



Ugidos, J.M., Barba, P., Valladares, M.I., Suarez, M., and [Ellam, R.M.](#)  
(2016) The Ediacaran-Cambrian transition in the Cantabrian Zone 1 (North  
Spain): sub-Cambrian weathering, K-metasomatism and provenance of  
detrital series. *Journal of the Geological Society*, 173(4), pp. 603-615.  
(doi:[10.1144/jgs2016-004](https://doi.org/10.1144/jgs2016-004))

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Deposited on: 1 February 2016

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1 **The Ediacaran-Cambrian transition in the Cantabrian Zone (North Spain): sub-**  
2 **Cambrian weathering, K-metasomatism and provenance of detrital series**

3  
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8  
9 **Abstract:** The Upper Ediacaran detrital succession in the Cantabrian Zone shows  
10 geochemical and mineralogical changes resulting from sub-Cambrian weathering during  
11 the Late Ediacaran worldwide sea-level fall. Relative to the unaltered rocks the altered  
12 ones show crosscutting rubefaction of varying thickness a remarkable increase in illite,  
13 K<sub>2</sub>O, Rb and Cs indicating K-metasomatism, and also depletion in MgO, CaO, Na<sub>2</sub>O,  
14 Be and Sr, but not in Zr, Nb, Y, Sc. The basal Cambrian siliciclastic rocks mostly  
15 consist of detritus derived from the Ediacaran materials, as demonstrated by the  
16 geochemical and Nd-isotope data [ $\epsilon_{Nd}(t)$  ranges: -3.4 to -2.1, and -3.6 to -1.8,  
17 respectively)]. However, these geochemical features of the basal Cambrian change  
18 upwards to more evolved compositions with lower  $\epsilon_{Nd}(t)$  values (-4.9 to -5.8). This  
19 change is the same as coinciding with the Ediacaran-Cambrian unconformity in the  
20 Central Iberian Zone and the Ediacaran siliclastic rocks of this zone and their unaltered  
21 equivalents of the Cantabrian Zone share the same geochemical features. This  
22 geochemical homogeneity rules out a significant coeval juvenile contribution to the  
23 Upper Ediacaran series. Thus, the juvenile supply took place in a geological setting  
24 different from the one where the Ediacaran series were finally deposited.

25 **Supplementary material:** [sampling, sample location, analytical techniques,  
26 diffractograms, tables of analyses, and figures] is available at.

27 Worldwide, the Ediacaran-Cambrian boundary corresponds to a major unconformity in  
28 most sections, suggesting a possible eustatic fall in sea level (Bartley *et al.* 1998; Saylor  
29 2003; Knoll *et al.* 2004; Pyle *et al.* 2004). The duration and magnitude of this  
30 unconformity vary between sections but its age appears to be the same within the  
31 precision afforded by biostratigraphic and chemostratigraphic correlations. Reflecting  
32 this sea-level fall, the Late Ediacaran-Early Cambrian interval is characterised by

33 elevated chemical weathering rates, as suggested by the steep rise in the normalised  
34 seawater strontium isotope curve (Shields 2007). Sub-Cambrian weathering profiles  
35 developed in many areas of Gondwana after the Pan-African orogeny on the  
36 Precambrian basement at continental scale (Avigad *et al.* 2005; Avigad & Gvirtzman  
37 2009; Johnson *et al.* 2011; Sandler *et al.* 2012 and references therein) and the age of this  
38 weathering ( $542.62 \pm 0.38$  Ma, Parnell *et al.* 2014 and references therein) is close to that  
39 of the Ediacaran-Cambrian boundary ( $541 \pm 0.13$  Ma, Bowring *et al.* 2007). Also, the  
40 Upper Ediacaran normal faulting and associated alkaline magmatism indicate a within-  
41 plate extensional tectonic setting in the northern margin of the West African craton  
42 (Ennih & Liégeois 2001; Piqué 2003; Thomas *et al.* 2004; Soulimani *et al.* 2004;  
43 Álvaro *et al.* 2010).

44 In the Cantabrian Zone (Fig. 1), the Upper Ediacaran series were deposited by  
45 turbidity currents (Pérez Estaún 1978) and show weathering and rubefaction below the  
46 Precambrian-Cambrian boundary (van der Bosch 1969; Gutiérrez-Alonso *et al.* 2004),  
47 which is an angular unconformity. In the Central Iberian Zone (Fig. 1A) the Ediacaran-  
48 Cambrian boundary is a disconformity, expressed as an irregular and strongly erosional  
49 surface related to a worldwide eustatic fall in sea level. A study of facies and facies  
50 associations of the Upper Ediacaran and Lower Cambrian detrital series revealed that  
51 the sedimentation resulted predominantly from gravity flows, such as turbidity currents  
52 and debris flows in slope and base-of-slope environments (Valladares *et al.* 2000, 2006)  
53 while in the Cantabrian Zone the setting of the basal part of the Lower Cambrian  
54 succession was a continental environment (Aramburu *et al.* 1992; Rubio-Ordoñez *et al.*  
55 2004). The aims of the present work are two-fold: to study the processes during the  
56 Ediacaran-Cambrian transition in the Cantabrian Zone and to determine the provenance  
57 of the Upper Ediacaran and Lower Cambrian detrital series. The results strongly suggest  
58 that the Upper Ediacaran detrital rocks derived from the same source as that of the  
59 equivalent rocks in the Central Iberian Zone while the Lower Cambrian siliciclastic  
60 rocks derived from mixtures of the unaltered and rubefacted Ediacaran materials. The  
61 contribution of an older crustal component increases upwards.

62

63 **Geological setting**

64 The Ediacaran rocks, known as the Narcea schists group (Fig. 2) or Mora formation,  
65 outcrop in the east part of the Narcea antiform (Fig. 1B), within the Cantabrian Zone,  
66 and its western boundary (Fig. 1B) is La Espina thrust (Pérez Estaún *et al.* 1990). In  
67 these rocks Pérez Estaún (1978) differentiated 1500-1700 m of alternating greywackes  
68 and shales with abundant sedimentary structures of turbiditic origin. The age of the  
69 Narcea schists is constrained by the presence of the acritarchs *Sphaerocongregus*  
70 *variabilis* and *Palaeogomphosphaeria caurensis* (Martín Parra *et al.* 1989; Palacios &  
71 Vidal 1992), both regarded as indicative of a late Vendian age (Vidal *et al.* 1994).  
72 Detrital zircons from the sandstones of the Narcea schists provided an age of  $553 \pm 4$  Ma  
73 (Fernández-Suaréz *et al.* 2014), which indicates a younger age for the Narcea  
74 succession in the Cantabrian Zone. The Upper Ediacaran series show weathering and  
75 crosscutting rubefaction up to 25 m thick below the Precambrian-Cambrian boundary  
76 (van der Bosch 1969).

77 In the Cantabrian and Central Iberian zones the Ediacaran siliciclastic successions  
78 were deposited predominantly from debris flows and turbidity currents in slope, base-  
79 of-slope and deep sea fan environments (Pérez Estaún 1978; Valladares *et al.* 2000).  
80 However, in the outcrops of the eastern-most area of the Central Iberian Zone, Montes  
81 de Toledo, and also in the Iberian Ranges (Fig. 1A, MT and IR, respectively), the upper  
82 part of the respective Ediacaran successions (Ibor and Paracuellos groups, Fig. 2) are  
83 mixed platforms (siliciclastic-carbonate), from offshore to shoreface, with sandy and  
84 ooidal-bioclastic shoals (Álvarez Nava *et al.* 1988; Calvet & Salas 1988; Álvaro &  
85 Blanc-Valleron 2002; Álvaro *et al.* 2008). The presence of the ichnospecies  
86 *Torrowangea roseis* (Liñán & Tejero 1988) and *Cloudina*-like shelly fossils (Álvarez &  
87 Blanc-Valleron 2002) in the Paracuellos group (Fig. 2), together with the presence *in*  
88 *situ* of skeletal fossil *Cloudina* in the platform carbonates of the Ibor group, indicates a  
89 latest Ediacaran age for the both groups (Vidal *et al.* 1994). In the Domo de las Hurdes  
90 (Fig. 1A, DH), this upper part of the Ediacaran succession exhibits laminated black  
91 shales with an intercalation of alternations of limestone breccias and stratified  
92 sandstone± limestone couplets (Unit IV. Valladares *et al.* 2000), which were deposited  
93 from debris flows and turbidity currents in a mixed slope apron (Valladares 1995). In  
94 the Cantabrian Zone and IR, the basal Lower Cambrian detrital series were deposited in  
95 continental environments (Aramburu *et al.* 1992; Rubio-Ordóñez *et al.* 2004; Álvaro *et*

96 *al.* 2008), while in the Central Iberian Zone the equivalent series were mainly deposited  
97 in slope and base-of-slope environments (Valladares *et al.* 2000).

98 The Herrería group (Fig. 2) predominantly consists of sandstones with some levels of  
99 conglomerates, shales and carbonates, with a thickness varying between 900 m in  
100 Barrios de Luna (Fig. 1B) and 1500 m in the Cangas de Narcea-Mieldes areas (Comte  
101 1959). The lower part of this group consists of conglomerates that locally displays  
102 volcanic clasts, and have a lenticular geometry (Fig. 1B), interpreted as a large  
103 deposition cone (Parga & Luque 1971). Palaeontological data have revealed two fossil  
104 assemblages that begin with the record of *Phycodes* (= *Treptichnus*) *pedum* 4 m above  
105 the angular unconformity of the base of the Herrería group (Palacios & Vidal 1992),  
106 followed by *Rusophycus* and *Cruziana* species, indicating that this member yielded  
107 associations of ichnofossils suggestive of a Late Corduban age (Liñan *et al.* 2002).

108 The Lower Cambrian begins with conglomerate and sandstone-shale alternations  
109 deposited in alluvial and braided-channel environments at the base, which evolved  
110 upwards to tidal environments. In a restricted area close to Mieldes (Fig. 1B) this  
111 conglomerate has completely rubefacted angular clasts of shales, and well-rounded  
112 quartzite and volcanic clasts of undeformed pyroclastic rocks, mostly rhyolites and  
113 dacites with sericitised K-feldspar, displaying a coating of rubefaction (Rubio-Ordóñez  
114 *et al.* 2004, 2015). Some of these volcanic clasts contain zircons whose morphology is  
115 typical of zircons observed in alkaline magmas (Rubio-Ordóñez *et al.* 2006; Rubio-  
116 Ordóñez 2010). The latter authors suggested a change in the tectonic conditions from  
117 compressive to extensional at the end of the Ediacaran to explain this alkaline  
118 volcanism. However, the volcanic clasts were intensely affected by hydrothermal  
119 processes that did not affect the other components of the conglomerate, and the  
120 geochemical data do not represent the original magmatic composition (Rubio-Ordóñez  
121 2010; Rubio-Ordóñez *et al.* 2015). In the rest of the Cantabrian Zone, at the sites where  
122 the base of the Herrería group outcrops the conglomerates lack volcanic clasts, and  
123 well-rounded clasts of white quartz prevail. The source area of the Lower Cambrian  
124 detrital series was to the east of the present location of the Cantabrian Zone (Aramburu  
125 *et al.* 1992 and references therein) but post-Cambrian sediments cover this area, and no  
126 other Ediacaran outcrops are observable. In the IR and Sierra de la Demanda (Fig. 1A,  
127 SD), the basal Lower Cambrian also has conglomerate beds but no volcanic clasts have

128 been reported; instead, subrounded clasts of white quartz prevail (Álvaro *et al.* 2008;  
129 Ábalos *et al.* 2011).

### 130 **Mineralogy**

131 XRD semi-quantitative data on the rock-forming minerals in the rubefacted and  
132 unaltered Ediacaran and the Cambrian shales reveal that albite and chlorite are relatively  
133 abundant only in the Ediacaran shales but albite is absent in the equivalent rubefacted  
134 and Cambrian shales while chlorite has varying but lower contents in these two groups  
135 of shales than in the Ediacaran ones (Table 1). Some samples have hematite ( $\leq 5\%$ ), this  
136 mineral being more abundant in the rubefacted shales, and only samples BEL-2 and  
137 BEL-9, both corresponding to the Cambrian, contain K-feldspar (12% and 4%,  
138 respectively). The Cambrian samples can be separated into two sub-groups: those  
139 having heterogeneous mineral contents and those with homogeneous contents (labels:  
140 BEL-2, 9 and LR-3, and BEL-3, PS-2, and 6, respectively) that correspond to different  
141 positions in the stratigraphic series (see below). Figure 3 shows the enrichment in illite  
142 and depletion in chlorite of the Ediacaran rubefacted shales relative to the unaltered  
143 ones, together with the overlapping plot of the rubefacted and Cambrian shales.

