1, 1, -2)$	$\mathbb{Z}_2 \times \mathbb{Z}_4$	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{4}(0, 0, 1, -1)$	
\mathbb{Z}_6 -I	$\frac{1}{6}(0, 1, 1, -2)$	$\mathbb{Z}_3 \times \mathbb{Z}_3$	$\frac{1}{3}(0, 1, 0, -1)$	$\frac{1}{3}(0, 0, 1, -1)$	
\mathbb{Z}_6 -II	$\frac{1}{6}(0, 1, 2, -3)$	$\mathbb{Z}_2 \times \mathbb{Z}_6$ -I	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$	
\mathbb{Z}_7	$\frac{1}{7}(0, 1, 2, -3)$	$\mathbb{Z}_2 \times \mathbb{Z}_6$ -II	$\frac{1}{2}(0, 1, 0, -1)$	$\frac{1}{6}(0, 1, 1, -2)$	
\mathbb{Z}_8 -I	$\frac{1}{8}(0, 1, 2, -3)$	$\mathbb{Z}_4 imes \mathbb{Z}_4$	$\frac{1}{4}(0, 1, 0, -1)$	$\frac{1}{4}(0, 0, 1, -1)$	
\mathbb{Z}_8 -II	$\frac{1}{8}(0, 1, 3, -4)$	$\mathbb{Z}_3 \times \mathbb{Z}_6$	$\frac{1}{3}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$	
\mathbb{Z}_{12} -I	$\frac{1}{12}(0, 1, 4, -5)$	$\mathbb{Z}_6 imes \mathbb{Z}_6$	$\frac{1}{6}(0, 1, 0, -1)$	$\frac{1}{6}(0, 0, 1, -1)$	
\mathbb{Z}_{12} -II	$\frac{1}{12}(0, 1, 5, -6)$				

Table D.3: Twist vectors for \mathbb{Z}_N and $\mathbb{Z}_N \times \mathbb{Z}_M$ orbifolds.

$V^{(\#)}$	Shift vector V	4D gauge group \mathcal{G}_{4D}	U sector	T_2 sector
0	$(0^8)(0^8)$	$E_8 \times E_8$		$27 \times 9(1, 1)$
1	$\left(\frac{2}{3}, \frac{1}{3}^2, 0^6\right) \left(0^8\right)$	$E_6 \times SU(3) \times E_8$	$3(27, \overline{3}, 1)$	$27 \times 1(27, 1, 1)$
	, , ,			$27 \times 3(1, 3, 1)$
2	$\left(\frac{2}{3}, \frac{1}{3}^2, 0^6\right) \left(\frac{2}{3}, \frac{1}{3}^2, 0^6\right)$	$E_6 \times SU(3) \times E_6 \times SU(3)$	$3(27, \overline{3}, 1, 1)$	$27 \times 1(1, 3, 1, 3)$
			$3(1, 1, 27, \overline{3})$	
3	$\left(\frac{2}{3},0^7\right)\left(\frac{1}{3}^2,0^6\right)$	$SO(14) \times E_7$	3(64, 1)	
			3(14, 1)	$27 \times 1(14, 1)$
			3(1 , 56)	$27 \times 4(1, 1)$
			3(1, 1)	
4	$\left(\frac{1}{3}^4, \frac{2}{3}, 0^3\right) \left(\frac{2}{3}, 0^7\right)$	$SU(9) \times SO(14)$	3(84, 1)	
			3(1 , 64)	$27 \times 1(\overline{9}, 1)$
			3(1, 14)	

Table D.4: Admissible shift vectors and corresponding matter content for the \mathbb{Z}_3 orbifold of the $\mathbb{E}_8 \times \mathbb{E}_8$ heterotic string.

V ^(#)	Shift vector V	$4\mathrm{D}$ gauge group \mathcal{G}_{4D}
1	$\frac{1}{4}\left(1^2,2,0^{13}\right)$	$SO(26) \times SU(2) \times U(1)_A \times U(1)$
2	$\frac{1}{4}\left(1^2,2^3,0^{11}\right)$	$SO(22) \times SU(4) \times SU(2) \times U(1)$
3	$\frac{1}{4} \left(1^6, 0^{10} \right)$	$SO(20) \times SU(6) \times U(1)_A$
4	$\frac{1}{4}\left(1^5,3,0^{10}\right)$	$SO(20) \times SU(6) \times U(1)_A$
5	$\frac{1}{4}\left(1^2,2^5,0^9\right)$	$SO(18) \times SO(10) \times SU(2) \times U(1)_A$
6	$\frac{1}{4}\left(1^6,2^2,0^8\right)$	$SO(16) \times SU(2)^2 \times SU(6) \times U(1)_A$
7	$\frac{1}{4}\left(1^2,2^7,0^7\right)$	$SO(14) \times SO(14) \times SU(2) \times U(1)$
8	$\frac{1}{4}\left(1^6,2^4,0^6\right)$	$SO(12) \times SO(8) \times SU(6) \times U(1)_A$
9	$\frac{1}{4}\left(1^{10},2,0^{5}\right)$	$SO(10) \times SU(10) \times U(1)_A \times U(1)$
10	$\frac{1}{4}\left(1^{10},2^3,0^3\right)$	$SU(4)^2 \times \times SU(10) \times U(1)$
11	$\frac{1}{4} \left(1^{14}, 0^2 \right)$	$SU(2)^2 \times SU(14) \times U(1)$
12	$\frac{1}{4}\left(1^{13},3,0^2\right)$	$SU(2)^2 \times SU(14) \times U(1)_A$
13	$\frac{1}{8} \left(1^{14}, -1, 3 \right)$	$SU(15) \times U(1)_A \times U(1)$
14	$\frac{1}{8} \left(1^{10}, -1, 3^5\right)$	$SU(11) \times SU(5) \times U(1)_A \times U(1)$
15	$\frac{1}{8} \left(1^6, -1, 3^9 \right)$	$SU(7) \times SU(9) \times U(1)_A \times U(1)$
16	$\frac{1}{8}$ (1 ² , -1, 3 ¹³)	$SU(3) \times SU(13) \times U(1)_A \times U(1)$

Table D.5: All admissible shift vectors for the \mathbb{Z}_4 orbifold of the SO(32) heterotic string. Shift vectors leading to the same four-dimensional gauge group correspond to models with different matter content. Further details are provided in our website [123].

Orbifold	Torsion ε	Shift vector V_1	Shift vector V_2
$\mathbb{Z}_2 imes \mathbb{Z}_2$	1	$\left(\frac{1}{2},0,-\frac{1}{2},0,0,0,0,0\right)$	$\left(0,\frac{1}{2},-\frac{1}{2},0,0,0,0,0\right)$
	-1	$\left(\frac{1}{2}, -1, -\frac{1}{2}, 1, 0, 0, 0, 0\right)$	$\left(1, \frac{1}{2}, -\frac{1}{2}, 0, 1, 0, 0, 0\right)$
$\mathbb{Z}_2 \times \mathbb{Z}_4$	1	$\left(0, \frac{1}{2}, -\frac{1}{2}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{4}, 0, -\frac{1}{4}, 0, 0, 0, 0, 0\right)$
	-1	$(2,\frac{1}{2},-\frac{1}{2},0,0,0,0,0)$	$\left(\frac{1}{4}, -1, -\frac{1}{4}, 1, 0, 0, 0, 0\right)$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -I	1	$\left(0, \frac{1}{2}, -\frac{1}{2}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{6}, 0, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
	-1	$(3, \frac{1}{2}, -\frac{1}{2}, 0, 1, 0, 0, 0)$	$\left(\frac{1}{6}, -1, -\frac{1}{6}, 1, 0, 0, 0, 0\right)$
$\mathbb{Z}_2 \times \mathbb{Z}_6$ -II	1	$\left(\frac{1}{2}, 0, -\frac{1}{2}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{6}, \frac{1}{6}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$
	-1	$\left(\frac{1}{2}, 3, -\frac{1}{2}, 1, 0, 0, 0, 0\right)$	$\left(-\frac{5}{6}, \frac{7}{6}, -\frac{1}{3}, 1, 1, 0, 0, 0\right)$
$\mathbb{Z}_3 \times \mathbb{Z}_3$	1	$\left(\frac{1}{3}, 0, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$	$\left(0, \frac{1}{3}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$
	$e^{2\pi i \frac{1}{3}}$	$\left(\frac{1}{3}, -1, -\frac{1}{3}, 1, 0, 0, 0, 0\right)$	$\left(1, \frac{1}{3}, -\frac{1}{3}, 0, 1, 0, 0, 0\right)$
	$e^{2\pi i \frac{2}{3}}$	$\left(\frac{1}{3}, -2, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$	$\left(2, \frac{1}{3}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$
$\mathbb{Z}_3 \times \mathbb{Z}_6$	1	$\left(0, \frac{1}{3}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{6}, 0, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
	$e^{2\pi i \frac{1}{3}}$	$\left(2, \frac{1}{3}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{6}, -1, -\frac{1}{6}, 1, 0, 0, 0, 0\right)$
	$e^{2\pi i \frac{2}{3}}$	$\left(4, \frac{1}{3}, -\frac{1}{3}, 0, 0, 0, 0, 0\right)$	$\left(\frac{1}{6}, -2, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
$\mathbb{Z}_4 \times \mathbb{Z}_4$	1	$\left(\frac{1}{4}, 0, -\frac{1}{4}, 0, 0, 0, 0, 0\right)$	$\left(0, \frac{1}{4}, -\frac{1}{4}, 0, 0, 0, 0, 0\right)$
	i	$\left(\frac{1}{4}, -1, -\frac{1}{4}, 1, 0, 0, 0, 0\right)$	$\left(1, \frac{1}{4}, -\frac{1}{4}, 0, 1, 0, 0, 0\right)$
	-1	$\left(\frac{1}{4}, -2, -\frac{1}{4}, 0, 0, 0, 0, 0\right)$	$\left(2, \frac{1}{4}, -\frac{1}{4}, 0, 0, 0, 0, 0\right)$
	-i	$\left(\frac{1}{4}, -3, -\frac{1}{4}, 1, 0, 0, 0, 0\right)$	$\left(3, \frac{1}{4}, -\frac{1}{4}, 0, 1, 0, 0, 0\right)$
$\mathbb{Z}_6 \times \mathbb{Z}_6$	1	$\left(\frac{1}{6}, 0, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$	$\left(0, \frac{1}{6}, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
	$e^{2\pi i \frac{1}{6}}$	$\left(\frac{1}{6}, -1, -\frac{1}{6}, 1, 0, 0, 0, 0\right)$	$\left(1, \frac{1}{6}, -\frac{1}{6}, 0, 1, 0, 0, 0\right)$
	$e^{2\pi i \frac{1}{3}}$	$\left(\frac{1}{6}, -2, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$	$\left(2, \frac{1}{6}, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
	-1	$\left(\frac{1}{6}, -3, -\frac{1}{6}, 1, 0, 0, 0, 0\right)$	$\left(3, \frac{1}{6}, -\frac{1}{6}, 0, 1, 0, 0, 0\right)$
	$e^{2\pi i \frac{2}{3}}$	$\left(\frac{1}{6}, -4, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$	$\left(4, \frac{1}{6}, -\frac{1}{6}, 0, 0, 0, 0, 0\right)$
	$e^{2\pi i \frac{5}{6}}$		$\left(5, \frac{1}{6}, -\frac{1}{6}, 0, 1, 0, 0, 0\right)$

Table D.6: $\mathbb{Z}_N \times \mathbb{Z}_M$ models with discrete torsion and standard embedding are equivalent to models without discrete torsion and non-standard embedding. We write the torsion phase factor as $\varepsilon = \mathrm{e}^{-2\pi\mathrm{i}\,V_2\cdot\Delta V_1}$. The components of the shifts within the second E₈ all vanish. This result also applies to orbifold models in SO(32), where the second half of the shift vector also vanishes.

Simple roots of SO(32)	Fundamental weights of SO(32)
$\alpha_1 = (1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \qquad \alpha_1^* = (1,0,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_2 = (0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \qquad \alpha_2^* = (1,1,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_3 = (0, 0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0)$	$(0,0) \qquad \alpha_3^* = (1,1,1,0,0,0,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_4 = (0,0,0,1,-1,0,0,0,0,0,0,0,0,0,0)$	$(0,0) \qquad \alpha_4^* = (1,1,1,1,0,0,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_5 = (0,0,0,0,1,-1,0,0,0,0,0,0,0,0,0)$	$(0,0) \qquad \alpha_5^* = (1,1,1,1,1,0,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_6 = (0, 0, 0, 0, 0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \qquad \alpha_6^* = (1,1,1,1,1,1,0,0,0,0,0,0,0,0,0,0,0)$
$\alpha_7 = (0, 0, 0, 0, 0, 0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \qquad \alpha_7^* = (1,1,1,1,1,1,1,0,0,0,0,0,0,0,0,0,0)$
$\alpha_8 = (0, 0, 0, 0, 0, 0, 0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \qquad \alpha_8^* = (1,1,1,1,1,1,1,1,0,0,0,0,0,0,0,0,0)$
$\alpha_9 = (0,0,0,0,0,0,0,1,-1,0,0,0,0)$	$(0,0) \qquad \alpha_9^* = (1,1,1,1,1,1,1,1,1,0,0,0,0,0,0,0,0)$
$\alpha_{10} = (0, 0, 0, 0, 0, 0, 0, 0, 0, 1, -1, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,$	$(0,0) \alpha_{10}^* = (1,1,1,1,1,1,1,1,1,0,0,0,0,0,0)$
$\alpha_{11} = (0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 1, -1, 0, 0)$	$(0,0) \alpha_{11}^* = (1,1,1,1,1,1,1,1,1,1,0,0,0,0,0)$
$\alpha_{12} = (0,0,0,0,0,0,0,0,0,0,1,-1,0)$	$(0,0) \alpha_{12}^* = (1,1,1,1,1,1,1,1,1,1,1,0,0,0,0)$
$\alpha_{13} = (0,0,0,0,0,0,0,0,0,0,0,1,-1)$	$(0,0) \alpha_{13}^* = (1,1,1,1,1,1,1,1,1,1,1,1,0,0,0)$
$\alpha_{14} = (0,0,0,0,0,0,0,0,0,0,0,0,0,0,0,0,0,0,0,$	$-1,0) \qquad \alpha_{14}^* = (1,1,1,1,1,1,1,1,1,1,1,1,1,0,0)$
$\alpha_{15} = (0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0$	$(3,-1) \alpha_{15}^* = (\frac{1}{2}, \frac{1}{2}, \frac{1}{2})$
$\alpha_{16} = (0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0$	$\alpha_{16}^* = \left(\frac{1}{2}, \frac{1}{2}, \frac{1}{2}\right)$