144

### 145 **Stratigraphy**

146 Nearly 2700 m of stratigraphic logs from five partial sections of the Ediacaran  
147 succession (Narcea group) in the Cantabrian Zone were studied. This succession is  
148 several hundred metres thick and consists of siliciclastic rocks, predominantly  
149 sandstones and shales, and less abundant paraconglomerates, of Late Ediacaran age. The  
150 lithological uniformity in all partial sections hindered the division of this succession  
151 into lithostratigraphic units. These lithologies form two types of sequence. The  
152 predominant one is a thinning- and fining-upward sequence, 8 to 12 m thick and  
153 occasionally up to 30 m, from paraconglomerates to shales that are interpreted as a  
154 channel-fills deposited by different sediment gravity flows (e.g. debris flows and high-  
155 and low-concentration turbidity currents) and vertical settling from the suspension cloud  
156 of the turbidity currents. The frequent presence of a more or less thick lag over a basal  
157 erosive surface is interpreted as a channel lag deposit and indicates that its deposition

158 could be coeval with the period of maximum activity in the channel and aggradation of  
159 the levees due to overflow of the finer sediments before their silting (Torres *et al.* 1997).

160 The other type consists of thickening- and coarsening-upward sequences, generally  
161 0.7 to 5 m thick. From bottom to top, these sequences have basal shales, followed by  
162 alternations of sandstone-shale and massive sandstones at the top. This facies  
163 association corresponds to lobe deposits. However, in the present case this type of  
164 sequences is scarce, and is associated with fine-grained facies (shales and sandstone-  
165 shale alternations) and hence interpreted as overbank deposits formed in interchannels  
166 areas adjacent to the channels during their filling phases as small lobes in the inner  
167 deep-sea fans.

168 In general, according to the classification of Folk (1980), the Ediacaran sandstones of  
169 the Cantabrian Zone are medium-grained litharenites consisting mostly of moderately-  
170 to poorly-sorted grains of angular quartz (35-43%), chert (19-35%), the clasts of black  
171 shales, sandstones and feldspars being less abundant (< 15%) and isolated (< 1%)  
172 microclasts of igneous groundmass consisting of plagioclase laths and quartz.

173 More than 800 m of stratigraphic logs from five partial sections of the basal 200 m of  
174 the Lower Cambrian succession of the Herrería group were studied. This succession  
175 consists predominantly of siliciclastic rocks, shales being more abundant in the area of  
176 Vega de los Caballeros (Fig. 1B), where the Herrería group has a carbonate level about  
177 30 m thick displaying a lenticular geometry at about 40-50 m above its base.  
178 Contrastingly, in the area of Miedes-Parada la Vieja (Fig. 1B) conglomerates and  
179 sandstones predominate, and the carbonate level has not been identified.

180 Figure 4 shows two stratigraphic logs having and lacking, respectively, the carbonate  
181 level. In both sections, three intervals were distinguished. The lower one is 40 m thick  
182 and its lower boundary is an erosive surface in both logs. At Parada la Vieja, overlying  
183 this surface there are fining- and thinning-upward sequences that begin with  
184 paraconglomerates with a red sandy matrix and clasts of white quartzite, and rubefacted  
185 and black shales, followed by very coarse- and coarse-grained red sandstones, finishing  
186 with red shales (Fig. 4). The clasts are disorganised and are sometimes vertical (debris  
187 flows), and occasionally there is inverse grading and horizons of clasts parallel to the  
188 stratification within the sandstones. At Vega de los Caballeros this lower interval begins

189 with a thin level of microconglomerate with clasts of rounded white quartzite. The rest  
190 of the interval consists of green shales with some thin levels of coarse- and medium-  
191 grained sandstones with clasts of green shales, and lenses of fine-grained sandstones  
192 with a ripple lamination.

193 The middle interval is carbonated at Vega de los Caballeros. It consists of  
194 recrystallised sandy limestones forming coarsening- and thickening-upward sequences  
195 between 1.5 and 2.7 m thick (Fig. 4) that sometimes finish with ferruginous crusts and  
196 that can be interpreted as oolitic and bioclastic shoals in an inner high energy platform,  
197 with the possible emersion of the tops of some shoals. At Parada la Vieja, this middle  
198 interval is formed by siliciclastic thickening-upward sequences 2 to 3 m thick consisting  
199 of coarse-grained sandstones with a planar cross stratification that alternate with green  
200 shales. The sandstones show herring-bone and sigmoidal stratifications. All these  
201 sedimentological features can be interpreted as tidal shoals that have migrated in the  
202 present SE-NW direction ( $130^{\circ}$  and more scarce to  $300^{\circ}$ ) according to the paleocurrents  
203 (Fig. 4). Towards the top of the interval, these sequences alternate with fining-upward  
204 sequences 1.5 to 3.5 m thick, with a basal erosive surface, overlaid by  
205 microconglomerates with a normal grading and sandstones with trough and planar cross  
206 stratification, and the sequences finish with green shales. These sequences are  
207 interpreted as tidal channels that would have transported the sediments from the present  
208 NE to the SW ( $230^{\circ}$ ) according to the paleocurrents crossing the tidal shoals (Fig. 4). In  
209 this middle interval, there are frequent flaser, wavy and lenticular stratifications. All  
210 these features indicate a subtidal environment in an inner platform.

211 Finally, the upper interval is siliciclastic in both logs (Fig. 4). At Parada la Vieja, this  
212 interval is similar to the lower one, coarse-grained sandstones and conglomerates  
213 predominating in thickening- and coarsening-upward sequences of 2 to 7 m thick in the  
214 lower part, while the uppermost metres are thinning- and fining-upward sequences with  
215 an erosive basal surface. The conglomerates have clasts of up to 4 cm of white quartzite  
216 and red shales. The sequences are interpreted as bars and channels deposited in a fluvial  
217 environment. At Vega de los Caballeros it is possible to distinguish a lower part that is  
218 65 m thick in which grey shales with scarce lenticular fine-grained sandstones with  
219 linguoid ripples and middle- to fine-grained sandstones prevail in fining-upward  
220 sequences of 1 m thick overlying an erosive base. In the upper 20 m of the section,



221 shales are rare except at the top, where they are red, although coarse- and middle-  
222 grained sandstones with erosive base and channel geometry predominate. As at the  
223 Parada la Vieja, this upper part corresponds to a fluvial environment, while the pelitic  
224 lower part is interpreted as a low-energy marine environment with predominant settling.

## 225 **Geochemistry**

### 226 *Ediacaran shales and sandstones*

227 Contents of alkaline and alkaline earth elements and their ratios do not show significant  
228 differences in the Ediacaran shales or in the sandstones. The shales show rough negative  
229 covariation in SiO<sub>2</sub>-major element diagrams only for Al<sub>2</sub>O<sub>3</sub> and Fe<sub>2</sub>O<sub>3</sub>. Also, TiO<sub>2</sub>-  
230 P<sub>2</sub>O<sub>5</sub>, and TiO<sub>2</sub>-Zr diagrams define positive correlations for shales ( $r^2 = 0.75$  and  $0.77$ ,  
231 respectively;  $n = 20$ ) suggesting homogeneous proportions of Ti-minerals, phosphates  
232 and zircon in these rocks. The Sc contents in the sandstones are relatively high and  
233 show rough positive covariations with TiO<sub>2</sub>, Yb and Y. Sc may be stored in Ti- and Fe-  
234 and rare earth element (REE)- minerals (Schock 1975). Thus, a plausible interpretation  
235 of the relatively high Sc contents in the sandstones and these covariations is that  
236 recycling and sorting favoured accumulation of heavy minerals storing Sc. The shales  
237 also show these covariations. The shales and sandstones exhibit parallel REE patterns  
238 (Fig. 5a,b) and the Ediacaran shales from Cantabrian Zone and NIBAS (Neoproterozoic  
239 Iberian Average Shale. Ugidos *et al.* 2010) show the same mean REE patterns and those  
240 of the corresponding mean sandstones (Ugidos *et al.* 1997b) are also the same (Fig. 5c).  
241 Normalised to the composition of the upper continental crust (UCC, Rudnick & Gao  
242 2005; Hu & Gao 2008), the mean shales of the Cantabrian Zone and NIBAS define the  
243 same multielement patterns and also the two corresponding mean sandstones (Fig. 5d).  
244 All these rocks exhibit a strong depletion in Sr, and the sandstones show enrichment in  
245 Zr with respect to the shales.

### 246 *Sub-Cambrian alteration and K-metasomatism of the Ediacaran rocks*

247 The abundances of SiO<sub>2</sub>, TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and Fe<sub>2</sub>O<sub>3</sub> in the rubefacted shales lack relevant  
248 changes that could be interpreted as being related to oxide gain or loss with respect to  
249 those of the unaltered ones. However, the abundances of other oxides and elements such  
250 as MgO, CaO, Na<sub>2</sub>O, P<sub>2</sub>O<sub>5</sub>, Be and Sr are lower, and those of K<sub>2</sub>O, Rb, Cs, and Ge are  
251 higher in the rubefacted rocks. To quantify these differences, the model compositional

252 changes (MCC) of element concentrations in the rubefacted shales were calculated with  
253 respect to the unaltered ones by using the expression:  $MCC = [(E_{G1} - E_{G2}) / E_{G2}] * 100$   
254 (Páez *et al.* 2010), in which  $E_{G1}$  and  $E_{G2}$  are the median and mean element  
255 concentrations in the group of the rubefacted and unaltered rocks respectively (Table 2).  
256 The abundances of some oxides such as CaO and Na<sub>2</sub>O are below the detection limit in  
257 many rubefacted shales. Thus, the parameters involving these oxides were calculated  
258 accepting the values of the detection limits (0.1% and 0.05%, respectively) as the  
259 maximum contents.

260 Although in Table 2 there are some differences in the results depending on whether  
261 the calculations were made using the median or the mean values of element contents  
262 (e.g., Sc enrichment is 10.9% or 3.2% in the rubefacted shales depending on whether  
263 the median or the mean abundance is used; similarly, the Cs enrichment is 226% or  
264 166%), in general there is good agreement between both groups of calculations. The  
265 major differences are probably a consequence of the relatively low number of analyses  
266 and ranges of some element contents with no or scarce intermediate values. Thus, the  
267 median values of some elements are strongly dependent on the number of data in each  
268 extreme group of values, and hence the compositional change can be exaggerated  
269 (either higher or lower) when the mean and median differ substantially. To minimise  
270 this effect, any change, depletion or enrichment of an element is accepted only if the  
271 results from the both calculations are comparable.

272 The strong depletion of MgO, CaO, and Na<sub>2</sub>O, and the strong increase in K<sub>2</sub>O  
273 contents are consistent with the trend parallel to the chlorite-illite join defined by the  
274 Ediacaran unaltered and rubefacted shales (Fig. 3). The Rb and Cs contents of the  
275 rubefacted ones increase remarkably, probably incorporated to illite (McLennan *et al.*  
276 1990), and those of Be and Sr decrease due to leaching during weathering. Although  
277 some samples of the rubefacted shales have lower REE contents, most of their REE  
278 patterns (Fig. 5e) overlap those of the unaltered ones and it is unclear whether the  
279 calculated loss of REE really occurred or whether it might be due to the low number of  
280 samples. Whatever, the possible loss of REE did not cause significant Sm-Nd  
281 fractionation given that the unaltered (0.194-0.211; mean = 0.203; st. dev. = 0.01) and  
282 rubefacted shales (0.197-0.220; mean = 0.209; st. dev. 0.01) have almost identical  
283 Sm/Nd. Furthermore, the Nd-isotope results also suggest that Sm/Nd was not affected

284 by weathering (see below). Y, Zr, V, Nb, Ta, and Cr (not included in Table 2) do not  
285 show significant changes. A loss of U occurred probably due to the oxidation of  $U^{+4}$  to  
286 the more soluble  $U^{+6}$  (McLennan *et al.* 1993), and also of Th (Table 2) as is suggested  
287 by the lower mean Th/Sc ratio of the rubefacted shales (Table 3). Probably, acidic  
288 waters resulting from sulphide oxidation (S content in six shales: 0.01 to 0.9%,  
289 unpublished data) during weathering affected the Ediacaran shales, favoured the  
290 precipitation of hematite, released elements such as Co, Ni, Cu, Zn, Pb, and As, and  
291 also favoured the dissolution of phosphates and the release of REE, Th and U, and the  
292 desorption of Be from clays (Grew 2002; Ryan 2002; Harlavan *et al.* 2009; Åström *et*  
293 *al.* 2010). Germanium is an element sequestered by secondary clays during weathering  
294 (Scribner *et al.* 2006) and its content is increased by about 20% in the rubefacted shales  
295 (Table 2).