Table D.7: Simple roots α_i and fundamental weights α_i^* of SO(32)

Simple roots of E ₈	Fundamental weights of E_8					
$\alpha_1 = (0, 1, -1, 0, 0, 0, 0, 0)$	$\alpha_1^* = (1, 1, 0, 0, 0, 0, 0, 0)$					
$\alpha_2 = (0, 0, 1, -1, 0, 0, 0, 0)$	$\alpha_2^* = (2, 1, 1, 0, 0, 0, 0, 0)$					
$\alpha_3 = (0, 0, 0, 1, -1, 0, 0, 0)$	$\alpha_3^* = (3, 1, 1, 1, 0, 0, 0, 0)$					
$\alpha_4 = (0,0,0,0,1,-1,0,0)$	$\alpha_4^* = (4, 1, 1, 1, 1, 0, 0, 0)$					
$\alpha_5 = (0,0,0,0,0,1,-1,0)$	$\alpha_5^* = (5, 1, 1, 1, 1, 1, 0, 0)$					
$\alpha_6 = (0, 0, 0, 0, 0, 0, 1, -1)$	$\alpha_6^* = (\frac{7}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, -\frac{1}{2})$					
$\alpha_7 = (\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, \frac{1}{2})$	$\alpha_7^* = (2, 0, 0, 0, 0, 0, 0, 0)$					
$\alpha_8 = (0,0,0,0,0,0,1,1)$	$\alpha_8^* = \left(\frac{5}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}\right)$					

Table D.8: Simple roots α_i and fundamental weights α_i^* of \mathbf{E}_8

Appendix E

Supersymmetric B-L Configurations

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$\begin{array}{c cccc} V & -\frac{1}{3} \\ A_3 & \frac{1}{2} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & -\frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_3 & 1 \\ A_5 & \frac{1}{2} \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_3 & 1 \\ A_5 & \frac{1}{2} \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & -\frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & -\frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_5 & 0 \\ t_Y & 0 \\ t_{B-L} & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ t_{B-L} & -1 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ A_5 & \frac{1}{4} \\ V & 0 \\ \hline V & \frac{1}{3} \\ V & 0 \\ \hline V & \frac{1}{3} \\ V & 0 \\ \hline V & \frac{1}{3} \\ V & 0 \\ \hline V & \frac{1}{3} \\ V & 0 \\ \hline V & \frac{1}{3} \\ V & 0 \\ \hline V & 0$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} 0 \\ -\frac{1}{6} \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{4}{3} \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ $	$\begin{array}{c} 0 \\ -\frac{1}{6} \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ 0 \\ 0 \\ -\frac{4}{3} \\ -1 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{4}{3} \\ -1 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ \frac{1}{2} \\ \frac{1}{2} \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ $	$\begin{array}{c} 0 \\ -\frac{1}{60} \\ -\frac{1}{32} \\ -\frac{2}{3} \\ 0 \\ -\frac{1}{60} \\ -\frac{1}{32} \\ -1$	$\begin{array}{c} 0 \\ -\frac{1}{60} \\ -\frac{1}{3223} \\ 0 \\ -\frac{1}{60} \\ -\frac{1}{3223} \\ 0 \\ -\frac{1}{3223} \\$	0 16 0 1323 0 16 0 1323 0 5312121323 0 5312121323 0 16 0 1323 0 16 0 1323 0 5312121323 0 16 0 1323 0 16 0 1323 0 16 0 16 0 1323 0 16 0 16 0 16 0 16 0 16 0 16 0 16 0 1	$\begin{array}{c} \frac{1}{2} \\ \frac{1}{3} \\ 4 \\ 0 \\ -2 \\ \hline \\ \frac{1}{2} \\ 4 \\ 0 \\ 0 \\ \hline \\ \frac{1}{2} \\ \frac{1}{2} \\ \frac{1}{3} \\ \frac{1}{2} \\ 0 \\ 0 \\ \frac{2}{3} \\ \hline \\ \frac{1}{2} \\ 0 \\ 0 \\ \frac{2}{3} \\ \hline \\ 1 \\ 0 \\ 0 \\ \end{array}$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{r} -\frac{1}{2} \\ 0 \\ -\frac{7}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ -4 \\ -\frac{7}{2} \\ 0 \\ -\frac{4}{2} \\ -\frac{1}{2} \\ 0 \\ -\frac{1}{2} \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{3}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ -\frac{3}{2} \\ 0 \\ -\frac{1}{2} \\ 0 \\ -\frac{3}{2} \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	12 0 72 0 0 12 2 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 2 0 12 1 3 3 1 3 1 2 0 12 1 3 3 1 3 1 3 0 12 1 3 3 1 3 1 3 0 12 1 3 3 1 3 1 3 1 3 1 3 1 3 1 3 1 3 1
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$\begin{array}{c cccc} t_{B-L} & -1 \\ \hline V & \frac{1}{3} \\ A_3 & -\frac{1}{2} \\ A_5 & \frac{1}{4} \\ t_Y & 0 \\ \hline t_{B-L} & -1 \\ \hline V & \frac{1}{3} \\ A_3 & -\frac{1}{2} \\ A_5 & -\frac{1}{4} \\ t_Y & 0 \\ \hline t_{B-L} & -1 \\ \hline \end{array}$	$ \begin{array}{r} -1 \\ -\frac{1}{2} \\ -\frac{1}{4} \\ 0 \\ -1 \\ -\frac{1}{2} \\ -\frac{1}{4} \\ 0 \\ -1 \end{array} $	$ \begin{array}{c} 0 \\ -\frac{1}{2} \\ \frac{1}{6} \\ -\frac{1}{4} \\ 0 \\ 0 \\ -\frac{1}{2} \\ \frac{1}{6} \\ -\frac{1}{4} \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 0 \\ -\frac{1}{6} \\ -\frac{1}{4} \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{4} \\ -\frac{1}{2} \\ 0 \\ 0 \end{array} $	$\begin{array}{c} 0 \\ -\frac{1}{6} \\ -\frac{1}{2} \\ 0 \\ 0 \\ -\frac{1}{6} \\ -\frac{1}{2} \\ 0 \\ \end{array}$	3 0 16141323 0 161413323	3 0 16144133223 0 166144133223		$\begin{array}{c} 0 \\ \frac{1}{2} \\ \frac{10}{3} \\ 1 \\ 0 \\ \frac{7}{2} \\ \frac{1}{3} \\ \frac{1}{3} \\ \frac{1}{4} \\ 0 \\ -1 \end{array}$	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{6} \\ 0 \\ -1 \\ 0 \\ -\frac{1}{2} \\ -\frac{1}{6} \\ 0 \\ -\frac{5}{4} \\ 0 \\ -\frac{1}{2} \end{array} $	$\begin{array}{c} 0 \\ -\frac{1}{2} \\ -6 \\ -\frac{5}{2} \\ 0 \\ 0 \\ -\frac{1}{2} \\ 0 \\ -\frac{3}{4} \\ 0 \\ 0 \\ \end{array}$	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{7}{3} \\ -\frac{3}{2} \\ 0 \\ \frac{3}{2} \\ -\frac{1}{2} \\ \frac{2}{3} \\ -\frac{5}{4} \\ 0 \\ 1 \end{array} $	$ \begin{array}{r} 0 \\ -\frac{1}{2} \\ -\frac{4}{3} \\ -\frac{1}{2} \\ 0 \\ -\frac{1}{2} \\ 0 \\ -\frac{5}{4} \\ 0 \\ 0 \end{array} $	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{2} \\ -5 \\ 0 \\ 0 \\ \hline -\frac{1}{2} \\ 0 \\ \hline -\frac{1}{2} \\ 0 \\ \hline 0 \\ -\frac{5}{4} \\ 0 \\ \hline 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ $	$ \begin{array}{r} \hline & -\frac{1}{2} \\ & -3 \\ & -\frac{3}{2} \\ & 0 \\ & 0 \\ & -\frac{1}{24} \\ & \frac{1}{3} \\ & \frac{1}{4} \\ & 0 \\ & -\frac{5}{2} \end{array} $	3 3 3 2 0 0 1 2 5 9 4 0 0 0

Table E.1: Shift vectors, Wilson lines, hypercharge generator and B-L generator of 15 B-L configurations of the MSSM candidates based on the shift vectors with local SO(10) structure. All properties of the corresponding models can be computed through this data.

Appendix F

An Orbifold-MSSM: Details

In this appendix we display in full detail the most important properties of the so-called orbifold-MSSM introduced in chapter 5. This model has been subject of analysis also in our previous works, refs. [75, 77, 78].

F.1 Model Definitions and Spectrum

The model is defined by its gauge embedding, i.e. shift and Wilson lines [75, 77]

$$V^{\text{SO}(10),1} = \left(\frac{1}{3}, -\frac{1}{2}, -\frac{1}{2}, 0, 0, 0, 0, 0\right) \left(\frac{1}{2}, -\frac{1}{6}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, -\frac{1}{2}, \frac{1}{2}\right), \quad (\text{F.1a})$$

$$A_3 = \left(-\frac{1}{2}, -\frac{1}{2}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}, \frac{1}{6}\right) \left(\frac{10}{3}, 0, -6, -\frac{7}{3}, -\frac{4}{3}, -5, -3, 3\right),$$
 (F.1b)

$$A_5 = \left(\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, -\frac{1}{4}, \frac{1}{4}, \frac{1}{4}, \frac{1}{4}, \frac{1}{4}\right) \left(1, -1, -\frac{5}{2}, -\frac{3}{2}, -\frac{1}{2}, -\frac{5}{2}, -\frac{3}{2}, \frac{3}{2}\right) . (F.1c)$$

The unbroken gauge group after compactification is

$$\mathcal{G}_{4D} = [SU(3) \times SU(2)] \times [SO(8) \times SU(2)] \times U(1)^{8}.$$
 (F.2)

The U(1) generators are chosen to be

$$t_1 = t_Y = (0, 0, 0, -\frac{1}{2}, -\frac{1}{2}, \frac{1}{3}, \frac{1}{3}, \frac{1}{3}) (0, 0, 0, 0, 0, 0, 0, 0, 0),$$
 (F.3a)

$$\mathsf{t}_2 \ = \ (1,0,0,0,0,0,0,0) \, (0,0,0,0,0,0,0,0) \,, \tag{F.3b}$$

$$t_3 = (0, 1, 0, 0, 0, 0, 0, 0) (0, 0, 0, 0, 0, 0, 0, 0),$$
 (F.3c)

$$t_4 \ = \ (0,0,1,0,0,0,0,0) \, (0,0,0,0,0,0,0,0) \, , \tag{F.3d} \label{eq:t4}$$

$$t_5 = (0, 0, 0, 1, 1, 1, 1, 1) (0, 0, 0, 0, 0, 0, 0, 0),$$
 (F.3e)

$$\mathsf{t}_6 = (0,0,0,0,0,0,0,0) \, (1,0,0,0,0,0,0,0) \,, \tag{F.3f}$$

$$\mathsf{t}_7 \ = \ (0,0,0,0,0,0,0,0) \, (0,1,0,0,0,0,0,0) \,, \tag{F.3g}$$

$$\mathsf{t}_8 \ = \ (0,0,0,0,0,0,0,0) \, (0,0,0,1,1,0,0,0) \, . \tag{F.3h}$$

The anomalous U(1) is generated by

$$t_A = \sum c_i t_i$$
, where $c_i = \left(0, \frac{7}{3}, -1, -\frac{5}{3}, \frac{1}{3}, \frac{2}{3}, -\frac{2}{3}, -\frac{2}{3}\right)$. (F.4)

The sum of anomalous charges is

$$\operatorname{tr} \mathsf{t}_A = \frac{170}{3} > 0.$$
 (F.5)

The model allows us to define a suitable B-L generator,

$$\mathbf{t}_{B-L} = \left(-1, -1, 0, 0, 0, \frac{2}{3}, \frac{2}{3}, \frac{2}{3}\right), \left(-\frac{1}{2}, -\frac{1}{2}, 0, -\frac{1}{2}, -\frac{1}{2}, 0, 0, 0\right). \tag{F.6}$$

with two important properties (cf. table F.1):

- the spectrum includes the families of quarks and leptons plus vectorlike exotics with respect to $G_{SM} \times \mathrm{U}(1)_{B-L}$, and
- there are SM singlets with B-L charge ± 2 , labeled as χ_i .

#	Representation	Label
3	$(3,2;1,1)_{(1/6,1/3)}$	q_i
3	$({f 1},{f 1};{f 1},{f 1})_{(1,1)}$	\bar{e}_i
4	$ig(\overline{f 3},{f 1};{f 1},{f 1}ig)_{(1/3,-1/3)}$	$ar{d}_i$
4	$({f 1},{f 2};{f 1},{f 1})_{(-1/2,-1)}$	ℓ_i
4	$({f 1},{f 2};{f 1},{f 1})_{(-1/2,0)}$	ϕ_i
3	$ig(\overline{f 3},{f 1};{f 1},{f 1}ig)_{(1/3,2/3)}$	$ar{\delta}_i$
20	$({f 1},{f 1};{f 1},{f 1})_{(1/2,*)}$	s_i^+
15	$({f 1},{f 1};{f 1},{f 1})_{(0,1)}$	\bar{n}_i
3	$(1,1;1,2)_{(0,1)}$	$ar{\eta}_i$
20	$({f 1},{f 1};{f 1},{f 2})_{(0,0)}$	h_i
2	$({f 1},{f 1};{f 1},{f 2})_{(1/2,1)}$	x_i^+
2	$({f 1},{f 1};{f 1},{f 1})_{(0,\pm 2)}$	χ_i
4	$ig(\overline{f 3},{f 1};{f 1},{f 1}ig)_{(-1/6,*)}$	$ar{v}_i$
2	$({f 1},{f 1};{f 8},{f 1})_{(0,-1/2)}$	f_i
5	$({f 1},{f 1};{f 8},{f 1})_{(0,0)}$	w_i

#	Representation	Label
3	$ig(\overline{f 3},{f 1};{f 1},{f 1}ig)_{(-2/3,-1/3)}$	$ar{u}_i$
4	$({f 1},{f 2};{f 1},{f 1})_{(0,*)}$	m_i
1	$(3,1;1,1)_{(-1/3,1/3)}$	d_i
1	$({f 1},{f 2};{f 1},{f 1})_{(1/2,1)}$	$ar{\ell}_i$
4	$({f 1},{f 2};{f 1},{f 1})_{(1/2,0)}$	$ar{\phi}_i$
3	$(3,1;1,1)_{(-1/3,-2/3)}$	δ_i
20	$({f 1},{f 1};{f 1},{f 1})_{(-1/2,*)}$	s_i^-
12	$({f 1},{f 1};{f 1},{f 1})_{(0,-1)}$	n_i
3	$({f 1},{f 1};{f 1},{f 2})_{(0,-1)}$	η_i
2	$({f 1},{f 2};{f 1},{f 2})_{(0,0)}$	y_i
2	$({f 1},{f 1};{f 1},{f 2})_{(-1/2,-1)}$	x_i^-
18	$({f 1},{f 1};{f 1},{f 1})_{(0,0)}$	s_i^0
4	$(3,1;1,1)_{(1/6,*)}$	v_i
2	$({f 1},{f 1};{f 8},{f 1})_{(0,1/2)}$	$ar{f_i}$

Table F.1: Spectrum summary. The quantum numbers under $SU(3) \times SU(2) \times [SO(8) \times SU(2)]$ are shown in boldface; hypercharge and B-L charge appear as subscripts. Note that the states s_i^{\pm} , m_i and v_i have different B-L charges for different i, which we do not explicitly list. Further details are given in section F.7.