296 The rubefacted shales are the only ones that show loss of Ca and Na, and a  
297 remarkable increase in K. This suggests a process of K-metasomatism related to  
298 weathering and rubefaction. A process of K-metasomatism may have different causes  
299 and affect siliciclastic and volcanic rocks, but in all cases the source of K is local rather  
300 than from externally derived fluids, as demonstrated by Hutcheon *et al.* (1998). In  
301 general, K-metasomatism is associated with extensional crustal settings and increases  
302 the illite contents in the siliciclastic rocks affected, causes Rb enrichment, and Ca, Na,  
303 Mg and Mn depletion (Beratan 1999; Ennis *et al.* 2000; Páez *et al.* 2010), as occurred in  
304 the area studied. Low-temperature K-rich brines resulting from the evaporation of  
305 marine waters in restricted environments and the percolation of basinal fluids have been  
306 proposed as causes of metasomatising processes (Munz *et al.* 1995; Leising *et al.* 1995;  
307 Ennis *et al.* 2000; Sandler & Harlavan 2006). Thus, it is possible that during the Upper  
308 Ediacaran sea-level fall marine water might have remained isolated in coastal  
309 environments and evolved to K-rich brines, which could have resulted in the K-  
310 metasomatism and rubefaction of the Ediacaran rocks of the Cantabrian Zone during the  
311 intense worldwide weathering processes in sub-Cambrian times.

### 312 ***Cambrian shales***

313 Two groups of Cambrian shales can be established (see also Table 1): those underlying  
314 and those overlying the carbonate level, referred to below as the underlying and  
315 overlying shales, respectively. The underlying and rubefacted shales show similar

316 ranges of SiO<sub>2</sub> contents; they lack or have minor contents of MnO, CaO and Na<sub>2</sub>O; they  
317 have relatively high contents of K<sub>2</sub>O, Rb, Cs and Ge, and most of the mean element  
318 ratios of both groups of shales are the same (Table 3). The rubefacted shales have the  
319 same Zr/Nb ratios as their parent rocks but increased Rb/Zr ratios that overlap those of  
320 the Cambrian shales (Fig. 6). These geochemical similarities, together with the presence  
321 of angular clasts of the rubefacted shales in the basal Cambrian conglomerate, reflect  
322 the contribution of the Ediacaran rocks to the Cambrian series. However, the underlying  
323 shales also have higher contents of Al<sub>2</sub>O<sub>3</sub>, K<sub>2</sub>O, REE and Th, and lower Fe<sub>2</sub>O<sub>3</sub> and  
324 MgO, and higher K<sub>2</sub>O/Al<sub>2</sub>O<sub>3</sub> ratios than the Ediacaran shales. The only major element  
325 covariations defined by the underlying shales are SiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub>, SiO<sub>2</sub>-K<sub>2</sub>O and Al<sub>2</sub>O<sub>3</sub>-K<sub>2</sub>O  
326 ( $r^2 = 0.98, 0.92$  and  $0.95$ , respectively;  $n = 5$ ). These results together, strongly support  
327 the notion that the increase in Al<sub>2</sub>O<sub>3</sub> and K<sub>2</sub>O contents, and in the K<sub>2</sub>O/Al<sub>2</sub>O<sub>3</sub> ratio  
328 (range: 0.353-0.386) is due to the relative increase in detrital illite derived from K-  
329 metasomatised and weathered Ediacaran rocks, although it is also possible that this high  
330 ratio may be partially due to the presence of undetected quantities of K-feldspar. The  
331 lower Fe<sub>2</sub>O<sub>3</sub> and MgO contents of the underlying shales are probably related to the  
332 illitisation of the chlorite of the rubefacted shales (Fig. 3), the leaching of Mg and a  
333 redistribution and loss of hematite during the sub-Cambrian weathering and erosion.  
334 The underlying shales also have higher contents of Be, REE and Th and a lower Eu/Eu\*  
335 ratio than the rubefacted ones probably due to the contribution of the Ediacaran  
336 weathered igneous rocks.

337 The overlying shales are characterised by lower contents of SiO<sub>2</sub> and higher contents  
338 of the other major elements, and higher Eu/Eu\* and (La/Yb)<sub>N</sub> ratios, but also lower  
339 contents of Rb, Cs and Zr. Mean element ratios such as Al<sub>2</sub>O<sub>3</sub>/TiO<sub>2</sub>, Ti/Nb, Th/Sc and  
340 Zr/Hf and others in Table 3 are the same in both groups of shales. Thus, the underlying  
341 and overlying shales share some geochemical features while other features, such as the  
342 lower contents of Rb, Cs and Zr, the higher contents of light REE and the relatively low  
343 Rb/Sr, Cs/Be and Zr/Nb ratios (Table 3) in the overlying shales, indicate a decreasing  
344 proportion of detritus from the rubefacted rocks and an increasing contribution of a new  
345 component that also supplied K-feldspar to these shales (Table 1). Plausibly, the higher  
346 Eu/Eu\* ratio of the overlying shales could be due to the presence of K-feldspar and the  
347 higher (La/Yb)<sub>N</sub> ratio could be related to a higher content of light REE and lower  
348 content of zircon, as suggested by the lower Zr of these shales.

### 349 *Nd isotopes*

350 The data in Table 4 strongly suggest a higher proportion of juvenile material in the  
351 Ediacaran and in the above-mentioned underlying shales than in the overlying ones. The  
352  $\epsilon_{Nd}(t)$ - $f_{Sm/Nd}$  diagram has the potential to identify mixtures of detritus derived from  
353 upper-crustal rocks and juvenile contributions coeval with sedimentation which would  
354 define sub-horizontal to diagonal trends. Actually, the shales show little evidence of  
355 correlation in the  $\epsilon_{Nd}(541)$ - $f_{Sm/Nd}$  diagram (Fig. 7a) approaching vertical arrays.  
356 Redistribution of REE could result in vertical arrays (Bock *et al.* 1994) but the range of  
357  $^{147}Sm/^{144}Nd$  (Table 4) within the various groups of shales, even the rubefacted ones, are  
358 narrow suggesting that no substantial fractionation of REE affected these rocks during  
359 exogenic processes (Jahn & Condie, 1995).

### 360 **Discussion and conclusions**

#### 361 *Provenance of the Ediacaran and Cambrian detrital series of the Cantabrian Zone:* 362 *major and trace elements and Nd-isotope results*

363 Mafic rocks contributing to siliciclastic sediments increase the contents of Sc, Cr and  
364 plagioclase while those of La and Th increase if the source rocks are felsic (Condie &  
365 Wronkiewicz, 1990; Eriksson *et al.* 1992; Crichton & Condie, 1993; McLennan *et al.*  
366 1993, 1995, 2006; Feng *et al.* 1993; Cox *et al.* 1995). According to these authors, the  
367 abundance of volcanic lithic fragments in sandstones, the relatively low values of La/Sc,  
368 La/Cr, Th/Cr and Th/Sc, and the high Eu/Eu\* ratios are expected in fine-grained  
369 siliciclastic rocks resulting directly from active margins and related settings, and  
370 considerable petrological and geochemical variability is also expected. By contrast, the  
371 homogeneity reflects the recycling and buffering of mixtures of different sources  
372 (Garrels, 1988; McLennan, 1989; Cullers, 1994; Brown *et al.* 2003; McLennan *et al.*  
373 2006). Moreover, sedimentary recycling causes an enrichment of Zr abundance in  
374 sandstones relative to contents in related fine-grained sediments, while there are no  
375 systematic differences in the sediments of active tectonic settings (McLennan *et al.*  
376 1990). Also, relatively high  $K_2O/Na_2O$  and Rb/Sr ratios, and high CIA (Nesbitt &  
377 Young, 1982) values are expected for shales (McLennan *et al.* 1990, 1993). Another  
378 useful diagram, Th/Sc- $\epsilon_{Nd}(t)$ , relates the composition of the sediments to the mean  
379 provenance age (McLennan *et al.* 1993) and potentially reflects the contribution of

380 felsic or mafic components to detrital rocks.

381 The Ediacaran shales studied have relatively high  $K_2O/Na_2O$  (1.24-2.81) and Rb/Sr  
382 (0.85-2.49) ratios and CIA values (65.10-73.16). The sandstones are enriched in Zr  
383 (mean: 257, range, 196-321, s.d. 44) with respect to the shales (197, 174-231, s.d. 16).  
384 The Eu/Eu\* ratios of the 80% of shales are lower than 0.70 and their mean Eu/Eu\* ratio  
385 (0.67) is lower than those of the UCC (Table 3). The Ediacaran shales of the Cantabrian  
386 Zone and NIBAS share element ratios (Table 3) regardless of whether such ratios (e.g.,  
387  $Al_2O_3/TiO_2$ , Ti/Zr, La/Sc, Zr/Sc, Zr/Nb, and Th/Nb) are dependent upon mineral  
388 proportions, or on diagenesis or other processes that could affect mobile/immobile  
389 element ratios (e.g., Rb/Th and Rb/Zr). Furthermore, the element ratios such as La/Cr,  
390 La/Sc and Th/Sc among others used as provenance indicators, are the same as those of  
391 the UCC (Table 3). The two mean sandstones also share most trace element ratios and  
392 both the shales and sandstones have the same mean Ti/Nb, La/Cr, Cr/Th and Th/Nb  
393 ratios (Table 3). The Ediacaran and Cambrian shales show little evidence of correlation  
394 in the  $\epsilon_{Nd}(541)-f_{Sm/Nd}$  and Th/Sc- $\epsilon_{Nd}(541)$  diagrams (Fig. 7a, b). Also, the  $\epsilon_{Nd}(541)$   
395 values of the Ediacaran shales presented here and those from the Central Iberian Zone (-  
396 3.8 to -1.8, Tassinari *et al.* 1996; Ugidos *et al.* 1997a, 2008, recalculated to 541Ma) are  
397 the same. Therefore, the Ediacaran detrital rocks have a remarkable geochemical  
398 homogeneity in these zones.

399 The geochemical uniformity is not compatible with a widespread volcanism coeval  
400 with the sedimentation of detrital successions and none of the criteria described above  
401 used to define a coeval contribution of magmatic arcs to detrital sediments are fulfilled  
402 by the Ediacaran detrital series studied here. Therefore, these series derived from  
403 recycled and homogenised mixtures of different materials. However, the relatively high  
404  $\epsilon_{Nd}(541)$  of the Ediacaran shales requires a juvenile contribution to these detrital  
405 sediments. Thus, it is concluded that this juvenile contribution occurred in a geological  
406 setting different from the one in which the Ediacaran series were finally deposited after  
407 the recycling and homogenisation of all the components. An immature passive margin is  
408 the most plausible setting for the basin of the Ediacaran detrital series.

#### 409 ***The Cambrian series: inheritance from the Ediacaran detrital and volcanic rocks***

410 The Cambrian conglomerate overlying the Ediacaran-Cambrian angular

411 unconformity has completely rubefacted angular clasts of shales and well-rounded  
412 volcanic clasts displaying a coating of rubefaction. The sources of these materials were  
413 the Ediacaran series and volcanic rocks affected by the worldwide weathering process in  
414 Late Ediacaran times ( $542.62 \pm 0.38$  Ma, Parnell *et al.* 2014). The volcanic clasts in the  
415 Cambrian conglomerate have been found in a relatively restricted area close to Miedes  
416 (Fig. 1), but no volcanic rocks have been found in the Ediacaran successions of the  
417 Cantabrian Zone or in the other basal Cambrian conglomerates of the region studied in  
418 this work. Also, in the IR and SD (Fig. 1) the Ediacaran successions lack igneous rocks,  
419 and no volcanic clasts from the basal Cambrian conglomerates have been reported  
420 (Álvaro *et al.* 2008; Ábalos *et al.* 2011, 2012). It may be inferred that the volcanic clasts  
421 in the basal conglomerate of the Cantabrian Zone resulted from volcanic activity of  
422 local importance, probably related to extensional faults in the Late Ediacaran.

423 Based on the observations presented here, the basal Cambrian series mostly consist  
424 of detritus derived from the Ediacaran rocks underlying the angular unconformity but  
425 the shales overlying the carbonate level also received a significant contribution of a  
426 component that supplied K-feldspar and increased the abundances of LREE and  
427 lowered the  $\epsilon_{Nd}(541)$  values. Therefore, the contribution of the inherited Ediacaran  
428 detritus to these shales decreased and the proportion of more evolved detrital  
429 components increased. Two main possibilities, perhaps coeval, could explain the  
430 geochemical change: a) The rise in sea-level gradually covered the outcropping  
431 Ediacaran rocks, which would have decreased their contribution to the Cambrian  
432 sediments; b) A decrease in the relative extension of Ediacaran rocks at the source due  
433 to erosion and peneplanation and an increase in subaerial exposures of more evolved  
434 compositions.