F.2 MSSM Configuration with *R*-Parity

Consider a vacuum configuration where the fields

$$\begin{cases}
\widetilde{s}_{i} \} = \begin{cases}
\chi_{1}, \chi_{2}, s_{3}^{0}, s_{5}^{0}, s_{8}^{0}, s_{9}^{0}, s_{12}^{0}, s_{15}^{0}, s_{16}^{0}, s_{22}^{0}, s_{24}^{0}, s_{35}^{0}, s_{41}^{0}, s_{43}^{0}, s_{46}^{0}, h_{2}, \\
h_{3}, h_{5}, h_{9}, h_{13}, h_{14}, h_{20}, h_{21}, h_{22} \end{cases}$$
(F.7)

develop a nonzero VEV while the expectation values of all other fields vanish. The emerging effective theory has the following properties:

1. the unbroken gauge symmetries are

$$G_{SM} \times G_{\text{hidden}}$$
 with $G_{\text{hidden}} = SO(8)$. (F.8)

- 2. since B-L is broken by two units, there is an effective matter parity $\mathbb{Z}_2^{\mathcal{M}}$.
- 3. the Higgs mass terms are

$$\bar{\phi}_{i} (\mathcal{M}_{\bar{\phi}\phi})_{ij} \phi_{j} \quad \text{where} \quad \mathcal{M}_{\bar{\phi}\phi} = \begin{pmatrix} \widetilde{s}^{4} & 0 & 0 & \widetilde{s} \\ \widetilde{s} & \widetilde{s}^{3} & \widetilde{s}^{3} & \widetilde{s}^{6} \\ \widetilde{s}^{5} & 0 & 0 & \widetilde{s}^{3} \\ \widetilde{s} & 0 & 0 & \widetilde{s}^{3} \end{pmatrix} . \tag{F.9}$$

The up-type Higgs h_u is a linear combination of $\bar{\phi}_1$, $\bar{\phi}_3$ and $\bar{\phi}_4$,

$$h_u \sim \tilde{s}^2 \bar{\phi}_1 + \bar{\phi}_3 + \tilde{s}^4 \bar{\phi}_4 ,$$
 (F.10)

while the down-type Higgs is composed out of ϕ_2 and ϕ_3 ,

$$h_d \sim \phi_2 + \phi_3$$
 (F.11)

The vacuum configuration is such that the μ -term, being defined as the smallest eigenvalue of $\mathcal{M}_{\bar{\phi}\phi}$,

$$\mu = \frac{\partial^2 W}{\partial h_d \partial h_u} \bigg|_{h_u = h_d = 0} \tag{F.12}$$

vanishes up to order \tilde{s}^6 , at which we work.

- 4. we check that switching on $\{\tilde{s}_i\}$ -fields allows us to cancel the FI term without inducing D-terms (cf. section F.3).
- 5. all exotics decouple (cf. section F.5).
- 6. neutrino masses are suppressed via the seesaw mechanism (cf. section F.6).

Therefore, we have obtained a supersymmetric vacuum with the precise matter content of the MSSM and R-parity.

The up-Higgs Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_u)_{ij}^{(k)} q_i \, \bar{u}_j \, \bar{\phi}_k \,,$$
 (F.13)

where

$$Y_{u}^{(1)} = \begin{pmatrix} 0 & 0 & \tilde{s}^{6} \\ 0 & 0 & \tilde{s}^{6} \\ \tilde{s}^{3} & \tilde{s}^{3} & 1 \end{pmatrix}, \quad Y_{u}^{(2)} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \tilde{s}^{6} \end{pmatrix},$$

$$Y_{u}^{(3)} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \tilde{s}^{6} \end{pmatrix}, \quad Y_{u}^{(4)} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \tilde{s}^{6} \end{pmatrix}.$$

$$(F.14a)$$

Thus, the physical 3×3 up-Higgs Yukawa matrix is

$$Y_u \sim \tilde{s}^2 Y_u^{(1)} + Y_u^{(3)} + \tilde{s}^4 Y_u^{(4)} = \begin{pmatrix} 0 & 0 & \tilde{s}^8 \\ 0 & 0 & \tilde{s}^8 \\ \tilde{s}^5 & \tilde{s}^5 & \tilde{s}^2 \end{pmatrix}.$$
 (F.15)

We find that at order six in SM singlet fields, Y_u does not have maximal rank, which means that the up-quark is massless at this level. However, at order seven in SM singlets Y_d takes the form

$$Y_u \sim \tilde{s}^2 Y_u^{(1)} + Y_u^{(3)} + \tilde{s}^4 Y_u^{(4)} = \begin{pmatrix} \tilde{s}^7 & \tilde{s}^7 & \tilde{s}^8 \\ \tilde{s}^7 & \tilde{s}^7 & \tilde{s}^8 \\ \tilde{s}^5 & \tilde{s}^5 & \tilde{s}^2 \end{pmatrix},$$
 (F.16)

providing then masses for all up-type quarks.

The down-Higgs Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_d)_{ij}^{(k)} q_i \, \bar{d}_j \, \phi_k \,,$$
 (F.17)

where

$$Y_{d}^{(1)} = \begin{pmatrix} \widetilde{s}^{4} & \widetilde{s}^{4} & \widetilde{s}^{5} & \widetilde{s}^{5} \\ \widetilde{s}^{4} & \widetilde{s}^{4} & \widetilde{s}^{5} & \widetilde{s}^{5} \\ \widetilde{s}^{5} & \widetilde{s}^{5} & \widetilde{s}^{6} & \widetilde{s}^{6} \end{pmatrix}, \quad Y_{d}^{(2)} = \begin{pmatrix} 1 & \widetilde{s}^{4} & 0 & 0 \\ \widetilde{s}^{4} & 1 & 0 & 0 \\ \widetilde{s} & \widetilde{s} & 0 & 0 \end{pmatrix},$$
 (F.18a)
$$Y_{d}^{(3)} = \begin{pmatrix} 1 & \widetilde{s}^{4} & 0 & 0 \\ \widetilde{s}^{4} & 1 & 0 & 0 \\ \widetilde{s} & \widetilde{s} & 0 & 0 \end{pmatrix}, \qquad Y_{d}^{(4)} = 0 \qquad .$$

The physical 3×3 down-Higgs Yukawa matrix emerges by integrating out a pair of vectorlike d- and \bar{d} -quarks,

$$Y_d = \begin{pmatrix} 1 & \widetilde{s}^3 & 0 \\ 1 & \widetilde{s}^3 & 0 \\ \widetilde{s} & \widetilde{s}^4 & 0 \end{pmatrix}. \tag{F.19}$$

We find that at order 6 in SM singlet fields, Y_d does not have maximal rank, which means that the down–quark is massless at this level. However, at order eight in SM singlets Y_d takes the form

$$Y_d = \begin{pmatrix} 1 & \widetilde{s}^3 & 0 \\ 1 & \widetilde{s}^3 & \widetilde{s}^8 \\ \widetilde{s} & \widetilde{s}^4 & \widetilde{s}^8 \end{pmatrix}. \tag{F.20}$$

providing then masses for all down-type quarks.

The charged lepton Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_e)_{ij}^{(k)} \ell_i \, \bar{e}_j \, \phi_k \,,$$
 (F.21)

where

$$Y_{e}^{(1)} = \begin{pmatrix} \widetilde{s}^{4} & \widetilde{s}^{4} & \widetilde{s}^{5} \\ \widetilde{s}^{4} & \widetilde{s}^{4} & \widetilde{s}^{5} \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad Y_{e}^{(2)} = \begin{pmatrix} 1 & \widetilde{s}^{4} & \widetilde{s} \\ \widetilde{s}^{4} & 1 & \widetilde{s} \\ 0 & 0 & \widetilde{s}^{6} \\ 0 & 0 & \widetilde{s}^{6} \end{pmatrix}, \quad (F.22a)$$

$$Y_{e}^{(3)} = \begin{pmatrix} 1 & \widetilde{s}^{4} & \widetilde{s} \\ \widetilde{s}^{4} & 1 & \widetilde{s} \\ 0 & 0 & \widetilde{s}^{6} \\ 0 & 0 & \widetilde{s}^{6} \\ 0 & 0 & \widetilde{s}^{6} \end{pmatrix}, \quad Y_{e}^{(4)} = \begin{pmatrix} 0 & 0 & \widetilde{s}^{5} \\ 0 & 0 & \widetilde{s}^{5} \\ 0 & 0 & \widetilde{s}^{6} \\ 0 & 0 & \widetilde{s}^{6} \end{pmatrix} .$$

The physical 3×3 matrix emerges by integrating out a pair of vectorlike ℓ - and $\bar{\ell}$ -leptons,

$$Y_e = \begin{pmatrix} 1 & 1 & \widetilde{s} \\ \widetilde{s} & \widetilde{s} & \widetilde{s}^2 \\ 0 & 0 & \widetilde{s}^6 \end{pmatrix}. \tag{F.23}$$

F.3 *D*-Flatness

One can write down gauge invariant monomials which carry net negative anomalous charge. An example for such a monomial involving all \widetilde{s}_i is

$$I(\widetilde{s}_{i}) = \chi_{1} \chi_{2} (s_{3}^{0})^{3} \begin{pmatrix} s_{5}^{0} \\ s_{9}^{0} \\ s_{12}^{0} \\ s_{16}^{0} \end{pmatrix}^{3} \begin{pmatrix} s_{8}^{0} \\ s_{15}^{0} \end{pmatrix} \begin{pmatrix} s_{22}^{0} \\ s_{24}^{0} \end{pmatrix} (s_{35}^{0})^{2} (s_{41}^{0})^{3} (s_{43}^{0})^{4} (s_{46}^{0})^{3}$$

$$\times h_{2}^{4} h_{3} h_{5}^{5} h_{9}^{2} h_{13}^{2} h_{14}^{2} h_{20} h_{21}^{3} h_{22}^{6} .$$
(F.24)

Here, gauge equivalent expressions are arranged vertically, e.g. s_{22}^0 and s_{24}^0 carry the same charges (cf. table F.2). The monomial carries anomalous charge $\sum_i q_i^A = -52/3$.

F.4 F-Flatness

Provided the superpotential at order six

$$W = s_{32}^{0}h_{5}(s_{5}^{0}h_{1} + s_{12}^{0}h_{2}) + (s_{15}^{0}h_{15} + s_{8}^{0}h_{13})(s_{42}^{0} + s_{43}^{0})(h_{23} + h_{25}) + (s_{22}^{0}h_{14} + s_{24}^{0}h_{16})(s_{42}^{0} + s_{43}^{0})(h_{18} + h_{20}) + h_{22}(s_{5}^{0}h_{3} + s_{12}^{0}h_{4})\left(s_{41}^{0}(s_{26}^{0} + s_{28}^{0}) + s_{32}^{0}(s_{42}^{0} + s_{43}^{0}) + s_{35}^{0}(s_{45}^{0} + s_{46}^{0})\right),$$
(F.25)

the F-terms of this model are

$$F_{s_{2}^{0}} = h_{1}h_{5}s_{32}^{0} + h_{22}h_{3}((s_{26}^{0} + s_{28}^{0})s_{41}^{0} + s_{32}^{0}(s_{42}^{0} + s_{43}^{0}) + s_{35}^{0}(s_{45}^{0} + s_{46}^{0})),$$
 (F.26a)

$$F_{s_{3}^{0}} = h_{13}(h_{23} + h_{25})(s_{42}^{0} + s_{43}^{0}),$$
 (F.26b)

$$F_{s_{12}^0} = h_2 h_5 s_{32}^0,$$
 (F.26c)

$$F_{s_{15}^0} = h_{15}(h_{23} + h_{25})(s_{42}^0 + s_{43}^0), (F.26d)$$

$$F_{s_{22}^0} = h_{14}(h_{18} + h_{20})(s_{42}^0 + s_{43}^0),$$
 (F.26e)

$$F_{s_{24}^0} = h_{16}(h_{18} + h_{20})(s_{42}^0 + s_{43}^0), (F.26f)$$

$$F_{s_{26}^0} = h_{22}h_3s_{41}^0s_5^0, (F.26g)$$

$$F_{s_{28}^0} = h_{22}h_3s_{41}^0s_5^0, (F.26h)$$

$$F_{s_{32}^0} = (h_{22}h_3(s_{42}^0 + s_{43}^0)s_5^0 + h_5(h_2s_{12}^0 + h_1s_5^0)),$$
 (F.26i)

$$F_{s_{35}^0} = h_{22}h_3(s_{45}^0 + s_{46}^0)s_5^0, (F.26j)$$

$$F_{s_{21}^0} = h_{22}h_3(s_{26}^0 + s_{28}^0)s_5^0, (F.26k)$$

$$F_{s_{12}^0} = (h_{18} + h_{20})(h_{14}s_{22}^0 + h_{16}s_{24}^0) + h_{22}h_3s_{32}^0s_5^0 + (h_{23} + h_{25})(h_{15}s_{15}^0 + h_{13}s_8^0), (F.261)$$

$$F_{s_{43}^0} = (h_{18} + h_{20})(h_{14}s_{22}^0 + h_{16}s_{24}^0) + h_{22}h_3s_{32}^0s_5^0 + (h_{23} + h_{25})(h_{15}s_{15}^0 + h_{13}s_8^0) (F.26m)$$

$$F_{s_{05}^0} = h_{22}h_3s_{35}^0s_5^0, (F.26n)$$

$$F_{s_{46}^0} = h_{22}h_3s_{35}^0s_5^0, (F.260)$$

$$F_{h_1} = h_5 s_{32}^0 s_5^0,$$
 (F.26p)

$$F_{h_2} = h_5 s_{12}^0 s_{32}^0 , (F.26q)$$

$$F_{h_3} = h_{22}((s_{26}^0 + s_{28}^0)s_{41}^0 + s_{32}^0(s_{42}^0 + s_{43}^0) + s_{35}^0(s_{45}^0 + s_{46}^0))s_5^0,$$
 (F.26r)

$$F_{h_5} = s_{32}^0 (h_2 s_{12}^0 + h_1 s_5^0),$$
 (F.26s)

$$F_{h_{13}} = (h_{23} + h_{25})(s_{42}^0 + s_{43}^0)s_8^0, (F.26t)$$

$$F_{h_{14}} = (h_{18} + h_{20})s_{22}^{0}(s_{42}^{0} + s_{43}^{0}), (F.26u)$$

$$F_{h_{15}} = (h_{23} + h_{25})s_{15}^{0}(s_{42}^{0} + s_{43}^{0}),$$
 (F.26v)

$$F_{h_{16}} = (h_{18} + h_{20})s_{24}^{0}(s_{42}^{0} + s_{43}^{0}), (F.26w)$$

$$F_{h_{18}} = (h_{14}s_{22}^0 + h_{16}s_{24}^0)(s_{42}^0 + s_{43}^0), (F.26x)$$

$$F_{h_{20}} = (h_{14}s_{22}^0 + h_{16}s_{24}^0)(s_{42}^0 + s_{43}^0), (F.26y)$$

$$F_{h_{22}} = h_3((s_{26}^0 + s_{28}^0)s_{41}^0 + s_{32}^0(s_{42}^0 + s_{43}^0) + s_{35}^0(s_{45}^0 + s_{46}^0))s_5^0,$$
 (F.26z)

$$F_{h_{23}} = (s_{42}^0 + s_{43}^0)(h_{15}s_{15}^0 + h_{13}s_8^0),$$

$$F_{h_{25}} = (s_{42}^0 + s_{43}^0)(h_{15}s_{15}^0 + h_{13}s_8^0).$$

At order six in the SM singlets, the vaccum configuration (F.7) leaves some nonzero Fterms which can only be cancelled if some singlets have trivial VEVs. However, as discussed in section 5.2.2, at order eight in the superpotential, one does find nontrivial solutions. We do not list them here due to their length.