435 The Cambrian shales of the IR exhibit an increase in the mean contents of illite (from  
436 39 to 59%) and  $K_2O$  (4.0 to 5.6%) and a decrease in the mean contents of chlorite (15 to  
437 5%) and feldspar (16 to 5%),  $CaO$  (1.0 to 0.6%) and  $Na_2O$  (1.9 to 0.4%) with respect to  
438 those of the Upper Neoproterozoic (Table 1, in Bauluz *et al.* 2000). Furthermore, the  
439 same geochemical parameters that discriminate the Ediacaran from Cambrian shales in  
440 the Central Iberian Zone also separate the equivalent shales in the IR (Valladares *et al.*  
441 2002b).

442 ***Paleogeography***

443 On the basis of the previous data mentioned in Geological Setting and of the  
444 stratigraphic data reported in this study, it is suggested that the upper part of the  
445 Ediacaran successions (Paracuellos and Ibor groups, Fig. 2) eastwards of the Iberian  
446 Massif (IR and MT, Fig.1) were deposited in a platform environment (Calvet & Salas,  
447 1988; Álvaro & Blanc-Valleron, 2002; Álvaro *et al.* 2008) while in the rest of the  
448 Central Iberian Zone the upper part of the Ediacaran succession (unit IV, Fig. 2) was  
449 deposited in slope and base-of-slope environment (Valladares *et al.* 2000) during Late  
450 Ediacaran times. However, in the Cantabrian Zone the Ediacaran succession was  
451 deposited in deep-sea fans (Pérez Estaún, 1978; this work) but geological processes  
452 such as extensional tectonics, differential block subsidence, and the development of  
453 high-relief areas favoured the erosion of the materials equivalent to those of Paracuellos  
454 and Ibor groups and the Unit IV, which have carbonate beds and *Cloudina*. Thus, these  
455 beds are absent in the Narcea succession although the remaining part is also of Late  
456 Ediacaran age (< 553 Ma). Therefore, it is unclear whether a shallow or a deep  
457 environment was present in the Cantabrian Zone in the latest Ediacaran times. In the  
458 Cantabrian Zone and IR, the basal detrital successions of the Herrería group and  
459 Bámbola formation (Fig. 2) were respectively deposited on alluvial plain sequences and  
460 braided channel environments (Aramburu *et al.* 1992; Rubio-Ordoñez *et al.* 2004;  
461 Álvaro *et al.* 2008) during the Early Cambrian, while the basal Cambrian succession in  
462 the Central Iberian Zone was mostly deposited in slope and base-of-slope environments  
463 (Valladares *et al.* 2000). Therefore, the source area of the Iberian basin was placed at  
464 the present ENE in the Late Ediacaran. During the Early Cambrian, the continental  
465 environments were also towards the present ENE and the deep basin towards SW. In the  
466 both cases the source area of the Iberian basin would have been located at the present  
467 NE.

#### 468 ***Geological processes around the time of the Ediacaran-Cambrian boundary***

469 According to the data presented here, the sequence of geological processes in the  
470 Cantabrian Zone close to the Ediacaran-Cambrian boundary is synthesised in Fig. 8 as  
471 follows:

472 a) The Narcea group was deposited in channels and small lobes in the interchannel areas  
473 of the inner deep-sea fans. This group is younger than  $553 \pm 4$  Ma (age of the youngest  
474 detrital zircons, Fernández Suárez *et al.* 2014).



475 b) After the sedimentation of the Narcea group and before the sub-Cambrian weathering  
476 ( $542.62 \pm 0.38$  Ma, Parnell *et al.* 2014), extensional tectonics and volcanism associated  
477 with related faults occurred. The morphological types of the zircons in the volcanic  
478 clasts from the basal Cambrian are typical of alkaline magmas (Rubio-Ordóñez *et al.*  
479 2006). Thus, the general geological setting is similar to that seen in the northern margin  
480 of Gondwana (see above).

481 c) The extensional regime resulted in differential block subsidence and high-relief areas  
482 (Valladares *et al.* 2002a; Rubio-Ordóñez *et al.* 2015). Then, the erosion of the Ediacaran  
483 materials began, and probably initiated the sea-level fall around the Ediacaran-  
484 Cambrian boundary, followed by the sub-Cambrian weathering of the Ediacaran rocks  
485 (Fig. 8).

486 d) During the sea-level fall weathering would have caused rubefaction and the K-rich  
487 brines resulting from the evaporation of marine water in restricted environments would  
488 have percolated and caused K-metasomatism of the Ediacaran rocks. The sub-Cambrian  
489 erosion would have favoured peneplanation and transport, and the rolling of volcanic  
490 clasts that favoured generation of their rubefacted coatings. Possibly, these processes  
491 continued during most or all of the early Corduban Series of the West Gondwana  
492 Standard (Geyer & Landing, 2004), equivalent to the Fortunian Stage of the  
493 International Chronostratigraphic Chart (Cohen *et al.* 2013), since *Rusophycus* (Jensen  
494 *et al.* 2010) appears close to *Treptichnus pedum* (Fig. 8) almost at the base of Herrería  
495 group. Thus, the sedimentation of the Herrería group began at around 529 Ma.

496 e) The Cambrian succession in the Cantabrian Zone began with relatively thick  
497 conglomeratic deposits related to a forced regression resulting from the sea-level fall of  
498 the latest Ediacaran and the extensional activity as indicated by the presence in Miedes  
499 (Fig. 1) of alluvial fans with volcanic and other rubefacted clasts filling valleys bounded  
500 by faults (Fig. 8).

501 f) The gradual sea-level rise began in Early Cambrian times and reached the Cantabrian  
502 Zone during the sedimentation in thickening-upward sequences of the carbonated level,  
503 or its siliciclastic equivalent, which is interpreted as the maximum flooding surface. The  
504 presence of acritarchs (Fig. 2, in Jensen *et al.* 2010) above these levels indicates the top  
505 of the Corduban or Terrenewian Series, and hence gives an age of around 521 Ma.

506 Underlying and overlying these levels there are geochemical differences suggestive of a  
507 gradual increasing upward contribution of older crustal materials.

#### 508 **Acknowledgements**

509 Financial support for this work was provided by the CGL2007-60035BTE and  
510 CGL2010-18905BTE projects of the MICINN (Spain). Comments by C. M. Fedo and  
511 G. Blanco improved the manuscript.

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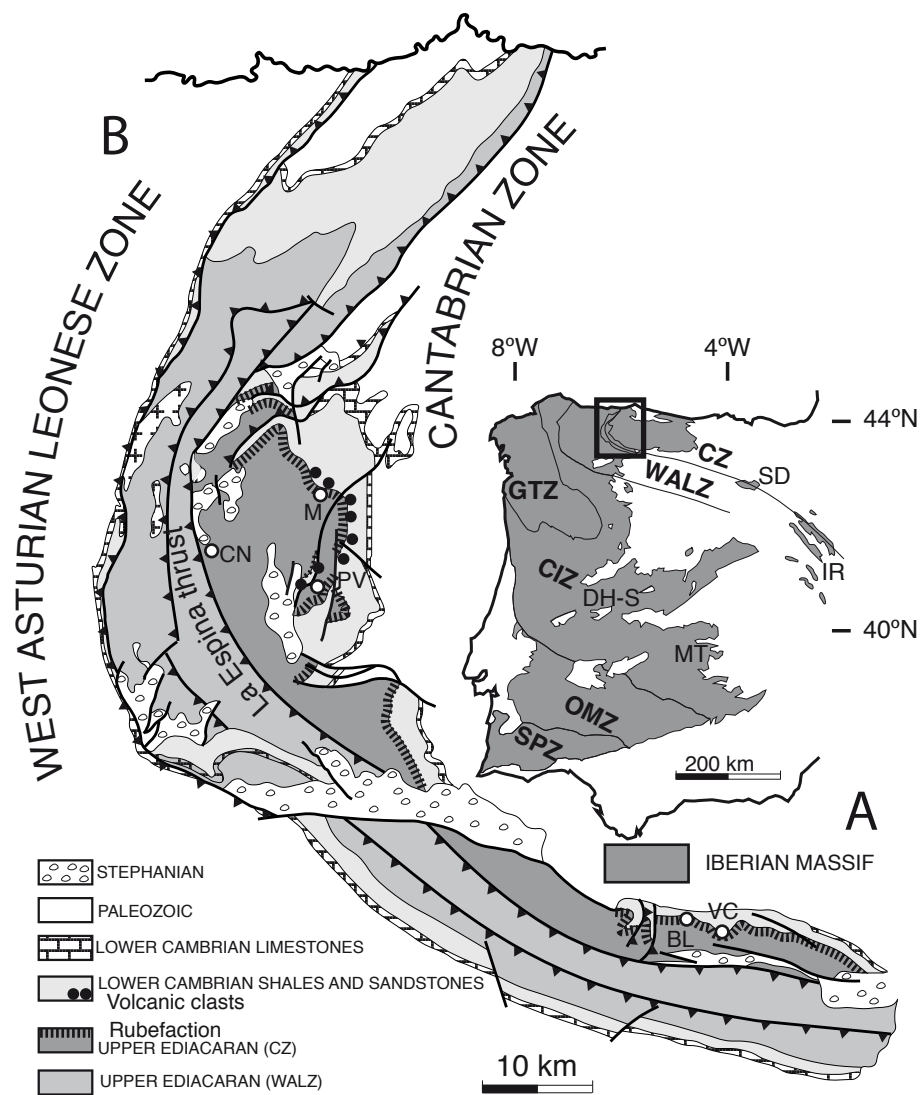
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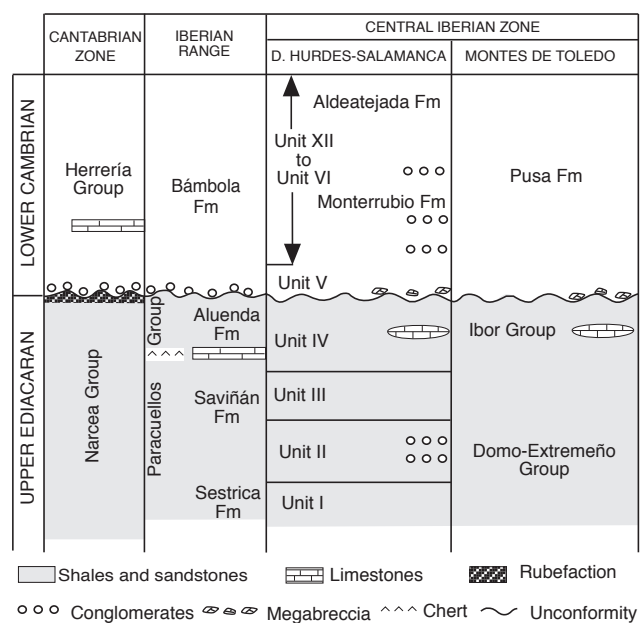
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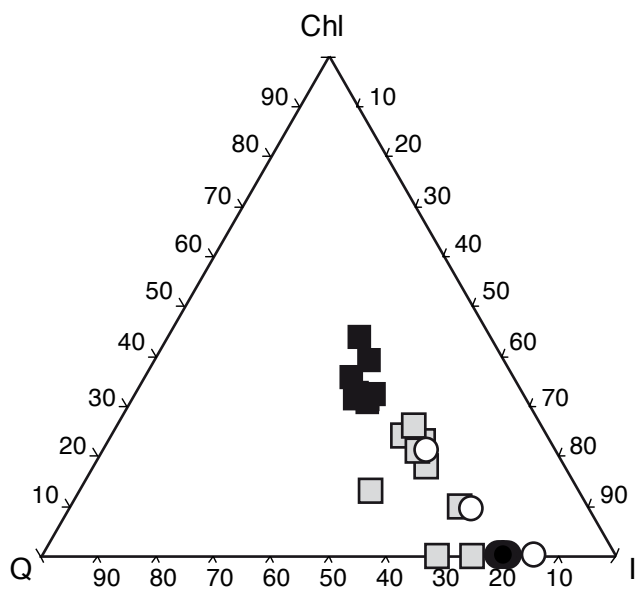
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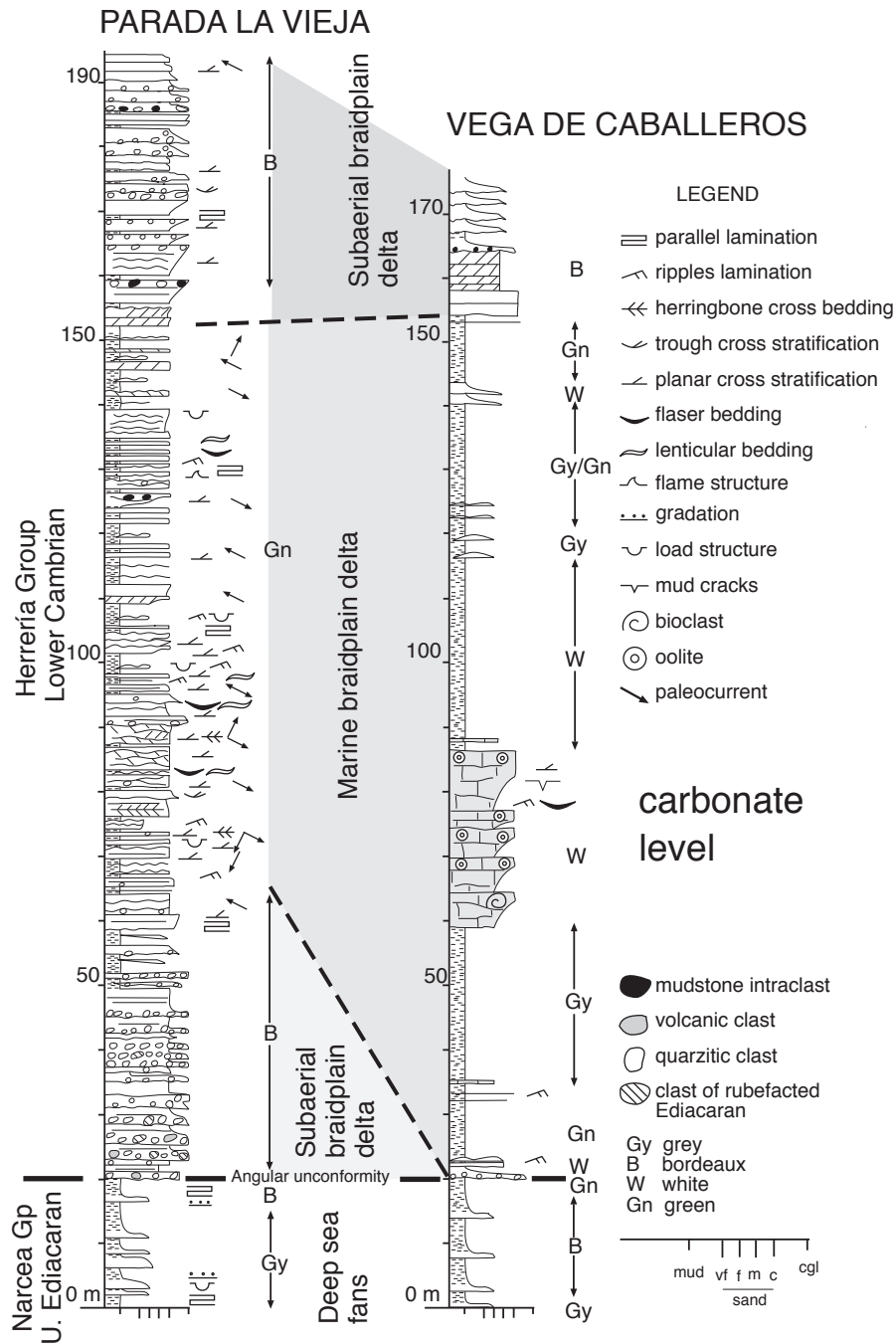
**Fig. 1.** (A) Zoning of the Iberian Massif, CZ: Cantabrian Zone; WALZ: West Asturian-Leonese Zone; GTZ: Galicia Tras os Montes Zone; CIZ: Central Iberian Zone; OMZ: Ossa Morena Zone; SPZ: South Portuguese Zone; IR: Iberian Range; SD: Sierra de la Demanda; DH-S: Domo de las Hurdes-Salamanca; MT: Montes de Toledo. (B) Geological map of the Narcea antiform in the Iberian Massif. BL: Barrios de Luna, CN: Cangas del Narcea, M: Mieldes, PV: Parada la Vieja, VC: Vega de los Caballeros.