F.5Mass Matrices

In this section, we show that that all exotics can be made massive. The exotics' mass terms are

$$x_i \mathcal{M}_{x\bar{x}}^{ij} \bar{x}_j$$
 (F.27)

In the following, we list the structure of the corresponding mass matrices.

$$\mathcal{M}_{\bar{\ell}\ell} = \begin{pmatrix} \widetilde{s}^2 & \widetilde{s}^2 & \widetilde{s}^3 & \widetilde{s}^3 \end{pmatrix}, \quad (F.28a) \mathcal{M}_{d\bar{d}} = \begin{pmatrix} \widetilde{s}^6 & \widetilde{s}^6 & \widetilde{s}^3 & \widetilde{s}^3 \end{pmatrix}, \quad (F.28b)$$

$$\mathcal{M}_{x^+x^-} = \begin{pmatrix} \widetilde{s}^5 & \widetilde{s}^5 \\ \widetilde{s}^5 & \widetilde{s}^5 \end{pmatrix}, \quad (F.28c)$$

$$\mathcal{M}_{mm} = \begin{pmatrix} 0 & 0 & \tilde{s}^{6} & \tilde{s}^{6} \\ 0 & 0 & \tilde{s}^{6} & \tilde{s}^{6} \\ \tilde{s}^{6} & \tilde{s}^{6} & 0 & \tilde{s}^{6} \\ \tilde{s}^{6} & \tilde{s}^{6} & \tilde{s}^{6} & 0 \end{pmatrix}, (F.28d) \qquad \mathcal{M}_{\delta\bar{\delta}} = \begin{pmatrix} \tilde{s}^{3} & \tilde{s}^{3} & \tilde{s}^{3} \\ \tilde{s}^{3} & \tilde{s}^{3} & \tilde{s}^{3} \\ 0 & \tilde{s}^{3} & \tilde{s}^{3} \end{pmatrix}, (F.28g)$$

$$\mathcal{M}_{v\bar{v}} = \begin{pmatrix} \tilde{s} & \tilde{s}^{5} & 0 & 0 \\ \tilde{s}^{5} & \tilde{s} & 0 & 0 \\ 0 & 0 & \tilde{s}^{5} & \tilde{s}^{5} \\ 0 & 0 & \tilde{s}^{5} & \tilde{s}^{5} \end{pmatrix}, (F.28e) \qquad \mathcal{M}_{ww} = \begin{pmatrix} \tilde{s} & \tilde{s}^{5} & 0 & \tilde{s}^{5} & \tilde{s}^{5} \\ \tilde{s}^{5} & \tilde{s} & 0 & \tilde{s}^{5} & \tilde{s}^{5} \\ 0 & 0 & 0 & \tilde{s}^{3} & \tilde{s}^{3} \end{pmatrix}, (F.28i)$$

$$\mathcal{M}_{ww} = \begin{pmatrix} \tilde{s} & \tilde{s}^{5} & 0 & \tilde{s}^{5} & \tilde{s}^{5} \\ \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} \\ \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{6} & \tilde{s}^{6} \\ \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{5} & \tilde{s}^{6} \end{pmatrix}, (F.28i)$$

$$\begin{pmatrix}
\widetilde{s} & \widetilde{s}^5 & \widetilde{s} & 0 & 0 \\
\widetilde{s}^5 & \widetilde{s} & 0 & 0
\end{pmatrix}$$

$$\mathcal{M}_{f\bar{f}} = \begin{pmatrix}
0 & \widetilde{s}^3 \\
0 & \widetilde{s}^3
\end{pmatrix}, \qquad (F.28h)$$

$$\mathcal{M}_{yy} = \begin{pmatrix} 0 & 0 & \widetilde{s}^5 & \widetilde{s}^5 \\ 0 & 0 & \widetilde{s}^5 & \widetilde{s}^5 \end{pmatrix}, (F.28i)$$

$$\mathcal{M}_{ww} = \begin{pmatrix} \widetilde{s}^1 & \widetilde{s}^5 \\ \widetilde{s}^5 & \widetilde{s}^5 & \widetilde{s}^3 & \widetilde{s}^6 & \widetilde{s}^6 \\ \widetilde{s}^5 & \widetilde{s}^5 & \widetilde{s}^3 & \widetilde{s}^6 & \widetilde{s}^6 \\ \widetilde{s}^5 & \widetilde{s}^5 & \widetilde{s}^3 & \widetilde{s}^6 & \widetilde{s}^6 \end{pmatrix}, (F.28i)$$

F.6 Neutrino Masses

We consider vacua where SU(2) is broken. This means that the η_i and $\bar{\eta}_i$ give rise to further SM singlets with $q_{B-L} = \pm 1$,

$$\bar{\eta}_1 = \begin{pmatrix} \bar{n}_{16} \\ \bar{n}_{17} \end{pmatrix}, \dots \bar{\eta}_3 = \begin{pmatrix} \bar{n}_{20} \\ \bar{n}_{21} \end{pmatrix} \quad \text{and} \quad \eta_1 = \begin{pmatrix} n_{13} \\ n_{14} \end{pmatrix}, \dots \eta_3 = \begin{pmatrix} n_{17} \\ n_{18} \end{pmatrix}.$$
 (F.29)

The neutrino mass matrix is then given by

$$\mathcal{M}_{\bar{\nu}\bar{\nu}} = \begin{pmatrix} \mathcal{M}_{\bar{n}\bar{n}} & \mathcal{M}_{n\bar{n}}^T \\ \mathcal{M}_{n\bar{n}} & \mathcal{M}_{nn} \end{pmatrix}, \tag{F.31}$$

and has full rank.

The neutrino Yukawa couplings decompose into

$$W_{\text{Yukawa}} \supset \sum_{k=1}^{4} (Y_n)_{ij}^{(k)} \ell_i \, n_j \, \bar{\phi}_k + (Y_{\bar{n}})_{ij}^{(k)} \, \ell_i \, \bar{n}_j \, \bar{\phi}_k$$
 (F.32)

where

$$Y_n^{(k>2)} = 0$$
, (F.33c)

$$Y_{\bar{n}}^{(1)} = \begin{pmatrix} 0 & \widetilde{s}^5 & 0 & 0 & \widetilde{s}^4 & 0 & 0 & 0 & \widetilde{s}^6 & \widetilde{s}^6 & \widetilde{s}^5 & \widetilde{s}^5 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^4 & 0 & 0 & \widetilde{s}^5 & 0 & 0 & 0 & \widetilde{s}^6 & \widetilde{s}^6 & \widetilde{s}^5 & \widetilde{s}^5 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^5 & 0 & 0 & \widetilde{s}^5 & 0 & 0 & 0 & 0 & 0 & 1 & \widetilde{s}^4 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^5 & 0 & 0 & \widetilde{s}^5 & 0 & 0 & 0 & 0 & 0 & \widetilde{s}^4 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \end{pmatrix}$$
 (F.33d)

$$Y_{\bar{n}}^{(3)} = \begin{pmatrix} 0 & \widetilde{s}^6 & 0 & 0 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & 0 & \widetilde{s} & \widetilde{s} & 0 & 0 & 0 & \widetilde{s}^5 & \widetilde{s}^5 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^6 & 0 & 0 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & 0 & \widetilde{s} & \widetilde{s} & 0 & 0 & 0 & \widetilde{s}^5 & \widetilde{s}^5 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^2 & \widetilde{s}^6 & 0 & \widetilde{s}^2 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & \widetilde{s}^6 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \widetilde{s}^2 & \widetilde{s}^6 & 0 & \widetilde{s}^2 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & \widetilde{s}^6 & \widetilde{s}^6 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \end{pmatrix}$$
(F.33f)

With these couplings, one can calculate the effective ℓ bilinear

$$\mathcal{M}_{\ell\ell}^{\text{eff}} = \frac{1}{2} Y_{\nu} \mathcal{M}_{\bar{\nu}\bar{\nu}}^{-1} Y_{\nu}^{T} , \qquad (F.34)$$

where $Y_{\nu} = (Y_n, Y_{\bar{n}})$. By integrating out the heavy ℓ , one arrives at the 3×3 effective neutrino mass operator κ , which is related to the light neutrino mass matrix via

$$m_{\nu} = v_u^2 \, \kappa \tag{F.35}$$

with v_u being the up-type Higgs VEV. By using the method explained in section 5.5.1, we find that m_{ν} is (very roughly) given by

$$M_{\text{eff}} \sim -\frac{v_u^2}{M_*} \begin{pmatrix} 1 & s & s \\ s & s^2 & s^2 \\ s & s^2 & s^2 \end{pmatrix} ,$$
 (F.36)

and the effective seesaw scale is given by

$$M_* \sim 0.1 \, s^5 \, M_{\rm str} \,,$$
 (F.37)

where $M_{\rm str} = 2 \cdot 10^{17} \, {\rm GeV}$ is the string scale taken as the overall scale of $\mathcal{M}_{\bar{\nu}\bar{\nu}}$.

F.7 Detailed Spectrum

In this section, we display the properties of each state in the spectrum. The spectrum listed here differs only aesthetically with respect to the results presented in ref. [78]. The reason being that the twist vector we have used in that work and here differ by a minus sign.

Table F.2: The spectrum of the orbifold-MSSM in terms of *left-chiral* states. The U(1) charges refer to the basis of generators, eq. (F.3). R_i denote bosonic R-charges. $\gamma(\vartheta)$ is the so-called γ -phase defined in sec. 2.5. $T_{k(n_1,n_2,n_3,n_4,n_5,n_6)}$ denotes the k-th twisted sector corresponding to the constructing element $g = (\vartheta^k, n_\alpha e_\alpha)$.

Sector	Representation	q_Y	q_1	q_2	q_3	q_4	q_5	q_6	q_7	q_{B-L}	R_1	R_2	R_3	$\gamma(\vartheta)$	
U_1	(1, 1, 1, 1)	0	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	$\frac{5}{2}$	0	0	0	2	-1	0	0	0	χ_2
	$({f 1},{f 1},{f 1},{f 1})$	0	1	1	0	0	0	0	0	-2	-1	0	0	0	χ_1
	$({f 3},{f 2},{f 1},{f 1})$	$\frac{1}{6}$	$-\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	0	0	0	$\frac{1}{3}$	-1	0	0	0	q_3
	(1, 1, 8, 1)	0	0	0	0	0	$\frac{1}{2}$	$-\frac{1}{2}$	-1	$\frac{1}{2}$	-1	0	0	0	\bar{f}_1
U_2	(1, 1, 1, 1)	1	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	0	0	0	1	0	-1	0	0	\bar{e}_3
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$-\frac{2}{3}$	$\begin{array}{c} \frac{1}{2} \\ \frac{1}{2} \\ 1 \end{array}$	$-\frac{1}{2} \\ -\frac{1}{2} \\ \frac{1}{2}$	$-\frac{1}{2} \\ -\frac{1}{2} \\ \frac{1}{2}$	$\frac{\frac{1}{2}}{\frac{1}{2}}$	0	0	0	$-\frac{1}{3}$	0	-1	0	0	\bar{u}_3
	$({f 1},{f 1},{f 1},{f 1})$	0	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{5}{2}$	0	0	0	0	0	-1	0	0	s_3^0
U_3	(1, 2, 1, 1)	$-\frac{1}{2}$	0	0	-1	1	0	0	0	0	0	0	-1	0	ϕ_1
	$({f 1},{f 2},{f 1},{f 1})$	$\frac{1}{2}$	0	0	1	-1	0	0	0	0	0	0	-1	0	$ar{\phi}_1$
$T_{2(0,0,0,0,0,0)}$	$({f 1},{f 2},{f 1},{f 1})$	$-\frac{1}{2}$	$-\frac{1}{3}$	0	0	1	0	$\frac{2}{3}$	0	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	0	ϕ_2
	$({f 3},{f 1},{f 1},{f 1})$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0	-1	0	$\frac{2}{3}$	0	$-\frac{2}{3}$	$-\frac{2}{3}$	$-\frac{1}{3}$	0	0	δ_1
	$({f 1},{f 1},{f 1},{f 2})$	0	$-\frac{1}{3}$ $-\frac{1}{3}$ $\frac{1}{32}$ $\frac{32}{32}$ $\frac{32}{32}$	0	0	0	0	$-\frac{\frac{2}{3}}{\frac{2}{3}}$	-1	0	- 2 3 2 3 2 3 2 3 2 3 2 3 2 3 2 3 2 3 2	$-\frac{1}{3}$	0	0	h_{18}
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{2}{3}$	0	0	0	0	$-\frac{1}{3}$	1	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	0	η_2
	(1, 1, 8, 1)	0	$\frac{2}{3}$	0	0	0	0	$-\frac{1}{3}$	0	$-\frac{1}{2}$		$-\frac{1}{3}$	0	0	f_1
$T_{2(0,0,1,1,0,0)}$	(1, 1, 1, 1)	0	$-\frac{5}{6}$	$\frac{1}{2}$	$-\frac{1}{6}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{\frac{2}{3}}{\frac{1}{3}}$	0	$-\frac{2}{3}$ $-\frac{2}{3}$ $-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{2}{3}$	s_{42}^{0}
	$({f 1},{f 1},{f 8},{f 1})$	0	$-\frac{5}{6}$ $\frac{1}{6}$	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$-\frac{1}{6}$ $-\frac{1}{6}$ $-\frac{1}{6}$	$-\frac{5}{6} \\ -\frac{5}{6}$	$-\frac{1}{3}$ $\frac{1}{6}$	$-\frac{1}{3}$ $\frac{1}{6}$ $-\frac{1}{3}$	$-\frac{1}{3}$	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{\frac{2}{3}}{\frac{2}{3}}$	w_4
	$({f 1},{f 2},{f 1},{f 1})$	$-\frac{1}{2}$	$\frac{1}{6}$	$\frac{1}{2}$	$-\frac{1}{6}$	$\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$\frac{2}{3}$	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{3}$	ℓ_3
												co	ntinu	ιed .	