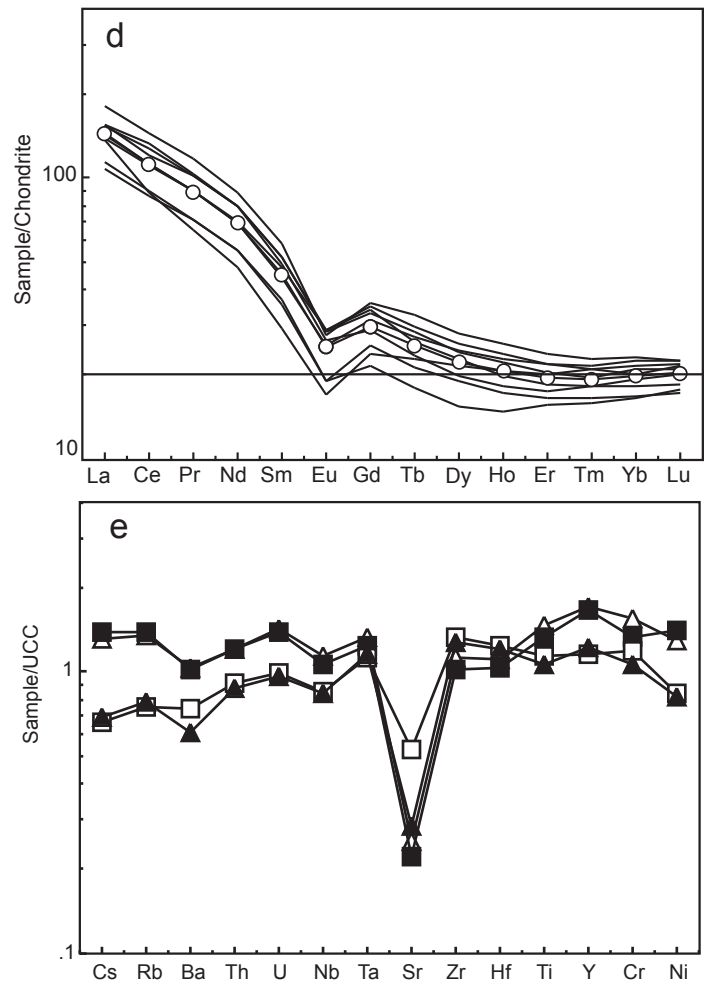
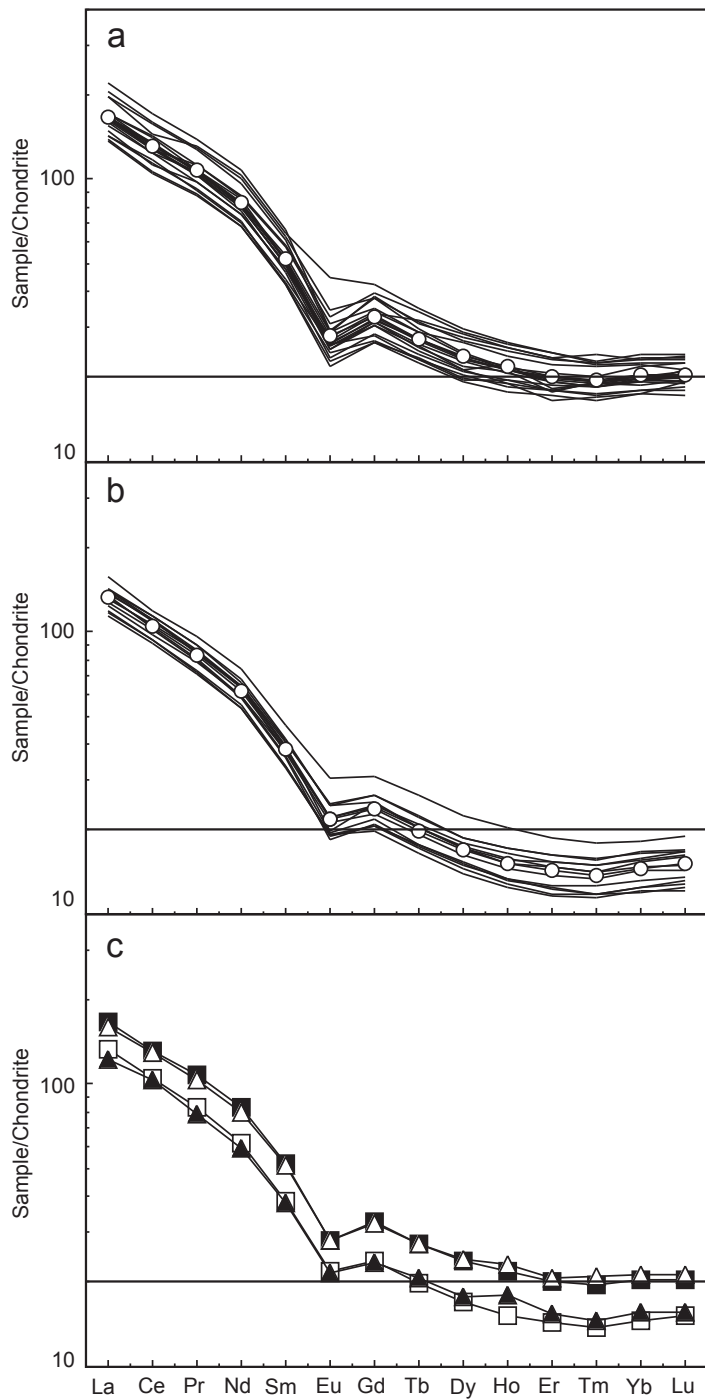
**Fig. 2.** Simplified correlation chart of the lithostratigraphic units of the Late Ediacaran-Early Cambrian of the Cantabrian, Central Iberian zones and Iberian Ranges (Valladares et al. 2002a and references therein).



**Fig. 3.** Plot of the main mineral components of the shales studied. The diagram separates the Upper Ediacaran shales (filled squares) from the rubefacted equivalents (grey squares) and Lower Cambrian (filled and open circles: underlying and overlying shales, respectively. See text). Chl: chlorite; I: illite; Q: quartz.

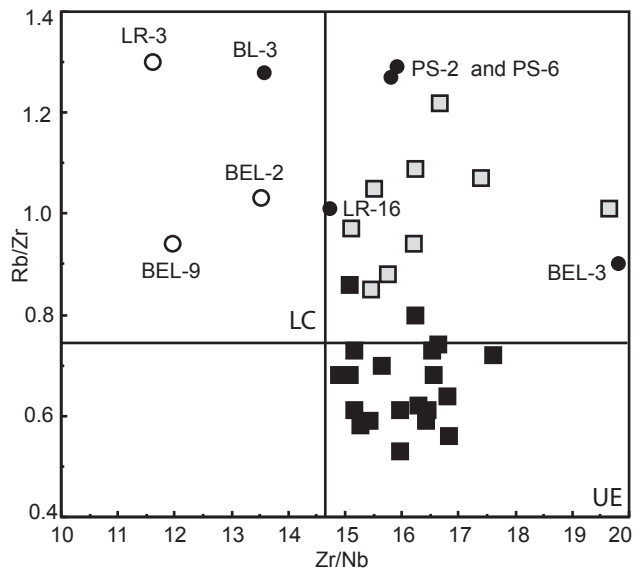


**Fig. 4.** Stratigraphic logs of the basal Herería group in the two sections studied: Parada la Vieja (PV, Fig. 1B) and Vega de Caballeros (VC, Fig. 1B) showing that the carbonated level present in VC is absent in PV. Both sections have three intervals showing braidplain delta features.

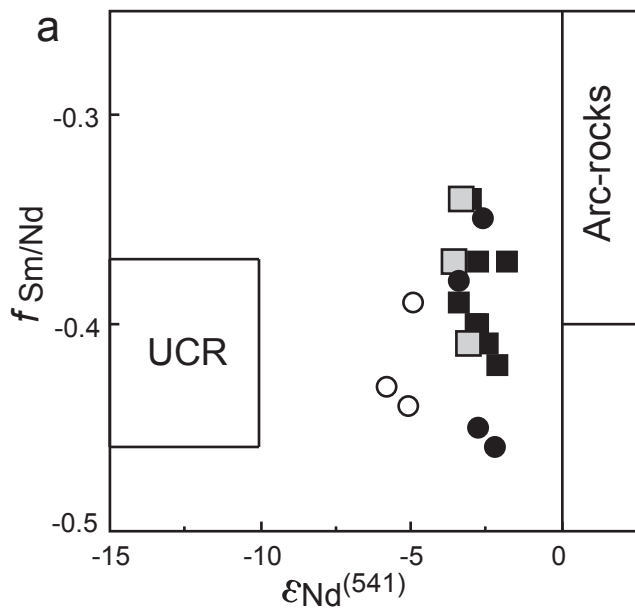


**Fig. 5.** Chondrite normalised REE patterns for the Ediacaran shales (a), sandstones (b), NIBAS and the mean (n=16) sandstone (c) of the Central Iberian Zone. Upper Continental Crust (UCC) normalised multi-element diagram (d). Chondrite normalised REE patterns for the rubefacted shales (e). Circles: mean patterns. Filled and open squares: mean Ediacaran shales and sandstones, respectively. Open and filled triangles: NIBAS and the mean sandstone, respectively.