Sector	Representation	q_Y	q_1	q_2	q_3	q_4	q_5	q_6	q_7	q_{B-L}	R_1	R_2	R_3	$\gamma(\vartheta)$
$T_{2(0,0,0,1,0,0)}$	(1, 1, 1, 1)	0		0			$\frac{1}{3}$			0			0	/ (/
(0,0,0,-,0,0)	$({f 1},{f 1},{f 1},{f 2})$	0	$-\frac{1}{3}$ $\frac{1}{6}$ $\frac{1}{6}$	$-\frac{1}{2}$	$\frac{\frac{2}{3}}{\frac{1}{6}}$	$-\frac{5}{3}$	$\frac{1}{3}$	- 132 31 32 32 31 31 31 31 31 31 31 31 31 31 31 31 31	$-\frac{2}{3}$ $\frac{1}{3}$ $\frac{1}{3}$	0	$-\frac{3}{3}$	$-\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3}$	0	$\begin{array}{c c} \frac{2}{3} & s{45}^0 \\ \frac{2}{3} & h_{23} \end{array}$
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{1}{6}$	$-\frac{1}{2} \\ -\frac{1}{2}$	$\frac{1}{6}$	$\frac{5}{6}$	$-\frac{\overline{3}}{\overline{3}}$	$-\frac{1}{3}$	$\frac{1}{3}$	1	$-\frac{\frac{3}{2}}{3}$	$-\frac{1}{3}$	0	$\bar{0}$ $\bar{\eta}_2$
$T_{2(1,0,0,0,0,0)}$	(1, 2, 1, 1)	$-\frac{1}{2}$		0	0	1	0	$\frac{2}{3}$	0	0	$-\frac{2}{3}$		0	$0 \phi_3$
(, , , , , , ,	$({f 1},{f 1},{f 1},{f 1})$	0	$-\frac{1}{3}$	0	1	0	0	$\frac{2}{3}$	0	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{2}$ s_{41}^0
	$({f 3},{f 1},{f 1},{f 1})$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0	-1	0	$\frac{2}{3}$	0	$-\frac{2}{3}$	$-\frac{2}{3}$	$-\frac{1}{3}$	0	δ_2
	$({f 1},{f 1},{f 1},{f 2})$	0	$ \begin{array}{r} -\frac{1}{3} \\ -$	0	0	0	0	$-\frac{1}{3}$	-1	0	$-\frac{2}{3}$	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	0	$0 h_{20}$
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{2}{3}$	0	0	0	0	$-\frac{1}{3}$	1	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$0 \mid \eta_3$
	$({f 1},{f 1},{f 1},{f 1})$	0	$-\frac{1}{3}$	1	0	0	0	$\frac{2}{3}$	0	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{2}$ n_{11}
	(1, 1, 8, 1)	0		0	0	0	0		0	$-\frac{1}{2}$	$-\frac{2}{3}$		0	$0 f_2$
$T_{2(1,0,1,1,0,0)}$	$({f 1},{f 1},{f 1},{f 1})$	0	$-\frac{5}{6}$ $\frac{1}{6}$ $\frac{1}{1}$	$\frac{1}{2}$	$-\frac{1}{6}$	$-\frac{5}{6}$	$-\frac{1}{3}$ $-\frac{1}{3}$ $\frac{2}{3}$ 1	$-\frac{1}{3}$	$\frac{2}{3}$	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{2}{3}$ s_{43}^0
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{1}{6}$	$-\frac{1}{2}$	$-\frac{1}{6}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$\frac{2}{3}$	$-\frac{1}{3}$	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{2}$ h_{22}
	(1, 1, 1, 2)	0	$\frac{1}{6}$	$-\frac{1}{2}$	$-\frac{1}{6}$	$-\frac{5}{6}$	$\frac{2}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{5}{6}$ h_{21}
	(1, 1, 1, 1)	0	$\frac{1}{6}$	$\frac{1}{2}$	<u>5</u> 6	$-\frac{5}{6}$	$-\frac{1}{3}$	$-\frac{1}{3}$	3	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{5}{6}$ n_{12}
	(1,1,8,1)	0	6 1 6 1 6	$ \begin{array}{r} \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	$\begin{array}{c} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{3} \end{array}$	5 65 65 65 65 61 62 3 2 3 5 31 65 65 65 65 6	$-\frac{3}{3}$ $\frac{1}{6}$ $\frac{1}{1}$	$ \begin{array}{r} -\frac{1}{32} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	2 31 31 31 32 31 32 32 32 32 31 32 31 3 	0	2 50 100 2 50 2 50 2 50 2 50 2 50 2 50 2	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	0	$\begin{array}{c} 3\\ s_{43}\\ s_{43}\\ s_{43}\\ s_{44}\\ s_{42}\\ s_{56}\\ s_{65}\\ s_{6$
	(1,2,1,1)	$-\frac{1}{2} \\ -\frac{1}{3}$	$\frac{1}{6}$	$\frac{1}{2}$	$-\frac{1}{6}$	$\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$\frac{2}{3}$	-1	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{3}$ ℓ_4
	(3,1,1,1)		$-{3}$	0	3 1	3	$-\frac{1}{3}$	$-\frac{3}{3}$	3	$\frac{1}{3}$	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{2}$ d_1
$T_{2(1,0,0,1,0,0)}$	(1, 2, 1, 1)	$-\frac{1}{2}$	$-\frac{1}{3}$	0	$-\frac{1}{3}$	$-\frac{2}{3}$	$\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{2}{3}$	0	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{1}{6}$ ϕ_4
	(1,1,1,1)	0	$-\frac{1}{3}$	0		$-\frac{3}{3}$	$\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{2}{3}$	0_{2}	$-\frac{2}{3}$	$-\frac{1}{3}$	0	$\frac{2}{3}$ s_{46}^0
	(3,1,1,1)	$-\frac{1}{3}$	$\frac{\overline{6}}{1}$	$\frac{1}{2}$	$\frac{\overline{6}}{1}$	$-\frac{1}{6}$	$\frac{\overline{3}}{1}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{2}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$\begin{bmatrix} 0 \\ 0 \end{bmatrix}$	$\frac{1}{2}$ δ_3
	(1,1,1,2)	-	$\frac{\overline{6}}{5}$	$-\frac{1}{2}$	$\frac{\overline{6}}{1}$	$\frac{\overline{6}}{5}$	$\frac{3}{1}$	$\frac{\overline{3}}{1}$	$\frac{3}{2}$	1	$-\frac{3}{2}$	$-\frac{3}{3}$		$\frac{2}{3} h_{25} $
	$(1,1,1,1) \ (1,1,1,2)$	$\begin{array}{c} 0 \\ 0 \end{array}$	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ \frac{1}{6} \\ -\frac{5}{6} \\ \frac{1}{6} \end{array} $	$\frac{\overline{2}}{1}$	-\frac{1}{32}\frac{2}{31}\frac{1}{61}\frac{1}{61}\frac{1}{6}	$\frac{\overline{6}}{5}$	$-\frac{\frac{3}{2}}{3}$	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	$-\frac{3}{1}$	1	$-\frac{3}{2}$	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	$\begin{bmatrix} 0 \\ 0 \end{bmatrix}$	
T				$ \begin{array}{r} \frac{1}{2} \\ -\frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $			1		$\frac{\overline{3}}{0}$	1		$-\frac{3}{3}$, -
$T_{3(0,0,0,0,1,0)}$	$(1,1,1,1) \ (1,1,1,1)$	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$\frac{\frac{1}{4}}{\frac{3}{4}}$	$-\frac{4}{1}$	$-\frac{\frac{1}{4}}{4}$	$-\frac{\frac{1}{4}}{4}$	$\frac{2}{1}$	$-\frac{1}{2} \\ \frac{1}{2}$	0	$\begin{vmatrix} 1 \\ -1 \end{vmatrix}$	$-\frac{1}{2} \\ -\frac{1}{2}$	0	$-\frac{1}{2} \\ -\frac{1}{2}$	$\begin{array}{c c} \frac{1}{2} & s_{14}^+ \\ 0 & s_{14}^- \end{array}$
T. (2. 2. 2. 2. 1. 1.)	(1,1,1,1) (1,1,1,1)		1				$\frac{-}{2}$		0	1		0		$\begin{array}{c c} 0 & s_{14}^- \\ \hline \frac{1}{2} & s_{18}^+ \end{array}$
$T_{3(0,0,0,0,1,1)}$	(1,1,1,1) $(1,1,1,1)$	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$\frac{4}{3}$	$-\frac{3}{4} \\ -\frac{1}{4}$	$-\frac{\frac{1}{4}}{4}$	$-\frac{\frac{1}{4}}{\frac{1}{4}}$	$-\frac{2}{1}$	$-\frac{1}{2} \\ \frac{1}{2}$	0	-1	$\frac{2}{1}$	0	$-\frac{1}{2} \\ -\frac{1}{2}$	$\begin{array}{c c} \bar{2} & s_{18} \\ 0 & s_{19}^{-} \end{array}$
$T_{\alpha(\alpha,1,\alpha,\alpha,\alpha)}$	(1,1,1,1) $(1,1,1,2)$	$\frac{\overline{2}}{0}$	$\frac{\overline{4}}{0}$			$\frac{\overline{4}}{0}$	$\frac{-\frac{1}{2}}{0}$	$\frac{\overline{2}}{0}$	1	0	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	0	$\frac{2}{-1}$	
$T_{3(0,1,0,0,0,0)}$	(1,1,1,2) $(1,1,1,2)$	0	0	$\frac{2}{1}$	$-\frac{\overline{2}}{1}$	0	0	0	-1	0	$-\frac{\overline{2}}{1}$	0	$-\frac{2}{1}$	$\frac{3}{2}$ h_{13}
	(1,1,1,1)	0	0	$\frac{2}{1}$	$-\frac{2}{1}$	0	1	0	0	-1	$-\frac{2}{1}$	0	$-\frac{2}{1}$	$\frac{3}{2} \begin{vmatrix} n_{13} \\ n_{9} \end{vmatrix}$
	(1,1,1,1)	0	0	$-\frac{2}{1}$	$\frac{2}{1}$	0	-1	0	0	1	$-\frac{2}{1}$	0	$-\frac{2}{1}$	$\frac{3}{2}$ \bar{n}_{13}
$T_{3(0,1,0,0,0,1)}$	(1,1,1,2)	0	0	$ \begin{array}{r} -\frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	$ \begin{array}{r} \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ \frac{1}{2} \end{array} $	0	0	0	1	0	$-\frac{2}{1}$	0	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	$\frac{3}{2}$ h_{16}
-3(0,1,0,0,0,1)	(1,1,1,2)	0	0			0	0	0	-1	0	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{3}{2}$ h_{15}
	(1,1,1,1)	0	0	$\frac{2}{1}$	$-\frac{1}{2} \\ -\frac{1}{2}$	0	1	0	0	-1	$-\frac{\frac{2}{1}}{2}$	0	$-\frac{1}{2}$	$\frac{3}{2}$ n_{10}
	(1,1,1,1)	0	0	$-\frac{\frac{1}{2}}{\frac{1}{2}}$ $-\frac{1}{2}$	$\frac{\overline{2}}{1}$	0	-1	0	0	1	$-\frac{1}{2}$ $-\frac{1}{2}$ $-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{3}{2}$ \bar{n}_{14}
$T_{3(0,1,0,0,1,0)}$	(1,1,1,1)			$\frac{\frac{2}{1}}{4}$					0	1		0	$-\frac{1}{2}$	$\frac{1}{3}$ s_{16}^{+}
3(0,1,0,0,1,0)	(1, 1, 1, 1)	$\frac{1}{2}$	$\frac{4}{1}$	$\frac{\frac{4}{1}}{4}$	$-\frac{3}{4}$	$\frac{4}{1}$	$-\frac{1}{2}$	$\frac{\frac{2}{1}}{2}$	0	0	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{6}$ s_{13}^{+}
	(1, 1, 1, 1)	$\frac{1}{2}$	$\frac{1}{4}$	$-\frac{\frac{4}{3}}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{\frac{2}{1}}{2}$	$-\frac{1}{2}$	0	1	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{2}$ s_{15}^{+}
	(1, 1, 1, 1)	12121212121616 	$-\frac{3}{4}$ $\frac{1}{4}$ $-\frac{1}{4}$ 1	$ \begin{array}{r} \frac{1}{4} \\ \frac{1}{4} \\ -\frac{3}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$ \begin{array}{r} \frac{1}{4} \\ -\frac{3}{4} \\ \frac{1}{4} \\ \frac{3}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ \frac{1}{1} \end{array} $	$ \begin{array}{r} \frac{1}{4} \\ \frac{1}{4} \\ \frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$-\frac{\frac{1}{2}}{\frac{1}{2}}$ $-\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$	$\begin{array}{c} -\frac{1}{2} \\ -1$	0	0	$ \begin{array}{c c} -\frac{1}{2} \\ -$	0	$ \begin{array}{r} -\frac{1}{2} \\ -$	$\begin{array}{c} \frac{2}{3} & h_{14} \\ \frac{2}{3} & h_{15} \\ \frac{2}{3} & h_{15} \\ \frac{2}{3} & \frac{1}{3} \\ \frac{2}{3} & h_{16} \\ \frac{2}{3} & h_{16} \\ \frac{2}{3} & \frac{1}{3} \\ \frac{1}{3} & \frac{1}{5} \\ \frac{1}{6} & \frac{1}{5} \\ \frac{2}{3} & \frac{1}{6} \\ \frac{1}{5} & \frac{1}{6} \\ \frac{2}{5} & \frac{1}{6} \\ \frac{2}{5} & \frac{1}{6} \\ \end{array}$
	(1, 1, 1, 1)	$-\frac{1}{2}$		$\frac{\frac{1}{3}}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{2}$	$\frac{1}{2}$	0	-1	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{5}{6}$ s_{16}^{-1}
	(1, 1, 1, 1)	$-\frac{1}{2}$	$\begin{array}{c} \frac{4}{3} \\ \frac{3}{4} \\ 1 \end{array}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$ $-\frac{3}{4}$ $-\frac{3}{4}$	$-\frac{1}{2}$ $-\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$	$\frac{1}{2}$	0	-1	$-\frac{1}{2}$	0	$-\frac{1}{2}$	0 s_{15}^-
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$-\frac{1}{6}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$\frac{3}{4}$	$\frac{1}{2}$	$-\frac{1}{2}$	0	$\frac{2}{3}$	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{2}{3}$ \bar{v}_3
	(3, 1, 1, 1)	$\frac{1}{6}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$-\frac{3}{4}$	$-\frac{1}{2}$		0	$\begin{array}{c} \frac{2}{3} \\ -\frac{2}{3} \end{array}$	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\begin{array}{c c} \frac{2}{3} & \bar{v}_3 \\ \frac{2}{3} & v_3 \end{array}$
$T_{3(0,1,0,0,1,1)}$	(1, 1, 1, 1)	$\frac{1}{2}$	$-\frac{3}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{2}$	$-\frac{1}{2}$	0	1	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{3}$ s_{20}^{+}
	-				<u> </u>					-		co	ntinu	$ed \dots$