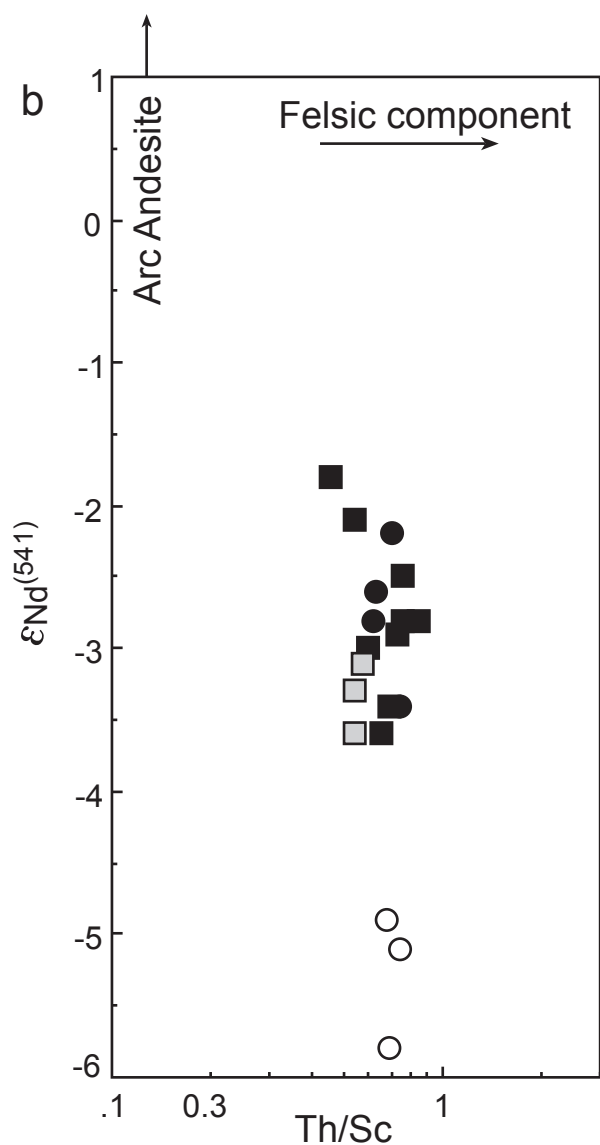




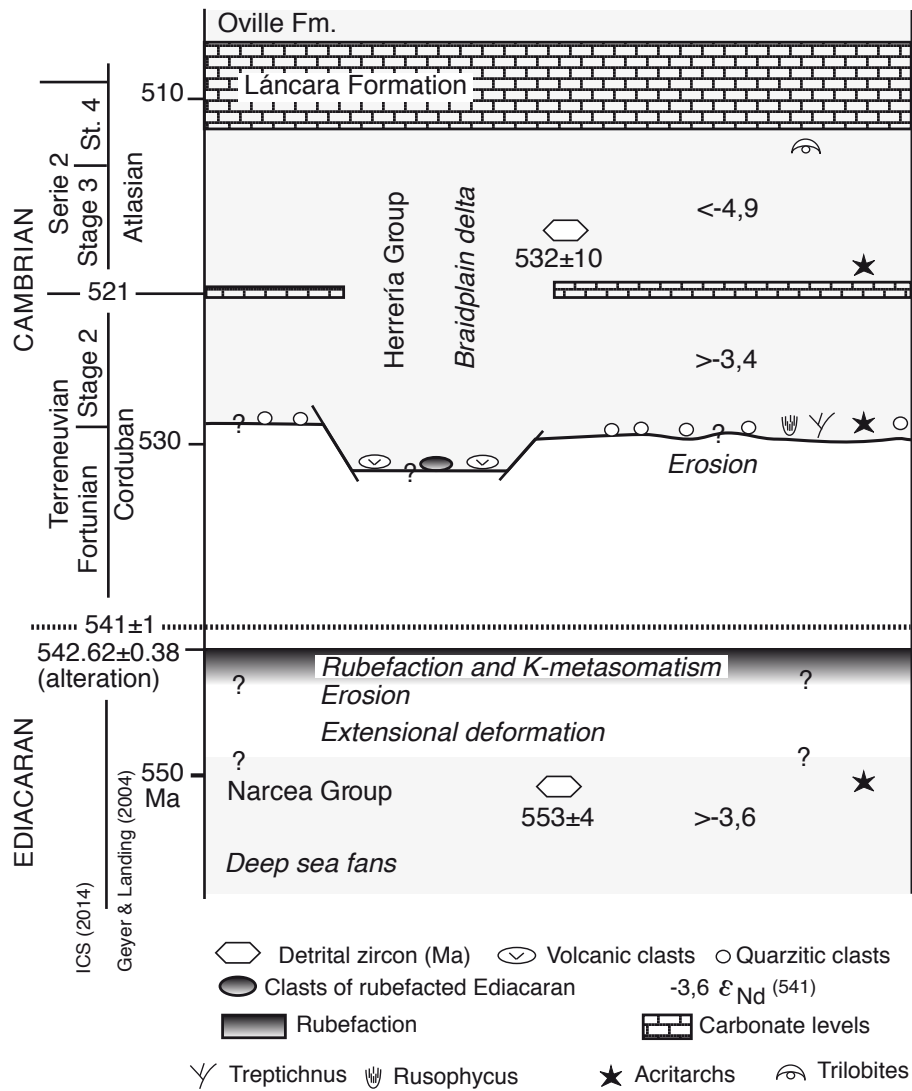
**Fig. 6.** Diagram Zr/Nb-Rb/Zr. The all the Ediacaran shales (filled squares), including the rubefacted equivalents (grey squares), show relatively high Zr/Nb ratios characteristic of the Ediacaran shales in the Central Iberian Zone (field UE). However, only the Lower Cambrian shales overlying the carbonate level (samples LR-3, BEL-2, BEL-9) plot in the Cambrian field (LC) while all the samples underlying this level (filled circles) but one show Zr/Nb ratios typical of the Ediacaran field but Rb/Zr ratios as high as those of the rubefacted shales. See text. UE and LC: fields of the Upper Ediacaran and Lower Cambrian shales in the Central Iberian Zone (Valladares et al. 2002b). Symbols as in Fig. 3.



**Fig. 7a.**  $\epsilon\text{Nd}(t)$ - $f\text{Sm}/\text{Nd}$  diagram (after Bock et al. 1994, simplified). Plots of Ediacaran and underlying Lower Cambrian shales are consistent with the inheritance of homogenised mixtures of upper-crustal and juvenile compositions and the increase of the proportion of the upper crustal rocks in the overlying Lower Cambrian shales. UCR: Upper-crustal rocks. Symbols as in Fig. 3.



**Fig. 7b.**  $\text{Th}/\text{Sc}$ - $\epsilon\text{Nd}(t)$  diagram (McLennan et al. 1993) consistently favouring the interpretation that Ediacaran detrital rocks would consist of inherited homogenised mixtures of juvenile and crustal components and a higher proportion of the latter ones in the Lower Cambrian sediments. Symbols as in Fig. 3.



**Fig. 8.** Chronostratigraphic framework of the Ediacaran-Lower Cambrian boundary in the Cantabrian Zone and sedimentary, geochemical and tectonic processes produced. Note that the vertical scale indicates time, not thickness.

**Table 1.** Mineral (%) semi-quantitative analyses (XRD) of the shales studied

Sample	Group	Quartz	Albite	Chlorite	Illite	K-feldspar	Other
BEL-2	OLC	12	0	0	76	12	-
BEL-9	OLC	19	0	9	68	4	-
LR-3	OLC	22	0	22	57	-	-
BL-3	ULC	20	0	0	80	-	-
PS-2	ULC	18	0	0	82	-	-
PS-6	ULC	19	0	0	81	-	Hematite
BEL-23	RUE	21	0	23	56	-	Hematite
BEL-7	RUE	24	0	0	76	-	Hematite
BL-5	RUE	35	0	13	52	-	Hematite
BL-6	RUE	21	0	26	52	-	Hematite
LR-14	RUE	23	0	21	56	-	Hematite
LR-15	RUE	21	0	10	69	-	Hematite
PS-16	RUE	30	0	0	70	-	Hematite
PS-3	RUE	24	0	24	51	-	-
PS-4	RUE	23	0	18	59	-	-
BEL-1	UE	17	12	30	41	-	Hematite
BEL-5	UE	19	14	37	29	-	-
BEL-6	UE	20	13	34	33	-	-
BL-1	UE	24	13	31	32	-	-
BL-13	UE	22	13	28	36	-	-
BL-14	UE	25	14	27	34	-	-
LR-5	UE	24	13	27	37	-	-
LR-7	UE	24	14	28	34	-	-
LR-8	UE	23	15	27	36	-	-

UCL and OCL Lower Cambrian, samples underlying and overlying, the carbonate level, respectively.

RUE: Rubefacted Upper Ediacaran.

UE: Upper Ediacaran.

**Table 2.** Model compositional changes (%) of element concentrations of the rubefacted Ediacaran shales from the Cantabrian Zone, N Spain. Change 1 and 2: % differences between RUE and UE median and mean values, respectively. Negative values indicate depletion.

	Median UE	Median RUE	Change 1	Mean UE	Mean RUE	Change 2
SiO <sub>2</sub>	60.45	60.46	0.02	59.84	61.02	1.98
TiO <sub>2</sub>	0.86	0.88	2.33	0.85	0.84	-1.39
Al <sub>2</sub> O <sub>3</sub>	18.14	18.81	3.69	18.32	18.36	0.20
Fe <sub>2</sub> O <sub>3</sub>	7.05	7.08	0.43	7.10	6.96	-1.99
MgO	2.96	2.09	-29.4	3.02	2.03	-32.8
CaO	0.43	0.11	-74.4	0.48	0.16*	-71.1
Na <sub>2</sub> O	2.22	0.07*	-96.9	2.22	0.10*	-95.5
K <sub>2</sub> O	3.60	5.73	59.2	3.77	5.63	49.5
P <sub>2</sub> O <sub>5</sub>	0.24	0.22	-8.33	0.24	0.20	-17.7
Rb	128	190	48.4	130	197	51.5
Cs	6.62	21.6	226	6.82	18.2	166
Be	3.00	2.62	-12.67	3.16	2.83	-10.3
Sr	67.0	15.7	-76.6	70.3	15.5	-77.9
La	40.7	35.7	-12.3	41.2	35.3	-14.3
Ce	81.6	71.3	-12.6	83.1	71.2	-14.3
Pr	10.2	8.70	-14.7	10.3	8.69	-16.0
Nd	39.1	34.0	-13.0	40.2	33.7	-16.1
Sm	7.95	7.49	-5.79	8.16	7.07	-13.4
Eu	1.58	1.58	0.00	1.67	1.47	-9.71
Gd	6.67	6.50	-2.55	6.86	6.22	-9.34
Tb	1.02	0.99	-2.94	1.04	0.96	-7.33
Dy	5.96	5.86	-1.68	6.09	5.74	-5.88
Ho	1.16	1.13	-2.59	1.20	1.13	-5.90
Er	3.24	3.36	3.70	3.35	3.26	-2.54
Tm	0.49	0.51	4.08	0.51	0.50	-1.94
Yb	3.33	3.41	2.40	3.44	3.35	-2.53
Lu	0.52	0.54	3.85	0.53	0.53	-0.93
Y	34.0	32.6	-4.12	34.8	33.1	-4.84
Sc	18.3	20.3	10.9	19.0	19.5	3.02
Th	12.8	11.1	-13.3	12.6	11.2	-10.9
U	3.60	3.21	-10.8	3.61	3.12	-13.5
Nb	12.7	13.0	2.36	12.4	12.0	-3.27
Co	19.0	14.6	-21.1	19.0	16.0	-14.4
Ni	49.0	41.0	-16.3	48.0	41.0	-14.4
Cu	41.0	4.22	-89.7	39.0	7.83*	-20.7
Zn	104	74.9	-27.9	107	71.0	-33.5
Ge	1.75	2.13	21.9	1.74	2.09	19.7
Pb	10.9	5.76	-47.2	14.9	5.30	-64.5
As	18.5	2.13	-88.5	20.0	3.03*	-84.8