Sector	Representation	q_Y	q_1	q_2	q_3	q_4	q_5	q_6	q_7	q_{B-L}	R_1	R_2	R_3	$\gamma(\vartheta)$	
	(1, 1, 1, 1)	$\frac{1}{2}$			$ \begin{array}{r} -\frac{3}{4} \\ \frac{1}{4} \\ \frac{3}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $				0	0	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{6}$	s_{17}^{+} s_{19}^{+}
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2}$ $-\frac{1}{2}$ $-\frac{1}{2}$ $-\frac{1}{2}$	$\frac{1}{4}$	$-\frac{3}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{2}$	$-\frac{1}{2}$	0	1	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{6}$ $\frac{1}{2}$ $\frac{1}{6}$ $\frac{5}{6}$	s_{19}^{+}
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$\frac{3}{4}$	$-\frac{1}{4}$	$\frac{1}{2}$	$-\frac{1}{2}$	0	0	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{1}{6}$	s_{17}^{-}
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2}$	$-\frac{1}{4}$	$\frac{3}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{2}$	$\frac{1}{2}$	0	-1	$-\frac{1}{2}$	0	$-\frac{1}{2}$		s_{20}^{-}
	(1,1,1,1)	$-\frac{1}{2}$	$ \begin{array}{r} \frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$ \begin{array}{r} \frac{1}{4} \\ -\frac{3}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$-\frac{1}{4}$	$ \begin{array}{r} 1\\ 4\\ -1\\ -1\\ 4\\ -1\\ 4\\ -1\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 3\\ 4\\ 4\\ 3\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\ 4\\$	$-\frac{1}{2} \frac{1}{2} 1$	$ \begin{array}{c} \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	0	-1	$ \begin{array}{r} -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	0	$ \begin{array}{c c} -\frac{1}{2} \\ -$	0	s_{18}^{-}
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$-\frac{1}{6}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$\frac{3}{4}$	$\frac{1}{2}$	$-\frac{1}{2}$	0	$-\frac{2}{3} - \frac{2}{3}$	$-\frac{1}{2}$	0	$-\frac{1}{2}$	$\frac{2}{3}$ $\frac{2}{3}$	\bar{v}_4
	(3,1,1,1)	$\frac{1}{6}$							0	$-\frac{2}{3}$		0	- 4		v_4
$T_{4(0,0,0,0,0,0)}$	(1,1,1,1)	0	$\frac{\frac{1}{3}}{\frac{1}{3}}$	0	-1	0	0	$-\frac{2}{3} \\ -\frac{2}{3}$	0	0	$-\frac{1}{3} \\ -\frac{1}{3}$	$-\frac{2}{3} \\ -\frac{2}{3}$	0	0	s_{26}^{0}
T	(1,1,1,1)	0		-1	0	0	0		0	1			0	0	\bar{n}_9
$T_{4(0,0,1,0,0,0)}$	(1,1,1,2)	0	$-\frac{1}{6}$ $-\frac{1}{6}$ $-\frac{1}{6}$ $\frac{1}{3}$	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$-\frac{\frac{1}{6}}{\frac{1}{6}}$	$\frac{\frac{3}{6}}{5}$	$-\frac{2}{3}$ $\frac{1}{3}$ $\frac{1}{3}$ $\frac{1}{3}$	$-\frac{\frac{1}{3}}{\frac{2}{3}}$	$\frac{1}{3}$	$\begin{bmatrix} 0 \\ 0 \end{bmatrix}$	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	$-\frac{1}{3}$	0	$\frac{2}{3}$	h_8
	$egin{array}{c} ({f 1},{f 1},{f 1},{f 2}) \ ({f 1},{f 1},{f 1},{f 1}) \end{array}$	0	$-\frac{6}{6}$	$\frac{\overline{2}}{1}$	<u>6</u> 5	$\frac{\overline{6}}{5}$	$\frac{3}{1}$	$-\frac{3}{1}$	$\frac{3}{2}$	$\begin{array}{c c} 0 \\ 1 \end{array}$	$-\frac{1}{3}$	$-\frac{3}{2}$	0	$\frac{2}{3}$	h_7 $\bar{\pi}$
	$(\overline{\bf 3}, {\bf 1}, {\bf 1}, {\bf 1})$	$\frac{1}{3}$	$-\frac{6}{1}$	$-\frac{1}{2}$	$-\frac{6}{6} \\ -\frac{1}{3}$	$\frac{6}{2}$	$\frac{3}{1}$	$\frac{3}{1}$	$-\frac{3}{2}$	$-\frac{1}{3}$	$-\frac{3}{3}$	$-\frac{3}{2}$	0	$\frac{3}{0}$	$ \bar{d}_3 $
$T_{4(0,0,1,1,0,0)}$	(1,2,1,1)	$\frac{3}{\frac{1}{2}}$	3 1	0	3 <u>1</u>	3 2		3 1	$\frac{3}{2}$	0	<u>3</u> <u>1</u>	$-\frac{3}{2}$	0		$\bar{\phi}_3$
4(0,0,1,1,0,0)	(1,1,1,1)	$\begin{array}{c} 2 \\ 0 \end{array}$	3 5		$-\frac{3}{1}$	$-\frac{3}{5}$	$-\frac{3}{2}$	$\frac{3}{\frac{1}{2}}$	$\frac{3}{2}$	-1	$-\frac{3}{1}$	$-\frac{3}{2}$	0	$\frac{1}{3}$ $\frac{2}{3}$	n_7
	$(\overline{\bf 3}, {f 1}, {f 1}, {f 1})$	$\frac{1}{3}$	$-\frac{\frac{1}{3}}{\frac{5}{6}}$ $-\frac{1}{6}$	$-\frac{1}{2} \\ -\frac{1}{2}$	$-\frac{\frac{1}{3}}{\frac{1}{6}}$ $-\frac{1}{6}$	5 65 65 62 32 35 61 6 	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	1 3 1 3 1 3	1 31 32 32 32 32 32 32 32 32 32 32	$\frac{2}{3}$	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	-23 -32 -32 -32 -32 -33 -32 -33 -33 -33	0	0	$ar{\delta}_2$
$T_{4(-1,1,0,0,0,0)}$		0	$-\frac{2}{3}$	0	0	0	0		0	$\frac{1}{2}$			0	$\frac{1}{2}$	\bar{f}_3
(, , , , , , ,	$({f 1},{f 2},{f 1},{f 1})$	$\frac{1}{2}$	$\frac{1}{3}$	0	0	-1	0	$ \begin{array}{r} \frac{1}{3} \\ -\frac{1}{3} \\$	0	0	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $		0	$\frac{1}{2}$ $\frac{1}{2}$	$\bar{\phi}_2$
	$({f 1},{f 1},{f 1},{f 1})$	0	$\frac{1}{3}$	0	-1	0	0	$-\frac{2}{3}$	0	0	$-\frac{1}{3}$	$-\frac{2}{3}$	0	0	s_{28}^{0}
	$({f 1},{f 1},{f 1},{f 2})$	0	$-\frac{2}{3}$	0	0	0	0	$\frac{1}{3}$	1	0	$-\frac{1}{3}$	$-\frac{2}{3}$	0	$\frac{\frac{1}{2}}{\frac{1}{2}}$	h_5
	$({f 1},{f 1},{f 1},{f 2})$	0	$-\frac{2}{3}$	0	0	0	0	$\frac{1}{3}$	-1	1	$-\frac{1}{3}$	$-\frac{2}{3}$	0	$\frac{1}{2}$	$\bar{\eta}_1$
	(1,1,1,1)	0	$\frac{1}{3}$	-1	0	0	0	$-\frac{2}{3}$	0	1	$-\frac{1}{3}$	$-\frac{2}{3}$	0	0	\bar{n}_{10}
T.	$(\overline{3}, 1, 1, 1)$	$\frac{1}{3}$	$\begin{array}{c} -\frac{2}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{6} \\ -1$	0	0	1 5	0	$-\frac{2}{3}$	0	$\frac{2}{3}$			0	$\frac{\frac{1}{2}}{\frac{5}{6}}$	$\bar{\delta}_1$
$T_{4(-1,1,1,0,0,0)}$	(1,1,1,1)	0	6 1	$ \begin{array}{r} -\frac{1}{2} \\ \frac{1}{2} \\ \frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \end{array} $	16161616561316 	6 5	$-\frac{1}{3}$ $-\frac{2}{3}$ $\frac{1}{3}$ $-\frac{1}{3}$	$ \begin{array}{r} \frac{1}{3} \\ \frac{1}{3} \\ \frac{2}{3} \\ \frac{1}{6} \\ \frac{1}{3} \\ \frac{1}{3} \\ \frac{1}{3} \\ \frac{1}{3} \end{array} $	$-\frac{2}{3}$	0	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \end{array} $		0	6 2	s_{32}^{0}
	(1,1,1,2)	0	$-\frac{1}{6}$	$\frac{1}{2}$	$\frac{1}{6}$	6 5	$-\frac{2}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	$\begin{bmatrix} 0 \\ 0 \end{bmatrix}$	$-\frac{1}{3}$	$-\frac{2}{3}$	0	$\frac{2}{3}$	h_{10}
	$egin{array}{c} (1,1,1,2) \ (1,1,8,1) \end{array}$	$\begin{array}{c} 0 \\ 0 \end{array}$	$-\frac{6}{6}$	$\frac{2}{1}$	$\frac{\overline{6}}{1}$	6 5	$\frac{3}{1}$	$-\frac{3}{3}$	$\frac{3}{1}$	0	$-\frac{3}{3}$	$-\frac{3}{2}$	0		h_9
	(1,1,0,1) (1,1,1,1)	0	$\frac{-6}{1}$	$-\frac{\overline{2}}{1}$	$-\frac{6}{5}$	$\frac{6}{5}$	$-\frac{6}{6}$	$-\frac{6}{6}$	$-\frac{3}{2}$	1	$\frac{-\frac{3}{3}}{1}$	$-\frac{3}{2}$	0	$\frac{5}{6}$ $\frac{2}{3}$	\overline{n}_{12}
	$(\overline{\bf 3}, {\bf 1}, {\bf 1}, {\bf 1})$		$\frac{6}{1}$	$\overset{2}{0}$	$-\frac{6}{1}$	$-\frac{6}{2}$	3 <u>1</u>	$\frac{3}{1}$	$-\frac{3}{2}$	$-\frac{1}{3}$	$-\frac{3}{1}$	$-\frac{3}{2}$	0	$\frac{3}{0}$	\bar{d}_4
	$({f 1},{f 2},{f 1},{f 1})$	$\frac{\frac{1}{3}}{\frac{1}{2}}$	$-\frac{\frac{1}{3}}{\frac{1}{6}}$	$-\frac{1}{2}$	$\frac{3}{\frac{1}{a}}$	$-\frac{3}{1}$	$\frac{\frac{1}{3}}{\frac{1}{3}}$	$\frac{3}{\frac{1}{2}}$	$-\frac{3}{2}$	$\frac{3}{1}$	$-\frac{3}{1}$	$-\frac{3}{2}$	0		$\bar{\ell}_1$
$T_{4(-1,1,1,1,0,0)}$		$\frac{1}{2}$	1 2	$\frac{2}{0}$		$\frac{6}{2}$		$\frac{3}{\frac{1}{2}}$		0			0	1 3 5 6 5 6 1 2 2 3	$\bar{\phi}_4$
1(1,1,1,1,0,0)	(1, 1, 1, 1)	$\stackrel{\scriptstyle 2}{0}$	$\frac{\frac{1}{3}}{\frac{1}{3}}$	0	$-\frac{\frac{1}{3}}{\frac{2}{3}}$	5 2	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	$\frac{3}{\frac{1}{2}}$	$\frac{3}{2}$	0	$-\frac{1}{3} \\ -\frac{1}{3}$	$-\frac{2}{3} \\ -\frac{2}{3}$	0	5 6	s_{35}^{0}
	$({f 1},{f 1},{f 1},{f 2})$	0		$\frac{1}{2}$	$-\frac{3}{6}$	$-\frac{5}{6}$	$-\frac{3}{3}$	$-\frac{3}{3}$	$-\frac{1}{3}$	0		$-\frac{3}{3}$	0	<u>5</u>	h_{12}
	$({f 1},{f 1},{f 1},{f 2})$	0	$-\frac{1}{6}$	$ \begin{array}{r} \frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	$ \begin{array}{r} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \end{array} $	$-\frac{5}{6}$	$\frac{2}{3}$	$\frac{1}{3}$	$-\frac{1}{3}$	-1	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	$-\frac{2}{3}$ $-\frac{2}{3}$ $-\frac{2}{3}$ $-\frac{2}{3}$	0	$\frac{1}{2}$	η_1
	$({f 1},{f 1},{f 1},{f 1})$	0	$\frac{5}{6}$	$-\frac{1}{2}$	$-\frac{1}{6}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$\frac{1}{3}$	$\frac{2}{3}$	-1	$-\frac{1}{3}$	$-\frac{2}{3}$	0	$\frac{2}{3}$	n_8
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$\frac{1}{3}$	$-\frac{1}{6}$	$-\frac{1}{2}$		$\frac{1}{6}$	$-\frac{1}{3}$	$\frac{1}{3}$	$\frac{2}{3}$	$\frac{2}{3}$ $\frac{1}{3}$	$-\frac{1}{3}$		0	0	$\bar{\delta}_3$
$T_{5(0,0,0,0,0,0)}$	$({f 3},{f 2},{f 1},{f 1})$	$\frac{1}{6}$	$\frac{1}{6}$	0	0	$\frac{1}{2}$	0	$-\frac{1}{3}$	0	$\frac{1}{3}$	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$	0	q_2
	(1,1,1,1)	1	$\frac{1}{6}$	0	0	$\frac{1}{2}$	0	$-\frac{1}{3}$	0	1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ $-\frac{1}{2}$ $-\frac{1}{2}$	0	\bar{e}_2
	$(\overline{3},1,1,1)$	$-\frac{2}{3}$	$\frac{1}{6}$	0	0	5 65 65 65 65 62 3+ 62 35 35 65 65 65 64 6+ 24 24 24 20	0	- 13-13-13-13-13-13-13-13-13-13-13-13-13-1	0	$-\frac{1}{3}$	$-\frac{1}{6}$	$ \begin{array}{r} -\frac{1}{3} \\ -$	$-\frac{1}{2} \\ -\frac{1}{2}$	0	\bar{u}_2
	(1,1,1,1)	0	$\frac{2}{3}$	$-\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$	$ \begin{array}{r} \frac{1}{2} \\ \frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \end{array} $		0	$-\frac{1}{3}$	0	0	6 11	$-\frac{1}{3}$	$-\frac{1}{2}$	0	$s_{8}^{0} \\ s_{9}^{0}$
	(1,1,1,1)	0	$-\frac{1}{3}$	$\frac{1}{2}$	$\frac{1}{2}$	0	0	$-\frac{1}{3}$	0	0	6	$-\frac{1}{3}$	$-\frac{1}{2}$	0	s_9^0
	(1,1,1,1)	$\begin{array}{c} 0 \\ 0 \end{array}$	$-\frac{1}{3}$	$\frac{1}{2}$	$\frac{\frac{2}{2}}{1}$	0 0	0	$-\frac{1}{3}$	0	$\begin{vmatrix} 0 \\ -1 \end{vmatrix}$	$-\frac{1}{6}$	$\frac{2}{3}$	$-\frac{1}{2}$	0	s_5^0
	$(1,1,1,1) \ (1,2,1,1)$	$-\frac{1}{2}$	$\begin{array}{c} -\frac{1}{6} \\ -\frac{1}{65} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{6} \\ -$	$\frac{\overline{2}}{0}$	$-\frac{1}{2}$	$-\frac{3}{2}$	0	$-\frac{3}{1}$	0	-1 -1	$\begin{array}{c} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ \frac{5}{6} \\ \frac{11}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ \end{array}$	$-\frac{3}{1}$	$-\frac{1}{2}$	0	$n_1 \\ \ell_2$
	(1 , 2 , 1 , 1)	$-\frac{1}{2}$	6	U	U	$-\frac{1}{2}$	U	$-{3}$	U	-1	6	$-{3}$	$\frac{-\frac{1}{2}}{ntinu}$		
												co	uuuuu	cu.	• •