UE: Ediacaran shales. RUE: Rubefacted UE. \*Calculated accepting detection limits as maximum abundances of the elements



**Table 3.** Synthesis of the mean element ratio values and standard deviations of the shales and sandstones from the Cantabrian and Central Iberian zones

Rock groups	1	2	3	4	5	6	7	8
Samples	(n=20)	(n=100)	(n=13)	(n=16)	UCC	(n=9)	(n=5)	(n=3)
Al <sub>2</sub> O <sub>3</sub> /TiO <sub>2</sub>	21.80(2.2)	20.48(1.43)	19.39(2.14)	18.85(1.03)	24.06	22.57(4.01)	24.46(2.87)	24.25(2.02)
K <sub>2</sub> O/Al <sub>2</sub> O <sub>3</sub>	0.223 (0.01)	0.215(0.02)	0.169(0.03)	0.166(0.02)	0.197	0.333(0.02)	0.370(0.01)	0.431(0.08)
Rb/Sr	1.93(0.37)	1.72(0.58)	0.45(0.13)	0.51(0.11)	0.29	13.75(3.85)	10.90(8.97)	3.21(0.64)
Cs/Be	2.18(0.48)	2.38(0.73)	1.67(0.33)	1.86(0.48)	2.58	6.69(3.37)	6.66(1.73)	3.03(1.38)
Ti/Zr	25.76(1.62)	26.08(1.86)	17.23(2.61)	16.90(3.12)	19.88	25.63(3.89)	24.34(2.82)	30.77(1.22)
Ti/Nb	411(21.80)	429(29.89)	440(25.22)	422(34.22)	331	418(47.38)	384(20.6)	379(19.0)
La/Sc	2.19(0.33)	1.93(0.31)	2.83(0.28)	nd	2.21	1.83(0.32)	2.13(0.16)	2.41(0.21)
Th/Sc	0.67(0.09)	0.62(0.10)	0.83(0.11)	nd	0.75	0.58(0.06)	0.68(0.05)	0.70(0.04)
La/Cr	0.43(0.03)	0.35(0.06)	0.38(0.03)	0.39(0.05)	0.42	0.38(0.07)	0.48(0.04)	0.50(0.04)
Cr/Th	7.69(0.76)	9.07(1.41)	9.09(0.81)	8.55(0.82)	6.95	8.39(1.28)	6.65(0.66)	6.92(0.37)
Zr/Hf	35.90(1.12)	36.96(1.46)	39.28(0.87)	38.56(2.49)	36.42	36.23(0.74)	35.80(0.86)	36.35(0.58)
Zr/Sc	10.48(1.08)	10.65(1.49)	22.36(4.63)	nd	10.9	10.11(1.29)	10.19(1.76)	7.82(0.68)
Zr/Nb	15.96(0.75)	16.49(1.23)	26.02(3.59)	25.54(3.65)	16.6	16.45(1.40)	15.96(2.36)	12.36(1.02)
Th/Nb	1.03(0.09)	0.96(0.10)	0.97(0.07)	0.94(0.09)	0.90	0.95(0.07)	1.06(0.08)	1.11(0.01)
(La/Yb) <sub>N</sub>	8.21(0.69)	7.44(0.99)	9.14(0.78)	7.84(1.05)	8.95	7.23(0.77)	7.33(0.83)	9.26(1.03)
Eu/Eu*	0.67(0.6)	0.70(0.05)	0.72(0.04)	0.73(0.03)	0.70	0.67(0.04)	0.62(0.04)	0.69(0.03)
Rb/Th	10.33(1.16)	10.01(1.18)	7.37(1.08)	8.04(1.36)	8.95	17.56(2.08)	17.01(1.34)	12.13(1.85)
Rb/Zr	0.66(0.08)	0.58(0.08)	0.28(0.05)	0.30(0.07)	0.49	1.01(0.11)	1.15(0.18)	1.09(0.19)

1: Ediacaran shales. 2: NIBAS (Ugidos *et al.* 2010), Sc: 72 analyses. 3: Ediacaran sandstones. 4 Ediacaran sandstones from the Central Iberian Zone (Ugidos *et al.* 1997b), nd: no data of Sc. 5: Upper Continental Crust (UCC, Rudnick & Gao 2005; Hu & Gao 2008). 6: Rubefacted shales. 7 and 8: Lower Cambrian shales underlying and overlying, respectively, the carbonate level (see text). N: chondrite (Pourmand *et al.* 2012) normalised.

**Table 4.** Sm and Nd contents (ppm) and Nd isotope data of the Ediacaran and Lower Cambrian shales.

Samples No.	Sm	Nd	$^{147}\text{Sm}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}$	2se	$\epsilon\text{Nd}(t)$	$f_{\text{Sm}/\text{Nd}}$	TDM (Ga)
<b>OLC</b>								
BEL-2*	8.52	45.8	0.1125	0.512041	6	-5.8	-0.43	1.61
LR-3*	8.07	40.8	0.1196	0.512268	5	-4.9	-0.39	1.62
BEL-9	7.97	43.4	0.1111	0.512069	11	-5.1	-0.44	1.55
<b>ULC</b>								
BL-3*	9.10	42.3	0.1301	0.512268	6	-2.6	-0.34	1.54
BEL-3*	6.44	34.2	0.1140	0.512236	6	-2.1	-0.42	1.35
LR-16*	8.43	41.5	0.1229	0.512199	5	-3.4	-0.38	1.54
PS-2*	8.23	45.7	0.1087	0.511798	6	-2.8	-0.45	1.36
<b>RUE</b>								
LR-14*	7.53	35.1	0.1298	0.512232	7	-3.3	-0.34	1.60
BEL-7*	4.52	23.5	0.1161	0.512196	5	-3.1	-0.41	1.44
PS-4*	5.82	28.5	0.1234	0.512192	6	-3.6	-0.37	1.56
<b>UE</b>								
BEL-1	9.53	46.8	0.1232	0.512195	6	-3.6	-0.37	1.55
BEL-5*	8.61	42.3	0.1231	0.512284	5	-1.8	-0.37	1.40
BEL-11*	9.91	52.5	0.1142	0.512238	7	-2.1	-0.42	1.35
CN-1	8.97	41.9	0.1295	0.512244	7	-3.0	-0.34	1.58
BL-11	7.68	38.8	0.1197	0.512191	3	-3.4	-0.39	1.50
LR-1	7.70	39.3	0.1185	0.512214	7	-2.9	-0.40	1.45
VIL-1	6.93	33.9	0.1237	0.512235	8	-2.8	-0.37	1.49
BL-1	8.06	41.5	0.1176	0.512215	7	-2.8	-0.40	1.43
BL-2	6.67	34.6	0.1166	0.512228	7	-2.5	-0.41	1.40

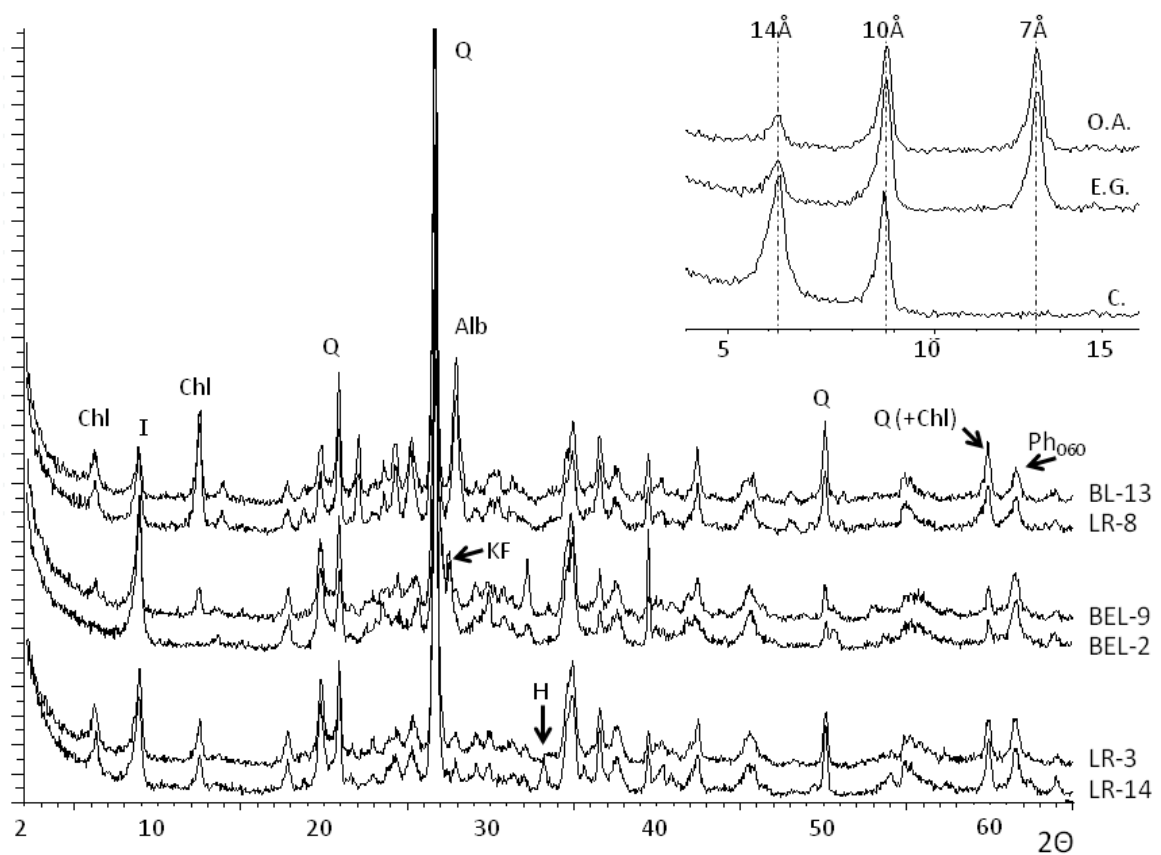
Nd isotope analyses at SGIker-Geochronology and Isotopic Geochemistry (\*), University of the Basque Country (Spain) and at SUERC, East Kilbride (UK). OCL and UCL: Lower Cambrian, samples overlying and underlying, respectively, the carbonate level. See text. RUE: Rubefacted Upper Ediacaran. UE: Upper Ediacaran.



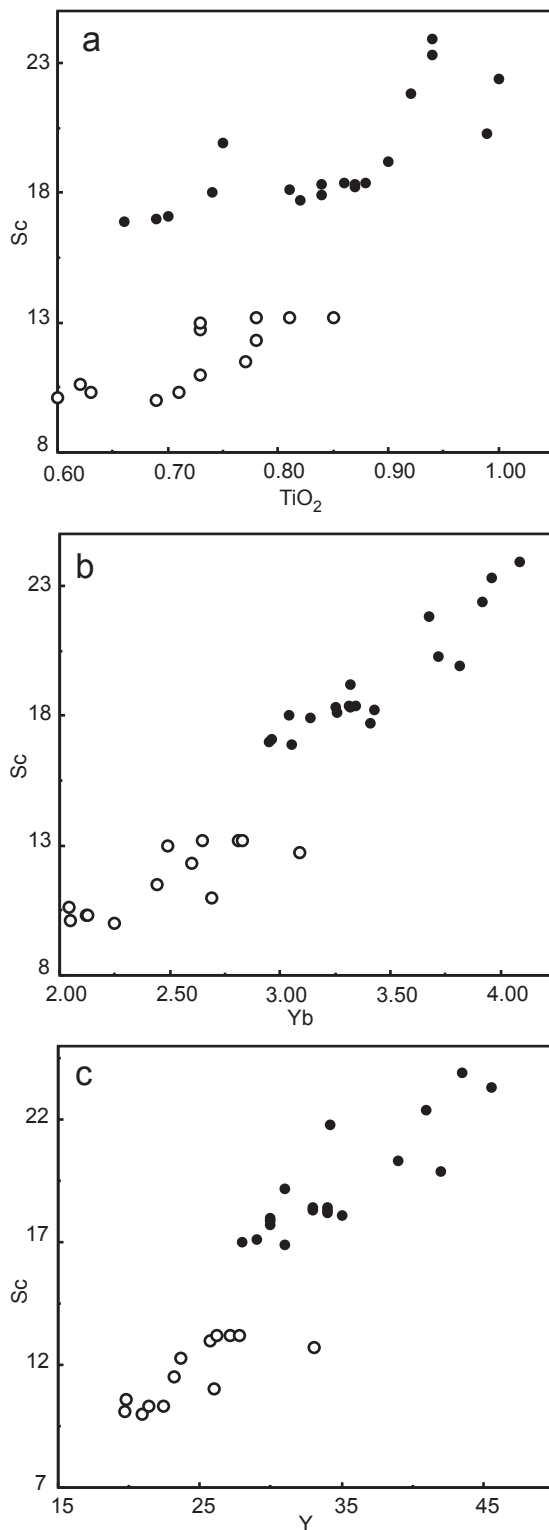
**Supplementary Fig. 1.** Representative diffractograms of the shales studied showing the main mineralogical results.