Sector	Representation	q_Y	q_1	q_2	q_3	q_4	q_5	q_6	q_7	q_{B-L}	R_1	R_2	$R_3 \gamma ($	ϑ)
	(1, 1, 1, 1)	0	$-\frac{1}{3}$	$-\frac{1}{2}$ $-\frac{1}{2}$	$-\frac{1}{2} \\ -\frac{1}{2}$	0	0	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	0	1	$ \frac{11}{6} $ $ -\frac{1}{6} $ $ -\frac{1}{6} $ $ -\frac{1}{6} $	$-\frac{1}{3}$	$-\frac{1}{2}$ (\bar{n}_3
	$({f 1},{f 1},{f 1},{f 1})$	0	$-\frac{1}{3}$	$-\frac{1}{2}$	$-\frac{1}{2}$	0	0	$-\frac{1}{3}$	0	1	$-\frac{1}{6}$	$-\frac{1}{3} \\ \frac{2}{3} \\ 1$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	\bar{n}_2
	$({f 1},{f 1},{f 1},{f 1})$	0	$\frac{1}{6}$	0	0	$-\frac{\frac{5}{2}}{\frac{3}{2}}$	0	$-\frac{1}{3}$	0	1	$-\frac{1}{6}$		$-\frac{1}{2}$ (-
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$\frac{1}{3}$	$\frac{1}{6}$	0	0	$-\frac{3}{2}$	0		0	$-\frac{1}{3}$	$-\frac{1}{6}$	$-\frac{\overline{3}}{\overline{3}}$		$ \bar{d}_2 $
$T_{5(0,0,0,0,0,1)}$	$({f 3},{f 2},{f 1},{f 1})$	$\frac{1}{6}$	$\frac{1}{6}$	0	0	$\frac{\frac{1}{2}}{\frac{1}{2}}$ $\frac{1}{2}$	0	$-\frac{1}{3}$	0	$\frac{1}{3}$	$-\frac{1}{6}$	$-\frac{1}{3} \\ -\frac{1}{3} \\ 1$	$-\frac{1}{2}$ (q_1
	$({f 1},{f 1},{f 1},{f 1})$	1	$\frac{1}{6}$	0	0	$\frac{1}{2}$	0	$-\frac{1}{3}$	0	1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$-\frac{2}{3}$	$\frac{1}{6}$	0	0	$\frac{1}{2}$	0	$-\frac{1}{3}$	0	$-\frac{1}{3}$	$-\frac{1}{6}$		$-\frac{1}{2}$ (_
	(1, 1, 1, 1)	0	$\frac{2}{3}$	$-\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$	$ \begin{array}{r} \frac{1}{2} \\ \frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \end{array} $	0	0	$-\frac{1}{3}$	0	0	5 6	$-\frac{3}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $\frac{2}{3}$	$-\frac{1}{2}$ (Τ.
	(1, 1, 1, 1)	0	$-\frac{1}{3}$	$\frac{1}{2}$	$\frac{1}{2}$	0	0	$-\frac{1}{3}$	0	0	$\frac{11}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (1(
	(1, 1, 1, 1)	0	$-\frac{1}{3}$	$\frac{1}{2}$	$\frac{1}{2}$	0	0	$-\frac{1}{3}$	0	0	$-\frac{1}{6}$	$\frac{2}{3}$	$-\frac{1}{2}$ (12
	(1, 1, 1, 1)	0	$\frac{2}{3}$	$\frac{1}{2}$		0	0	$-\frac{1}{3}$	0	-1	6	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	(1, 2, 1, 1)	$-\frac{1}{2}$	6	0	0	$-\frac{3}{2} \\ 0$	0	$-\frac{1}{3}$	0	-1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	(1,1,1,1)	0	$-\frac{1}{3}$	$-\frac{1}{2}$	$-\frac{1}{2} \\ -\frac{1}{2}$		0	$-\frac{1}{3}$	0	1	6	$-\frac{1}{3}$	$-\frac{1}{2}$ (Ŭ
	(1,1,1,1)	0	$-\frac{1}{3}$	$-\frac{1}{2} \\ -\frac{1}{2} \\ 0$		0	0	$-\frac{1}{3}$	0	1	$\begin{array}{c} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ \frac{5}{6} \\ \frac{11}{6} \\ -\frac{1}{6} \\ \frac{11}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \end{array}$	$\frac{\frac{2}{3}}{1}$	$ \begin{array}{c} -\frac{1}{2} \\ -\frac{1}{2} \end{array} $	
	(1,1,1,1)	0 ₁	-1316161616233131323316613161616	0	0	$-rac{rac{5}{2}}{2}$	0	- 13131313131313131313131313131313131313	0	1	$-\frac{1}{6}$	$ \begin{array}{r} -\frac{3}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \end{array} $	$-\frac{1}{2}$	
T	$ \begin{array}{c c} (\overline{\bf 3},{\bf 1},{\bf 1},{\bf 1}) \\ \hline ({\bf 1},{\bf 1},{\bf 1},{\bf 1}) \end{array}$	$\frac{1}{3}$						$-\frac{3}{5}$	0	$-\frac{1}{3}$	$-\frac{1}{6}$	$-\frac{3}{3}$		
$T_{5(0,0,0,0,1,0)}$	(1,1,1,1) $(1,1,1,1)$	$\begin{array}{c} \frac{1}{2} \\ \frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{6} \end{array}$	$-\frac{12}{1}$	$\frac{4}{1}$	$\frac{\overline{4}}{1}$	$\frac{\overline{4}}{1}$	$-\frac{1}{2}$	$-\frac{6}{6}$	0	0	$\frac{-6}{11}$	$-\frac{3}{3}$	$-\frac{1}{2}$ $\frac{1}{2}$ $-\frac{1}{2}$ $\frac{1}{2}$ $-\frac{1}{2}$ $\frac{1}{2}$	$egin{array}{c c} ar{s}_1^+ & s_1^+ \ s_2^+ & s_2^+ \ \hline s_2^+ & s_2^+ \end{array}$
	(1,1,1,1) $(1,1,1,1)$	$\frac{2}{1}$	$-\frac{12}{1}$	$\frac{4}{1}$	$\frac{\overline{4}}{1}$	$\frac{4}{1}$	$\frac{2}{1}$	$\frac{6}{1}$	0	0	6 _ 1	$\frac{3}{2}$	$-\frac{\overline{2}}{1}$	2 3
	(1,1,1,1) $(1,1,1,1)$	$-\frac{2}{1}$	$-\frac{12}{7}$	$-\frac{4}{1}$	$-\frac{4}{1}$	$-\frac{4}{1}$	$\frac{2}{1}$	$\frac{6}{1}$	0	0	6 <u>5</u>	$-\frac{3}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	$-\frac{2}{1}$ (
	(1,1,1,1) $(1,1,1,1)$	$\frac{2}{1}$	$\frac{12}{5}$	$-\frac{4}{1}$	$\frac{4}{3}$	$-\frac{4}{1}$	$\frac{2}{1}$	$\frac{6}{1}$	0	-1	$-\frac{6}{1}$		$-\frac{2}{1}$ (
	(1,1,1,1)	$-\frac{2}{1}$	$\frac{12}{5}$	$\frac{4}{3}$	$-\frac{\frac{4}{1}}{1}$	$-\frac{4}{1}$	$\frac{2}{1}$	$\frac{6}{1}$	0	-2	$-\frac{6}{1}$	$-\frac{3}{1}$	$-\frac{2}{1}$ (
	$(\overline{\bf 3},{\bf 1},{\bf 1},{\bf 1})$	$-\frac{2}{1}$	$\frac{12}{5}$	$-\frac{4}{1}$	$-\frac{\frac{4}{1}}{4}$	$\frac{4}{3}$	$\frac{1}{2}$	$\frac{6}{1}$	0	$-\frac{1}{3}$	$-\frac{6}{1}$	$-\frac{3}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	$-\frac{1}{2}$	
	$({f 1},{f 2},{f 1},{f 1})$	0	$ \begin{array}{r} -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{7}{12} \\ -\frac{7}{12} \\ \frac{5}{12} $	$-\frac{\frac{4}{1}}{4}$	$ \begin{array}{c} \frac{1}{4} \\ \frac{1}{4} \\ \frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$ \begin{array}{c} \frac{1}{4} \\ \frac{1}{4} \\ \frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{$	$\frac{\frac{2}{1}}{2}$	- 561616161616161616	0	-1^{3}	$ \begin{array}{r} -\frac{1}{6} \\ \underline{11} \\ 6 \\ -\frac{1}{6} \\ \underline{6} \\ -\frac{1}{6} \\ -\frac{1}{$	$-\frac{3}{2}$	$ \begin{array}{c} -\frac{1}{2} \\ -\frac{1}{2} \\ -\frac{1}{2} \end{array} $ $ -\frac{1}{2} \\ -\frac{1}{2} \\$	$\frac{1}{2}$ m_1
$T_{5(0,0,0,0,1,1)}$	$({f 1},{f 1},{f 1},{f 1})$	$\frac{1}{2}$		$\frac{\frac{4}{1}}{4}$	$\frac{\frac{4}{1}}{4}$		$-\frac{1}{2}$	$-\frac{5}{6}$	0	1	$-\frac{1}{6}$	$-\frac{1}{3}$		$\frac{1}{2} s_4^+$
- (-,-,-,-, , ,	$({f 1},{f 1},{f 1},{f 1})$	$ \begin{array}{c} 1 \\ 2 \\ 1 \\ 2 \\ 1 \\ 2 \\ - 1 \\ 2 \\ - 1 \\ 2 \end{array} $	$-\frac{1}{12}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{2}$	$\frac{1}{6}$	0	0	1 <u>1</u>	$-\frac{1}{3}$	$ \begin{bmatrix} -\frac{1}{2} & \frac{1}{2} \\ -\frac{1}{2} & \frac{1}{2} \\ -\frac{1}{2} & \frac{1}{2} \end{bmatrix} $	$\frac{1}{2} s_6^+$
	$({f 1},{f 1},{f 1},{f 1})$	$\frac{1}{2}$	$-\frac{1}{12}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{4}$	$\frac{1}{2}$	$\frac{3}{6}$	0	0	$-\frac{1}{6}$	$\frac{\frac{2}{2}}{3}$	$-\frac{1}{2}$	$\left s_5^+\right $
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{\overline{1}}{2}$	$-\frac{7}{12}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$\frac{\overline{1}}{2}$	$\frac{1}{6}$	0	0	$\frac{5}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	$({f 1},{f 1},{f 1},{f 1})$		$\frac{5}{12}$	$-\frac{1}{4}$	$\frac{3}{4}$	$-\frac{1}{4}$	$\frac{1}{2}$	$\frac{1}{6}$	0	-1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2} \\ -\frac{1}{6}$	$ \begin{array}{r} -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{7}{12} \\ 5 \\ \hline 12 \\ \hline \\ 12 \\ 12 \\ 12 \\ 12 \\ 12 \\ 12 \\ 12 \\$	$\begin{array}{c} 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 1 \\ 4 \\ 4$	$ \begin{array}{c} 1 \\ 4 \\ 1 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4 \\ 4$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$			0	-2	$ \begin{array}{r} -\frac{1}{6} \\ \frac{11}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \end{array} $	$ \begin{array}{r} -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ -\frac{1}{3} \\ 1 \end{array} $	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ s_4^- $
	$({f \overline{3}},{f 1},{f 1},{f 1})$	$-\frac{1}{6}$	$\frac{5}{12}$	$-\frac{1}{4}$	$-\frac{1}{4}$	$\frac{3}{4}$	$\frac{1}{2}$	$\frac{1}{6}$	0	$-\frac{1}{3}$	$-\frac{1}{6}$		$-\frac{1}{2}$ $\frac{1}{2}$ $-\frac{1}{2}$ $\frac{1}{2}$	$ \bar{v}_2 $
	$({f 1},{f 2},{f 1},{f 1})$	0				$-\frac{5}{4}$			0	-1		$-\frac{\frac{3}{1}}{\frac{1}{3}}$		
$T_{5(0,0,1,1,0,0)}$	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{\frac{1}{6}}{\frac{1}{6}}$	0	$-\frac{2}{3}$ $-\frac{2}{3}$ $-\frac{2}{3}$	$-\frac{5}{6}$	$-\frac{1}{3} \\ -\frac{1}{3} \\ \frac{2}{3}$	$-\frac{1}{3}$ $\frac{2}{3}$ $-\frac{1}{3}$	$-\frac{1}{3}$	0	$-\frac{1}{6}$ $-\frac{1}{6}$ $-\frac{1}{6}$	$-\frac{1}{3}$	$ \begin{array}{c} -\frac{1}{2} \\ -$	$\frac{2}{3}$ h_1
	(1, 1, 1, 1)	0	$\frac{1}{6}$	0	$-\frac{2}{3}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	-1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$	$\frac{1}{3}$ n_4
	(1,1,1,1)	0	$\frac{\frac{1}{6}}{1}$	0	$-\frac{2}{3}$	$-\frac{5}{6}$	$\frac{\frac{2}{3}}{1}$	$-\frac{1}{3}$	$\frac{\frac{2}{3}}{1}$	-1		$-\frac{1}{3}$	$-\frac{1}{2}$	n_3
$T_{5(0,0,1,1,0,1)}$	(1, 1, 1, 2)	0	$\frac{1}{6}$	0	$-\frac{2}{3}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$	$\frac{2}{3}$ h_2
	(1,1,1,1)	0	6	0	$-\frac{2}{3}$	$-\frac{5}{6}$	$-\frac{1}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	-1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$	$\frac{1}{8}$ n_6
T	(1,1,1,1)	0	$\frac{1}{6}$	0	$-\frac{2}{3}$	$-\frac{5}{6}$	$\frac{2}{3}$	$-\frac{1}{3}$	$\frac{2}{3}$	-1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$	$\frac{2}{3}$ n_5
$T_{5(0,0,1,1,1,0)}$	(1,1,1,2)	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$-\frac{1}{12}$	$-\frac{1}{4}$	$\frac{1}{12}$	$-\frac{1}{12}$	- 6 1	1 6	$-\frac{1}{3}$	1	$-\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ (
	(1,1,1,2)		$-\frac{1}{12} \\ \frac{1}{5}$	$\frac{1}{4}$	$-\frac{3}{12}$	$-\frac{13}{12}$	$-\frac{1}{3}$ $\frac{1}{3}$ $\frac{1}{3}$ $\frac{1}{6}$ $\frac{1}{6}$ $\frac{1}{6}$	$\frac{1}{6}$	$-\frac{1}{3}$	-1	$ \begin{array}{r} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \end{array} $	$-\frac{1}{3}$	$ \begin{array}{c} -\frac{1}{2} \\ -$	x_1^{-1}
T	(1,2,1,2)	0	$\frac{3}{12}$	$-\frac{1}{4}$	$\frac{12}{1}$	$\frac{3}{12}$		$\frac{\frac{1}{6}}{1}$	$-\frac{1}{3}$	0		$-\frac{1}{3}$		y_1
$T_{5(0,0,1,1,1,1)}$	(1,1,1,2)	$-\frac{\frac{1}{2}}{\frac{1}{2}}$	$-\frac{1}{12}$	$-\frac{1}{4}$	$\frac{12}{5}$	$-\frac{12}{13}$	$\frac{1}{6}$	$\frac{1}{6}$	$-\frac{1}{3}$	1	$-\frac{1}{6} \\ -\frac{1}{6}$	$-\frac{1}{3}$	$-\frac{1}{2}$ ($-\frac{1}{2}$ $\frac{5}{6}$	
	(1,1,1,2)		$ \begin{array}{r} \frac{1}{6} \\ \frac{1}{6} \\ -\frac{7}{12} \\ -\frac{1}{12} \\ -\frac{7}{12} \\ -\frac{7}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \\ \frac{5}{12} \end{array} $	$ \begin{array}{r} -\frac{1}{4} \\ \frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \\ -\frac{1}{4} \end{array} $	$ \begin{array}{r} -\frac{2}{3} \\ -\frac{2}{3} \\ -\frac{2}{3} \\ -\frac{1}{12} \\ -\frac{5}{12} \\ -\frac{1}{12} \\ -\frac{5}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \end{array} $	$ \begin{array}{r} -\frac{5}{6} \\ -\frac{5}{6} \\ -\frac{5}{6} \\ -\frac{5}{6} \\ -\frac{5}{6} \\ -\frac{5}{6} \\ -\frac{7}{7} \\ -\frac{13}{12} $	$\frac{\frac{1}{6}}{\frac{1}{6}}$	$-\frac{1}{3} \underbrace{\frac{2}{3}}_{1} \underbrace{\frac{1}{6}}_{1} \underbrace{\frac{1}{6}}$	-132132132132131313131313131313131313131	-1	$\frac{-\frac{1}{6}}{1}$	$-\frac{1}{3}$	$-\frac{1}{2}\begin{vmatrix} \frac{1}{2} & \frac{1}{2} \\ -\frac{1}{2} & \frac{1}{2} \end{vmatrix}$	$\begin{bmatrix} x_2^- \\ y_2 \end{bmatrix}$
	$({f 1},{f 2},{f 1},{f 2})$	0	$\frac{5}{12}$	$-\frac{1}{4}$	12	$\frac{5}{12}$	$\frac{\bar{6}}{6}$	6	$-\frac{1}{3}$	0	$-\frac{1}{6}$	$-\frac{1}{3}$	$\frac{-\frac{1}{2}}{ntinue}$	

$T_{5(0,0,0,1,0,0)}$	$({f 1},{f 1},{f 1},{f 1})$				q_3	q_4	q_5	q_6		q_{B-L} I	$R_1 R_2$	$R_3 \gamma(i)$	
0(0,0,0,1,0,0)	(-,-,-,-,	0	$\frac{1}{6}$	0	$-\frac{1}{3}$	5 65 65 65 6	$\frac{1}{3}$	$-\frac{\frac{2}{3}}{\frac{1}{6}}$	$-\frac{2}{3}$ $\frac{1}{3}$ $\frac{1}{3}$ $-\frac{2}{3}$	0	$ \begin{array}{ccc} $	$-\frac{1}{2} \frac{2}{3}$	s_{22}^{0}
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{\frac{1}{6}}{\frac{1}{6}}$	0	$-\frac{1}{3}$	$\frac{5}{6}$	$\frac{1}{3}$	$-\frac{1}{3}$	$\frac{1}{3}$	0	$\frac{\frac{5}{6}}{\frac{5}{6}}$ $-\frac{1}{3}$	$ \begin{array}{c c} -\frac{1}{2} & \frac{2}{3} \\ -\frac{1}{2} & 0 \\ 1 & 2 \end{array} $	
	$({f 1},{f 1},{f 8},{f 1})$	0	$\frac{1}{6}$	0	$-\frac{1}{3}$	$\frac{5}{6}$	$-\frac{1}{6}$	$\frac{1}{6}$	$\frac{1}{3}$	0 –	$-\frac{1}{6}$ $-\frac{1}{3}$		w_1
	$({f 1},{f 1},{f 1},{f 1})$	0	$\frac{1}{6}$	0	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	$\frac{\tilde{5}}{6}$	$-\frac{1}{3}$ $-\frac{1}{6}$ $-\frac{2}{3}$	$-\frac{1}{3}$	$-\frac{2}{3}$	1	$\frac{5}{6}$ $-\frac{1}{3}$	$-\frac{1}{2} = 0$	\bar{n}_7
$T_{5(0,0,0,1,0,1)}$	(1, 1, 1, 1)	0		0				$\frac{2}{3}$	$-\frac{2}{3}$	0	$\frac{5}{6}$ $-\frac{1}{3}$	$-\frac{1}{2} \frac{2}{3}$	s_{24}^{0}
	$({f 1},{f 1},{f 1},{f 2})$	0	$\frac{\frac{1}{6}}{\frac{1}{6}}$	0	$-\frac{1}{3}$	<u>5</u>	$\frac{1}{3}$	$-\frac{1}{3}$	$\frac{1}{3}$	0	5 1		h_4
	(1, 1, 8, 1)	0	$\frac{1}{6}$	0	$-\frac{1}{3}$	$\frac{5}{6}$	$-\frac{3}{6}$	$\frac{1}{6}$	$\frac{1}{3}$	0 –		$-\frac{1}{2} \frac{2}{3}$	w_2
	(1, 1, 1, 1)	0	$\frac{1}{6}$	0	$-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$ $-\frac{1}{3}$	5 65 <mark> </mark> 65 65 6	$ \begin{array}{r} \frac{1}{3} \\ \frac{1}{3} \\ -\frac{1}{6} \\ -\frac{2}{3} \end{array} $	$-\frac{1}{3}$	$-\frac{2}{3}$ $\frac{1}{3}$ $\frac{1}{3}$ $-\frac{2}{3}$	1	$-\frac{1}{6}$ $-\frac{1}{3}$ $\frac{5}{6}$ $-\frac{1}{3}$	$-\frac{1}{2} = 0$	
$T_{5(0,0,0,1,1,0)}$	(1, 1, 1, 1)	$\frac{1}{2}$	$-\frac{7}{12}$	$-\frac{1}{4}$		13			$-\frac{2}{3}$	2 –		$-\frac{1}{2}$ $\frac{1}{6}$	s_8^+
	(1, 1, 1, 1)	$\frac{1}{2}$	$-\frac{1}{12}$	$\frac{1}{4}$	$-\frac{1}{12}$	$-\frac{17}{12}$	$-\frac{1}{6}$	$\frac{1}{6}$	$-\frac{3}{2}$	0	$\frac{1}{6}$ $-\frac{1}{3}$ $\frac{5}{6}$ $-\frac{1}{3}$	$-\frac{1}{2} \left \begin{array}{c} \frac{1}{6} \\ \frac{1}{6} \end{array} \right $	s_9^+
	$({f 1},{f 1},{f 1},{f 1})$	$ \begin{array}{r} $	$\frac{12}{5}$	$-\frac{1}{4}$ $-\frac{1}{4}$ $-\frac{1}{4}$ $\frac{1}{4}$ $\frac{1}{4}$ $\frac{1}{4}$	$-\frac{12}{12}$	$\frac{13}{12}$	$ \begin{array}{r} -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \\ -\frac{1}{6} \end{array} $	$\frac{1}{6}$	$-\frac{3}{2}$	1 –	$\begin{array}{ccc} \frac{1}{6} & -\frac{1}{3} \\ \frac{5}{6} & -\frac{1}{3} \\ \frac{1}{6} & -\frac{1}{3} \\ \frac{11}{6} & -\frac{1}{3} \end{array}$	$ \begin{array}{c cccc} -\frac{1}{2} & \frac{1}{6} \\ -\frac{1}{2} & \frac{1}{6} \\ -\frac{1}{2} & \frac{1}{3} \\ -\frac{1}{2} & \frac{1}{3} \\ -\frac{1}{2} & \frac{1}{3} \end{array} $	s_7^+
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2}$	$-\frac{1}{12}$	$\frac{1}{4}$	$-\frac{1}{12}$	$\frac{\frac{12}{7}}{12}$	$-\frac{1}{6}$	$\frac{1}{6}$	$-\frac{3}{2}$	0	$\frac{11}{6}$ $-\frac{1}{2}$	$-\frac{1}{2} \frac{1}{2}$	s_9^-
	$({f 1},{f 1},{f 1},{f 1})$		$-\frac{1}{10}$	$\frac{4}{1}$	$-\frac{1}{10}$	$\frac{\frac{12}{7}}{\frac{19}{19}}$	$-\frac{1}{6}$	$\frac{1}{c}$	$-\frac{3}{2}$	0 –			s_8^-
	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{1}{2} \\ -\frac{1}{2}$	$-\frac{1}{1}$	$\frac{4}{1}$	$-\frac{1}{1}$	$\frac{\frac{12}{7}}{12}$	$-\frac{1}{6}$	$\frac{1}{c}$	$\frac{3}{4}$	-1 -		1 0	
	$({f 3},{f 1},{f 1},{f 1})$	$\frac{\frac{2}{1}}{6}$	$ \begin{array}{r} -\frac{7}{12} \\ -\frac{1}{12} \\ \frac{5}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{1}{12} \\ -\frac{5}{12} \\ \frac{5}{12} \end{array} $	$-\frac{\frac{4}{1}}{4}$	5 12	$\frac{12}{19}$	$-\frac{1}{6}$	$\frac{1}{c}$	$-\frac{3}{2}$	$\frac{1}{3}$ -	$\begin{bmatrix} \frac{1}{6} & -\frac{1}{3} \\ \frac{1}{6} & -\frac{1}{3} \end{bmatrix}$	$-\frac{1}{2}$ 0 $-\frac{1}{2}$ $\frac{5}{6}$	v_1
	$({f 1},{f 2},{f 1},{f 1})$	0	$-\frac{1}{12}$	$-\frac{\frac{4}{3}}{4}$	$-\frac{1}{10}$	$-\frac{5}{10}$	$-\frac{1}{6}$	$\frac{1}{c}$	$-\frac{3}{2}$	$\stackrel{\circ}{1}$ $-$	$-\frac{1}{6}$ $-\frac{1}{3}$	$-\frac{1}{2} \begin{vmatrix} \frac{1}{2} \\ \frac{1}{2} \end{vmatrix}$	m_3
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	$({f 1},{f 1},{f 1},{f 1})$	$-\frac{2}{1}$	$-\frac{12}{1}$	$\frac{4}{1}$	$-\frac{12}{1}$	$\frac{12}{7}$	$-\frac{6}{1}$	$\frac{6}{1}$	$-\frac{3}{2}$	0 3	$\frac{1}{6}$ $-\frac{1}{3}$ $\frac{11}{6}$ $-\frac{1}{3}$	$ \begin{array}{c cccc} -\frac{1}{2} & \frac{1}{6} \\ -\frac{1}{2} & \frac{1}{3} \\ -\frac{1}{2} & \frac{1}{3} \\ -\frac{1}{2} & \frac{1}{3} \end{array} $	s_{12}^{-10}
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