Chl: chlorite; I: illite; Ph: phyllosilicates; Q: quartz; Alb: albite; H: Hematite; KF: K-feldspar; Ph060: 060 reflection of phyllosilicates.

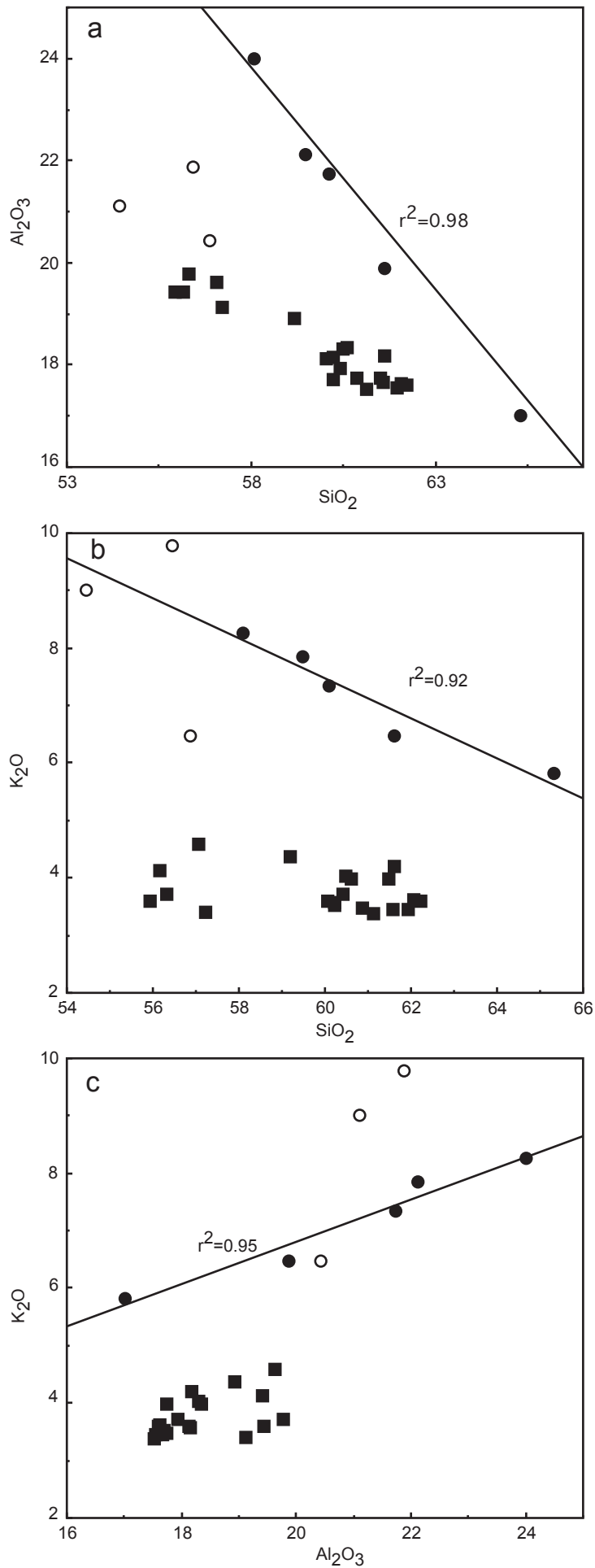
Top right corner: Detail of the XRD patterns of the <2 $\mu$ m fraction of a representative sample (LR-8) containing chlorite (14Å and 7Å) and illite (10 Å). OA: oriented aggregates; EG: treated with ethylene-glycol; C: calcinated.



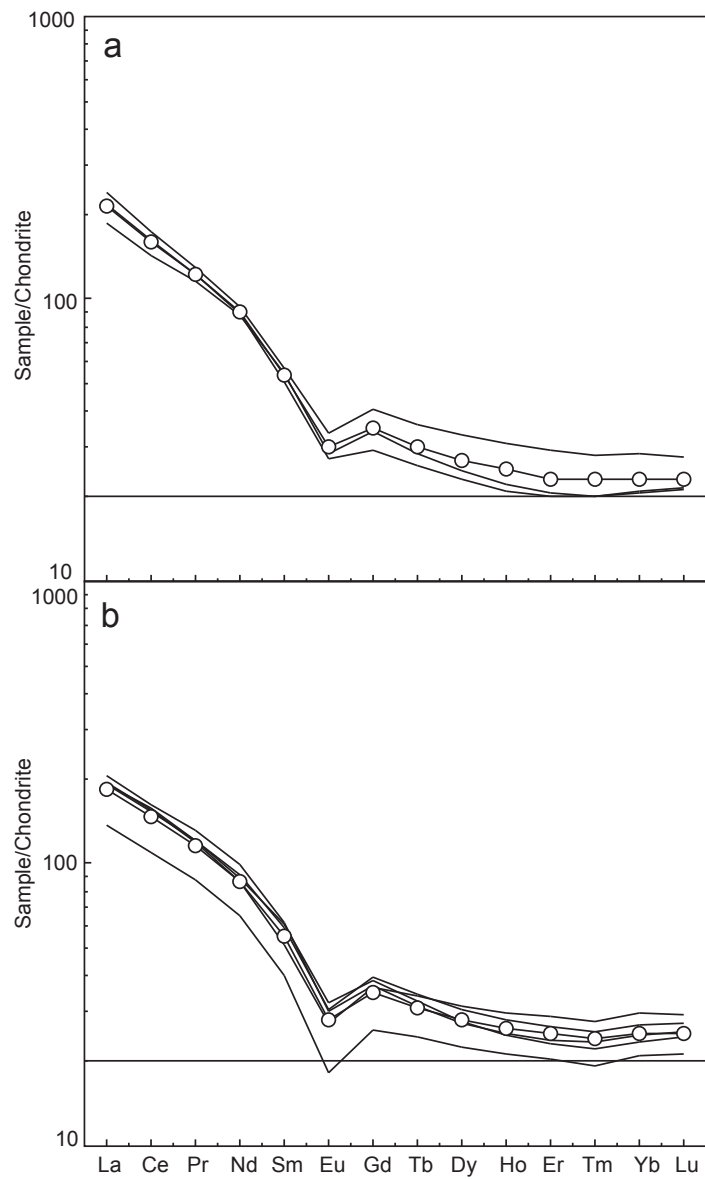
**Supplementary Fig. 2.** Chondrite-normalised REE patterns for the Ediacaran shales (a) and sandstones (b). Open circles: mean patterns. (c) REE patterns of the mean Ediacaran shales (filled squares) and sandstones (open squares) from the Cantabrian Zone, NIBAS (open triangle) and the mean (n=16) sandstone (filled triangles) of the Central Iberian Zone. Normalising chondrite values after Pourmand *et al.* 2012. Symbols as in Fig. 3.



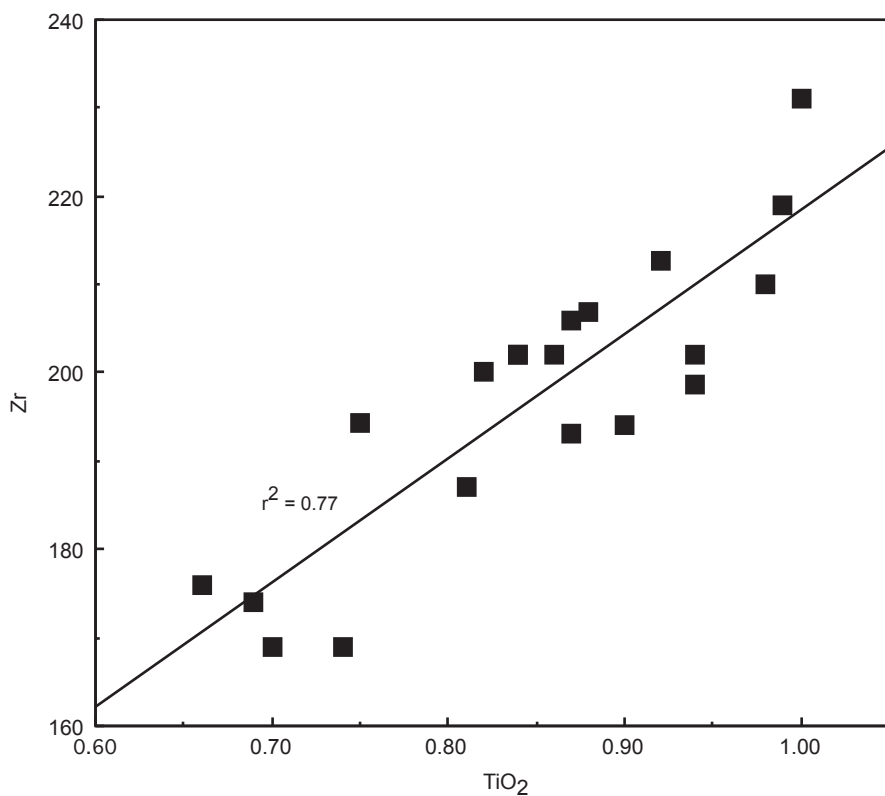
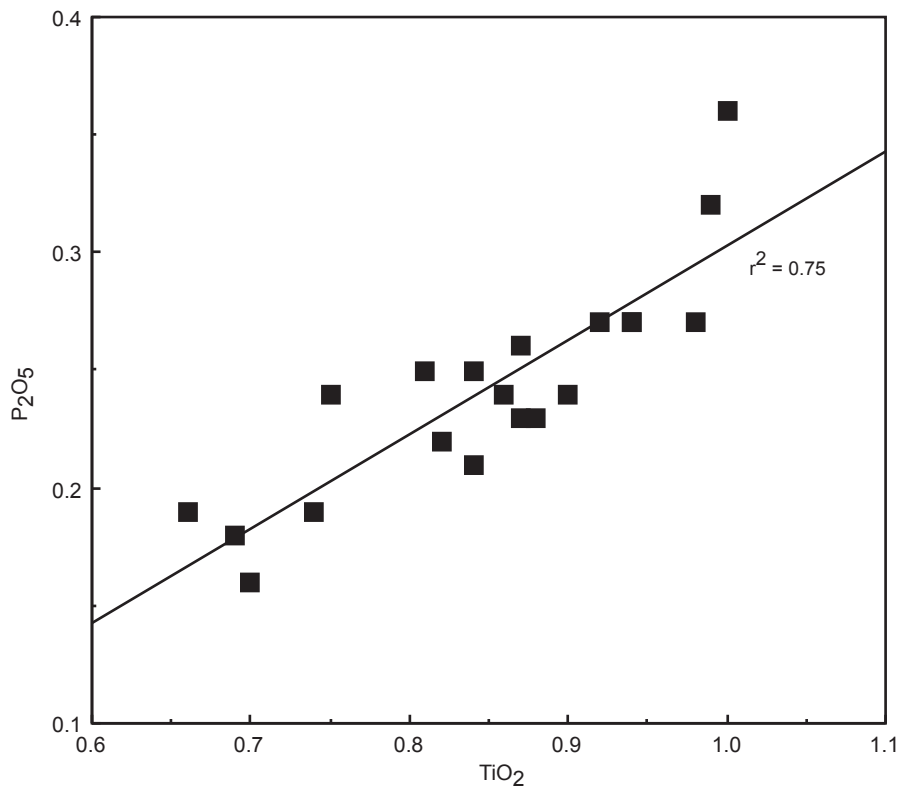
**Supplementary Fig. 3:** The covariations of SiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub>, SiO<sub>2</sub>-K<sub>2</sub>O and Al<sub>2</sub>O<sub>3</sub>-K<sub>2</sub>O suggest that illiteAl<sub>2</sub>O<sub>3</sub>2O<sub>3</sub> would be the main K-mineral in the underlying shales. Symbols as in Fig. 3.



**Supplementary Fig. 4.** Chondrite-normalised REE patterns for the Lower Cambrian shales overlying (a) and underlying (b) the carbonate level. See text. Symbols as in Fig. 3.



**Supplementary Fig. 5:** Diagrams TiO<sub>2</sub>-P<sub>2</sub>O<sub>5</sub> and TiO<sub>2</sub>-Zr for the Ediacaran shales. The positive covariations of these elements probably reflect similar proportions of Ti-minerals, phosphates and zircon in all samples. Symbols as in Fig. 3.



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